

Optimized evaluation approach for inflationary power spectra

Hayato Motohashi

Kavli Institute for Cosmological Physics
University of Chicago

HM and Wayne Hu, [arXiv:1503.04810](https://arxiv.org/abs/1503.04810)

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Planck TT, TE, EE + lowP

$$n_s = 0.9644 \pm 0.0049 \text{ (68\%CL)}$$

$$\alpha = -0.009 \pm 0.008 \text{ (68\%CL)}$$

$$r_{0.002} < 0.15 \text{ (95\%CL)}$$

$$|\alpha| \lesssim |n_s - 1| \\ = O(10^{-2})$$

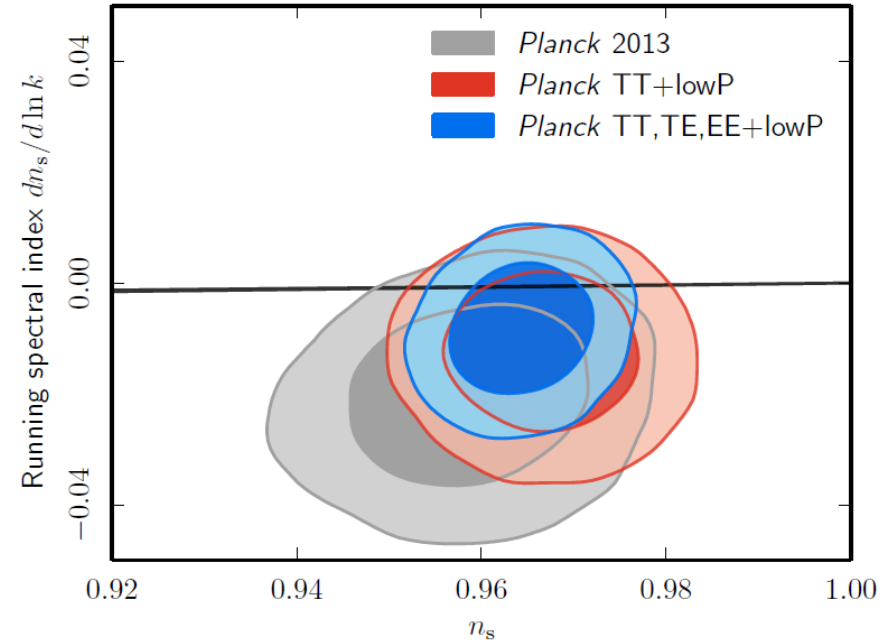


Fig. 5. Marginalized joint confidence contours for $(n_s, dn_s/d \ln k)$, at the 68% and 95% CL, in the presence of a non-zero tensor contribution, and using *Planck* TT+lowP or *Planck* TT,TE,EE+lowP. Constraints from the *Planck* 2013 data release are also shown for comparison. The thin black stripe shows the prediction of single-field monomial inflation models with $50 < N_* < 60$.

First order slow-roll approximation

Consistency relation

$$n_s - 1 = -6\epsilon_V + 2\eta_V$$

$$\alpha = 16\epsilon_V\eta_V - 24\epsilon_V^2 - 2\xi_V$$

$$r = 16\epsilon_V$$

Assumption

$$\epsilon_V, \eta_V = O(N^{-1})$$

$$\xi_V = O(N^{-2})$$

Suppose $|\alpha|, |n_s - 1| = O(10^{-2})$

→ $\epsilon_V, \eta_V, \xi_V = O(10^{-2})$

→ $n_s - 1$ needs ξ_V term

→ Inappropriate order counting



Generalized slow-roll (GSR) approximation

Stewart, astro-ph/0110322

Choe, Gong, Stewart, hep-ph/0405155

Dvorkin, Hu, 0910.2237

Generalized slow-roll

Solve the Mukhanov-Sasaki equation

$$\frac{d^2 y}{dx^2} + \left(1 - \frac{2}{x^2}\right) y = \frac{f'' - 3f'}{f} \frac{y}{x^2}$$

de Sitter

Deviation from dS

$$y = \sqrt{2k} a \delta\varphi, \quad x = k\eta, \quad f = 2\pi a \eta \sqrt{2\epsilon_H}, \quad ' = \frac{d}{d \ln \eta}$$

iteratively with Green function technique starting with the **de Sitter solution**

$$y_0 = \left(1 + \frac{i}{x}\right) e^{ix}$$

and using **deviations from de Sitter** as external source.

(cf. Born approximation)

Generalized slow-roll

First order of iteration

$$\ln \Delta_R^2(k) = - \int_0^\infty \frac{dx}{x} W'(x) G(\ln x)$$

Window function Source function

$$W(x) = \frac{3 \sin 2x}{2x^3} - \frac{3 \cos 2x}{x^2} - \frac{3 \sin 2x}{2x}$$

$$G(\ln x) = -2 \ln f + \frac{2}{3} (\ln f)'$$

$$\approx \ln \left(\frac{H^2}{8\pi^2 \epsilon_H} \right) - \frac{10}{3} \epsilon_H - \frac{2}{3} \delta_1$$

Optimized evaluation

HM, Hu, 1503.04810

First order of iteration

$$\ln \Delta_R^2(k) = - \int_0^\infty \frac{dx}{x} W'(x) G(\ln x)$$

Window function Source function

$$= G(\ln x_f) + \sum_{p=1}^{\infty} q_p(\ln x_f) G^{(p)}(\ln x_f)$$

Taylor expansion of G around evaluation point $\ln x_f$

$$q_p(\ln x_f) = - \frac{1}{p!} \int_0^\infty \frac{dx}{x} W'(x) \left(\ln \frac{x}{x_f} \right)^p$$

$$\lim_{p \rightarrow \infty} \frac{q_p}{q_{p-1}} = -\frac{1}{2} \rightarrow \text{Convergence criteria } \lim_{p \rightarrow \infty} \left| \frac{G^{(p)}}{G^{(p-1)}} \right| = \frac{1}{\Delta N} < 2$$

→ Low frequency oscillation

Optimized evaluation

HM, Hu, 1503.04810

$$\ln \Delta_R^2(k) = G(\ln x_f) + \sum_{p=1}^{\infty} q_p(\ln x_f) G^{(p)}(\ln x_f)$$

We can optimize the **evaluation point** $\ln x_f$ to minimize correction to a truncation.

$$x \equiv k\eta$$

Optimized evaluation

HM, Hu, 1503.04810

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$$x \equiv k\eta$$

First order slow-roll approximation (no optimization)

$$\ln \Delta_R^2 = G(\mathbf{0}) : \text{correction} = O(q_1(\mathbf{0})G'(\mathbf{0}))$$

$\ln x_f = 0$: horizon exit

$q_1(\mathbf{0}) \approx 1.06$

Optimized evaluation

HM, Hu, 1503.04810

$$\ln \Delta_R^2(k) = G(\ln x_f) + \sum_{p=1}^{\infty} q_p(\ln x_f) G^{(p)}(\ln x_f)$$

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$$x \equiv k\eta$$

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$$\ln \Delta_R^2 = G(\mathbf{0}) : \text{correction} = O(q_1(\mathbf{0})G'(\mathbf{0}))$$

$\ln x_f = \mathbf{0}$: horizon exit

$q_1(\mathbf{0}) \approx 1.06$

Optimized leading order approximation

$$\ln \Delta_R^2 = G(\ln x_1) : \text{correction} = O(q_2(\ln x_1)G''(\ln x_1))$$

$\ln x_f = \ln x_1 \approx 1.06$ with $q_1(\ln x_1) = 0$

$q_2(\ln x_1) \approx -0.36$

~ 1 efold before horizon exit

Optimized evaluation

HM, Hu, 1503.04810

$$\ln \Delta_R^2 = G(\ln x_f) + \sum_{p=1}^{\infty} q_p(\ln x_f) G^{(p)}(\ln x_f)$$

$$n_s - 1 = -G'(\ln x_f) - \sum_{p=1}^{\infty} q_p(\ln x_f) G^{(p+1)}(\ln x_f)$$

$$\alpha = G''(\ln x_f) + \sum_{p=1}^{\infty} q_p(\ln x_f) G^{(p+2)}(\ln x_f)$$

We can optimize the **evaluation point** $\ln x_f = \ln x_p$ where $q_p(\ln x_p) = 0$ and make a correction to a truncation at q_{p-1} as small as $O\left(q_{p+1}(\ln x_p) G^{(p+1)}(\ln x_p)\right)$.

Optimized evaluation

HM, Hu, 1503.04810

Second order of iteration

$$\ln \Delta_R^2 = I_0 + \ln \left[\left(1 + \frac{I_1^2}{4} + \frac{I_2}{2} \right)^2 + \frac{I_1^2}{2} \right]$$

$$I_0 = - \int_0^\infty \frac{dx}{x} W'(x) G(\ln x)$$

$$I_1 = \frac{1}{\sqrt{2}} \int_0^\infty \frac{dx}{x} X(x) G'(\ln x)$$

$$I_2 = -4 \int_0^\infty \frac{dx}{x} \left(X + \frac{X'}{3} \right) \frac{f'}{f} \int_0^x \frac{du}{u^2} \frac{f'}{f}$$

Differs from standard second-order slow-roll approximation as order counting is different.

Second order in $1/\Delta N$ expansion

$$\begin{aligned} \ln \Delta_R^2 \approx & G(\ln x_f) + q_1 G'(\ln x_f) + q_2 G''(\ln x_f) \\ & + \frac{\pi^2}{8} [G'(\ln x_f)]^2 - 4 \left[\frac{f'}{f}(\ln x_f) \right]^2 \end{aligned}$$

Monodromy case study

We consider monodromy inflation

Silverstein et al, 0803.3085

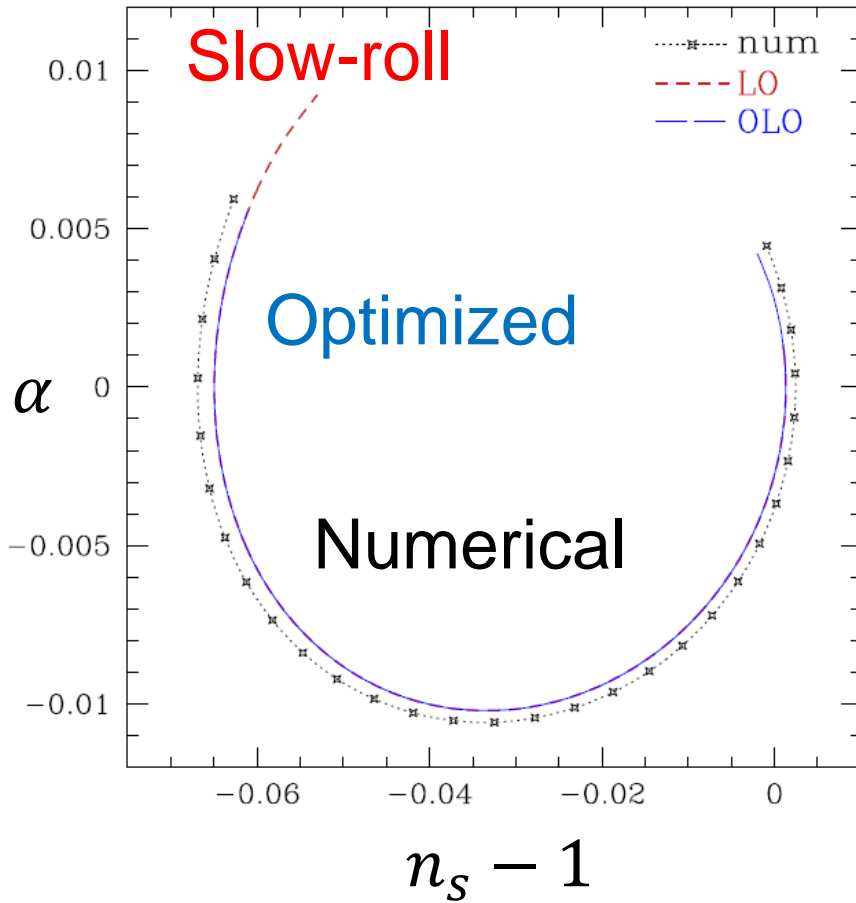
$$V(\phi) = \bar{V}(\phi) + \delta V(\phi)$$

$$\bar{V}(\phi) = \lambda\phi$$

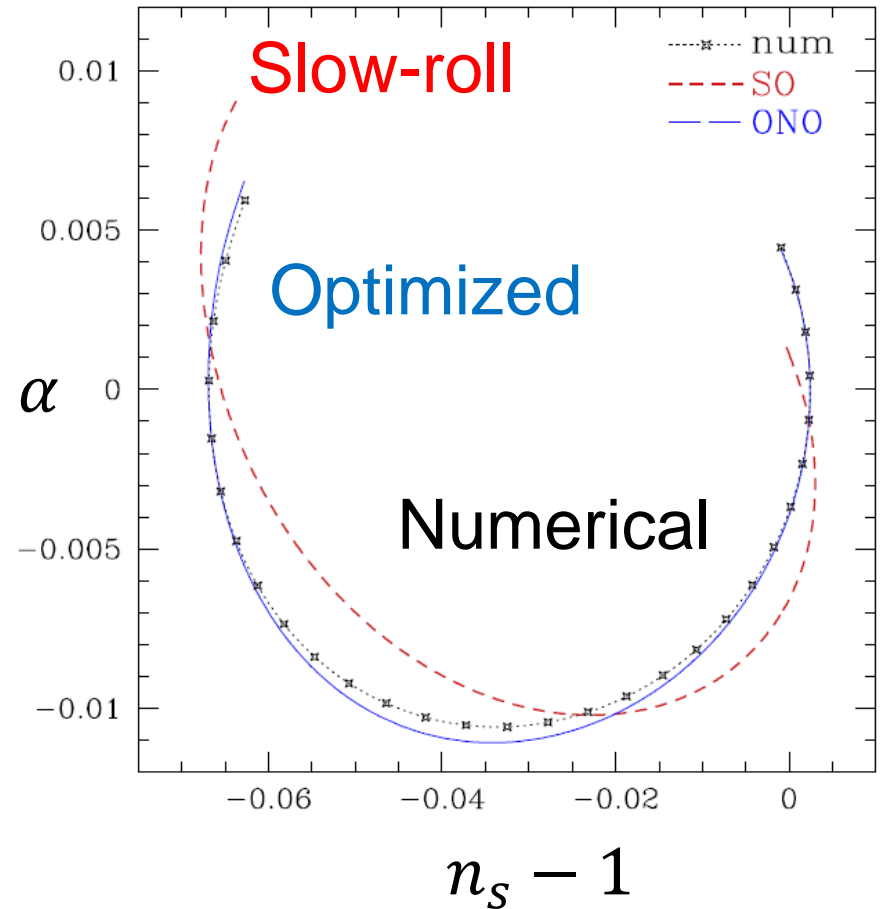
$$\delta V(\phi) = \Lambda^4 \cos\left(\frac{\phi}{f} + \theta\right)$$

Monodromy case study – low frequency

Leading order $\omega \sim \Delta N^{-1} \sim 0.3$ Next-to-leading order



Same formula with different evaluation point



Different formulae with different evaluation point

Monodromy case study – high frequency

Evaluate integrals with stationary phase approximation

$$\ln \Delta_R^2 = I_0 + \ln \left[\left(1 + \frac{I_1^2}{4} + \frac{I_2}{2} \right)^2 + \frac{I_1^2}{2} \right]$$

$$\delta \ln \Delta_R^2 \approx \delta I_0 + 2\bar{I}_1 \delta I_1 + \delta I_1^2$$

$$\bar{I}_1 = \frac{\pi}{2\sqrt{2}} (1 - \bar{n}_s)$$

$$\delta I_0 = -A \left[\cos(\psi - \beta) - \frac{3f}{2\phi_*} \sin(\psi - \beta) \right]$$

$$\delta I_1 = \frac{A}{\sqrt{2}} \tanh \left(\frac{\pi\omega_*}{2} \right) \left[\sin(\psi - \beta) + \frac{3f}{2\phi_*} \cos(\psi - \beta) \right]$$

- $O(A^2)$ cf. Flauger et al, 0907.2916 → - $O(A)$
Flauger et al, 1412.1814 - Phase is free parameter

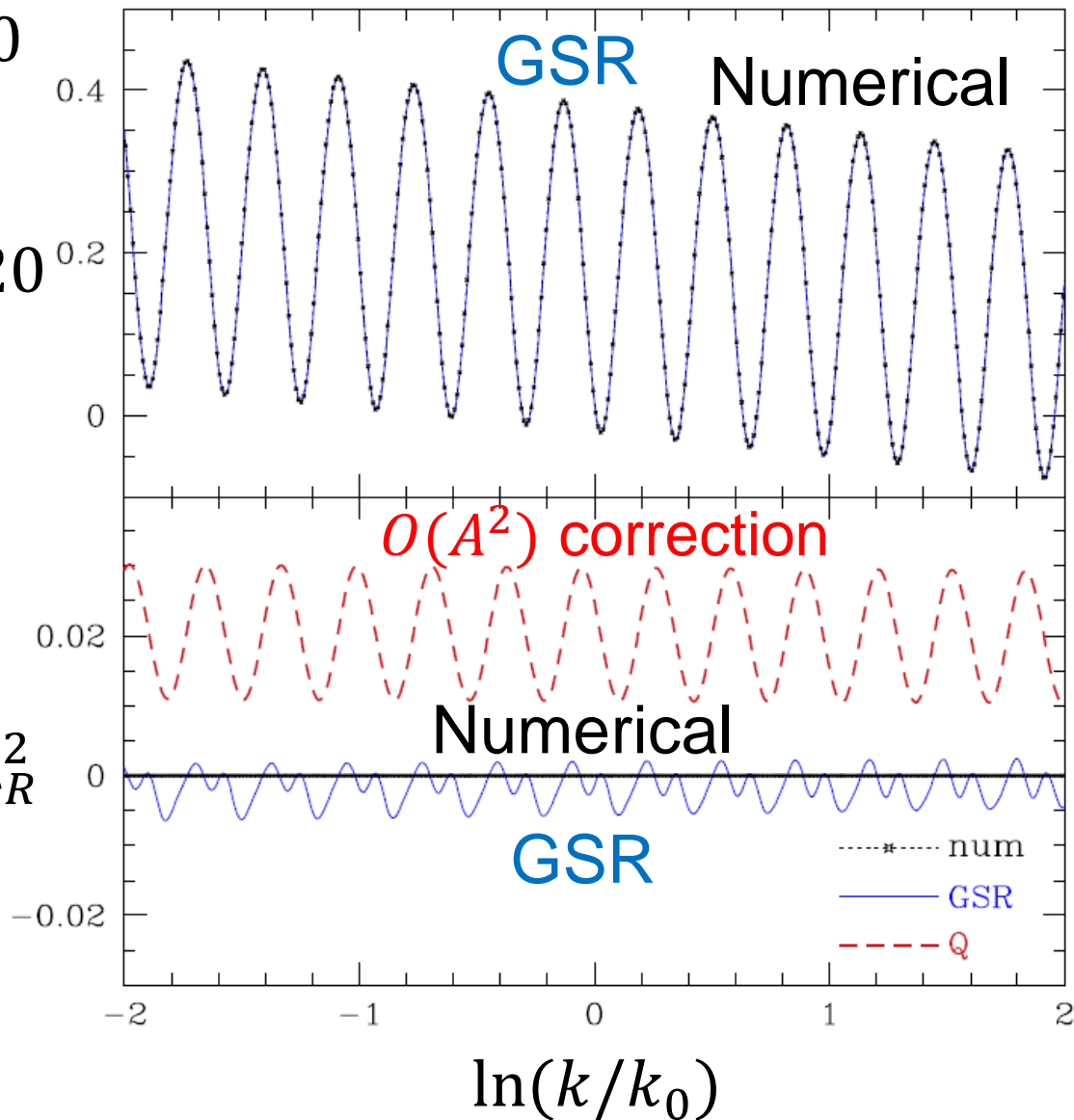
- All terms are written in terms of model parameters

Monodromy case study – high frequency

$\omega \sim \Delta N^{-1} \sim 20$
 $A \sim 0.2$

$\ln \Delta_R^2 + 20$

$\Delta \ln \Delta_R^2$



Summary

- Using GSR approximation, we developed optimized evaluation of inflationary power spectra that are written by Taylor expansion of features.
- Optimized leading order approximation reabsorb next-to-leading order corrections in Taylor expansion by evaluating the same formulae as the standard slow-roll at ~ 1 efold before horizon exit.
- We revealed that standard second-order slow-roll approximation fails while optimized next-to-leading order leads to 10^{-3} accuracy.
- For high frequency monodromy, we obtained $\ln \Delta_R^2$ up to $O(A^2)$ and clarified model parameter dependencies.