

HEAVY NEUTRINOS IN PARTICLE PHYSICS AND COSMOLOGY

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2013 review [arXiv:1303.6912 \[hep-ph\]](#)
update [arXiv:1502.06891 \[hep-ph\]](#)

23. 7. 2015
EPS-HEP meeting Vienna, Austria

Three Generations of Matter (Fermions) spin $\frac{1}{2}$

	I	II	III
mass →	2.4 MeV	1.27 GeV	171.2 GeV
charge →	$\frac{2}{3}$	$\frac{2}{3}$	$\frac{2}{3}$
name →	u up	c charm	t top
Quarks	4.8 MeV $-\frac{1}{3}$ d down	104 MeV $-\frac{1}{3}$ s strange	4.2 GeV $-\frac{1}{3}$ b bottom
	0 eV 0 ν_e electron neutrino	0 eV 0 ν_μ muon neutrino	0 eV 0 ν_τ tau neutrino
	0.511 MeV -1 e electron	105.7 MeV -1 μ muon	1.777 GeV -1 τ tau
Leptons			

0
0
g
gluon

0
0
 γ
photon

91.2 GeV
0
Z⁰
weak
force

80.4 GeV
 ± 1
W[±]
weak
force

125 GeV
0
0
H
Higgs
boson

spin 0

Bosons (Forces) spin 1

Neutrino masses: Seesaw mechanism

$$\mathcal{L} = \mathcal{L}_{SM} + i\bar{\nu}_R \not{\partial} \nu_R - \bar{L}_L F \nu_R \tilde{H} - \bar{\nu}_R F^\dagger L \tilde{H}^\dagger - \frac{1}{2}(\bar{\nu}_R^c M_M \nu_R + \bar{\nu}_R M_M^\dagger \nu_R^c)$$

Minkowski 1979, Gell-Mann/Ramond/Slansky 1979, Mohapatra/Senjanovic 1979, Yanagida 1980

$$\Rightarrow \frac{1}{2}(\bar{\nu}_L \quad \bar{\nu}_R^c) \begin{pmatrix} 0 & m_D \\ m_D^T & M_M \end{pmatrix} \begin{pmatrix} \nu_L^c \\ \nu_R \end{pmatrix}$$

two sets of Majorana mass states with mixing $\theta = m_D M_M^{-1} = v F M_M^{-1}$

- **three light "active" neutrinos** $\nu \simeq U_\nu(\nu_L + \theta \nu_R^c)$
 - mostly "active" SU(2) doublet
 - light masses $m_\nu \simeq \theta M_M \theta^T = v^2 F M_M^{-1} F^T$
- **three heavy "sterile" neutrinos** $N \simeq \nu_R + \theta^T \nu_L^c$
 - mostly "sterile" singlets
 - heavy masses $M_N \simeq M_M$
- Majorana masses M_M introduce **new mass scale(s)**
- new heavy states only interact via **small mixing** $\theta \ll 1$

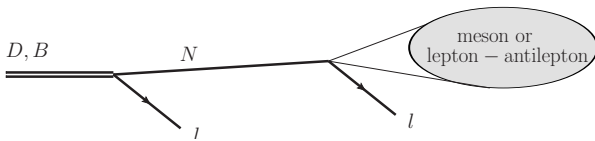
Where to see the N_i

● Indirect searches

- neutrino oscillation data
- LFV in rare lepton decays
- violation of lepton universality,
- (apparent) violation of CKM unitarity
- neutrinoless double β -decay
- EW precision data

● Direct searches

- LNV and LFV in gauge boson or meson decays

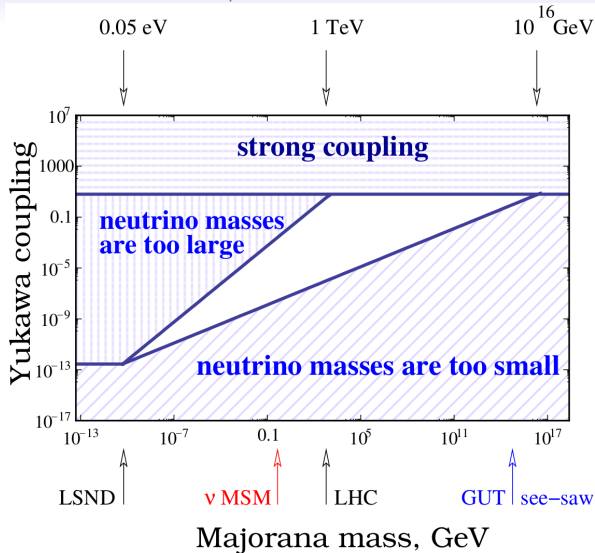


- displaced vertices
- peak searches, missing 4-momentum

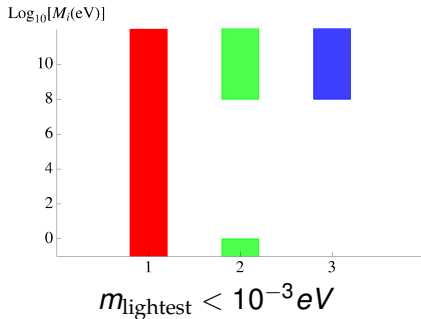
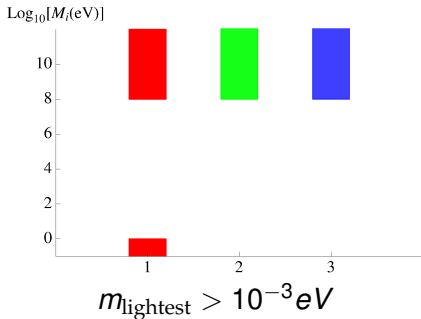
● Cosmology: BBN and N_{eff}

ν -oscillation data and the seesaw scale

plot from 1204.5379

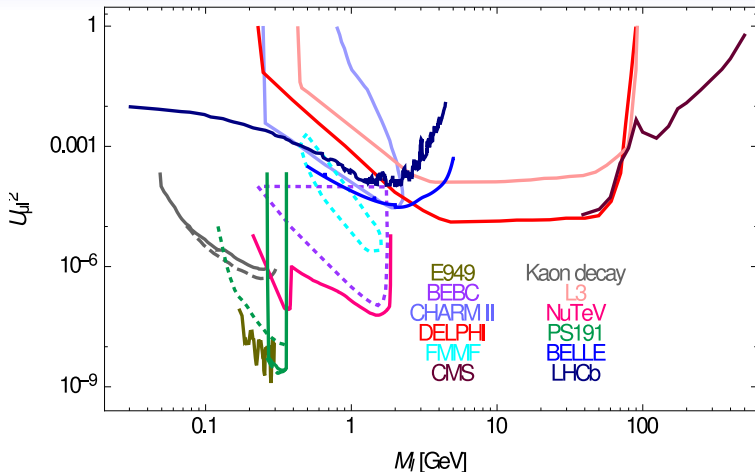


Bounds from cosmology: N_{eff} and BBN



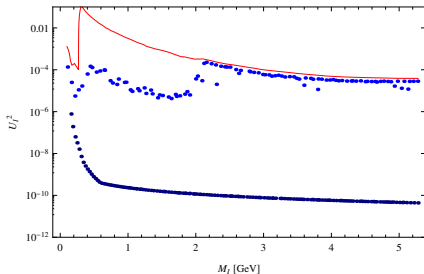
from Hernandez/Kevic/Lopez-Pavon 1406.2961

Bounds from Colliders



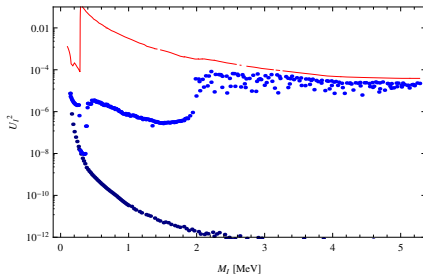
plot from MaD/Garbrecht 1502.00477

Combining direct and indirect constraints



$$m_{\text{lightest}} = 0.23\text{eV}$$

plot from MaD/Garbrecht 1502.00477

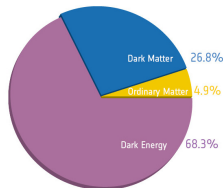


$$m_{\text{lightest}} = 0\text{eV}$$

What can RH neutrinos do for you?

Neutrino masses and...

- $M > 100 \text{ MeV}$: **Leptogenesis**
CP-violating interactions of RH neutrinos can generate a matter-antimatter asymmetry in the early universe.
- $M \sim \text{keV}$: **sterile neutrino Dark Matter**
RH neutrinos with tiny mixing θ are long lived massive particles and obvious DM candidates.
- $M \lesssim \text{eV}$: **oscillation anomalies and Dark Radiation**
Light sterile neutrinos could explain oscillation anomalies (LSND, Gallium, reactor) and contribute to N_{eff} in the early universe.



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HEAVY NEUTRINOS IN PARTICLE PHYSICS AND COSMOLOGY

- **GUT seesaw**

- naturally fits into GUTs
- naturally gives neutrino masses
- naturally does leptogenesis

- not observable
- adds to hierarchy problem

- **electroweak or TeV seesaw**

- common origin with EW scale?
- gives neutrinos masses
- allows for leptogenesis
- accessible to LHC

- requires some "tuning" or an approximate L -conservation
- not "natural"

- **GeV seesaw**

- gives neutrino masses
- does leptogenesis
- accessible to LHC, BELLE, SHiP
- part of minimal ν MSM

- requires some "tuning" or an approximate L -conservation
- not "natural"
- origin of this scale?

- **keV seesaw**

- viable DM candidate

- cannot give neutrino masses & be DM at the same time
- origin of this scale & mixing?

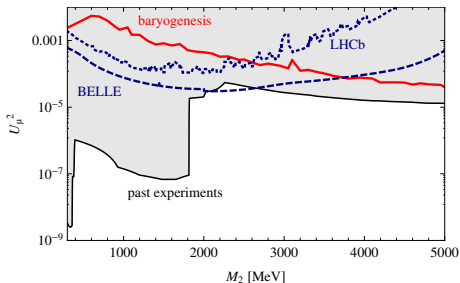
- **eV seesaw**

- LSND, gallium, reactor anomaly
- "Dark Radiation" N_{eff}

- cannot explain neutrino mass anomalies at the same time
- origin of this scale & mixing?

Low scale leptogenesis

- baryon asymmetry can be produced in the early universe
 - during N_i decay Fukugita/Yanagida 1986, Pilaftsis 2004
 - during N_i production Akhmedov/Rubakov/Smirnov 1998, Asaka/Shaposhnikov 2005
- N_i can be found at LHC, BELLE II or SHiP plot: Canetti/MaD/Garbrecht 1404.7114



- relevant CP-violation can be observable
 - can work with PMNS-phases alone 1208.4607
 - sterile sector CP-violation may also be observable 1403.2555

keV Masses: Sterile Neutrino Dark Matter?

- Where is the decay line? Very active discussion of 3.5 keV excess...
- How were they produced?
- Are they consistent with structure formation?

Upcoming White Paper:

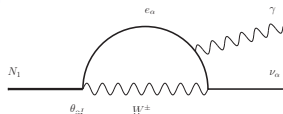
http://irfu.cea.fr/en/Phoce/Vie_des_labos/Ast/ast_visu.php?id_ast=3446

HEAVY NEUTRINOS IN PARTICLE PHYSICS AND COSMOLOGY

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- radiative decay $N \rightarrow \nu_L \gamma$
- Search for X-ray line!



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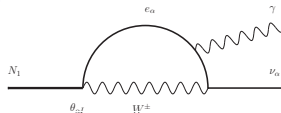
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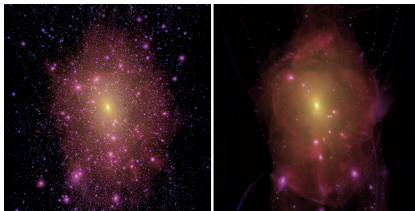
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- How were they produced?
- Are they consistent with structure formation?
 - DM is absolutely essential to form structures in the universe
 - DM is “cold” , i.e. $\langle \mathbf{k} \rangle < M$ at freezeout

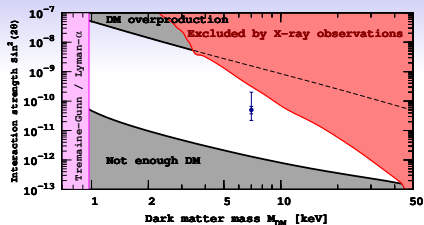


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astro/cosmology status early 2014

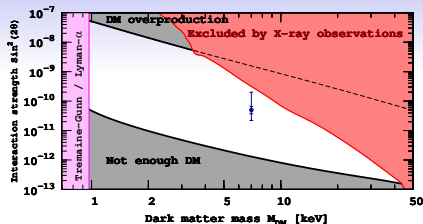
plot from 1402.4119, see also 1402.2301



Now: very active discussion 1405.7943, 1408.1699, 1408.3531, 1408.4388 and many more

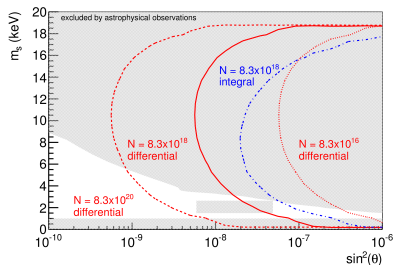
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plot from 1402.4119, see also 1402.2301



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Potential of KATRIN 1409.0920, see also 1404.5955



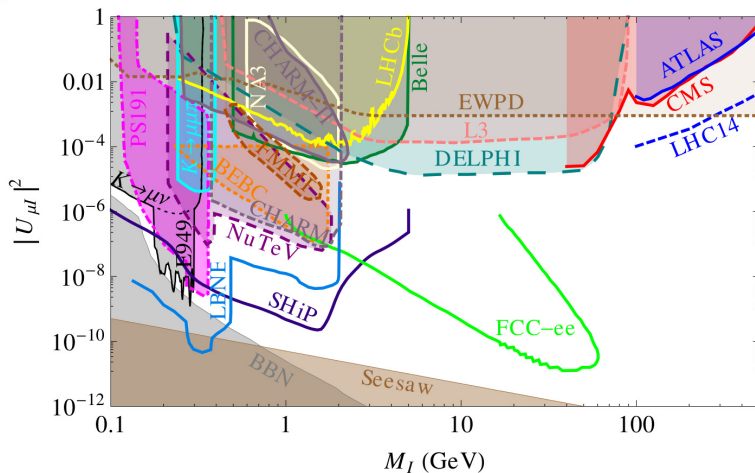
What's the most promising experiment? ... of course it depends on the mass scale...

Some selected comments:

- **neutrinoless double β -decay**
 - very sensitive, clear BSM signal
 - but can be hidden if approx. B-L conserving scenarios
- **lepton flavour violation**
 - may be observable in approx. B-L conserving scenario
- **lepton universality**
 - may be observable in approx. B-L conserving scenario
- **direct searches**
 - EW/TeV scale: LHC, FCC-ee
 - GeV scale: SHiP, LHCb, BELLE II

Also important: absolute mass scale, hierarchy, Dirac-phase

Future collider searches



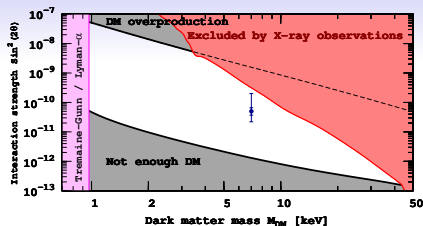
Plot from arXiv:1504.04855 [hep-ph]

Summary

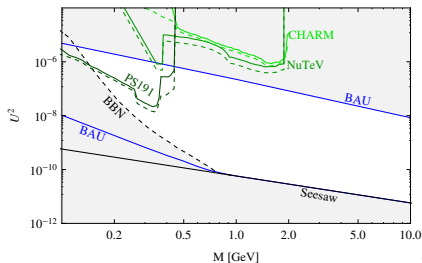
- ν -oscillations are the only BSM signal seen in the lab
definitely require new BSM degrees of freedom!
- the new particles are RH neutrinos, they may be related to
cosmological puzzles (Dark Matter, baryogenesis, Dark Radiation)
- if new particles are below the electroweak scale, they can be found experimentally \Rightarrow **experimental search for exciting New Physics!**
- even if they are heavier, indirect probes involve
 - neutrino oscillation experiments
 - neutrinoless double β -decay
 - lepton flavour violation
 - lepton universality violation
 - unitarity of the observed CKM matrix

We are looking forward to exciting new data...

The ν MSM: heavy neutrinos solve all problems! Asaka/Shaposhnikov



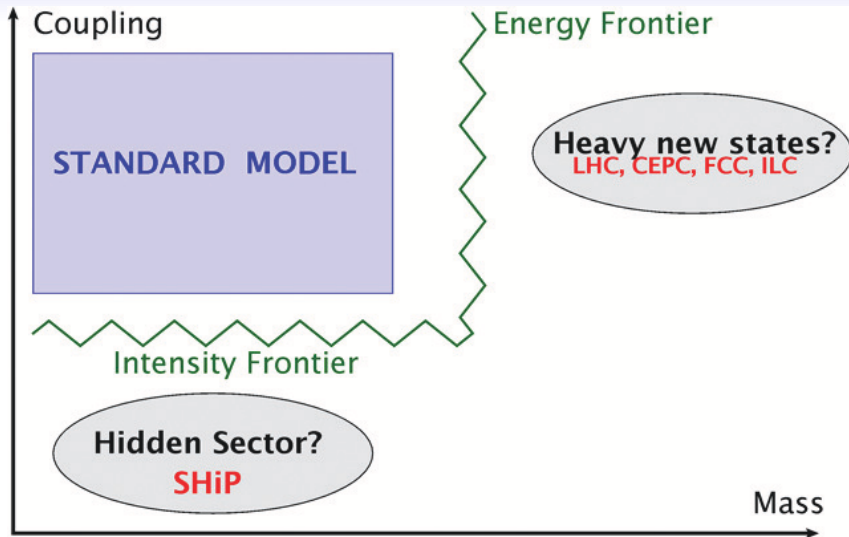
Boyarsky/Ruchayskiy/Iakubovskiy/Franse 1402.4119



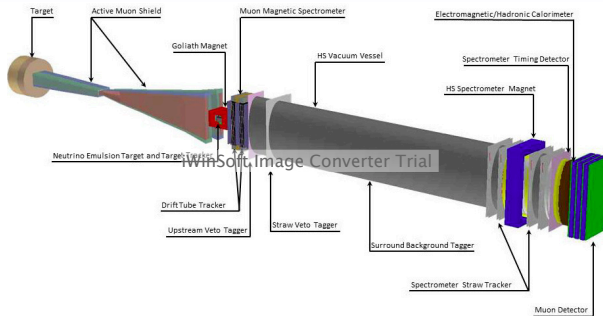
Canetti/MaD/Frossard/Shaposhnikov 1204.3902, 1208.4607

DM, Baryogenesis and neutrino masses from RH neutrinos!

Where is the New Physics hiding?



The SHiP Experiment



- intensity frontier experiment using CERN SPS beam
- fixed target experiment with strong shield
- technical report arXiv:1504.04956 [physics.ins-det]

The SHiP Experiment

Search for **H**idden **P**articles

- neutrino portal
- scalar portal
- vector portal
- axion-like particles
- ν_τ physics
- LFV in τ -decays
- very light neutralino?
- **your proposal!**

see arXiv:1504.04855 [hep-ph] for details

great opportunity at the intensity frontier - also for China