



Search for a Dark Photon at BABAR

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Search for a Dark Photon at BABAR



- Motivation
- The BABAR Experiment
- Search for a Dark Photon, A'

PRL **113**, 201801 (2014)

05 Aug 2015

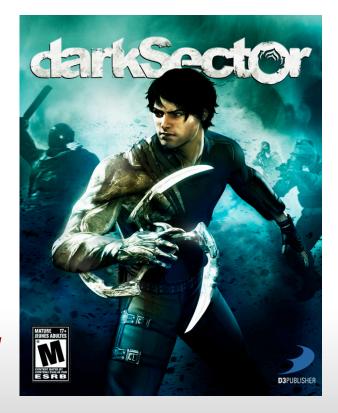


Motivation – Search for the Dark Sector



- Existence of dark matter is well-established from astrophysical evidence, but its nature is not known
- Collider experiments allow for:
 - Direct searches for dark matter particles through decays to Standard Model (SM) particles
- Electron-positron collider experiments are particularly clean environments and in many cases provide the best reach in searches for new physics

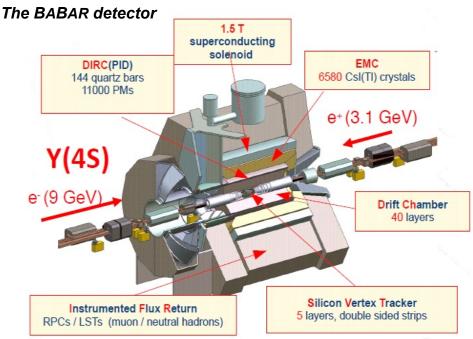
Zwicky, AcHPhys 6 (1933); et al.



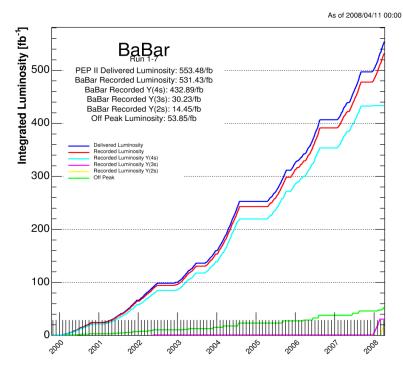


The BABAR Experiment





- Primarily designed for study of *CP*-violation in *B* meson decays
- Quality and general-purpose design make it suitable for a large variety of studies



BABAR data sample contains

~120 x
$$10^6 \Upsilon(3S)$$
 (10x Belle)

~100 x
$$10^6 \text{ Y(2S)}$$
 (10x CLEÓ)

~ 23 x 10⁶
$$\Upsilon(2S,3S) \rightarrow \Upsilon(1S) \pi^{+}\pi^{-}$$



Dark Sector Overview

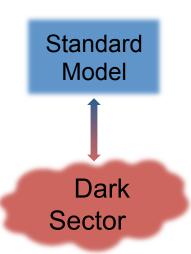


- Theory introduces idea of a new force corresponding to a new U(1)'.
 PLB 166, 196 (1986)
- Applied to the dark sector, the gauge boson, the so-called dark photon (A'), may be light (MeV – GeV mass) in these models.

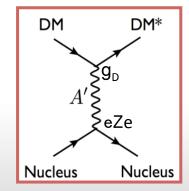
PLB 662, 53 (2008); PRD 79, 015014 (2009); arXiv:1311.0029, for example

- Dark sector particles do not couple directly to the SM content. Interaction dark sector SM via kinetic mixing between the dark photon and photon/Z with a mixing strength & among other "portals".
- In other words, there is a dark photon SM fermion coupling $\alpha' = \varepsilon^2 \alpha$ Strength small, but how small?

Currently favor $10^{-7} < \varepsilon < 10^{-2}$



[Slatyer, Schuster&Toro,...]



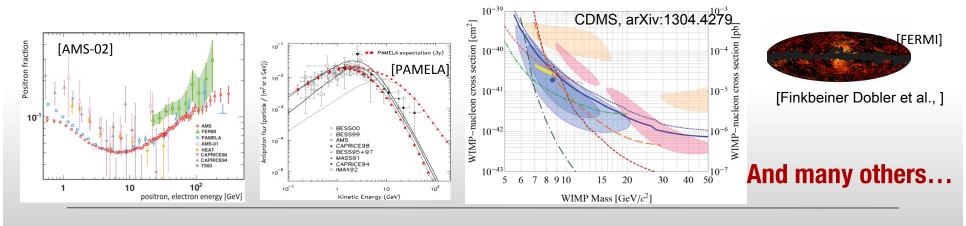
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Dark Sector Overview contd.



- In this framework, wimp-like TeV-scale dark matter particles can annihilate into dark photons, which subsequently decay to SM fermions.
- If the dark photon is light \rightarrow can only decay to light states. Could explain recent observations in cosmic rays (electron excess but no antiprotons) and by ground experiments.
- Other explanations of these anomalies have been proposed, but the possibility of a hidden MeV/GeV-scale sector is poorly constrained and really worth exploring.



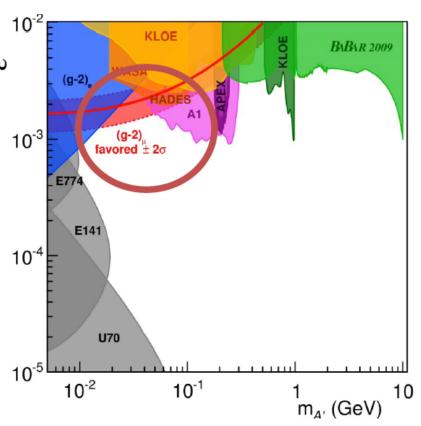
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Particle Physics Implications



- Can produce dark photons. In fact, photons ω in any process can be replaced by a dark photon (with an extra factor of ε).
- Decays back to lepton/quark pairs → search for resonances
- Dark photon decay can be prompt or displaced (long-lived)
- Current bounds on the mixing parameter *ε* are shown as a function of the dark photon mass.
- Constraints from electron/muon g-2, beam dump and fixed target experiments and e⁺e⁻ colliders (some constraints reinterpreted from limits of other measurements by theorists, e.g. BABAR)



Plot courtesy B. Echenard

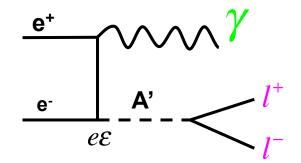


Dark Photon Production



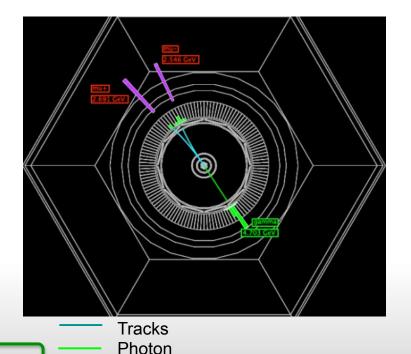
A dark photon can be produced in

$$e^+e^- \rightarrow \gamma A', A' \rightarrow e^+e^-, \mu^+\mu^-$$



Select events with one photon and two oppositely charged leptons*. Look at spectrum of dilepton mass. Use reduced mass for muons.

$$m_{\rm Red} = \sqrt{m_{\mu\mu}^2 - 4m_{\mu}^2}$$



*with further cuts to reduce radiative Bhabhas

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Signal in muon/hadron detector



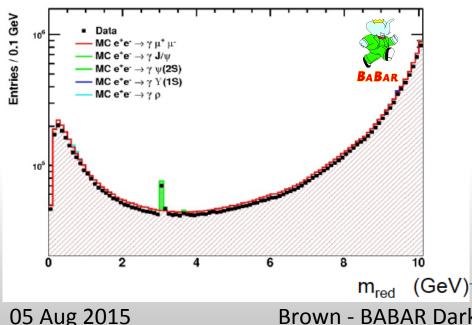
Analysis Technique



Scan the dielectron and dimuon reduced mass spectra and fit a background plus signal function at each step. Background includes resonances - ρ^0 , ϕ $J/\psi, \psi(2S), \Upsilon(nS)$

Mass resolution varies between 1.5 and 8 MeV Window size 30x mass resolution Step size approximately half the mass resolution

Electrons: 5704 fits, 0.02 - 10.2 GeV 5370 fits, 0.212 - 10.2 GeV Muons:



Assign a statistical significance for each fit:

$$S_S = \sqrt{2\log(L/L_0)}$$

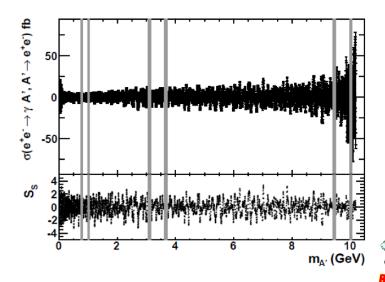
likelihood w/background + signal

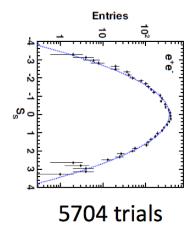
L₀: likelihood w/background only



Results – both channels



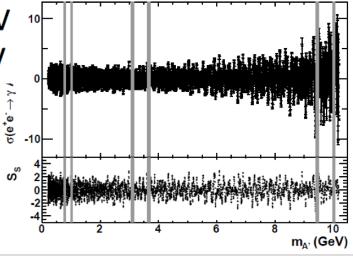


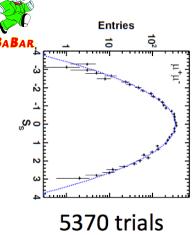


Highest fluctuations

> 3.4 σ (e) M=7.02 GeV

 \geq 2.9 σ (μ) M=6.09 GeV



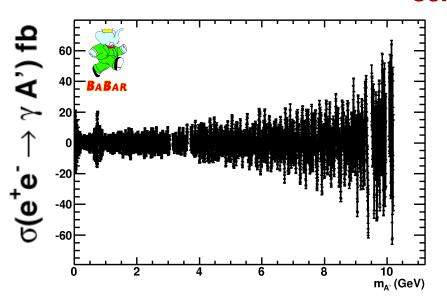


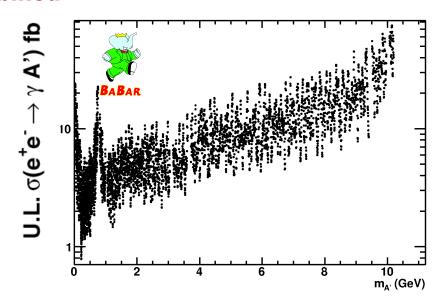


Cross-section Results



Combined





90% Confidence Level Upper Limit

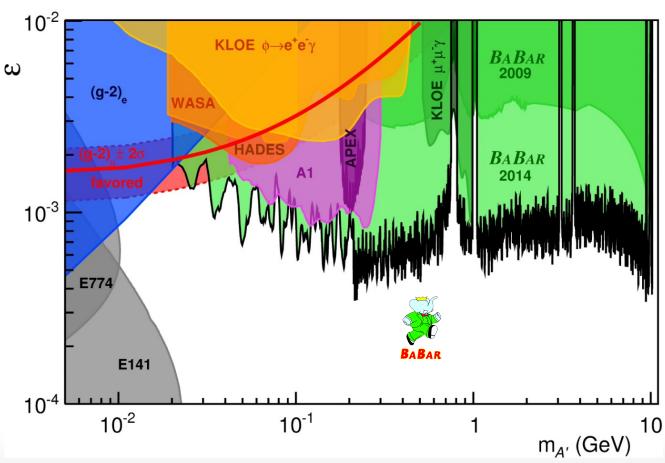
Including Trial factors,

Most significant excursion from null hypothesis for electrons: 0.6σ Most significant excursion from null hypothesis for muons: 0.1σ



Dark Sector Mixing Results





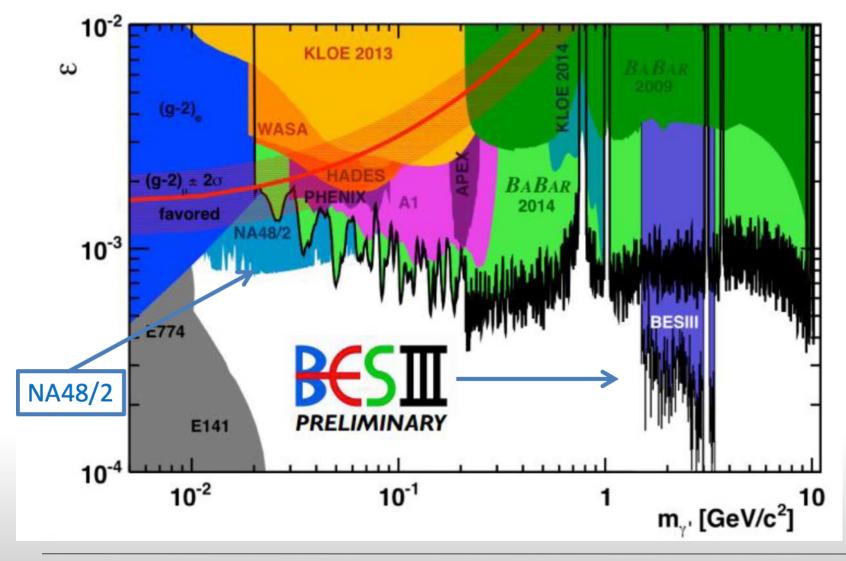
201801 (2014) 113, Phys. Rev. Lett.

• Further exclude the region favored by the g-2 measurement and improve the existing constraints over a wide range of masses.



Recent Addition to the Story





05 Aug 2015



New Search for Dark Photon



arXiv:0808.0017

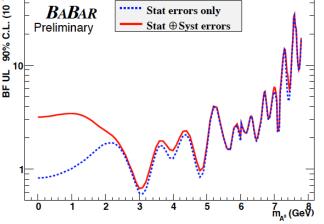
The Dark photon could decay to dark particles, which would be invisible to a SM detector.

$$e^+e^- \rightarrow \gamma A', A' \rightarrow \text{invisible}$$

Experimentally, search for single photon events, look for peak in Energy spectrum. Preliminary result based on small portion of BABAR data taken at $\Upsilon(3S)$ energy. Will be extended to full dataset.

$$\mathcal{B}(\Upsilon(3S) \to \gamma A^0) \times \mathcal{B}(A^0 \to invisible) < 31 \times 10^{-6}$$

At 90% CL **BABAR Preliminary**



Belle II could push this limit down significantly





- Dark matter is well-established but mysteries remain
- BABAR has completed searches for the dark photon
 - No evidence for it, but
 - Tighten constraints on dark sector models
- The dark sector is an exciting field, to which electron-positron collider experiments can make significant contributions



Thank you to B. Echenard for use of several slides







Thanks!

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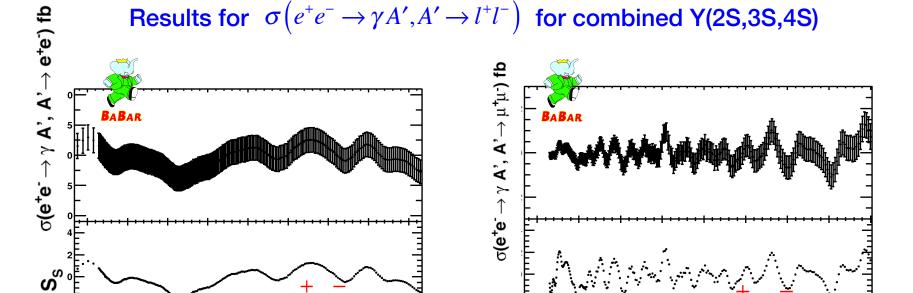


Back Up Slides



Dark Photon Results





$$S_{S} = \sqrt{2\log(L/L_{0})}$$

Here, look in 200 MeV region as suggested by excess in HyperCP results

No excess observed

05 Aug 2015

PRL 94, 021801 (2005)

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 $m_{\Delta^{+}}(GeV)$

0.235