

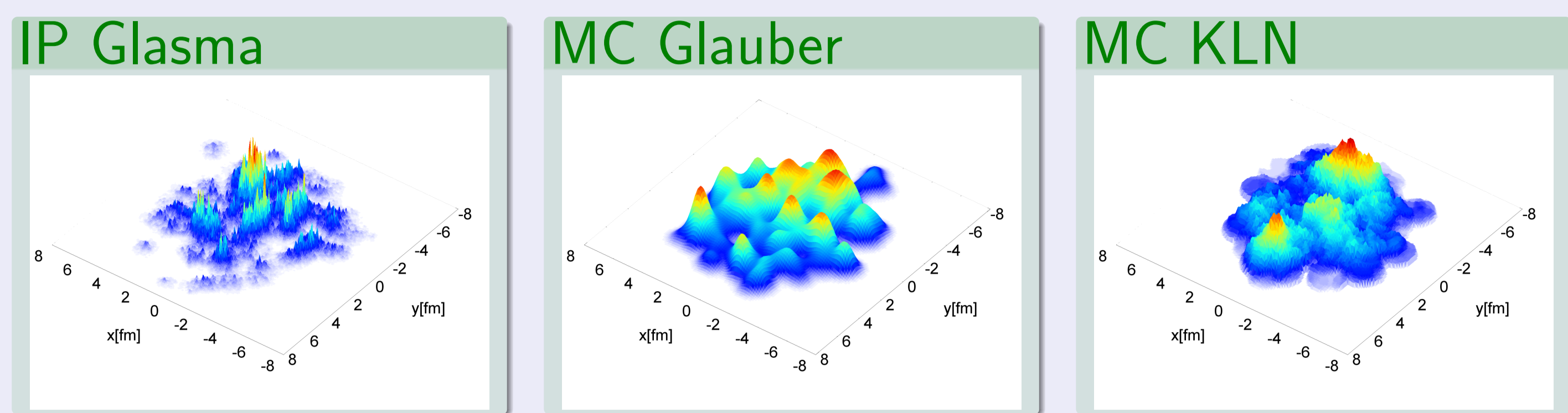
Non-Gaussian eccentricity fluctuations

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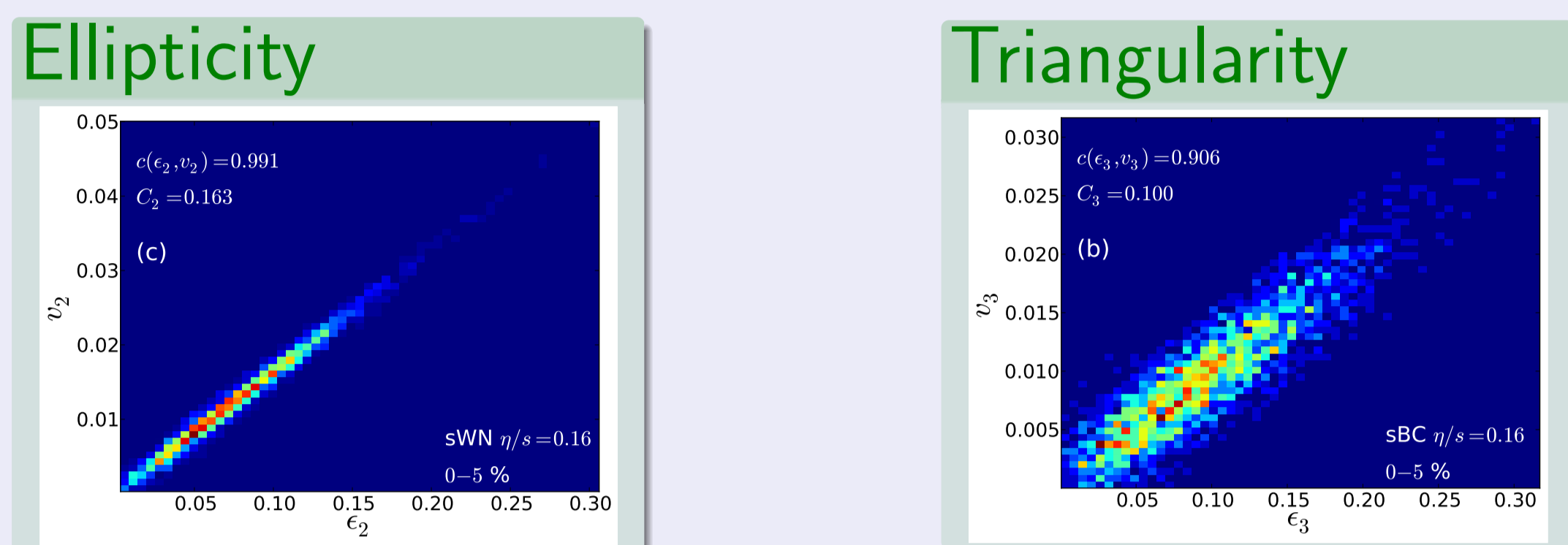
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1. Motivation

- Fluctuations in the initial density profile result in flow fluctuations [1].



- Anisotropic flow is approximately proportional to the initial anisotropy [2].



- Therefore the statistics of v_n fluctuations observed in experiments tell us about ε_n fluctuations. In particular, if ε_n fluctuations are Gaussian, one expects Gaussian v_n fluctuations [3].

- The second and fourth cumulants are

$$\varepsilon_n\{2\}^2 \equiv \langle |\varepsilon_n|^2 \rangle$$

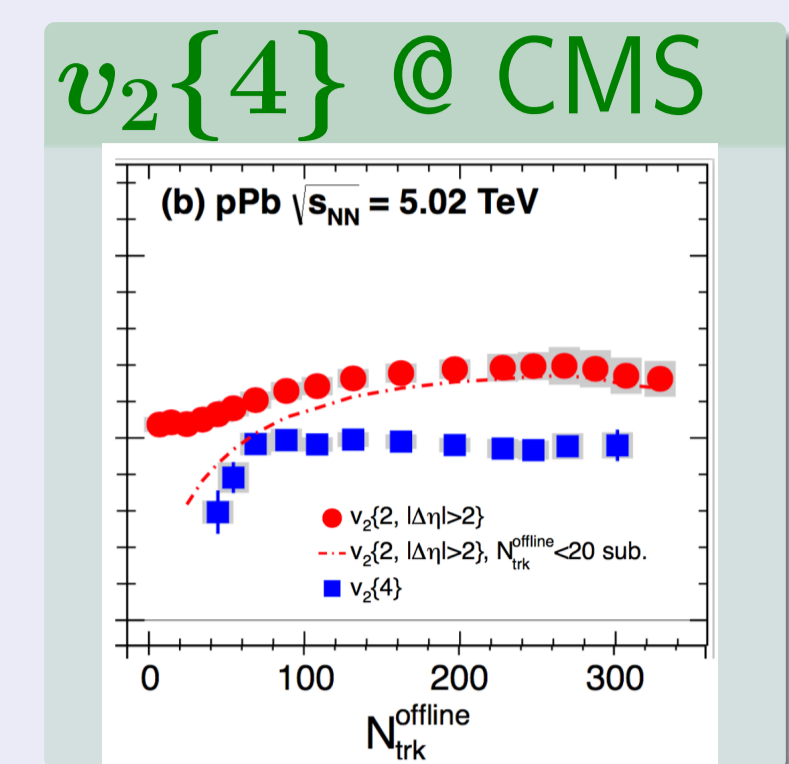
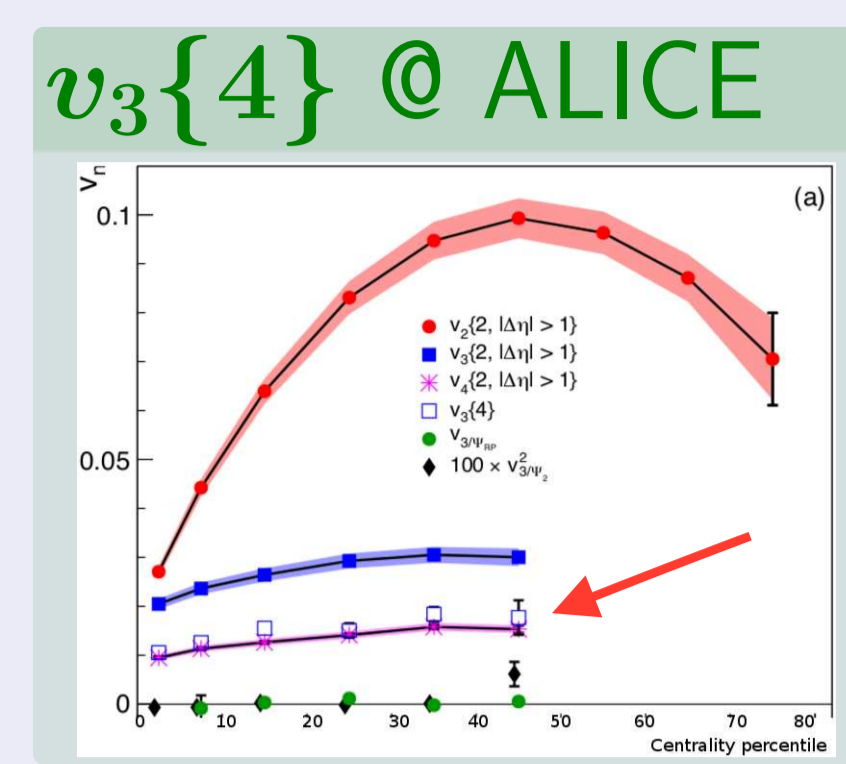
$$\varepsilon_n\{4\}^4 \equiv 2 \langle |\varepsilon_n|^2 \rangle^2 - \langle |\varepsilon_n|^4 \rangle,$$

- We consider only fluctuation-driven anisotropies for simplicity (no reaction plane eccentricity). We note that $\varepsilon_n\{4\} = 0$ for Gaussian fluctuations.

2. Measurements

We know that v_n fluctuations are **not** Gaussian:

- non-zero $v_3\{4\}$ in Pb+Pb [4]
- non-zero $v_2\{4\}$ in p+Pb [5]



Therefore ε_n fluctuations are not Gaussian. I study these non-Gaussianities systematically in order to figure out what we can learn from them.

3. State of the art

- Eccentricity fluctuations have been studied in the "independent source model" where the initial energy density profile is a superposition of N identical, pointlike sources. Analytic results have been obtained for $N \gg 1$ [6].

$$\left. \begin{aligned} \varepsilon_n\{2\} &\propto N^{-1/2} \\ \varepsilon_n\{4\} &\propto N^{-3/4} \end{aligned} \right\} \Rightarrow \varepsilon_n\{4\} \propto \varepsilon_n\{2\}^{3/2},$$

where the proportionality coefficient depends on the energy density profile.

- It has been argued that the non-Gaussianity of eccentricity fluctuations is mostly driven by the condition $|\varepsilon_n| < 1$ [7]. A parameterization of the distribution of $|\varepsilon_n|$ has been proposed which satisfies this constraint and which predicts

$$\varepsilon_n\{4\} \simeq 2^{1/4} \varepsilon_n\{2\}^{3/2}.$$

- There is a contradiction between a) (not universal) and b) (universal).

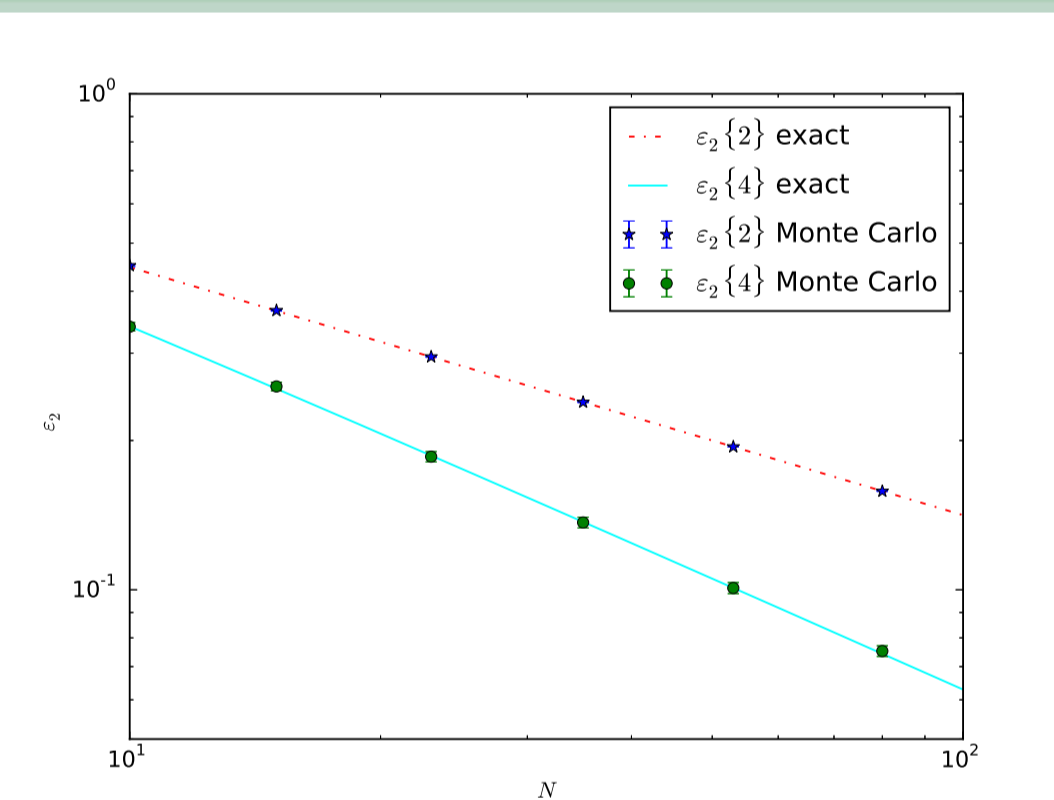
4. My work

Monte Carlo simulations

N independent, pointlike sources. We test the effect of changing

- the density profile in the transverse plane (2d Gaussian versus uniform in a disk)
- the energy of each source (identical sources versus negative binomial fluctuations fitted to LHC data)

Gaussian distribution, identical sources



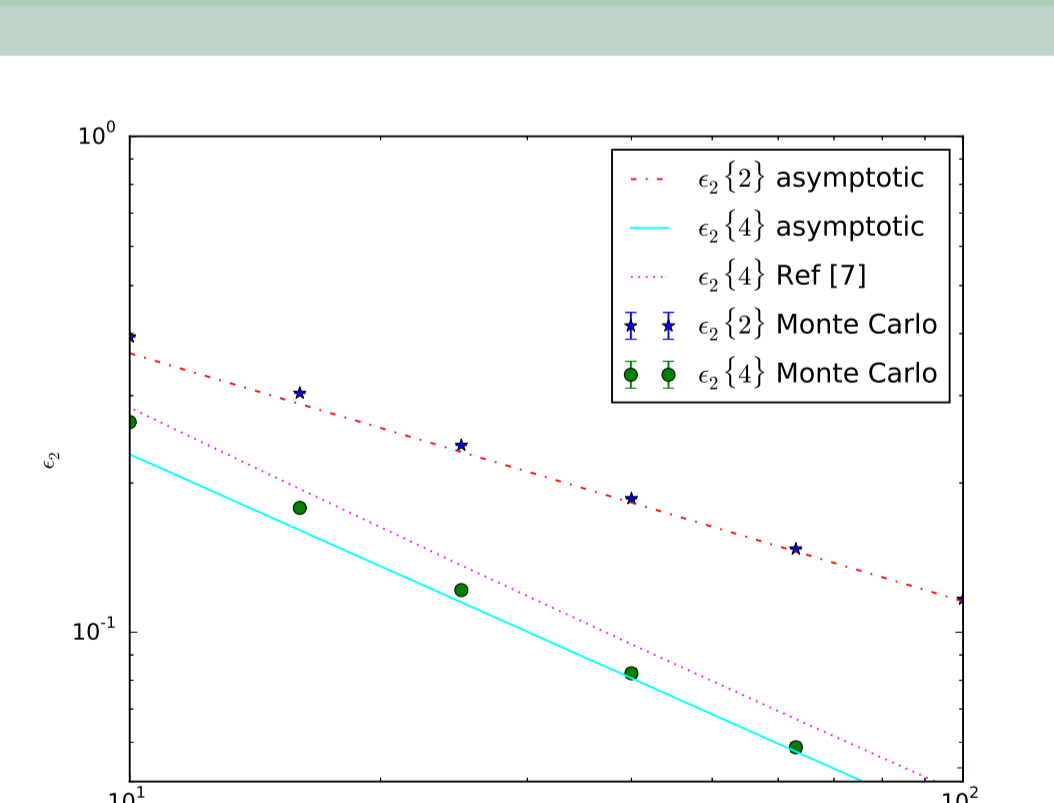
- exactly solvable

$$\varepsilon_2\{2\} = \sqrt{2/N}$$

$$\varepsilon_2\{4\} = \left(\frac{16}{N^2(N+2)} \right)^{1/4}$$

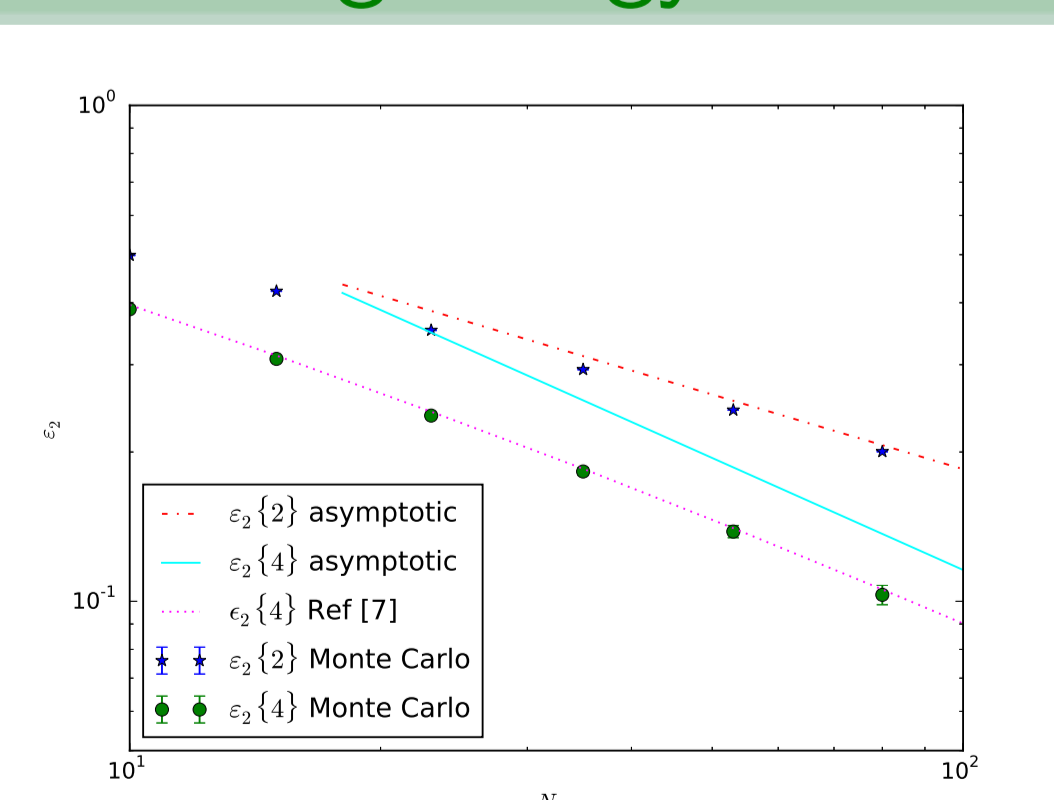
MC simulations fit the exact result perfectly.

Uniform distribution, identical sources



- The universal statistics from Ref. [7] overestimates the non-Gaussianity $\varepsilon_2\{4\}$. The asymptotic result from Ref. [6] is better for $N > 15$.

Uniform distribution, fluctuating energy

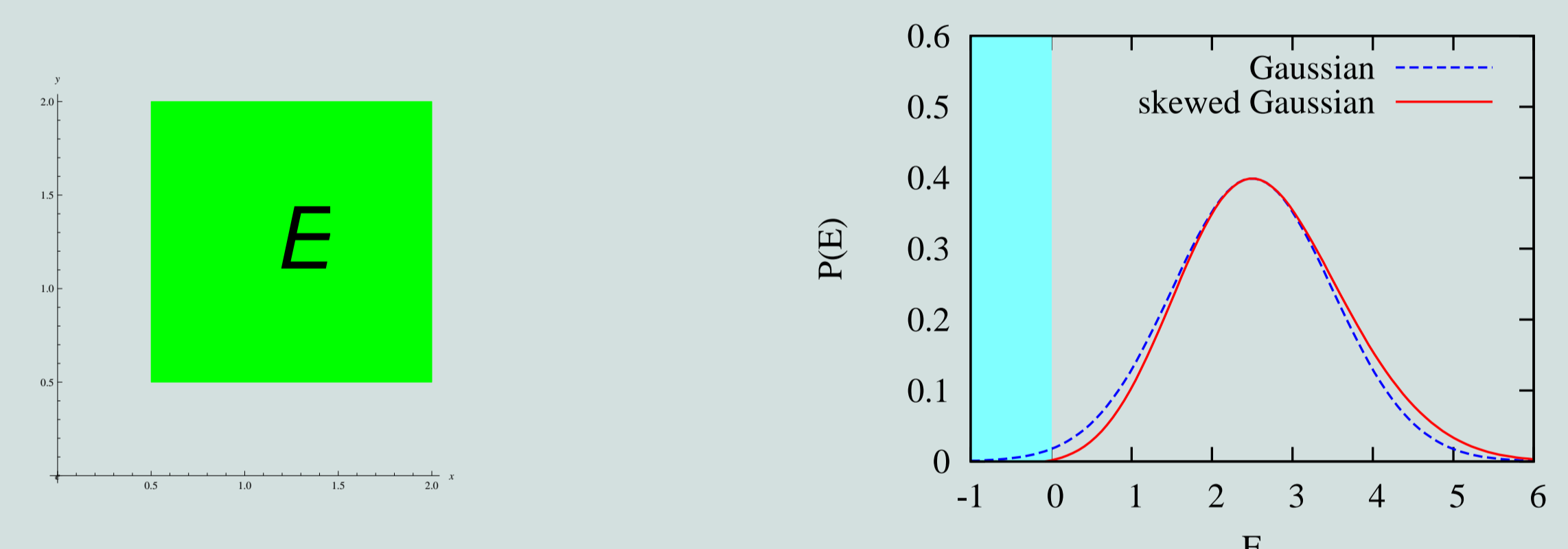


- Fluctuations in the energy of each source increase both anisotropies ($\varepsilon_2\{2\}$) and non-Gaussianities (smaller splitting between $\varepsilon_2\{2\}$ and $\varepsilon_2\{4\}$). These trends are also present in asymptotic results (generalization of Ref. [6]), yet convergence to the asymptotic regime is very slow. The universal statistics from [7] works significantly better than asymptotic results.

Analytical results

Results from the independent-source model generalized to a continuous fluctuating energy density profile [9].

The energy E contained in a given area of the transverse plane has non-Gaussian event-to-event fluctuations because of the constraint $E \geq 0$.



The distribution is typically right-skewed $\langle (E - \langle E \rangle)^3 \rangle > 0$. We treat non-gaussianities as small perturbations and obtain the general expression of $\varepsilon_n\{4\}$ to leading order (not shown). For identical pointlike sources we recover the result of [6]; each term has a different physical interpretation.

$$\varepsilon_n\{4\}^4 = \frac{1}{N^3} \left(\underbrace{\frac{\langle r^{2n} \rangle^3}{\langle r^n \rangle^6}}_{\text{Gaussian}} + \underbrace{\frac{8 \langle r^{3n} \rangle \langle r^{2n} \rangle}{\langle r^n \rangle^5}}_{\text{Skewness}} - \underbrace{\frac{\langle r^{4n} \rangle}{\langle r^n \rangle^4}}_{\text{Kurtosis}} + \underbrace{\frac{2 \langle r^{2n} \rangle^2}{\langle r^n \rangle^4}}_{\text{E conserv.}} \right)$$

- Gaussian density fluctuations yield slight non-Gaussianities in ε_n fluctuations but they have the wrong sign [10].

- Non gaussian density fluctuations restore positivity.

Conclusion

Fluctuations in the energy of each source yield strong non-Gaussianities which are well described by the universal statistics of Ref. [7]. This may explain why CMS p+Pb results [8] on $v_2\{6\}$ and $v_2\{8\}$ are also well described by this universal statistics.

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