

# The SST-1M prototype camera for the Cherenkov Telescope Array

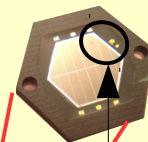
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## The Hamamatsu large area hexagonal sensor

**Challenge #1:** we use hexagonal hollow light concentrators → we custom designed a large area (93.56 mm<sup>2</sup>) hexagonal sensor

**Challenge #2:** the sensor has a high capacitance (3.4 nF) → complex fine-tuning of the preamp

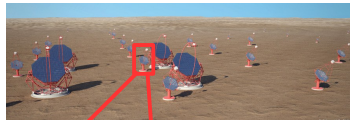
**Challenge #3:** we have to stabilize the operational point of the sensor → we implemented a compensation logic in the front end



NTC temperature sensor

## The Cherenkov Telescope Array (CTA)

1400 scientists (28 countries) propose to build an array of more than 100 telescopes deployed in two sites (North and South hemisphere) to study gamma-rays with 10 GeV < E < 300 TeV.

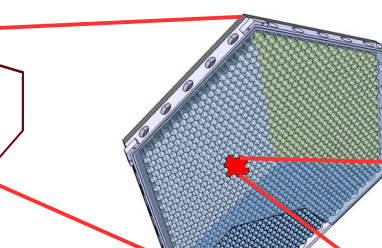


## The single mirror Small Size Telescopes (SST-1M)

Among the three types of telescopes (Large, Medium and Small size – LST, MST and SST), the SSTs are meant for covering the 5-300 TeV energy region.

## The SST-1M camera

Designed to be very compact and light (<200 kg). The photo-detection plane and the readout are physically separated.

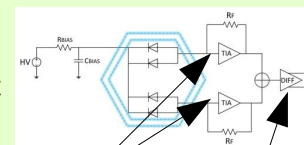
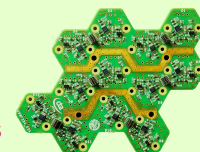


## The photo-detection plane

A hexagonal matrix of 1296 Silicon photomultipliers (SiPMs) coupled to hollow light concentrators, arranged on 108 modules (12 pixels each).

## The preAmp Board

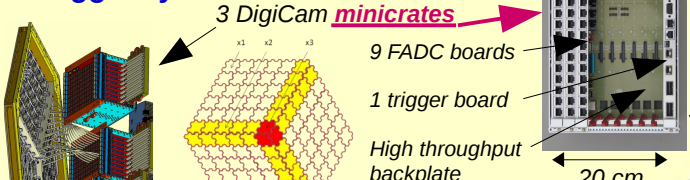
The PreAmp Board reads out the **12 pixels** of a module. The **4 channels** of each sensor are biased with a single high voltage and their signal is re-summed to provide 1 readout. The sensor is **DC coupled** to the front-end, to estimate the background light contribution from the baseline level. The board has been designed with the aim of minimizing **power consumption**.



Texas Instruments OPA846 low-noise trans-impedance

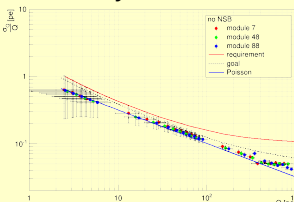
Differential output LMH6551

## DigiCam: a fully digital readout and trigger system



The photo-detection plane is divided in **3 logical sectors**. Each minicrate controls one sector.

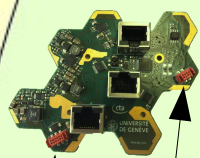
**The FADC board** digitizes the signals at 250 MS/s on 48 channels (4 modules). The samples are sent through high speed multi-gigabit serial links to the Xilinx XC7VX415T **FPGA**, where they are pre-processed and stored in **ring buffers**, so that the system is **dead-time free**.



Although we use a non-linear input stage, the charge resolution of the full system complies with the CTA requirements.

**The trigger board** receives data from the FADC boards, resampled in sets of 3 pixels. A highly parallelized trigger algorithm implemented in the Xilinx XC7VX485T **FPGA** is applied, based on the recognition of geometrical patterns on pixels above threshold. Selected events are sent to the central system via a 10Gb fiber link.

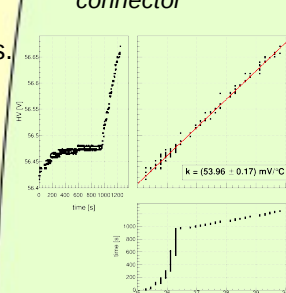
## The Slow Control Board



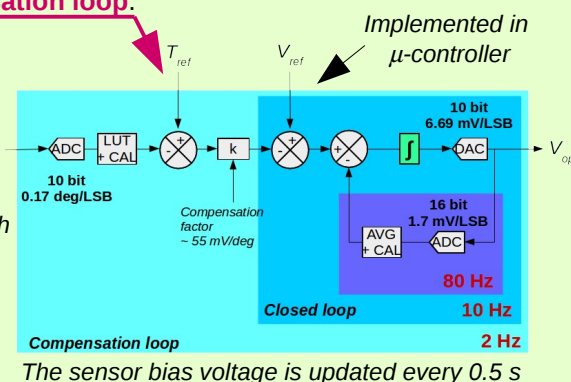
CANbus connector  
Power connector

The Slow Control Board functions are:

- Route the **analog signals** from the photo-detection plane to DigiCam via 3 RJ45 connectors per module;
- Read/Set the **bias voltage** of each sensor individually;
- Read the **temperature** of each sensor individually;
- Stabilize the operational point of each sensor via the **compensation loop**.



The bias voltage of each of the 1296 pixels is adjusted as a function of temperature



The sensor bias voltage is updated every 0.5 s

## Acknowledgements

We gratefully acknowledge support from the agencies and organizations listed under Funding Agencies at this website: <http://www.cta-observatory.org/>. In particular we are grateful for support from the CN grant DEC-2011/01/M/ST9/01891 380 and the MNiSW grant 498/L/FNITP/FNITP/2010 in Poland and the University of Geneva and the Swiss National Foundation.

## References

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- J. A. Aguilar et al. for the SST-1M sub-Consortium, "Design, optimization and characterization of the light concentrators of the single-mirror small size telescopes of the Cherenkov Telescope Array.", Astroparticle Physics, Vol. 60, pp. 32-40, 2015.