

# Perturbative Unitarity and Dark Matter

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with

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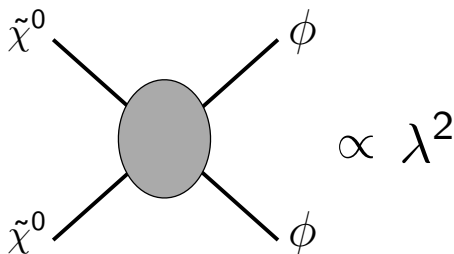
arXiv:1407.0395  
arXiv:1410.1534  
arXiv:1412.5660  
arXiv:1501.03153



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# Unitarity and Thermal Dark Matter

- ▶ Heavy Dark Matter needs to annihilate efficiently
- ▶ Annihilation cross section bounded by unitarity  
Griest and Kamionkowski, 1990



- ▶ Generic Unitarity Bound : 120 TeV for  $\lambda$  non perturbative
- ▶ **Can we identify lower characteristic scales using unitarity?**

# Perturbative Unitarity

Given a scattering matrix

$$S = 1 + iT$$

Optical theorem

$$S^\dagger S = I \Rightarrow -i(T - T^\dagger) = T^\dagger T$$

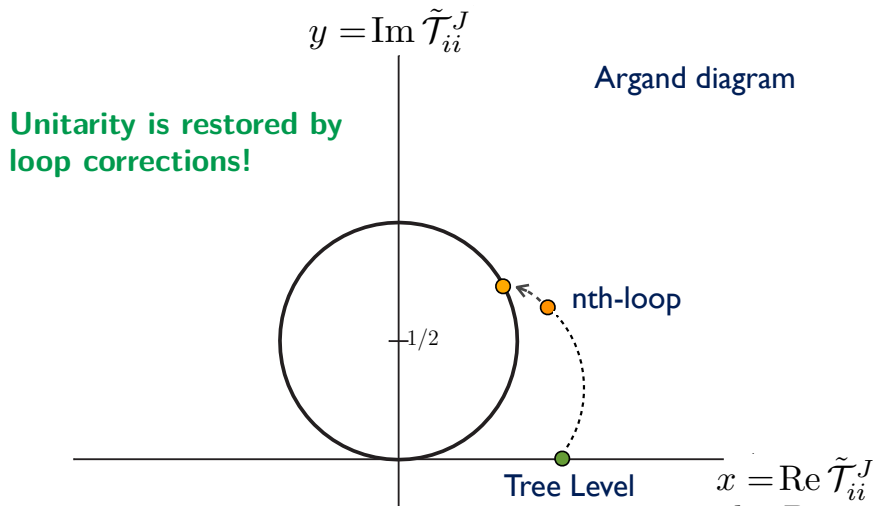
Use Partial Wave Decomposition

$$\tilde{T}_{ij}^J = \frac{\lambda_i^{1/4} \lambda_f^{1/4}}{32\pi s} \int_{-1}^1 T_{ij} P_J(\cos \theta) d \cos \theta$$

We find

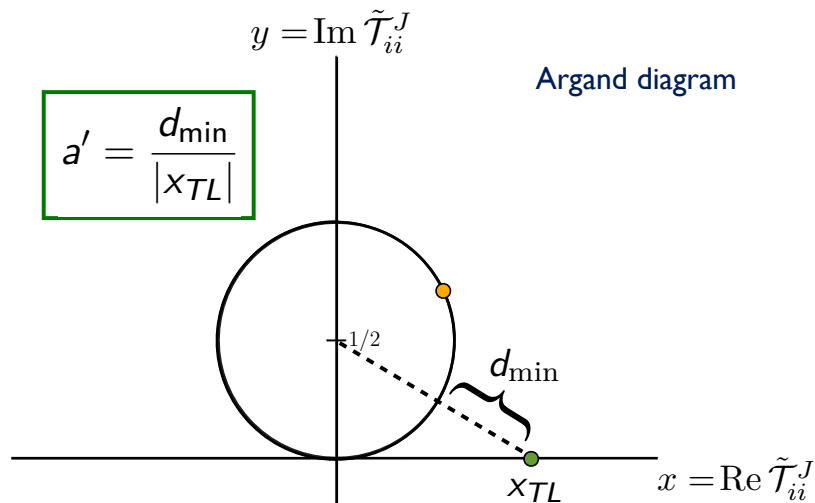
$$\text{Im } \tilde{T}_{ii} = |\tilde{T}_{ii}|^2 \Rightarrow |\text{Re } \tilde{T}_{ii}| < \frac{1}{2}$$

# Unitarity in the Complex Plane



Schuessler and Zeppenfeld [arXiv:07105175, Schuessler's thesis (2005)]  
Aydemir, Anber, Donoghue [arXiv:1203.5153]

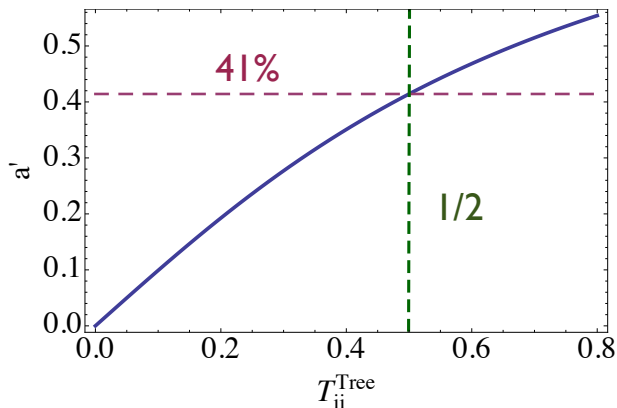
# Unitarity and Loop Corrections



Schuessler and Zeppenfeld [arXiv:07105175, Schuessler's thesis (2005)]

# Perturbativity breakdown

- ▶ For large  $\tilde{T}_{ii}$  restoring unitarity violates perturbativity!



Schuessler and Zeppenfeld [arXiv:07105175, Schuessler thesis (2005)]

- ▶ **New perturbativity bounds on Dark Matter models**

## Application: Squark-Dark Matter simplified model

$$\mathcal{L} \supset \frac{1}{2} M_\chi \bar{\chi} \chi + \frac{1}{2} M_{\tilde{u}}^2 \tilde{u}^* u + \lambda_{\text{dark}} \tilde{u}^* \bar{\chi} P_R u$$

- ▶ Example: Majorana fermionic Dark Matter  
DiFranzo, Nagao, Rajaraman, Tait [[arXiv:1308.2679](https://arxiv.org/abs/1308.2679)]
- ▶ MFV – Consider one degenerate flavor triplet

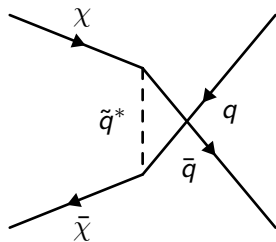
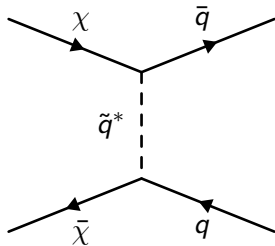
$$\tilde{u}_R = (\tilde{u}_R, \tilde{c}_R, \tilde{t}_R)$$

- ▶ Three parameters – Two mass scales

$$\lambda_{\text{dark}}, M_\chi, M_{\tilde{u}}$$

# Squark-Dark Matter – Strategy

- ▶ Same diagrams for Unitarity, Direct Detection and Relic Density constraints

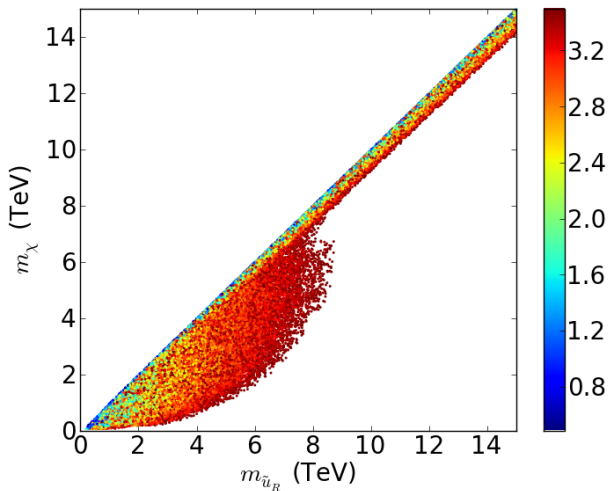


- ▶ Unitarity  $\Rightarrow \lambda_{\text{dark}} \lesssim 3.5$
- ▶ Negative interference  $\Rightarrow$  suppressed annihilation rate



# The Perturbative Sector

- ▶ Suppressed annihilation cross section

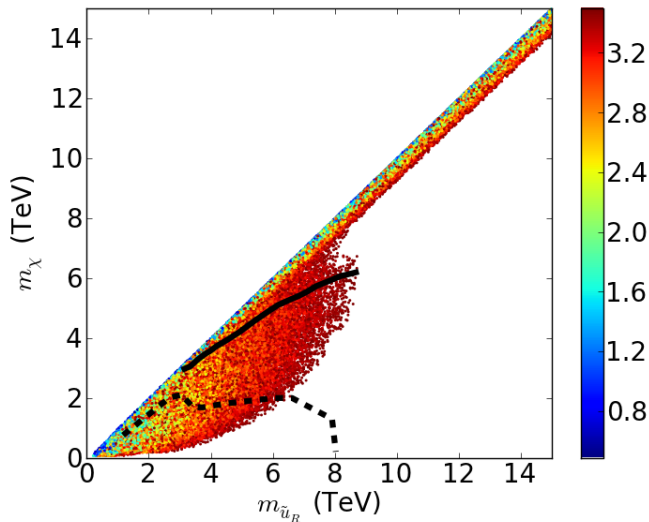


- ▶ 10 TeV bound on Squark and Dark Matter masses!

# Squark Portal at 100 TeV – Exclusion reach

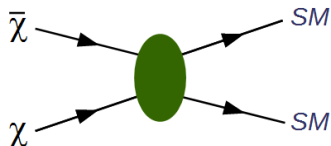
Snowmass 100 TeV search recasted for large  $\lambda_{\text{dark}}$

Cohen, Golling, Hance, Henrichs, Howe, Loyal, Padhi, Wacker [arXiv:1311.6480]

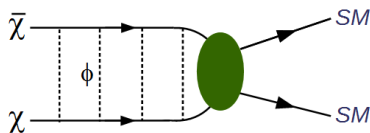


# Consequences of breaking perturbativity

The Dark Matter candidate forms bound states



Elementary DM annihilation



Bound state decay

Completely different physics in the dark sector!

- ▶ Different depletion mechanism
- ▶ Different collider/indirect/DD phenomenology

**Find new energy scales by considering bound states?**

# Massive Mediators – Gauge boson exchange

Exchange of a **massive**  $Z'$  – Vector and axial-vector interaction

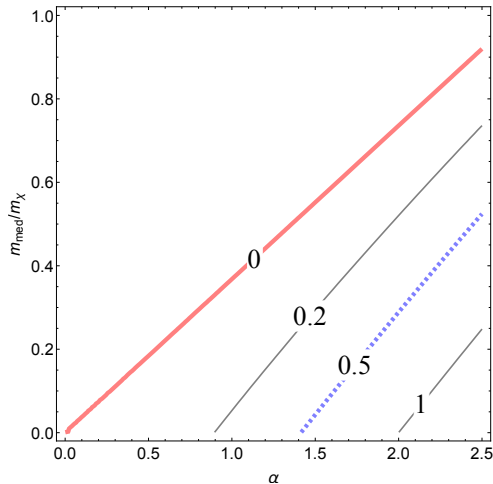
$$\mathcal{L} \supset g' \bar{\chi} \gamma_\mu (\lambda_0 + i\lambda_5 \gamma_5) \chi Z'^\mu$$

- ▶ Yukawa Potential – Spin-spin interaction

$$V(r) = \frac{e^{-m_{Z'} r}}{4\pi r} \left( \lambda_0^2 - \lambda_5^2 \vec{\sigma}_1 \cdot \vec{\sigma}_2 \right)$$

- ▶ Short range interaction
- ▶ Binding energy depends on the mediator mass
- ▶ Set a new upper bound on the Dark Matter mass using the binding energy

# Finding Characteristic Energy Scales – Yukawa potential



$\Delta/m_\chi$  contours

$$V(r) = \alpha \frac{e^{-m_{\text{med}} r}}{r}$$

$$T_{\text{freeze}} \sim \frac{m_\chi}{20}$$

- ▶ For  $|\Delta| \gg T_{\text{freeze}}$  (blue), Dark Matter dynamics is dominated by bound states formation!

# Summary and Outlook

- ▶ Perturbative Unitarity coupled with relic density sets powerful constraints on the scales of thermal Dark Matter models
- ▶ Combining unitarity constraints with current and future experimental results will allow to "corner" weakly interacting Dark Matter for large classes of models
- ▶ The unitarity bound corresponds to a fundamental change in the Dark Matter dynamics due to the formation of bound states
- ▶ For portals, the bound state dynamics can be fully derived  
⇒ Compute a physically motivated bound!
- ▶ New refined bounds can be computed for our models.  
Calculation of the potentials for the Higgs and squark portals in progress.