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Emergence of **Power Law** in Statistical Hadronization

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Content

- 1 Temperature and energy do fluctuate
- 2 Finite Heat Bath and Fluctuation Effects
- 3 RHIC and LHC spectra

Content

- 1 Temperature and energy do fluctuate
 - Gauss approximation
 - Gaussian is insufficient
 - Beta- and Gamma-distribution
- 2 Finite Heat Bath and Fluctuation Effects
- 3 RHIC and LHC spectra

Theoretical equation of state: $S(E)$

Product of the spreads of energy and temperature

$$\Delta E \cdot \Delta \beta = 1 \quad (1)$$

Connection to the (absolute) temperature:

$C = dE/dT$ heat capacity, $\beta = 1/T$

$$|C| \Delta T \cdot \frac{\Delta T}{T^2} = 1 \quad (2)$$

The relative spread in temperature is the one over square root of the heat capacity!

$$\frac{\Delta T}{T} = \frac{\Delta \beta}{\beta} = \frac{1}{\sqrt{|C|}} \quad (3)$$

The heat capacity C is proportional to the size of the heat bath – mostly.

Gauss distributed reciprocal temperature, β

$$w(\beta) = \frac{1}{\sqrt{2\pi}\sigma} e^{-\frac{(\beta-1/T_0)^2}{2\sigma^2}} \quad (4)$$

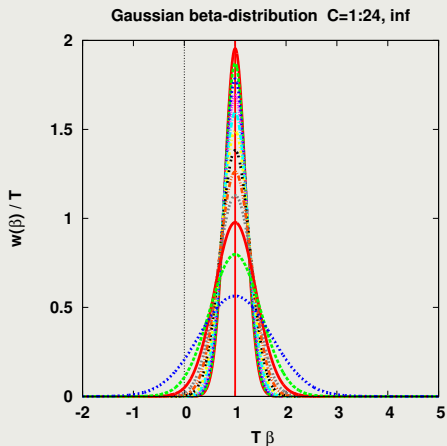
Average

$$\langle \beta \rangle = \frac{1}{T_0}$$

Spread (square root of variance)

$$\Delta\beta = \sigma = \frac{1}{T_0\sqrt{|C|}}$$

Gaussian Fluctuations; Figure



Superstatistics: single particle energy distribution

Canonical distribution in additive thermodynamics:

$$p_i = p(E_i) = e^{\beta(\mu - E_i)}. \quad (5)$$

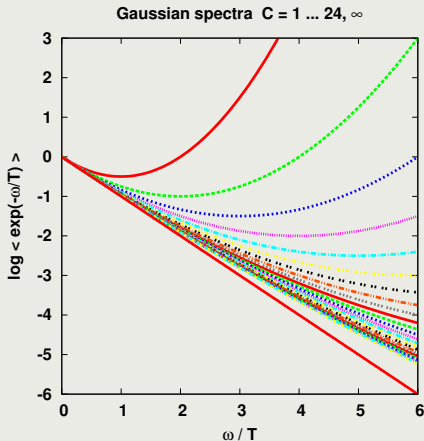
If β fluctuates according to Gauss, then the exponential weight factor averages to the **characteristic function**

$$\langle e^{-\beta\omega} \rangle = e^{-\omega/T_0} e^{\sigma^2\omega^2/2}. \quad (6)$$

Turning point: the largest single particle energy, where this can make a sense...

$$E_i^{\max} - \mu = \omega^{\max} = \frac{1}{\sigma^2 T_0} = |C| T_0. \quad (7)$$

Figure on Thermal Spectra with Gaussian β -distribution



Deficiencies of the Gaussian Picture

- 1 There is a finite probability, $w(\beta) > 0$, for $\beta < 0$
- 2 $\langle e^{-\beta\omega} \rangle$ does not diminish for large ω (this cannot be a canonical spectrum)

Ideal Gas: Thermodynamics

EoS

$$p = nT, \quad e = \frac{1}{3}p.$$

Heat Capacity:

$$E = \frac{1}{3}pV = \frac{1}{3}NT; \quad C = \frac{dE}{dT} = \frac{1}{3}N.$$

fix N : $C(T)$ constant; $C(S)$ constant; We have to solve:

$$\frac{dT}{dE} = \left(\frac{1}{S'} \right)' = -\frac{S''}{(S')^2} = \frac{1}{C}. \quad (8)$$

Ideal Gas: Thermodynamics

Constant Heat Capacity

$$-\frac{S''}{(S')^2} = \frac{1}{C}.$$

Integrals: temperature and entropy

$$T = \frac{1}{S'} = \frac{E}{C} + T_0, \quad S = C \ln \left(1 + \frac{E}{CT_0} \right) + S_0.$$

Mutual info based probability

(phase volume product)

$$\mathfrak{P}(E_1) = e^{S_1 + S_2 - S_{12}} \propto \left(1 + \frac{E_1}{C_1 T_0} \right)^{C_1} \left(1 + \frac{E - E_1}{C_2 T_0} \right)^{C_2}$$

Ideal Gas: temperature distribution

It is an **Euler-Beta distribution**

$$\mathfrak{P}(T_1) \propto T_1^{C_1} \left(T_* - \frac{C_1}{C_2}(T_1 - T_*) \right)^{C_2}$$

in the scaling variable: $x = C_1 T_1 / (C_1 + C_2) T_* = C_* T_1 / C_2 T_*$

$$\mathfrak{B}(x) = \frac{\Gamma(C_1 + C_2 + 2)}{\Gamma(C_1 + 1)\Gamma(C_2 + 1)} x^{C_1} (1 - x)^{C_2}$$

Beta distribution in x , binomial in C_1 at fix $C_1 + C_2$, NBD at fix C_2 .

Ideal Gas: limits

- Huge reservoir ($C_2 \rightarrow \infty$: with $t = C_1 T_1 / T_*$)
Euler-Gamma

$$\lim_{C_2 \rightarrow \infty} \mathfrak{B}(x) dx = \frac{1}{\Gamma(C_1 + 1)} t^{C_1} e^{-t} dt.$$

Euler fitted to Gaussian Uncertainty

$$\text{Average: } \langle \beta \rangle = \frac{v}{a} = \frac{1}{T}, \quad \text{Spread: } \frac{\Delta\beta}{\langle \beta \rangle} = \frac{1}{\sqrt{v}} = \frac{\Delta T}{T} = \frac{1}{\sqrt{|C|}}$$

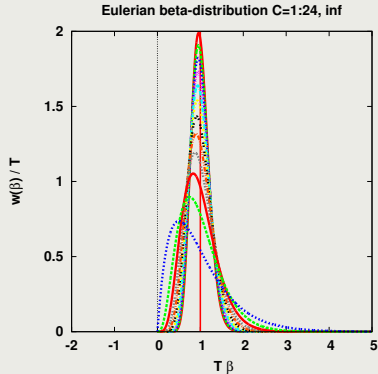
The corresponding Euler-Gamma distribution for $\beta = 1/T_*$:

$$w(\beta) = \frac{(|C|T)^{|C|}}{\Gamma(|C|)} \beta^{|C|-1} e^{-|C|T\beta}. \quad (9)$$

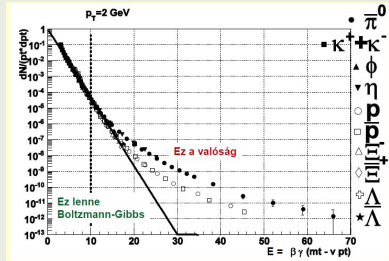
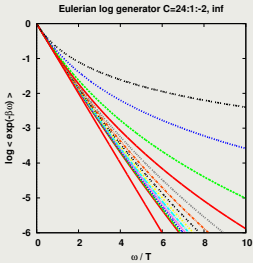
Characteristic function = **spectrum**

$$\langle e^{-\beta\omega} \rangle = \left(1 + \frac{\omega}{|C|T} \right)^{-|C|} \xrightarrow{|C| \rightarrow \infty} e^{-\omega/T}. \quad (10)$$

Plot Eulerian Fluctuations



Plot Eulerian Spectra and RHIC results as blast wave



Content

- 1 Temperature and energy do fluctuate
- 2 **Finite Heat Bath and Fluctuation Effects**
 - The physics behind the power law
 - Particle Number Fluctuations
- 3 RHIC and LHC spectra

Finite Reservoirs

- Avogadro number (atoms in classical matter) $\sim 10^{24}$
- Neurons in human brain $\sim 10^{12}$
- Internet users in 2014 $\sim 10^9$
- New particles from heavy ion collisions $\sim 10^3$
- From elementary high energy collisions (pp) $\sim 10^1$

General expectation:

smaller size \rightarrow **larger *relative* fluctuations.**

Ideal Gas: microcanonical statistical weight

The one-particle energy, ω , out of total energy, E , is distributed in a one-dimensional relativistic jet according to a statistical weight factor which depends on the number of particles in the reservoir, n :

$$P_1(\omega) = \frac{\Omega_1(\omega) \Omega_n(E - \omega)}{\Omega_{n+1}(E)} = \rho(\omega) \cdot \frac{(E - \omega)^n}{E^n} \quad (11)$$

HEP Superstatistics: E fix, n has a distribution (based on the physical model of the reservoir and on the event by event detection of the spectra).

Phase Volume Ratio is a $q < 1$ Tsallis–Pareto

Thermodynamic limit:

$$\lim_{n \rightarrow \infty} \lim_{E \rightarrow nT} \left(1 - \frac{\omega}{E}\right)^n = e^{-\omega/T}. \quad (12)$$

Compare with Tsallis distribution:

$$\left(1 - \frac{\omega}{E}\right)^n = \left(1 + (q-1)\frac{\omega}{T}\right)^{-\frac{1}{q-1}}, \quad (13)$$

If and only if

$$T = \frac{E}{n}, \quad q = 1 - \frac{1}{n}. \quad (14)$$

Ideal Reservoir: (Negative) binomial n -distribution

n particles among k cells: bosons $\binom{n+k}{n}$ fermions $\binom{k}{n}$

A subspace (n, k) out of (N, K)

Limit: $K \rightarrow \infty, N \rightarrow \infty$; average occupancy $f = N/K$ is fixed.

$$B_{n,k}(f) := \lim_{K \rightarrow \infty} \frac{\binom{n+k}{n} \binom{N-n+K-k}{N-n}}{\binom{N+K+1}{N}} = \binom{n+k}{n} f^n (1+f)^{-n-k-1}. \quad (15)$$

$$F_{n,k}(f) := \lim_{K \rightarrow \infty} \frac{\binom{k}{n} \binom{K-k}{N-n}}{\binom{K}{N}} = \binom{k}{n} f^n (1-f)^{k-n}. \quad (16)$$

Norm and Pascal triangle

Binomial expansion:

$$(a + b)^k = \sum_{n=0}^{\infty} \binom{k}{n} a^n b^{k-n} \quad (17)$$

Replace k by $-k - 1$ and a by $-a$, noting that

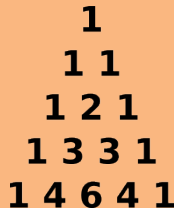
$$\binom{-k-1}{n} = \frac{(-k-1)(-k-2)\dots(-k-n)}{n!} = (-1)^n \frac{(k+1)(k+2)\dots(k+n)}{n!} = (-1)^n \binom{n+k}{n}.$$

we arrive at

$$(b - a)^{-k-1} = \sum_{n=0}^{\infty} \binom{n+k}{n} a^n b^{-n-k-1} \quad (18)$$

BD recursion: Pascal Triangle

$$F_{n,k} = f F_{n-1,k-1} + (1-f) F_{n,k-1} \quad (19)$$



NBD recursion: Tilted Pascal Triangle

$$B_{n,k} = \frac{f}{1+f} B_{n-1,k} + \frac{1}{1+f} B_{n,k-1} \quad (20)$$

1 1 1 1 1
1 2 3 4
1 3 6
1 4
1

Bosonic reservoir

Reservoir in hep: E is fixed, n fluctuates according to NBD.

$$\sum_{n=0}^{\infty} \left(1 - \frac{\omega}{E}\right)^n B_{n,k}(f) = \left[(1+f) - f \left(1 - \frac{\omega}{E}\right) \right]^{-k-1} = \left(1 + f \frac{\omega}{E}\right)^{-k-1} \quad (21)$$

Note that $\langle n \rangle = (k+1)f$ for NBD. Then with $T = E / \langle n \rangle$ and $q - 1 = \frac{1}{k+1}$ we get

$$\left(1 + (q-1) \frac{\omega}{T}\right)^{-\frac{1}{q-1}}$$

This is **exactly** a $q > 1$ Tsallis-Pareto distribution.

Fermionic reservoir

E is fixed, n is distributed according to BD:

$$\sum_{n=0}^{\infty} \left(1 - \frac{\omega}{E}\right)^n F_{n,k}(f) = \left[(1-f) + f \left(1 - \frac{\omega}{E}\right) \right]^k = \left(1 - f \frac{\omega}{E}\right)^k \quad (22)$$

Note that $\langle n \rangle = kf$ for BD. Then with $T = E / \langle n \rangle$ and $q - 1 = -\frac{1}{k}$ we get

$$\left(1 + (q - 1) \frac{\omega}{T}\right)^{-\frac{1}{q-1}}$$

This is **exactly** a $q < 1$ Tsallis-Pareto distribution.

Boltzmann limit

In the $k \gg n$ limit (low occupancy in phase space)

$$\begin{aligned} \binom{n+k}{n} f^n (1+f)^{-n-k-1} &\rightarrow \frac{k^n}{n!} \left(\frac{f}{1+f}\right)^n \dots \\ \binom{k}{n} f^n (1-f)^{k-n} &\rightarrow \frac{k^n}{n!} \left(\frac{f}{1-f}\right)^n \dots \end{aligned} \quad (23)$$

After normalization this is the **Poisson** distribution:

$$\Pi_n = \frac{\langle n \rangle^n}{n!} e^{-\langle n \rangle} \quad \text{with} \quad \langle n \rangle = k \frac{f}{1 \pm f} \quad (24)$$

The result is **exactly** the Boltzmann-Gibbs weight factor:

$$\sum_{n=0}^{\infty} \left(1 - \frac{\omega}{E}\right)^n \Pi_n(\langle n \rangle) = e^{-\omega/T}. \quad (25)$$

Summary of reservoir fluctuation models

$$\begin{aligned}\left\langle \left(1 - \frac{\mathcal{E}}{E}\right)^n \right\rangle_{\text{Bernoulli}} &= \text{Tsallis}(q < 1) \\ \left\langle \left(1 - \frac{\mathcal{E}}{E}\right)^n \right\rangle_{\text{Poisson}} &= \text{Boltzmann}(q = 1) \\ \left\langle \left(1 - \frac{\mathcal{E}}{E}\right)^n \right\rangle_{\text{NBD}} &= \text{Tsallis}(q > 1)\end{aligned}\quad (26)$$

In all the three above cases

$$T = \frac{E}{\langle n \rangle}, \quad \text{and} \quad q = \frac{\langle n(n-1) \rangle}{\langle n \rangle^2}\quad (27)$$

Ideal gas with general n -fluctuations

Canonical approach: expansion for small $\omega \ll E$.

Tsallis-Pareto distribution as an approximation:

$$\left(1 + (q-1)\frac{\omega}{T}\right)^{-\frac{1}{q-1}} = 1 - \frac{\omega}{T} + q\frac{\omega^2}{2T^2} - \dots \quad (28)$$

Ideal reservoir phase space up to the subleading canonical limit:

$$\left\langle \left(1 - \frac{\omega}{E}\right)^n \right\rangle = 1 - \langle n \rangle \frac{\omega}{E} + \langle n(n-1) \rangle \frac{\omega^2}{2E^2} - \dots \quad (29)$$

To subleading in $\omega \ll E$

$$T = \frac{E}{\langle n \rangle}, \quad q = \frac{\langle n(n-1) \rangle}{\langle n \rangle^2} = 1 - \frac{1}{\langle n \rangle} + \frac{\Delta n^2}{\langle n \rangle^2}. \quad (30)$$

General system with general reservoir fluctuations

Canonical approach: expansion for small $\omega \ll E$.

$$\begin{aligned} \left\langle \frac{\Omega_n(E-\omega)}{\Omega_n(E)} \right\rangle &= \left\langle e^{S(E-\omega)-S(E)} \right\rangle = \left\langle e^{-\omega S'(E) + \omega^2 S''(E)/2 - \dots} \right\rangle \\ &= 1 - \omega \langle S'(E) \rangle + \frac{\omega^2}{2} \langle S'(E)^2 + S''(E) \rangle - \dots \end{aligned} \quad (31)$$

Compare with expansion of Tsallis

$$\left(1 + (q-1)\frac{\omega}{T}\right)^{-\frac{1}{q-1}} = 1 - \frac{\omega}{T} + q\frac{\omega^2}{2T^2} - \dots \quad (32)$$

Interpret the parameters

$$\frac{1}{T} = \langle S'(E) \rangle, \quad q = 1 - \frac{1}{C} + \frac{\Delta T^2}{T^2} \quad (33)$$

$$\langle S''(E) \rangle = -1/CT^2$$

expressed via the heat capacity of the reservoir, $1/C = dT/dE$

Understanding the parameter q

in terms fluctuations

Opposite sign contributions from $\langle S'^2 \rangle - \langle S' \rangle^2$ and from $\langle S'' \rangle$.
In all cases approximately

$$q = 1 - \frac{1}{C} + \frac{\Delta T^2}{T^2}.$$

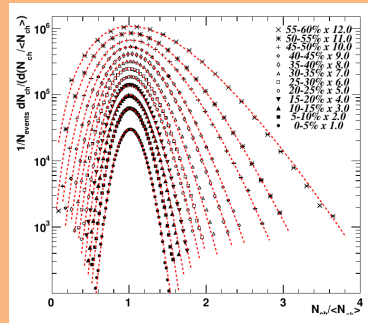
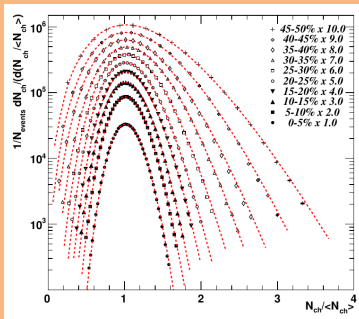
- $q > 1$ and $q < 1$ are both possible
- for any relative variance $\Delta T/T = 1/\sqrt{C}$ it is exactly $q = 1$
- for $nT = E/\text{dim} = \text{const}$ it is $\Delta T/\langle T \rangle = \Delta n/\langle n \rangle$.
- for ideal gas and n distributed as NBD or BD, the Tsallis form is exact

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Experimental NBD distributions PHENIX PRC 78 (2008) 044902

Au + Au collisions at $\sqrt{s_{NN}} = 62$ (left) and 200 GeV (right). Total charged multiplicities.



$$k \approx 10 - 20 \quad \rightarrow \quad q \approx 1.05 - 1.10.$$

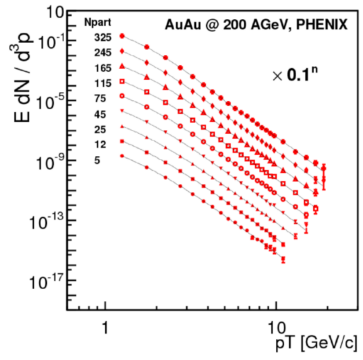
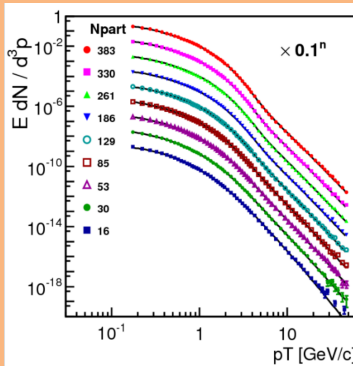
Statistical vs QCD power-law

The experimental fact for hadrons is **NBD!**

- QCD power-law: size-independent power $(k + 1) > 4$
- statistical power: $(k + 1) = \langle n \rangle / f \propto$ reservoir size
- consider data fits: $k + 1$ powers vs N_{part}
- soft and hard power-laws should differ for large N_{part}

Soft and Hard Tsallis fits:

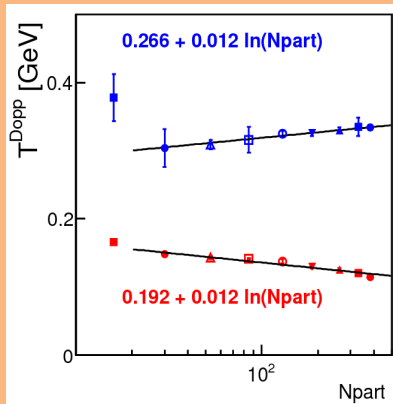
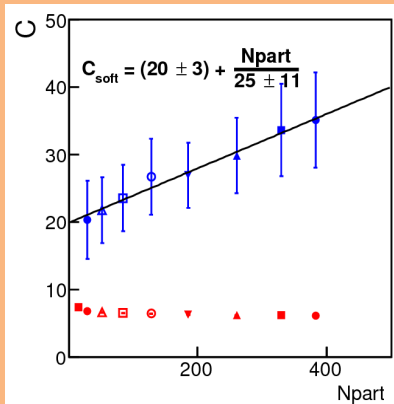
ALICE PLB 720 (2013) 52; PHENIX PRL 101 (2008) 232301



the knick is around $p_T \approx 4 - 5$ GeV.

Hard and Soft Trends with N_{part}

arxiv: 1405.3963

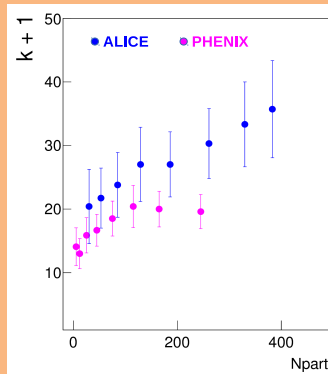


$C = k + 1$ powers of the power law and fitted T parameters (ALICE).

Soft Powers vs N_{part}

arxiv: 1405.3963

Only the soft ("statistical") branches for **PHENIX** and **ALICE**:



Summary

- The (measured and calculated) temperature surely fluctuates, but is **not Gaussian**.
- Ideal Gas prefers the **Beta or Gamma** distribution for T estimators.
- **NBD** particle number fluctuations generate **exact Tsallis** distribution with $q > 1$.