

Laser-plasma interactions

Reference material:

"Physics of laser-plasma interactions" by W.L.Kruer

"Scottish Universities Summer School Series on Laser Plasma Interactions" number 1 edited by R. A. Cairns and J.J.Sanderson number 2 edited by R. A. Cairns number 3 - 5 edited by M.B. Hooper

Outline

- 1. Introduction Ionisation processes
- 1. Laser pulse propagation in plasmas Resonance absorption Inverse bremsstrahlung Not-so-resonant resonance absorption (Brunel effect)
- 3. Ponderomotive force
- 4. Instabilities in laser-produced plasmas Stimulated Raman scatter (SRS) instability Stimulated Brillouin scatter (SBS) instability Two plasmon decay Filamentation

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1. Introduction

Ionisation processes

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Structure of a laser produced plasma

Laser systems at the CLF

Laser systems for laser plasma research

Vulcan: Nd: glass system

- long pulse (100ps-2ns), 2.5 kJ, 8 beams
- 100TW pulse (1ps) and 10TW pulse (10ps) synchronised with long pulse
- Petawatt (500fs, 500J)
- Upgrading to 10 PW peak power

UK Flagship 2002-2008: Vulcan facility

Increased complexity, accuracy & rep-rate

UK Flagship 2008-2013: Astra-Gemini facility

Dual Beam Petawatt Upgrade of Astra (factor 40 power upgrade) 10²² W/cm² on target irradiation 1 shot every 20 seconds Opened by the UK Science Minister (Ian Pearson MP) Dec 07 Significantly over-subscribed. Assessing need for 2nd target area.

Ian Pearson, MP *Minister of State for Science and Innovation* , *Dec 2007*

The next step: Vulcan 10 PW (>10²³ W/cm²)

Introduction

Power and Intensity

 $1 \text{PW} = 10^{15} \text{W}$

Power = Energy = $500 \text{ J} = 1 \times 10^{15} \text{ W}$ pulse duration 500 fs

To maximise the intensity on target, the beam must be focused to a small spot. The focal spot diameter is $7 \mu m$ and is focused with an f/3.1 off-axis parabolic mirror. 30% or the laser energy is within the focal spot which means

diffraction gratings

In these lecture notes, the units are defined in cgs. Conversions to SI units are given here in the hand-outs.

But note that in the literature

Energy – Joules Pulse duration τ - ns nanoseconds 10^{-9} sec (ICF) ps picoseconds 10-12 sec fs femtoseconds 10-15 sec (limit of optical pulses) as attoseconds 10^{-18} sec (X-ray regime)

Distance – microns (μm)

Power – Gigawatts 10^9 Watts

Intensity – W $cm⁻²$ Terawatts 10^{12} Watts Irradiance – W cm⁻² μ m² Petawatts 10¹⁵ Watts Pressure – Mbar – 10¹¹ Pa Exawatts 10^{18} Watts Magnetic field – MG – 100 T

FORMULA CONVERSION⁸

Here $\alpha = 10^2$ cm m⁻¹, $\beta = 10^7$ erg J⁻¹, $\epsilon_0 = 8.8542 \times 10^{-12}$ F m⁻¹, $\mu_0 = 4\pi \times 10^{-7}$ H m⁻¹, $c = (\epsilon_0 \mu_0)^{-1/2} = 2.9979 \times 10^8$ m s⁻¹, and $\hbar = 1.0546 \times$ 10^{-34} J s. To derive a dimensionally correct SI formula from one expressed in Gaussian units, substitute for each quantity according to $\overline{Q} = \overline{k}Q$, where \overline{k} is the coefficient in the second column of the table corresponding to Q (overbars denote variables expressed in Gaussian units). Thus, the formula $\bar{a}_0 = \bar{\hbar}^2 / \bar{m} \bar{e}^2$ for the Bohr radius becomes $\alpha a_0 = (\hbar \beta)^2/[(m \beta/\alpha^2)(e^2 \alpha \beta/4\pi \epsilon_0)]$, or $a_0 =$ $\epsilon_0 h^2/\pi m e^2$. To go from SI to natural units in which $\hbar = c = 1$ (distinguished by a circumflex), use $Q = \hat{k}^{-1}\hat{Q}$, where \hat{k} is the coefficient corresponding to Q in the third column. Thus $\hat{a}_0 = 4\pi\epsilon_0\hbar^2/[(\hat{m}\hbar/c)(\hat{e}^2\epsilon_0\hbar c)] = 4\pi/\hat{m}\hat{e}^2$. (In transforming from SI units, do not substitute for ϵ_0 , μ_0 , or c.)

MAXWELL'S EQUATIONS

In a plasma, $\mu \approx \mu_0 = 4\pi \times 10^{-7}$ Hm⁻¹ (Gaussian units: $\mu \approx 1$). The permittivity satisfies $\epsilon \approx \epsilon_0 = 8.8542 \times 10^{-12}$ Fm⁻¹ (Gaussian: $\epsilon \approx 1$) provided that all charge is regarded as free. Using the drift approximation $v_{\perp} = E \times B/B^2$ to calculate polarization charge density gives rise to a dielectric constant $K \equiv \epsilon/\epsilon_0 = 1 + 36\pi \times 10^9 \rho/B^2$ (SI) = $1 + 4\pi \rho c^2/B^2$ (Gaussian). where ρ is the mass density.

The electromagnetic energy in volume V is given by

$$
W = \frac{1}{2} \int_{V} dV (\mathbf{H} \cdot \mathbf{B} + \mathbf{E} \cdot \mathbf{D})
$$
 (SI)

$$
= \frac{1}{8\pi} \int_{V} dV (\mathbf{H} \cdot \mathbf{B} + \mathbf{E} \cdot \mathbf{D})
$$
 (Gaussian).

Poynting's theorem is

$$
\frac{\partial W}{\partial t} + \int_S \, \mathbf{N} \cdot d\mathbf{S} = - \, \int_V \, dV \, \mathbf{J} \cdot \mathbf{E},
$$

where S is the closed surface bounding V and the Poynting vector (energy flux across S) is given by $N = E \times H$ (SI) or $N = cE \times H/4\pi$ (Gaussian).

http://wwppd.nrl.navy.mil/formulary/NRL_FORMULARY_07.pdf

• The laser can propagate in a plasma until what is known as the critical density surface. The dispersion relationship for electromagnetic waves in a plasma is $\omega^2 = \omega_{pe}^2 + k^2 c^2$ so that when $\omega = \omega_{pe}$ the wave cannot propagate

Critical density

$$
n_{critical} = 1.1 \times 10^{21} \text{ cm}^{-3}
$$
\n
$$
\lambda_{\mu m} \text{ is the laser wavelength in } \mu\text{m}
$$

• An electron oscillates in a plane electromagnetic wave according to the Lorentz force $F = -e(\underline{E} + c^{-1}\underline{v} \times \underline{B})$ Neglecting the $v \times B$ term, then the <u>quiver velocity</u> v_{osc} is

$$
v_{osc} = \frac{eE_L}{m\omega}
$$

c v • The dimensionless vector potential is defined as $a = \frac{v_{osc}}{v_{osc}}$

Note that a can be > 1. This is because $a = \gamma \beta = \gamma v_{electron} / a$ and

$$
a = \left[\frac{I\lambda^2}{1.4 \times 10^{18} Wcm^{-2}}\right]^{1/2} \sim 1
$$
 if relativity is important and $\gamma_{electron} = \left[1 + a^2\right]^{1/2}$

Laser-plasmas can access the high energy density regime

• The ratio of the quiver energy to the thermal energy of the plasma is given by

TeV I kT $\frac{1}{2}mv_{osc}^2 = \frac{1.8 \times 10^{-18} I \lambda^2}{1.8 \times 10^{-18} I}$ 2 $\times 10^{-18} I \lambda^2$ $=$.

< 1 for ICF plasmas >> 1 for relativistic intensities

• The laser electric field is given by

$$
E_0 = \frac{3.2a_0}{\lambda} \left[10^{10} \right] \text{ Vcm}^{-1}
$$

• The Debye length λ_D for a laser-produced plasma is $\sim \mu m$ for gas targets and \sim nm for solids. n λ_D - 50 -100

• Strongly coupled plasma regime can be accessed

Map of the HED Universe

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• **Multi-photon ionisation**

Non linear process Very low rate Scales as I^n n = number of photons

• **Collisional ionisation (avalanche ionisation)**

Seeded by multiphoton ionisation Collisions of electrons and neutrals \Leftrightarrow plasma

$$
v \propto \frac{{n_e}^2 T_e^{-3/2}}{\omega^2}
$$

•**Tunnel and field ionisation**

Laser electric field "lowers" the Coloumb barrier which confines the electrons in the atom

Tunnelling

Isolated atom Atom in an electric field

Atom in an oscillating field

Keldysh parameter γ = time to tunnel out / period of laser field

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EM wave propagation in plasma

 $j = \sigma E$ where the conductivity σ is $\sigma = \frac{pe}{4\pi\omega}$ ω $\sigma = \frac{1}{4}$ $i\omega_{pe}^2$ The current density can be written $i = \sigma E$ where the conductivity σ is $\sigma =$ and $\omega_{pe}^2 = 4\pi e^2 n_0 / m_e$ $\omega_{ne}^2 = 4\pi e^2$ Start with an electric field of the form $\underline{E} = \underline{E}(k) \exp(-i\omega t)$ (1)

 $\overline{}$ $\overline{}$ \int $\bigg)$ $\overline{}$ $\overline{}$ \setminus $\sqrt{2}$ $=\left(1-\frac{\omega_{pe}}{a^2}\right)$ 2 1 ω ω Where we have defined the dielectric function of the plasma $\mathcal{E} = \left| 1 - \frac{\omega_{pe}}{n^2} \right|$ *E c i E c i E c* $\underline{B} = \frac{4\pi}{\sigma} \sigma \underline{E} - \frac{i\omega}{\sigma} \underline{E} = -\frac{i\omega}{\sigma} \varepsilon$ σ $\nabla \times \underline{B} = \frac{4\pi}{\sigma} \sigma \underline{E} - \frac{i\omega}{\sigma} \underline{E} = -$ 4 *B c i* Substituting (1) into Faraday's law $\nabla \times \underline{E} = -\frac{1}{2} \frac{\partial B}{\partial x}$ gives $\nabla \times \underline{E} = -\frac{1}{2} \frac{\partial B}{\partial y}$ Substituting (1) and **j** into Ampere's law $\nabla \times \underline{B} = \frac{m}{n} J + \frac{c}{n}$ gives (2) *t B c E* ∂ ∂ $\nabla \times \underline{E} = -$ 1 gives *t E c J c B* ∂ ∂ $\nabla \times \underline{B} = \frac{-\pi}{\mu} I +$ 4π ₁ 1 (3) Taking the curl of Faraday's law (2) and substituting Ampere's law (3) with a

standard vector identity gives

$$
\nabla^2 \underline{E} - \nabla(\nabla \cdot \underline{E}) + \frac{\omega^2}{c^2} \varepsilon \underline{E} = 0 \qquad \text{i.e.} \qquad \nabla^2 \underline{E} + \frac{\omega^2}{c^2} \varepsilon \underline{E} = 0 \qquad (4)
$$

EM wave propagation in a plasma with a linear density gradient

Consider plane waves at normal incidence. Assume the density gradient is in the propagation direction Z. This

$$
n_0 = n_0(z)
$$

\n
$$
\varepsilon = \varepsilon(\omega, z)
$$
 (5)
\n
$$
\underline{E}(x) = \underline{E}(z) \exp(-i\omega t)
$$

Substituting (5) into (4) gives
$$
\frac{d^2}{dz^2} E_{x,y} + \frac{\omega^2}{c^2} \varepsilon E_{x,y} = 0
$$
(6)

$$
\varepsilon E_z = 0
$$

We can rewrite the dielectric function as
$$
\varepsilon = 1 - \frac{n(x)}{n_c} = 1 - \frac{\omega_{pe}^2}{\omega^2} = 1 - \frac{z}{L}
$$
 (7)

when
$$
n(x) = \frac{z}{L} n_c
$$
 and the critical density is $n_c = \frac{m\omega^2}{4\pi e^2}$

Equation (6) can then be written as $\frac{u}{dx^2} + \frac{dv}{dx^2} \left(1 - \frac{2}{l}\right)E = 0$ 2 2 2 $E=$ \int \backslash $\overline{}$ \setminus $\bigg($ $+\frac{\omega}{2}$ 1 - $\frac{2}{\tau}$ | E *L z* dz^2 *c* $\frac{d^2 E}{dr^2} + \frac{\omega^2}{r^2} \left(1 - \frac{z}{r}\right) E = 0$ (8)

EM wave propagation in a plasma with a linear density gradient continued (1)

Change variables
$$
\eta = \left(\frac{\omega^2}{c^2 L}\right)^{1/3} (z - L)
$$

Gives $\frac{d^2 L}{dx^2} - \eta E = 0$ whose solution is an Airy function 2 $-\eta$ *E* = *d* d^2E η η

$$
E(\eta) = \alpha A_i(\eta) + \beta B_i(\eta) \qquad (9)
$$

 α and β are constants determined by the boundary condition matching.

Physically, E should represent a standing wave for $\eta < 0$ and to decay at $\eta \rightarrow \infty$. Since $B_i(\eta) \rightarrow \infty$ as $\eta \rightarrow \infty$, β is chosen to $\beta = 0$. α is chosen by matching the E-fields at the interface between the vacuum and the plasma at *z*=0. This gives $\eta = (\omega L/c)^{2/3}$. If we assume that $\omega L/c >>1$ and represent $A_i(\eta)$ by the expression

$$
A_i(-\eta) = \frac{1}{\sqrt{\pi} \eta^{1/4}} \cos \left(\frac{2}{3} \eta^{2/3} - \frac{\pi}{4} \right) \tag{10}
$$

EM wave propagation in a plasma with a linear density gradient continued (2)

Then
$$
E(z=0) = \frac{\alpha}{2\sqrt{\pi}(\omega L/c)^{1/6}} \left[\exp i \left(\frac{2}{3} \frac{\omega L}{c} - \frac{\pi}{4} \right) + \exp - i \left(\frac{2}{3} \frac{\omega L}{c} - \frac{\pi}{4} \right) \right]
$$
 (11)

Now *E(z=0)* can be represented as the sum of the incident wave with amplitude E_{FS} with a reflected wave with the same amplitude but shifted in phase

$$
E(z=0) = E_{FS} \left[1 + \exp{-i(\frac{4}{3} \frac{\omega L}{c} - \frac{\pi}{2})} \right]
$$
 (12)

Provided that $\alpha = 2\sqrt{\pi} \left(\frac{\omega L}{c} \right)^2 E_{FS} e^{i\phi}$ *c* L $\big\rangle^{1/6}$ 2 / $\overline{}$ \int $\bigg)$ $\overline{}$ \setminus $\sqrt{2}$ $=$

E not affect E_{FS} is the free space value of the incident light and ϕ is an arbitrary phase that does

$$
\therefore E(\eta) = 2\sqrt{\pi} \left(\frac{\omega L}{c}\right)^{1/6} E_{FS} e^{i\phi} A_i(\eta) \tag{13}
$$

This function is maximised at $\eta = 1$ corresponding to (z - L)= -(c² L / ω^2)^{1/3}

$$
\therefore \left| \frac{E_{\text{max}}}{E_{FS}} \right|^2 \approx 3.6 \left(\frac{\omega L}{c} \right)^{1/3} \approx 3.6 \left(\frac{2\pi L}{\lambda} \right)^{1/3} \quad (14)
$$
(valid for L/ λ >1)

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Since ε is a function only of *z*, k_y must be conserved and hence $k_y = (\omega/c)\sin\theta$ and

$$
E_x = E(z) \exp\left(\frac{-i\omega y \sin \theta}{c}\right) \qquad (16)
$$

Substituting (16) into (15) gives (z) $\left(\varepsilon(z)-\sin^2\theta\right)E(z)=0$ 2 2 2 2 $+\frac{\omega}{z}\big(\varepsilon(z)-\sin^2\theta\big)E(z)=$ dz^2 *c* $\frac{d^2E(z)}{dz} + \frac{\omega^2}{2} (\varepsilon(z) - \sin^2 \theta)$ (17)

Reflection of the light occurs when $\varepsilon(z) = \sin^2 \theta$. Remember that $\varepsilon = (1 - \omega_{pe}^2(z) / \omega^2)$ and therefore reflection takes place when $\omega_{pe}^{\text{}}$ = ω cos θ at a density lower than critical given by $n_e = n_c \cos^2 \theta$.

Oblique incidence – resonance absorption

p-polarised

In this case, there is a component of the electric field that oscillates along the density gradient direction, i.e. \underline{E} . $\nabla n_e \neq 0$ Part of the incident light is transferred to an electrostatic interaction (electron plasma wave) and is no longer purely electromagnetic.

This means that there is a resonant response when $\varepsilon = 0$, i.e. when $\omega = \omega_{pe}$

We seen previously that the light does not penetrate to $\bm{{\mathsf n}}_{\rm c}.$ Instead,the field penetrates through to $\omega = \omega_{pe}$ to excite the resonance.

Oblique incidence – resonance absorption continued

 \wedge $\underline{B} = B_x \hat{x}$ and use the conservation of $k_y = \left(\frac{\omega}{c}\right) \sin \theta$ $\left.\rule{0pt}{10pt}\right)$ I \setminus ſ $=$ *c* Note that the magnetic field $\underline{B} = B_x \hat{x}$ and use the conservation of k_y

The B-field can be expressed as
$$
\underline{B} = B(z) \exp\left(-i\omega t + \frac{i\omega y}{c}\sin\theta\right) \hat{x}
$$
 (20)

Substituting into Ampere's law (equation (3)) $\nabla \times \underline{B} = -\frac{\mu \omega}{\sigma} \varepsilon \underline{E}$ gives *c i* $\nabla \times \underline{B} = -\frac{i\omega}{\epsilon} \varepsilon$

$$
E_z = \frac{\sin \theta}{\varepsilon(z)} B_z \approx \frac{E_d}{\varepsilon(z)}
$$
 (21)

This implies that E_z is strongly peaked at the critical density surface and the resonantly driven field E_d is evaluated there by calculating the B-field at $\boldsymbol{{\rm n}}_{\rm c.}$

From Faraday's law
$$
B = -\frac{ic}{\omega} \frac{\partial E}{\partial z}
$$
 or $B = \frac{-ic}{\omega} \left(\frac{\partial \eta}{\partial z}\right) \left(\frac{\partial E}{\partial \eta}\right)$ (22)
\n
$$
E(\eta) = 2\sqrt{\pi} \left(\frac{\omega L}{c}\right)^{1/6} E_{FS} e^{i\phi} A_i(\eta) \qquad B(\eta) = -i2\sqrt{\pi} \left(\frac{c}{\omega L}\right)^{1/6} E_{FS} e^{i\phi} A_i(\eta) \qquad (23)
$$

B decreases as E gets larger near the reflection point

At the reflection point,
$$
|B(\eta = 0)| \approx 0.9 \left(\frac{c}{\omega L}\right)^{1/6} E_{FS}
$$
 (24)

The B field decays like exp (- β) where $\beta = \int_{L \cos^2 \theta}^{L} k(z) dz$. *L* $\beta = \int_{L\cos^2\theta} k(z)dz$ (z)

$$
\beta = \int_{L\cos^2\theta}^{L} \frac{1}{c} \left(\omega_p^2 - \omega^2 \cos^2\theta \right)^{1/2} dz \qquad \frac{\omega_{pe}^2}{\omega^2} = \frac{z}{L} \tag{25}
$$

Solution: $\beta = \left(\frac{2\omega L}{2}\right) \sin^3 \theta$ 3 2 sin \backslash $\overline{}$ \setminus $\bigg($ $=$ *c L* (26)

$$
\therefore B(z = L) \approx 0.9 E_{FS} \left(\frac{c}{\omega L}\right)^{1/6} \exp\left(-\frac{2\omega L \sin^3 \theta}{3c}\right) \tag{27}
$$

 $\tau = \left(\frac{\omega L}{\omega}\right)$ sin θ 1/3 $\bigg\}$ \int \setminus $\overline{}$ \setminus $\bigg($ \equiv *c L* Defining (z) sin (z) $\varepsilon(z)$ *B z E* $E_z = \frac{L_d}{4}$ $z = \frac{\varepsilon}{\varepsilon}$ $\left(z \right) = \frac{\varepsilon}{\varepsilon}$ θ ${\cal E}$ and remember that $E_z = \frac{L_d}{c(s)}$

$$
\therefore E_d = \frac{E_{FS}}{(2\pi\omega L/c)} \phi(\tau) \qquad \phi(\tau) \approx 2.3\tau \exp\left(-\frac{2}{3}\tau^3\right) \qquad (28)
$$

The driver field E_d vanishes as $\tau \rightarrow 0$, as the component along z varies as sin θ .

Also, E_d becomes very small when $\tau \to \infty$, because the incident wave has to tunnel through too large a distance to reach n_c

The simple estimate for $\phi(\tau)$ from (28) is plotted here against an numerical solution of the wave equation.

Note that the maximum value of $\phi(\tau)$ is ~1.2 (needed to estimate the absorption fraction in the next section)

Resonance absorption – energy transfer

 $E_z \approx \frac{L_d}{4}$ $z \sim c$ Start from the electrostatic component of the electric field. \approx

Now include a damping term ν in the dielectric function of the plasma that can arise form collisions, wave particle interactions etc.

$$
\varepsilon(z) = 1 - \frac{\omega_p^2}{\omega(\omega - i\nu)}
$$

(*z*)

E

$$
I_{\text{abs}} = \int_0^\infty v \frac{E_z^2}{8\pi} dz = \frac{v}{8\pi} \int_0^\infty \frac{E_d^2(z)}{\left|\varepsilon\right|^2} dz
$$

 ν can be considered as the rate of energy dissipation and $E_z^2/8\pi$ the incident energy density*.*

For a linear density profile,
$$
(n_e = n_{cr} z / L)
$$
 $|\varepsilon|^2 = \left(1 - \frac{z}{L}\right)^2 + \left(\frac{v}{\omega}\right)^2 \frac{z^2}{L^2}$

Substituting and approximating that E_d is constant over a narrow width of the resonance function gives

$$
I_{\text{abs}} \approx \frac{vE_d^2(z=L)}{8\pi} \int_0^\infty \frac{dz}{(1-z/L)^2 + (v/\omega)^2}
$$

The integral gives $\pi \omega / v$, $\therefore I_{abs} \approx \omega L E_d^2 / 8$ 8π 2 *FS* $I_{abs} = f_A \frac{cE_{FS}^2}{8\tau}$ $f_A \approx \phi^2(\tau)/2$

I_{abs} peaks at ~0.5

The absorption is maximised at an angle given by

$$
\theta_{\text{max}} \approx \sin^{-1} \left[0.8(c/\omega L)^{1/3} \right]
$$

This is the characteristic signature of resonance absorption – the dependence on both angle of incidence and density scale-length.

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Collisional absorption

Principal energy source for direct drive inertial fusion plasmas

From the Vlasov equation, the linearized force equation for the electron fluid is

$$
\frac{\partial u_e}{\partial t} = -\frac{e}{m} \underline{E} - v_{ei} \underline{u_e}
$$

where v_{ei} is the damping term.

Since the field varies harmonically in time

$$
u_e = \frac{-ie\underline{E}}{m(\omega + i\nu_{ei})}
$$

The plasma current density is $J = -n_e e u_e = \frac{e^{i\omega_p} e^{-i\omega_p}}{4\pi\epsilon_0} \frac{E}{\sqrt{2\pi}}$ *i i* $J = -n_e e u$ *ei p e* $e^{\mathcal{L} u} e^{-\mathcal{L} u} \frac{d\mathcal{L} u}{4\pi(\omega + i v_{ei})}$ ω $\ddot{}$ $=-n_{e}e u_{e}$ = 4 2

The conductivity σ (J= σ E) σ = $\frac{mg_{pe}}{1.6 \times 10^{-4} m}$ and dielectric function $(\omega + i v_{ei})$ *pe i i* $\pi(\omega + i \nu)$ ω $\sigma = \frac{1}{4\pi(\omega + \pi)}$ $=$ 4 2 $(\omega + i v_{ei})$ *pe* $\omega(\omega+i\nu)$ ω $\mathcal{E} = 1 - \frac{1}{\omega(\omega + \omega)}$ $=1-$ 2 1

Collisional absorption

Faraday's law (cf (eqn (2)) Ampere's law (cf eqn (3))

$$
\nabla \times \underline{E} = -\frac{i\omega}{c} \underline{B}
$$

$$
\nabla \times \underline{B} = -\frac{i\omega}{c} \varepsilon \underline{E}
$$

Combining them gives

$$
\nabla^2 \underline{E} + \frac{\omega^2}{c^2} \varepsilon \underline{E} = 0
$$

To derive the dispersion relation for EM waves in a spatially uniform plasma, take $\underline{E}(x) \sim e^{i\underline{k} . \underline{x}}$ and substitute for ε and assume that $v_{\rm ei} / \omega$ << 1 gives

$$
\omega^2 = k^2 c^2 + \omega_{pe}^2 (1 - \frac{i v_{ei}}{\omega})
$$

Expressing $\omega = \omega_r - i \nu/2$ gives
$$
\omega_r = \left(\omega_{pe}^2 + k^2 c^2\right)^{1/2}
$$

$$
v = \frac{\omega_{pe}^2}{\omega_r^2} v_{ei}
$$

where v is the energy damping term.

collisional damping in linear density gradient

The rate of energy loss from the EM wave $[=(vE^2/8\pi)]$ must balance the oscillatory velocity of the electrons is randomised by electron- ion scattering. Assume again that $n_e = n_{cr} z / L$ and write down equation (4) again

$$
\frac{\partial^2 E}{\partial z^2} + \frac{\omega^2}{c^2} \varepsilon(z) E = 0
$$

 $(1+i\overline{v}_{e1}^*/\omega)$ $(1 + i v_{e1}^* /)$ $=1-\frac{2}{\sqrt{1-\frac{1}{\sqrt{1-\frac$

$$
\frac{d^2E}{dz^2} + \frac{\omega^2}{c^2} \left(1 - \frac{z}{L(1 + i v_{ei}^*/\omega)} \right) E = 0
$$

 $L(1+i \nu_e)$

 $\ddot{}$

Changing variables again, gives Airy's equation $\frac{u}{dx^2} - \eta \underline{E} = 0$ *d* $d^2 \underline{E}$ η η

 ${\cal E}$

1

The dielectric function $\varepsilon \left(z \right)$ is now

$$
\eta = \left(\frac{\omega^2}{c^2 L(1 + i v_{ei}^* / \omega)}\right)^{1/3} \left(z - L\left(1 + \frac{i v_{ei}^*}{\omega}\right)\right)
$$

$$
E(\eta) = \alpha \, A_i(\eta)
$$

 $\overline{}$ \int \backslash L \setminus \int $-\eta$) = $\frac{1}{\sqrt{2}} \cos \frac{2}{3} \eta^{2/3}$ – $3'$ 4 1 $2 \times 2^{2/3}$ $1/4$ π η $\pi\eta$ $(\eta) = \frac{1}{\sqrt{1-\frac{1}{4}}\cos{\frac{2}{3}}\eta^{2/4}}$ $A_i(-\eta) = \frac{1}{\sqrt{\pi n^{1/4}}} \cos$ Again, match the incident light wave to the vacuum plasma boundary interface at *z=0*

Energy absorption

Thus at *z=0*, *E* can be represented as an incident and reflected wave whose amplitude is multiplied by *e i*

$$
\phi = \frac{4}{3}(-\eta(z=0)) - \frac{\pi}{2}
$$

As η is now complex, there is both a phase shift and a damping term for the reflected wave

$$
\eta(0) = -\left(\left(\frac{\omega L}{c} \right) \left(1 + \frac{i v_{ei}^*}{\omega} \right) \right)^{2/3}
$$

For
$$
v^*_{ei}/\omega \ll 1
$$
 $\phi_{real} = \frac{4\omega L}{3c} - \frac{\pi}{2}$ $\phi_{imaginary} = \frac{4v^*_{ei}L}{3c}$

Energy ~ (amplit ude)² ~
$$
\exp\left(-\frac{8 v_{ei} L}{3 c}\right)
$$

$$
\therefore f_A = 1 - \exp\left(-\frac{8v_{ei}^*L}{3c}\right) \qquad \qquad v_{ei} \approx 3 \times 10^{-6} \ln \Lambda \frac{n_e Z}{T_{eV}^{3/2}} \sec^{-1}
$$

Fig. II-2 Absorption coefficient as a function of laser intensity and laser wavelength

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- 1. Introduction Ionisation processes
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Collisionless "Brunel" absorption

- •Dominant mechanism is $L/\lambda < 1$
- •Quiver motion of the electrons in the field of a p-polarised laser pulse
- Laser energy is absorbed on each cycle
- Can result in higher absorption than classical resonance absorption (up to 0.8)
- •Not as dependent on angle

From P.Gibbon *Phys. Rev. Lett.* **76**, 50 (1996)

Harmonic generation: the moving mirror model

see references:

- D. von der Linde and K. Rzazewski: *Applied Physics B* **63,** 499 (1996).
- R. Lichters, J. Meyer-ter-Vehn and A. Pukhov *Phys. Plasmas* **3**, 3425 (1996).

The moving mirror model

Ignoring retardation effects, the phase shift of the reflected wave resulting from a sinusoidal displacement of the reflecting surface in the z direction $s(t)=s_0 \sin w_m t$ is

$$
\phi(t) = (2\omega_0 s_0) \cos \theta \sin \omega_m t
$$

 θ is the angle of incidence, ω_m is the modulation frequency. The electric field of the reflected wave is

$$
E_R \propto e^{-i\omega_0 t} e^{i\phi(t)} = e^{-i\omega_0 t} \sum_{n=-\infty}^{n=\infty} J_n(\chi) e^{-in\omega_m t}
$$

 $J_n(\chi)$ is the Bessel function of order n and $\chi = (2\omega_0 s_0/c)\cos\theta$

where we've made use of the Jacobi expansion \sum $=\infty$ *n*=−∞ $e^{-i z \sin \omega_m t} = \sum J(\chi) e^{-i \pi \chi}$ *n* $e^{-iz\sin\omega_m t} = \sum J(\chi)e^{-in\omega_m t}$

Shot 270502

Conversion efficiency dependence on intensity on target

Energy conversion efficiencies of 10-6 in X-UV harmonics up to the 68th order were measured and scaled as $E(\omega) = E_L(\omega/\omega_0)^{-q}$

Good agreement was found with particle in cell simulations

P.A.Norreys *et al.* Physical Review Letters **76**, 1832 (1996)

Complex dynamics of n_c motion

For $I\lambda^2$ ⁼ 1.2×10¹⁹ Wcm⁻²µm² and $n/n_c = 3.2$ only ω_0 oscillations are observed. The power spectrum extends to ~35 and still shows modulations. Over dense regions of plasma are observed at the peak of the pulse.

 $I\lambda^{2} = 1.2 \times 10^{19}$ Wcm⁻² μ m² and $n/n_c = 6.0$, both ω_0 and $2\omega_0$ oscillations appear.

Additional order results in the modulation of the harmonic spectrum

I.Watts *et al*., Physical Review Letters **88**, 155001 (2002).

- Harmonic efficiency changes with increasing density scalelength and mirrors the change in the absorption process
- As the density scale-length increases, the absorption process changes from Bruneltype to classical resonance absorption.

Maximum absorption at L/λ $~10.2$

M. Zepf et al., Physical Review E 58, R5253 (1998).

Intense laser-plasma interaction physics

High harmonic generation – towards attosecond interactions

B.Dromey, M.Zepf….P.A.Norreys *Nature Physics* **2** 546 (2006)

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Ponderomotive force

Neglecting electron pressure, the force equation on the electron fluid is

$$
\frac{\partial u_e}{\partial t} + \underline{u}_e \cdot \nabla \underline{u}_e = -\frac{e}{m} E(x) \sin \omega t
$$

To a first approximation $u_e = \underline{u}^h$ where $\underline{u}^h = \frac{eL}{m} \cos \omega t$ *m eE* $\underline{u}^h = \frac{eE}{me} \cos \omega h$ ω $=\frac{eE}{m}$ cos

$$
\frac{\partial u^h}{\partial t} - \frac{e}{m} E(x) \sin \omega t
$$

Averaging the force equation over the rapid oscillations of the electric field,

$$
m\frac{\partial u^s}{\partial t} = -e\underline{E}^s - m < \underline{u}^h \cdot \nabla \underline{u}^h > t
$$

<> denotes the average of the high frequency oscillation, $\underline{u}^{\circ}=_{t}$ $\underline{u}^s = < u_e >_t \underline{E}^s = < \underline{E} >_t$

$$
m\frac{\partial \underline{u}^s}{\partial t} = -e\underline{E}^s - \frac{1}{4}\frac{e^2}{m\omega^2}\nabla \underline{E}^2(x)
$$

Ponderomotive force

- An electron is accelerated from its initial position x_0 to the left and stops at position x_1 .
- It is accelerated to the right until its has passed its initial position x_0 .
- From that moment, it is decelerated by the reversed electric field and stops at x_2
- If x_0' is the position at which the electric field is reversed, then the deceleration interval x_2 - x_0 ' is larger than that of the acceleration since the E field is weaker and a longer distance is needed to take away the energy gained in the former $\frac{1}{4}$ period.

• On its way back, the electron is stopped at position x_3

The result is a drift in the direction of decreasing wave amplitude.

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Summary of laser-plasma instabilities

Stimulated Brillouin Scatter - SBS. Scattered light is downshifted by $\omega_{i} \sim \omega_{I}$

Source of large losses for ICF experiments

Very dangerous for laser systems

Stimulated Raman Scatter – SRS. Scattered light downshifted by $\omega_{\rm p}$ Produces large amplitude plasma waves Hot electrons (Landau damping) Preheating of fuel

Two Plasmon Decay - TPD. Decays into two plasma waves Not as serious Localised near n_{cr} / 4.

Filamentation Can cause beam non-uniformities problematic for direct drive ICF

Regimes of applicability

$$
\omega \rightarrow \omega_{sc} + \omega_{ia}
$$

TPD

$$
\omega \to \omega_p + \omega_p
$$

Parametric instability growth

Large amplitude light wave couples to a density fluctuation E_{L} δn *Ep* $E_{L}E_{p}$ *n* Enhances the density perturbation

> The electric fields combine to generate a fluctuation in the field pressure

Drives an electric field associated with an electron plasma wave

 The threshold for the instabilities is determined by spatial inhomogeneities and damping of plasma waves – this reduces the regions where the waves can resonantly interact

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Stimulated Raman Scatter

$$
\omega_0 = \omega_s + \omega_{pe}
$$

$$
\underline{k}_0 = \underline{k}_s + k_{pe}
$$

The instability requires

$$
\omega_0 \geq 2\omega_{pe}
$$

i.e
$$
n \leq \frac{n_{cr}}{4}
$$

 \geq 10 % in ICF plasmas. The instability causes heating of the plasma, due to damping

Can have high phase velocity – produces high energy electrons

Stimulated Raman scattering

General instability analysis

Derive an expression that relates the creation of the scattered EM waves (by coupling the density perturbations generated by the laser field) to the transverse current that produces the scattered light wave.

Start from Ampere's law
$$
\nabla \times \underline{B} = 4\pi \underline{J} + \frac{1}{c} \frac{\partial \underline{E}}{\partial t}
$$
 (1)

Remember that $\nabla \cdot A = 0$ $B = \nabla \times A$

$$
\underline{E} = \frac{1}{c} \frac{\partial \underline{A}}{\partial t} - \nabla \phi \qquad \qquad \nabla \times (\nabla \times \underline{A}) = \nabla (\nabla \cdot \underline{A}) - \nabla^2 \underline{A}
$$

$$
\Rightarrow \left(\frac{1}{c^2}\frac{\partial^2}{\partial t^2} - \nabla^2\right) A = \frac{4\pi}{c} \, \underline{j} - \frac{1}{c} \frac{\partial}{\partial t} \nabla \phi \tag{2}
$$

General instability analysis continued…1

t l Separate J into a EM part j_t and a longitundinal part j_l $j = j + j$ (3)

Poisson's equation $\nabla^2 \phi = -4\pi \rho$ (4) ρ is the charge density Conservation of charge $\frac{op}{\gamma} + \nabla . j = 0$ ∂ ∂ *j t* $\frac{\rho}{\gamma} + \nabla.$ (5)

Taking $\partial/\partial t$ of (4) and substituting for $\partial \rho/\partial t$ from (5)

$$
\nabla \cdot \left(\frac{\partial}{\partial t} \nabla \phi - 4\pi \underline{j}\right) = 0 \tag{6}
$$

$$
\therefore \frac{\partial}{\partial t} \nabla \phi = 4\pi \underline{j}_{l} \quad \text{because} \quad \nabla \cdot \underline{j}_{t} = 0 \quad (7)
$$

Equation (2) becomes
$$
\left(\frac{1}{c^2} \frac{\partial^2}{\partial t^2} - \nabla^2 \right) \underline{A} = \frac{4\pi}{c} \underline{j}_{t} \quad (8)
$$

If we restrict the problem to A *.* ∇n_{e} = 0, the transverse current becomes j_{t} = - n_{e} e u_{t} where u_t is $v_{\rm osc}$. For $u_t < c$, $u_t = eA/mc$ because

$$
\frac{\partial \underline{u}_t}{\partial t} = -\frac{e}{m} \underline{E}_t = \frac{e}{mc} \frac{\partial \underline{A}}{\partial t}
$$
(9)

This gives eqn for propagation of a light wave in a plasma

$$
\left(\frac{\partial^2}{\partial t^2} - c^2 \nabla^2\right) A = -\frac{4\pi e^2}{m} n_e A \tag{10}
$$

Substitute for $\underline{A} = \underline{A}_L + \underline{A}_1$ and $n_e = n_0 + n_1$

$$
\left(\frac{\partial^2}{\partial t^2} - c^2 \nabla^2 + \omega_{pe}^2\right) A_1 = -\frac{4\pi e^2}{m} n_1 A_L \tag{11}
$$

The right hand side is simply the transverse current ($\propto n_{1}v_{L}$) that produces the light wave.

Now we need an equation for the density perturbation

Continuity equation
$$
\frac{\partial n_e}{\partial t} + \nabla \cdot (n_e \underline{u}_e) = 0
$$
 (12)
Force equation
$$
\frac{\partial \underline{u}_e}{\partial t} + \underline{u}_e \cdot \nabla \underline{u}_e = -\frac{e}{m} \left(E + \frac{\underline{u}_e \times B}{c} \right) - \frac{\nabla p_e}{n_e m}
$$
 (13)

Here the ions are fixed and provide a neutralising background

General instability analysis continued…2

$$
\underline{u}_e = \underline{u}_L + \frac{e\underline{A}}{mc} \qquad \text{transverse}
$$
\n
$$
\uparrow
$$
\n
$$
\frac{\partial}{\partial t} \left(\underline{u}_L + \frac{e\underline{A}}{mc} \right) + \left(\underline{u}_L + \frac{e\underline{A}}{mc} \right) \nabla \cdot \left(\underline{u}_L + \frac{e\underline{A}}{mc} \right) = -\frac{\nabla p_e}{n_e m} - \frac{e}{m} \left(\underline{E} + \frac{1}{c} \left(\underline{u}_L + \frac{e\underline{A}}{mc} \right) \times B \right)
$$
\n(14)

substituting

$$
\underline{B} = \nabla \times \underline{A} \qquad \underline{E} = \frac{1}{c} \frac{\partial \underline{A}}{\partial t} - \nabla \phi \qquad (\underline{A} \cdot \nabla) \underline{A} + \underline{A} \times (\nabla \times \underline{A}) = \frac{1}{2} \nabla (\underline{A})^2
$$
\ngives\n
$$
\frac{\partial \underline{u}_L}{\partial t} = \frac{e}{m} \nabla \phi - \frac{1}{2} \nabla \left(\underline{u}_L + \frac{e \underline{A}}{mc} \right)^2 - \frac{\nabla p_e}{n_e m} \qquad (15)
$$
\n
$$
\underline{E} = \frac{1}{c} \frac{\partial \underline{A}}{\partial t} - \nabla \phi \qquad (15)
$$
\n
$$
\underline{u}_L + \frac{e \underline{A}}{mc} \qquad \frac{\partial \underline{v}_L}{\partial t} - \frac{\nabla p_e}{mc} \qquad (16)
$$
\n
$$
\underline{u}_L + \frac{e \underline{A}}{mc} \qquad \frac{\partial \underline{v}_L}{\partial t} - \frac{\nabla \phi}{mc} \qquad (17)
$$
\n
$$
\underline{u}_L + \frac{e \underline{A}}{mc} \qquad \frac{\partial \underline{v}_L}{\partial t} - \frac{\nabla \phi}{mc} \qquad (18)
$$

General instability analysis continued…3

Use adiabatic equation of state p_e/n_e^3 = constant (16)

Linearise (12) and (15). Do this by taking $u_L = u$, $n_e = n_o + n_1$, $\underline{A} = \underline{A}_L + \underline{A}_1$

$$
\nabla \left(\frac{p_e}{n_e^3} \right) = 0
$$
\n
$$
\nabla p_e = \frac{3p_e}{n_e} \nabla n_1
$$
\n
$$
\frac{\nabla p_e}{n_e m_e} = \frac{3v_{th}^2}{n_0} \nabla n_1
$$
\n(17)

$$
\frac{\partial \underline{u}}{\partial t} = \frac{e}{m} \nabla \phi - \frac{1}{2} \nabla \left(\underline{u}_1 + \frac{e}{mc} \left(\underline{A}_L + \underline{A}_1 \right) \right)^2 - \frac{3v_{th}^2}{n_0} \nabla n_1 \tag{18}
$$

General instability analysis finalised

To 1st order ignore A_1^2 , u_1^2 and $u_1 \cdot \underline{A} = 0$

$$
\frac{\partial u}{\partial t} = \frac{e}{m} \nabla \phi - \frac{e^2}{m^2 c^2} \nabla (\underline{A}_L \cdot \underline{A}_1) - \frac{3v_{th}^2}{n_0} \nabla n_1 \tag{19}
$$

$$
\frac{\partial n_1}{\partial t} + n_0 \nabla \cdot \underline{u}_1 = 0 \tag{20}
$$

Combining by taking time derivative of (20), a divergence of (19) gives

$$
\left(\frac{\partial^2}{\partial t^2} + \omega_{pe}^2 - 3\upsilon_e^2 \nabla^2\right) n_1 = \frac{n_0 e^2}{m^2 c^2} \nabla^2 \left(\underline{A}_L \cdot \underline{A}_1\right) \tag{21}
$$

Where $\nabla^2 \phi = 4\pi n_1 e$

This is now the equation for the density perturbations generated by fluctuations in the intensity of the EM wave

Dispersion relationship for SRS

Start with $\underline{A}_L = \underline{A}_0 \cos(k_0 \cdot \underline{x} - \omega_0 t)$

Fourier analyse

$$
\left(\omega^2 - k^2 c^2 - \omega_{pe}^2\right) A_1(k,\omega) = \frac{4\pi e^2}{2m} A_0 \left[n_1(k - k_0, \omega - \omega_0) + n_1(k + k_0, \omega + \omega_0)\right]
$$
(22)

$$
\left(\omega^2 - \omega_{ek}^2\right) n_1(\underline{k}, \omega) = \frac{k^2 e^2 n_0}{2m^2 c^2} \underline{A}_0 \cdot \left[\underline{A}_1(k - k_0, \omega - \omega_0) + \underline{A}_1(k + k_0, \omega + \omega_0)\right]
$$
(23)

Where $\omega_{ek} = (\omega_{pe}^2 + 3k^2v_e^2)^{1/2}$ is the Bohm-Gross frequency and ω_0 and k_0 are the frequency and wave number of the laser wave. Also

$$
\underline{A}_1 = A_1 \exp(ikx - i\omega t)
$$

\n
$$
n_1 = n_1 \exp(ikx - i\omega t)
$$

\n
$$
\underline{A}_0 = A_0 \cos(k_0x - w_0t)
$$

\n
$$
\underline{A}_0 = A_0 [\exp(i(k_0x - w_0t) + \exp(-i(k_0x - w_0t))]
$$

(24)

 $\overline{}$

Growth rates

Use (19) to eliminate A_1

$$
\left(\omega^2 - \omega_{ek}^2\right) = \frac{\omega_p^2 k^2 v_{os}^2}{4} \cdot \left[\frac{1}{D(\omega - \omega_0, k - k_0)} + \frac{1}{D(\omega + \omega_0, k + k_0)}\right] \tag{25}
$$

 $D(\omega, k) = \omega^2 - k^2 c^2 - \omega_{pe}^2$ *v*_{os} is the oscillatory velocity (*eA*₀/*mc*)

Note: in deriving (25), it was assumed that $\omega \sim \omega_{pe}$ and terms n_1 (*k-2k₀,* ω *-2* ω_0 *)* and n_1 (k+2k $_0$, ω +2 ω_0) were neglected as non-resonant.

From this dispersion relation, the growth rates can be found. Neglecting the upshifted light wave as non-resonant gives

$$
\left(\omega^2 - \omega_{ek}^2\right) \left[(\omega - \omega_0)^2 - \left(\underline{k} - \underline{k}_0\right)^2 c^2 - \omega_{pe}^2 \right] = \frac{\omega_{pe}^2 k^2 \omega_{os}^2}{4} \tag{26}
$$

Maximum growth rate

Take $\omega_{ek} + \delta \omega$ where $\delta \omega \ll \omega_{ek}$. Note that maximum growth rate occurs when the scattered light is resonant, i.e. when

$$
(\omega_{ek} - \omega_0)^2 - (\underline{k} - \underline{k}_0)^2 c^2 - \omega_{pe}^2 = 0
$$
 (27)

Then take
$$
\delta \omega = i\gamma
$$
 where $\gamma = \frac{k v_{os}}{4} \left[\frac{\omega_{pe}^2}{\omega_{ek} (\omega_0 - \omega_{ek})} \right]^{1/2}$ (28)

The wave number *k* is given by (27). For backscattered light the growth rate maximises for

$$
k = k_0 + \frac{\omega_0}{c} \left(1 - \frac{2\omega_{pe}}{\omega_0} \right)^{1/2}
$$
 (29)

The wave numbers start at $k = 2k_0$ for $n \ll n_c/4$ and goes to $k = k_0$ for $n \sim n_{cr}/4$

Growth rate in low density plasmas

For forward scattered light at low density *k<<⁰ /c,* both the up-shifted and down-shifted light waves can be nearly resonant

$$
D(\omega \pm \omega_0, \underline{k} \pm \underline{k}_0) \approx 2(\omega_{pe} \pm \omega_0) \delta \omega \tag{30}
$$

Substitute into (25), the maximum growth rate $\delta \omega = i\gamma$

$$
\gamma \approx \frac{\omega_{pe}^2}{2\sqrt{2}\omega_0} \frac{v_{os}}{c}
$$
 (31)

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Stimulated Brillouin Scatter

Stimulated Brillouin Scattering (SBS)

- density fluctuations associated with high frequency ion acoustic waves causes scattering of light
- •> 60 % has been observed in experiment

 $\omega_{0} \rightarrow \omega_{s} + \omega_{ia}$ $\underline{k}_0 \rightarrow \underline{k}_s + \underline{k}_{ia}$

• Similar analysis to SRS, except n_1 is density fluctuation associated with an ion acoustic wave. We write

$$
\left(\frac{\partial^2}{\partial t^2} - c^2 \nabla^2 + \omega_{pe}^2\right) A_1 = -\frac{4\pi e^2}{m} n_1 A_L \tag{32}
$$

• Force equation gives $\frac{U h_1}{\Delta t^2} - c_s^2 \nabla n_1 = \frac{\Sigma h_0 \varepsilon}{\Delta t^2} \nabla^2 (\underline{A}_L \cdot \underline{A}_1)$ 2 2 0 1 2 2 1 2 $A_{\iota}\cdot A$ *mMc* Zn_0e $c_s^2 \nabla n$ *t n* $-c_s^2 \nabla n_1 = \frac{\Sigma n_0 c}{m \Lambda L^2} \nabla^2 (\underline{A}_L \cdot$ ∂ ∂ $c_{s} = (ZkT_{e}/M)^{1/2}$ ion acoustic velocity (33)

Growth rate

Fourier analyse (32) and (33)

$$
\omega = kc_s + i\gamma
$$

$$
\gamma << kc_s
$$

gives

$$
\underline{k} = 2\underline{k}_0 - \frac{2\omega_0 c_c}{c^2} \tag{34}
$$

$$
\gamma = \frac{1}{2\sqrt{2}} \frac{k_0 v_{os} \omega_{pi}}{(\omega_0 k_0 c_s)^{1/2}}
$$
(35)

for backscattered light

Holhraum irradiation uniformity by SBS

S. Glenzer et al., Science **327**, 1228 (2010) P. Michel et al., Phys. Rev. Lett. **102**, 025004 (2009)

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• Occurs in a narrow density region at n_{cr} / 4

$$
\omega \to \omega_{pe} + \omega_{pe}
$$

$$
\underline{k} \to \underline{k}_1 + \underline{k}_2
$$

• Similar analysis as SRS except the equations describe the coupling of the electron plasma waves with wave number \underline{k} and $\underline{k} - \underline{k}_0$ with the large light wave

- For k $>>$ k₀, $\gamma = \frac{\kappa_0 U_{os}}{4}$ for plasma waves at 45° to both <u>k</u> and v_{os} 4 $\gamma = \frac{k_0 v_{os}}{4}$
- Threshold is much less than SRS at n_{cr} / 4 unless the plasma is hot

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Filamentation instability

- Same dispersion relation as SBS also describes filamentation
- Zero frequency density perturbation corresponding to modulations of the light intensity
- Occurs in a plane orthogonal to propagation vector of light wave $k \cdot k_o = 0$
 $\omega = i\gamma << \omega_0$

$$
\omega = i\gamma \ll \omega_0
$$

$$
\gamma = \frac{1}{8} \left(\frac{v_{os}}{v_t} \right)^2 \frac{\omega_p^2}{\omega_0}
$$

- Filamentation can also be caused by thermal or relativistic effects:
	- 1. enhanced temperature changes refractive index locally
	- 2. relativistic electron mass increase has same effect as reducing the density

$$
P_{critical} = 20 \frac{\omega^2}{\omega_{pe}^2} \text{GW} \quad \text{for self-focusing}
$$

Efficient Raman amplification into the petawatt regime

R. Trines , F. Fiuza, R. Bingham, R.A. Fonseca, L.O.Silva, R.A. Cairns and P.A. Norreys

Nature Physics **7**, 87 (2011)

•A long laser pulse (pump) in plasma will spontaneously scatter off Langmuir waves: Raman scattering

Stimulate this scattering by sending in a short, counter propagating pulse at the frequency of the scattered light (probe pulse)

Because scattering happens mainly at the location of the probe, most of the energy of the long pump will go into the short probe: efficient pulse compression

Miniature pulse compressor

Solid state compressor (Vulcan)

Image: STFC Media Services

Volume of a plasmabased compressor

Simulations versus theory

RBS growth increases with pump amplitude and plasma density, but so do pump RFS and probe filamentation

Optimal simulation regime corresponds to at most 10 e-foldings for pump RFS and probe filamentation

A bad result

For a 2*10¹⁵ W/cm² pump and $\omega_0/\omega_p = 10$, the probe is still amplified, but also destroyed by filamentation

A good result

For a 2*10¹⁵ W/cm² pump and ω_0/ω_p = 20, the probe is amplified to 8*10¹⁷ W/cm² after 4 mm of propagation, with limited filamentation

 $10 \text{ TW} \rightarrow 2 \text{ PW}$ and transversely extensible!