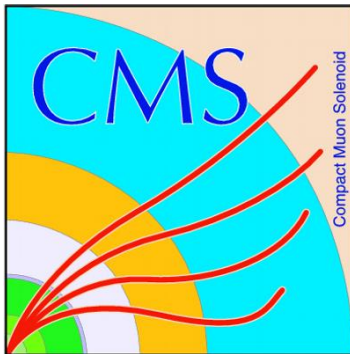


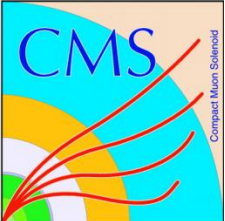
The CMS Beam Halo Monitor detector system



Kelly Stifter, University of Minnesota

08/04/15

2015 DPF Meeting, Ann Arbor



Machine-induced background



CMS Experiment at the LHC, CERN

2009-Nov-20 21:29

Run/ Event: 121964 / 62138

Data taken 2009-Nov-20 21:29:16.241110 GMT

Run_no 121964

Event_no 162138

Lumi_sec 74

Orbit 76711661

Crossing 2607

<http://lhclog.cern.ch/>

L1 Triggers:

L1_SingleMuBeamHalo

Tech

2

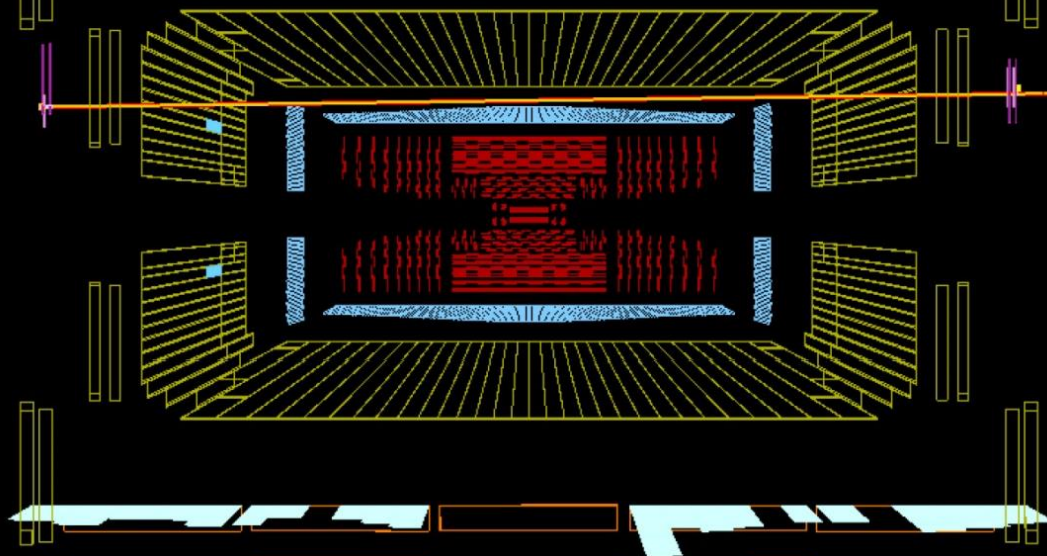
6

60

61

62

63



CMS BHM Group



A. Dabrowski¹, M. Giunta¹, R. Loos¹, S. Orfanelli^{1,2}, D. Stickland³, F. Fabbri⁴, A. Manna⁴, A. Montanari⁴, N. Tosi⁴, M. Ambrose⁵, J. Mans⁵, R. Rusack⁵, K. Stifter⁵



1. CERN,
Switzerland



2. National
Technical
University of
Athens, Greece



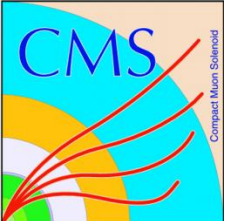
3. Princeton
University,
USA



4. INFN
Bologna, Italy



5. University
of Minnesota,
USA

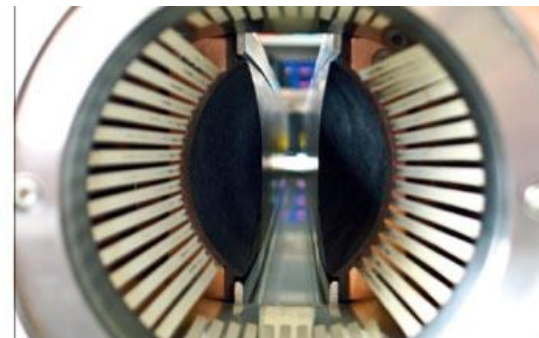
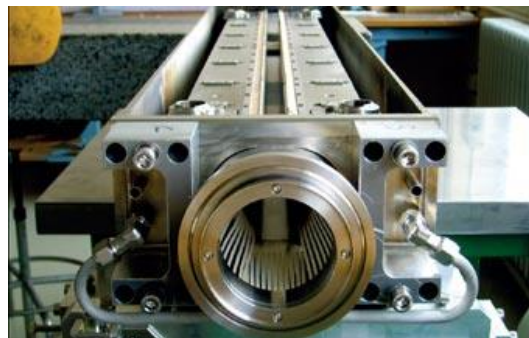
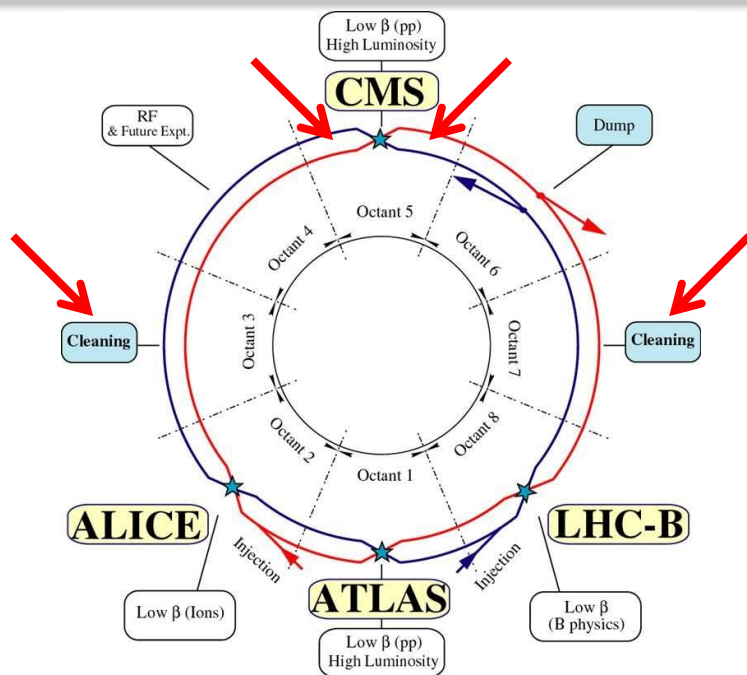


Outline

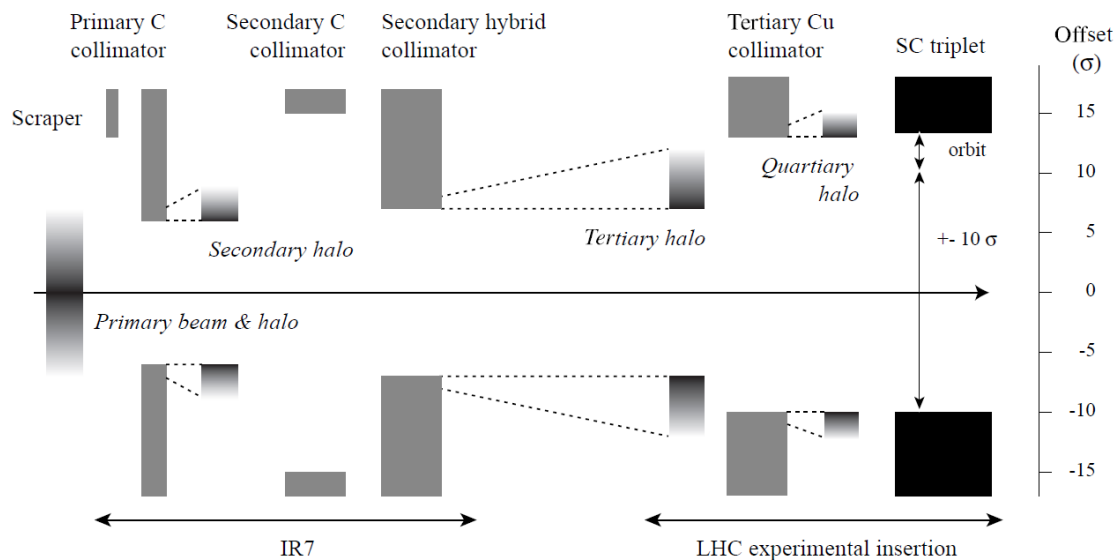


1. Machine-induced background
2. Beam Halo Monitor detector
3. Detector system
 - Detector units
 - Electronics
4. First results
 - Splashes
 - Correlation to collimator movement
 - Commissioning data

Machine-induced background (MIB)

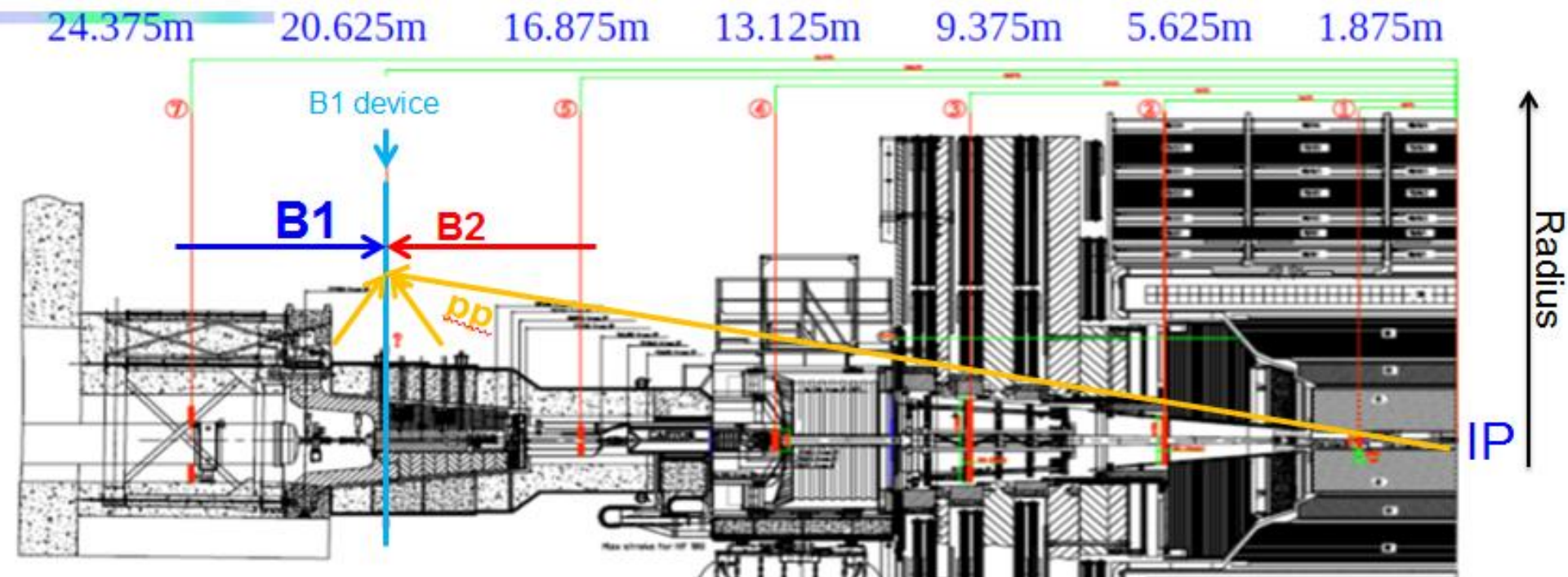


- Interactions with collimators
- Beam gas
- UFOs



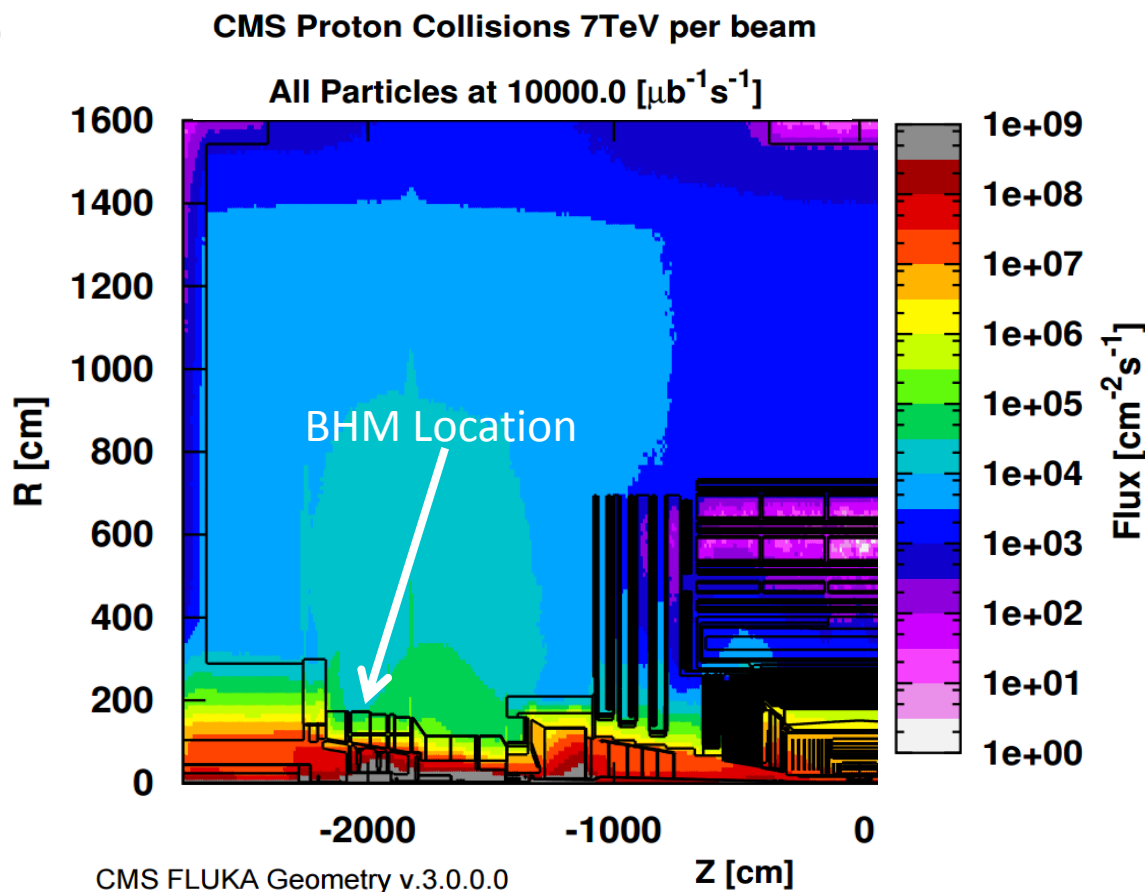
Beam Halo Monitor

- Purpose:** Provide an online, bunch-by-bunch, per beam MIB rate arriving at CMS at high radius



Problem

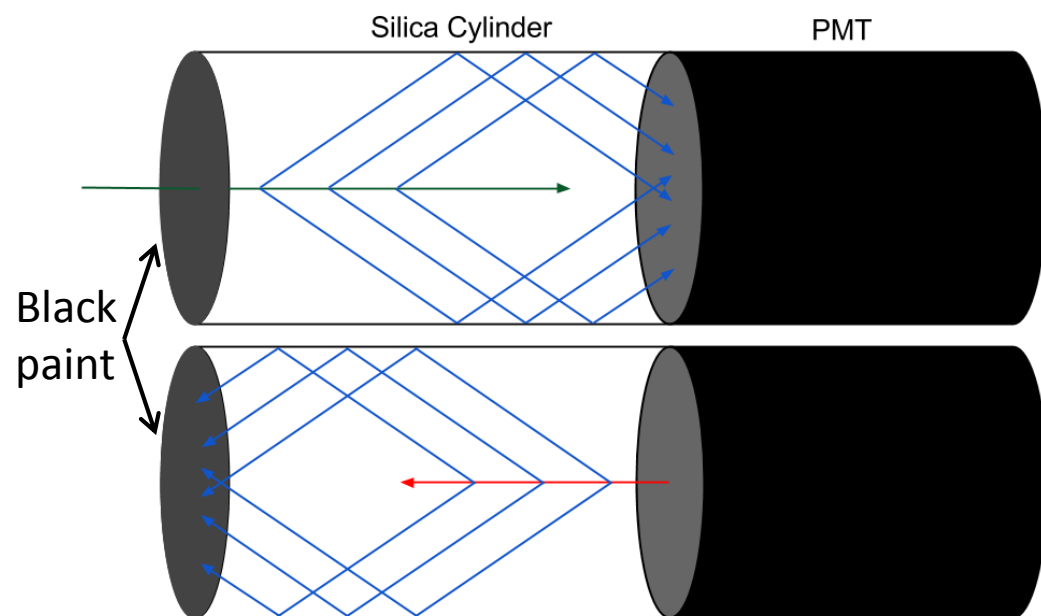
- Many more collision products ($\sim 10^4 \text{ Hz/cm}^2$) than MIB particles ($\sim 1 \text{ Hz/cm}^2$)



- Must find MIB signal amidst collision products \rightarrow factor of $\geq 10^4$ suppression required
- Solutions: **DIRECTION-SENSITIVITY** and **PRECISION TIMING**

Direction sensitivity

MIB muon: Arrives with incoming beam.
Cherenkov radiation is seen by PMT.



Cherenkov radiation:

- ✓ Insensitive to neutron and gamma backgrounds
- ✓ Prompt signal in time with incoming particle
- ✓ Use quartz, radiation hard & UV transmissive
- ✓ Large signals – ~ 60 p.e./cm for forward particle

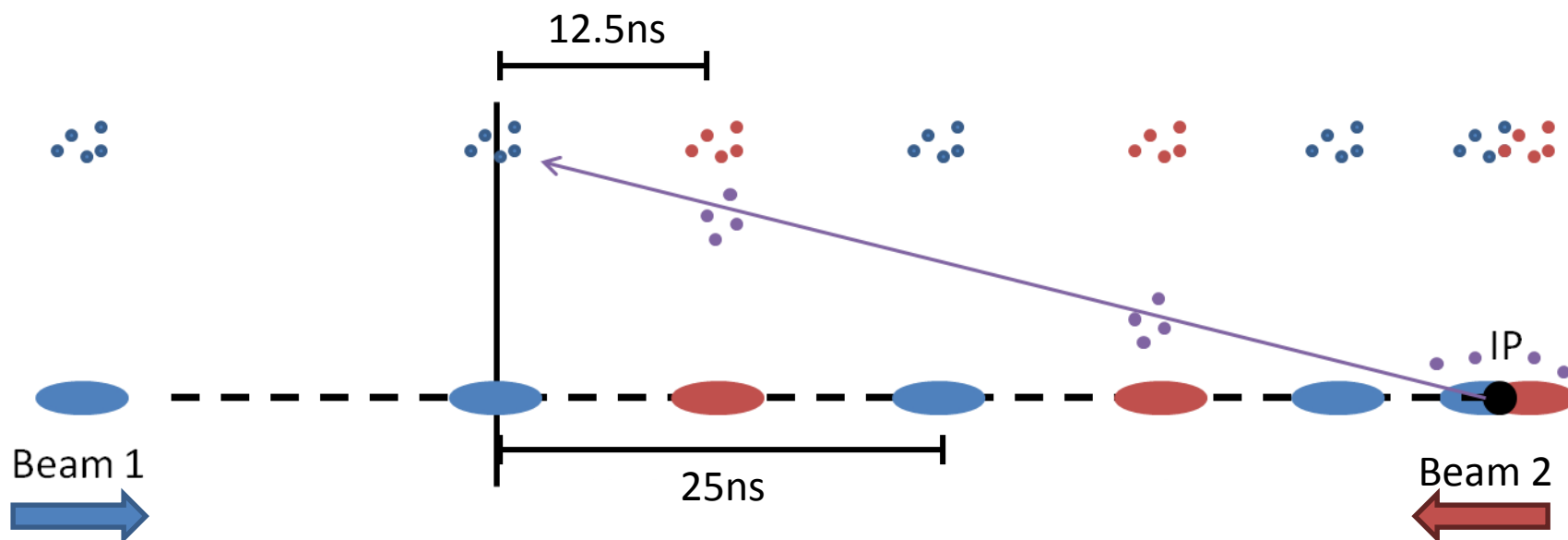
Collision product: Arrives in opposite direction. Cherenkov radiation is absorbed by black paint.

2014 DESY test beam:
Background rejection of $>99.99\%$
With forward acceptance of $>98\%$

Timing

- Golden locations allow for maximum separation in timing

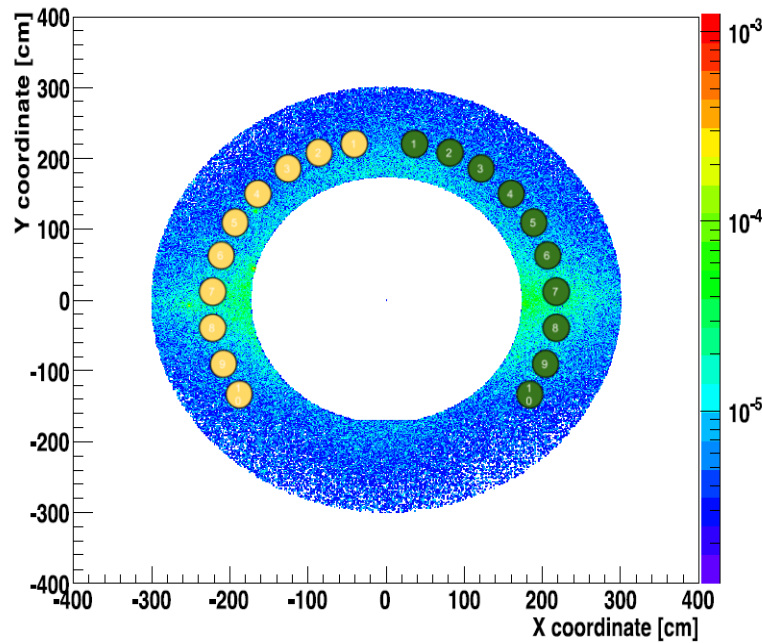
$$GL_{k+1} = \left(\frac{1}{4} + \frac{1}{2}k \right) \cdot (\text{BX spacing}) \approx 1.875m + 3.75m \cdot k$$



Detector placement

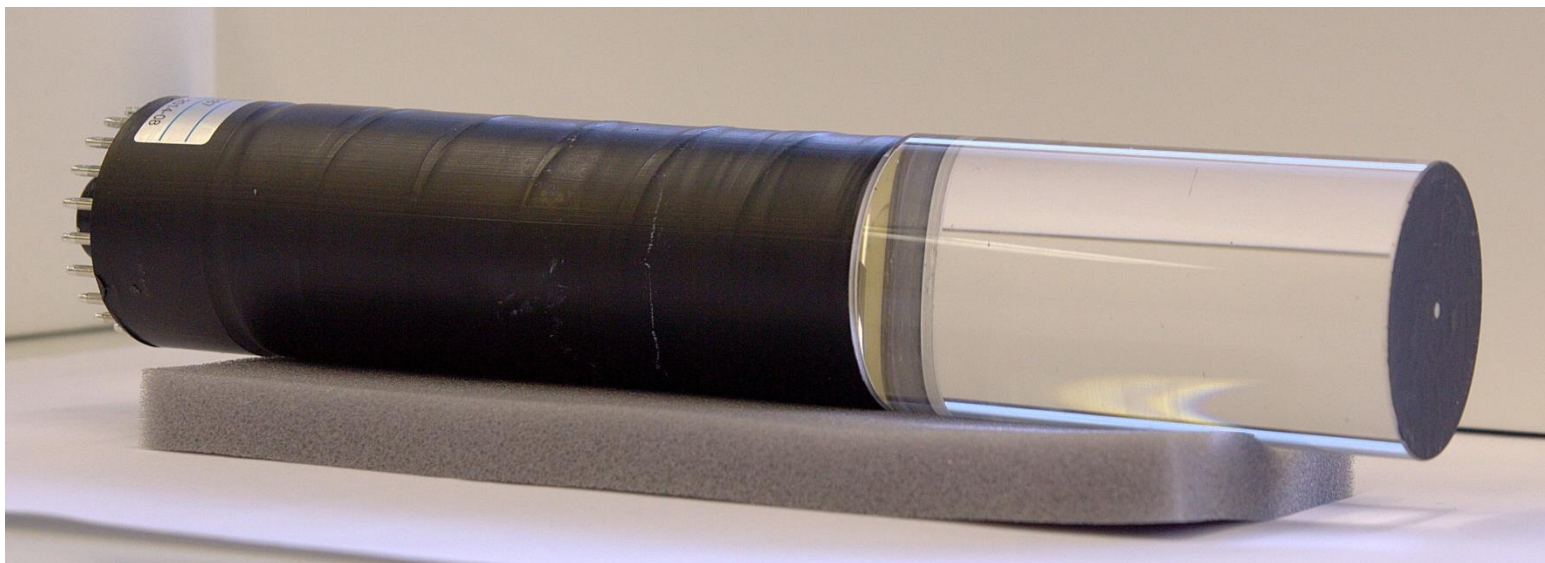


- Golden location 6 – 20.6m from IP
- 40 detectors, 20 at each end
- Acceptance of $21.2 \text{ cm}^2/\text{unit}$, $424 \text{ cm}^2/\text{beam}$
- Installed at radius of 1.8m from beam
- Placed in ϕ -region of highest flux



Detector units

- SQ0 synthetic fused silica: 10 cm long, 5.2 cm diameter, UV transmissive, radiation hard
- Optically coupled to UV sensitive Hamamatsu R2059 PMT
- 3 layers of magnetic shielding: Permalloy, mumetal, iron

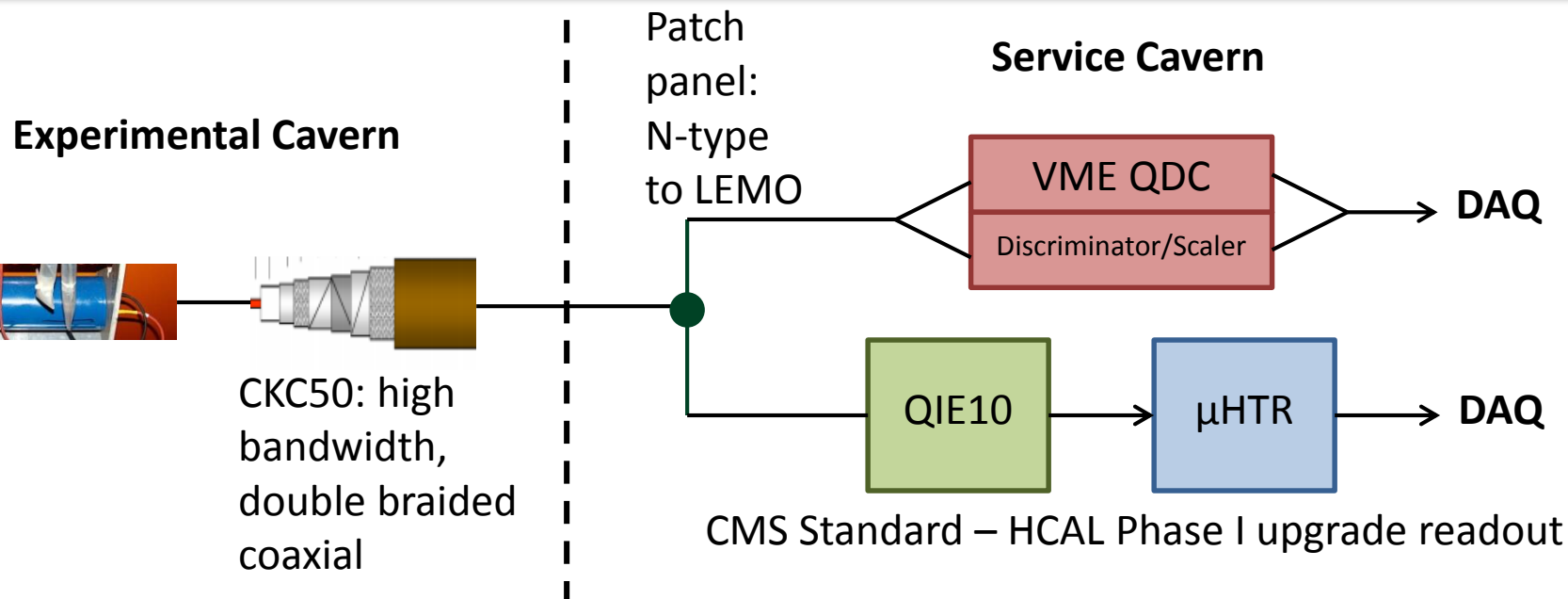


Detector units

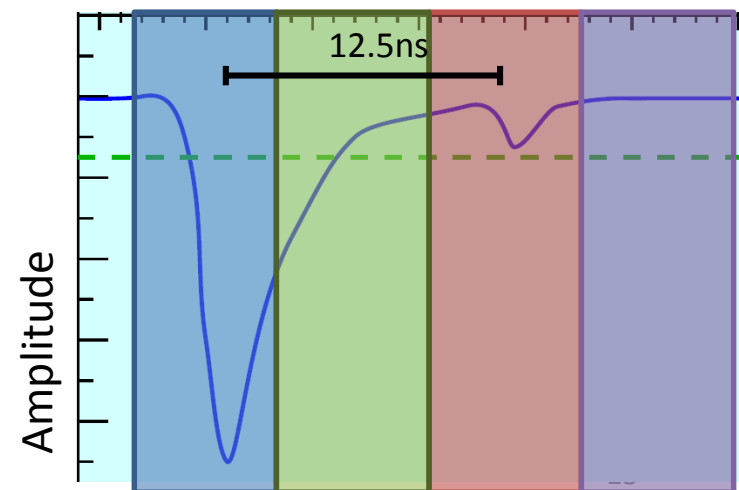
- SQ0 synthetic fused silica: 10 cm long, 5.2 cm diameter, UV transmissive, radiation hard
- Optically coupled to UV sensitive Hamamatsu R2059 PMT
- 3 layers of magnetic shielding: Permalloy, mumetal, iron



Read-out chain

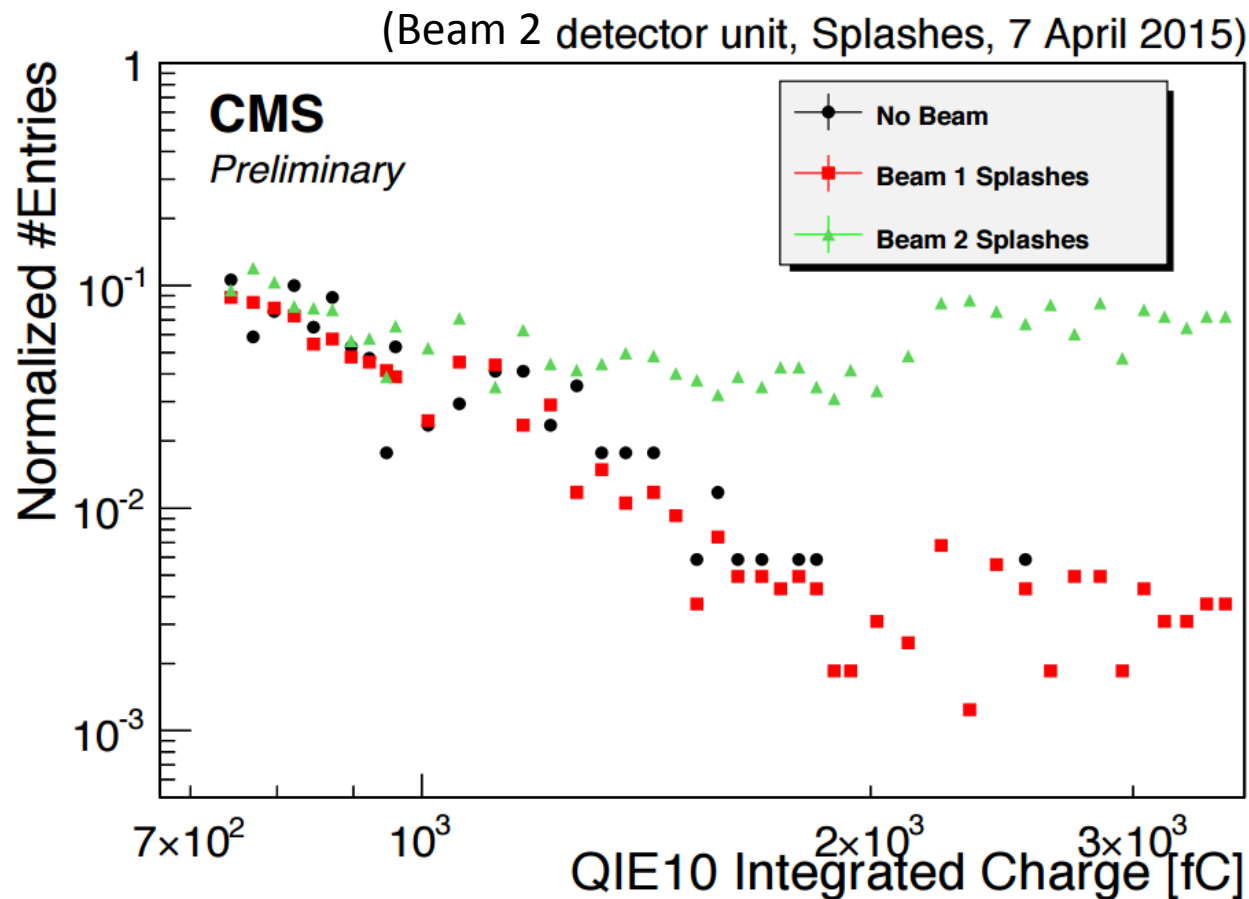


- Every BX, every detector → get 8-bit charge and 6-bit timing information
 - Histogram (μ HTR)
- Calculate flux
 - Published to CMS and LHC operations every ~ 23 s



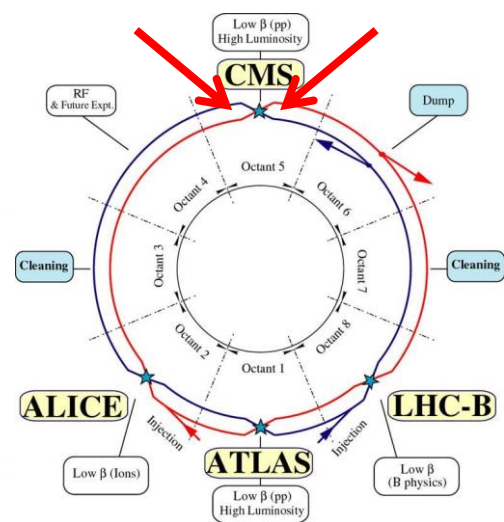
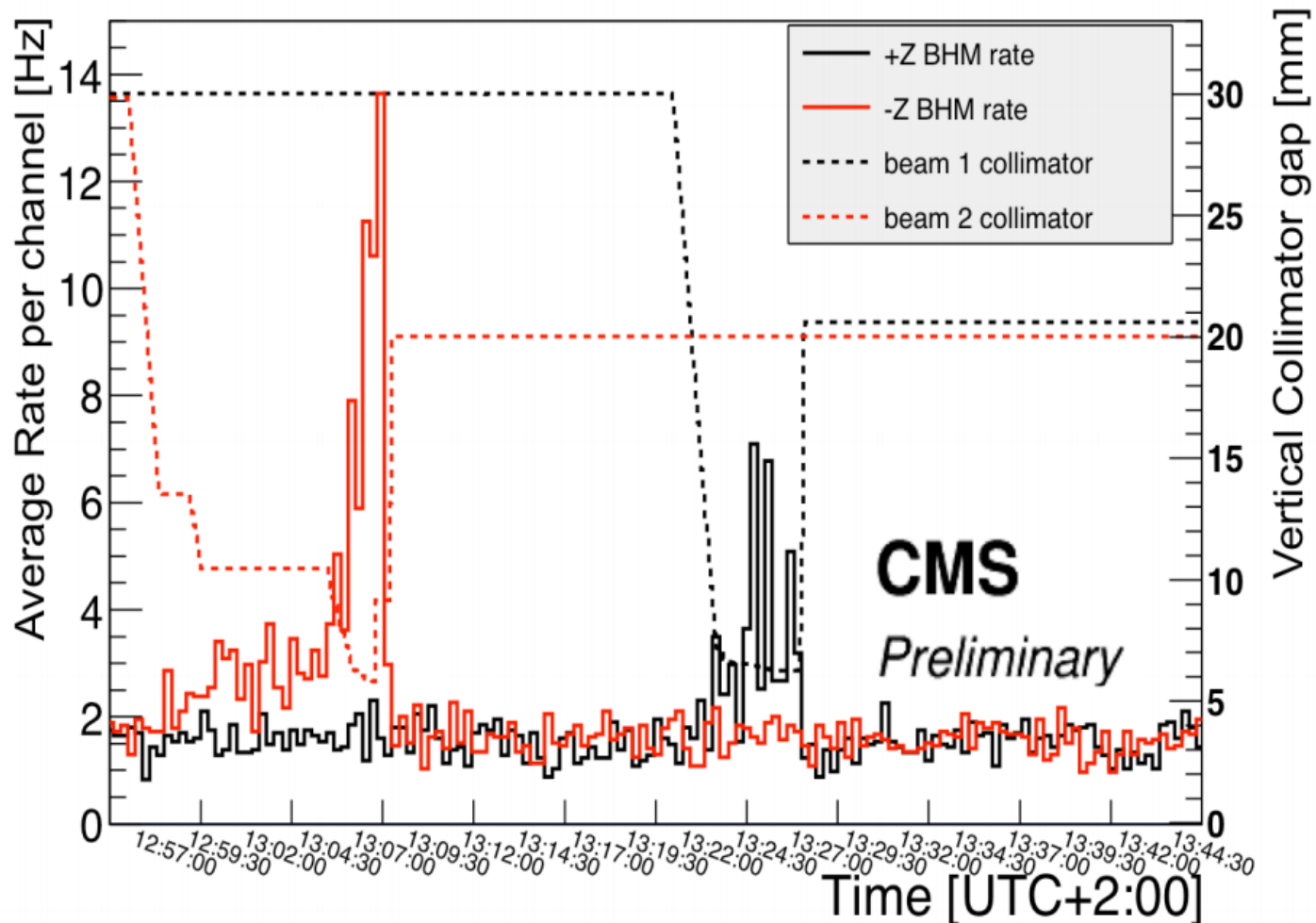
Splashes

- **Indication of directionality**: a unit measuring beam 2 sees only beam 2 splashes, no beam 1 splashes



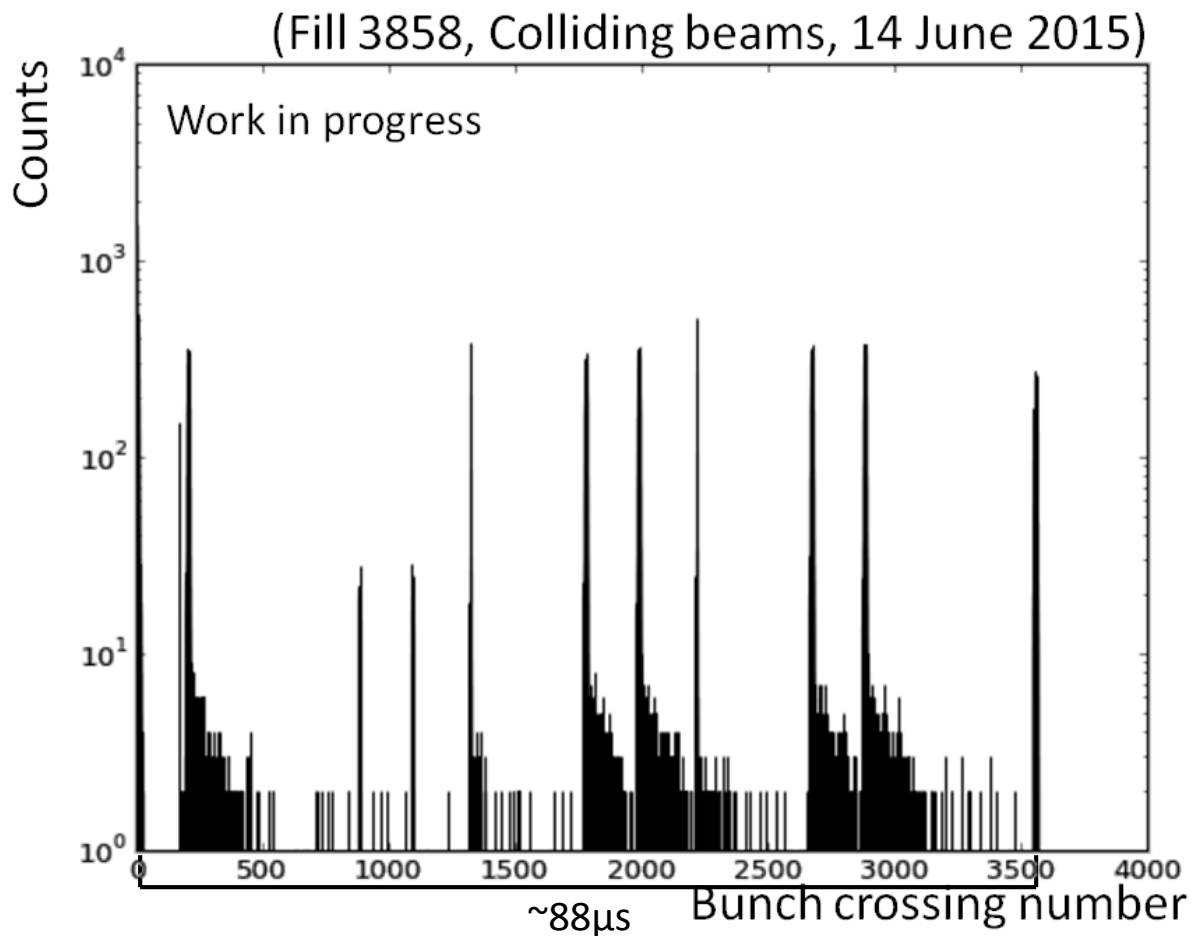
Correlation to collimator movement

(Fill 3679, Circulating Beams, 5 May 2015)



Detector commissioning

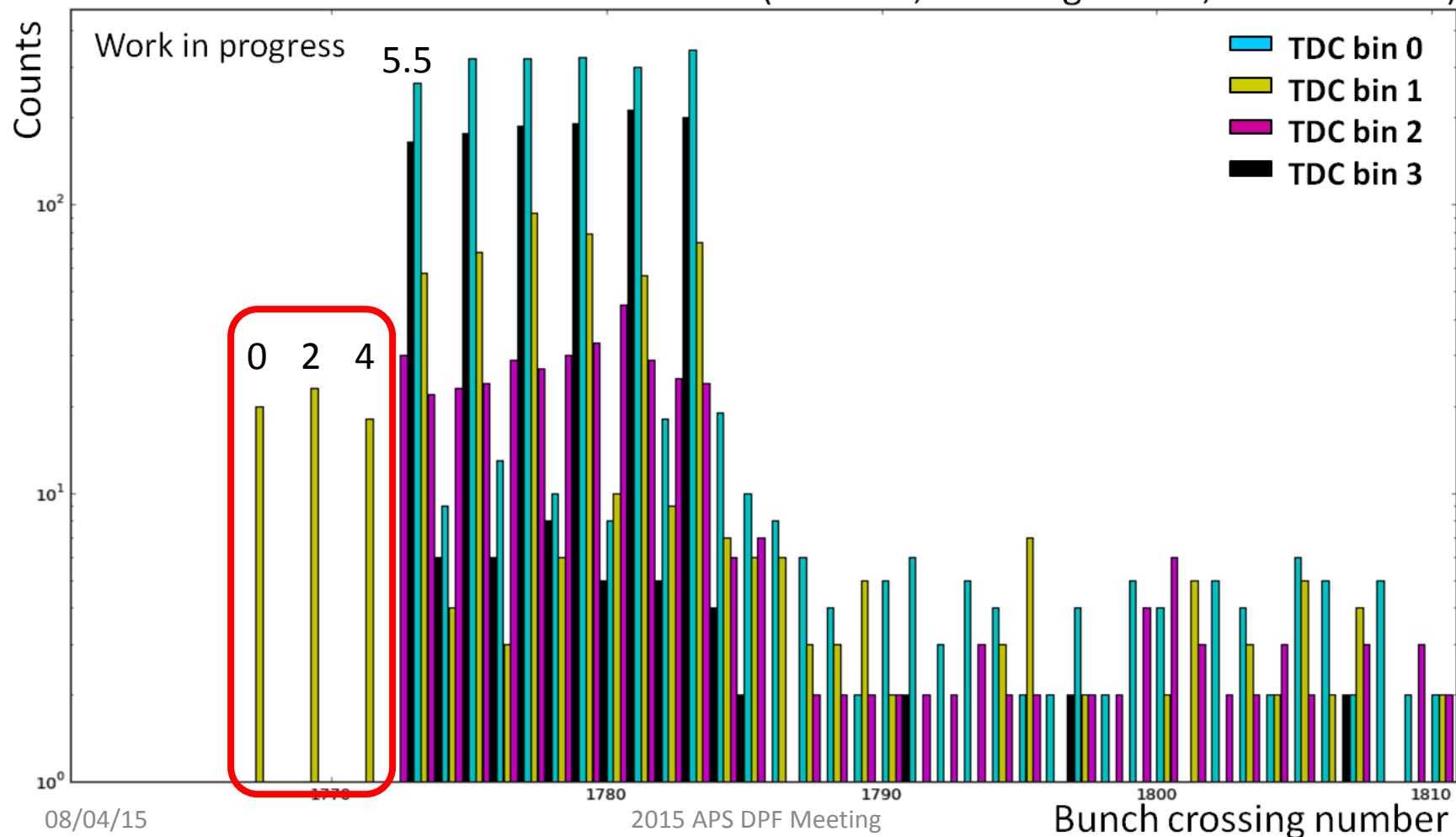
- Commissioned with low detector thresholds
- Detailed bunch structure seen
- Tail consistent with albedo from cavern



Detector commissioning

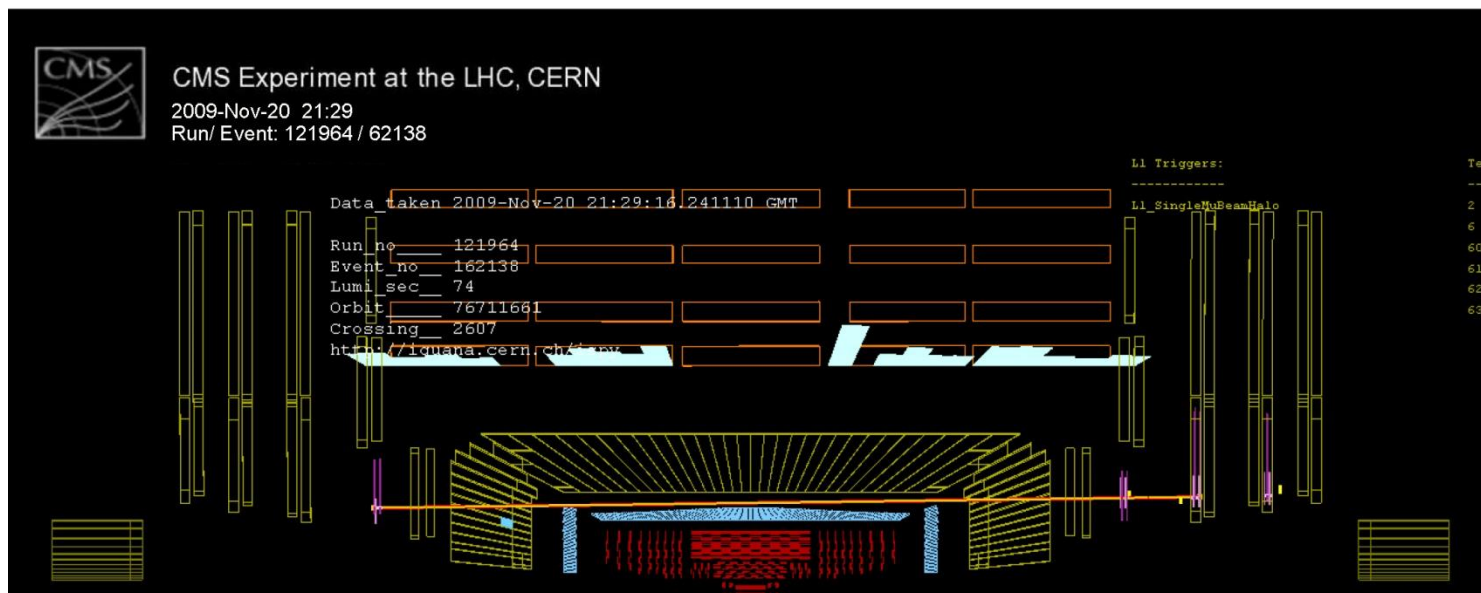
- Three bunches of pure background

(Fill 3858, Colliding beams, 14 June 2015)



Summary

- New Beam Halo Monitor (BHM) will provide an online, bunch-by-bunch, per beam MIB rate arriving at CMS at high radius
- Takes advantage of directional nature of Cherenkov radiation and golden location timing to separate MIB from collisions product signals

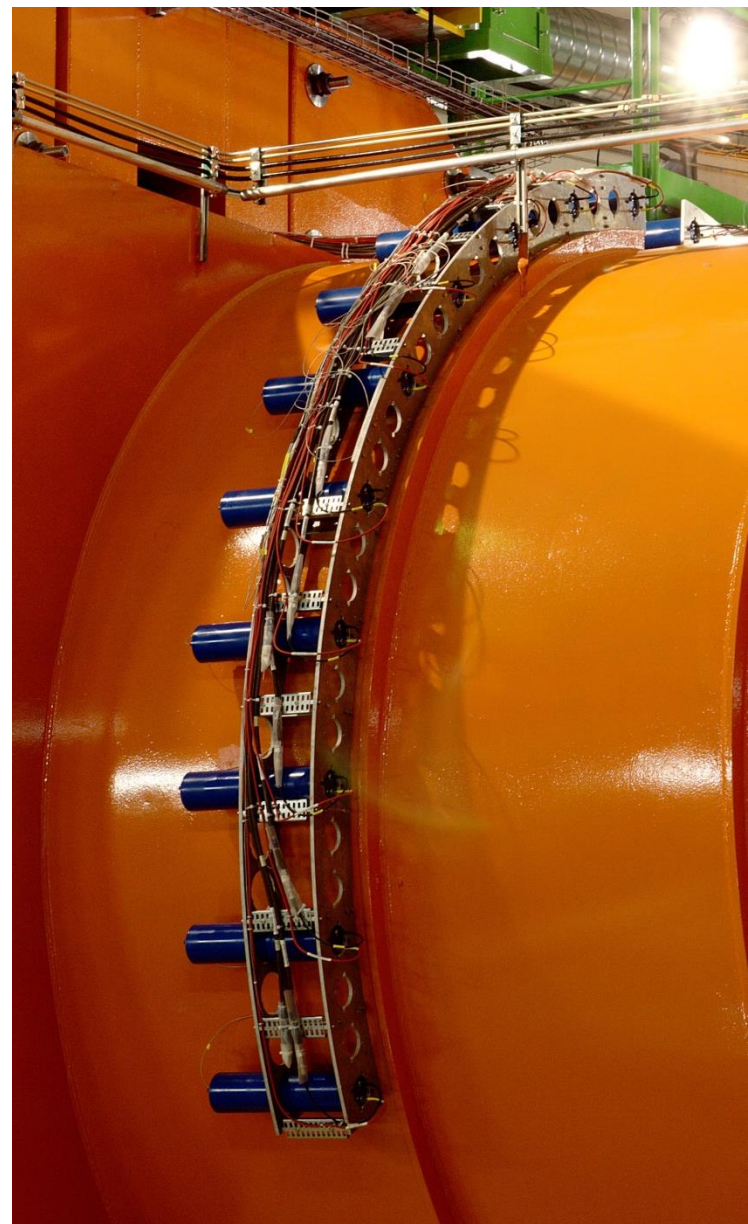


Acknowledgements

- University of Minnesota: R. Rusack, J. Mans
- BRIL: A. Dabrowski
- BHM: N. Tosi, S. Orfanelli
- Other UMN, CERN, and Bologna students and technicians



ADDITIONAL SLIDES



The CMS detector



CMS DETECTOR

Total weight : 14,000 tonnes
 Overall diameter : 15.0 m
 Overall length : 28.7 m
 Magnetic field : 3.8 T

STEEL RETURN YOKE
 12,500 tonnes

SILICON TRACKERS
 Pixel ($100 \times 150 \mu\text{m}$) $\sim 16\text{m}^2 \sim 66\text{M}$ channels
 Microstrips ($80 \times 180 \mu\text{m}$) $\sim 200\text{m}^2 \sim 9.6\text{M}$ channels

SUPERCONDUCTING SOLENOID
 Niobium titanium coil carrying $\sim 18,000\text{A}$

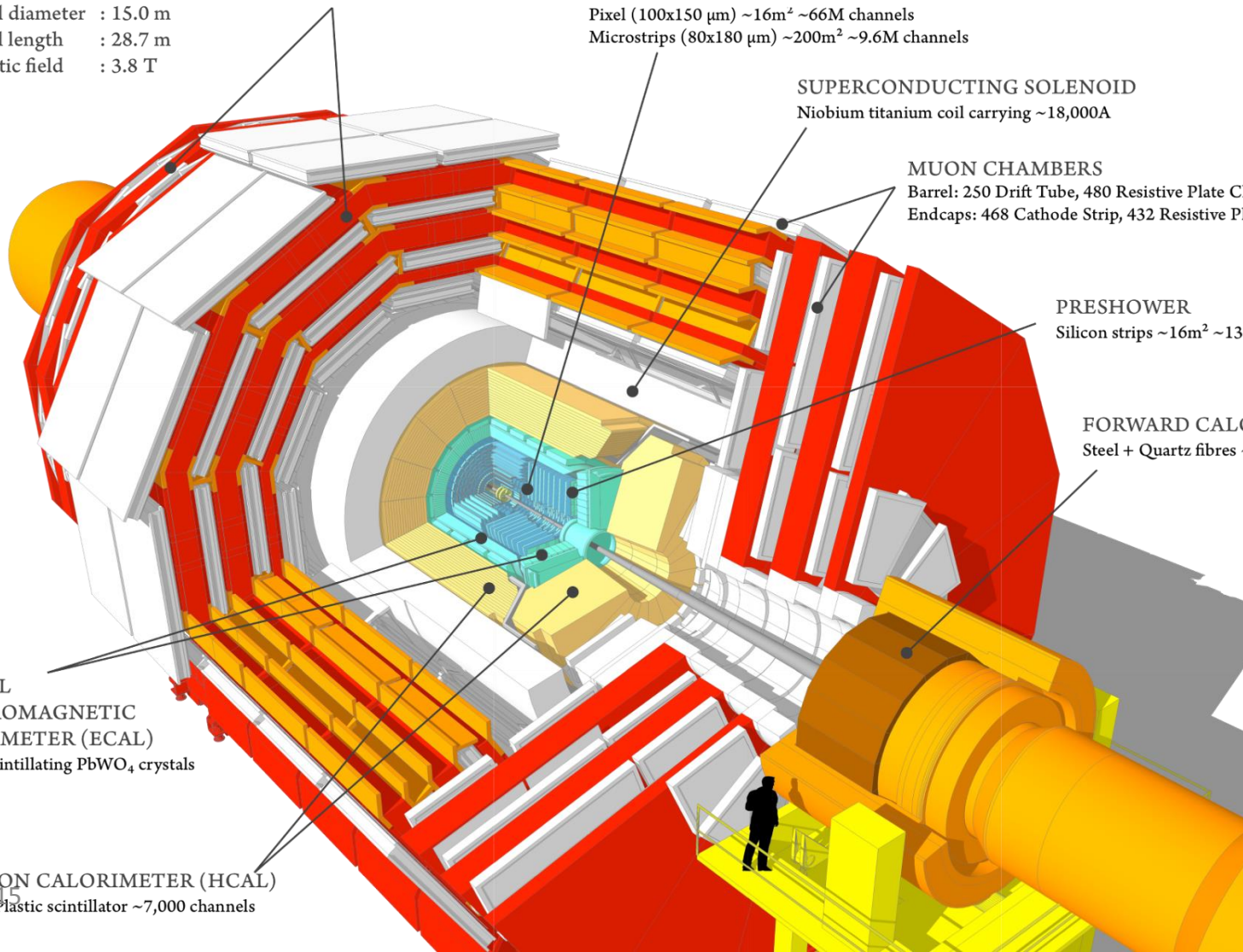
MUON CHAMBERS
 Barrel: 250 Drift Tube, 480 Resistive Plate Chambers
 Endcaps: 468 Cathode Strip, 432 Resistive Plate Chambers

PRESHOWER
 Silicon strips $\sim 16\text{m}^2 \sim 137,000$ channels

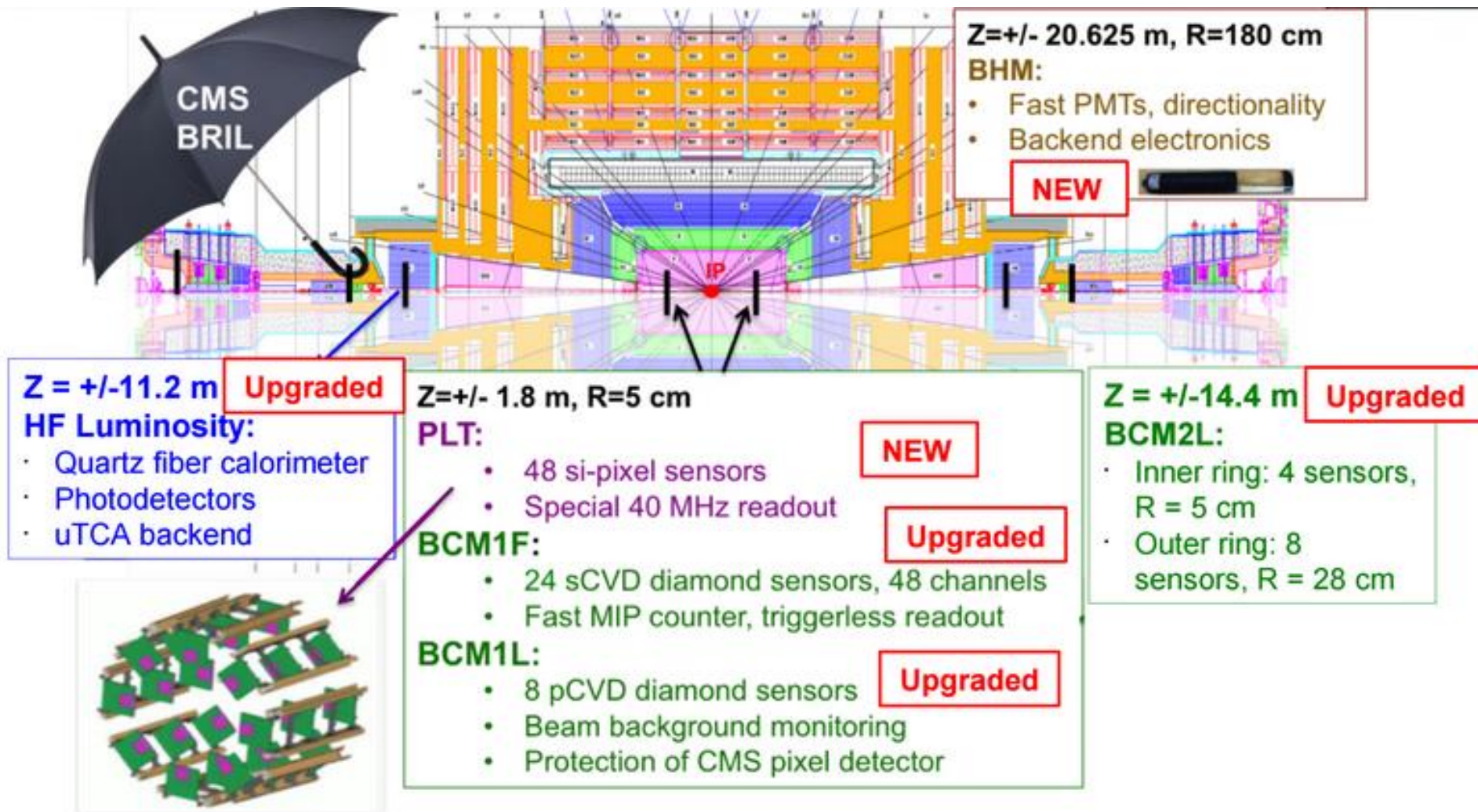
FORWARD CALORIMETER
 Steel + Quartz fibres $\sim 2,000$ Channels

CRYSTAL
 ELECTROMAGNETIC
 CALORIMETER (ECAL)
 $\sim 76,000$ scintillating PbWO_4 crystals

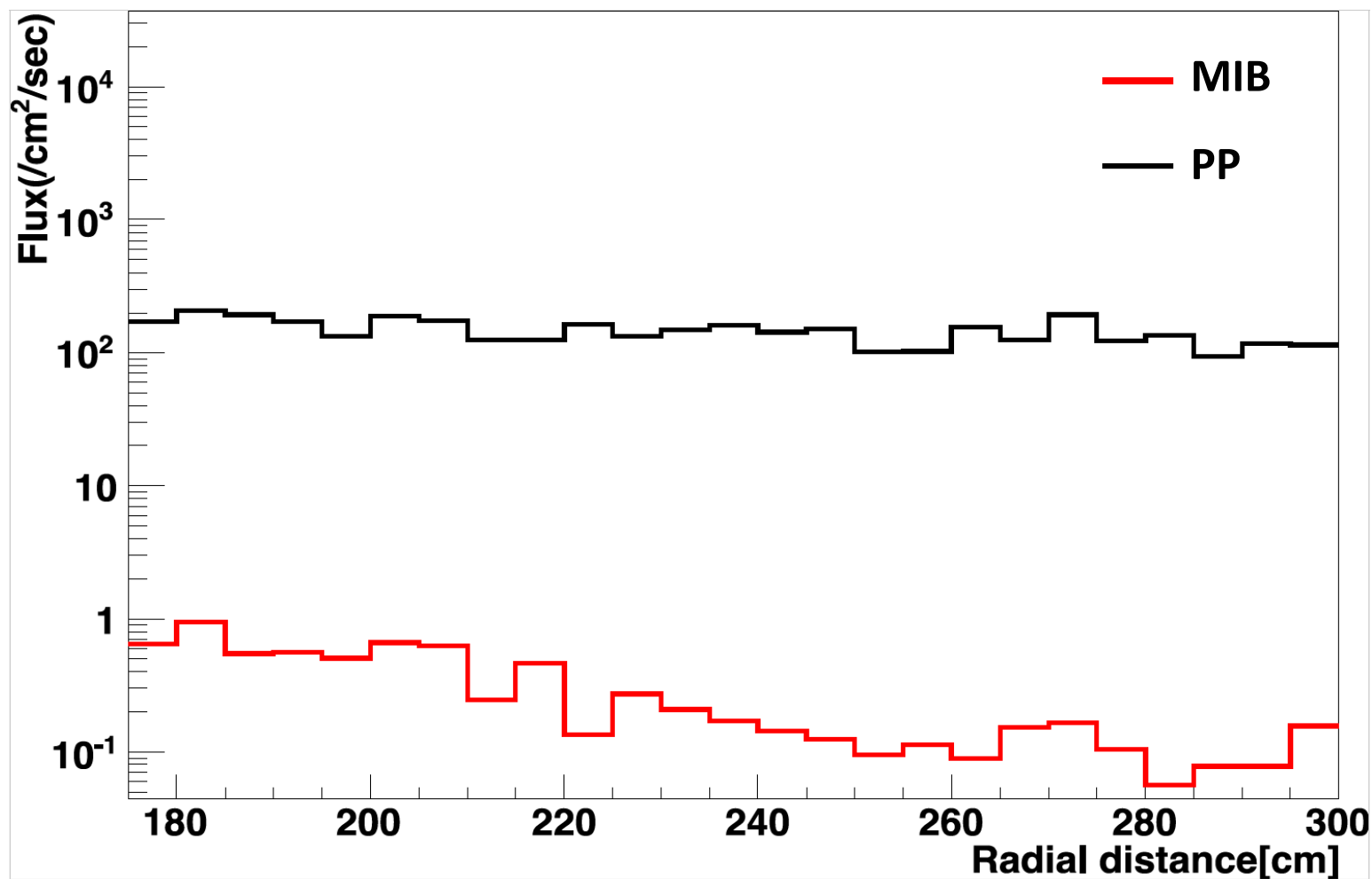
HADRON CALORIMETER (HCAL)
 Brass + Plastic scintillator $\sim 7,000$ channels



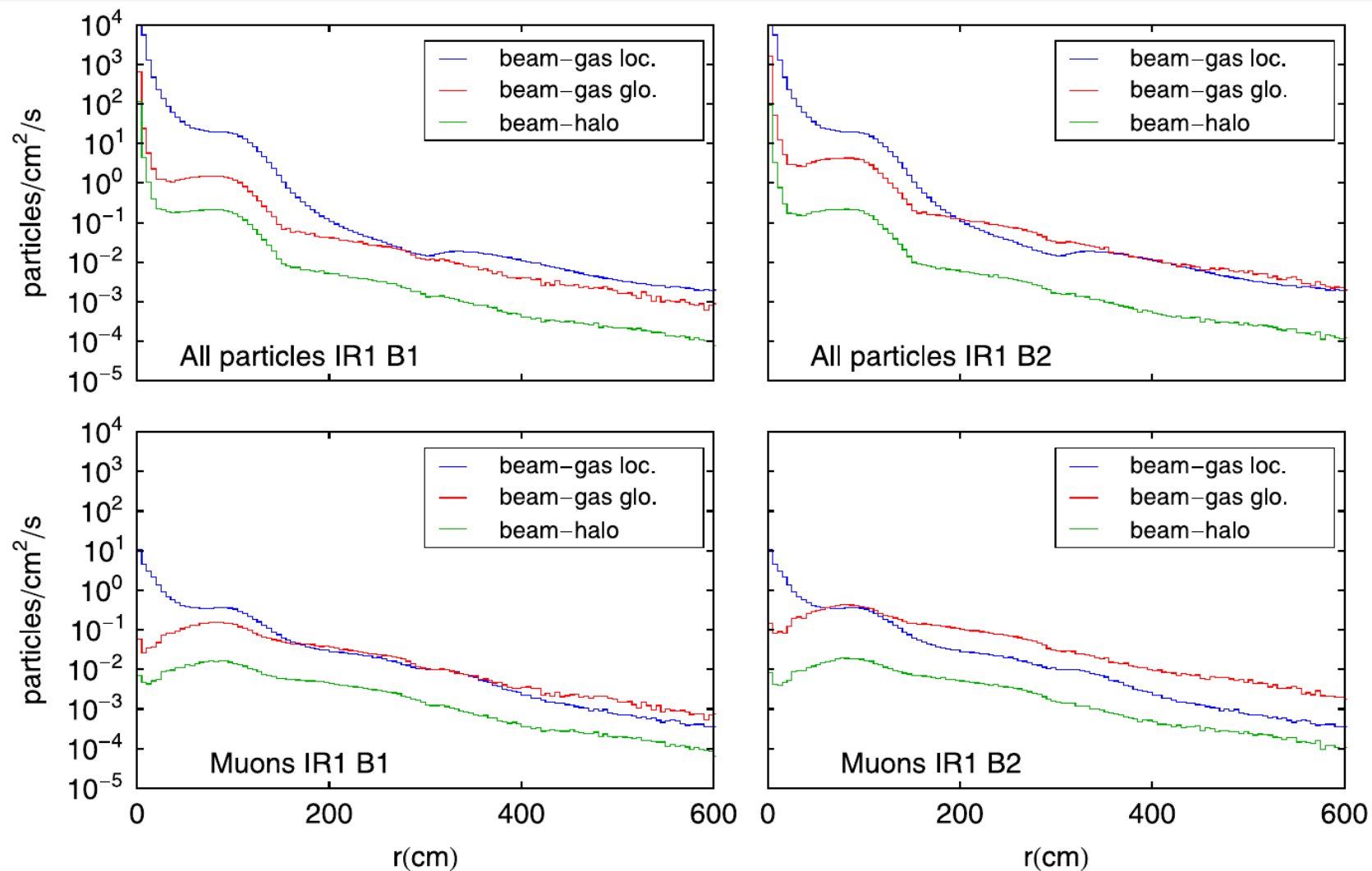
08/04/15



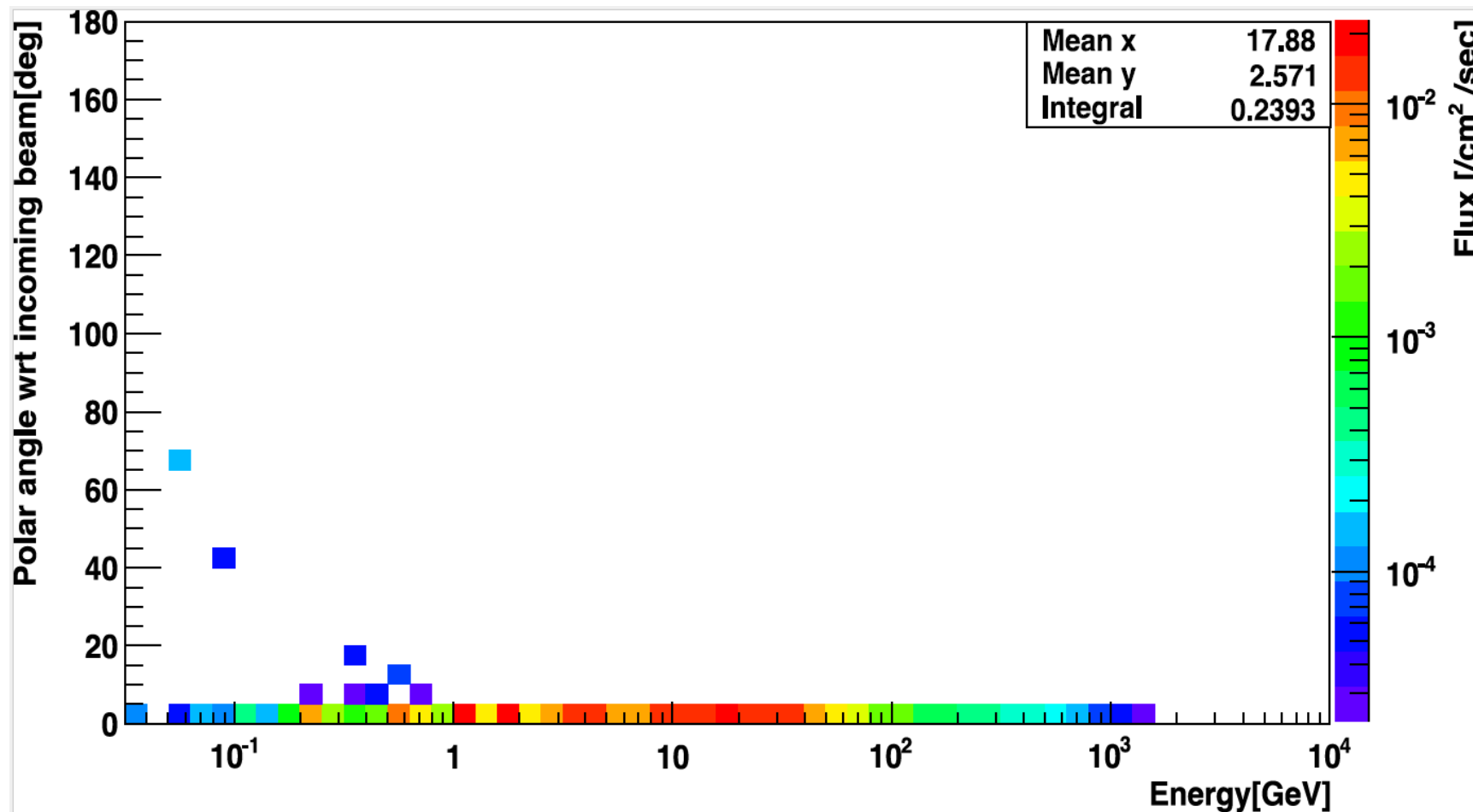
MIB vs. PP



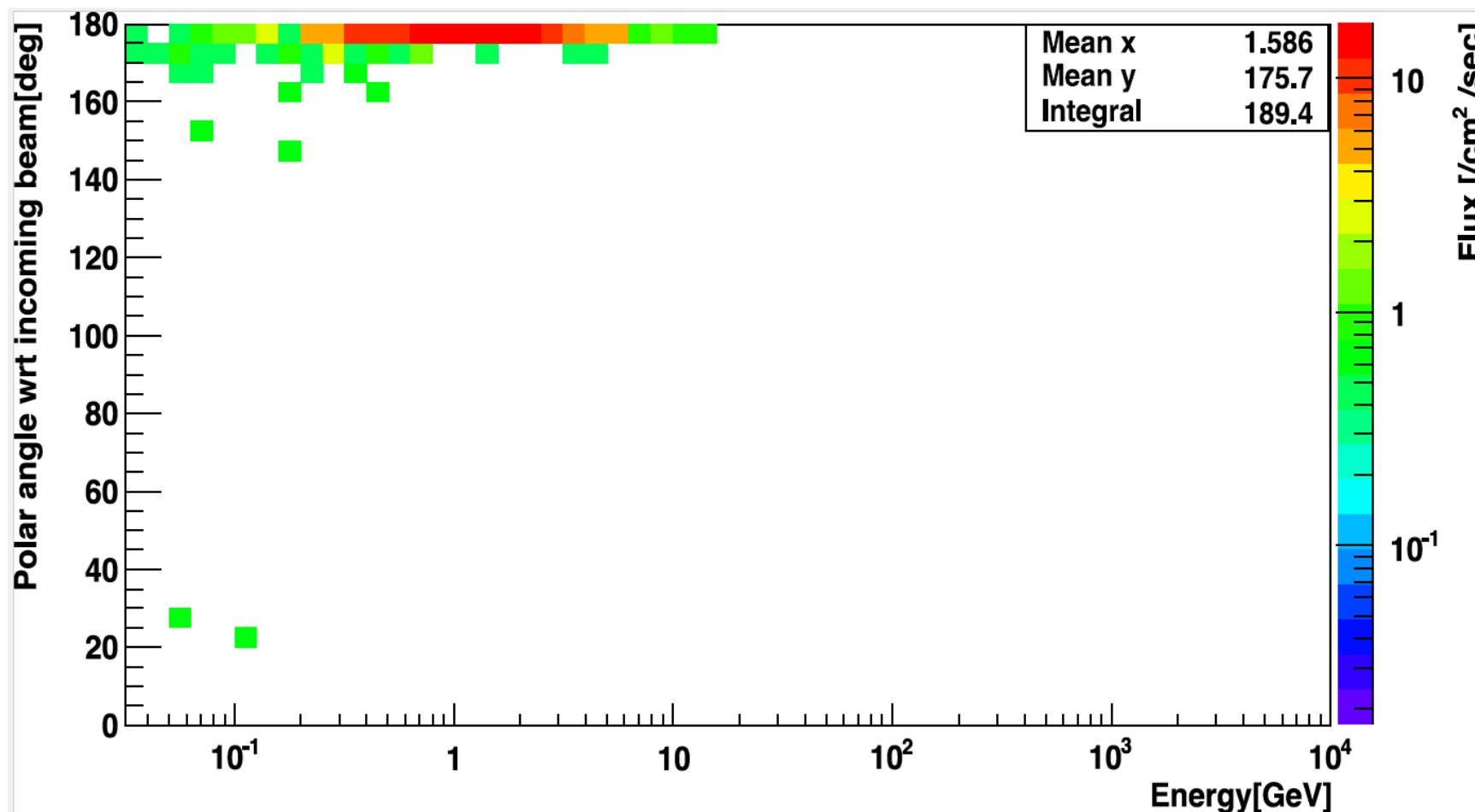
Beam halo contributions



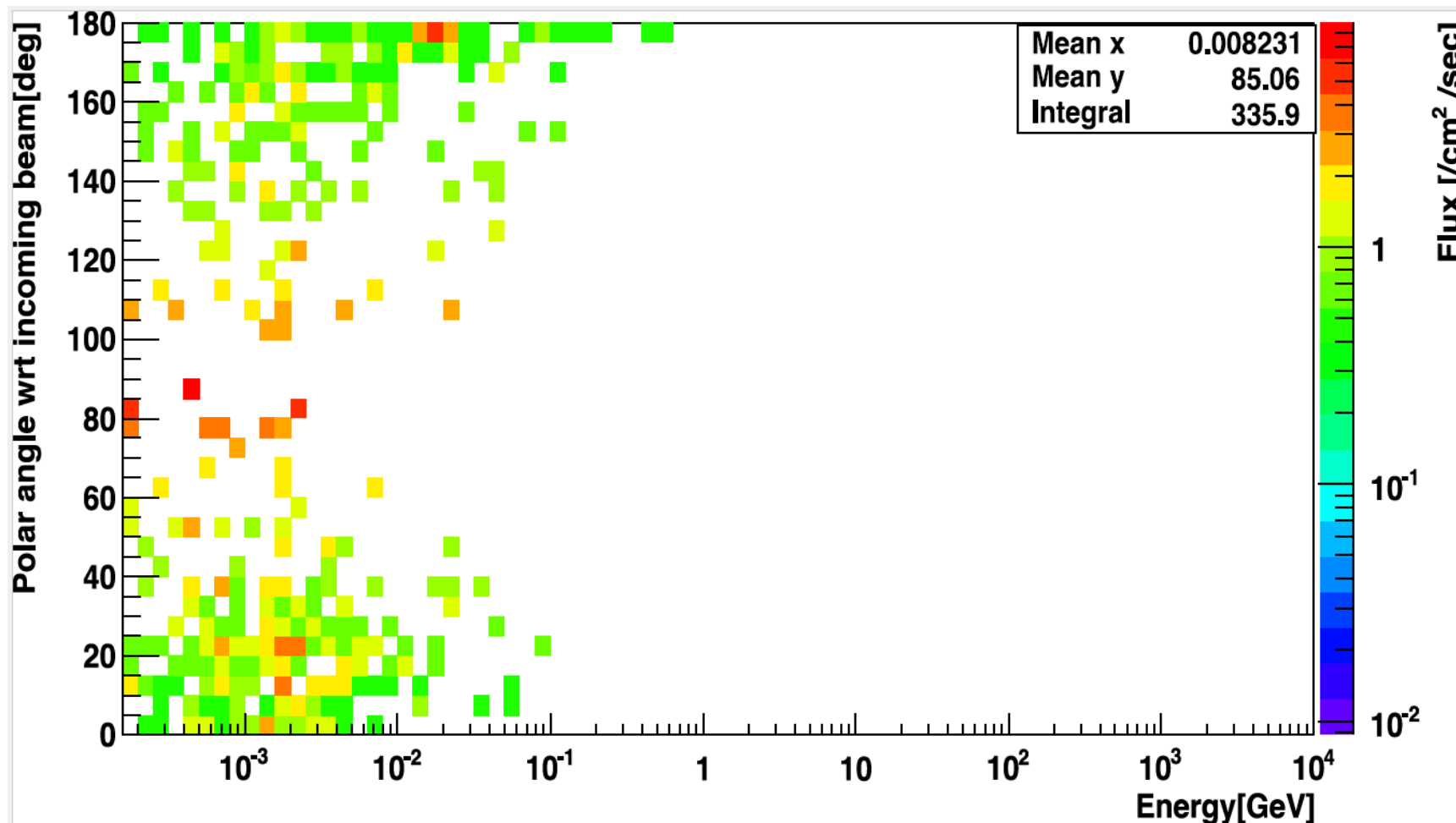
MIB muons



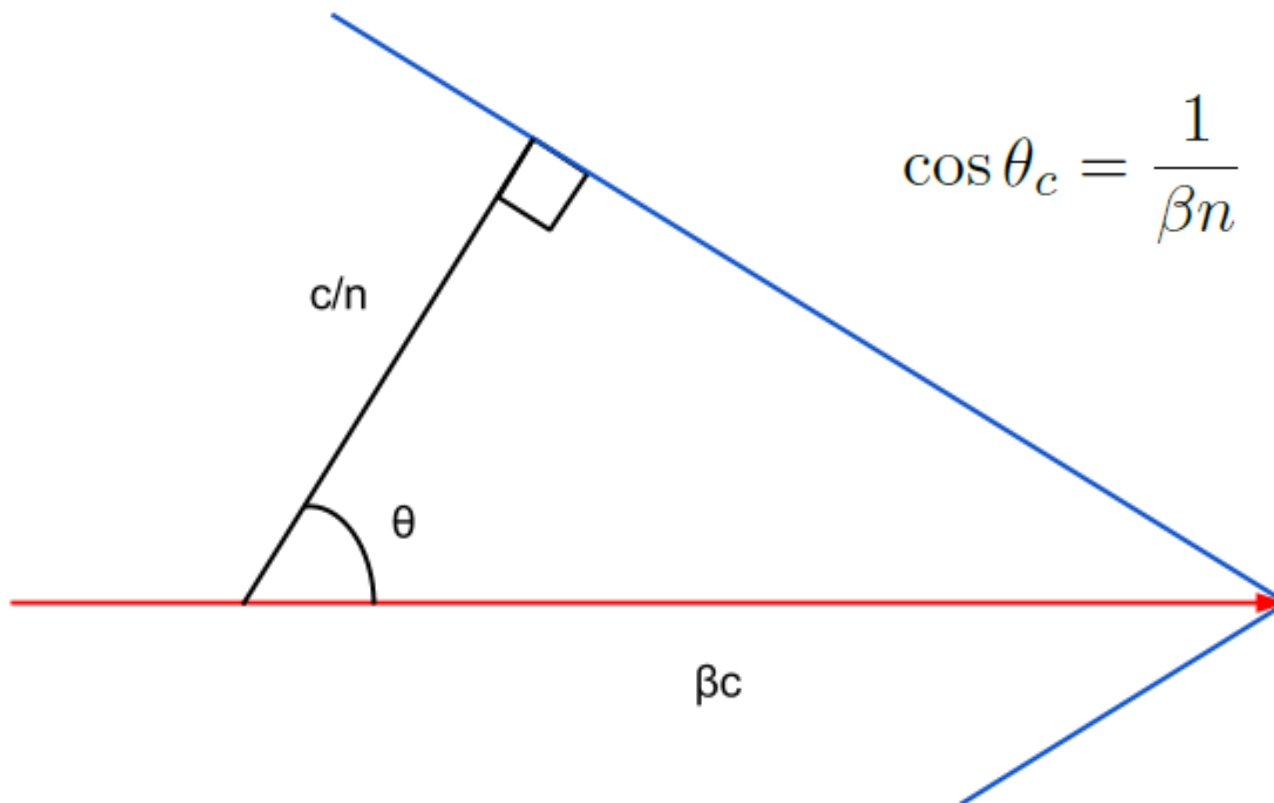
PP muons



PP electrons



Cherenkov angle



$$E_{th} = \frac{nmc^2}{\sqrt{n^2 - 1}} \approx 142 \text{ MeV for muons}$$

Cherenkov radiation

$$\frac{\partial^2 N}{\partial x \partial \lambda} = \frac{2\pi\alpha}{\lambda^2} \left(1 - \frac{1}{\beta^2 n(\lambda)^2}\right)$$

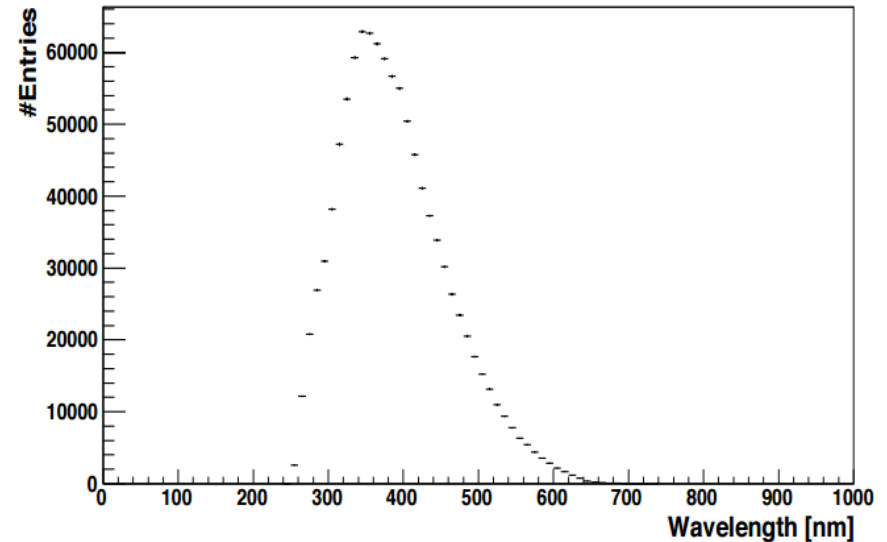
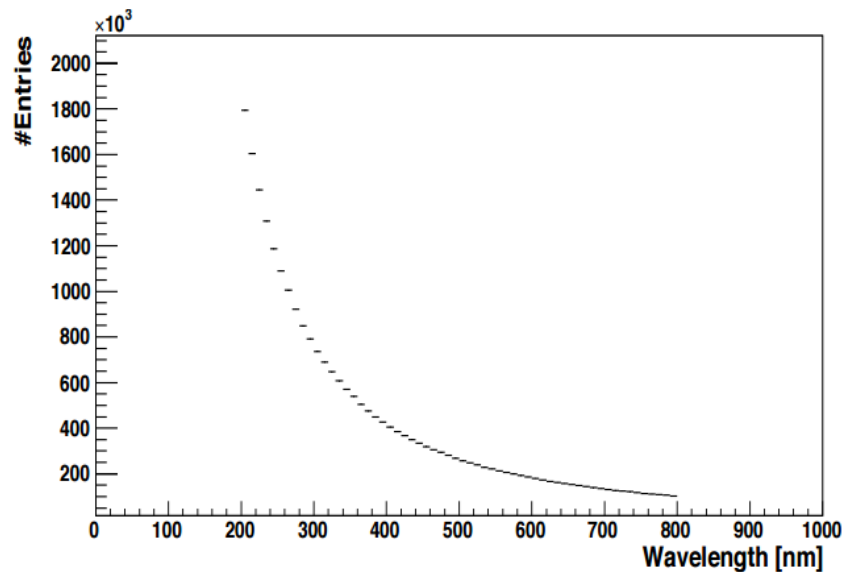
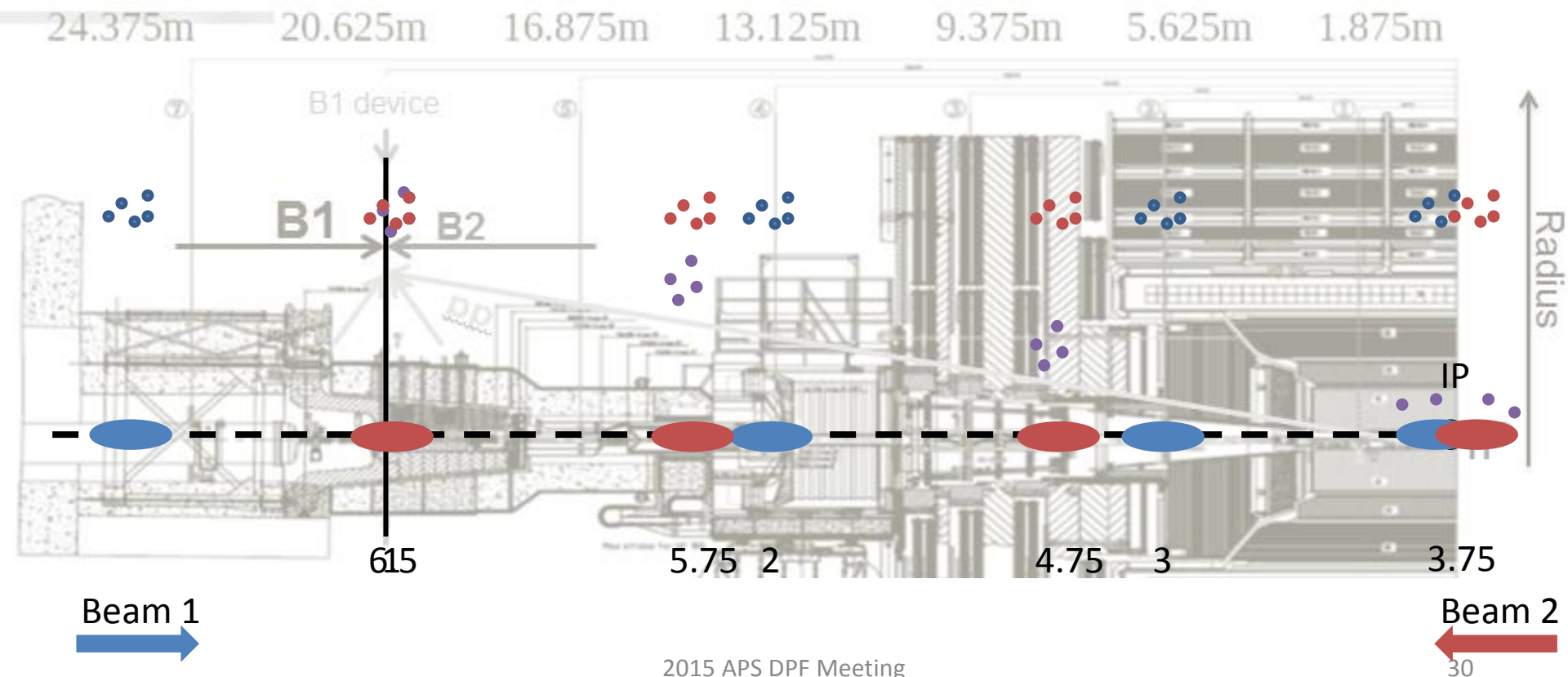


Figure 9 The wavelength of the Cherenkov light produced (left) and detected by the photocathode (right) as simulated when a muon of 4 GeV crosses 10 cm long quartz radiator, entering from the centre of the front face of the bar.

Timing

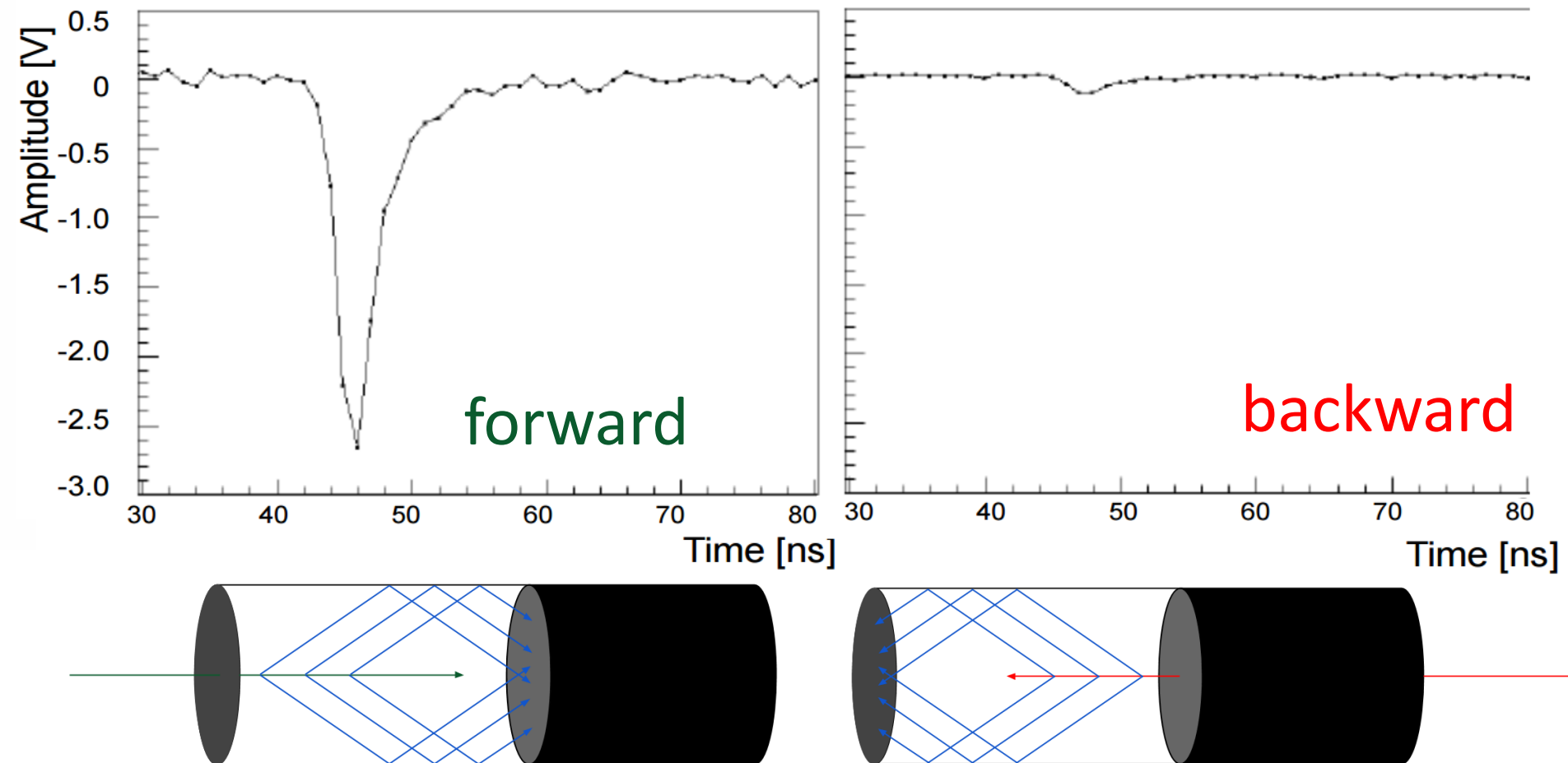
- Golden locations allow for maximum separation in timing

$$GL_{k+1} = \left(\frac{1}{4} + \frac{1}{2}k \right) \cdot (\text{BX spacing}) \approx 1.875\text{m} + 3.75\text{m} \cdot k$$



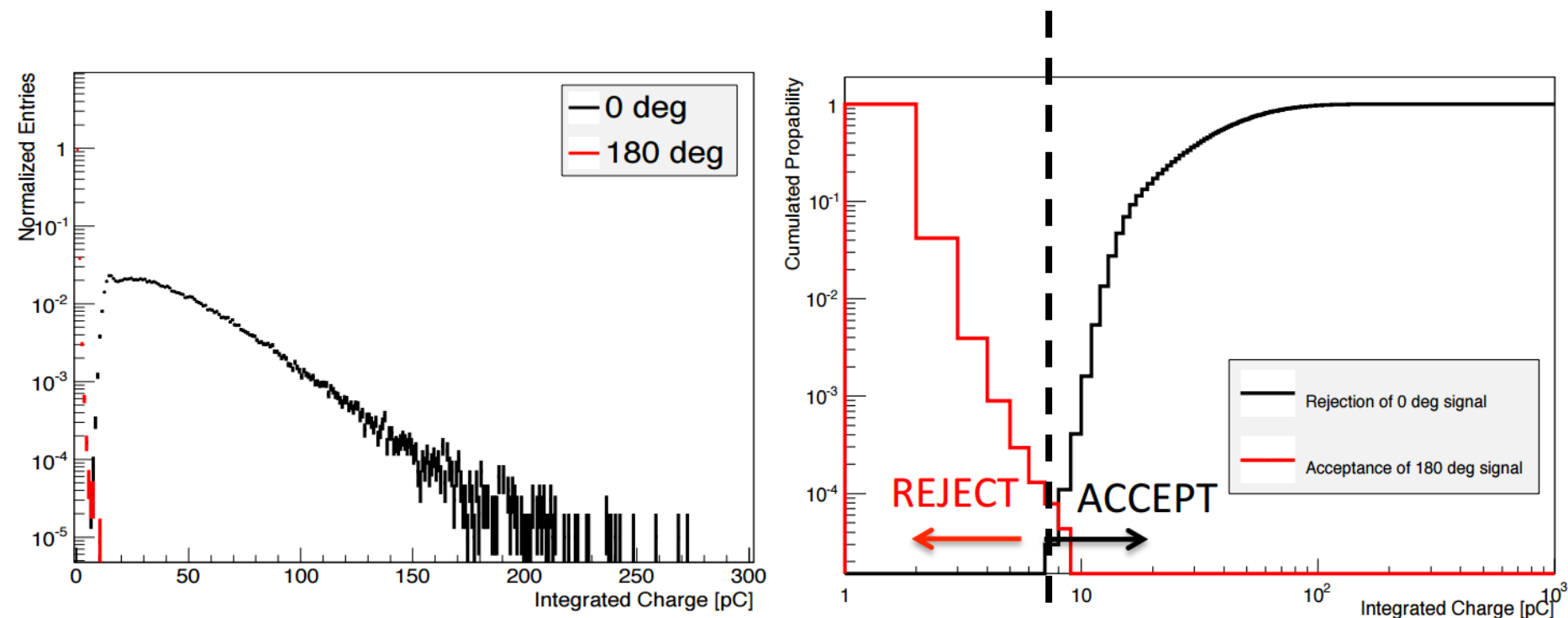
Test beam results

- Test beam studies performed at CERN in 2012



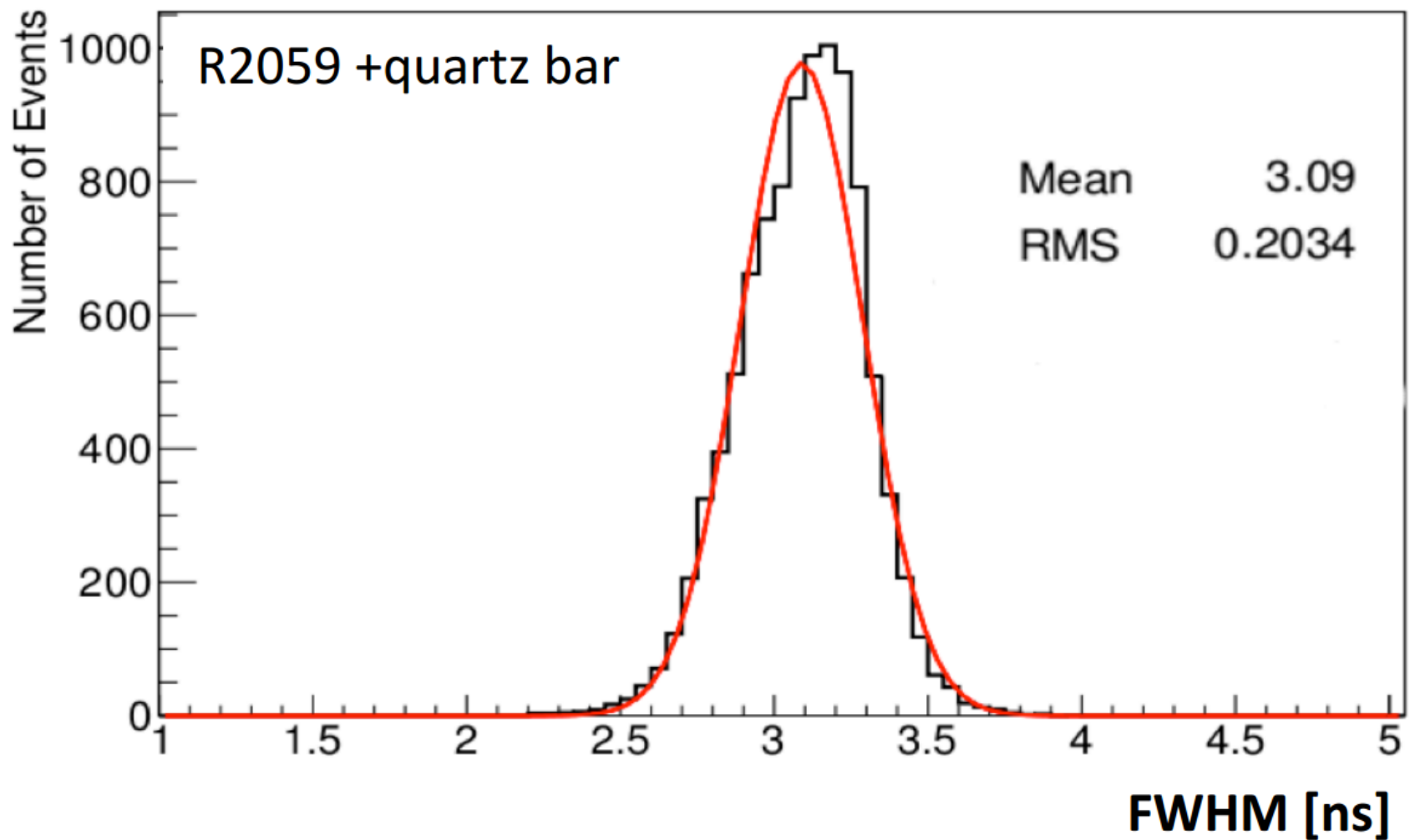
Test beam results

- Test beam studies performed at DESY in 2014

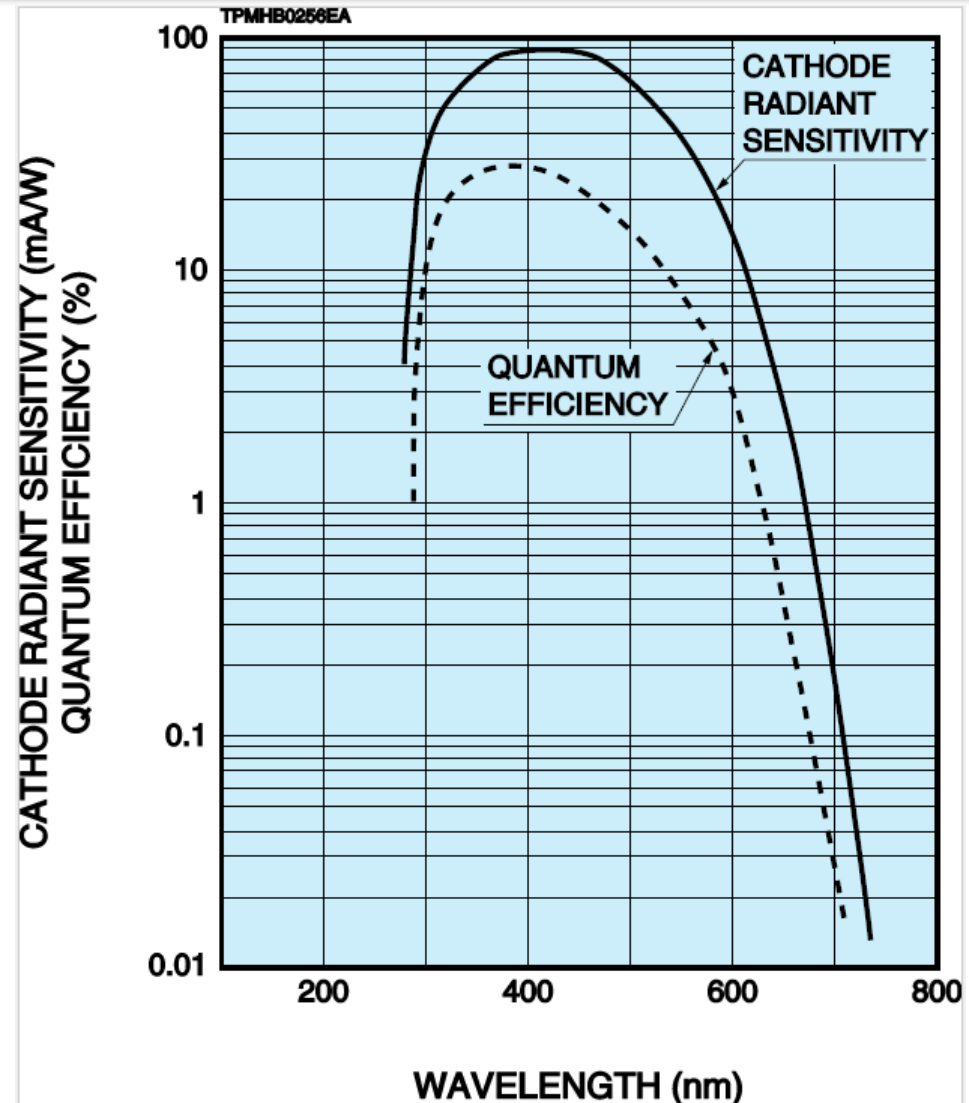
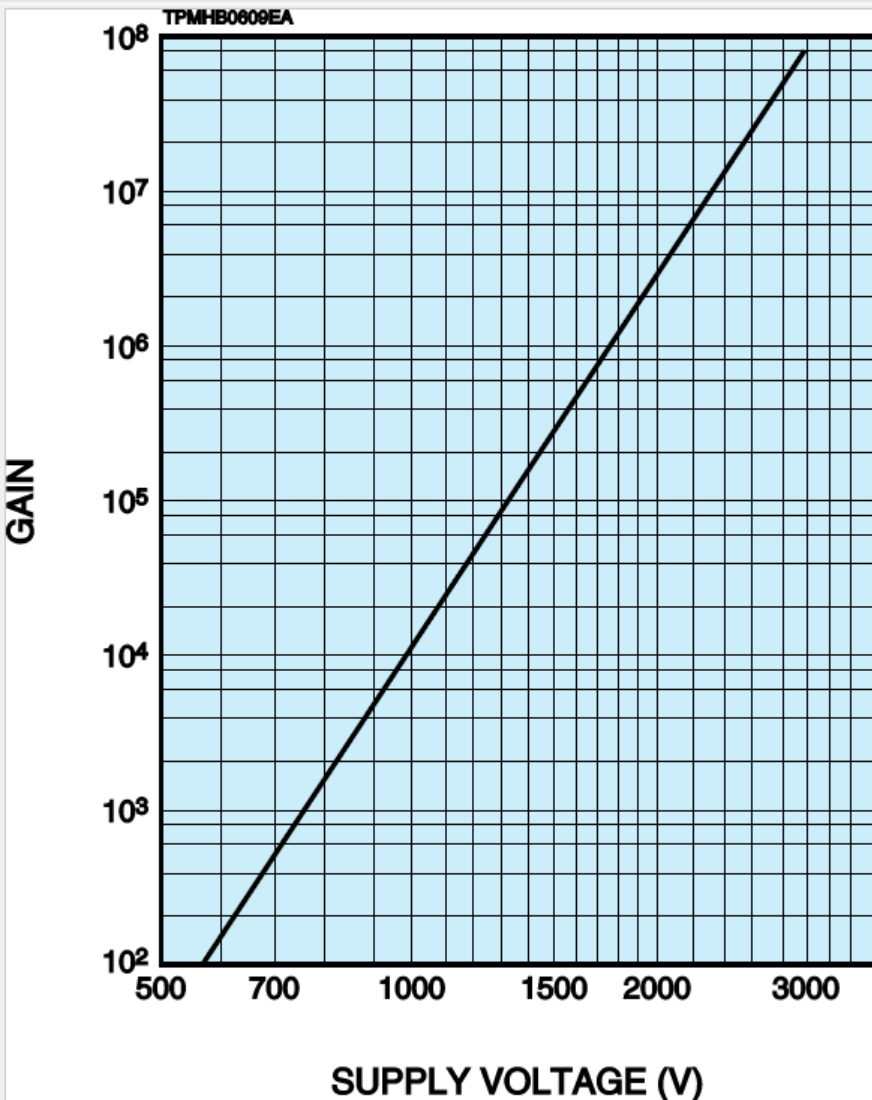


Amplitude cut: Backward suppression to **0.01%**, forward acceptance 100%

Test beam timing

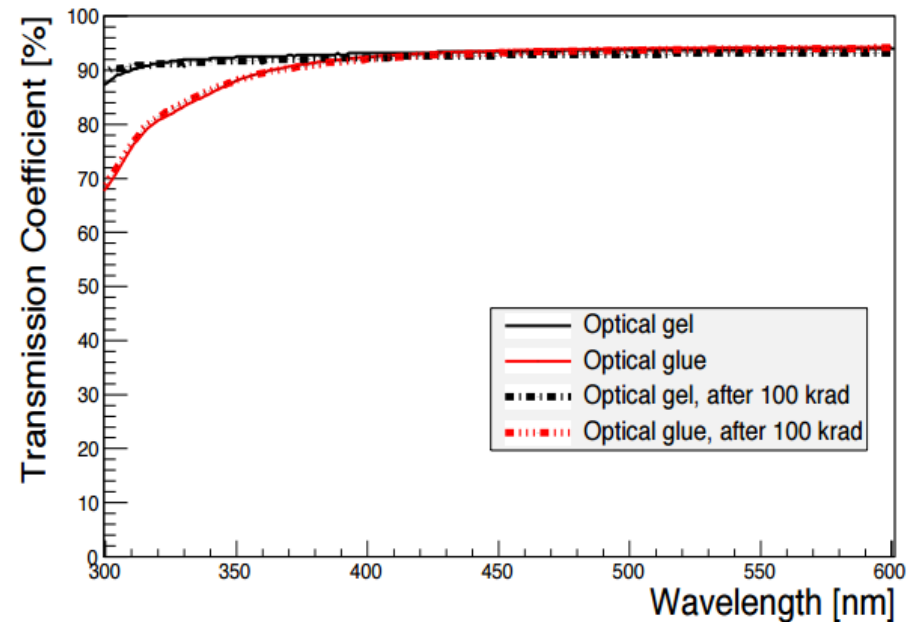
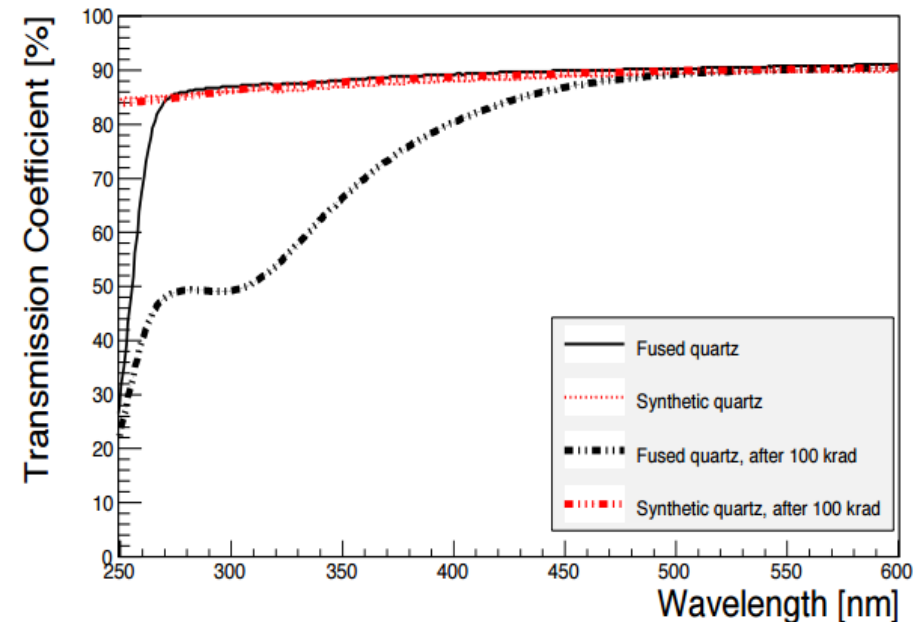


Hamamatsu R2059



Radiation studies

- Unit irradiated with 3000 fb^{-1} of γ rays



Detector units

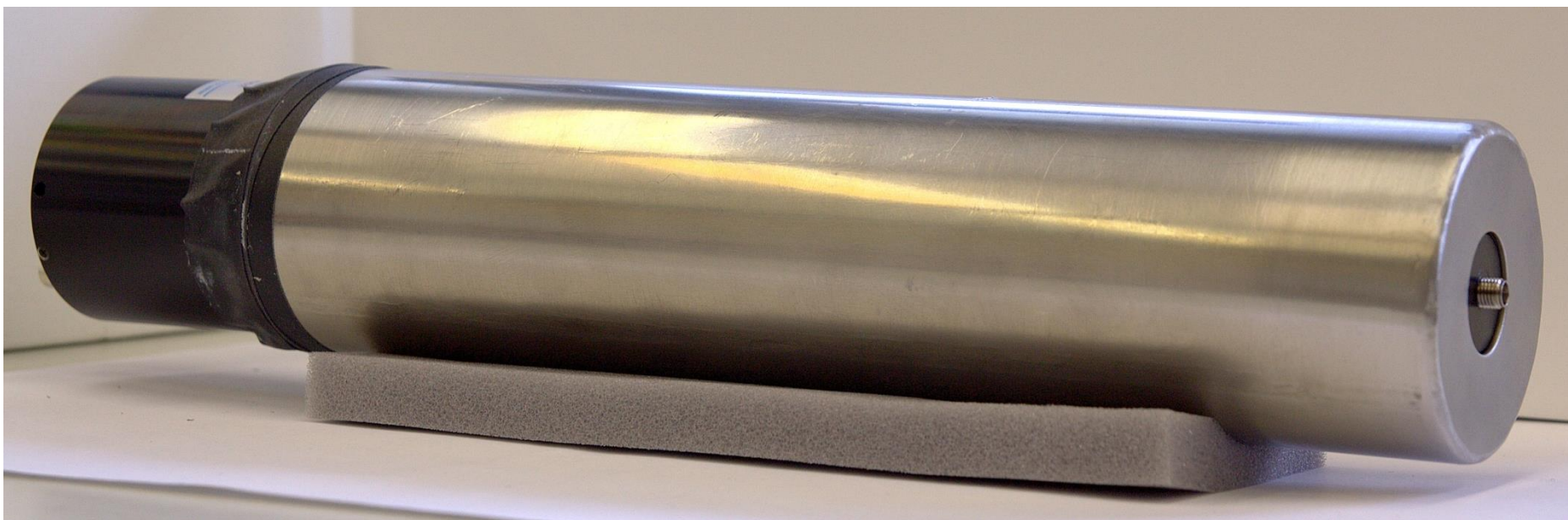


- Permally layer

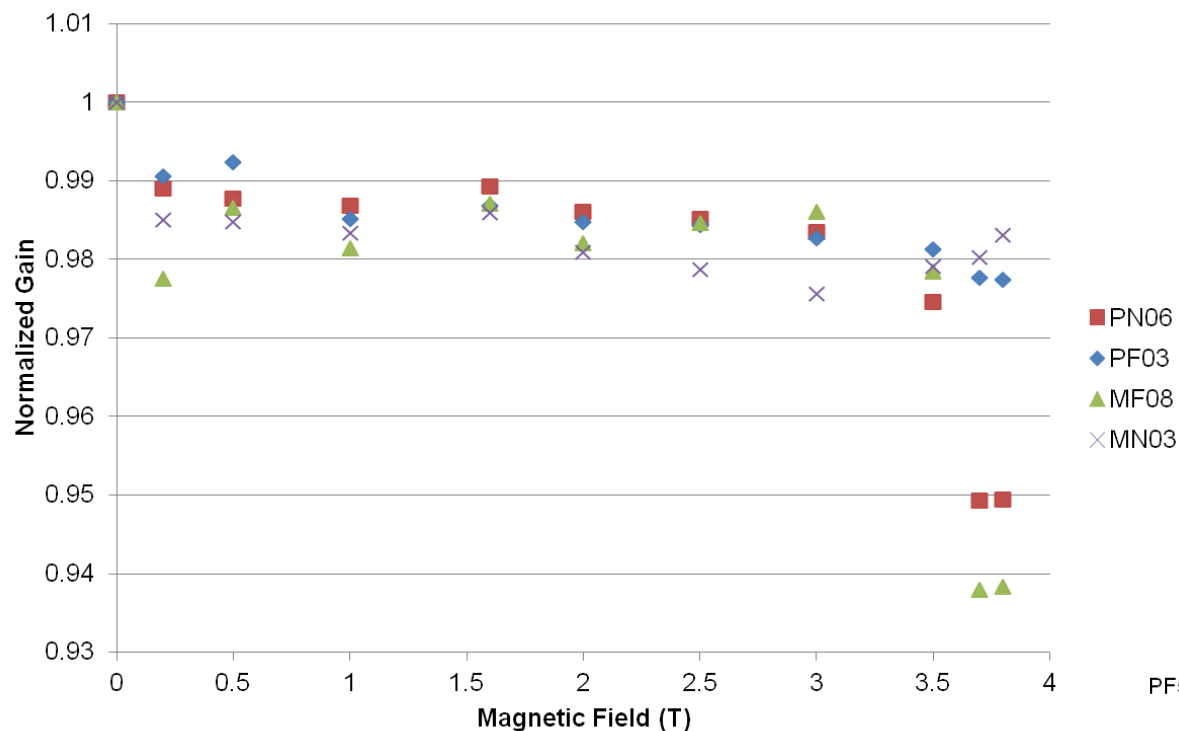


Detector units

- Mumetal layer

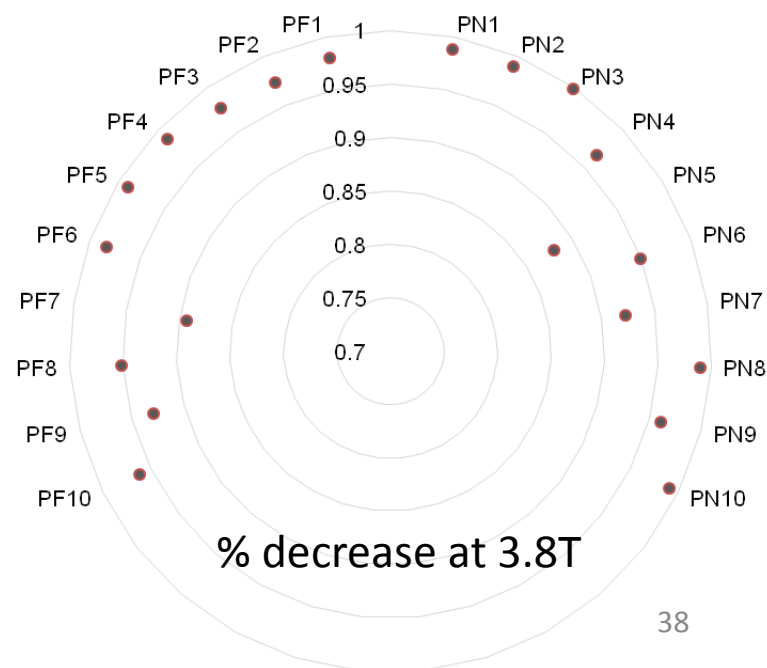


Magnetic shielding efficiency



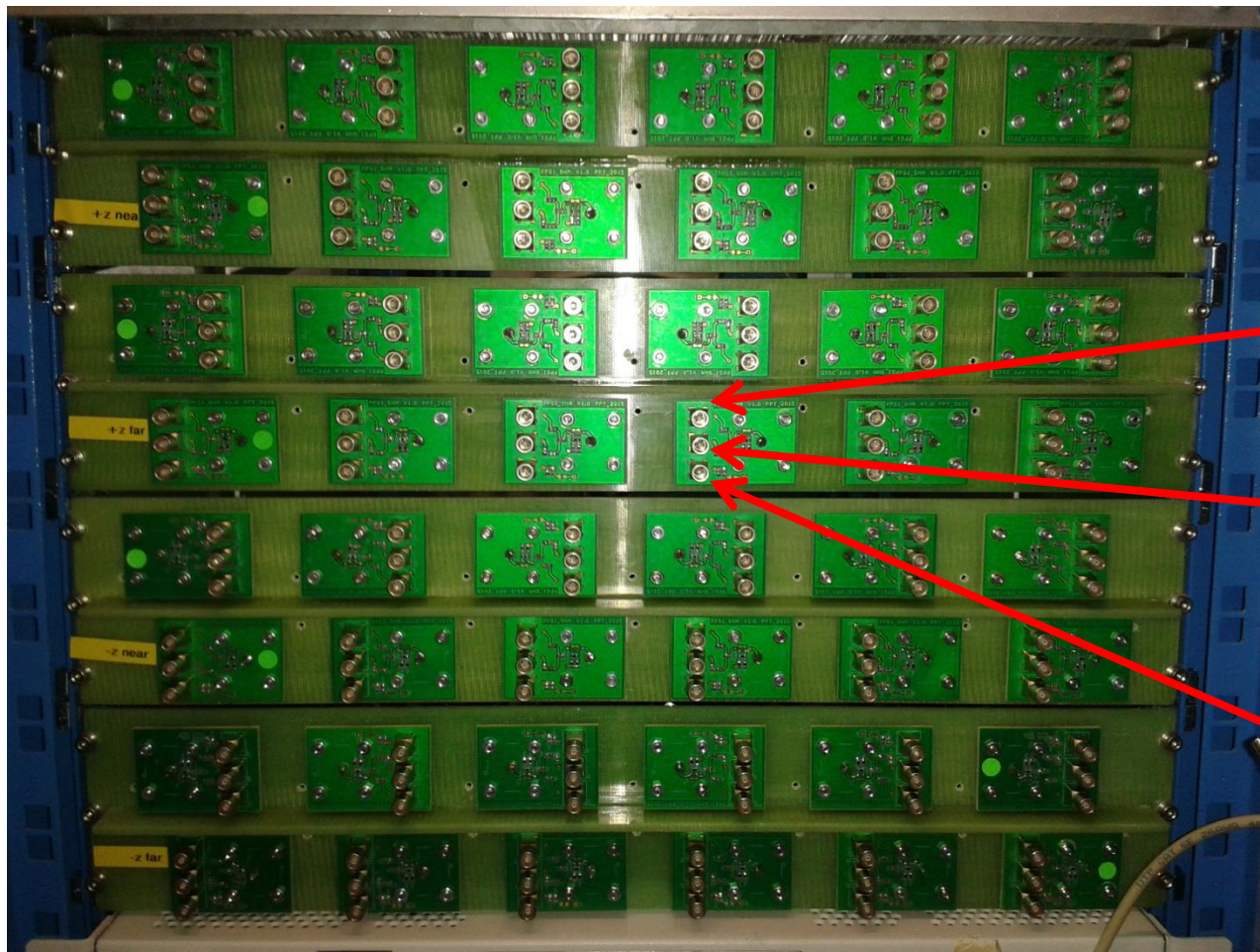
• Gain decreases as field increases

• Largest decrease ~10%

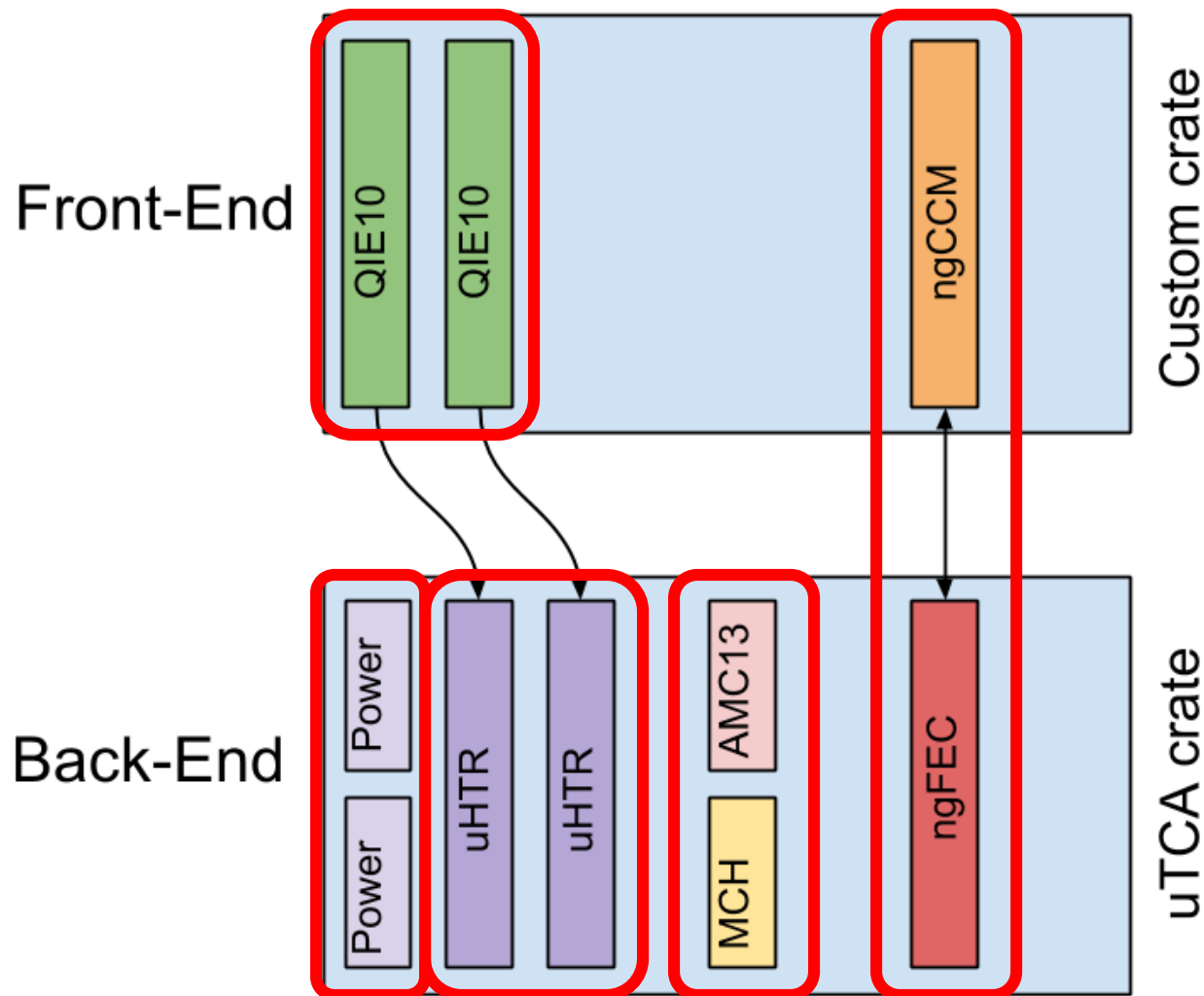


Patch panel

- Acts as passive splitter and attenuator

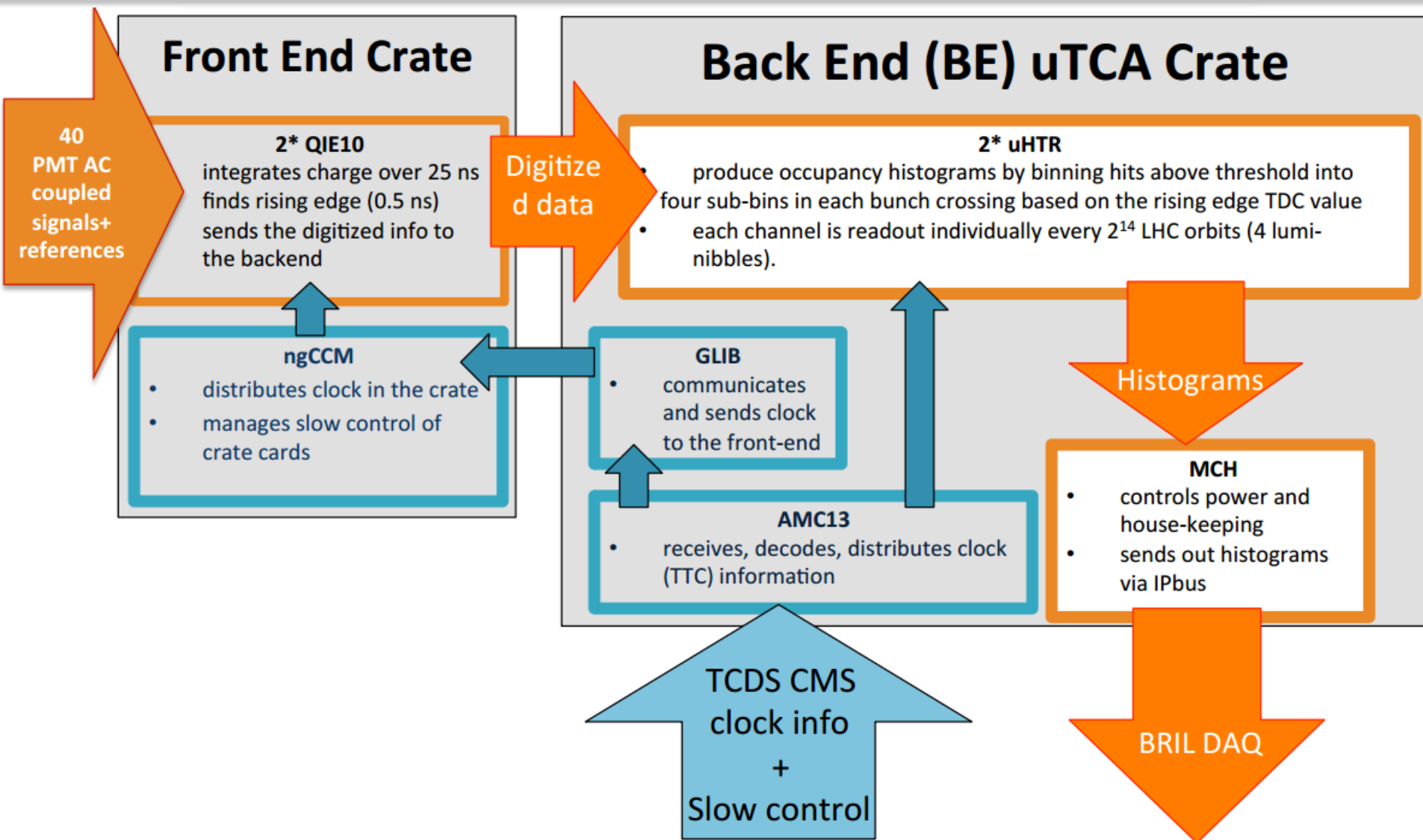


Read-out electronics



- QIE10:
digitizer
- uHTR:
histogramming
unit
- Other units:
power, clocks,
slow control,
data read-out

Electronics Overview



QIE10

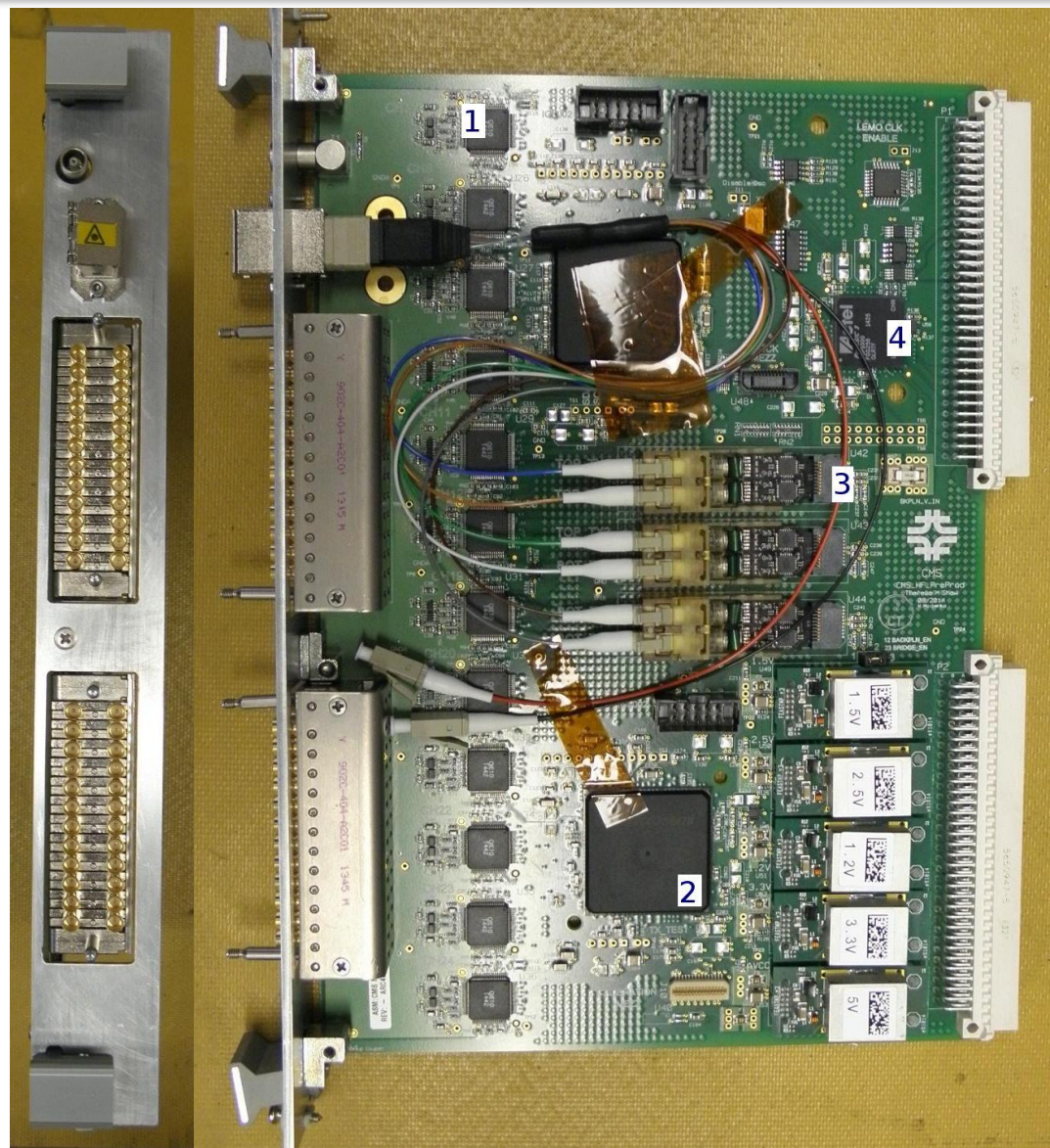
- Reads analog PMT signal

24 QIE10 ASICs:

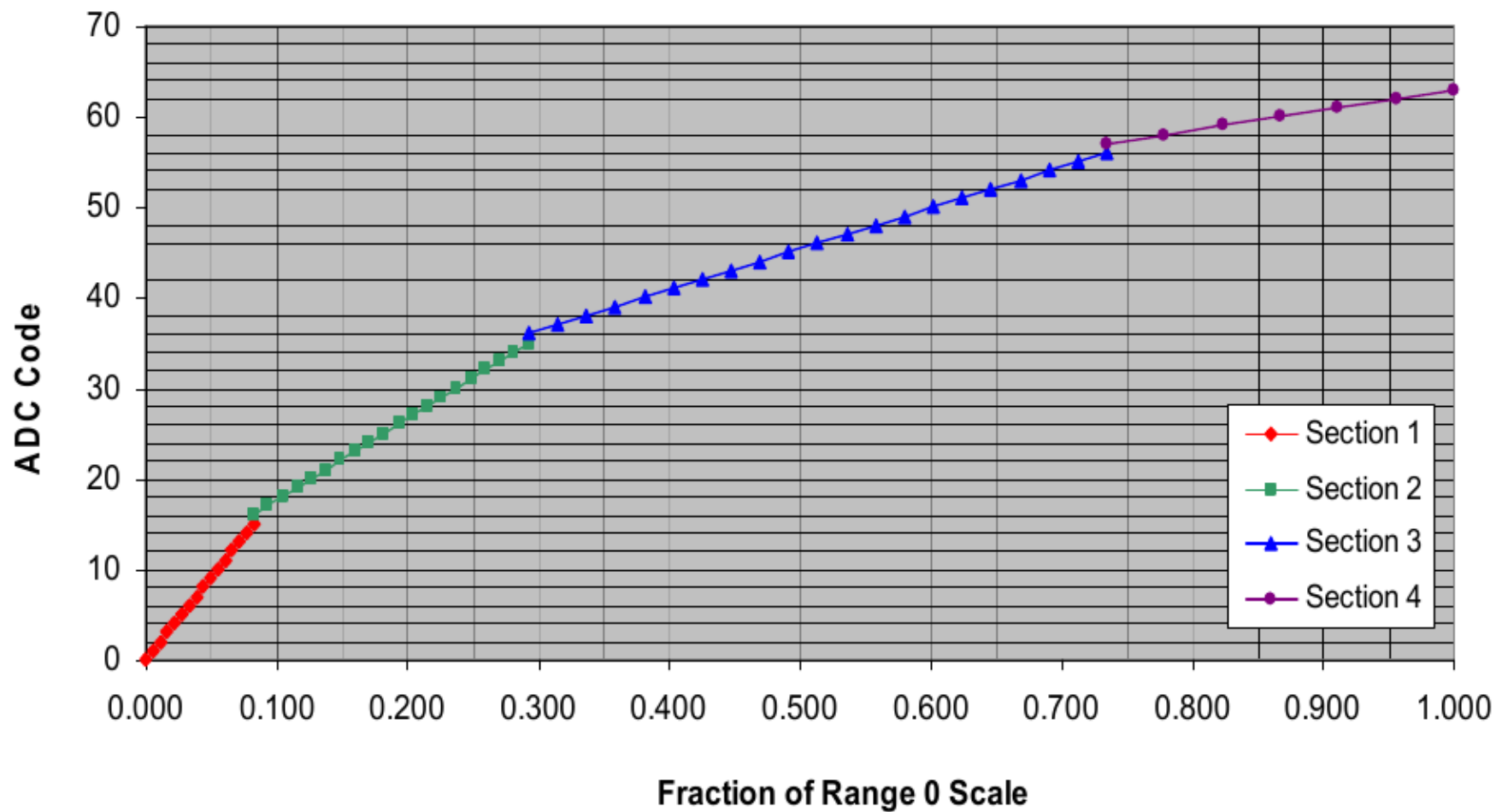
- Integrates charge over 25ns, produces 8-bit ADC value (0-340pC)
- Produces 6-bit TDC value based on fixed-threshold leading edge measurement (.5ns resolution)

Igloo2 FPGAs:

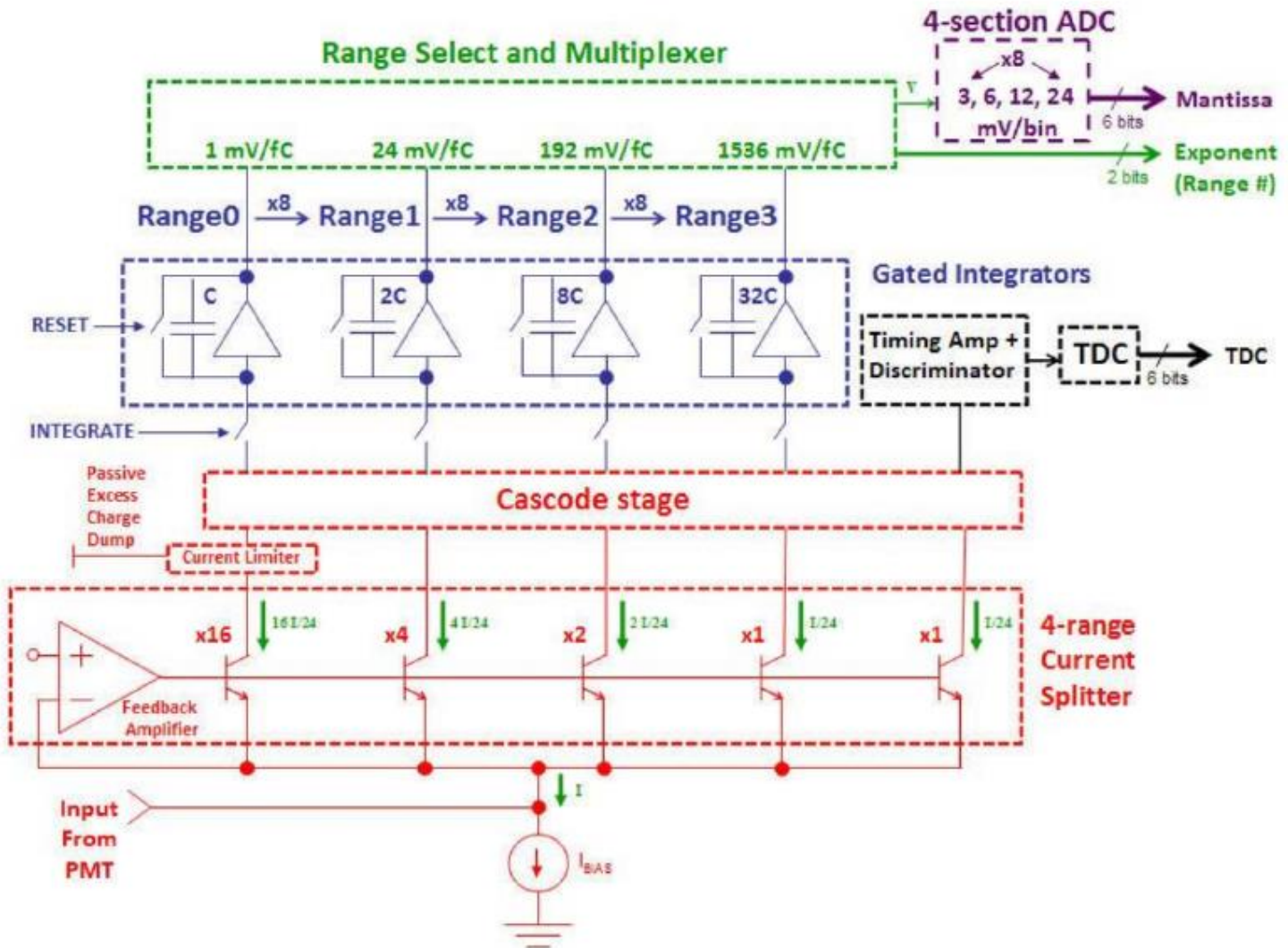
- Collect and format data
- Data sent to back-end via 5Gbps asynchronous optical link



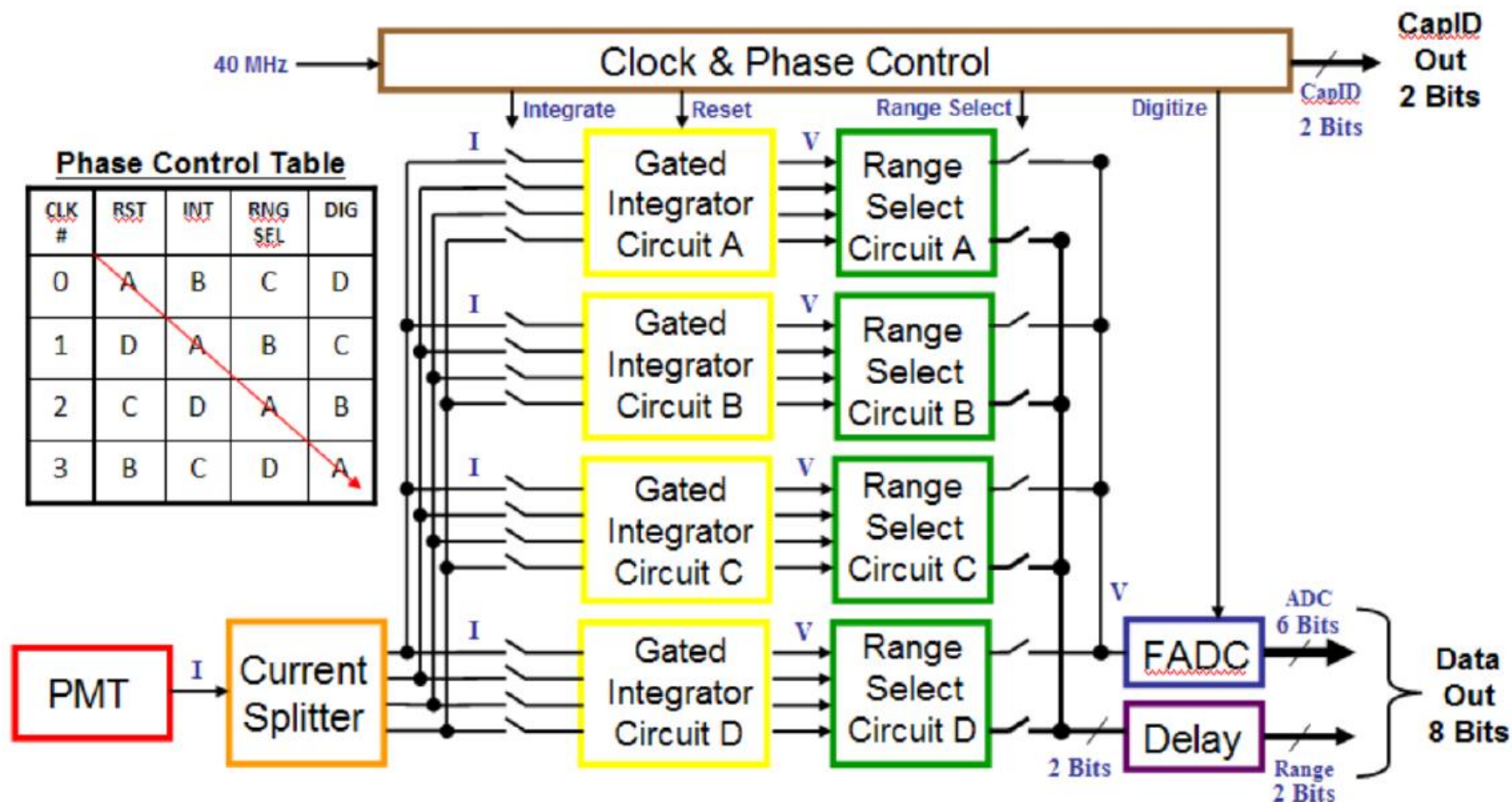
QIE10 Range



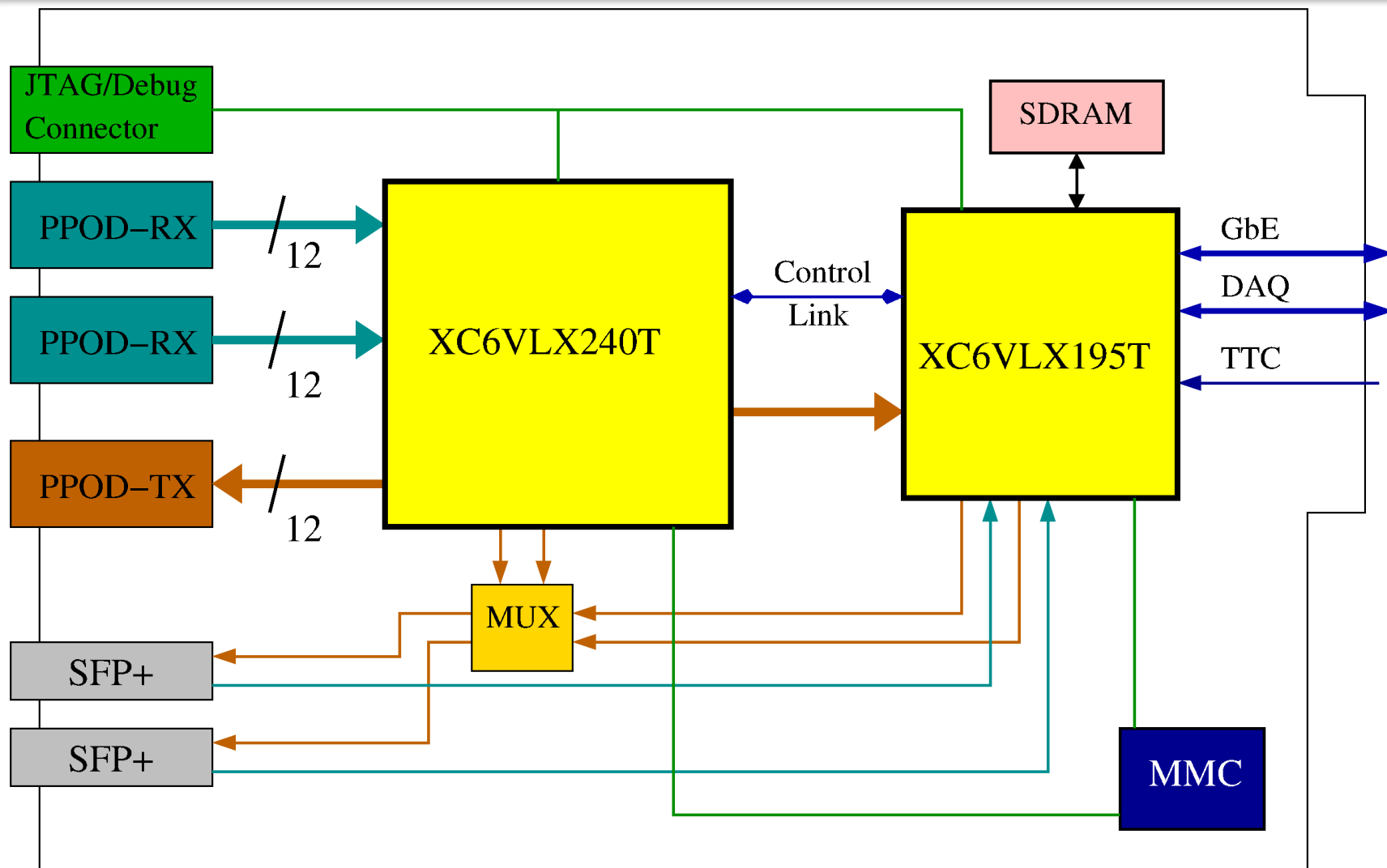
QIE10 input



QIE10 block diagram

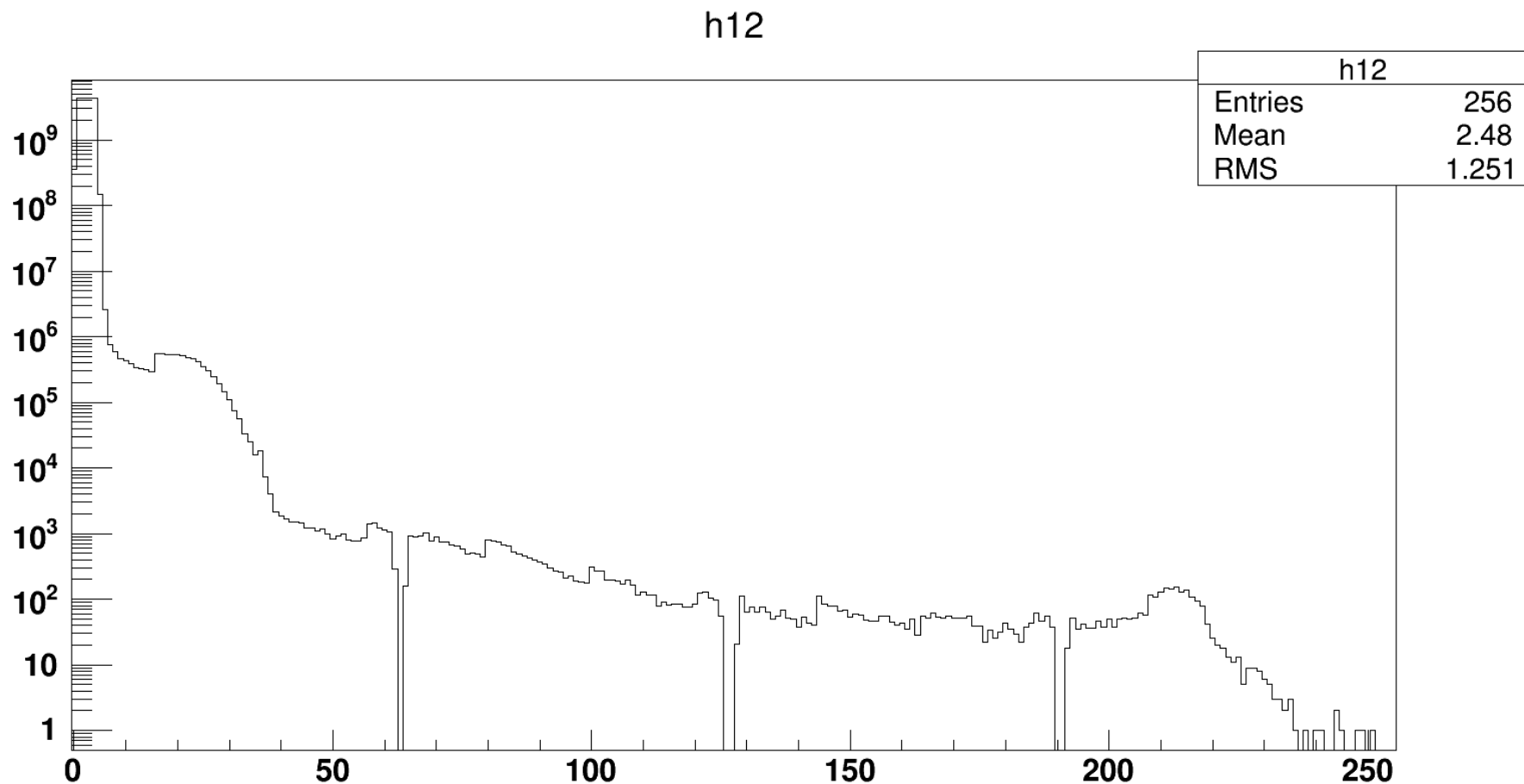


uHTR block diagram



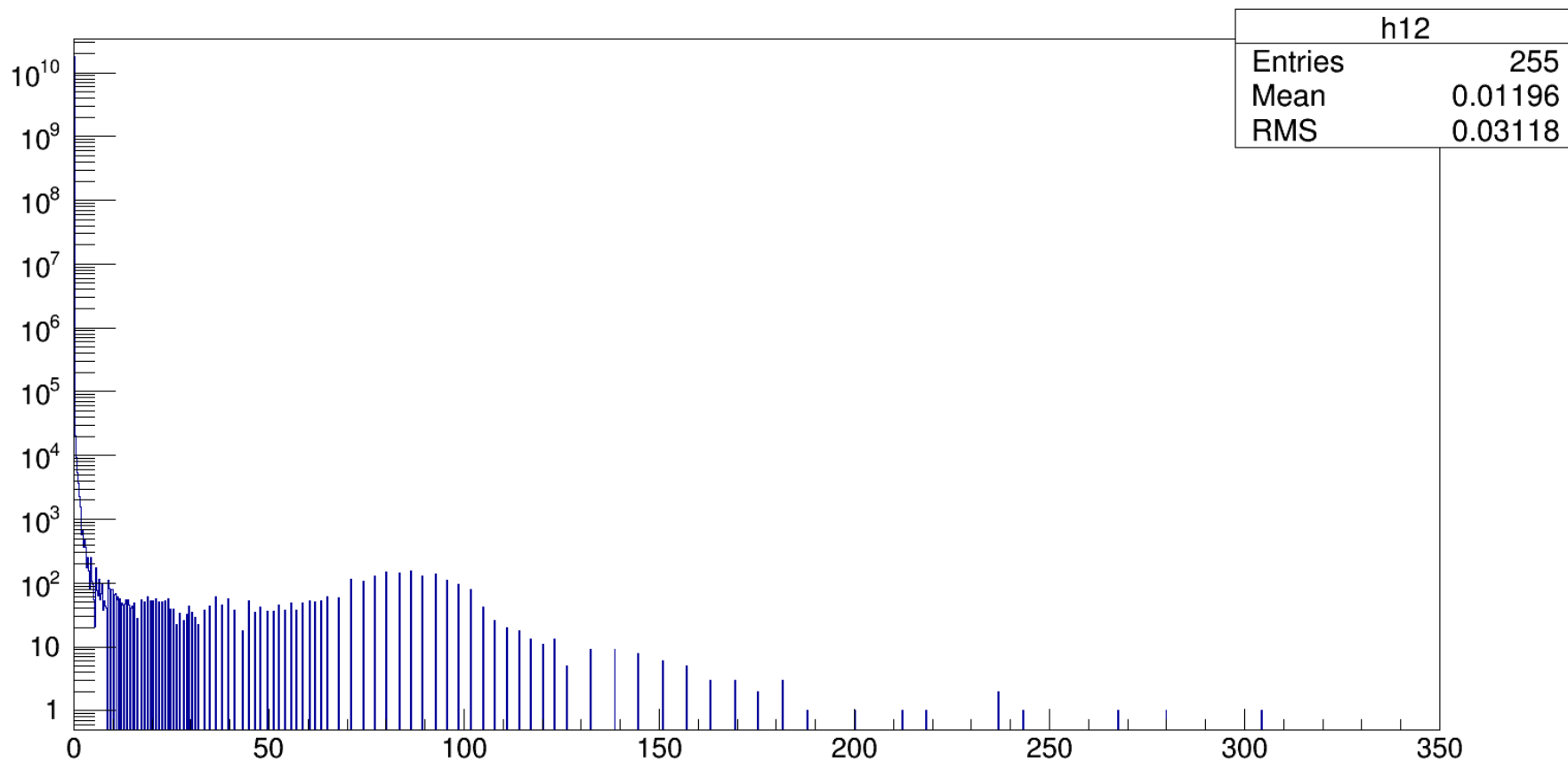
Setting thresholds

- Cut out majority of PP events, in addition to cosmics



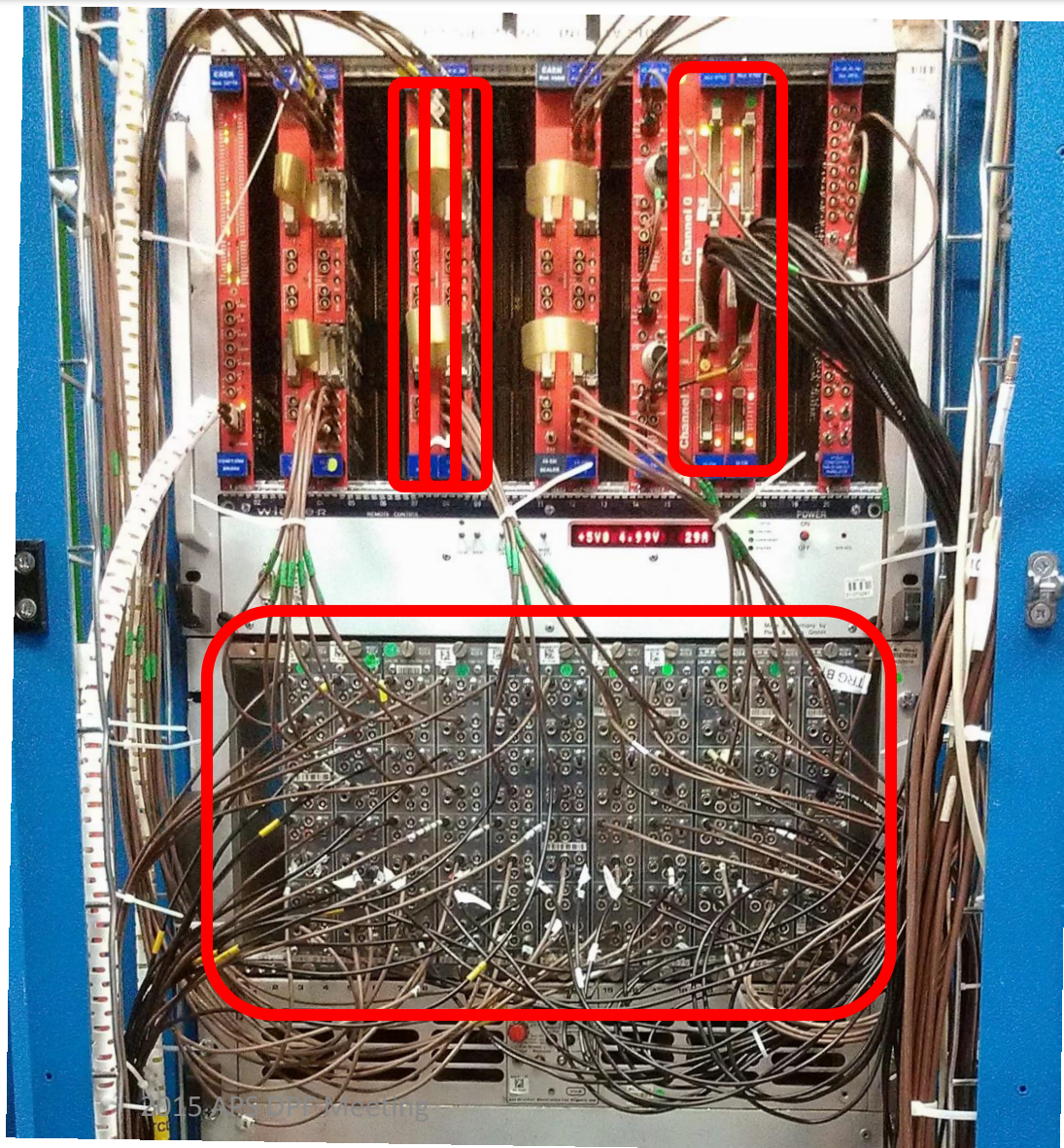
Setting thresholds

h12



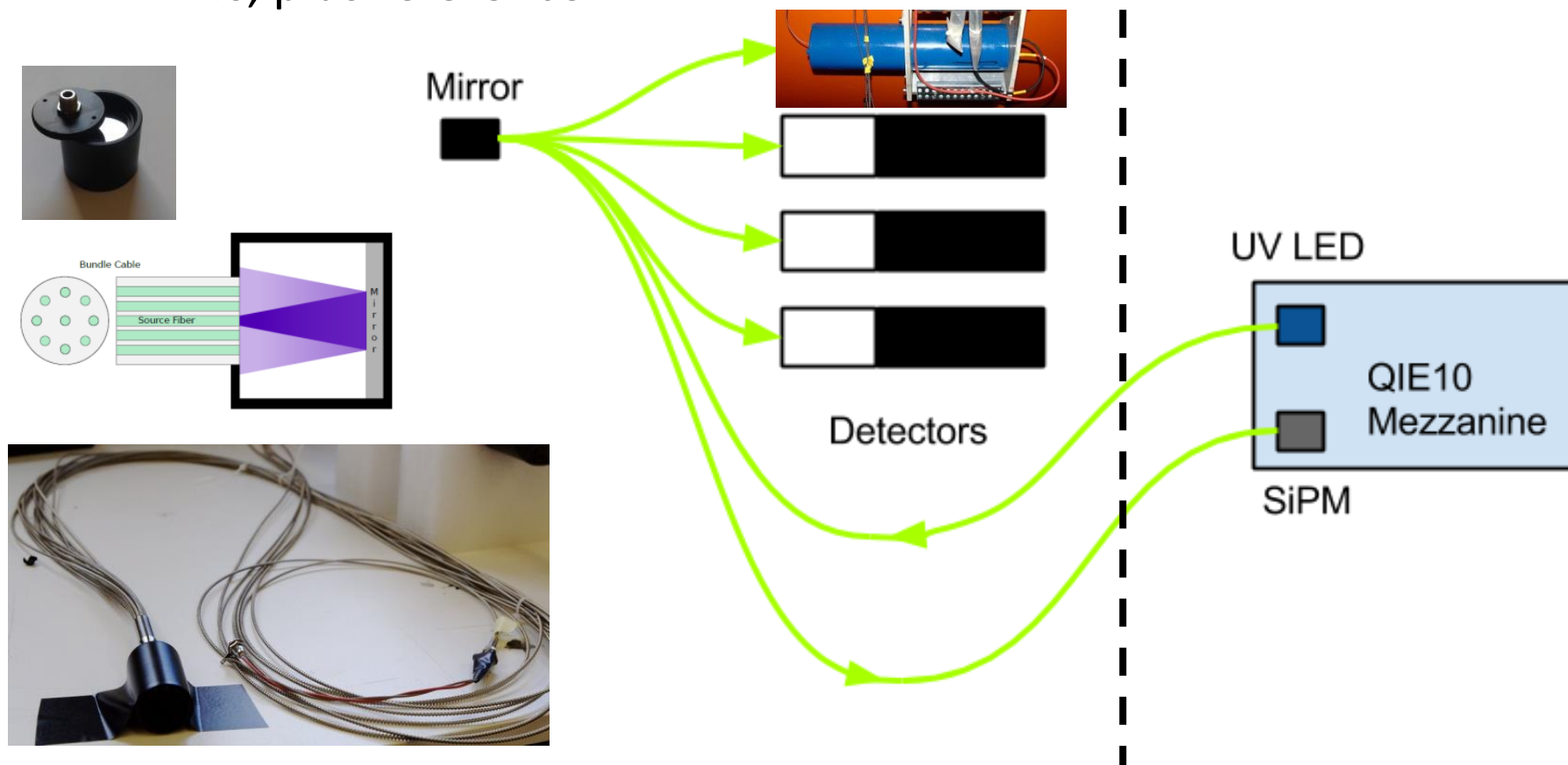
Read-out electronics: VME

- Background rate:
 - Fan out
 - ↳ Discriminator
 - ↳ Scaler
- Amplitude measurement:
 - Fan out
 - ↳ QDC



Calibration system

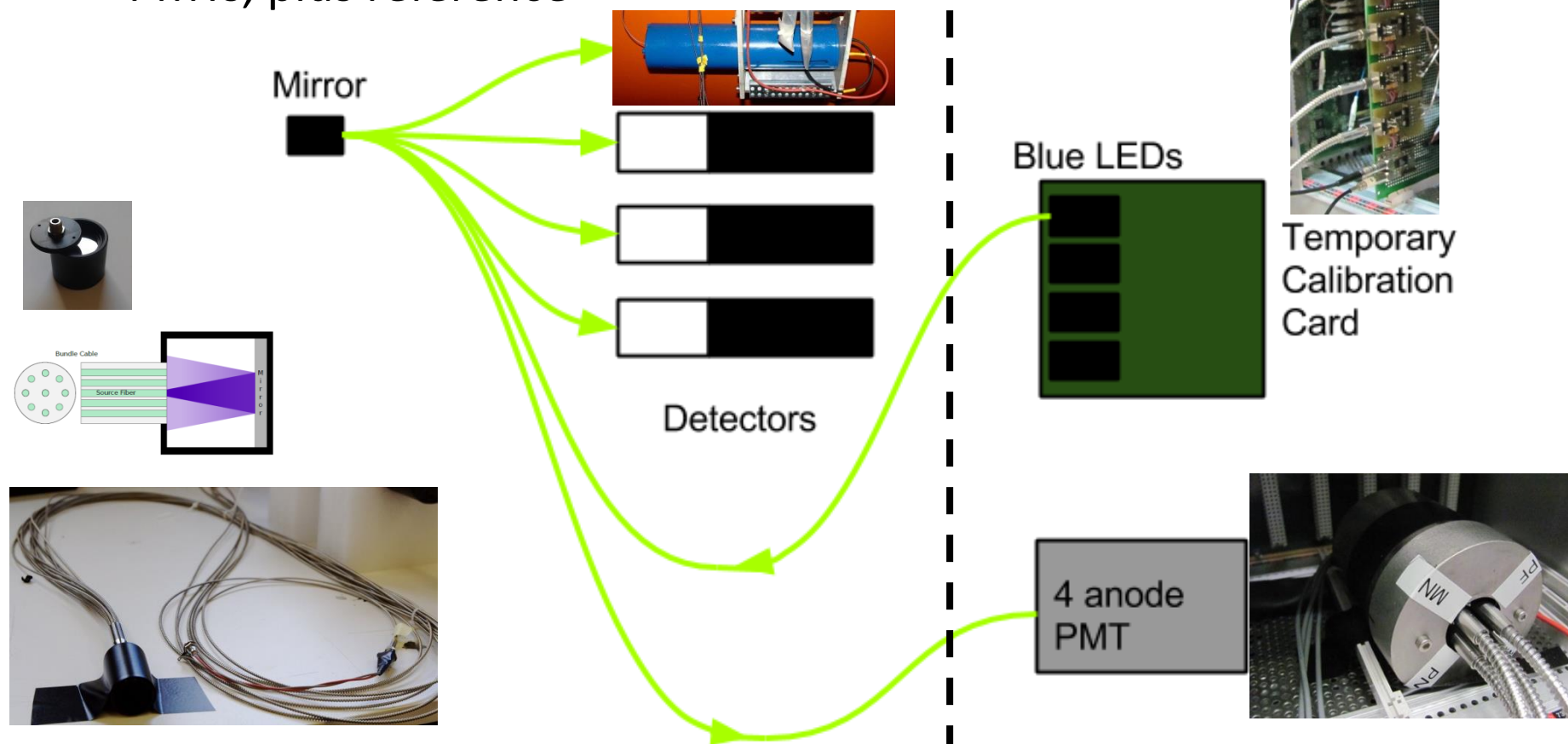
- Measures health of system over time
- UV LED pulses of known timing and amplitude distributed to all PMTs, plus reference



Temporary calibration system

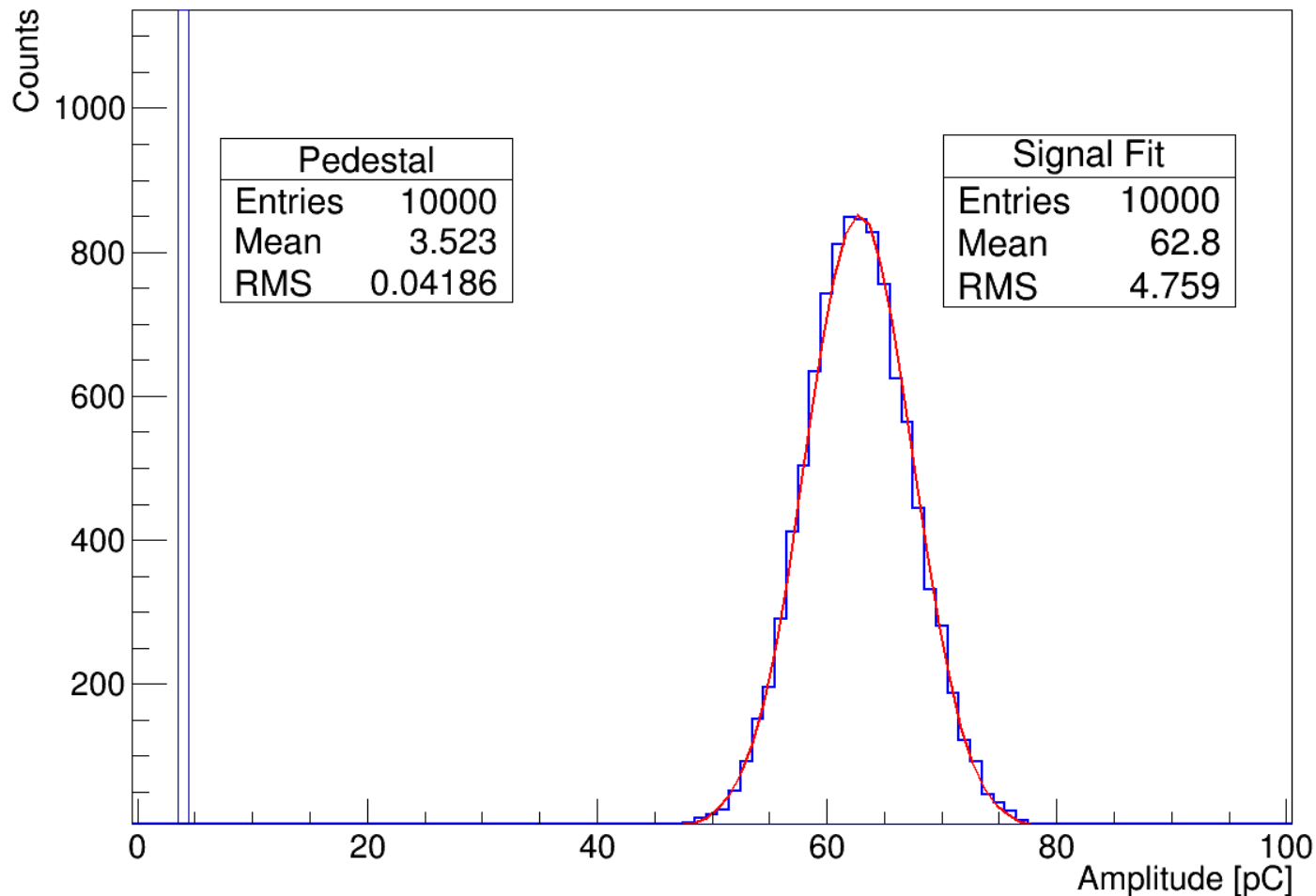


- Measures health of system over time
- UV LED pulses of known timing and amplitude distributed to all PMTs, plus reference



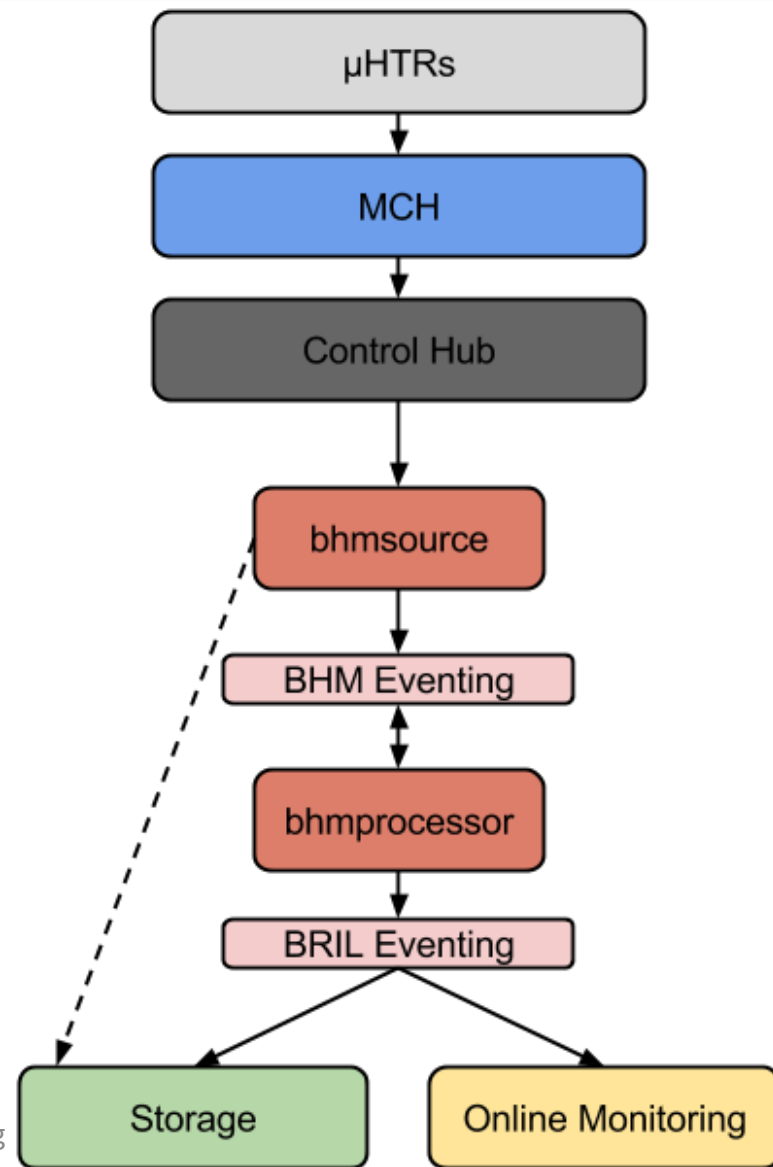
Calibration system

- Delivers pulses of expected signal size, ~ 600 photoelectrons

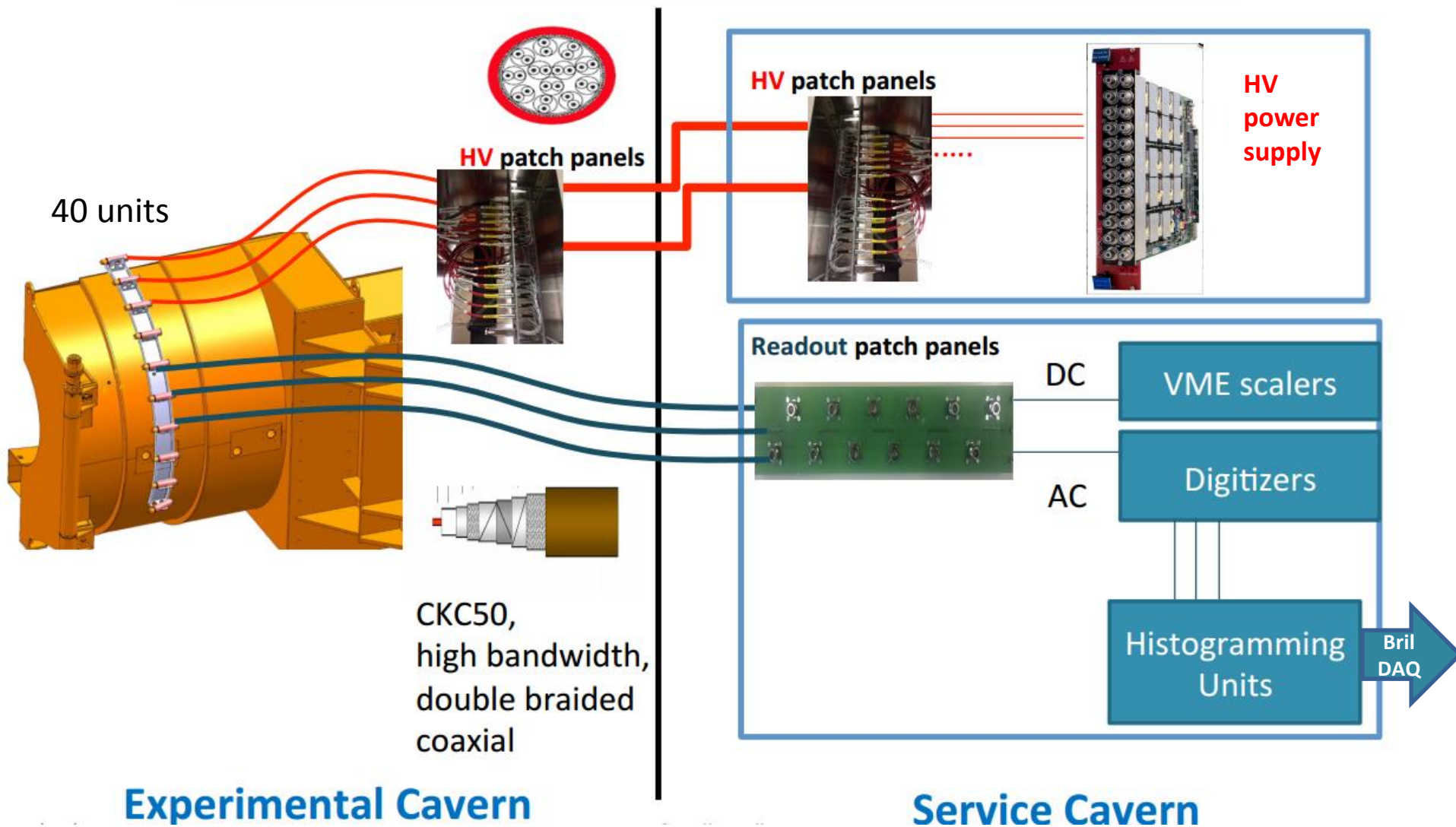


BHM in BrilDAQ

- BrilDAQ: based on xDAQ publisher/subscriber framework
- bhmsource – reads out histograms from uHTRs
- bhmprocessor – calculates background rate for each beam, publishes every lumi section (2^{18} orbits = ~ 23 s)



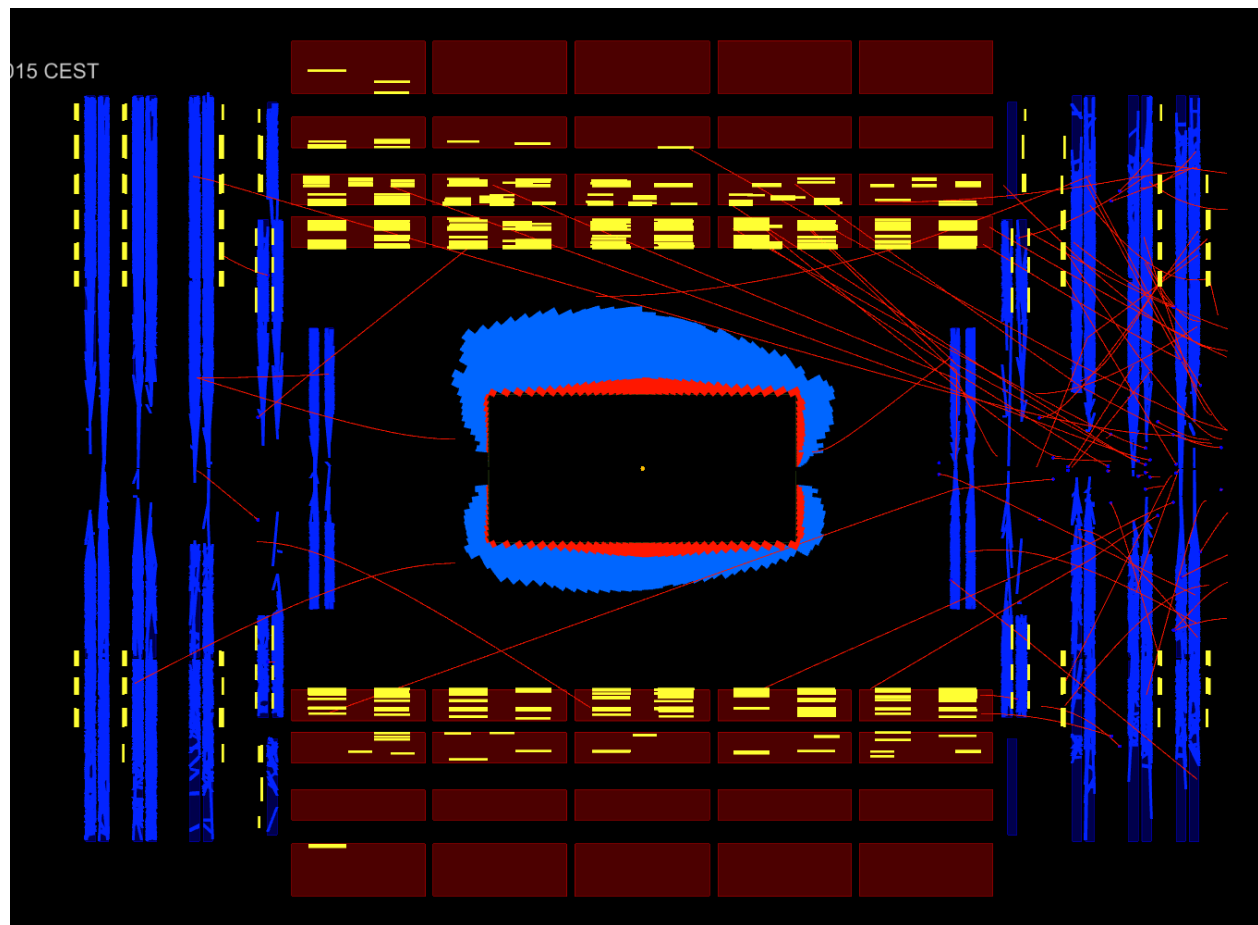
System overview



Splash events

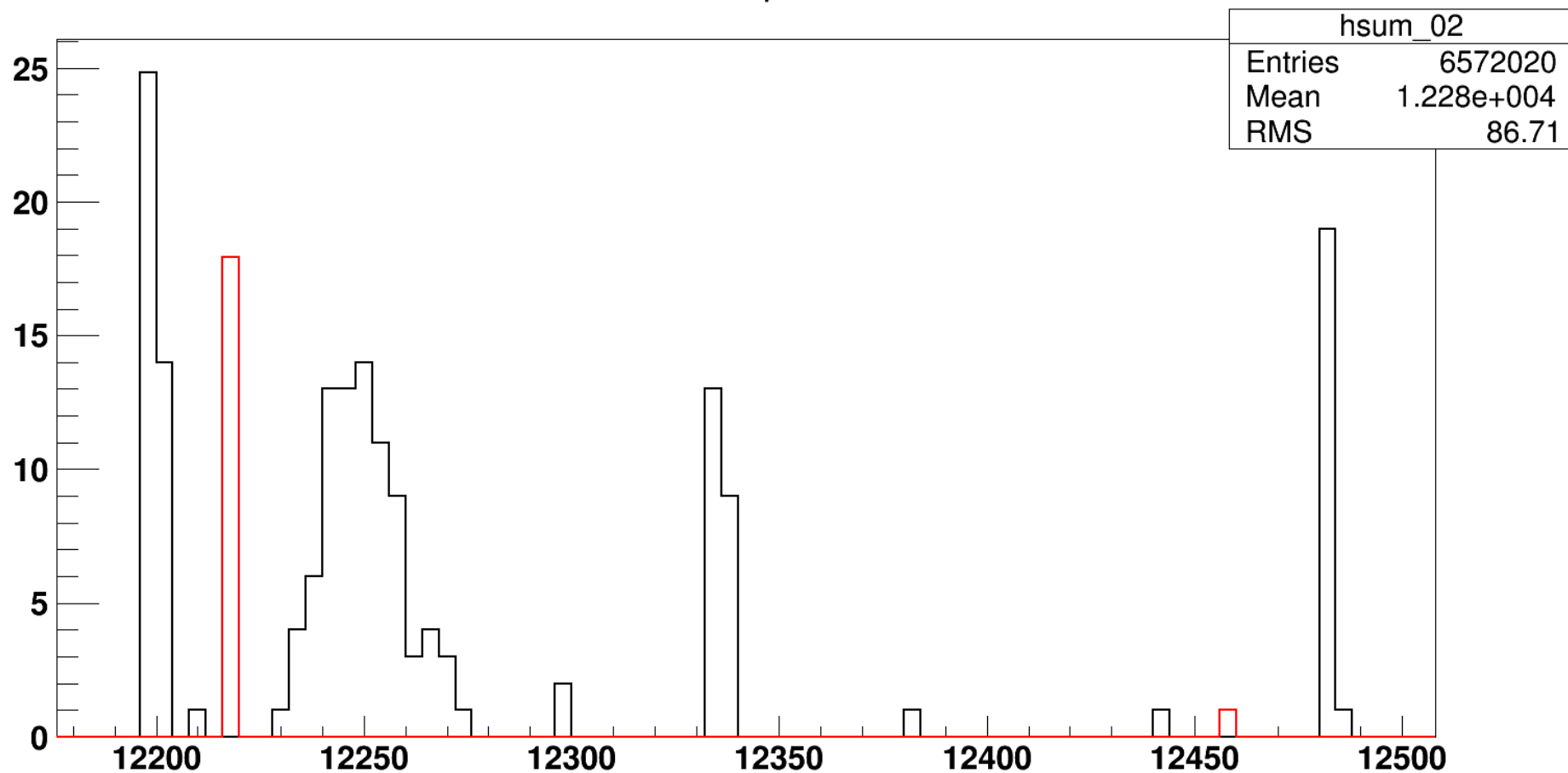


- Send a bunch directly into the TCTs – creates a ‘splash’ of particles in one direction

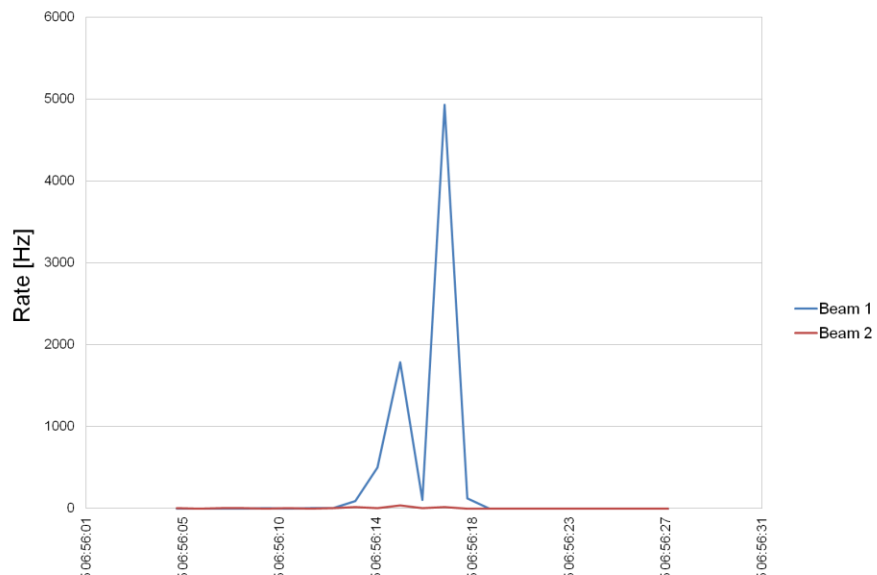
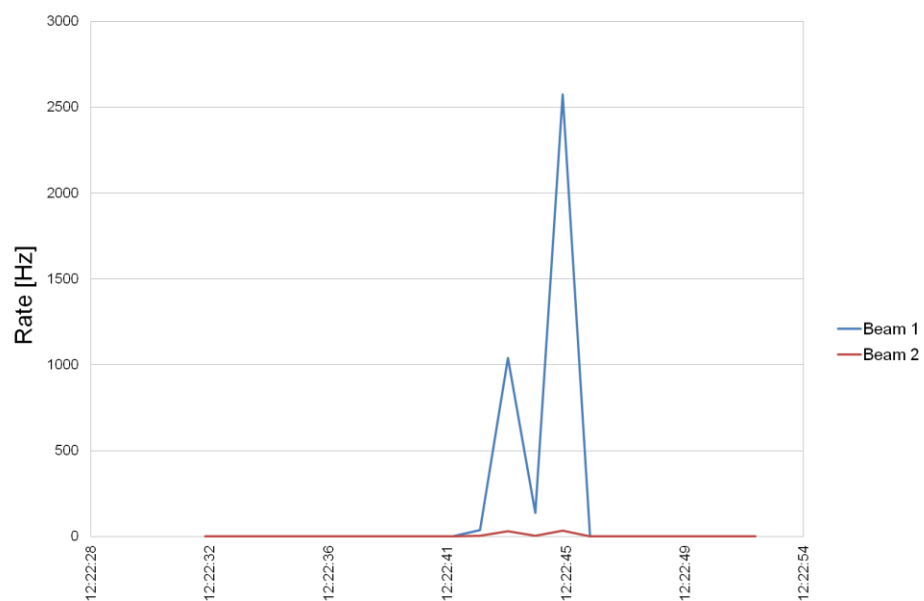
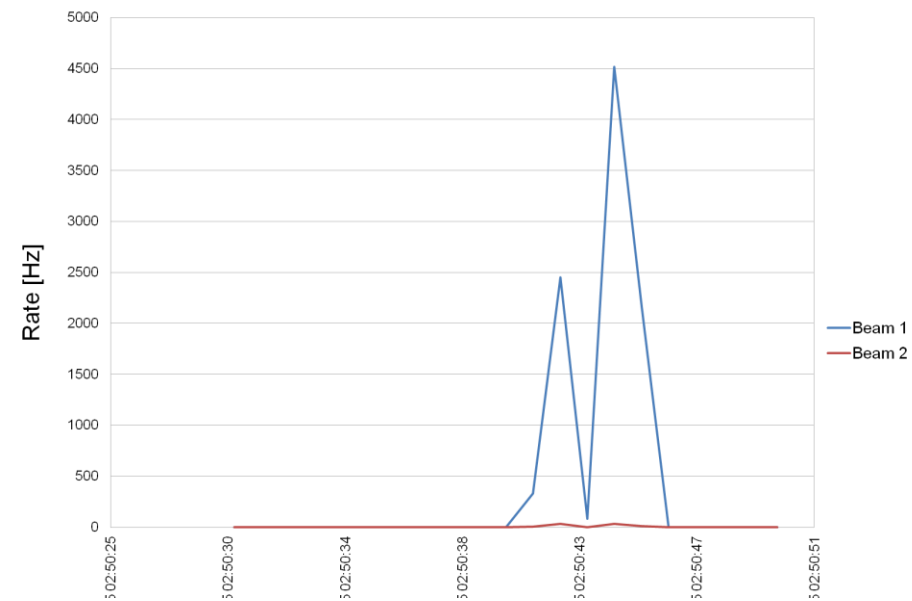


Splashes

Beam 1 splashes, PF10 and MN07



Beam losses



VME results

