Fluid dynamic perturbations in heavy ion collisions

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mainly based on

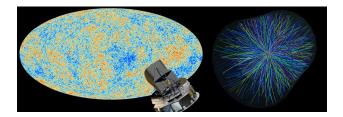
- Mode-by-mode fluid dynamics for relativistic heavy ion collisions [Phys. Lett. B, 728, 407 (2014), with U. A. Wiedemann]
- Statistics of initial density perturbations in heavy ion collisions and their fluid dynamic response [JHEP 1408 (2014) 005, with U. A. Wiedemann]
- Fluctuations of baryonic number around Bjorken background [work in progress, with M. Martinez]

Introduction

What perturbations are interesting and why?

- Initial fluid perturbations: Event-by-event fluctuations around a background or average of fluid fields at time τ_0 :
 - energy density ϵ
 - ${\scriptstyle \bullet}\,$ fluid velocity u^{μ}
 - shear stress $\pi^{\mu\nu}$
 - more general also: baryon number density *n*, electric charge density, electromagnetic fields, ...
- governed by universal evolution equations
- can be used to constrain thermodynamic and transport properties
- contain interesting information from early times

Similarities to cosmological fluctuation analysis



- fluctuation spectrum contains info from early times
- many numbers can be measured and compared to theory
- can lead to detailed understanding of evolution
- to learn something about the evolution one needs to know some universal properties of initial state, for example $P(k) \sim k^{n_s-1}$

A program to understand fluid perturbations

- Oharacterize initial perturbations
- Propagated them through fluid dynamic regime
- Oetermine influence on particle spectra and harmonic flow coefficients
- Take also perturbations from non-hydro sources (jets) into account [see work with K. Zapp, EPJC 74 (2014) 12, 3189]

Characterization of initial conditions

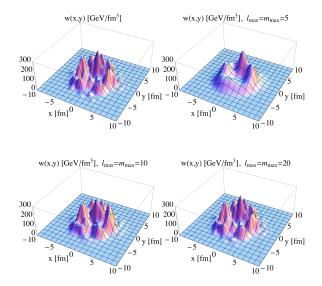
Transverse enthalpy density

Based on Bessel-Fourier expansion and background density [Floerchinger & Wiedemann 2013, see also Coleman-Smith, Petersen & Wolpert 2012, Floerchinger & Wiedemann 2014]

$$w(r,\phi) = w_{\mathsf{BG}}(r) + w_{\mathsf{BG}}(r) \sum_{m,l} w_l^{(m)} e^{im\phi} J_m\left(z_l^{(m)}\rho(r)\right)$$

- azimuthal wavenumber m, radial wavenumber l
- $w_l^{(m)}$ dimensionless
- \bullet higher m and l correspond to finer spatial resolution
- coefficients $w_l^{(m)}$ can be related to eccentricienies
- works similar for vectors (velocity) and tensors (shear stress)

Transverse density from Glauber model



Event ensembles

• Initial conditions at beginning of fluid dynamic regime are governed by event-by-event probability distribution

 $p_{\tau_0}[w, u^\mu, \pi^{\mu\nu}, \ldots]$

• Moments / correlation functions

$$\left\langle w_{l_1}^{(m_1)} w_{l_2}^{(m_2)} \dots w_{l_n}^{(m_n)} \right\rangle$$

contain information from initial state physics / early dynamics
universal (model independent) properties would be nice to have
same information in cumulants (connected correlation functions)

Statistics of initial density perturbations

Independent point-sources model (IPSM)

$$w(\vec{x}) = \left[\frac{1}{\tau_0} \frac{dW_{\mathsf{BG}}}{d\eta}\right] \frac{1}{N} \sum_{j=1}^N \delta^{(2)}(\vec{x} - \vec{x}_j)$$

- random positions \vec{x}_j , independent and identically distributed
- probability distribution $p(\vec{x}_j)$ reflects collision geometry
- possible to determine correlation functions analytically for *central* and *non-central* collisions [Floerchinger & Wiedemann (2014)]
- Long-wavelength modes (small *m* and *l*) that don't resolve differences between point-like and extended sources have *universal statistics*.

Solution of IPSM and scaling with number of sources

• for IPSM one can exactly determine the correlation functions

 $\langle w_{l_1}^{(m_1)} \cdots w_{l_n}^{(m_n)} \rangle$

[Floerchinger & Wiedemann, JHEP 1408 (2014) 005]

• connected correlation functions (cumulants) scale with N like [see also Ollitrault & Yan (2014), Bzdak & Skokov (2014)]

$$\langle w_{l_1}^{(m_1)} \cdots w_{l_n}^{(m_n)} \rangle_c \sim \frac{1}{N^{n-1}}$$

(implies that distribution is non-Gaussian)

- scaling broken for non-central collisions
- impact parameter dependence of terms that break scaling is known

Fluid dynamic response

Response to density perturbations

For a single event

$$V_m^* = v_m e^{-i m \psi_m}$$

= $\sum_l S_{(m)l} w_l^{(m)} + \sum_{\substack{m_1, m_2, \ l_1, l_2}} S_{(m_1, m_2)l_1, l_2} w_{l_1}^{(m_1)} w_{l_2}^{(m_2)} \delta_{m, m_1 + m_2} + \dots$

- $S_{(m)l}$ is linear dynamic response function
- $S_{(m_1,m_2)l_1,l_2}$ is quadratic dynamic response function etc.
- Symmetries imply conservation of azimuthal wavenumber
- Response functions depend on thermodynamic and transport properties, in particular viscosity.

Flow correlations from initial density correlations Moments of flow coefficients

$$\left\langle V_{m_1}^* \cdots V_{m_n}^* \right\rangle = S_{(m_1)l_1} \cdots S_{(m_n)l_n} \left\langle w_{l_1}^{(m_1)} \cdots w_{l_n}^{(m_n)} \right\rangle$$

+ non-linear terms

- combination of dynamical response coefficients and correlation functions of initial density perturbations
- linear, quadratic and higher-order terms
- $\bullet~{\rm For}~N$ independent sources and central collisions

$$v_m \{n\}^n \sim \frac{1}{N^{n-1}}$$
 or $v_m \{n\} \sim \frac{1}{N^{1-\frac{1}{n}}}$

- explains why $v_m\{4\} \sim v_m\{6\} \sim v_m\{8\}$ are of almost equal size
- holds also for extended sources
- holds also for non-linear response contributions
- can be extended to other correlation functions, e. g. $\langle V_2 V_3 V_5^* \rangle \sim \frac{1}{N^2}$
- gets broken for non-central collisions
- impact parameter dependence of corrections is known

Scaling with system size

- Large (PbPb) and small systems (pPb) may have different number of independent sources N and response functions $S_{(m)l}$
- For linear dynamics one has parametrically

$$v_m\{n\} \sim \frac{S_{(m)l}}{N^{1-\frac{1}{n}}}$$

 $\bullet\,$ To have $v_m\{n\}|_{\mathsf{PbPb}}=v_m\{n\}|_{\mathsf{pPb}}$ one needs

$$\frac{S_{(m)l}|_{\mathsf{pPb}}}{S_{(m)l}|_{\mathsf{PbPb}}} = \left(\frac{N_{\mathsf{pPb}}}{N_{\mathsf{PbPb}}}\right)^{1-\frac{1}{n}}$$

• For comparison at equal multiplicity one may have $N_{\rm pPb}\approx N_{\rm PbPb}$ so that response functions must be equal

$$S_{(m)l}|_{\rm pPb} \approx S_{(m)l}|_{\rm PbPb}$$

• $S_{(m)l}$ depends on system size only via initial background $w_{BG}(r)$. Precise dependence can be investigated more closely.

$Hydrodynamic \ evolution$

$Perturbative \ expansion$

Write the hydrodynamic fields $h = (w, u^{\mu}, \pi^{\mu\nu}, \pi_{\text{Bulk}}, \ldots)$

 \bullet at initial time τ_0 as

 $h = h_0 + \epsilon h_1$

with background h_0 , fluctuation part ϵh_1

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• at later time \tau > \tau_0 as
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$$h = h_0 + \epsilon h_1 + \epsilon^2 h_2 + \epsilon^3 h_3 + \dots$$

Solve for time evolution in this scheme

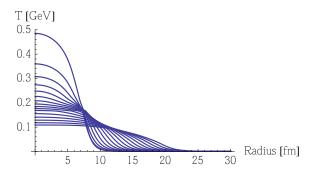
- h_0 is solution of full, non-linear hydro equations in symmetric situation: azimuthal rotation and Bjorken boost invariant
- h_1 is solution of linearized hydro equations around h_0 , can be solved mode-by-mode
- h_2 can be obtained by from interactions between modes etc.

Background evolution

System of coupled $1 + 1 \mbox{ dimensional non-linear partial differential equations for$

- enthalpy density $w(\tau,r)$ (or temperature $T(\tau,r)$)
- fluid velocity $u^\tau(\tau,r), u^r(\tau,r)$
- $\bullet\,$ two independent components of shear stress $\pi^{\mu\nu}(\tau,r)$

Can be easily solved numerically

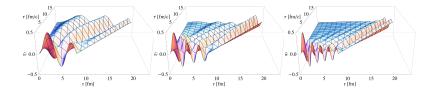


Evolving perturbation modes

- Linearized hydro equations: set of coupled 3 + 1 dimensional, linear, partial differential equations.
- Use Fourier expansion

$$h_j(\tau, r, \phi, \eta) = \sum_m \int \frac{dk_\eta}{2\pi} h_j^{(m)}(\tau, r, k_\eta) e^{i(m\phi + k_\eta \eta)}.$$

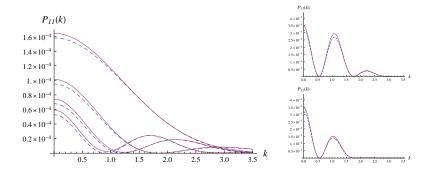
- Reduces to 1+1 dimensions.
- Can be solved numerically for each initial Bessel-Fourier mode.



Evolution of spectrum of density perturbations

Density-density spectrum

$$P_{11}(\vec{k}) = \int d^2x \, e^{-i\vec{k}(\vec{x}-\vec{y})} \, \langle \, d(\vec{x}_1) \, d(\vec{x}_2) \, \rangle_c$$

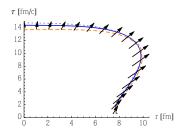


dashed: linear evolution, solid: including first non-linear correction left: $\eta/s = 0.08$, $\tau = 1.5, 2.5, 3.5, 4.5$ fm/c, right: $\eta/s = 0.08$ and $\eta/s = 0.8$, $\tau = 7.5$ fm/c [Brouzakis, Floerchinger, Tetradis & Wiedemann, arXiv:1411.2912]

Kinetic freeze-out

Freeze-out surface

- Perturbative expansion can be used also at freeze-out. [Floerchinger, Wiedemann 2013]
- Freeze-out surface is azimuthally symmetric as background.
- Generalization to kinetic hadronic scattering and decay phase possible.



(solid: $\eta/s = 0.08$, dotted: $\eta/s = 0$, dashed: $\eta/s = 0.3$)

Particle distribution

for single event

$$\ln\left(\frac{dN^{\text{single event}}}{p_T dp_T d\phi dy}\right) = \underbrace{\ln S_0(p_T)}_{\text{from background}} + \underbrace{\sum_{m,l} w_l^{(m)} e^{im\phi} \theta_l^{(m)}(p_T)}_{\text{from fluctuations}}$$

 $\bullet\,$ each mode comes with an angle, $w_l^{(m)} = |w_l^{(m)}|\,e^{im\psi_l^{(m)}}$

- each mode has different p_T -dependence, $\theta_l^{(m)}(p_T)$
- quadratic order correction

r

$$\sum_{n_1,m_2,l_1,l_2} w_{l_1}^{(m_1)} w_{l_2}^{(m_2)} e^{i(m_1+m_2)\phi} \kappa_{l_1,l_2}^{(m_1,m_2)}(p_T)$$

• non-linearities from hydro evolution and freeze-out

Differential harmonic flow coefficients

Double differential harmonic flow coefficient (to lowest order)

$$v_m \{2\}^2(p_T^a, p_T^b) = \sum_{l_1, l_2} \theta_{l_1}^{(m)}(p_T^a) \; \theta_{l_2}^{(m)}(p_T^b) \; \langle w_{l_1}^{(m)} w_{l_2}^{(m)*} \rangle$$

- intuitive matrix expression
- in general no factorization
- can be generalized to higher order flow cumulants

Baryon number density fluctuations

Fluctuations around vanishing baryon number

• Evolution of baryon number density

$$u^{\mu}\partial_{\mu}n + n\nabla_{\mu}u^{\mu} + \nabla_{\mu}\nu^{\mu} = 0$$

with diffusion current ν^{α} determined by heat conductivity κ

$$\nu^{\alpha} = -\kappa \left[\frac{nT}{\epsilon + p}\right]^2 \Delta^{\alpha\beta} \partial_{\beta} \left(\frac{\mu}{T}\right)$$

- Consider situation with $\langle n(x) \rangle = \langle \mu(x) \rangle = 0$ but event-by-event fluctuation $\delta n \neq 0$
- Concentrate now on Bjorken flow profile for u^{μ}

$$\partial_{\tau}\delta n + \frac{1}{\tau}\delta n - \kappa \left[\frac{nT}{\epsilon+p}\right]^2 \left(\frac{\partial(\mu/T)}{\partial n}\right)_{\epsilon} \left(\partial_x^2 + \partial_y^2 + \frac{1}{\tau^2}\partial_{\eta}^2\right)\delta n = 0$$

• Structures in transverse and rapidity directions are "flattened out" by heat conductive dissipation

Baryon number correlations experimentally

• Two-particle correlation function of baryons minus anti-baryons

$$C_{\mathsf{Baryon}}(\phi_1 - \phi_2, \eta_1 - \eta_2) = \langle n(\phi_1, \eta_1) n(\phi_2, \eta_2) \rangle_{c}$$

• In Fourier representation

$$C_{\mathsf{Baryon}}(\Delta\phi,\Delta\eta) = \sum_{m=-\infty}^{\infty} \int \frac{dq}{2\pi} \, \tilde{C}_{\mathsf{Baryon}}(m,q) \, e^{im\Delta\phi + iq\Delta\eta}$$

heat conductivity leads to exponential suppression

$$\tilde{C}_{\mathsf{Baryon}}(m,q) = e^{-m^2 I_1 - q^2 I_2} \left. \tilde{C}_{\mathsf{Baryon}}(m,q) \right|_{\kappa=0}$$

• I_1 and I_2 can be approximated as

$$I_1 \approx \int_{\tau_0}^{\tau_f} d\tau \, \frac{2}{R^2} \, \kappa \left[\frac{nT}{\epsilon + p} \right]^2 \left(\frac{\partial(\mu/T)}{\partial n} \right)_{\epsilon}$$
$$I_2 \approx \int_{\tau_0}^{\tau_f} d\tau \, \frac{2}{\tau^2} \, \kappa \left[\frac{nT}{\epsilon + p} \right]^2 \left(\frac{\partial(\mu/T)}{\partial n} \right)_{\epsilon}$$

• $I_2 \gg I_1$ would lead to long-range correlations in rapidity direction ("baryon number ridge")

Remarks on baryon number fluctuations

- Initial ("primordial") baryon number fluctuations are poorly understood so far but presumably non-vanishing.
- Heat conductivity of QCD also poorly understood theoretically so far
 - from perturbation theory [Danielewicz & Gyulassy, PRD 31, 53 (1985)]

$$\kappa \sim \frac{T^4}{\mu^2 \alpha_s^2 \ln \alpha_s} \qquad (\mu \ll T)$$

• from AdS/CFT [Son & Starinets, JHEP 0603 (2006)]

$$\kappa = 8\pi^2 \frac{T}{\mu^2} \eta = 2\pi \frac{sT}{\mu^2} \qquad (\mu \ll T)$$

- More refined study needed to take transverse expansion properly into account.
- Seems to be interesting topic for further experimental and theoretical studies.

Summary and Conclusions

Conclusions

- Systematic expansion in initial fluid perturbations is possible (good convergence properties) and very useful.
- Formalism works in praxis (see backup slides for results of "proof of principle" study).
- Initial density perturbations have some universal properties that can help to better constrain thermodynamic and transport properties.
- Fluid dynamic response allows to access correlation functions of initial perturbations.
- Baryon number correlations could allow to constrain heat conductivity.

Backup

Characterization of transverse density via eccentricities

Fluctuations in initial transverse enthalpy density $w(r, \phi)$ can be characterized in terms of eccentricities $\epsilon_{n,m}$ and angles $\psi_{n,m}$ [Ollitrault, Teaney, Yan, Luzum, and others]

$$\epsilon_{n,m} e^{im \psi_{n,m}} = \frac{\int dr \int_0^{2\pi} d\varphi r^{n+1} e^{im\varphi} w(r,\varphi)}{\int dr \int_0^{2\pi} d\varphi r^{n+1} w(r,\varphi)}$$

- $w(r,\phi)$ completely determined by set of all $\epsilon_{n,m}$ and $\psi_{n,m}$
- closely related method is based on cumulants [Teaney, Yan]
- no positive transverse density can be associated to small set of cumulants (beyond Gaussian order) such that higher order cumulants vanish
- generalization to velocity and shear fluctuations not known

Scaling tests

• Start with single enthalpy density mode (m=2,l=1) on top of background

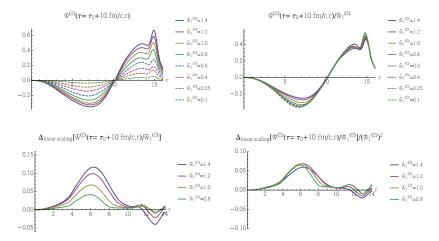
$$w(\tau_0, r, \phi) = w_{\mathsf{BG}}(\tau_0, r) \left[1 + 2 \,\tilde{w}_1^{(2)} J_2(k_1^{(2)} r) \, \cos(2\phi) \right].$$

- Evolve this with hydro solver ECHO-QGP [Del Zanna *et al.*, EPJC 73, 2524 (2013), see also following talk]
- Determine Fourier components

$$\tilde{w}^{(m)}(\tau, r) = \frac{1}{w_{\mathsf{BG}}(r)} \frac{1}{2\pi} \int d\phi \ e^{-im\phi} \ w(\tau, r, \phi)$$

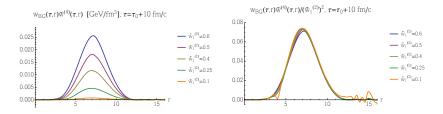
Scaling tests at first order

Compare enthalpy $\tilde{w}^{(2)}(\tau,r)$ at fixed τ for different initial weights



Scaling tests at second order

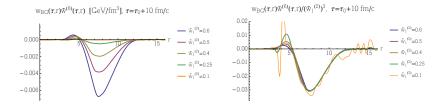
From symmetry considerations one expects that modes with m = 0 and m = 4 receive mainly quadratic contributions $\sim (\tilde{w}_1^{(2)})^2$.



- Hydrodynamic response to initial enthalpy density fluctuations is perturbative.
- Non-linearities can be understood order-by-order and lead to characteristic "overtones".
- Results motivate more thorough development of fluid dynamic perturbation theory.

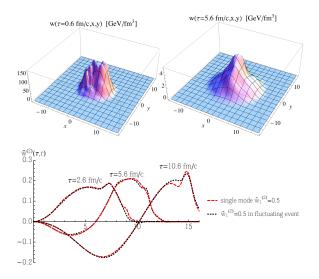
Scaling tests at third order

From symmetry considerations one expects that modes m = 6 receive mainly cubic contributions $\sim (\tilde{w}_1^{(2)})^3$.



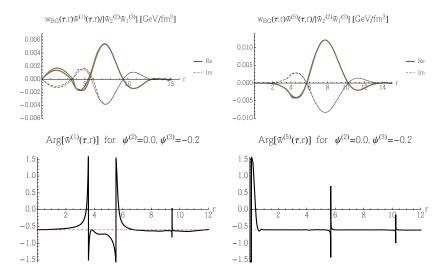
Scaling tests embedded in realistic event

Embed mode (m = 2, l = 1) into realistic fluctuating event and compare to embedding into pure background.



Scaling tests with several initial modes

Start with linear combination of (m = 2, l = 2) and (m = 3, l = 1) modes and test scaling for m = 1 and m = 5 response.



Generalized Glauber model

• Fluctuations due to nucleon positions: used so far

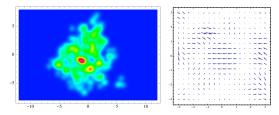
$$\epsilon(\tau, \mathbf{x}, y) = \sum_{i=1}^{N_{\text{part}}} \epsilon_w(\tau, \mathbf{x} - \mathbf{x}_i, y), \qquad u^{\mu} = (1, 0, 0, 0)$$

• can be generalized to include also velocity fluctuations

$$T^{\mu\nu}(\tau, \mathbf{x}, y) = \sum_{i=1}^{N_{\text{part}}} T^{\mu\nu}_w(\tau, \mathbf{x} - \mathbf{x}_i, y)$$

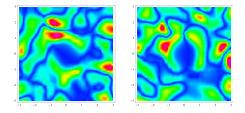
- More generally describe primordial fluid fields by
 - expectation values $\langle \epsilon(\tau_0, \mathbf{x}, y) \rangle, \langle u^{\mu}(\tau_0, \mathbf{x}, y) \rangle, \langle n_B(\tau_0, \mathbf{x}, y) \rangle$
 - correlation functions $\langle \epsilon(\tau_0, \mathbf{x}, y) \, \epsilon(\tau_0, \mathbf{x}', y') \rangle$, etc.
- Origin of this fluctuations is initial state physics and early-time, non-equilibrium dynamics.

Velocity fluctuations



- ullet also the velocity field will fluctuate at the initialization time τ_0
- $\bullet\,$ take here transverse velocity for every participant to be Gaussian distributed with width 0.1c

• vorticity
$$|\partial_1 u^2 - \partial_2 u^1|$$
 and divergence $|\partial_1 u^1 + \partial_2 u^2|$



"Proof of principle" study: One-particle spectrum Initial conditions from Glauber Monte Carlo Model

 $S(p_T) = dN/(2\pi p_T dp_T d\eta d\phi)$

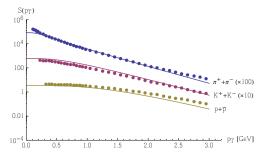


Figure: *

Points: 5% most central collisions, ALICE [PRL 109, 252301 (2012)] Curves: Our calculation, no hadron rescattering and decays after freeze-out.

Triangular flow for charged particles

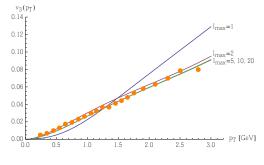


Figure: *

Points: 2% most central collisions, ALICE [PRL 107, 032301 (2011)] Curves: Different maximal resolution l_{max}

Elliptic flow for charged particles

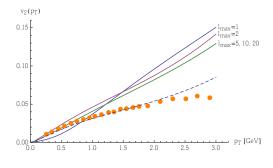


Figure: *

Points: 2% most central collisions, ALICE [PRL 107, 032301 (2011)] Solid curves: Different maximal resolution l_{max} Dashed curve: Mode (m = 2, l = 1) suppressed by factor 0.7

Flow coefficient v_4 for charged particles

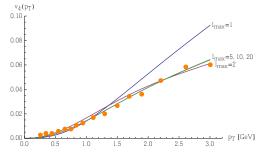


Figure: *

Points: 2% most central collisions, ALICE [PRL 107, 032301 (2011)] Curves: Different maximal resolution l_{max}

Flow coefficient v_5 for charged particles

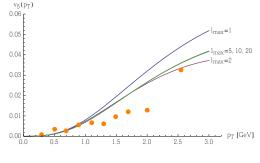


Figure: *

Points: 2% most central collisions, ALICE [PRL 107, 032301 (2011)] Curves: Different maximal resolution l_{max}

Harmonic flow coefficients, central, particle identified

