

# Identifying the Theory of Dark Matter with Direct Detection

Sam McDermott

[arXiv:1506.04454](https://arxiv.org/abs/1506.04454) and [ascl.net/1506.002](https://ascl.net/1506.002)

with [Vera Gluscevic](#), [Moira Gresham](#), [Annika H. G. Peter](#), and [Kathryn Zurek](#)



# Motivation

- Dark matter exists
- Large ongoing experimental effort to discover its interactions with the Standard Model
- Imagine we detect dark matter: what information will we be able to extract?

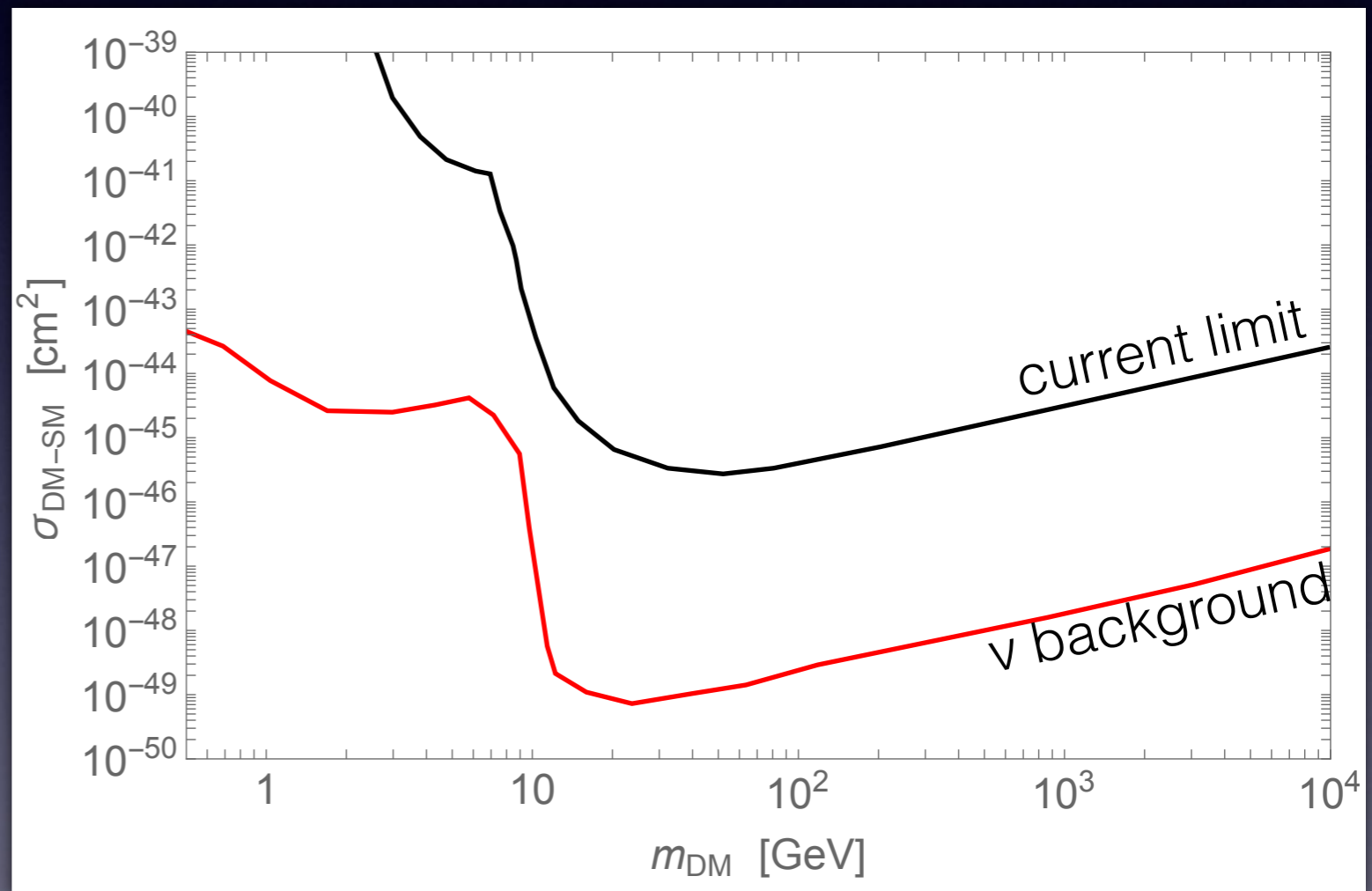
also, McDermott, Yu, Zurek 1110.4281,  
Gluscevic, Peter 1406.7008,  
Catena, 1406.0524, 1407.0127

# Parameter Space

Direct detection experiments are probing several orders of magnitude in  $m_{\text{DM}}$

Improving in  $\sigma_{\text{DM}}$  about x2/yr

Could make a discovery soon!



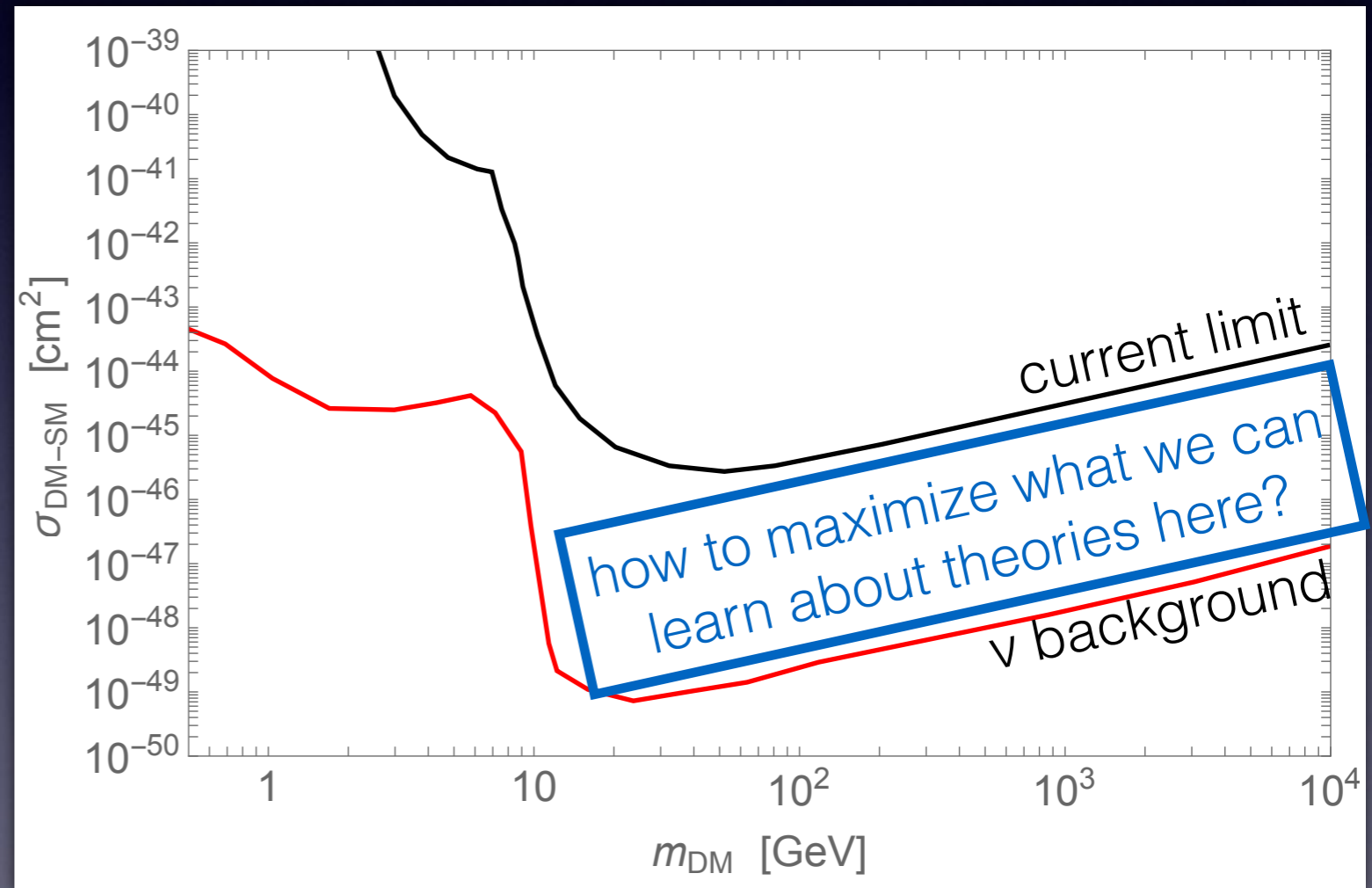
adapted from Snowmass document

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# Observables

$$\frac{dR}{dE_R}(E_R) = \frac{\rho_\chi}{m_T m_\chi} \int_{v_{\min}}^{v_{\text{esc,lab}}} v f(\mathbf{v}) \frac{d\sigma_T}{dE_R}(E_R, v) d^3v$$

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recoil  
spectrum

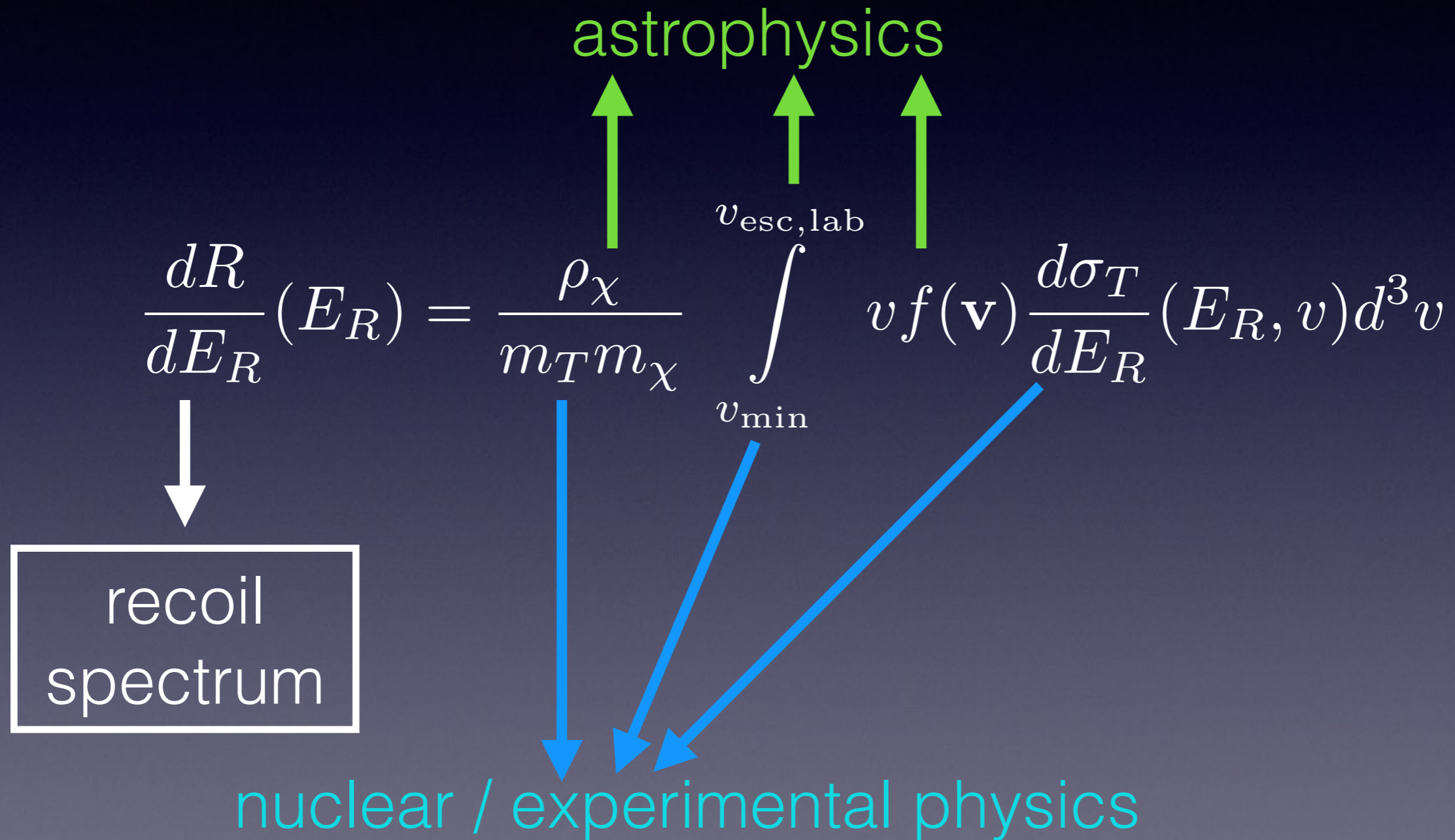
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astrophysics

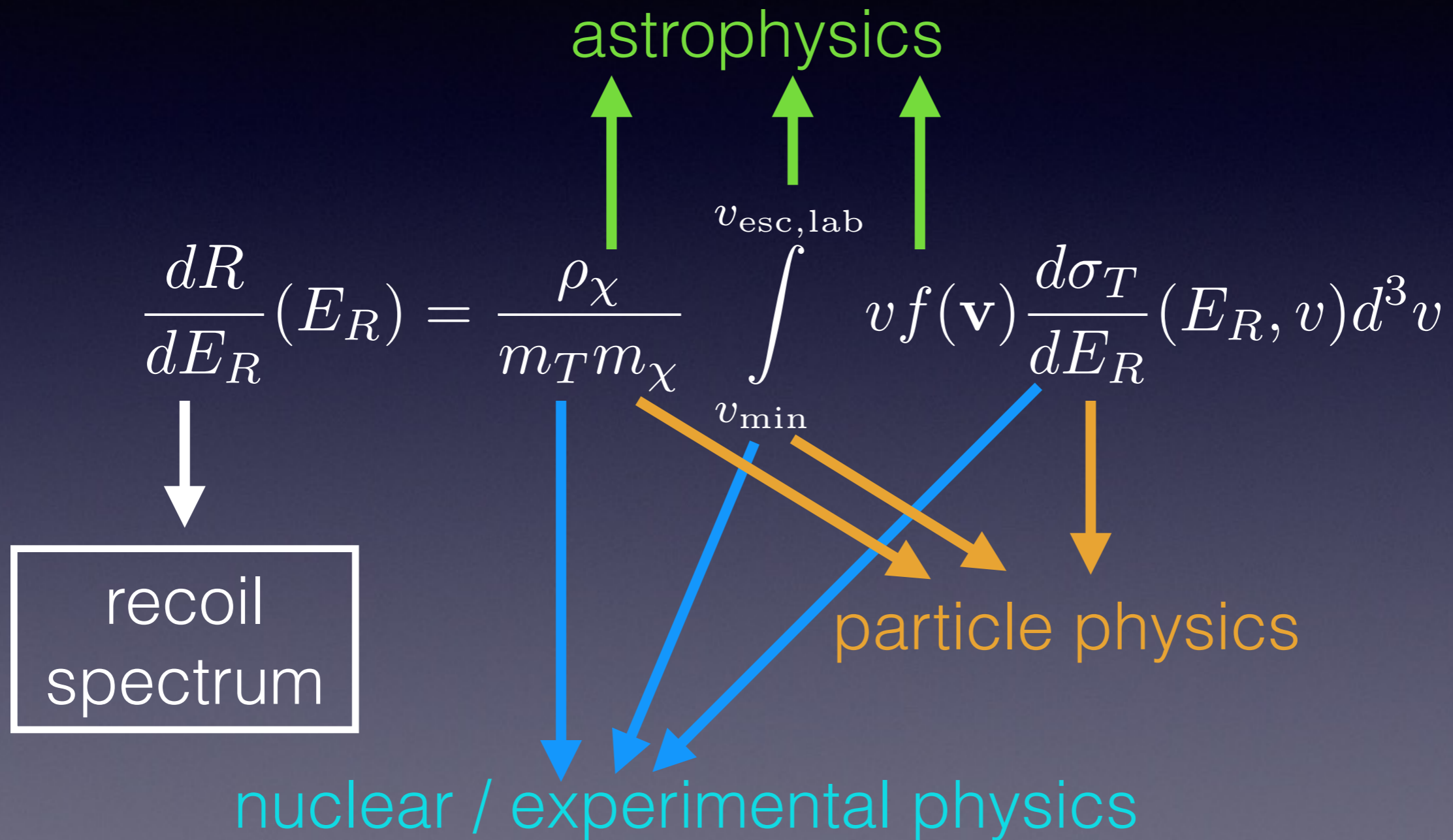
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# Requirements

- A set of models (hypotheses) & corresponding phenomenology of scattering off nuclei
- A statistical representation of experiments
- An analysis framework for evaluating how well a given hypothesis fits a single data realization

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Bayesian model  
selection, cf. Gluscevic and  
Peter 1406.7008

# EFT of DD

$$\frac{d\sigma_T}{dE_R}(E_R, \nu) = \frac{m_T}{2\pi\nu^2} \sum_{(N, N')} \sum_X R_X \left( E_R, \nu, c_i^{(N)}, c_j^{(N')} \right) \widetilde{W}_X^{(N, N')}(y)$$

“particle physics”

given a Lorentz-invariant theory,  
calculate the low energy, non-  
relativistic cross section

nuclear form factors

given a Lorentz-invariant theory,  
nuclear physics measurements  
predict nuclear responses

$X = M, \Sigma', \Sigma'', \Phi'', \Delta, M\Phi'', \Delta\Sigma'$  (responses)

$y \equiv m_T E_R b^2 / 2, \quad (b/\text{fm})^2 \equiv 41.467 / (45A^{-1/3} - 25A^{-2/3})$

# Particle Physics: $R_\chi$

lowest-dimension, least-suppressed Lorentz-invt.  
products of DM fermion bilinears with SM fields

- “standard” — SI ( $\bar{\chi}\chi\bar{f}f$ ), SD ( $\bar{\chi}\gamma^\mu\gamma_5\chi\bar{f}\gamma_\mu\gamma_5f$ )
- photon-mediated — millicharged ( $\bar{\chi}\gamma^\mu\chi A_\mu$ ),  
anapole ( $\bar{\chi}\gamma^\mu\gamma_5\chi\partial^\mu F_{\mu\nu}$ ), magnetic dipole  
( $\bar{\chi}\sigma^{\mu\nu}\chi F_{\mu\nu}$ ), electric dipole ( $\bar{\chi}\sigma^{\mu\nu}\gamma_5\chi F_{\mu\nu}$ )

UV theory  $\Rightarrow$  overall momentum and velocity  
dependence, triggered responses

# “Nuclear Physics”: $\widetilde{W}_X$

- form factor = how rate falls off at higher energy
- depends on target, response, and energy

number of form factors =

= number of targets  $\times$  number of responses



# Statistical Methodology

$$\mathcal{E}(\{E_R\}|\mathcal{M}) = \int d\Theta \overset{\text{likelihood}}{\mathcal{L}(\{E_R\}|\Theta, \mathcal{M})} \overset{\text{prior}}{p(\Theta|\mathcal{M})}$$

observed (noisy) energy spectrum      given model

$$\text{Pr}(\mathcal{M}_j) = \frac{\mathcal{E}(\{E_R\}|\mathcal{M}_j)}{\sum_i \mathcal{E}(\{E_R\}|\mathcal{M}_i)}$$

free parameters ( $\sigma_{\text{DD}}$  and  $m_{\text{DM}}$ )

“What are the odds” of extracting the true underlying model from the data?

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(Note: not maximizing a  $\chi^2$  statistic or finding a best fit)

# Mock data

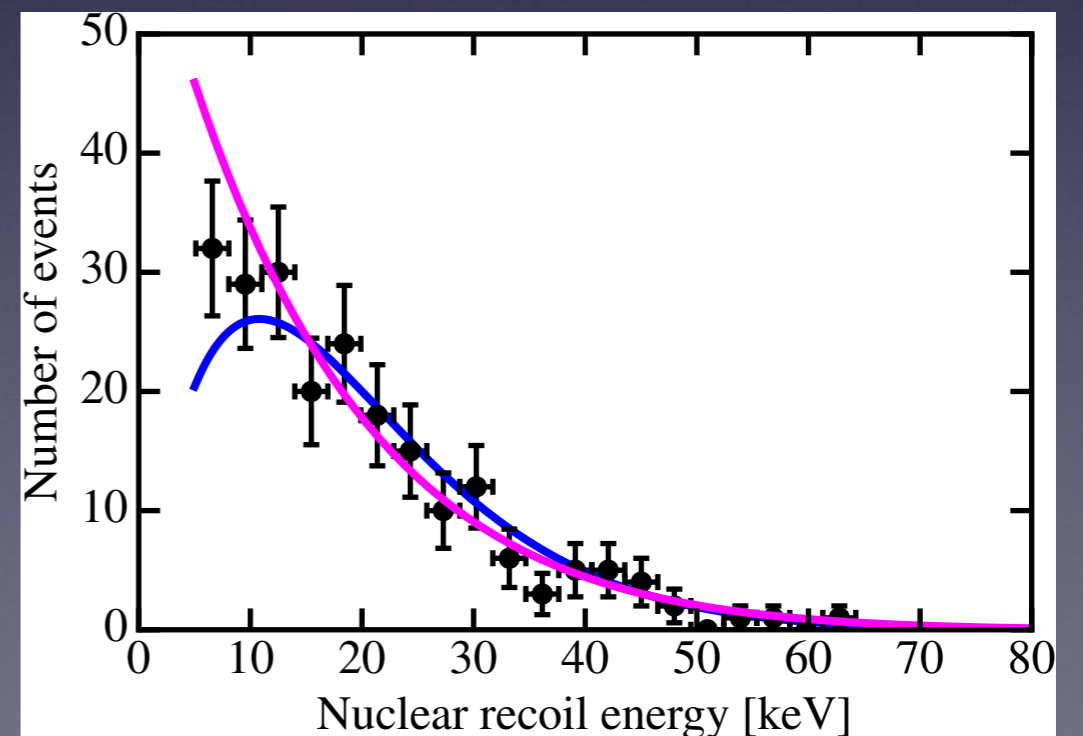
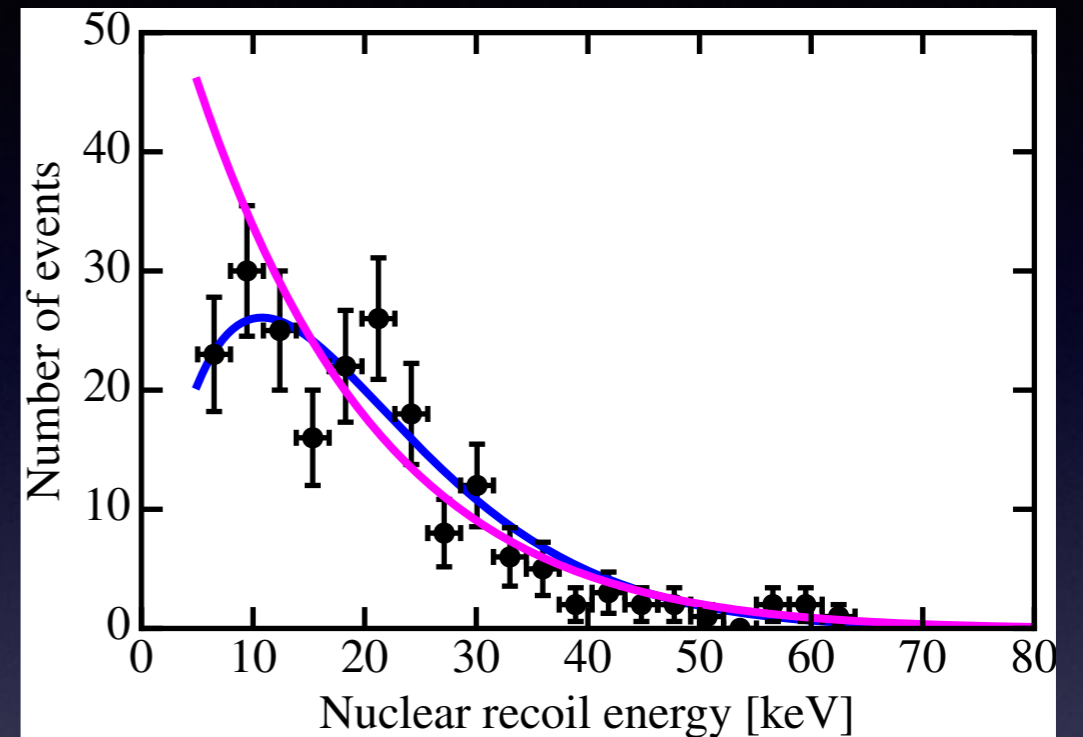
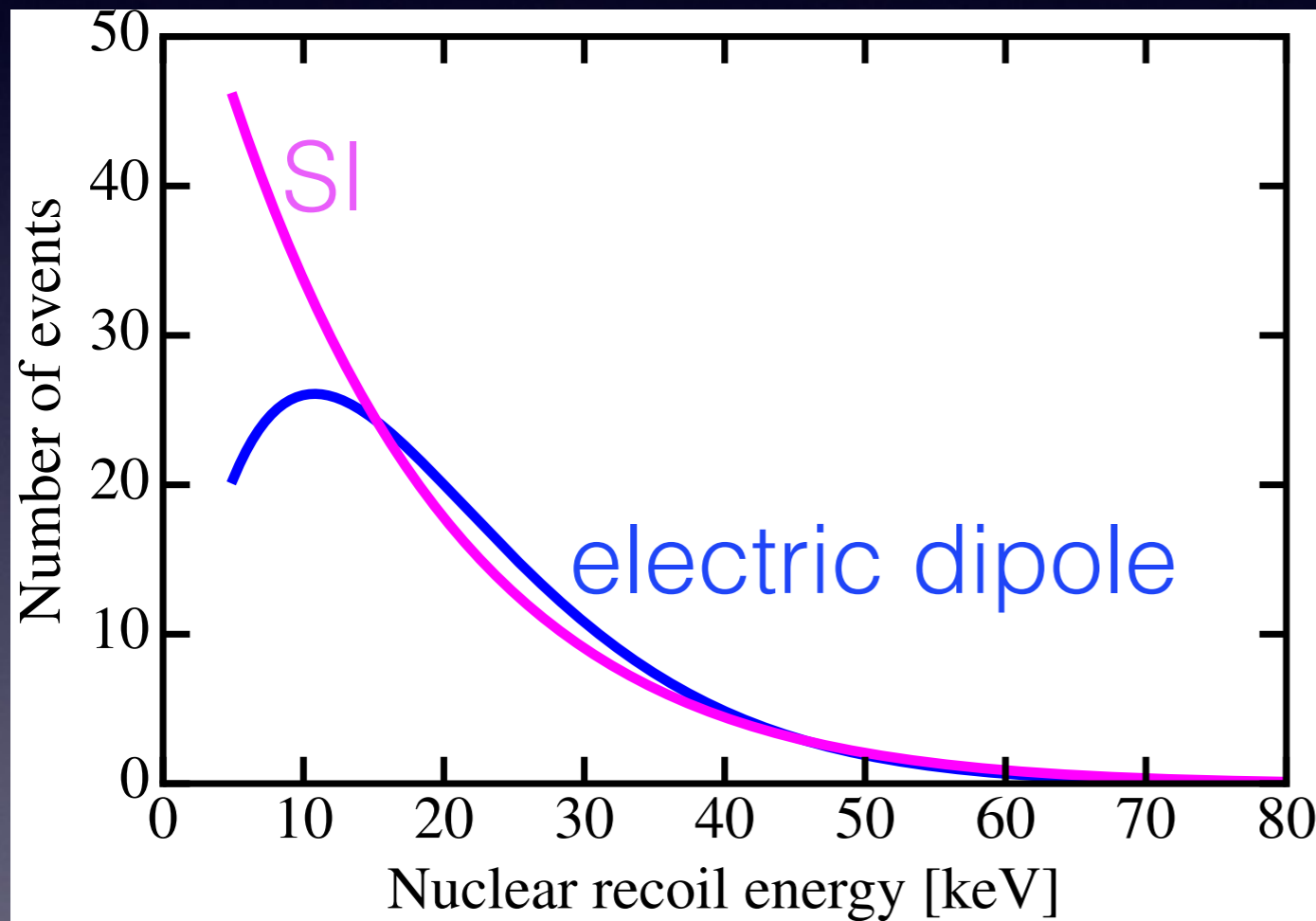
need to make mock data sets  $\{E_R\}$

simulate G2-like experiments

Label	A (Z)	Energy window [keVnr]	Exposure [kg-yr]
Xe	131 (54)	5-40	2000
Ge	73 (32)	0.3-100	100
I	127 (53)	22.2-600	212
F	19 (9)	3-100	606
Na	23 (11)	6.7-200	38

choose  $m_{\text{DM}}$ , set  $\sigma_{\text{DD}}$   
just below current limits

# A Single Simulation





# Criterion for Success

$$\Pr(\mathcal{M}_j) = \frac{\mathcal{E}(\{E_R\}|\mathcal{M}_j)}{\sum_i \mathcal{E}(\{E_R\}|\mathcal{M}_i)} > 90\%$$

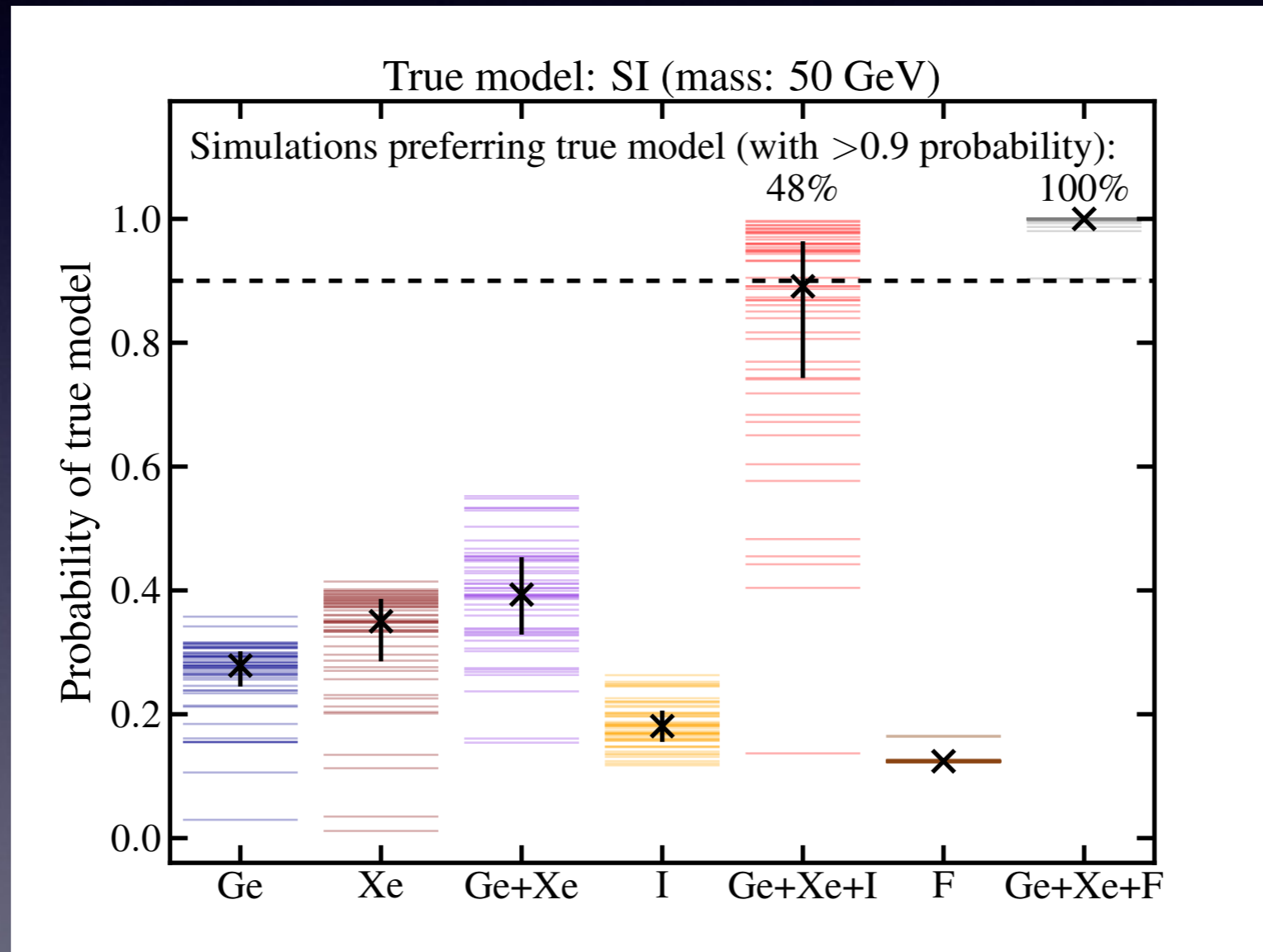
true underlying model is  
“confidently selected” if  
 $\Pr(M_{\text{true}}) > 90\%$



this depends on the  
Poisson realization

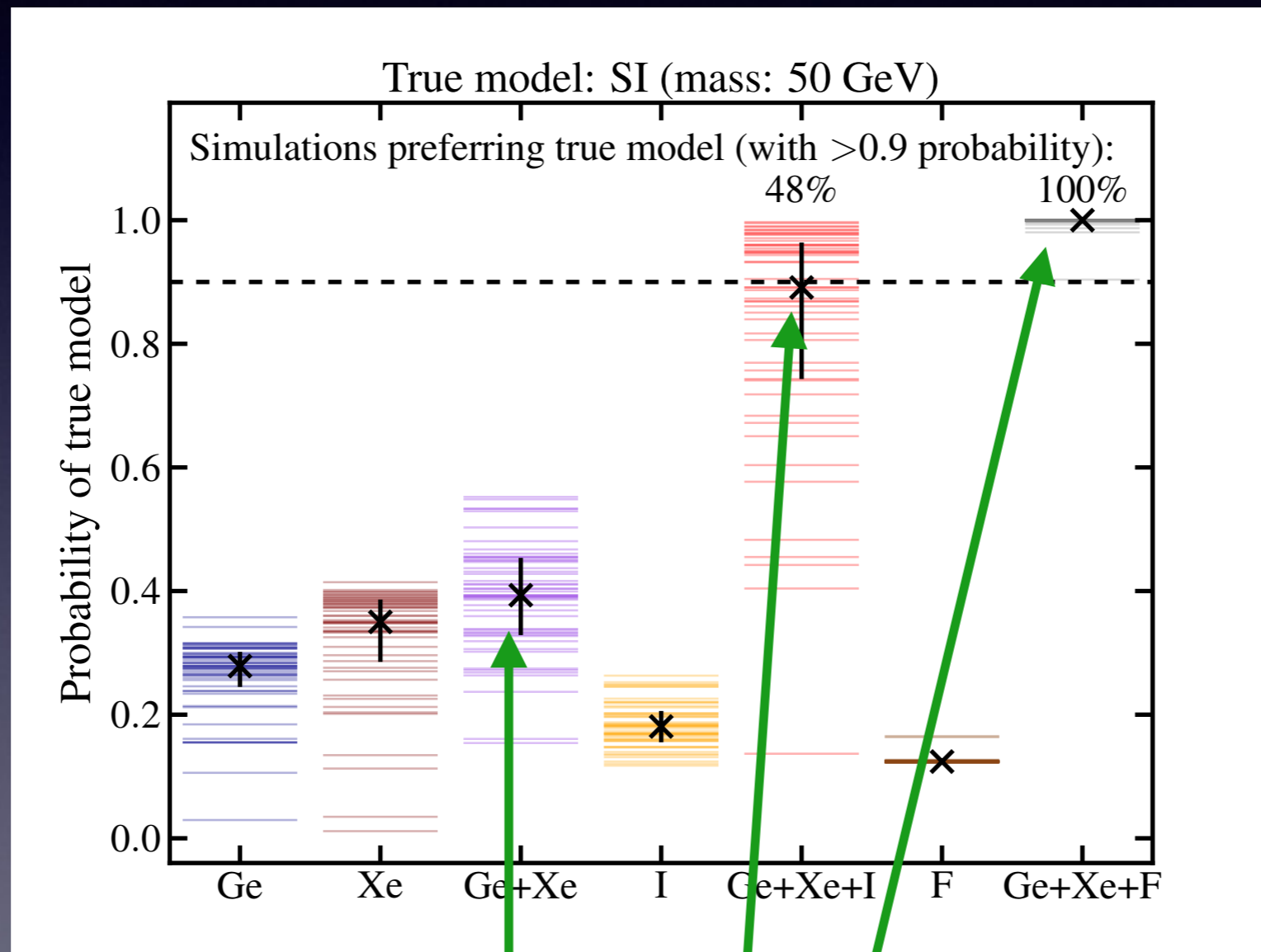
create many  
Poisson  
realizations to  
test robustness

# Results (example)



single elements not so good

# Results (example)



Complementarity!

# Conclusions

- A conclusive direct observation of DM will just be the beginning of the work
- Target complementarity will be a critical requirement for learning about DM physics