

# Neutrino Physics Prospects with PINGU

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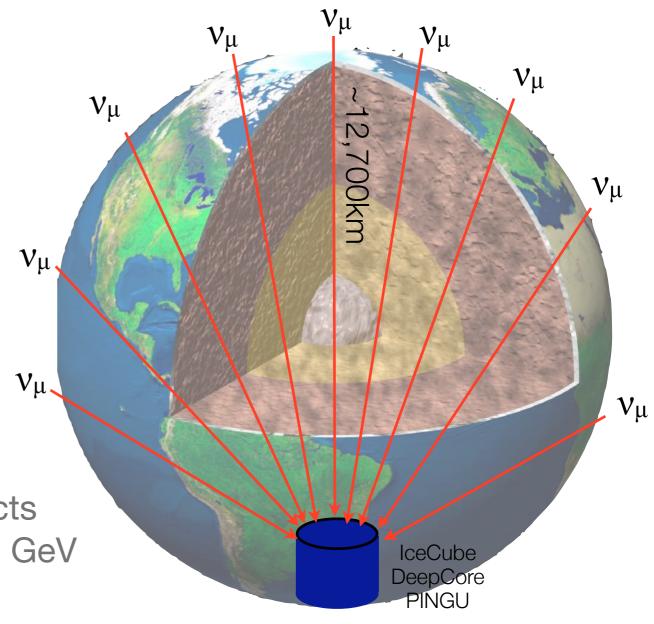


PRECISION ICECUBE NEXT
GENERATION UPGRADE

DPF 2015 Ann Arbor, Michigan August 7, 2015

# Oscillation Physics with Atmospheric Neutrinos

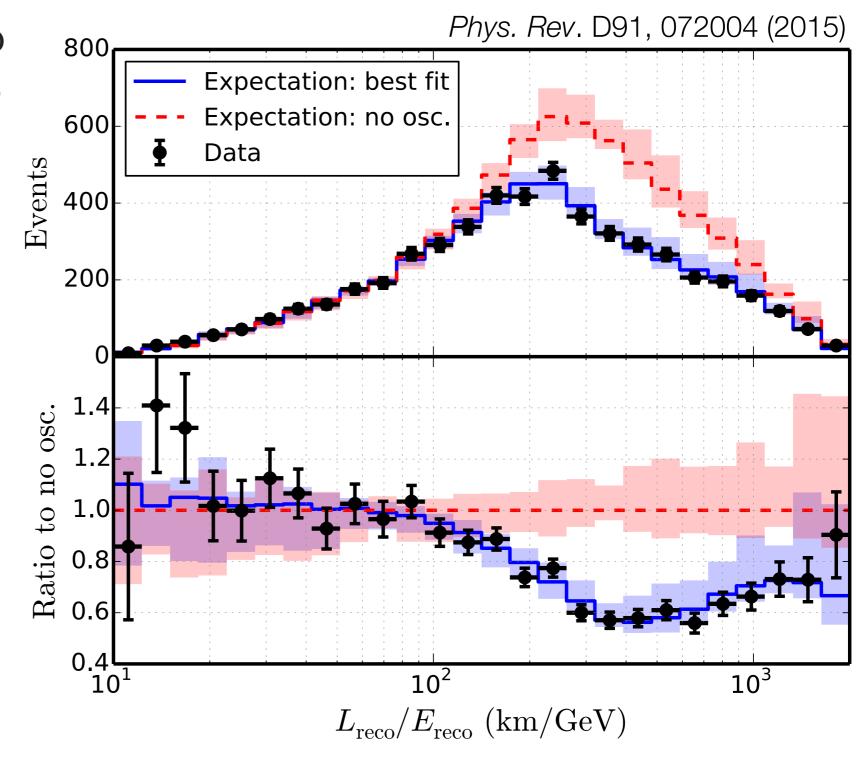
- Neutrinos available over a wide range of energies and baselines
  - Oscillations produce distinctive pattern in energy-angle space
  - Approach: control systematics using events in "side band" regions – trade statistics for constraints on systematics
- Neutrinos oscillating over one Earth diameter have a  $v_{\mu}$  survival minimum at ~25 GeV
  - Hierarchy-dependent matter effects on v or v̄ (MSW etc.) below 10-20 GeV





#### Current IceCube Oscillation Results

- Real analysis is 2D to maximize constraints on systematics
  - Data projected onto reconstructed (L/E<sub>v</sub>) here for illustration
  - Shaded range shows systematic uncertainties allowed by IceCube data
- Details and results in J. Hignight's talk (next)





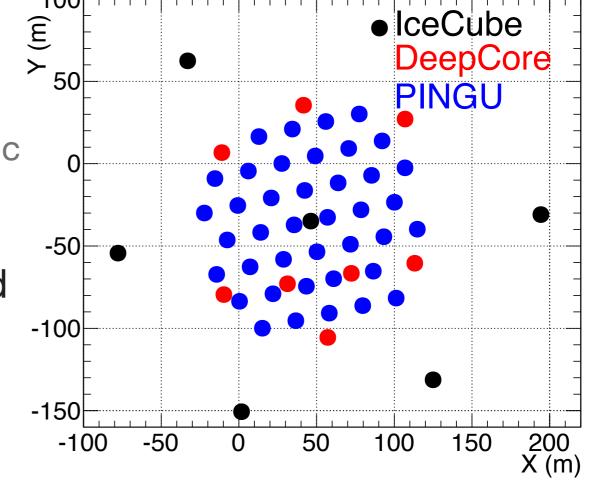
#### PINGU



PRECISION ICECUBE NEXT

- Baseline 4 MTon detector: 40 additional strings at 22 m spacing, with Digital Optical Modules spaced 3 m vertically, deployed inside IceCube DeepCore
  - Compare to 72 m string spacing and 7 m DOM spacing for DeepCore
  - ~25x higher photocathode density
  - Additional in situ calibration devices will better control detector systematic (not included in projections)
- Achieve few GeV energy threshold
- Engineering and costs are well understood from IceCube

Top view of the PINGU new candidate detector

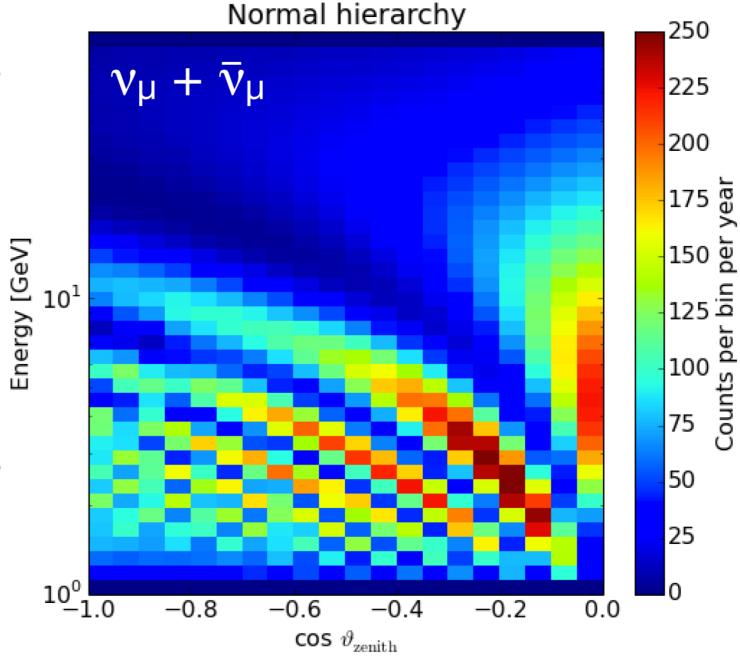




## Signatures of the Neutrino Mass Hierarchy

• Matter effects alter oscillation probabilities for v or  $\bar{v}$  traversing the Earth – exploit differences in cross section to distinguish

- Effects vary with E<sub>v</sub> and L (= zenith angle) due to Earth's density profile
- Neutrino oscillation probabilities affected if hierarchy is normal, antineutrinos if inverted
- Rates of all flavors are affected
- Note: effect of detector resolution not shown here
- Distinct signatures observable in both track (v<sub>μ</sub> CC) and cascade (v<sub>e</sub> and v<sub>τ</sub> CC, v<sub>x</sub> NC) channels
  - At higher energies, v<sub>µ</sub> CC events distinguishable by the presence of a muon track

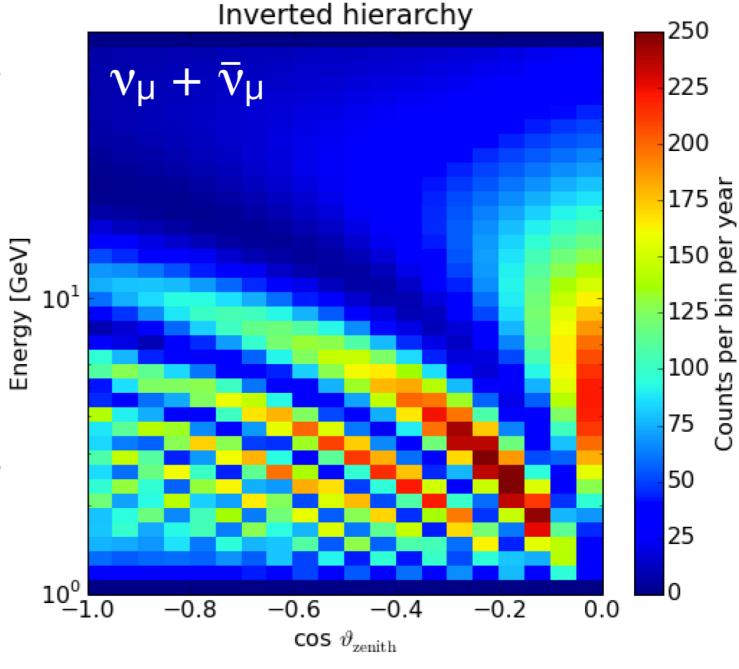




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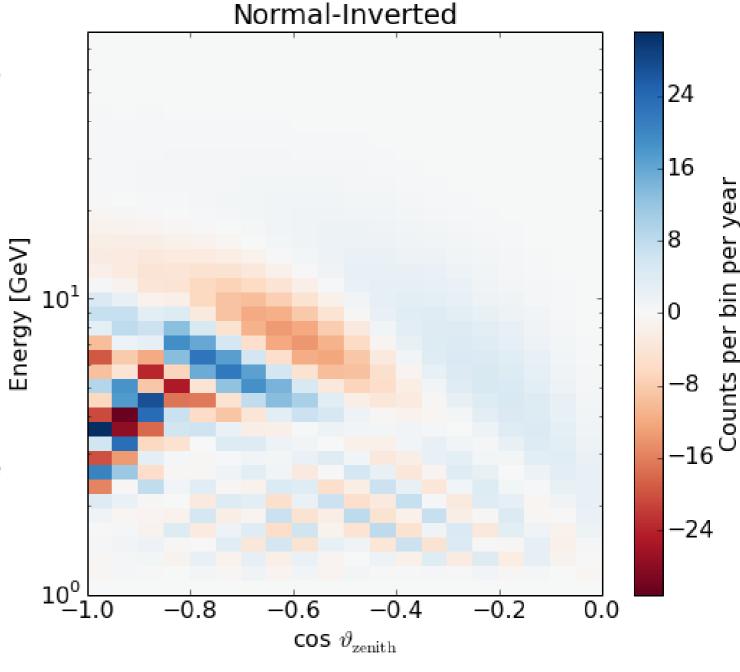




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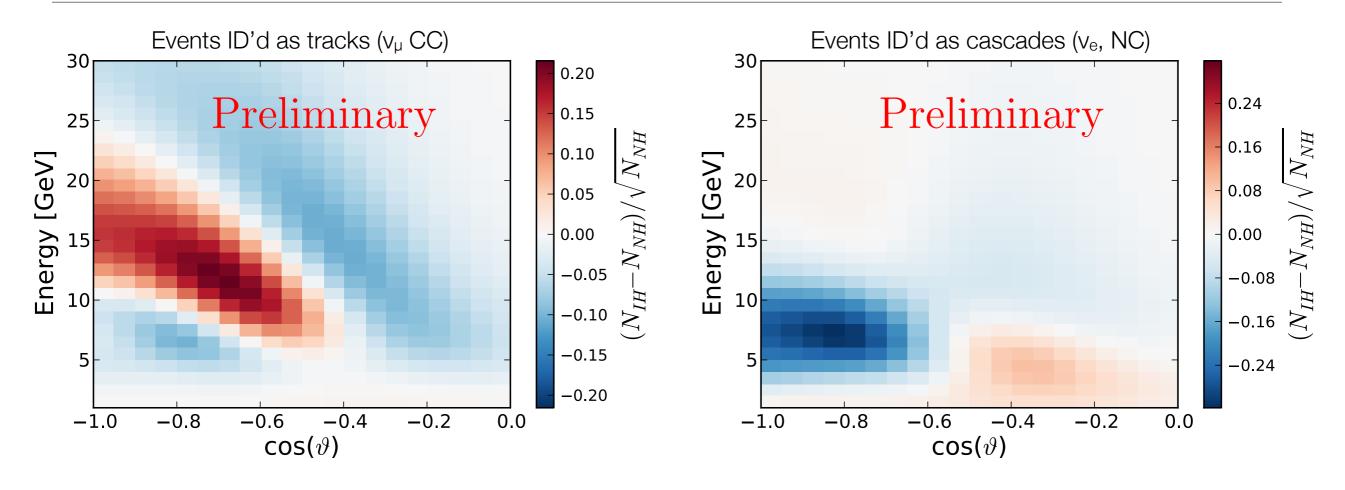
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# Hierarchy Signature: Observables

arXiv:1401.2046



- Simple visualization of statistical signal, with full detector response included
- Distinctive (and quite different) hierarchy-dependent signatures visible in both the track and cascade channels
  - Parametrized rates, detector resolutions and efficiencies from full detector Monte Carlo used to eliminate statistical fluctuations – statistical distributions checked with MC



#### Effects of Systematics

- Oscillation physics produces distinctive patterns unlike those of other effects
- Uncertainties in oscillation parameters (mainly  $\theta_{23}$ ) dominate systematics
  - No prior placed on  $\theta_{23}$  or  $\Delta m^2_{atm}$  fit jointly with NMH
  - θ<sub>13</sub> fit with prior, solar parameters and  $\delta_{CP}$  (=0) held fixed
- Flux: v<sub>e</sub>/v<sub>µ</sub> ratio (3%), v/v̄ ratio (10%), spectral index (5%), detailed flux uncertainties from Barr et al. 2006\*

Туре	3y σ (NH)	3 y σ (IH)
stat only	4.84	4.82
flux only	4.55	4.56
det only	4.06	3.99
$\theta_{23}$ only	3.52	3.26
osc only	2.96	2.53
All	2.90	2.51

 Detector: rate/normalization (free), energy scale (10%), detailed cross-section systematics from GENIE\* \*only with  $\Delta \chi^2$  method

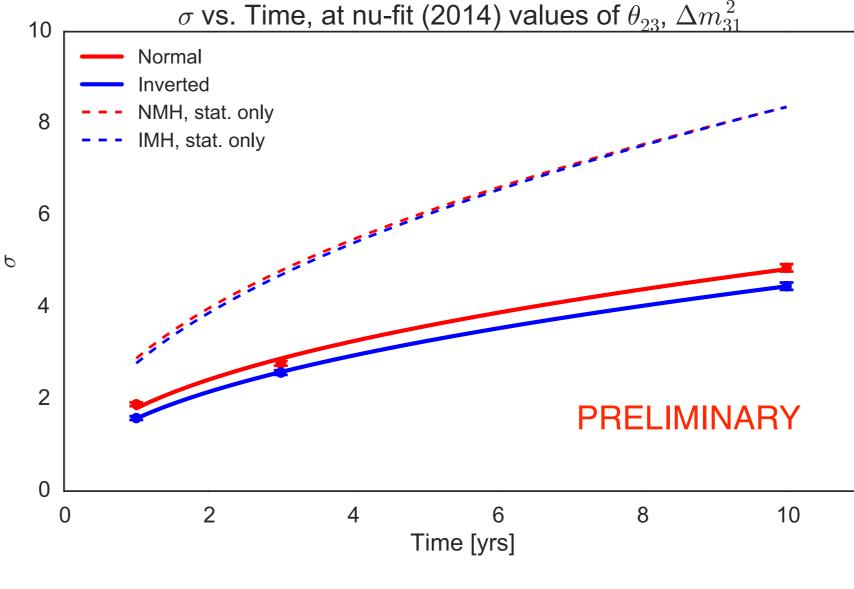


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## Significance vs. Time

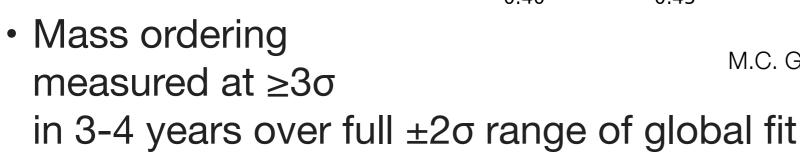
- Measurement strongly affected by systematics, but continues to improve with time
- Systematics are constrained by same data set
  - Increased statistics means gradually better control of systematics

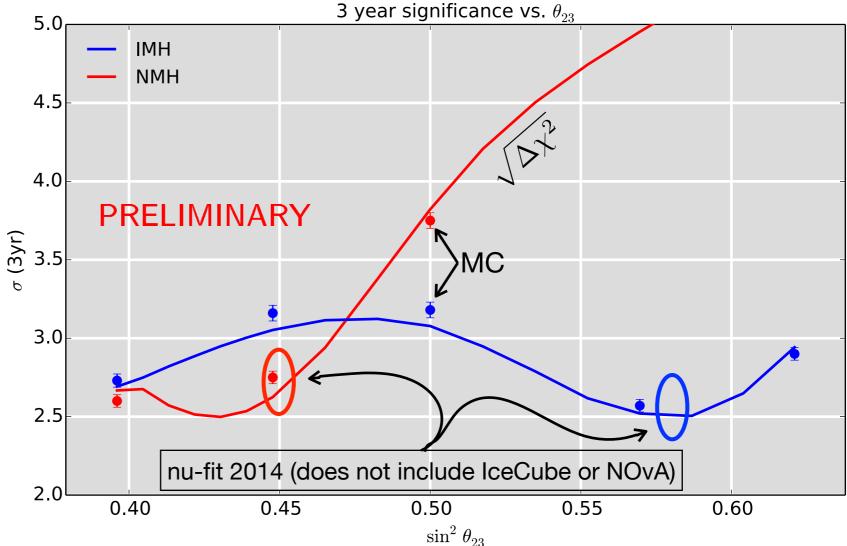




# Dependence on Mixing Angle

- Most values of θ<sub>23</sub>
   would give higher
   significance for
   mass ordering
  - Drift toward
     maximal mixing
     since PINGU Lol
     has increased both
     matter effects and
     degeneracies



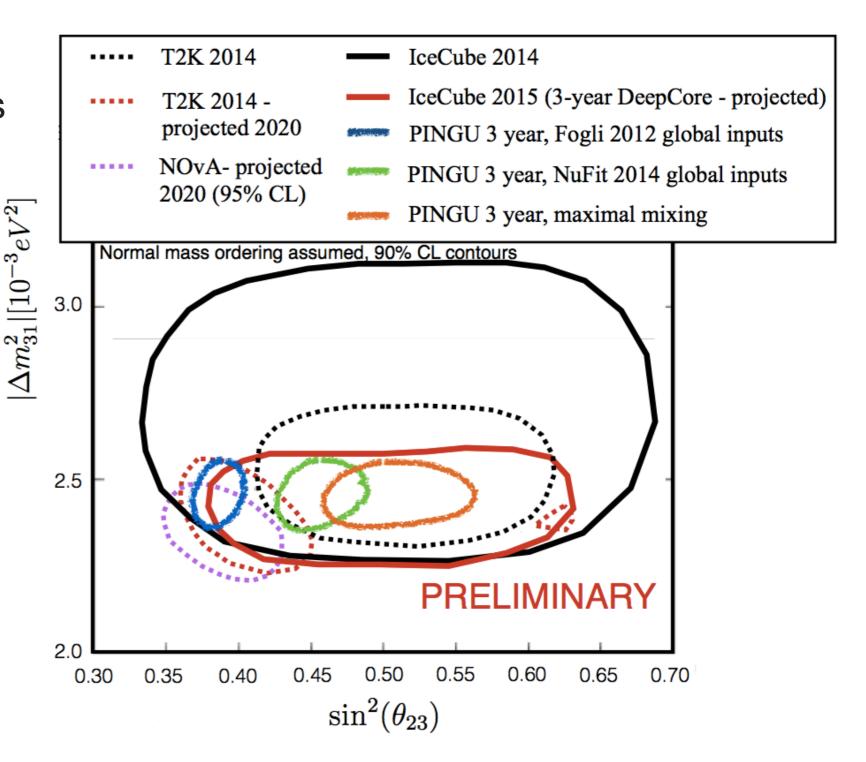


M.C. Gonzalez-Garcia et al. JHEP 11, 052 (2014)



#### Oscillation Parameters with PINGU

- Significantly improve IceCube measurements of  $\theta_{23}$  and  $\Delta m^2_{atm}$
- Comparable precision to NOvA, T2K
- Complementary to other measurements – interesting tests of standard oscillations
  - Higher energies, joint disappearancetau appearance measurement

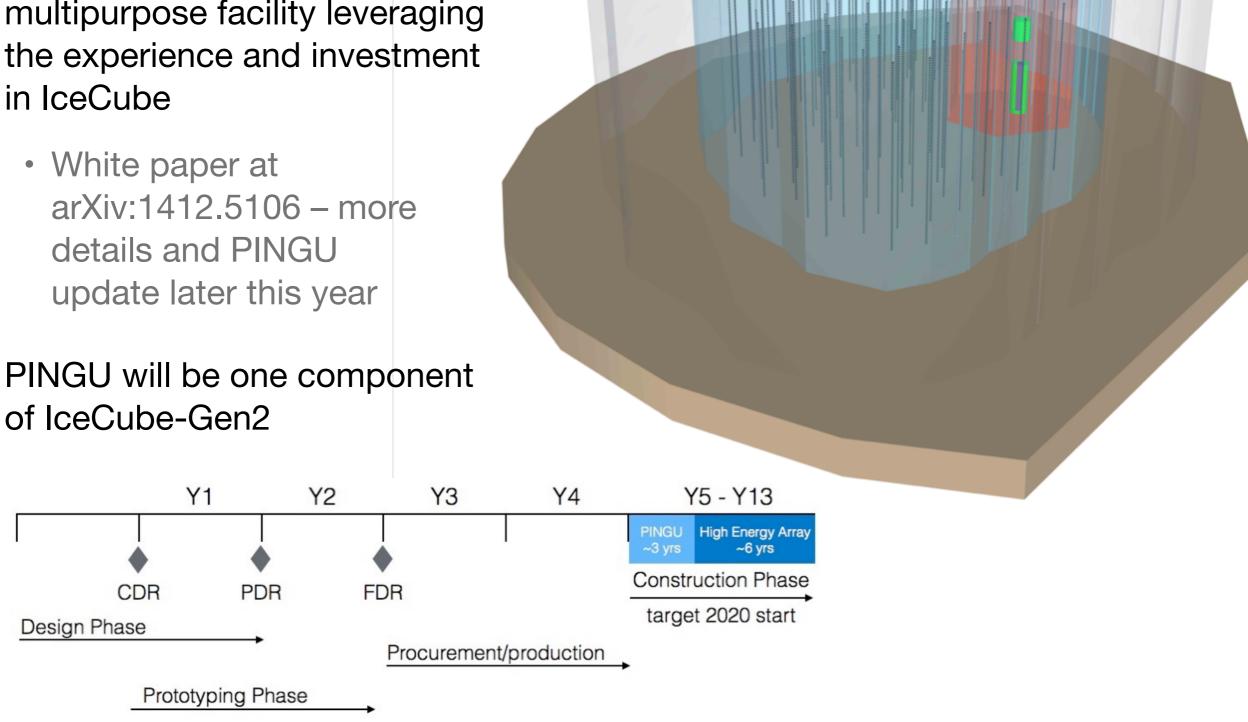




#### IceCube-Gen2

 Planning underway for a multipurpose facility leveraging the experience and investment in IceCube

PINGU will be one component



#### Outlook

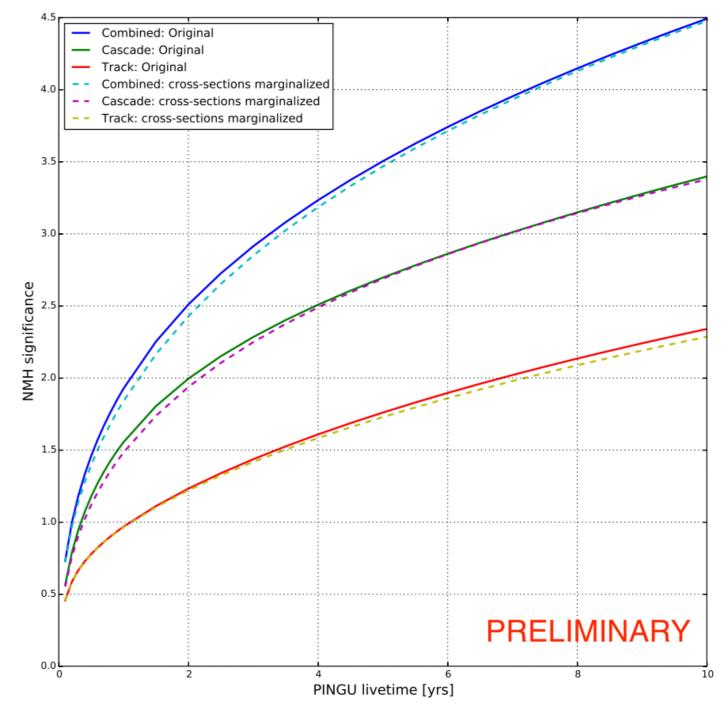
- PINGU has a unique place in the world-wide neutrino program
  - Measurements at a range of higher energies/longer baselines, with high statistics
- Opportunity to discover new physics is greatly enhanced by PINGU's complementarity with other experiments
- PINGU will be a natural part of the IceCube-Gen2 Observatory
  - Closely based on IceCube technology low technical and cost risk
  - PINGU will use the same hardware as high energy extensions of IceCube –
     common design gives flexibility to optimize based on progress of the field
- Focused here on neutrino physics, but also interesting potential in searches for low mass dark matter and other exotica



### Backup Slides

#### Neutrino-Nucleon Interaction Uncertainties

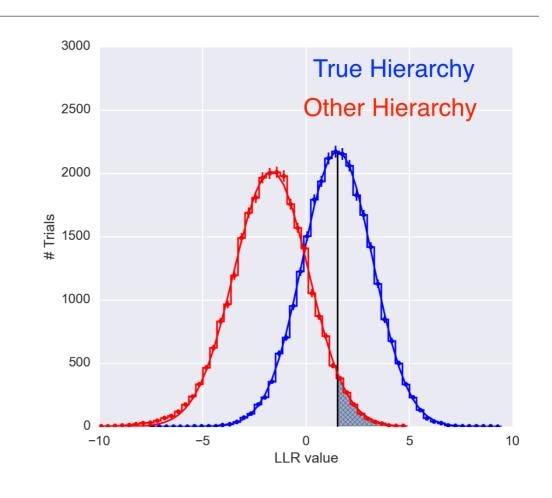
- Comparison of impact of GENIE uncertainties to original ad hoc treatment
- Net impact of full treatment is negligible – oscillation uncertainties dominate
  - Largest impacts from m<sub>A</sub> in CCQE and resonance interactions, higher twist parameters in Bodek-Yang DIS model

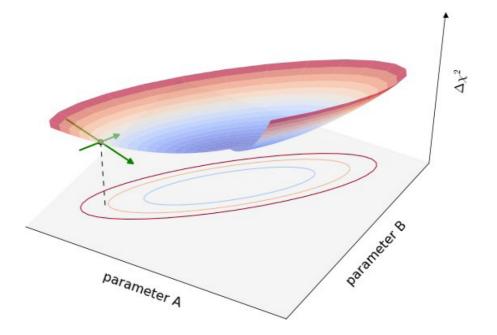




# Estimating Sensitivity to the Mass Hierarchy

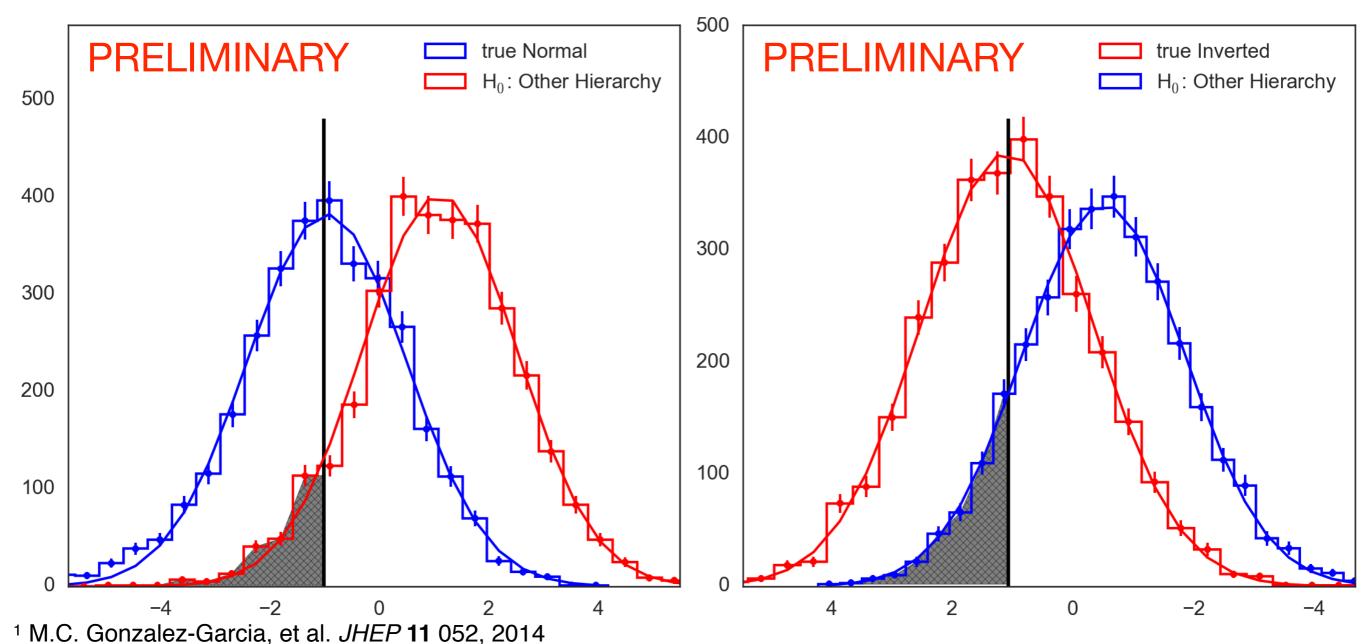
- Comparison of two methods of calculating expected significance
- Log-likelihood ratio method
  - Large ensemble of pseudo-data sets, best-fit physics and nuisance parameters determined numerically
  - Build up distribution of test statistic and integrate tail for expected significance
- Penalized  $\Delta \chi^2$  (pull) method
  - Asimov data set rather than ensemble
  - Linear error propagation for linear parameters, scan over nonlinear ones
  - Fast: analytic minimization of  $\Delta \chi^2$





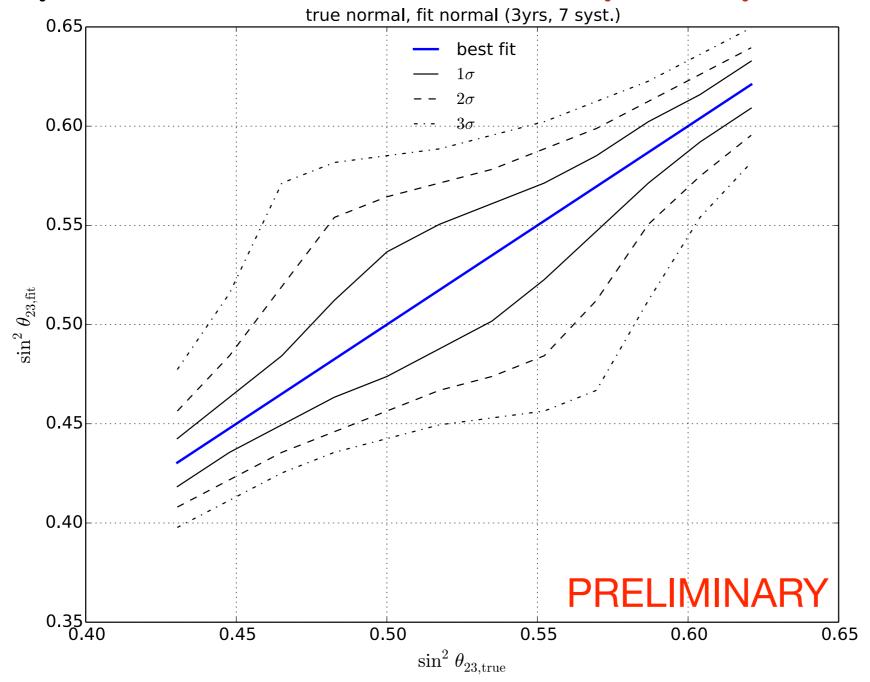


- example: LLR distributions for PINGU for True NH and True IH
  - 1 year significance: 1.83 (NH) and 1.55 (IH) for the NuFit<sup>1</sup> values of oscillation parameters



#### Oscillation Parameters with PINGU

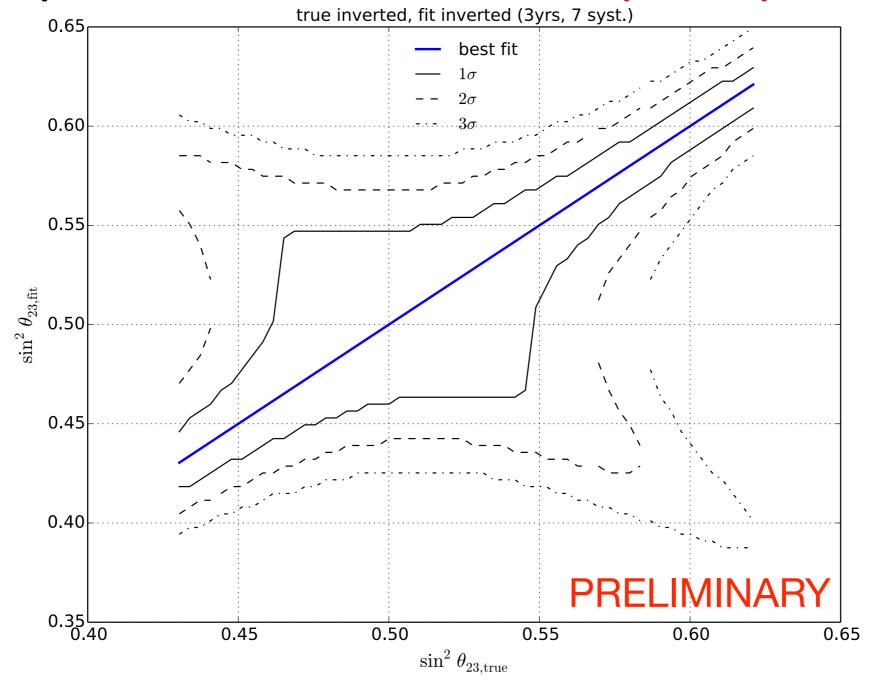
after 3 years of livetime, with normal hierarchy correctly identified



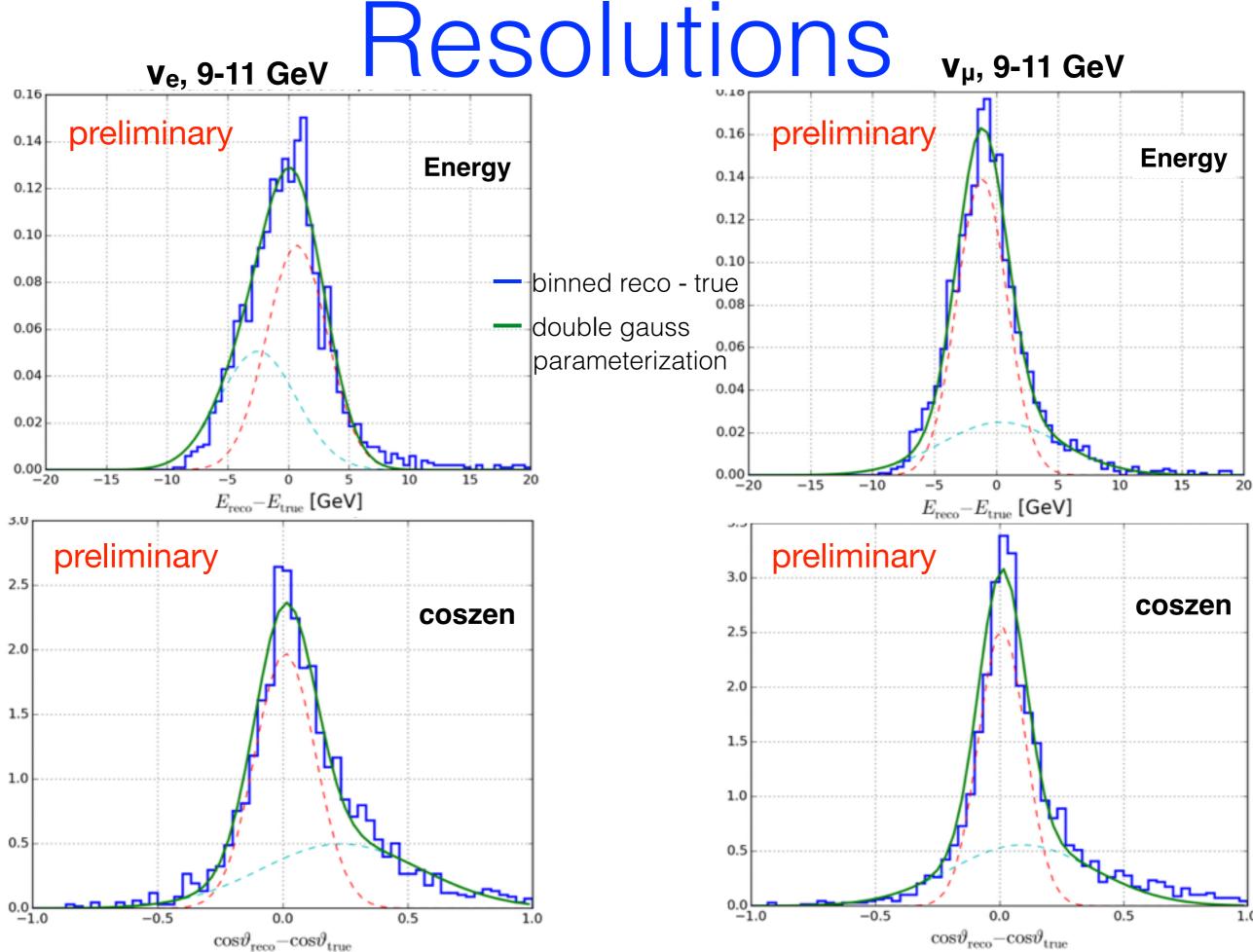


#### Oscillation Parameters with PINGU

after 3 years of livetime, with inverted hierarchy correctly identified







#### Cost and Schedule

 Primary US funding source for IceCube-Gen2 would be NSF

Total cost comparable to IceCube

 Many items common to PINGU and other elements (drill, engineering, etc.)

 Marginal cost of PINGU within larger IceCube-Gen2 is \$88M, with anticipated non-US contributions of \$25M

 Gen2 conceptual design document and PINGU performance update to be released later this year Cost for PINGU Component

Hardware \$48M

Logistics \$23M

Contingency \$16M

Expected non-US \$25M contributions

Total US Cost

\$63M

(elements do not sum to total due to rounding)

