



# Recent improvements on Monte Carlo modelling at ATLAS



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# Introduction



- The ATLAS Run 2 experimental environment will be challenging more where many aspects becoming gradually more stressful, notably hard event and pile-up modelling :

## Run 1 conditions:

$\sqrt{s} = 900 \text{ GeV, 7 \& 8 TeV}$   
 $L \sim 7 \times 10^{33} \text{ cm}^{-2} \text{ s}^{-1}$   
Pile-up: up to 35  
Int. Lum  $\sim 30 \text{ fb}^{-1}$

## More Energy !



## Run 2 conditions:

$\sqrt{s} = 13 \text{ TeV}$   
 $L \sim 1.6 \times 10^{34} \text{ cm}^{-2} \text{ s}^{-1}$   
Pile-up: up to 50  
Int. Lum  $\sim 150 \text{ fb}^{-1}$

- MC event modelling is needed for both experimental understanding of QCD, V+jets etc as backgrounds to new physics; and to allow theory/data comparison as much as possible.
- ATLAS is using a wide range of Monte Carlo Generators where the MC parameters are expected to be **common** and describe **various processes** in **several types of collider** at **different energies**.
- Basically, the non-perturbative MC parameters of the models have to be tuned since they are not making predictions (Hadronization & UE). However, some parts of MC simulation like ME (especially NLO), parton shower, QED, etc can derive first principles.

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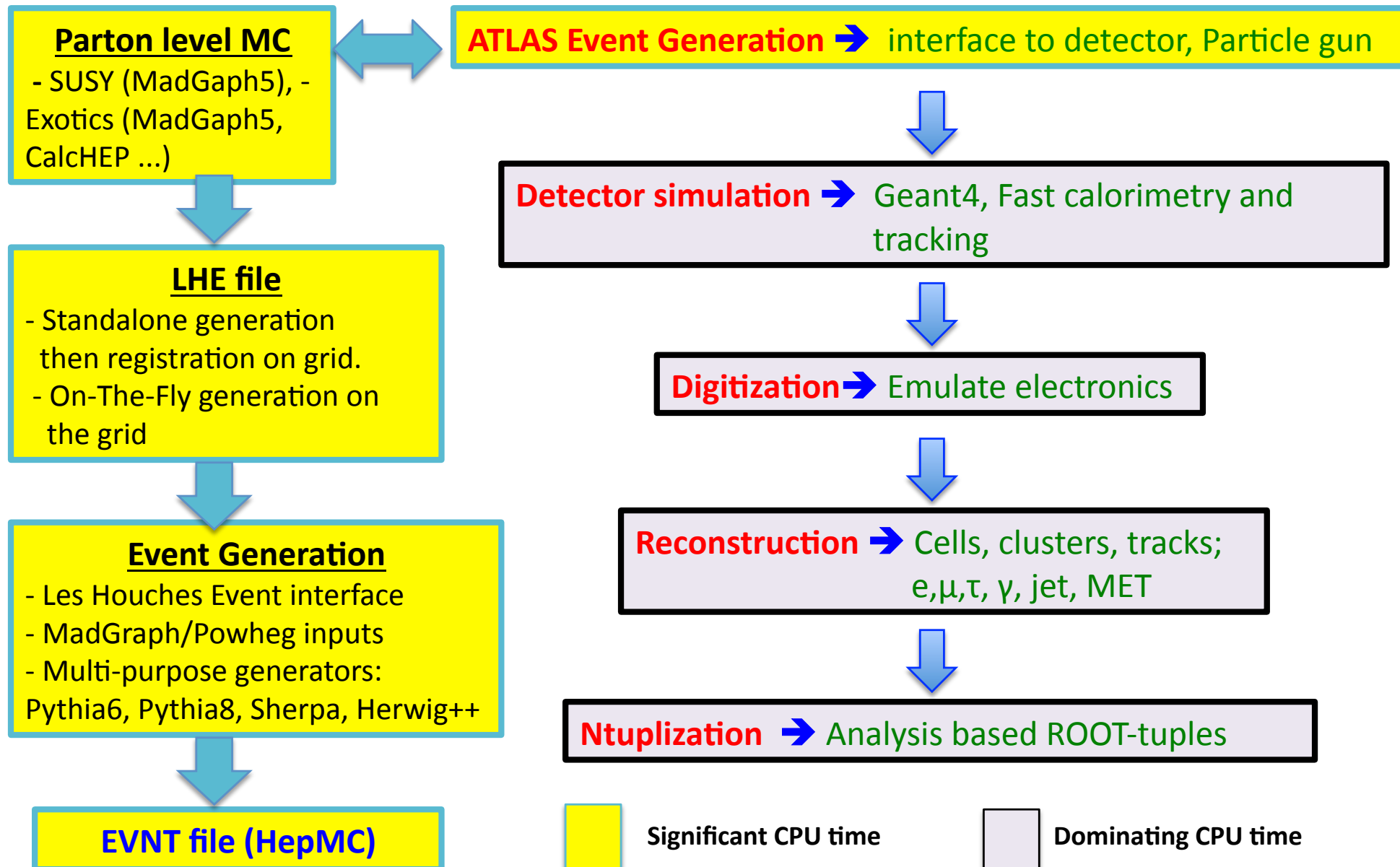
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**Today:** Just the recent tuning results of MC modeling @ ATLAS will be presented: A14 and ATTBAR tunes; hdamp setting for ttbar modeling

# The ATLAS simulation: state of the art

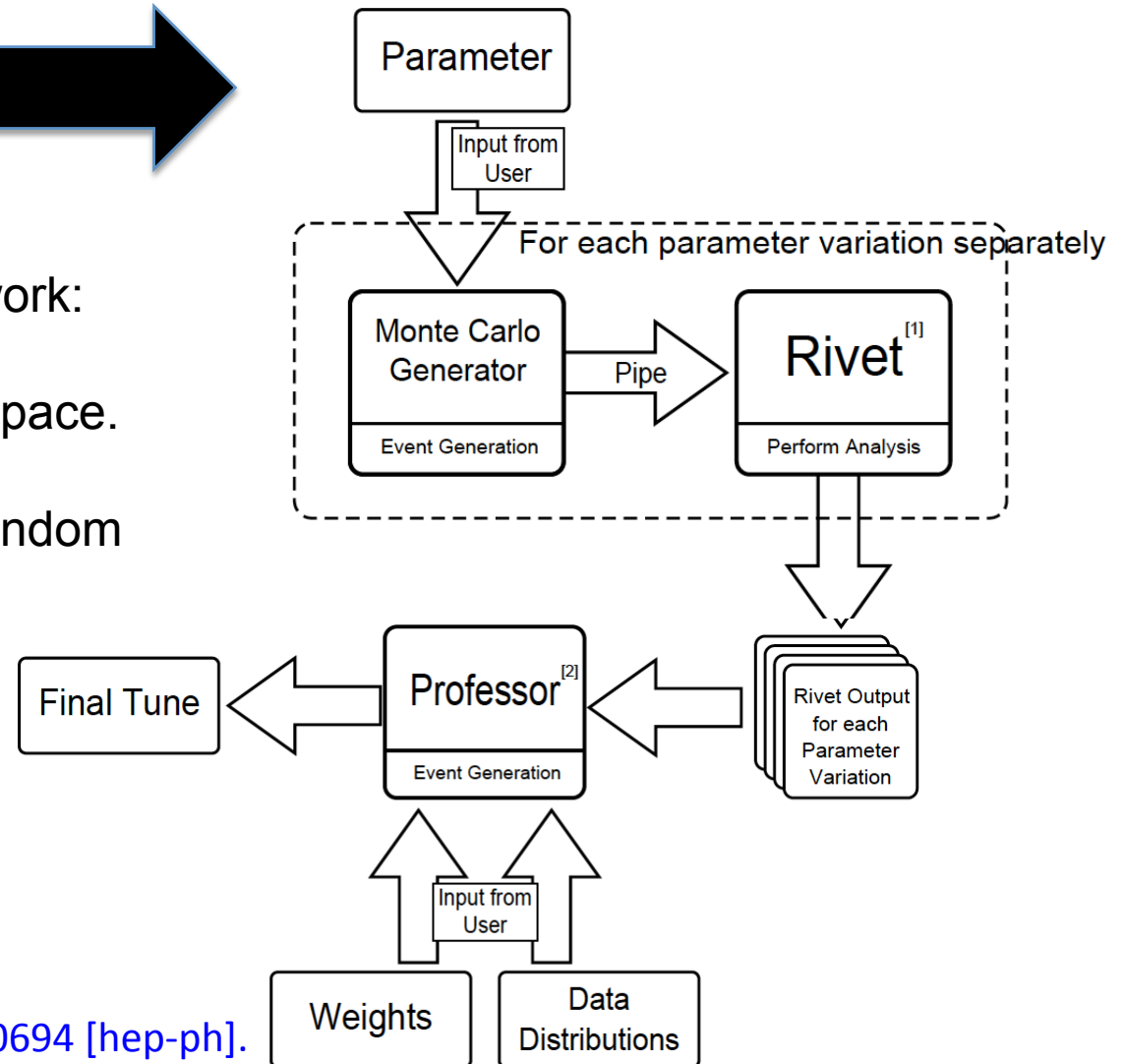


# MC tuning workflow in ATLAS



The main steps for the tuning framework:

- 1 - Definition of tuning parameter space.
- 2 - Event samples generation for random parameter points.
- 3 - Performing analyses for each event using Rivet ( see [1] ).
- 4 - Tuning using Professor ( see [2] ).



[1] - A. Buckley et al, Rivet user manual, 1003.0694 [hep-ph].

[2] - A. Buckley et al, Systematic event generator tuning for the LHC, 0907.2973 [hep-ph]

# A14: global tune of Shower and MPI



- A14 (ATLAS 14) is a set of Pythia8 tunes combining in a single step ISR, FSR and MPI tunes (500 points of 3 M events in 10 parameters hypercube).
- Exploiting all ATLAS 7 TeV data.
- Performing four separate tunes to LO PDFs: CTEQ6L1, MSTW2008LO, NNPDF23LO and HERAPDF15LO leads to four tunes designated as A14 tune series (at the end, there is no preference of any PDF, so we choose NNPDF23LO as the primary).
- The data used in this tuning includes ATLAS observables sensitive to UE, Jet structure and observables sensitive to additional jet emissions above lowest-order process ( ttbar gap fraction, 3/2 jet ratio, Z-pT ...etc).



| Parameter  | Definition                                  | Sampling range |
|--|---|----------------|
| <code>SigmaProcess:alphaSvalue</code>            | The $\alpha_S$ value at scale $Q^2 = M_Z^2$ | 0.12 – 0.15    |
| <code>SpaceShower:pT0Ref</code>                  | ISR $p_T$ cutoff                            | 0.75 – 2.5     |
| <code>SpaceShower:pTmaxFudge</code>              | Mult. factor on max ISR evolution scale     | 0.5 – 1.5      |
| <code>SpaceShower:pTdampFudge</code>             | Factorisation/renorm scale damping          | 1.0 – 1.5      |
| <code>SpaceShower:alphaSvalue</code>             | ISR $\alpha_S$                              | 0.10 – 0.15    |
| <code>TimeShower:alphaSvalue</code>              | FSR $\alpha_S$                              | 0.10 – 0.15    |
| <code>BeamRemnants:primordialkThard</code>       | Hard interaction primordial $k_\perp$       | 1.5 – 2.0      |
| <code>MultipartonInteractions:pT0Ref</code>      | MPI $p_T$ cutoff                            | 1.5 – 3.0      |
| <code>MultipartonInteractions:alphaSvalue</code> | MPI $\alpha_S$                              | 0.10 – 0.15    |
| <code>BeamRemnants:reconnectRange</code>         | CR strength                                 | 1.0 – 10.0     |

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| Param  | CTEQ  | MSTW  | NNPDF | HERA  |
|--|-------|-------|-------|-------|
| <code>SigmaProcess:alphaSvalue</code>            | 0.144 | 0.140 | 0.140 | 0.141 |
| <code>SpaceShower:pT0Ref</code>                  | 1.30  | 1.62  | 1.56  | 1.61  |
| <code>SpaceShower:pTmaxFudge</code>              | 0.95  | 0.92  | 0.91  | 0.95  |
| <code>SpaceShower:pTdampFudge</code>             | 1.21  | 1.14  | 1.05  | 1.10  |
| <code>SpaceShower:alphaSvalue</code>             | 0.125 | 0.129 | 0.127 | 0.128 |
| <code>TimeShower:alphaSvalue</code>              | 0.126 | 0.129 | 0.127 | 0.130 |
| <code>BeamRemnants:primordialKThard</code>       | 1.72  | 1.82  | 1.88  | 1.83  |
| <code>MultipartonInteractions:pT0Ref</code>      | 1.98  | 2.22  | 2.09  | 2.14  |
| <code>MultipartonInteractions:alphaSvalue</code> | 0.118 | 0.127 | 0.126 | 0.123 |
| <code>BeamRemnants:reconnectRange</code>         | 2.08  | 1.87  | 1.71  | 1.78  |

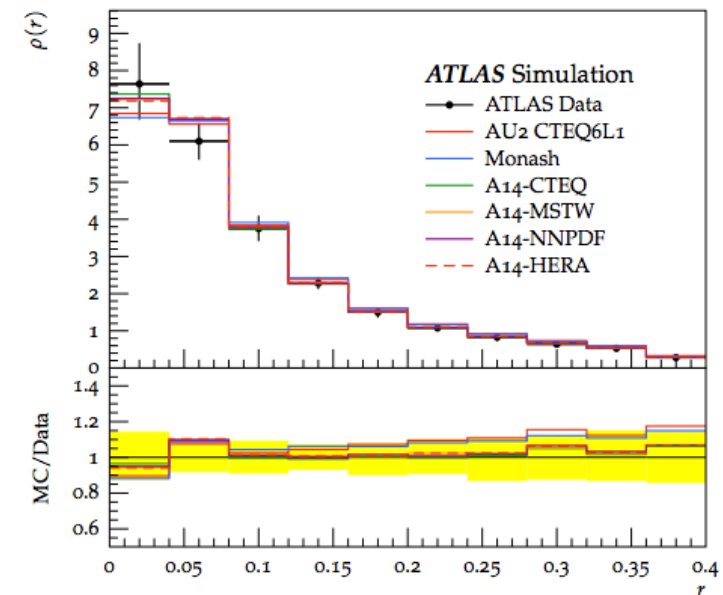
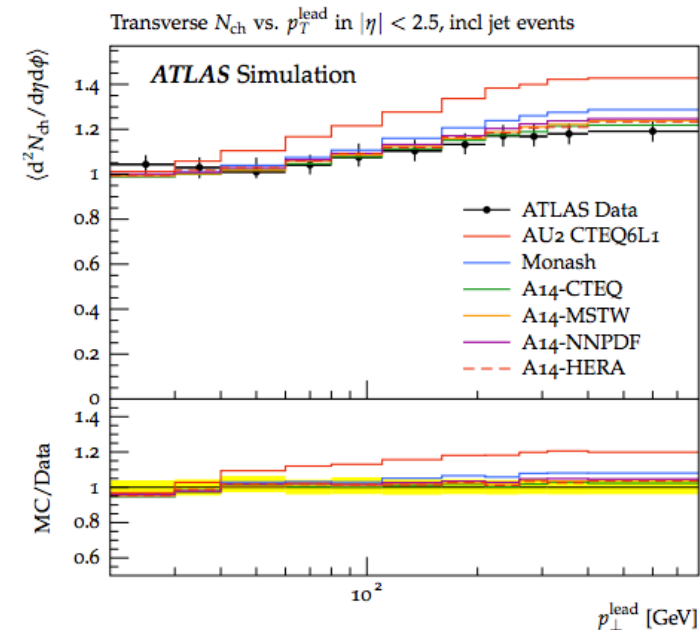
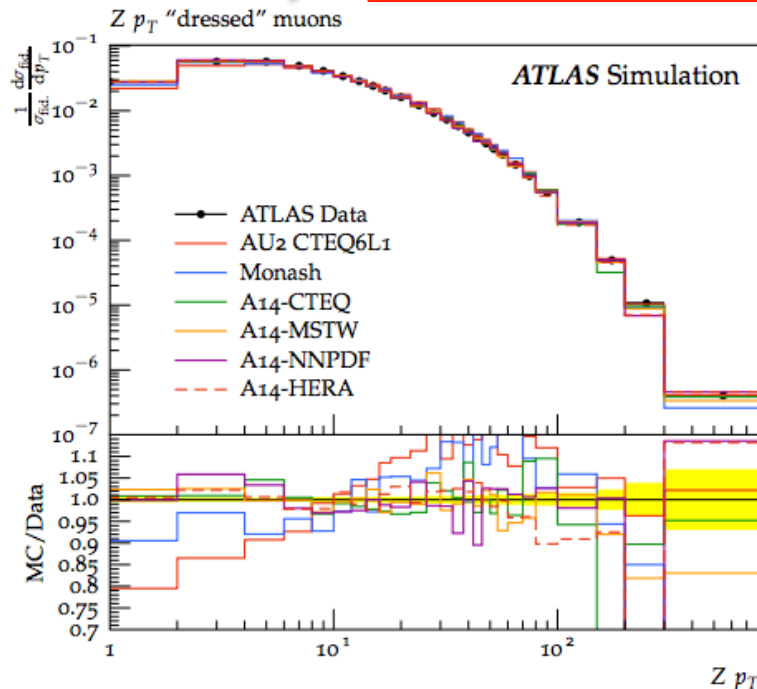
# A14: global tune of Shower and MPI



- This tune uses Monash as starting point and thus builds on the work of the others.
- A14-NNPDF tune gives a good performance on jets, W/Z jets and ttbar processes.
- More work is on-going for a matched shower-NLO tune, but that is accounting for the additional emission in higher order ME and it is not adding higher order corrections to the shower.



**ATL-PHYS-PUB-2014-021**





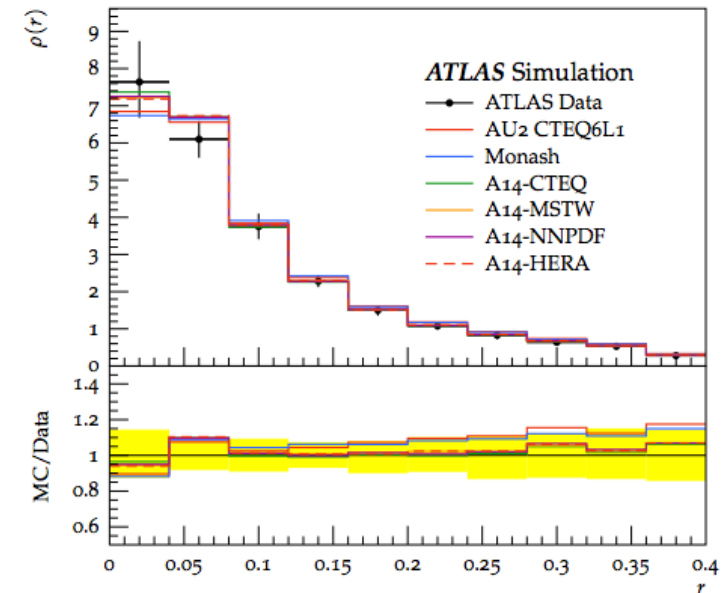
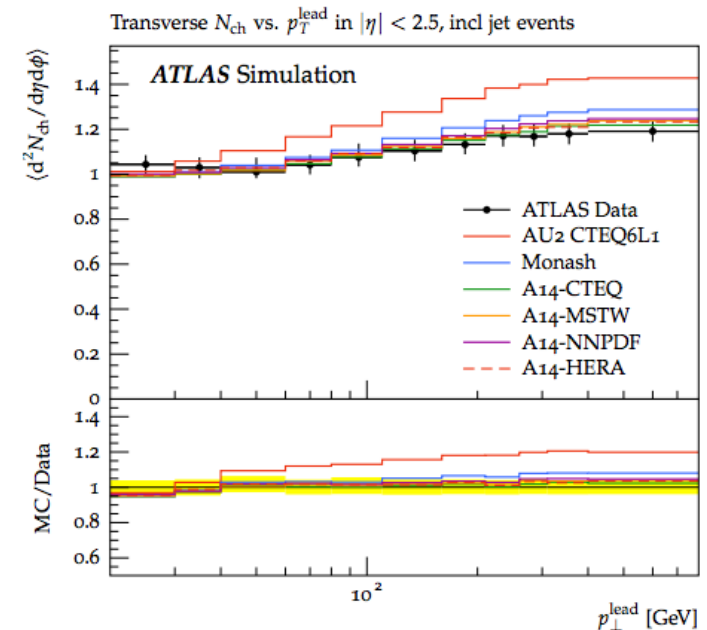
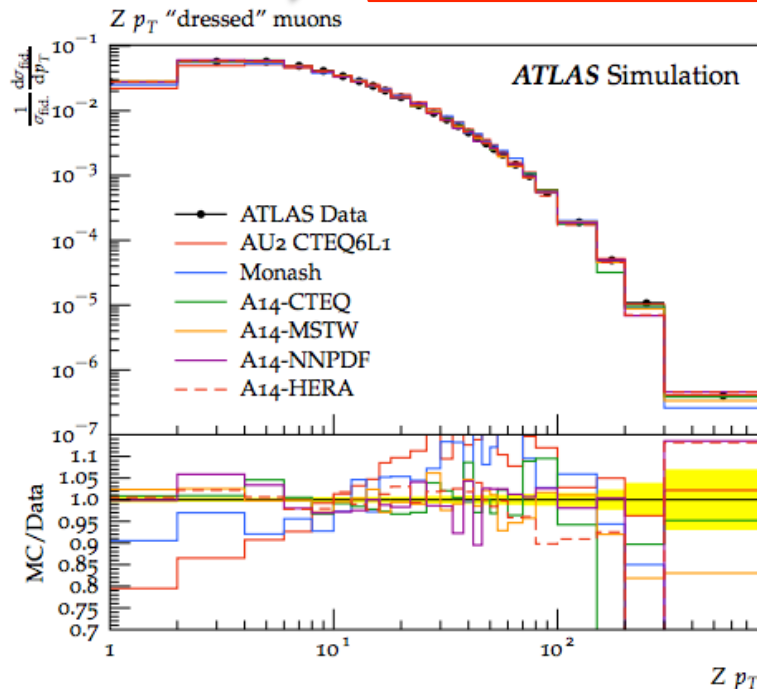
# A14: global tune of Shower and MPI



- A14 → suitable for use in high  $p_T$  processes and using four different PDFs. These tunes simultaneously optimised MPI and parton shower (ISR,FSR) parameters since these model features affect data observables in combination.




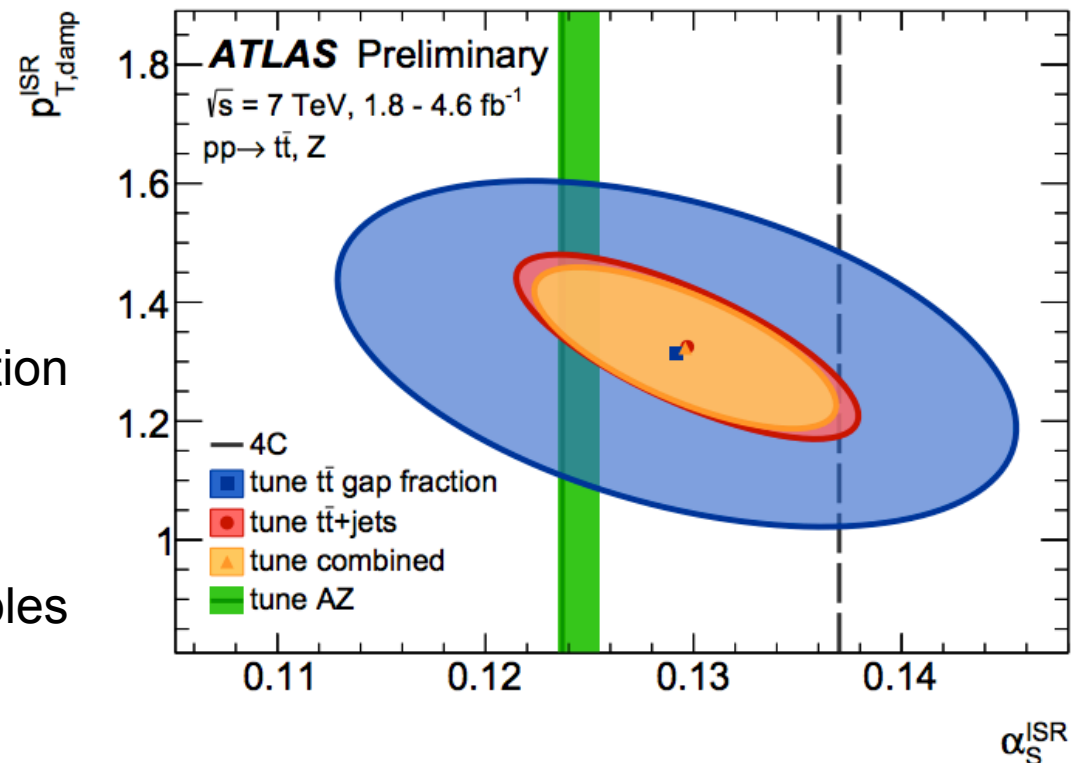
**ATL-PHYS-PUB-2014-021**



# ATTBAR tune of $t\bar{t}$ observables



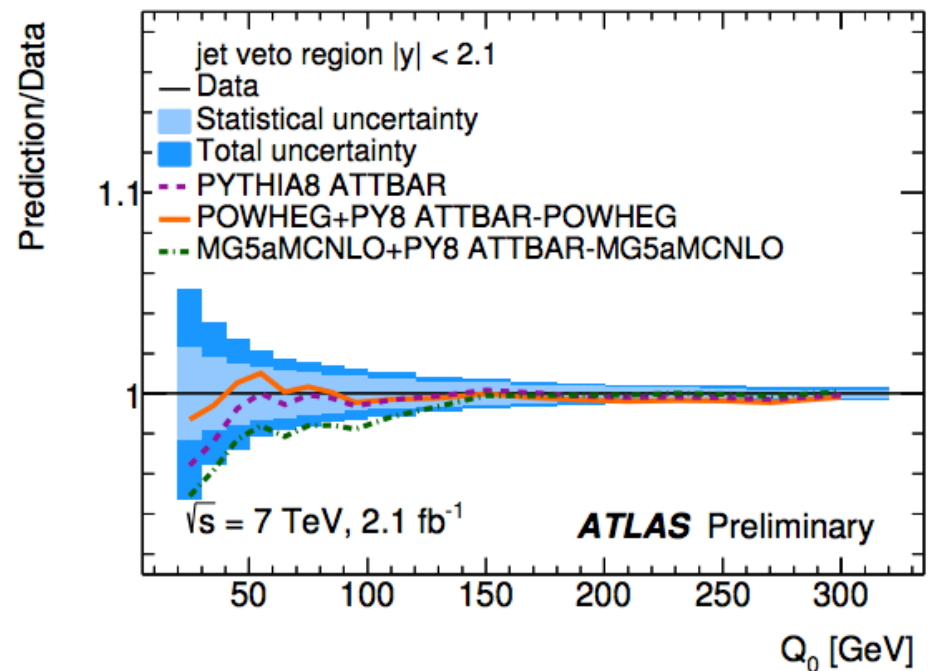
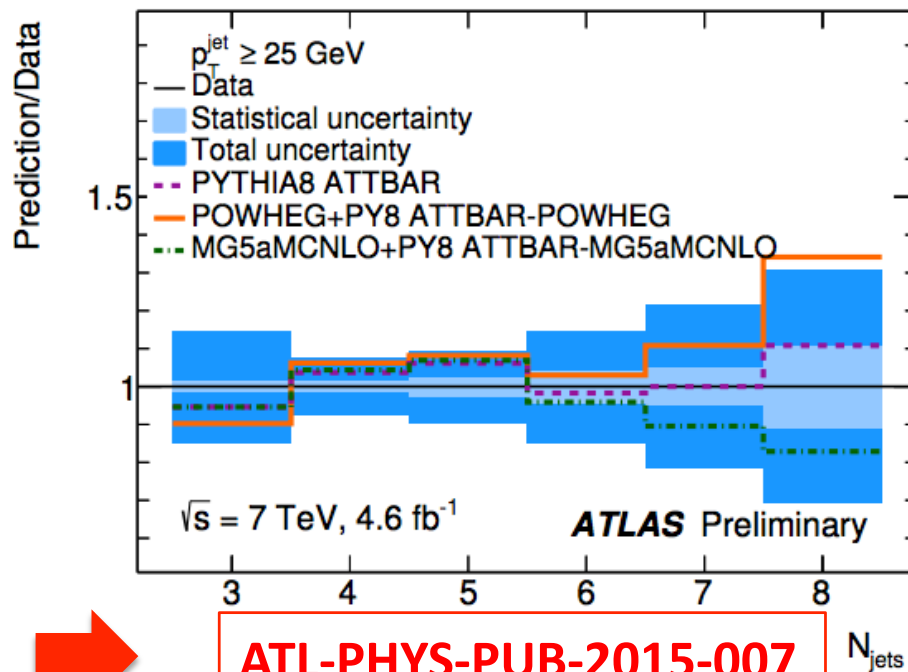
- In many  $t\bar{t}$  measurements, the ISR & FSR modelling is crucial since it is one of the dominant uncertainties.
  - Significant reduction of the modelling uncertainty can be achieved by verifying the universality of the parton shower model between Z boson and  $t\bar{t}$  production.
  - A simultaneous tune for ISR & FSR parameters is applied : **ATTBAR**
-  **ATL-PHYS-PUB-2015-007**
- For the first time in the context of PS tuning in MC simulations, the correlation of the experimental uncertainties is included in  $\chi^2$  definition.
  - The sensitivity of the various observables to the PS parameters is improved.



# ATTBAR tune of ttbar observables



- The ATTBAR tune is applied to the NLO PS generators POWHEG & MadGraph5\_aMC@NLO.
- By lowering hdamp parameter (a variable which effectively regulates the high-pT radiation in POWHEG) to  $1.8 m_t$ , good agreement between ttbar data and Powheg+Pythia8 predictions.
- However this opens-up more important points: FSR cut-off in Pythia8, Global recoil in aMC@NLO, b-jet shapes.



**ATL-PHYS-PUB-2015-007**

# ttbar modeling with hdamp and scale variation

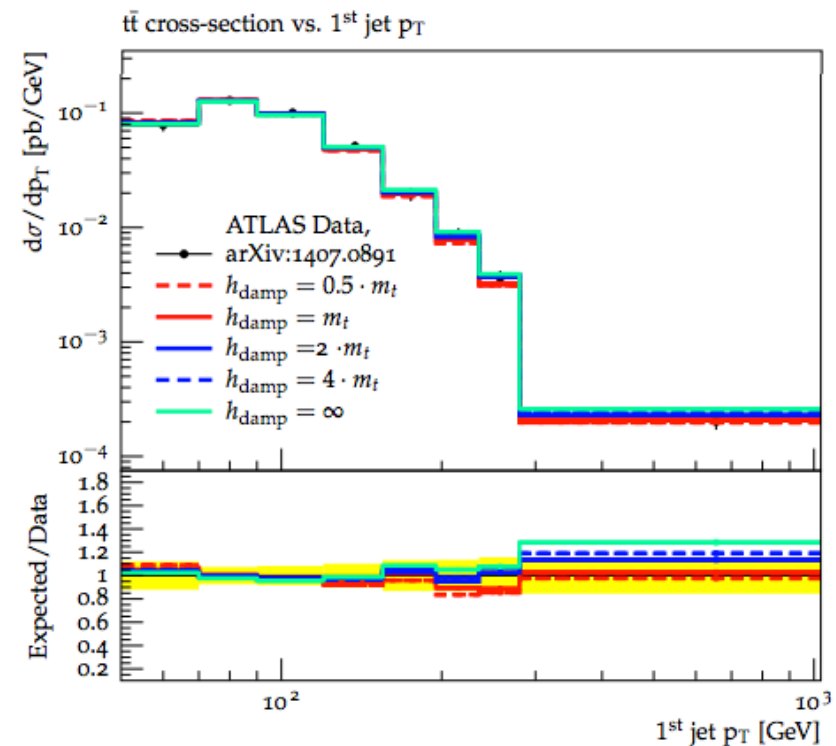
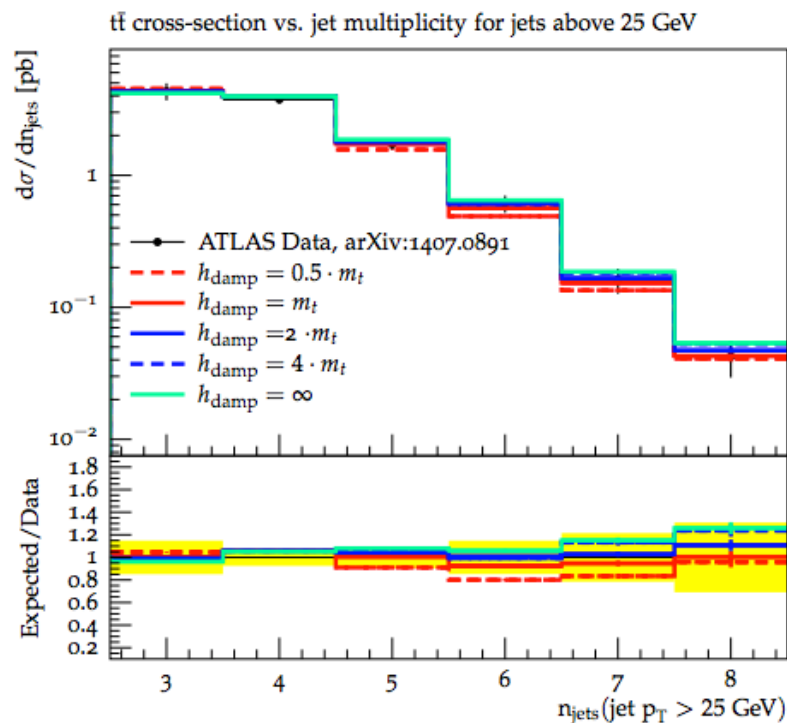


## • @ Run1:

- Powheg(CT10)+Pythia6 with P2011C tune and CTEQ6L1 is being the default.
- Switching  $h_{\text{damp}} = \text{infinity}$  to  $h_{\text{damp}} = m_{\text{top}}$  is found best to describe the data and improves  $T(\text{ttbar})$  but not  $p_T(\text{top})$ .



ATL-PHYS-PUB-2015-002

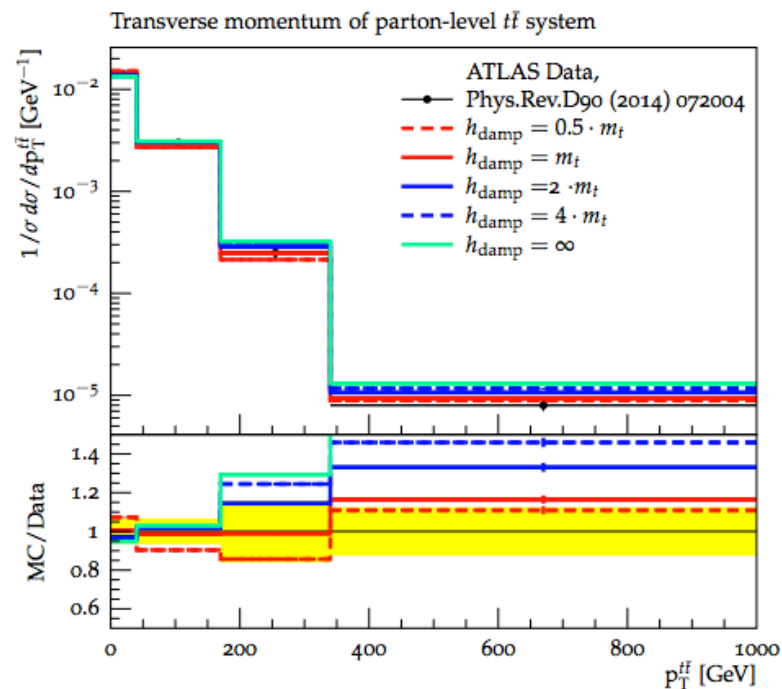
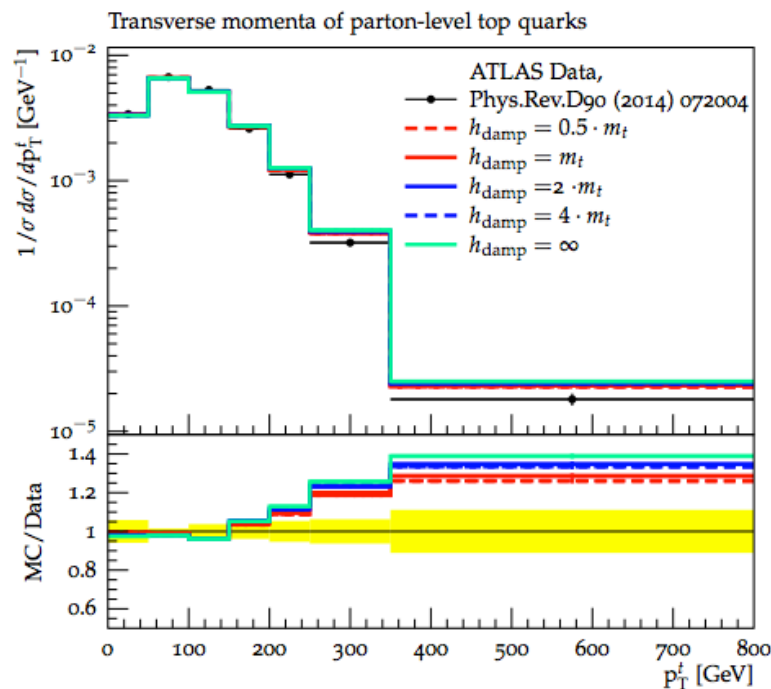


# ttbar modeling with hdamp and scale variation



## • @ Run1:

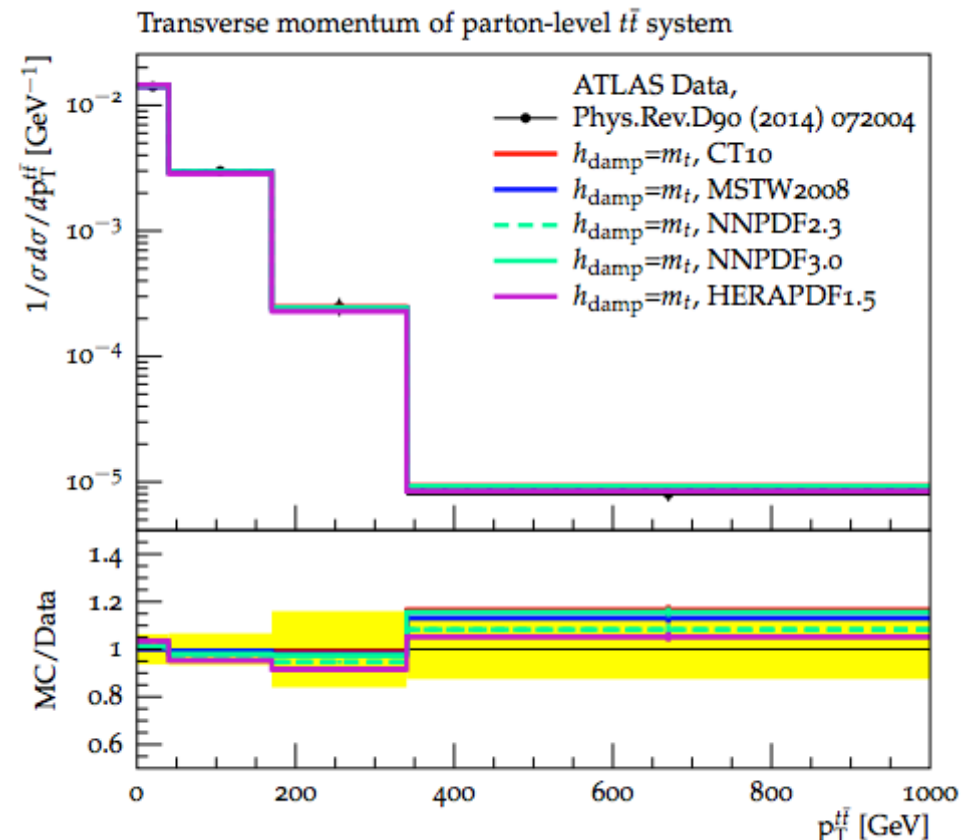
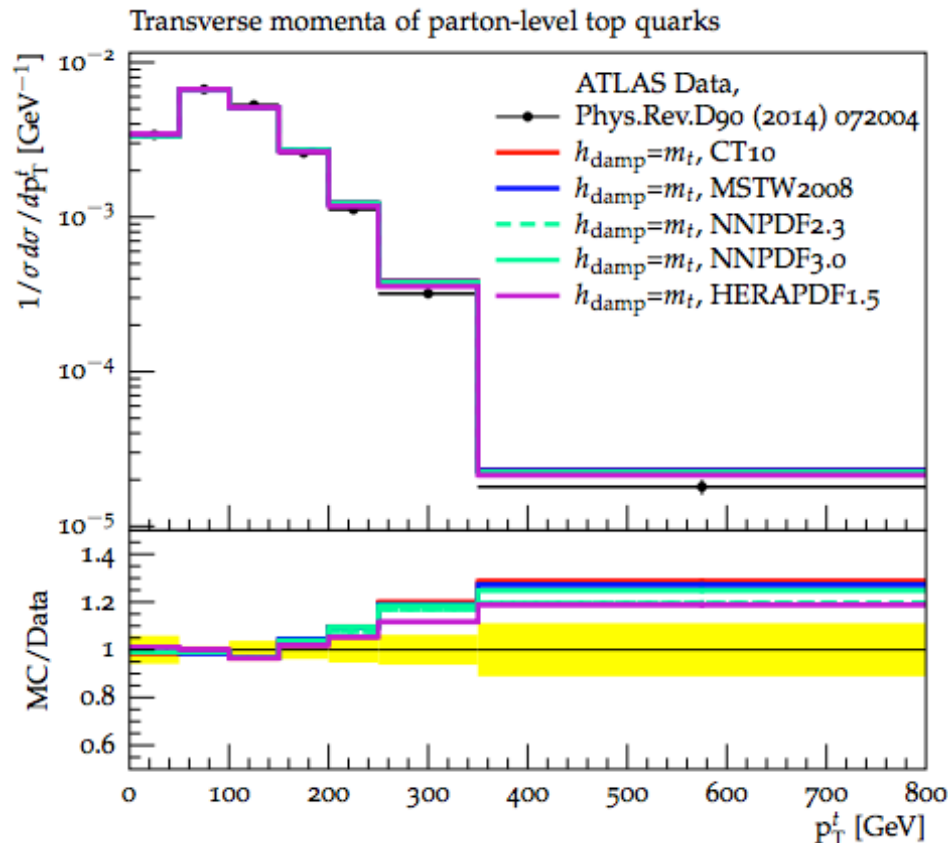
- Powheg(CT10)+Pythia6 with P2011C tune and CTEQ6L1 is being the default.
  - Switching  $h_{\text{damp}} = \text{infinity}$  to  $h_{\text{damp}} = m_{\text{top}}$  is found best to describe the data and improves  $T(\text{ttbar})$  but not  $p_T(\text{top})$ .
- The main improvement here is ttbar  $p_T$ , the  $p_T$  top distribution (which is highly correlated with  $m_{\text{tt}}$ ) is less improve.



# ttbar modeling with hdamp and PDF variation



- The choice of the PDF can have quite significant impact on the kinematics of the ttbar production.
- For this reason, a variety of NLO PDF sets have been studied.
- But again, the discrepancy between the prediction and the experimental data for high top quark pT cannot be solved by any of the used PDF sets !



# ttbar modeling with new C++ generators

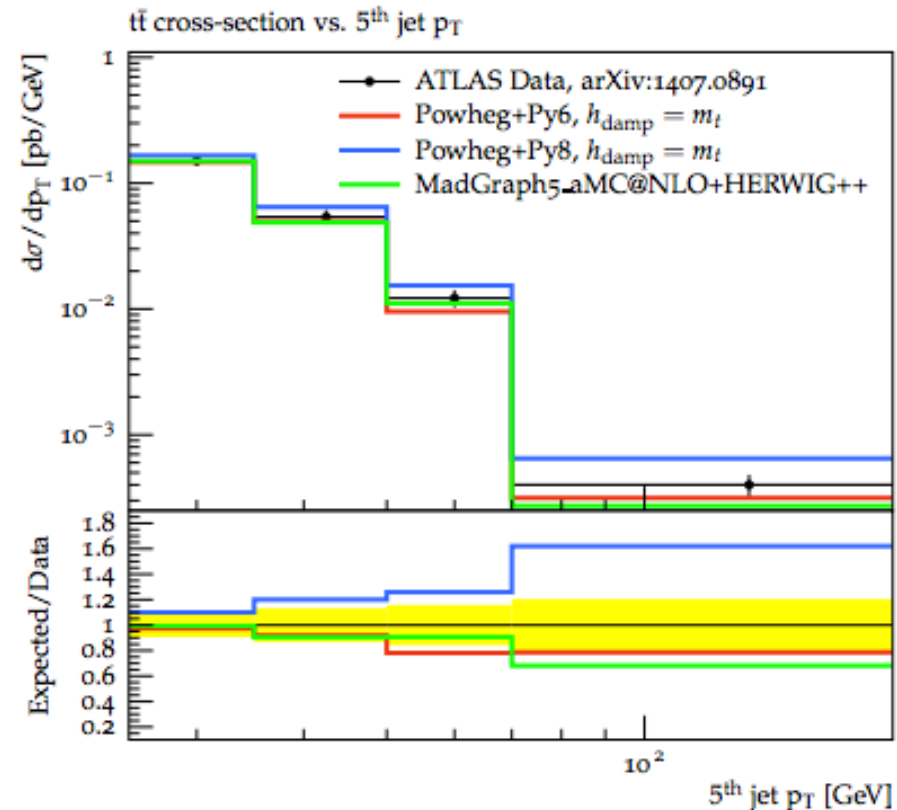
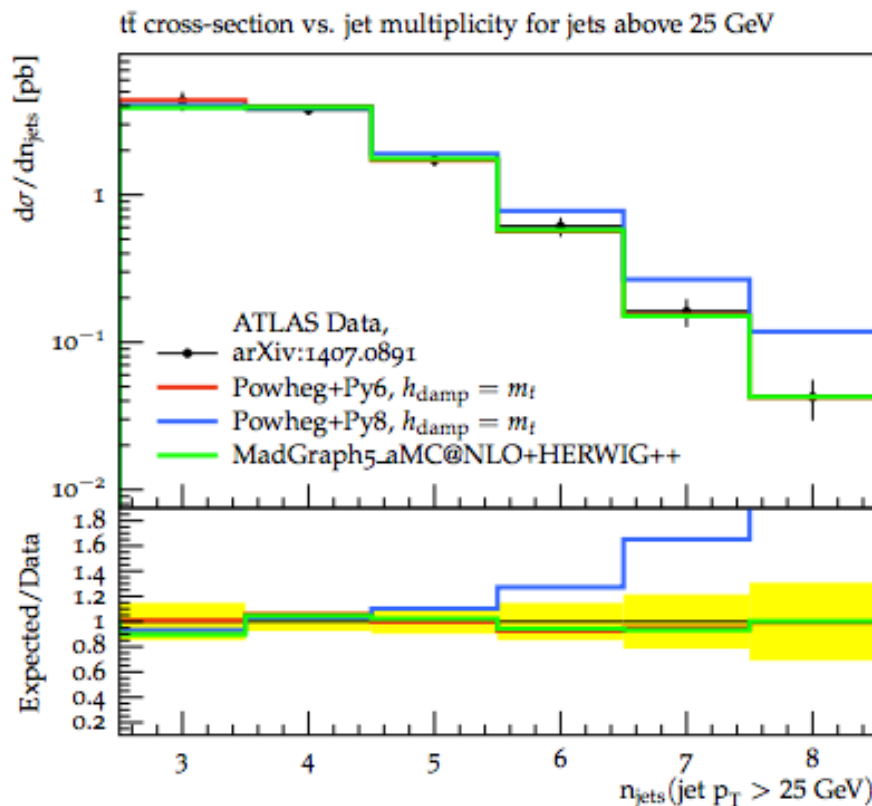


- The ATLAS baseline ttbar sample (Powheg+Pythia6 with  $h_{\text{damp}} = m_{\text{top}}$ ) is compared to new generator setups: Powheg+Pythia8 using AU2 tune (also with  $h_{\text{damp}} = m_{\text{top}}$ ) and MadGraph5\_aMC@NLO+Herwig++ with UEEE5 tune.



Powheg+Pythia8 → produces too many additional jet

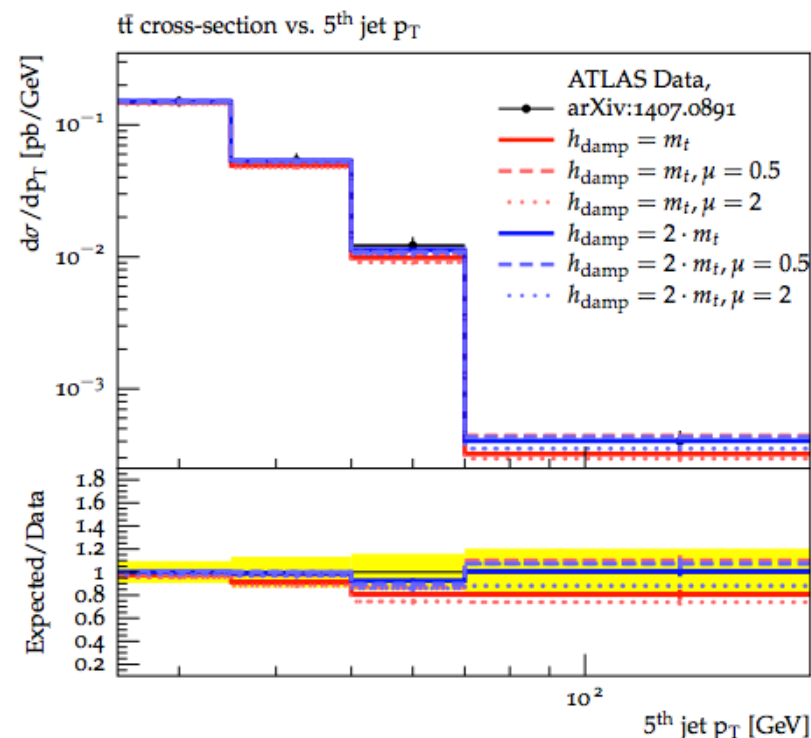
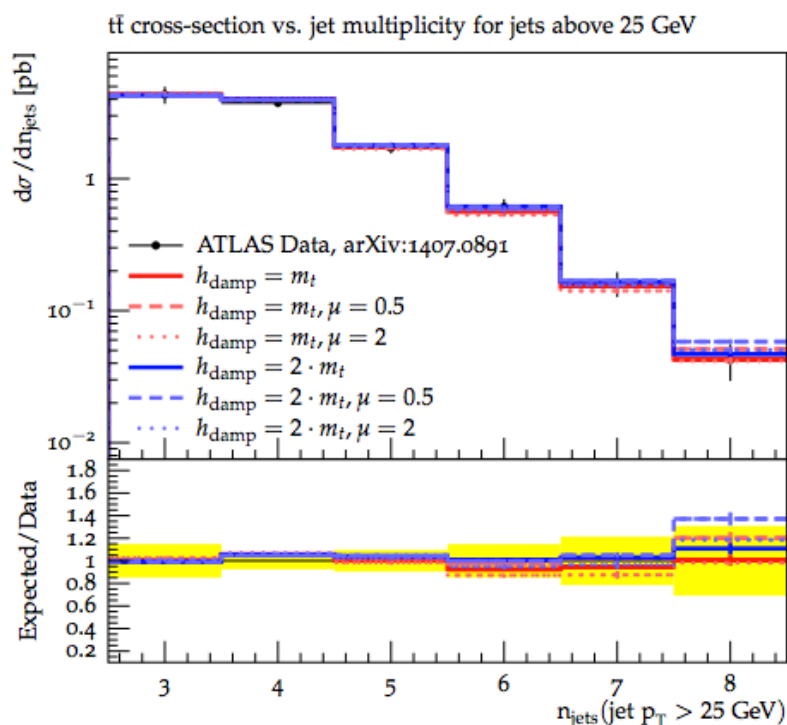
MadGraph5\_aMC@NLO+Herwig++ → give a better description of the data



# New radiation systematics for $t\bar{t}$ production



- 2 samples of Powheg+Pythia6 (with  $h_{\text{damp}} = m_{\text{top}}$  and  $h_{\text{damp}} = 2 m_{\text{top}}$ ) are compared to access the systematic variations connected to the extra radiation activity.
- For each sample the factorization and renormalization scale are simultaneously ( $\mu = \mu_f = \mu_r$ ) varied by a factor of 0.5 (for  $h_{\text{damp}} = m_{\text{top}}$ ) and 2 (for  $h_{\text{damp}} = 2 m_{\text{top}}$ ).
- Variation of the scale in the shower: use of Perugia2012radLo and radHi





# Conclusions



- Due to time limit, I didn't cover the MC studies in Higgs production for ATLAS Run 2. See **ATL-PHYS-PUB-2014-022**
- The ATLAS recent performed tunes: AT14 as baseline tune and ATTBAR as specific tune for ttbar production.
- The new  $h_{\text{damp}} = m_{\text{top}}$  setting is found to best describe the data, although none of the settings provide a good description of  $p_T(\text{top})$ .
- The MC modeling & production is a very expensive and time-consuming activity. Therefore one have to be precise from the beginning of MC generator level.
- With the LHC-Run2 we will have completely new data and new era starting.
- A lot of efforts were/are needed to provide proper MC estimates for signal and backgrounds at 13 (14) TeV.



# Thanks for your attention !

Backup

# A14 tunes



- Observable–weight combinations used for the tuning

| Observable   | Fit range                | Weight |
|--|--------------------------|--------|
| <b>Track jet properties [6]</b>  |                          |        |
| Charged jet multiplicity (50 distributions)  |                          | 10.0   |
| Charged jet $z$ (50 distributions)   |                          | 10.0   |
| Charged jet $p_T^{\text{rel}}$ (50 distributions)  |                          | 10.0   |
| Charged jet $\rho_{\text{ch}}(r)$ (50 distributions)                                       |                          | 10.0   |
| <b>Jet shapes [7]</b>  |                          |        |
| Jet shape $\rho$ (59 distributions)  |                          | 10.0   |
| <b>Dijet decorr [10]</b>   |                          |        |
| Decorrelation $\Delta\phi$ (9 distributions)   | $\Delta\phi > 0.75$      | 20.0   |
| <b>Multijets [12]</b>  |                          |        |
| 3-to-2 jet ratios (8 distributions)  |                          | 100.0  |
| <b><math>p_T^Z</math> [13, 14]</b>   |                          |        |
| Z-boson $p_T$ (20 distributions)   | $p_T^Z < 50 \text{ GeV}$ | 10.0   |
| <b>Substructure [9]</b>  |                          |        |
| Jet mass, $\sqrt{d_{12}}$ , $\sqrt{d_{23}}$ , $\tau_{21}$ , $\tau_{23}$ (36 distributions) |                          | 5.0    |
| <b><math>t\bar{t}</math> gap [11]</b>  |                          |        |
| Gap fraction vs $Q_0$ , $Q_{\text{sum}}$ for $ y  < 0.8$                                   |                          | 100.0  |
| Gap fraction vs $Q_0$ , $Q_{\text{sum}}$ for $0.8 <  y  < 1.5$                             |                          | 80.0   |
| Gap fraction vs $Q_0$ , $Q_{\text{sum}}$ for $1.5 <  y  < 2.1$                             |                          | 40.0   |
| Gap fraction vs $Q_0$ , $Q_{\text{sum}}$ for $ y  < 2.1$                                   |                          | 10.0   |

and more. See [ATL-PHYS-PUB-2014-021](#) for further details.