

Neutrino Mass Measurements Asia

Second International Meeting for
Large Neutrino Infrastructures

Apr. 21, 2015

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Neutrino Mass

Neutrino oscillation \longrightarrow finite small masses and mixing

$$|\nu_\alpha\rangle = \sum_{i=1}^3 U_{\alpha i} |\nu_i\rangle \quad (\alpha = e, \mu, \tau) \quad m_i > 0$$

Possible mass terms

neutrino has no electric charge

$$L_{mass} = \overline{\psi}_R m_D \psi_L + \frac{1}{2} \overline{(\psi_L)^c} m_L \psi_L + \frac{1}{2} \overline{\psi}_R m_R (\psi_R)^c$$

$$\begin{array}{c} \longrightarrow \longrightarrow \\ \nu_R \quad \nu_L \\ \Delta L = 0 \end{array}$$

$$\begin{array}{c} \longrightarrow \longleftarrow \\ \nu_L \quad \bar{\nu}_L \\ \Delta L = 2 \end{array}$$

$$\begin{array}{c} \longrightarrow \longleftarrow \\ \nu_R \quad \bar{\nu}_R \\ \Delta L = 2 \end{array}$$

lepton number violation

Majorana particle $\nu = \bar{\nu}$

- Heavy neutrino (just below GUT scale) naturally explains **“finite but light neutrino mass”** $m \sim \frac{m_D^2}{M_R}$ (Seesaw mechanism)
- CP violating decay of heavy neutrino explains **“matter dominance in the universe”** (Leptogenesis theory)

Neutrinoless Double-Beta Decay

Nucleus $\beta\beta$ -decay emitting two electrons

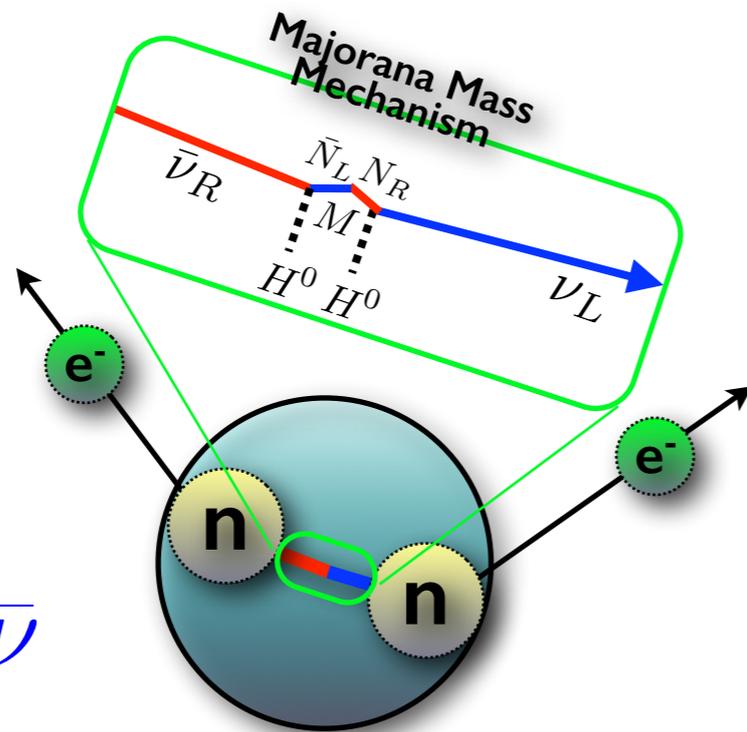
$$(Z, A) \rightarrow (Z + 2, A) + 2e^-$$

lepton number violation

$$\Delta L = 2$$

$0\nu\beta\beta$

neutrinoless double beta-decay



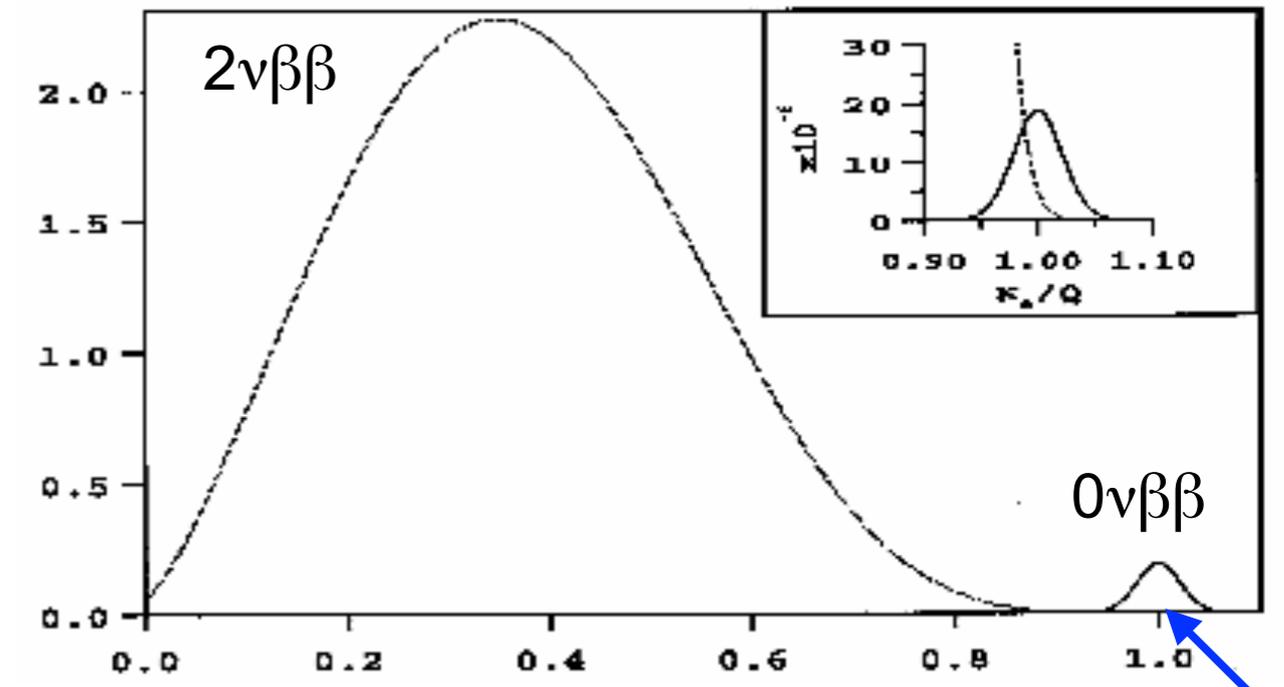
$$\nu = \bar{\nu}$$

Require the neutrino **Majorana nature**

amplitude $\sim \langle m_{\beta\beta} \rangle \equiv \left| \sum_{i=1}^3 U_{ei}^2 m_i \right|$

↑
effective neutrino mass

Possible contributions also from Majoron, SUSY, right-handed current, ...



sum energy for two electrons / Q peak at Q-value

decay rate

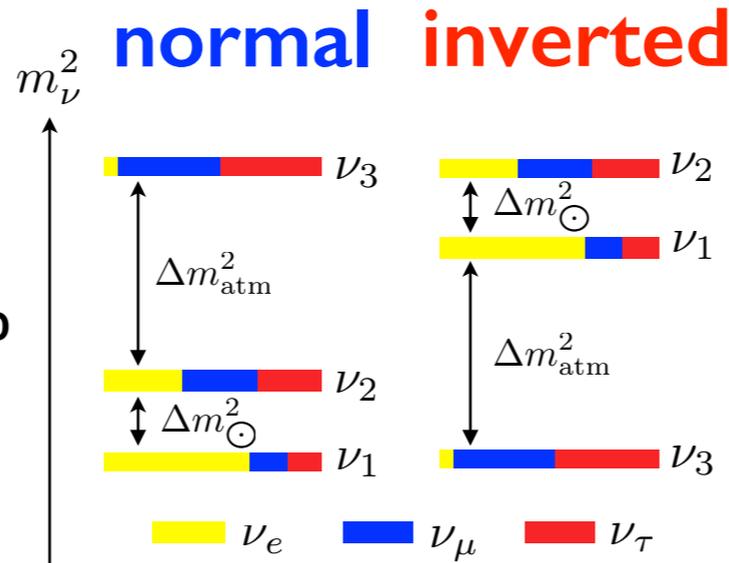
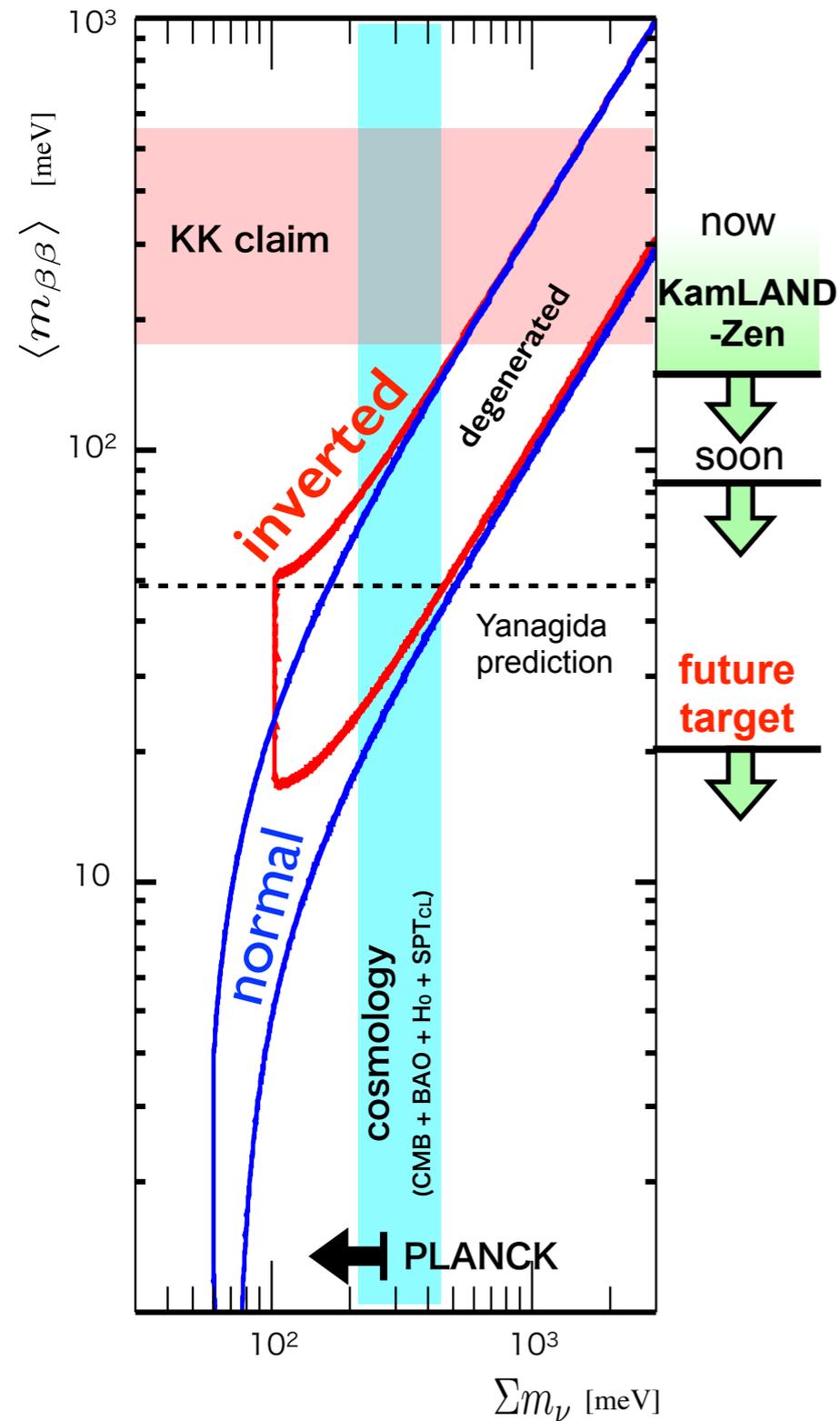
$$[T_{1/2}^{0\nu}]^{-1} = G^{0\nu}(Q_{\beta\beta}, Z) |M^{0\nu}|^2 \langle m_{\beta\beta} \rangle^2$$

nuclear matrix element
(nuclear physics)

phase space factor

effective neutrino mass
(particle physics)

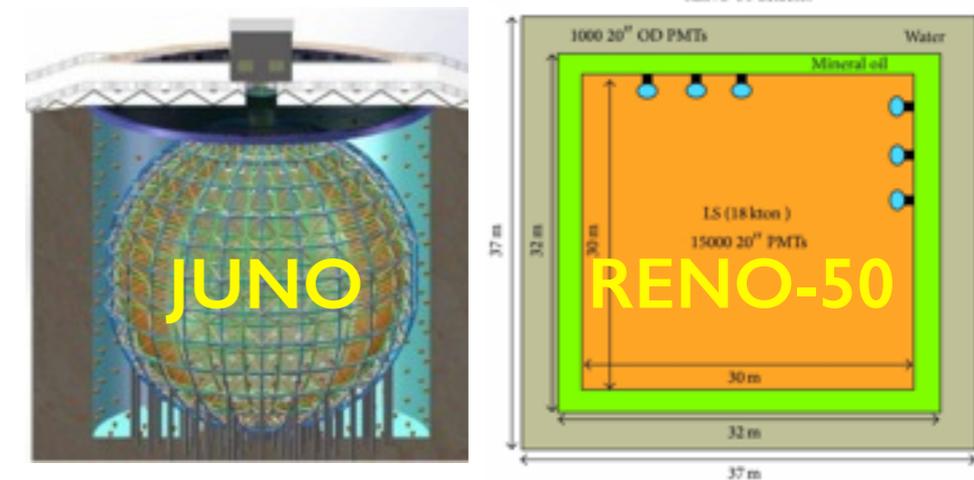
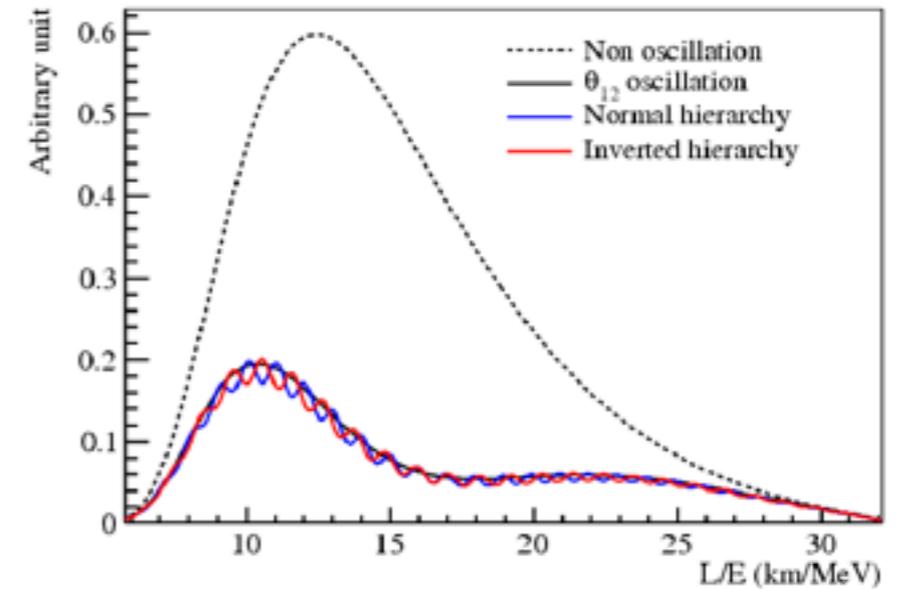
Neutrino Mass Hierarchy



Mass hierarchy (MH) determination

reactor neutrino experiments in China and Korea are competitive with other MH projects (atmospheric / accelerator neutrinos)

Reactor neutrino oscillation



(a) inverted hierarchy case

find $0\nu\beta\beta$ signal in the next generation experiments! 😊

(b) normal hierarchy case

need to set the next target down to 1-5 meV ... 😞

If we don't find the signal, "Majorana neutrino" will be rejected

Double Beta Decay Experiments

Asia

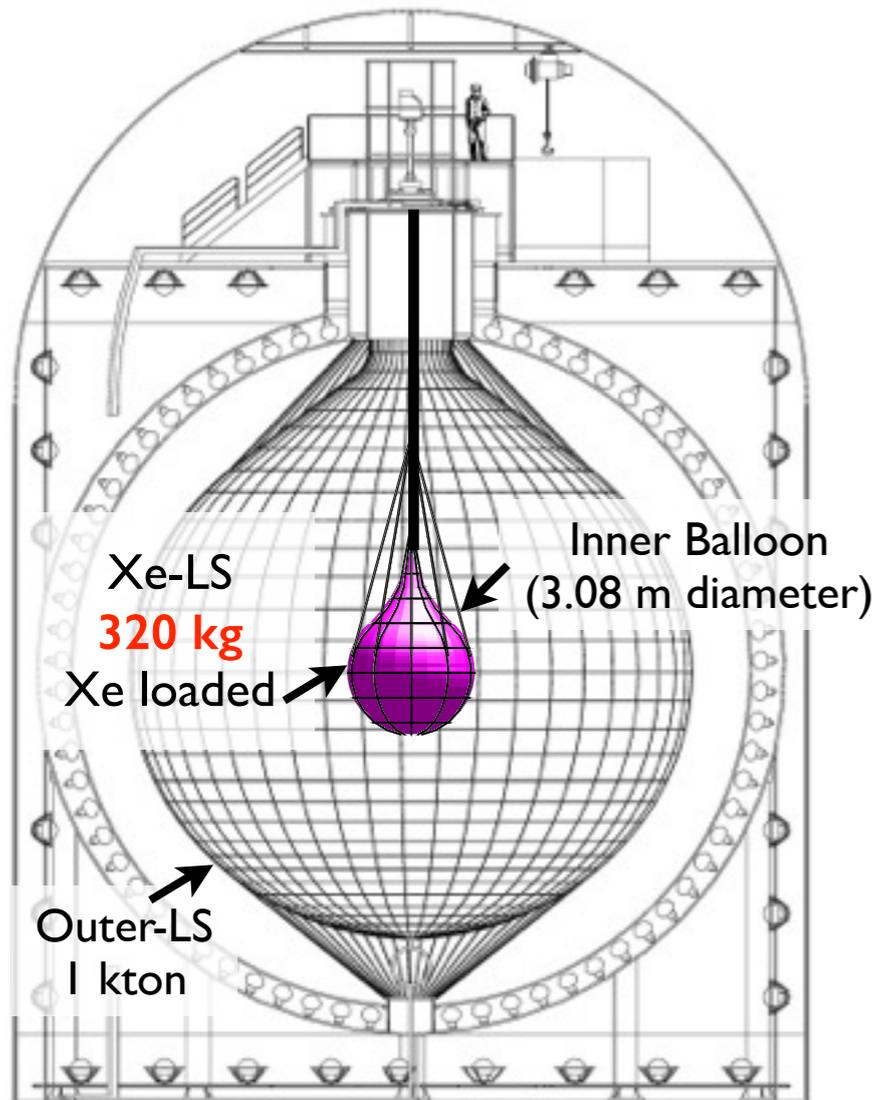
Experiment	$\beta\beta$ candidate	Q-value[keV]	Enrichm.	$N_{\beta\beta} \times 10^{26}$	Start [y]	$\langle m_{ee} \rangle$ [meV]@5y
GERDA	^{76}Ge	2039	yes	3.2	2013	73-203
Majorana	^{76}Ge	2039	yes	2.4	2014	106-295
MaGe	^{76}Ge	2039	yes	68	2020	43-120
CUORE	^{130}Te	2527.5	no	9.6	2014	40-94
Lucifer	^{82}Se	2995	yes	1.3	2014	35-94
AMore	^{100}Mo	3034	yes	3	?	27-63
SNO+	^{150}Nd	3370	no	1.8	2014	172-180
Kamland-Zen	^{136}Xe	2476	yes	4	2013-2015	25
Candles	^{48}Ca	4270	no	0.04	2011	500
Candles-enr	^{48}Ca	4270	yes	1	?	IH
Exo-200	^{136}Xe	2476	yes	2.3	2011	87-221 @2y
Exo-Full	^{136}Xe	2476	yes	20	?	16-40
Next-100	^{136}Xe	2476	yes	4	2015	90 @6y
Next-1t	^{136}Xe	2476	yes	30	?	38 @(3+3)y
COBRA	^{116}Cd	2809	yes	nd	?	50
SuperNemo	^{82}Se	2995	yes	7.3	2014	40-105
Moon	$^{82}\text{Se}/^{100}\text{Mo}$	2995/3134	yes	30	?	IH
DCBA	^{150}Nd	3370	yes	10	?	30

KamLAND-Zen

KamLAND-Zen

Kamioka Liquid Scintillator Anti-Neutrino Detector
Zero Neutrino Double Beta

KamLAND-Zen
Phase I



Xenon loaded LS (Xe-LS)

decane	82%
pseudo-cumene	18%
PPO	2.7 g/liter
xenon	2.44 wt%

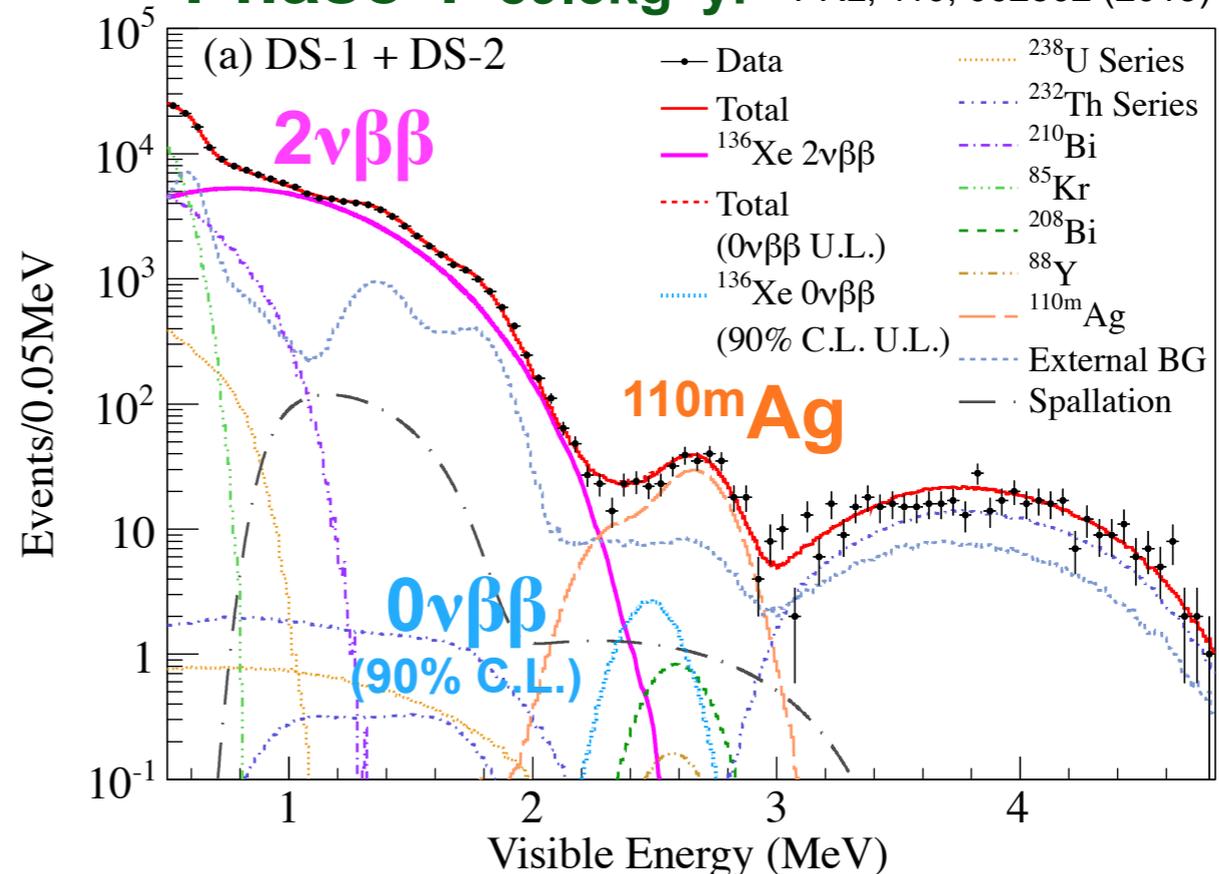
$$\sigma_E(2.5\text{MeV}) = 4\%$$

Advantage of KamLAND

- running detector : start quickly with relatively low cost
 - big and pure : no BG from external gamma-rays
 - purification of LS, replacement of mini-balloon are possible
- **high scalability** (a few ton of Xe)

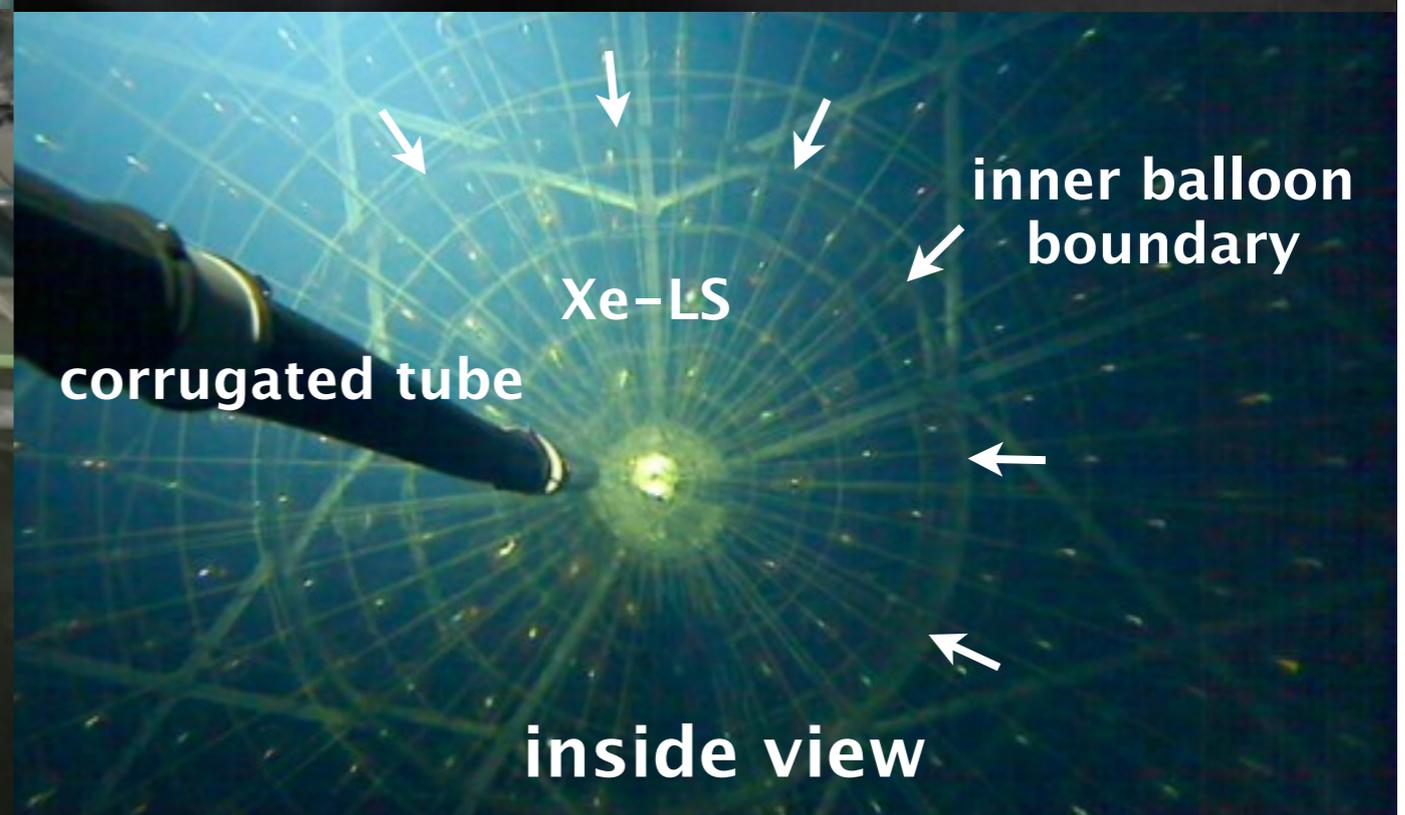
realize double beta-decay search with **low background**

Phase 1 89.5kg·yr PRL, 110, 062502 (2013)



$$T^{0\nu}_{1/2} > 1.9 \times 10^{25} \text{ yr (90\% C.L.)}$$

Construction in 2011



Improvement Efforts after Phase 1

1. Remove radioactive impurities by Xe-LS purification

candidates of ~2.6 MeV peak

→ only 4 nuclei ^{110m}Ag (250 d), ^{208}Bi (3.68×10^5 yr), ^{88}Y (107 d), ^{60}Co (5.27 yr)
lifetime longer than 30 days ↗ detected in Fukushima fallout

Two possible sources:

- (1) contamination by Fukushima-I reactor fallout
- (2) cosmogenic Xe spallation while above ground

**“primary” background source (^{110m}Ag)
can be removed by Xe-LS purification**

2. Increase amount of Xenon

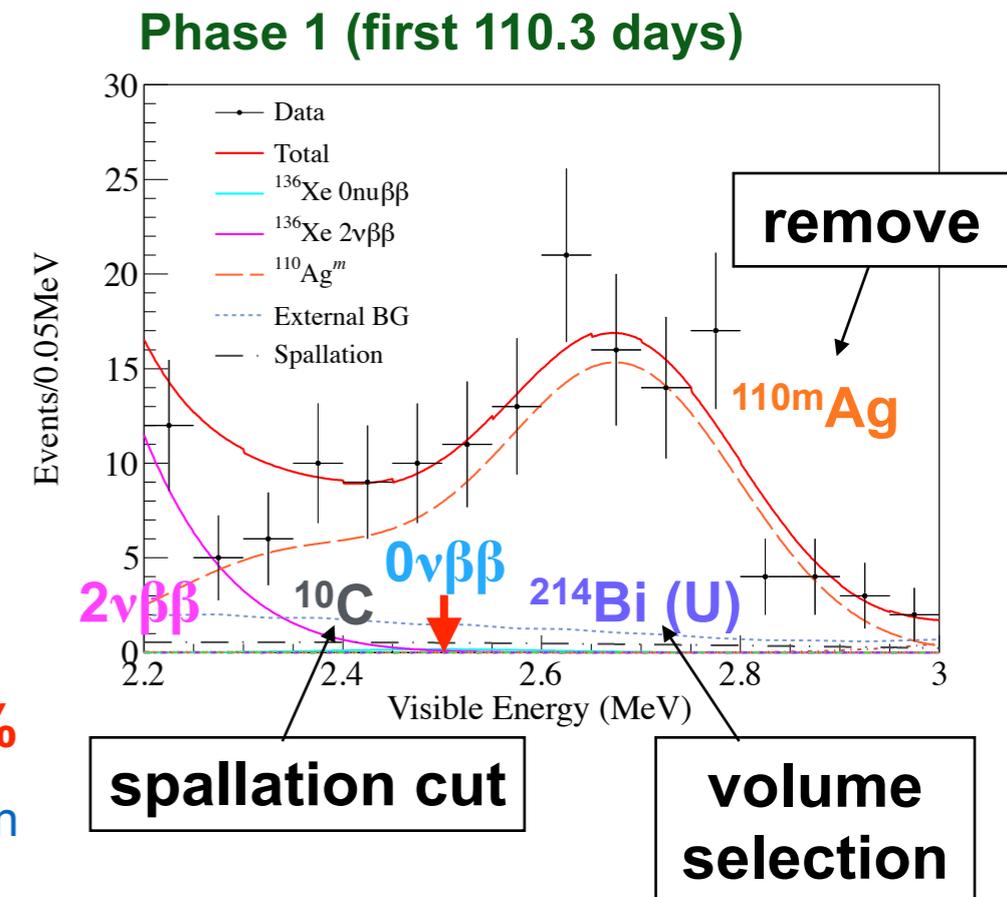
Xe concentration (2.44 ± 0.01) wt% \rightarrow (2.96 ± 0.01) wt%
increase of S/N ~ 1.2 Xe-pressurized phase is a future option

3. Spallation cut after muon

muon-neutron- ^{10}C ($\tau = 27.8$ s) triple coincidence \rightarrow ^{10}C background rejection

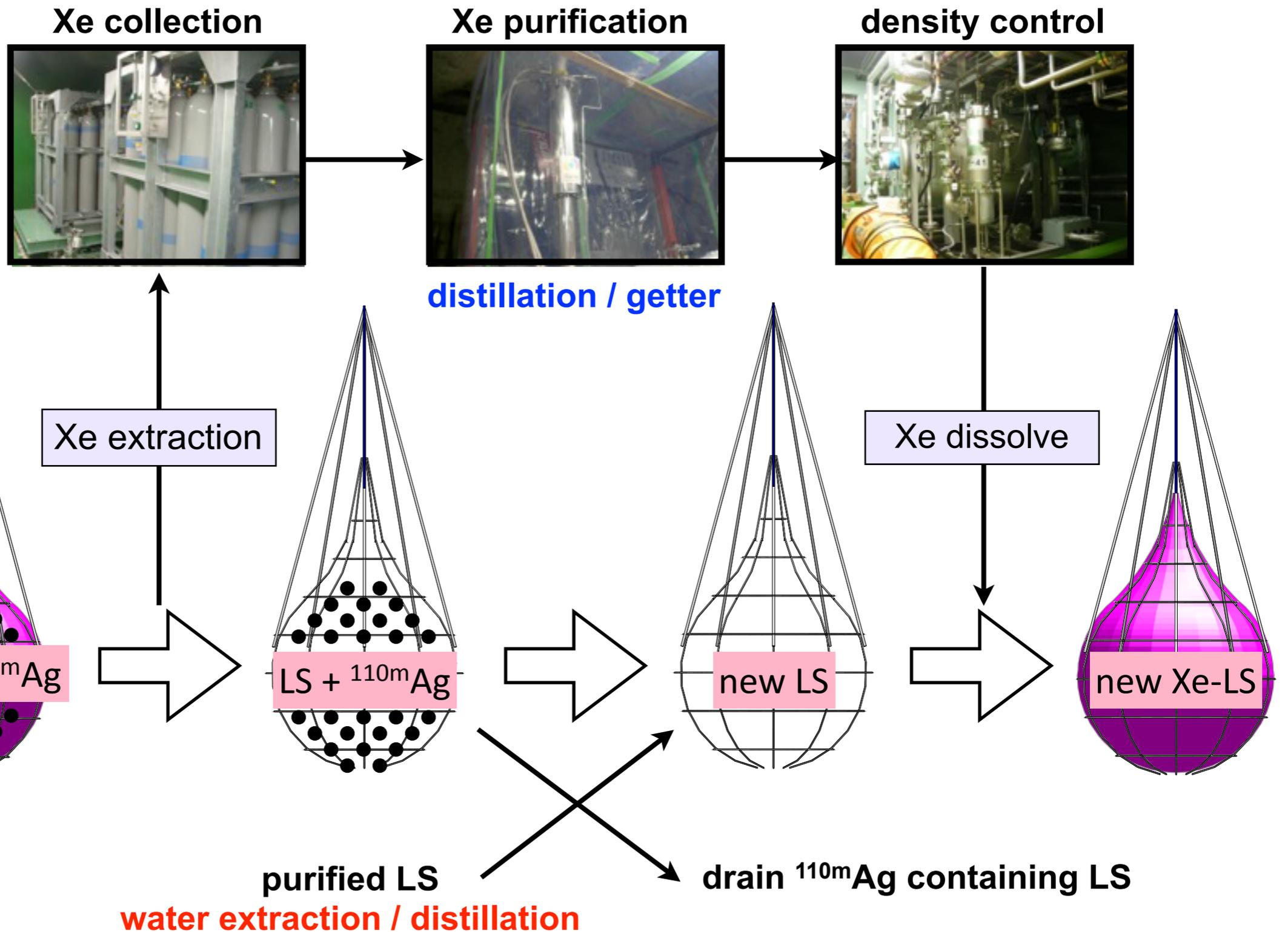
4. Optimization of volume selection

fiducial volume limitation by ^{214}Bi (U) on the balloon film \rightarrow multi-volume selection



Purification Strategy

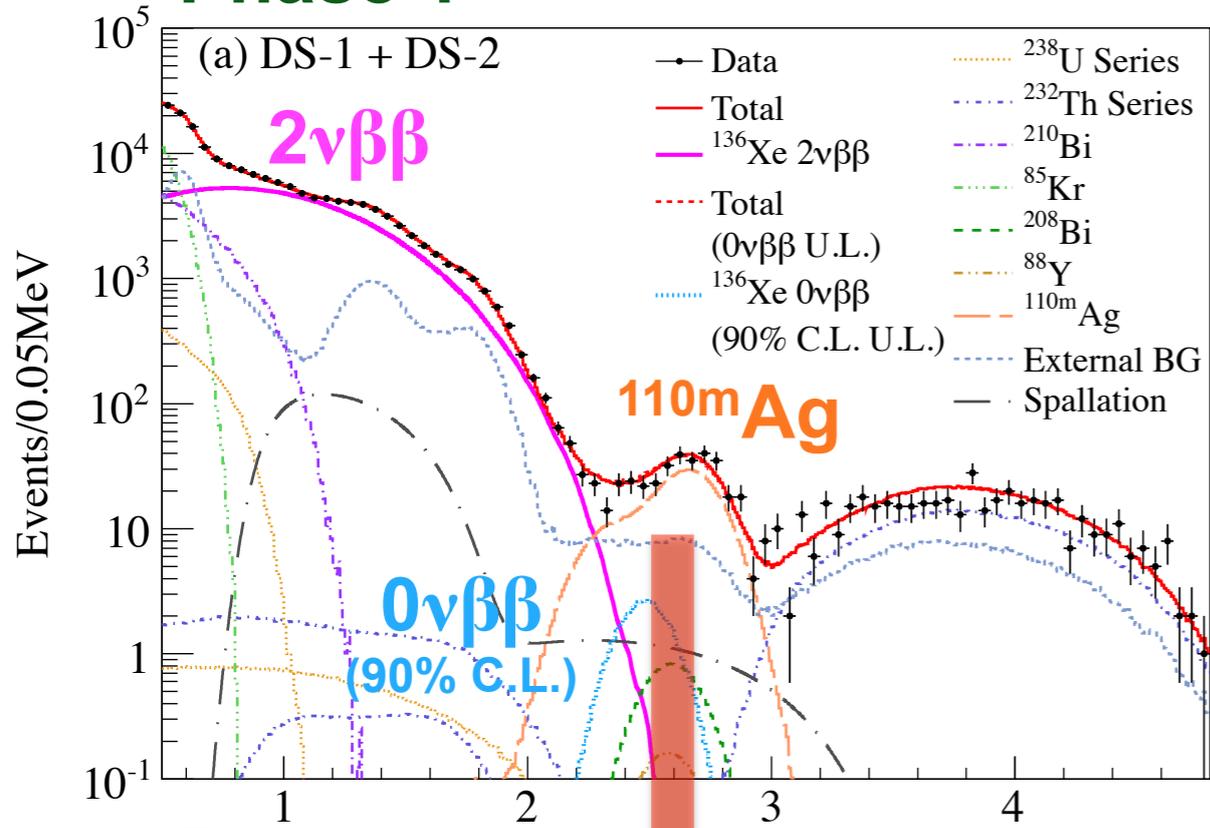
**Xenon
purification**



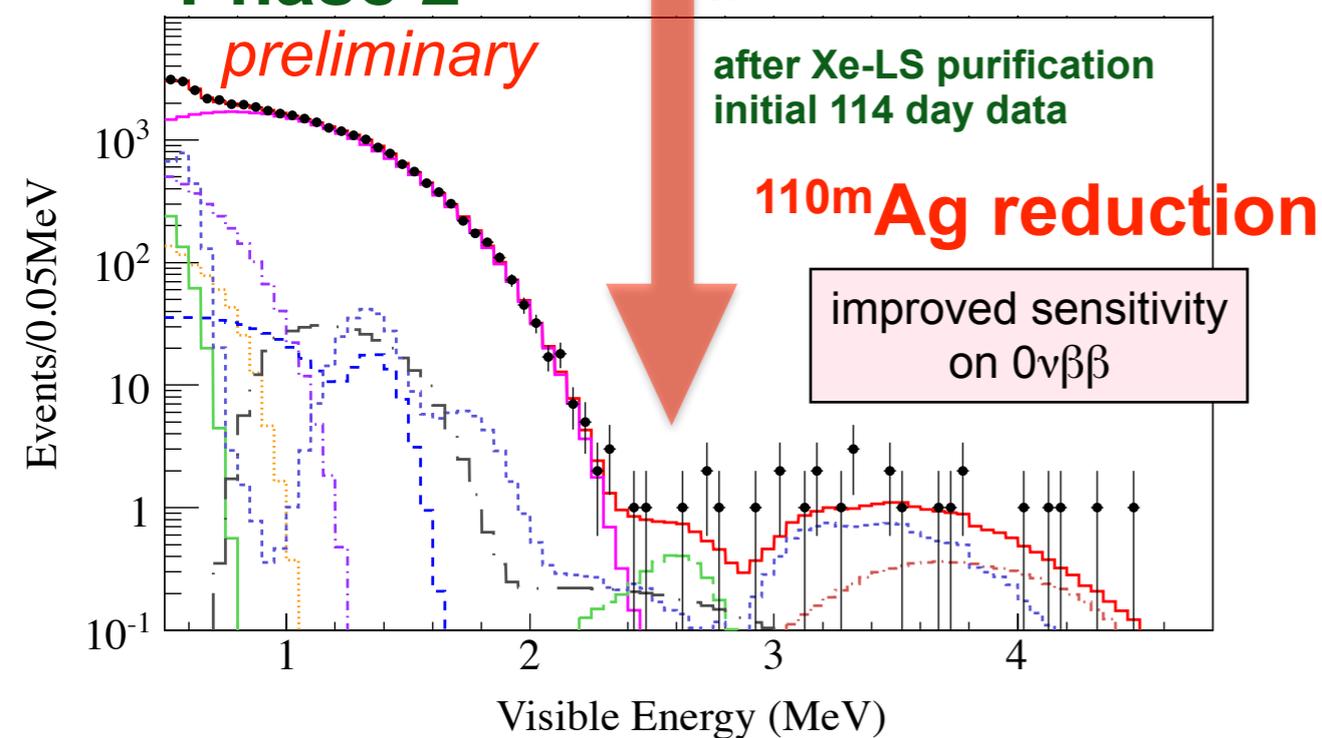
Xenon and LS purification to reduce radioactive impurities

KamLAND-Zen Phase 2 start

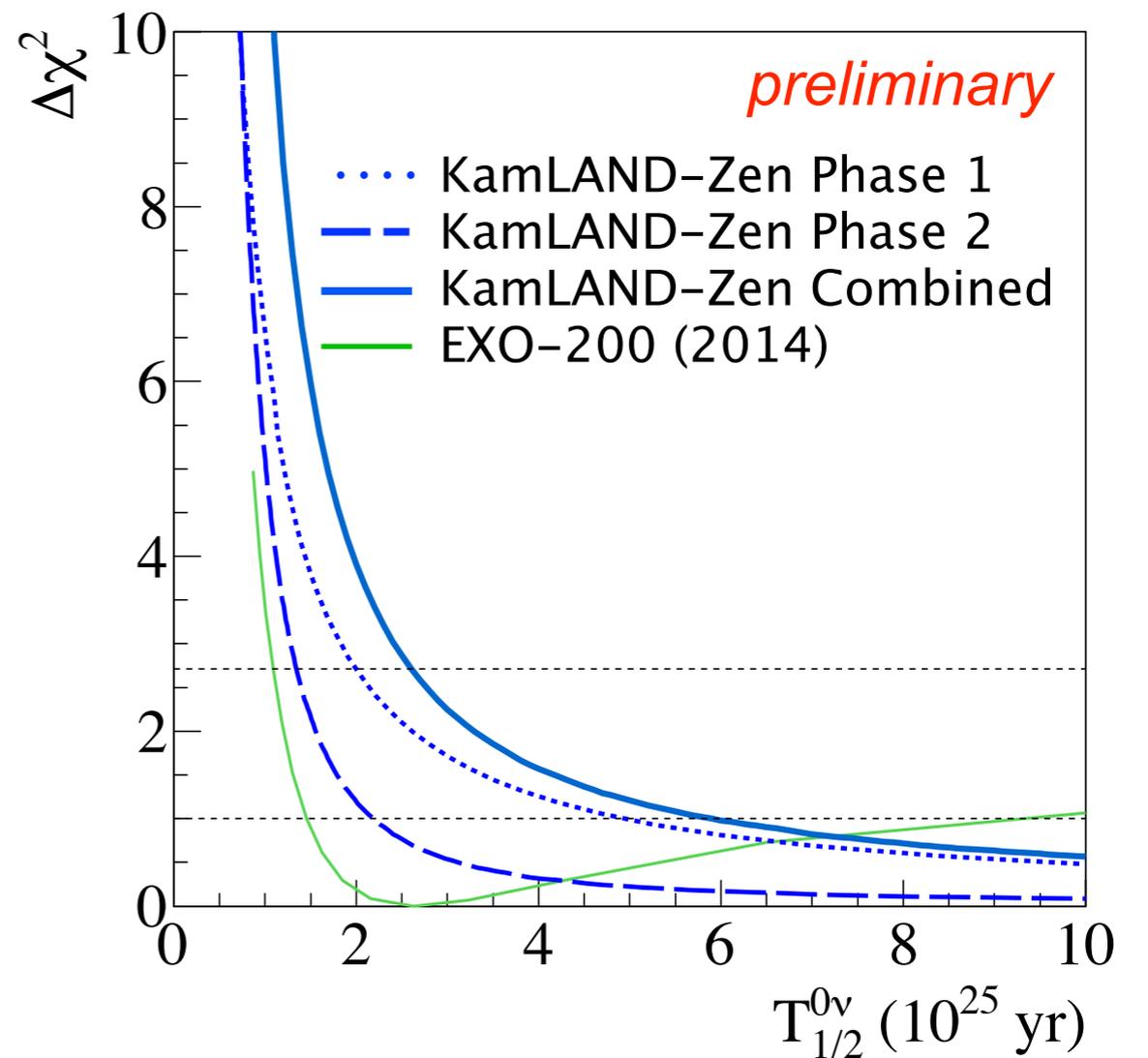
Phase 1



Phase 2



combined result (Phase 1 + 2)



KamLAND-Zen limit at 90% C.L.

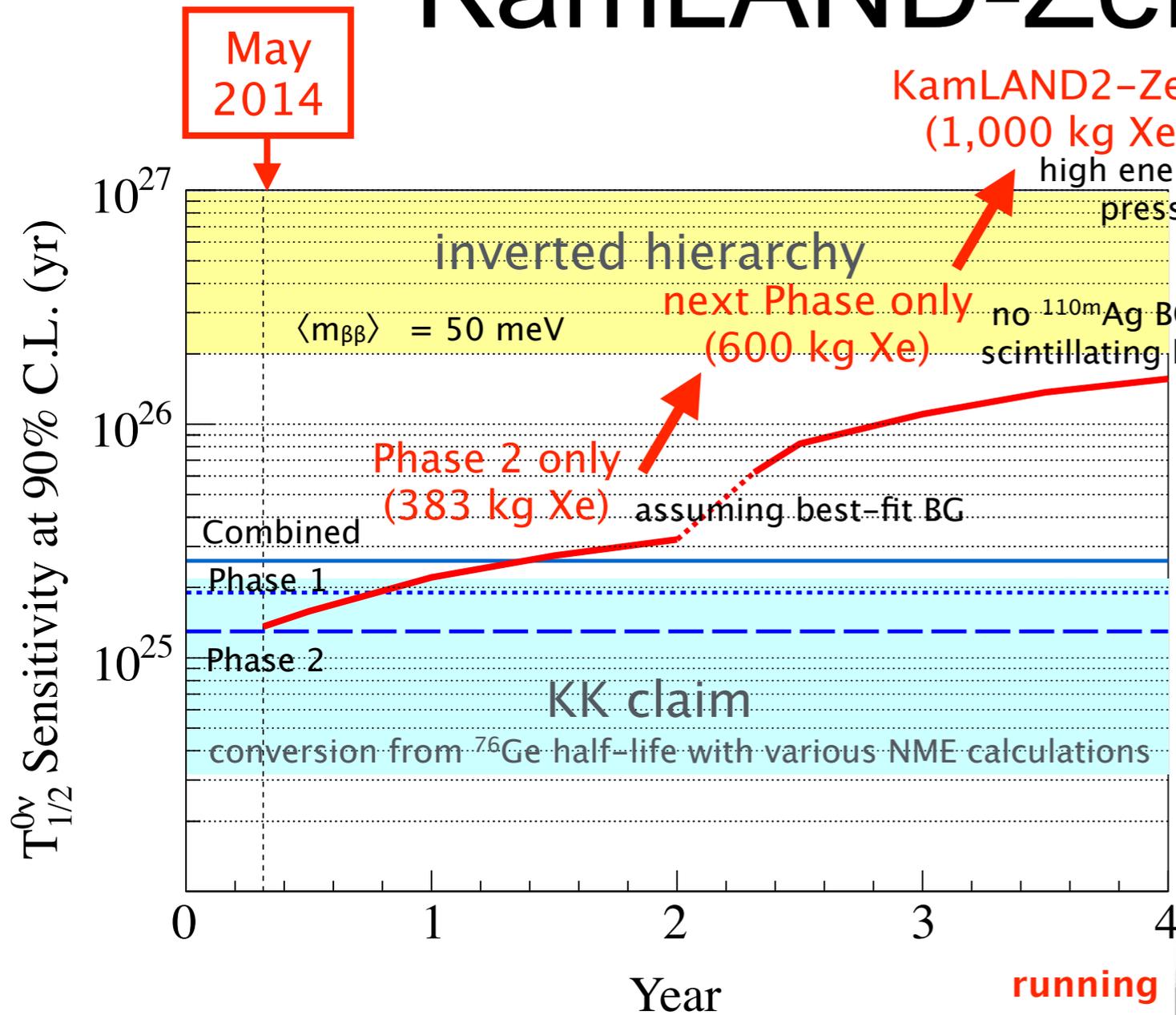
Phase 1 (213 days) + Phase 2 (115 days)

$$T_{1/2}^{0\nu} > 2.6 \times 10^{25} \text{ yr}$$

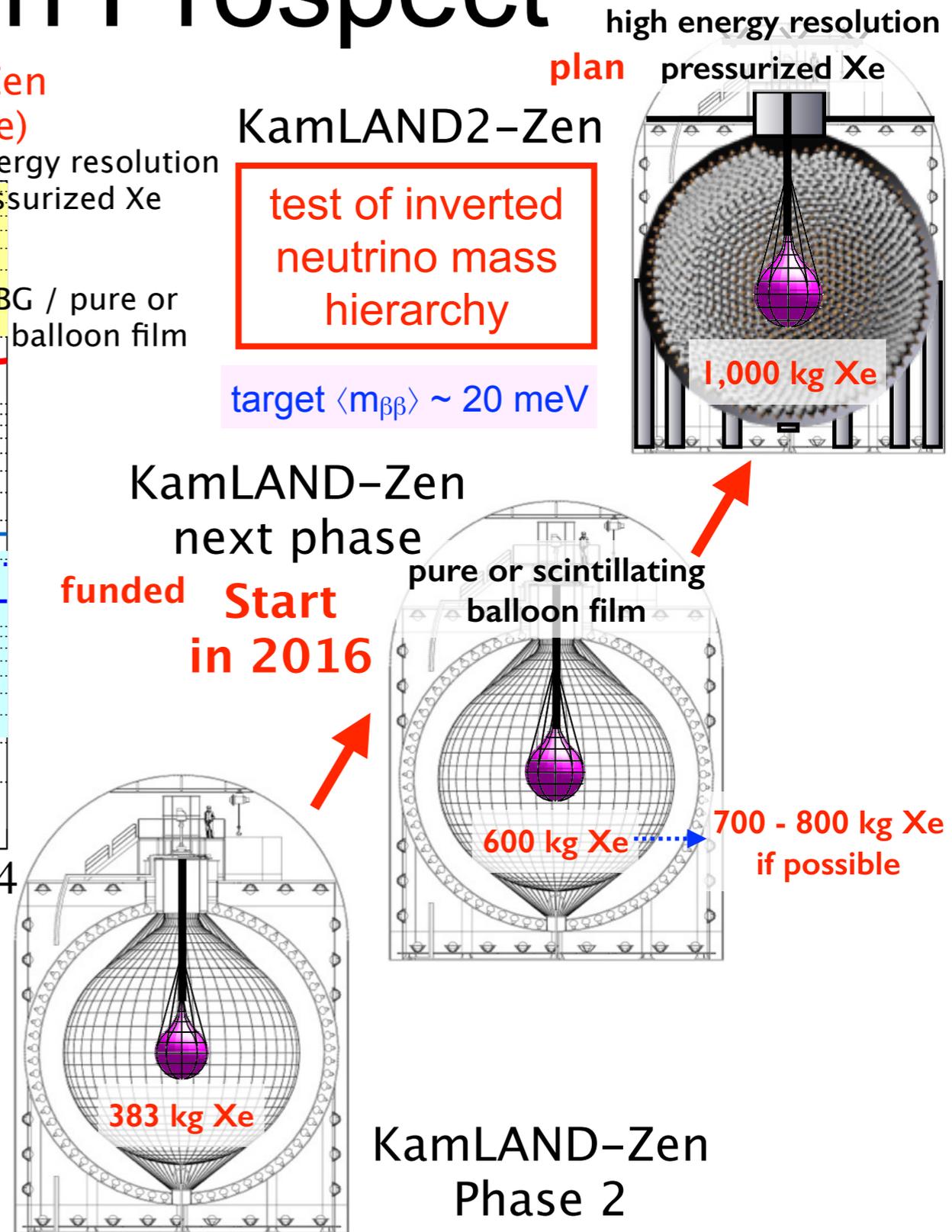
QRPA NME model
J. Phys. G 39 124006 (2012)

$$\langle m_{\beta\beta} \rangle < 140 - 280 \text{ meV}$$

KamLAND-Zen Prospect



- Future sensitivity
- ⋯ KamLAND-Zen Phase 1
- - - KamLAND-Zen Phase 2
- KamLAND-Zen Combined



Detector improvements are planned in the near future

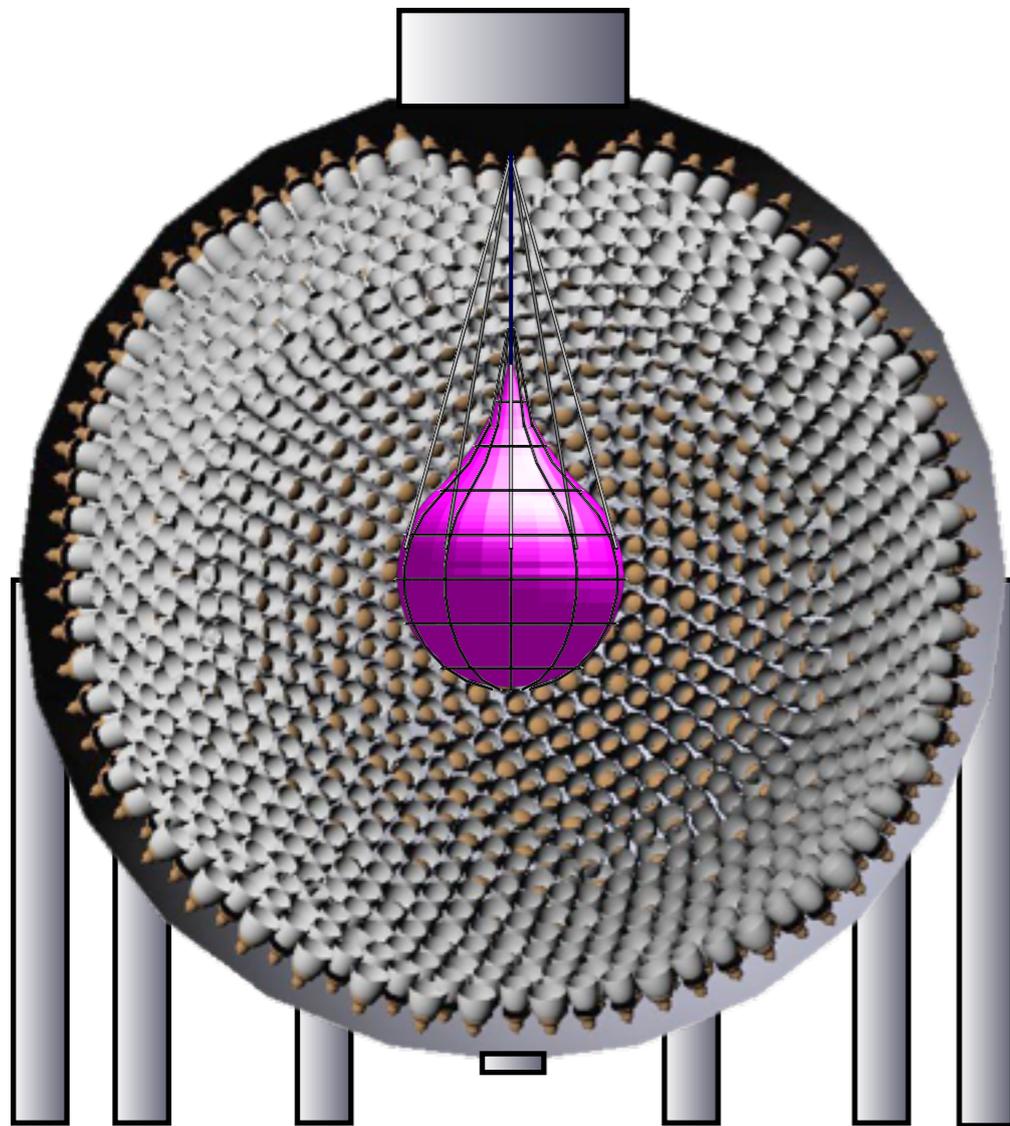
KamLAND2-Zen

General-purpose

larger crane
strengthen floor
enlarge opening

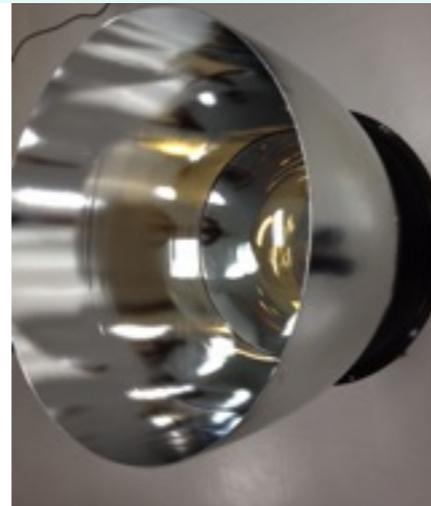
accommodate various devices
CaF₂, CdWO₄, NaI, ...

1000 kg enriched Xe



High performance

Winstone Cone



High Q.E. PMT



17"Φ → 20"Φ, ε=22% → 30%

Photo-coverage > **x2**

Light Collection Eff. > **x1.8**

x1.9

New Liquid Scintillator

x1.4

KamLAND liquid scintillator	8,000 photon/MeV
typical liquid scintillator	12,000 photon/MeV

$\sigma(2.6\text{MeV}) = 4\% \rightarrow < 2.5\%$

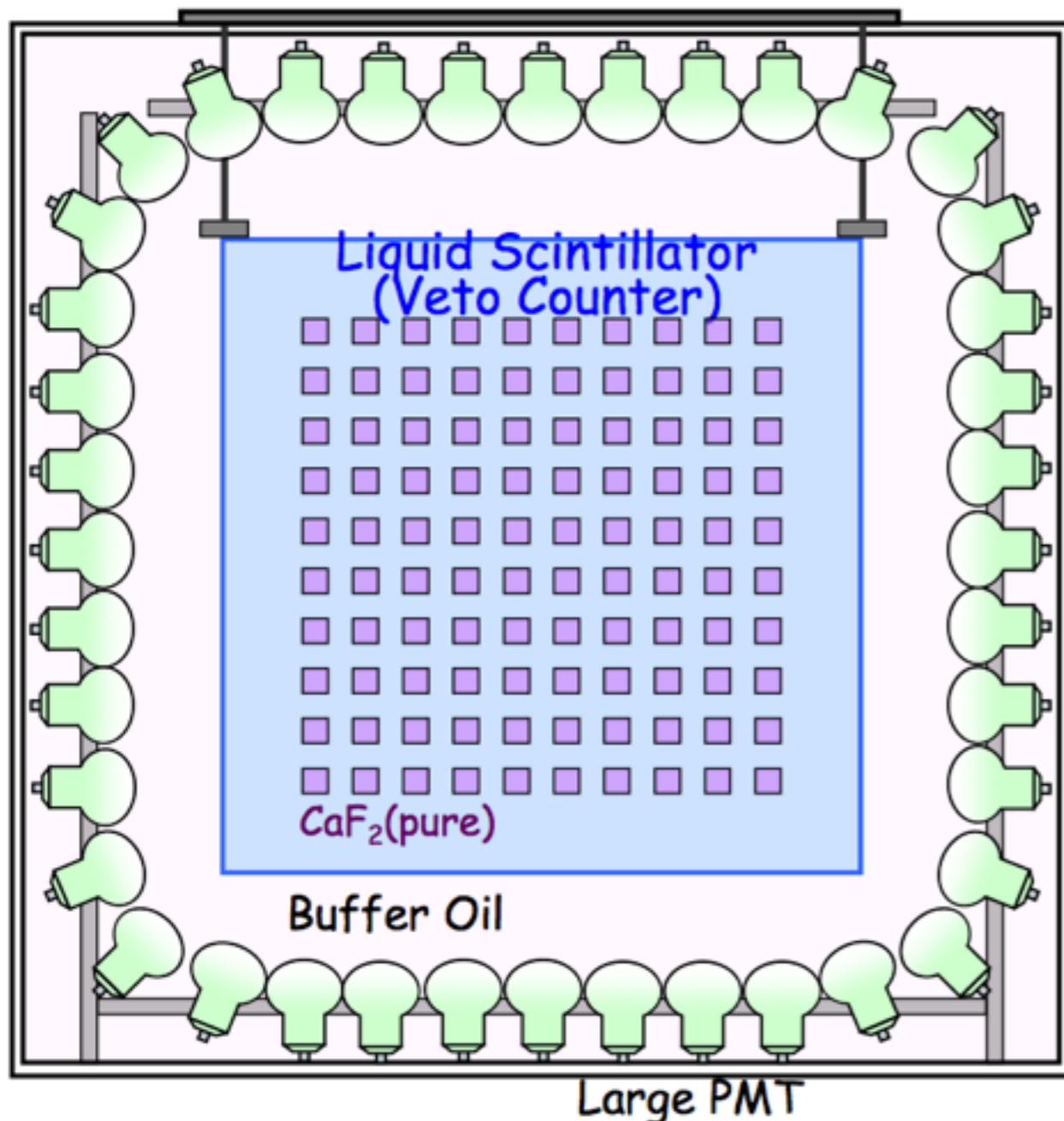
naive calc. < 2%

target $\langle m_{\beta\beta} \rangle \sim$ **20 meV / 5 year**

CANDLES

CANDLES

CAlcium fluoride for studies of Neutrino and DArk matters
by L ow E nergy S pectrometer



Why ^{48}Ca ?

- highest Q-value (^{48}Ca $\beta\beta$ -decay : 4.27 MeV)
- little background from natural radioactivities
- natural abundance of ^{48}Ca : 0.187% (small)
→ need isotope enrichment

- **CaF₂ (pure) scintillator**

transparent, ultra-pure crystal

200 kg, 300 kg, 2 ton, ... enrichment of ^{48}Ca

- **Liquid scintillator**

wavelength shifter, CaF₂ emission ~ 280 nm

4 π active shield (surrounded by passive shield)

- **Large PMT**

high energy resolution

pulse shape of scintillation light from CaF₂ & LS

CANDLES Prospect

- ELEGANT VI
- CANDLES I, II
- CANDLES III
- Next CANDLES

now

future option

background free measurement with 4π active shield

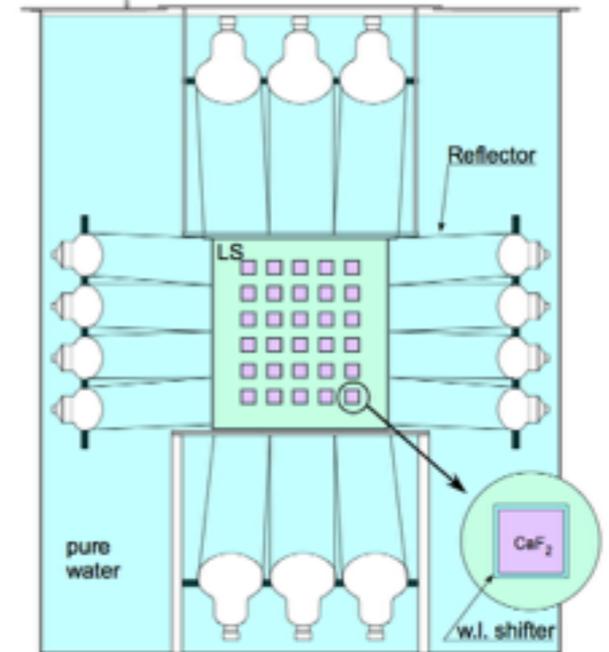


larger amount of ^{48}Ca in CANDLES

performance of 96 CaF_2 in LS active shield will be tested at underground enrichment of ^{48}Ca with Crown Ether chromatography is in R&D stage

increase CaF_2 in large LS detector

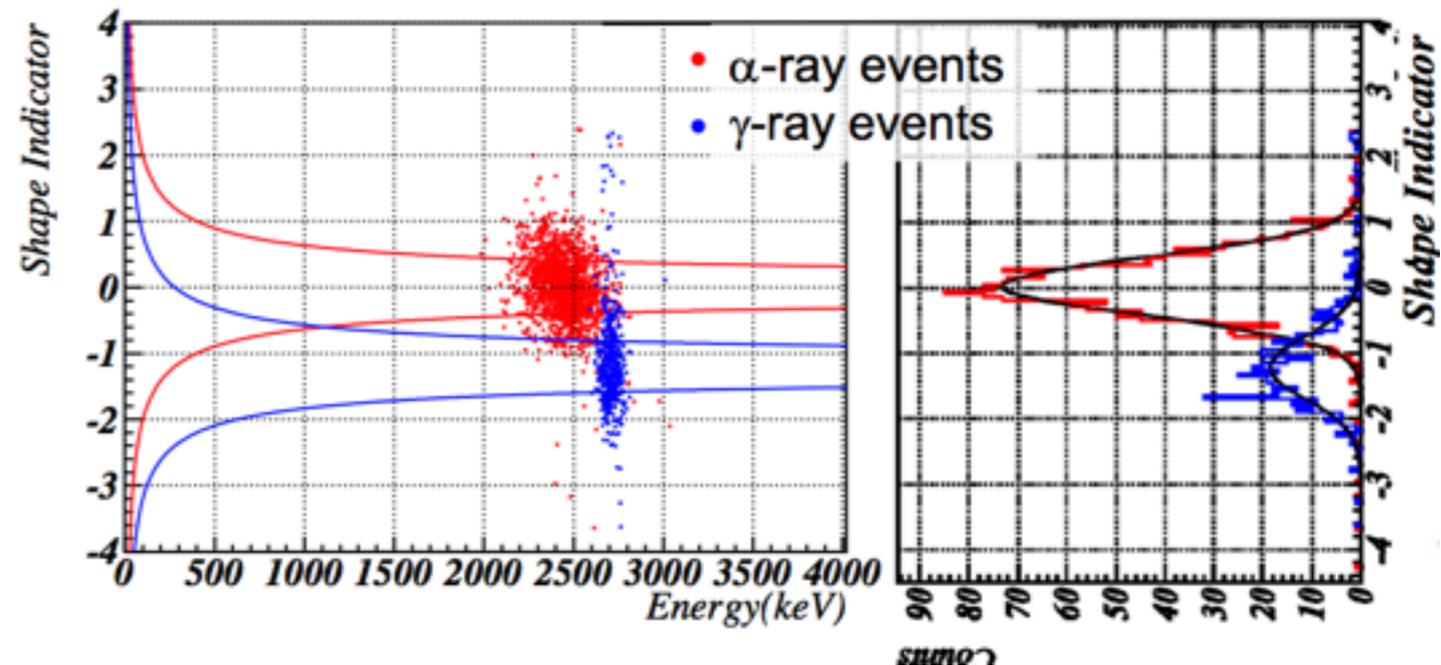
CANDLES III
96 CaF_2 scintillators



CANDLES III @ Kamioka



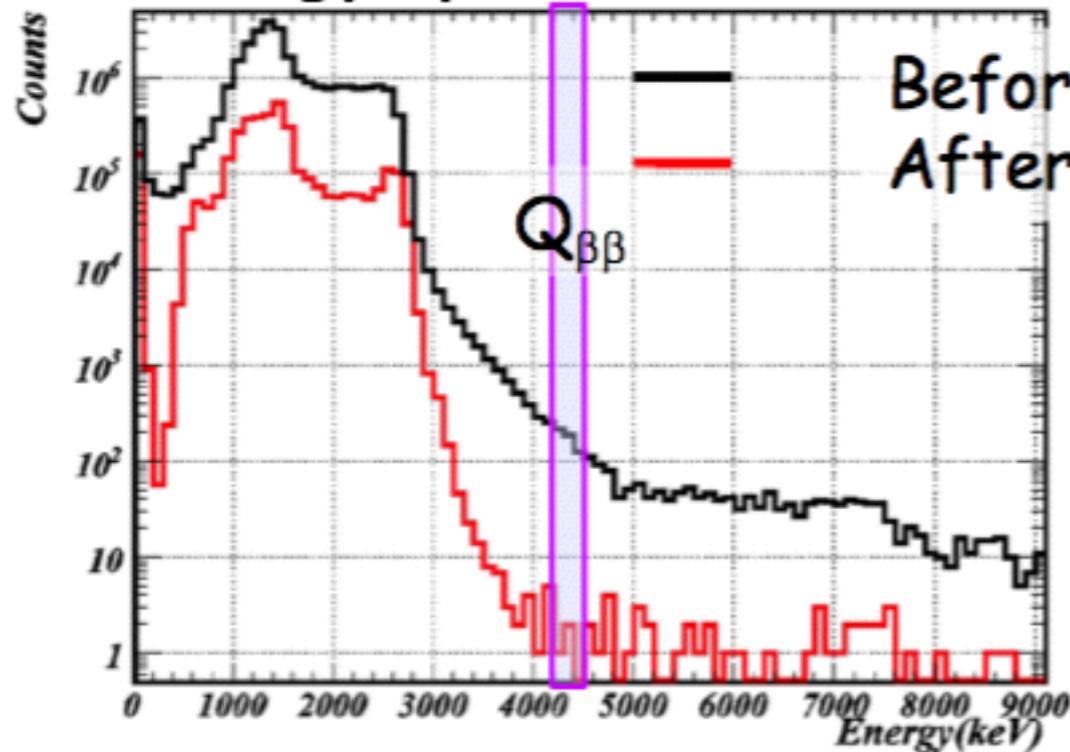
Pulse shape discrimination between α and γ



Energy Spectrum in CANDLES III

~8 week measurement with high purity $^{26}\text{CaF}_2$

Energy spectrum (~8 weeks)



S. Yoshida - DBD14

Before event selection
After event selection (BG rejection)

	Pilot run data
Measurement time	4987 kg · days
Number of events	6
Expected BG	~1 (CaF ₂ crystal) 3.4(γ-rays) (n, γ) reaction at surrounding materials
Sensitivity	0.8 × 10 ²² year materials

Internal BG : almost OK

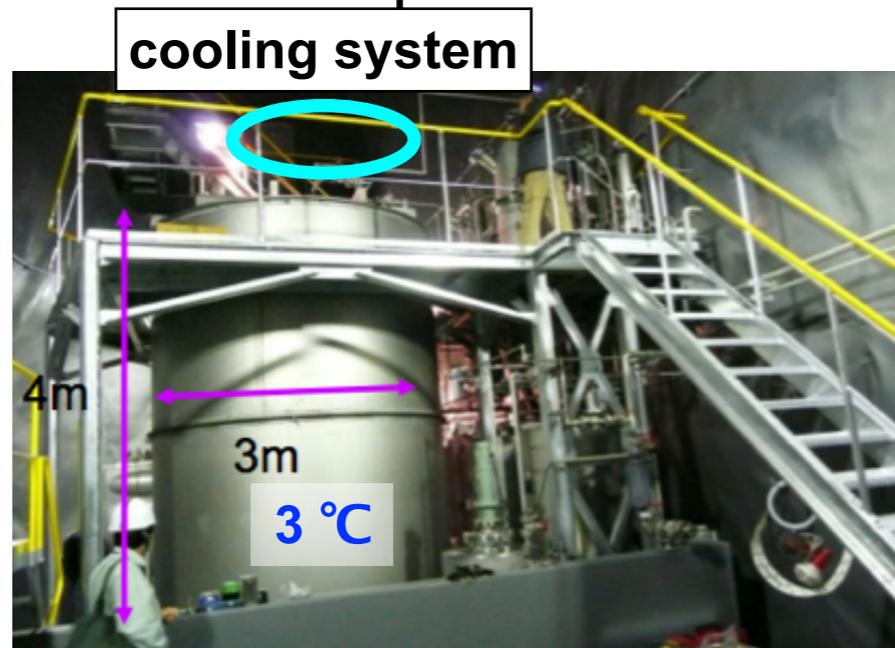
External BG : need to be reduced

newly install shield system for external γ-rays (neutron origin)

→ expected $\langle m_{\beta\beta} \rangle$ sensitivity ~ 0.5 eV

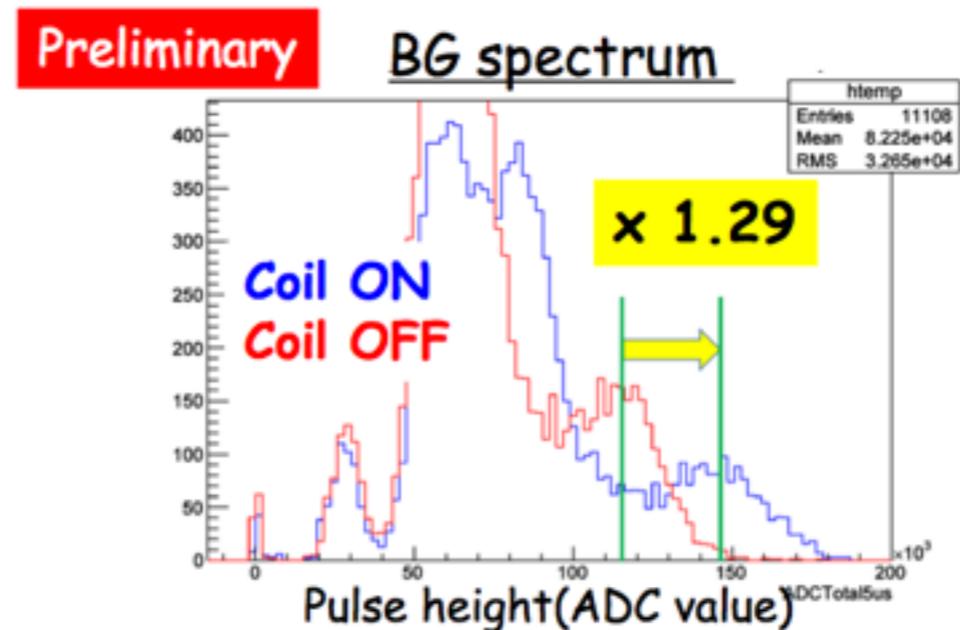
Further Development

1. Installation of the cooling system at the top of the detector



light output increase at low temperature

2. Installation of the “Geomagnetic cancellation coil” for PMTs

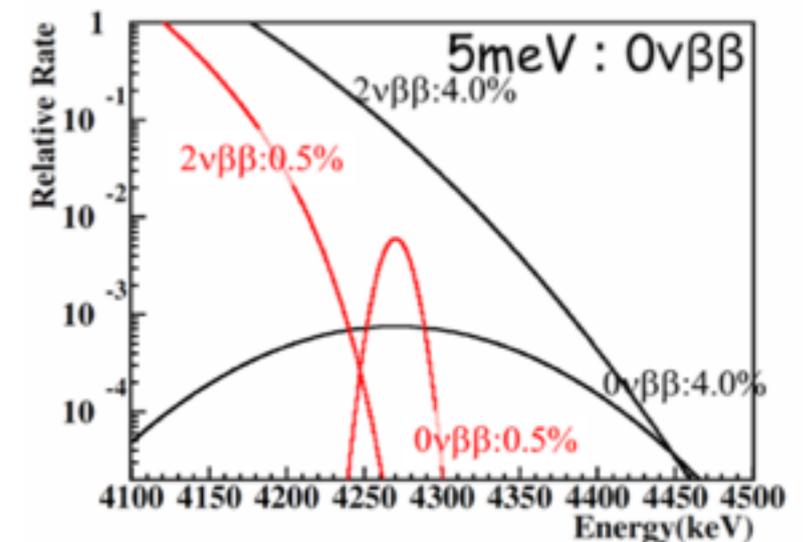


improved p.e. collection efficiency

expected energy resolution ~ 4% (FWHM) @ Q-value

Sensitivity of CANDLES

	CANDLES III	Next CANDLES	50% ⁴⁸ Ca 2 ton(610kg)
Crystal	3.2kg × 96 crystals	2% ⁴⁸ Ca	50% ⁴⁸ Ca 2 ton(610kg)
Total Mass	305kg (350g)	2 ton (25kg)	2 ton(610kg)
Energy Resolution	(4.0%)	2.8%(Req.)	0.5%(Req.)
2νββ	0.01	0.1	0.01
²¹² Bi, ²⁰⁸ Tl	0.26	~0.1	~0.01
Expected BG	0.27/year	< 0.7/3year	< 0.2/9year
<m _ν >	0.5 eV	0.08	0.009
	Current system	~2% enriched ⁴⁸ Ca and cooling system	scintillating bolometer



S. Yoshida - DBD14

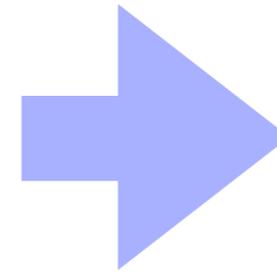
DCBA

DCBA

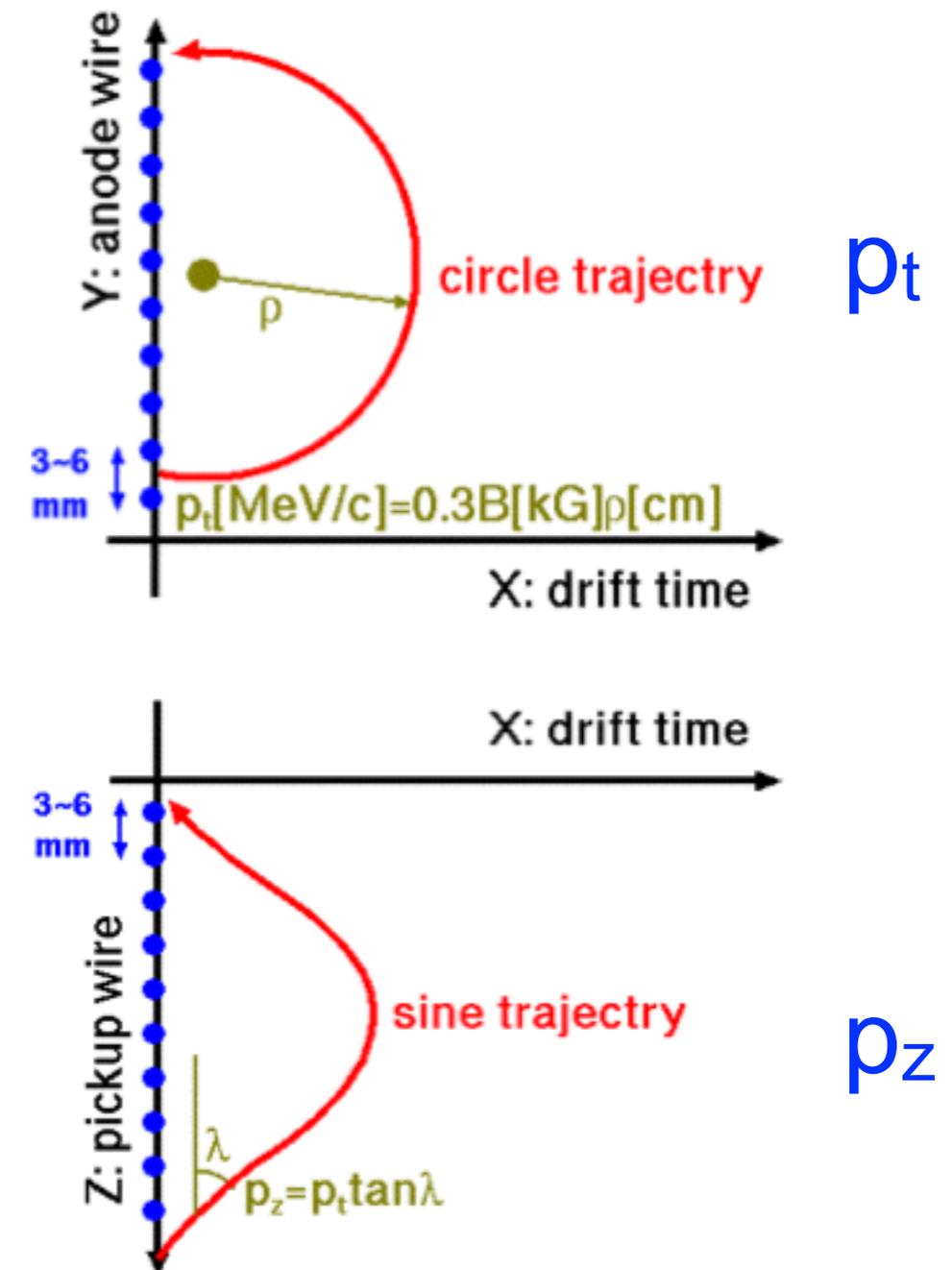
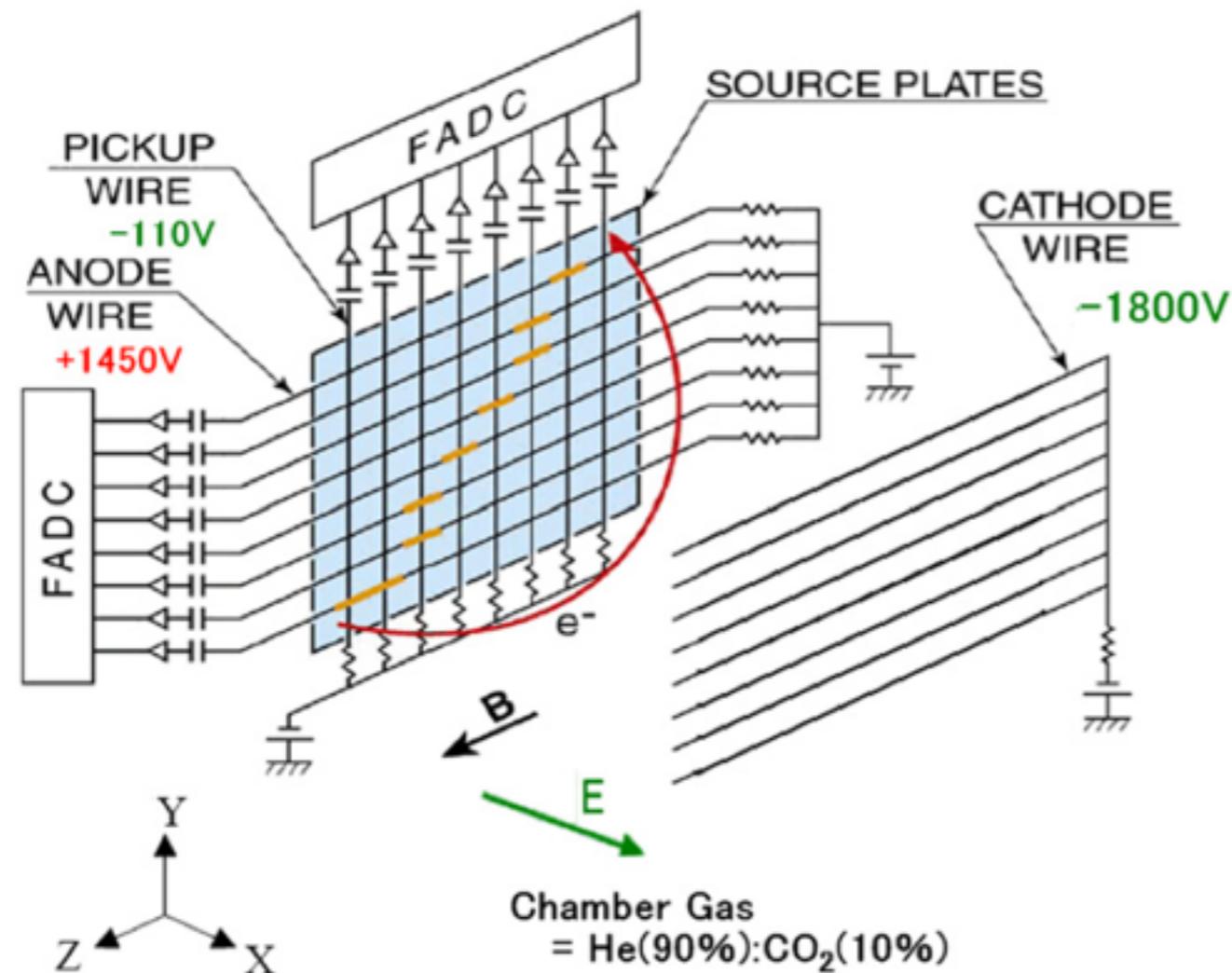
Drift Chamber Beta-ray Analyzer

Track / Position reconstruction

X : drift time
 Y : hit position of anode wire
 Z : hit position of pickup wire



Momentum reconstruction



DCBA Prospect

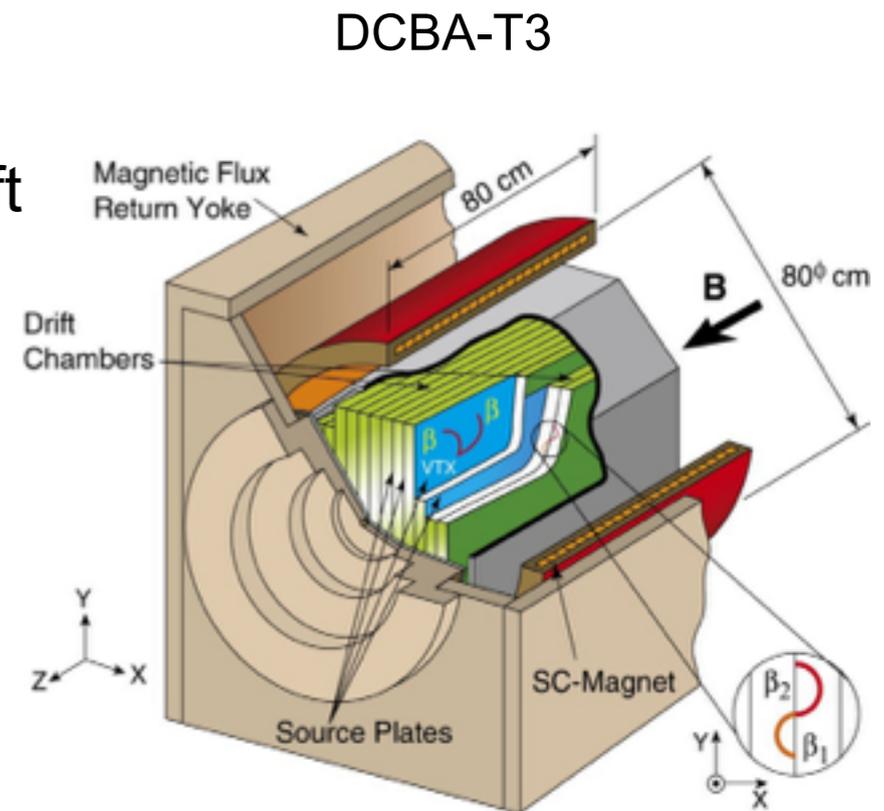
- 2005 DCBA
- 2007 DCBA-T2 demonstrate the principal of drift chamber tracking technique
- 2011 DCBA-T2.5
- 2014 DCBA-T3 prototype for MTD
- now
- 2017 **MTD** search for $0\nu\beta\beta$

$\beta\beta$ isotope ^{150}Nd (natural, enriched)
 DCBA-T3: 0.18 mol ^{150}Nd $Q_{\beta\beta} = 3370$ keV
 MTD-full: **several 10 kg ^{150}Nd (60% enriched)**

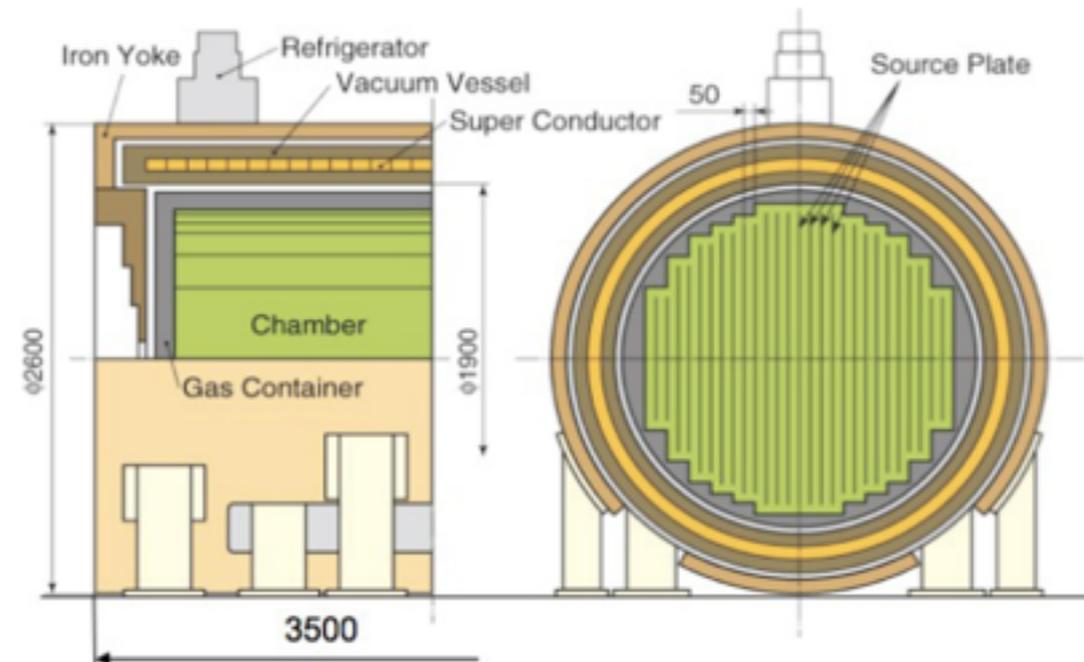
Resolution FWHM
 DCBA-T3: ~3.4% @ ROI DCBA-T2: ~6.2%
 (foreseen: 2.4 kG magnetic field)

Background
 $2\nu\beta\beta$: reduction with better energy resolution
 ^{214}Bi : tagged by ^{214}Po alpha

Projected sensitivity @ 90% C.L.
 DCBA-T3: $\langle m_{\beta\beta} \rangle < \sim 4$ eV
 MTD-full: **$\langle m_{\beta\beta} \rangle < \sim 30$ meV**



MTD (Magnetic Tracking Detector)

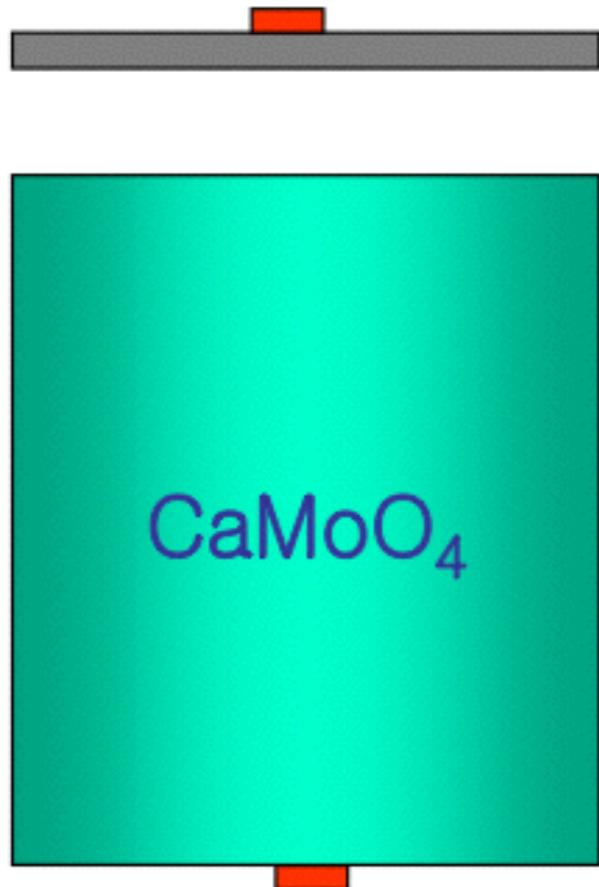


AMoRE

AMoRE

Advanced **Mo**-based **R**are process **E**xperiment

Si or Ge TES/MMC/NTD



10 - 50 mK

Phonon sensor (MMC)

V. N. Kornoukhov - TAUP 2011

Why CaMoO₄?

- highest light yield among Mo-crystals
- high Q-value (¹⁰⁰Mo ββ-decay : 3.03 MeV)
- less background from natural radioactivities
- enrichment of ¹⁰⁰Mo > 90% not so difficult
- ⁴⁸Ca needs to be depleted

- CaMoO₄ scintillator

transparent, ultra-pure crystal

1 kg, 10 kg, 200 kg, ... ⁴⁰Ca¹⁰⁰MoO₄

- Phonon + Light sensor (MMC)

scintillating bolometer (Debye temp. = 438 K)

metallic magnetic calorimeter, fast signals

light yield at low temp. ~ 30,000 photon / MeV

alpha / gamma separation

$^{40}\text{Ca}^{100}\text{MoO}_4$ Crystal

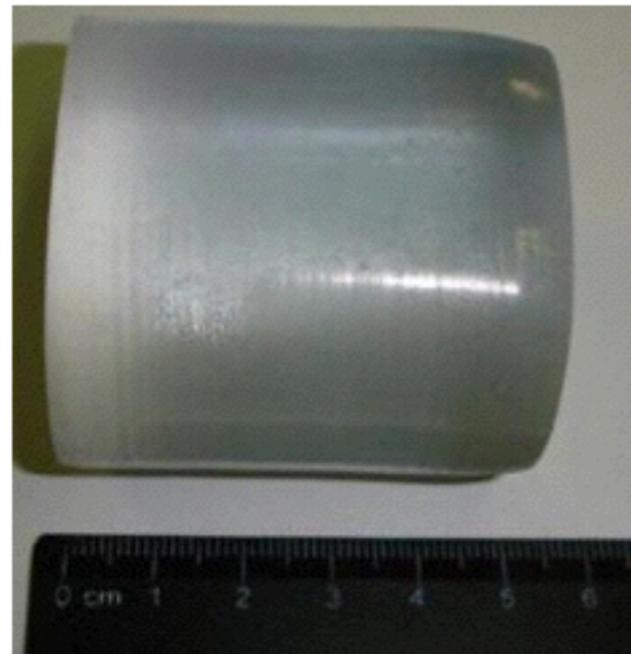
- **SB28**

weight 196 g



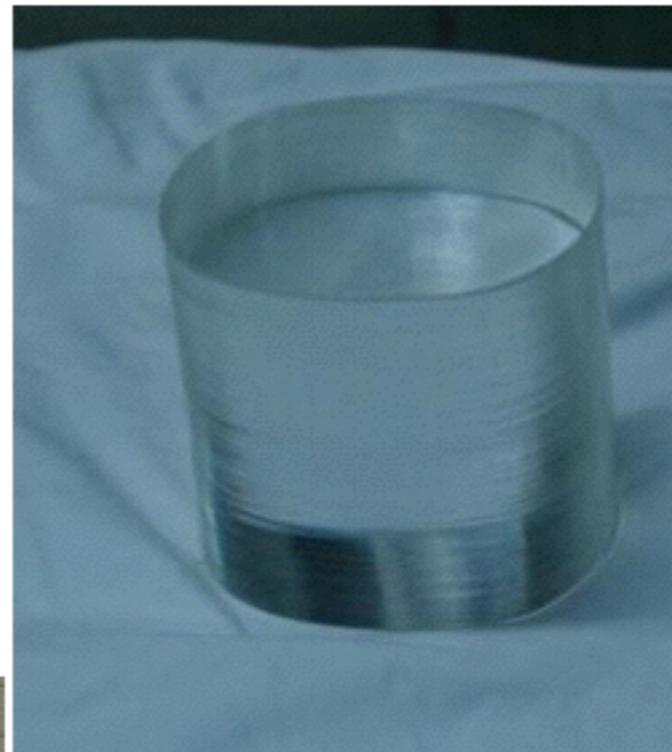
- **SB29**

weight 390 g



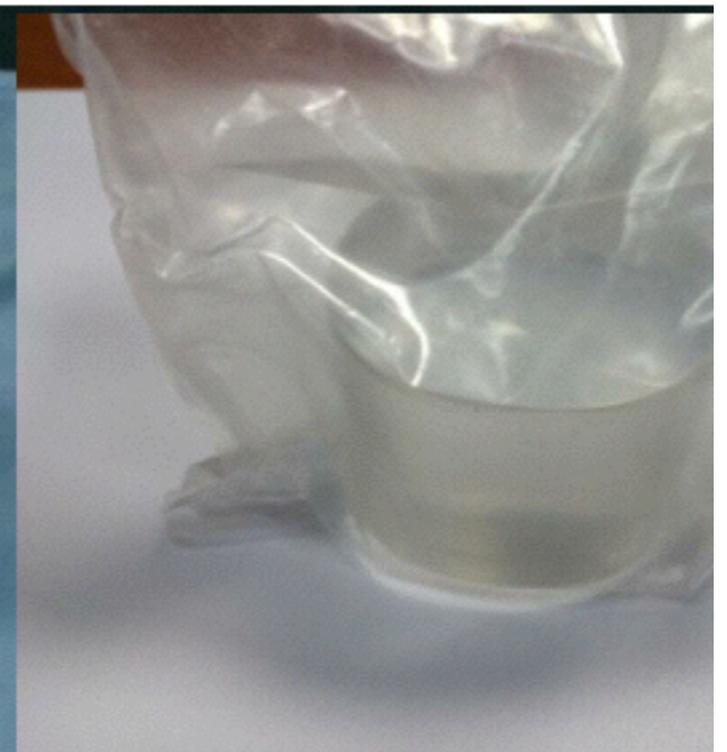
- **S35**

weight ~300 g



- **SS68**

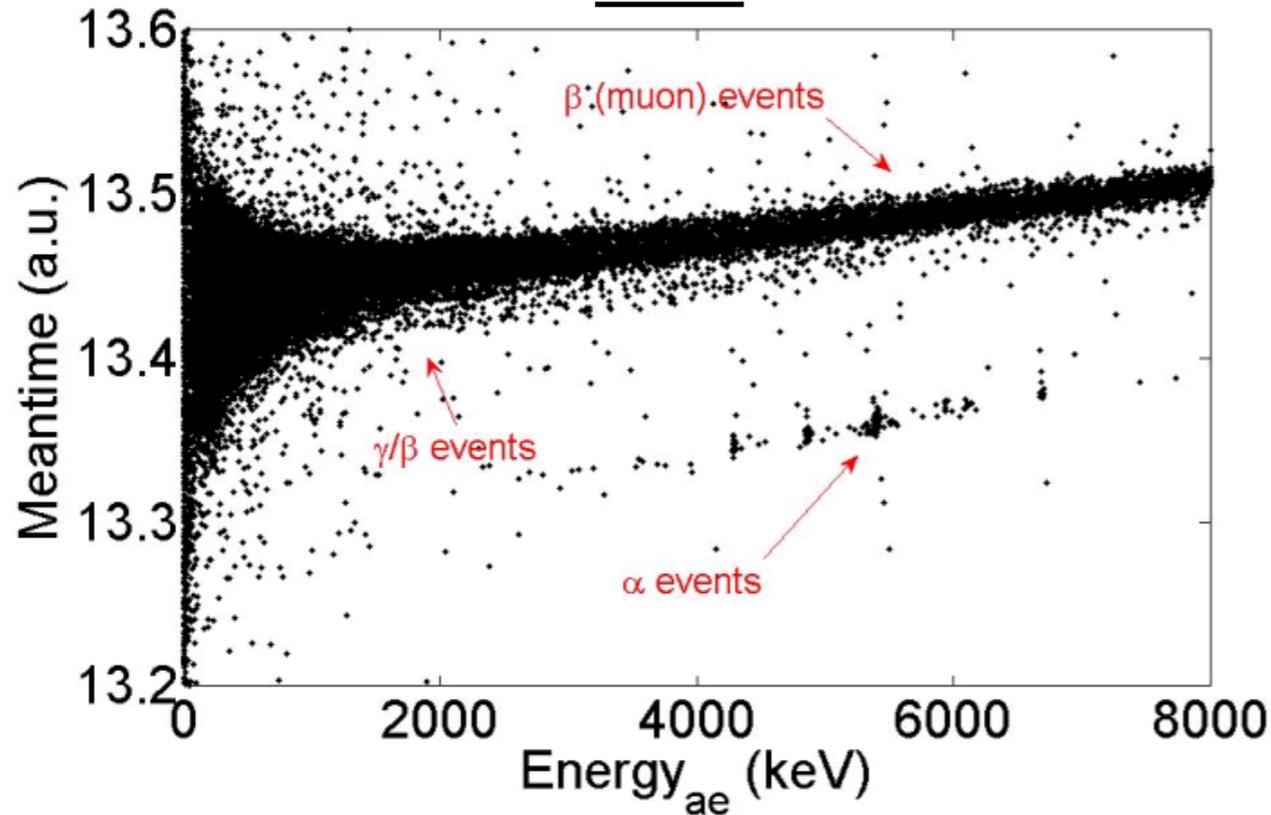
weight ~300 g



**New S68
Under test
(Nb doped)**

Detector Performance

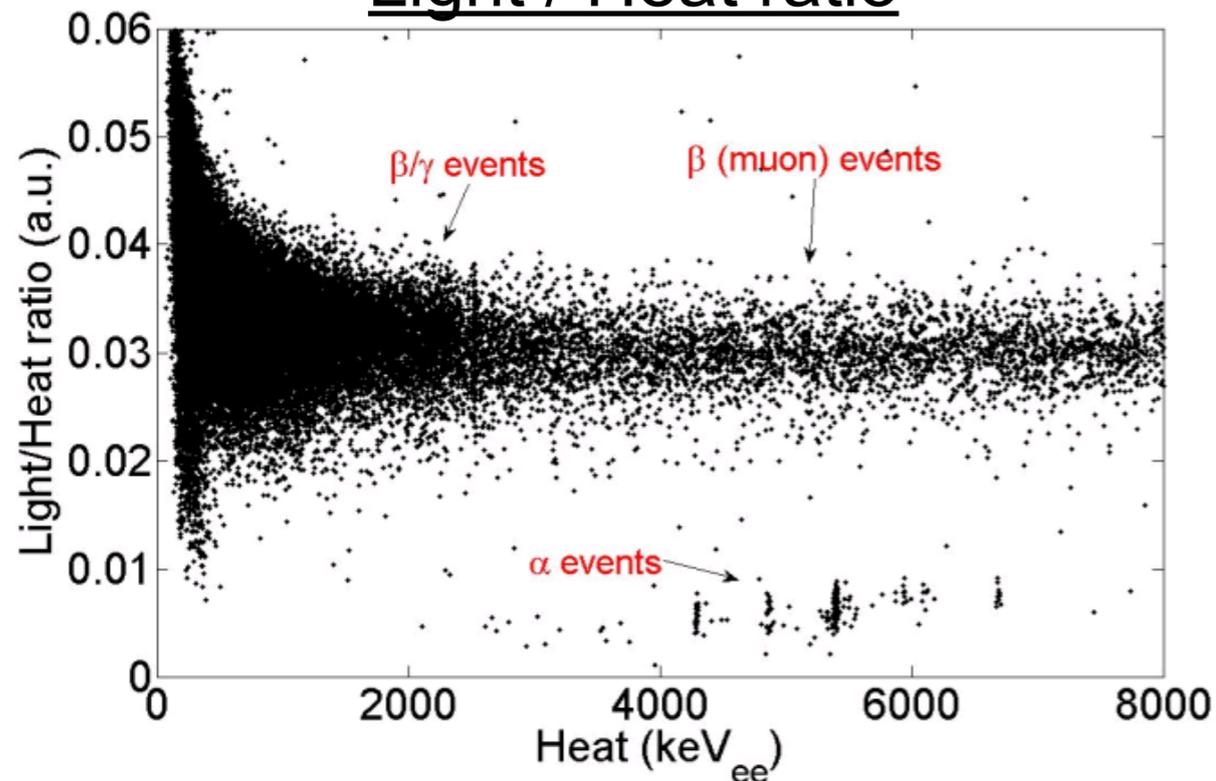
PSD



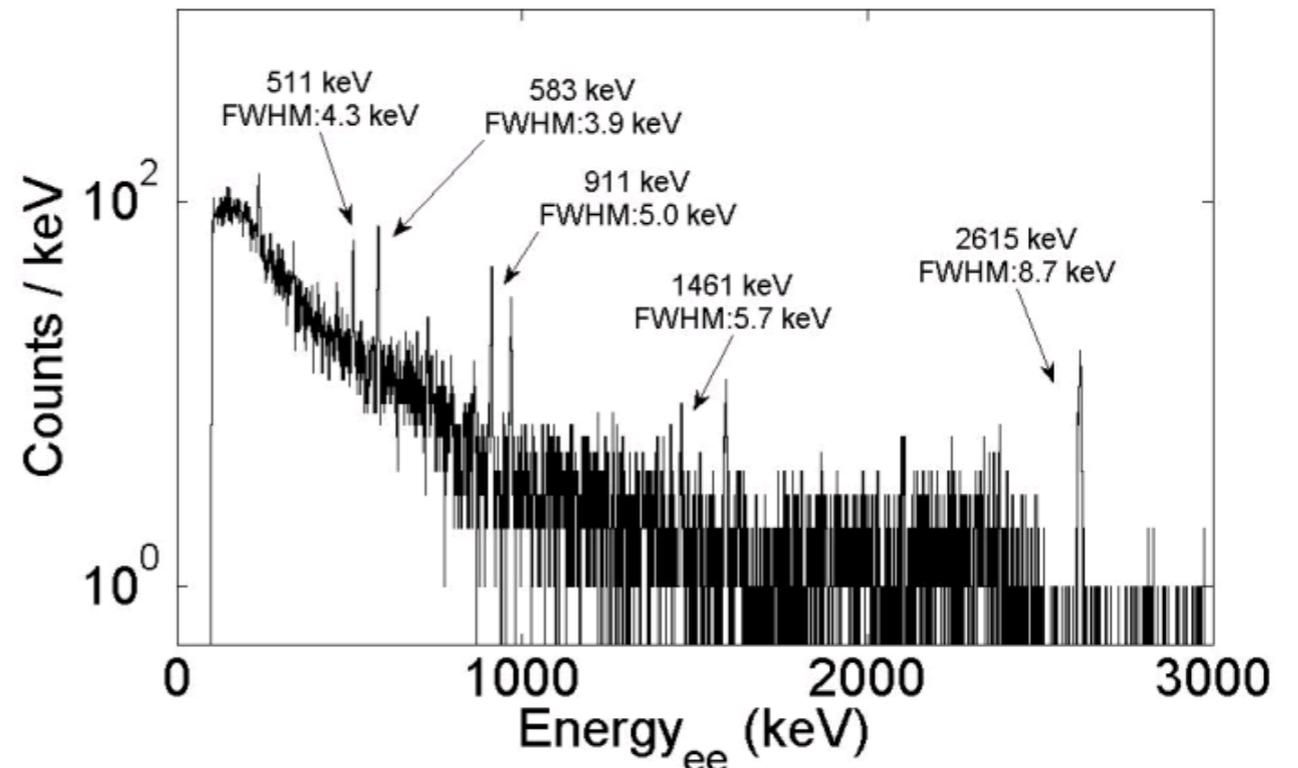
measurement @ over-ground

- α / e separation in **PSD and Light / Heat ratio**
- **high energy resolution** with cryogenic technique

Light / Heat ratio



Energy Spectrum @ 10 mK

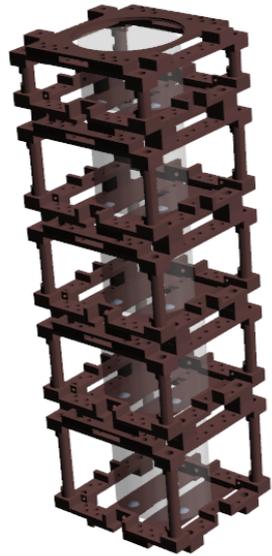


energy resolution :

FWHM = 8.7 keV @ 2615 keV

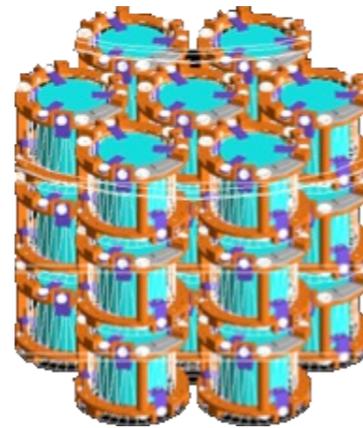
AMoRE Prospect

Pilot



5 CMOs ~ 1 kg

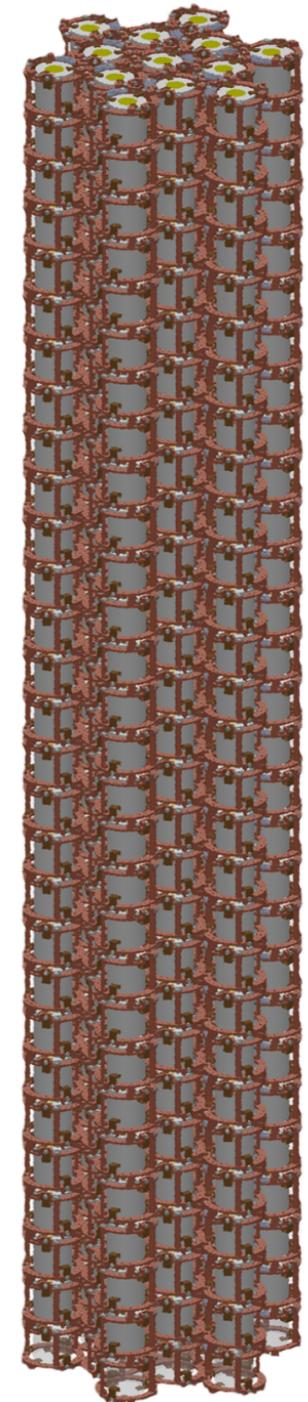
Phase I



CMO ~ 300 g

5 layers - 7 columns
10 kg

Phase II



200 kg

	Pilot	Phase I	Phase II
Mass	1 kg	10 kg	200 kg
Sensitivity(m_{ee}) (meV)	300-900	60-180	10-30
Location	Y2L	Y2L	New Lab
Schedule	2015	2016	2019

Summary

- Neutrino mass measurements in Asia are reviewed

Double Beta Decay

Neutrino Oscillation

Japan

KamLAND-Zen ^{136}Xe

Super-K / T2K **MH test**

CANDLES ^{48}Ca

KamLAND

DCBA ^{150}Nd , ^{82}Se , ^{100}Mo

Hyper-K (future)

AXEL (future) ^{136}Xe

ZICOS (future) ^{96}Zr

Korea

AMoRE ^{100}Mo

RENO

RENO-50 (future)

China

Daya Bay

JUNO (future)

India

INO (future)

- Several double-beta decay experiments will reach the **sensitivity to test the inverted hierarchy** in the near future