

Baryon Number Violation and the Dark Sector

Hooman Davoudiasl

HET Group, Brookhaven National Laboratory

Based on:

H. D., Phys. Rev. Lett. 114, no. 5, 051802 (2015), [1409.4823]

H.D. and Y. Zhang, arXiv:1504.07244 [hep-ph]

Pheno 2015, University of Pittsburgh, May 4-6, 2015

- SM an impressive description of natural phenomena, yet incomplete
- Neutrino masses and dark matter (DM)
- Both can be addressed with “dark” states X uncharged under SM
 - X : Heavy Majorana neutrino N for seesaw or DM state
 - X may couple to “dark” forces
- X also could mediate violations of SM accidental symmetries
- We will consider B violating interactions
 - (I) $m_X \lesssim 1$ GeV: $\frac{(Xudd)_R}{M^2}$ (light quarks)
 - (II) $m_X \gtrsim 100$ GeV ($X = N$): $\frac{(Ntbb)_R}{\Lambda^2}$ (top and bottom)

Sample of recent works on related subjects: [Cheung, Ishiwata 1304.0468](#); [Baldes, Bell, Millar, Petraki, Volkas, 1410.0108](#); [Monteux, Shin, 1412.5586](#); [Dev, Mohapatra, 1504.07196](#)

Case (I): $m_X \lesssim 1 \text{ GeV}$

- “Neutron portal” $\frac{(Xudd)_R}{M^2}$ HD, Morrissey, Sigurdson, Tulin, 2010
- Potentially contributing to nucleon decay, e.g. $p \rightarrow X\pi^+$
- For $m_X \ll m_N - m_{\text{meson}}$ hard to distinguish from $p \rightarrow \pi^+\nu$
- Focus on $(\text{few} \times 100 \text{ MeV}) \lesssim m_X \lesssim m_N$
- For $\mathcal{O}(100 \text{ MeV})$ meson momenta can use chiral perturbation theory
Claudson, Wise, Hall, 1982

$$\mathcal{L}_{\Delta B=0} = \frac{D+F}{2f_\pi} \left[\left(\frac{3F-D}{D+F} \right) \frac{\partial_\mu \eta}{\sqrt{3}} - \partial_\mu \pi^0 \right] \bar{n} \gamma^\mu \gamma_5 n + \left(\frac{D+F}{\sqrt{2}f_\pi} \right) \partial_\mu \pi^- \bar{n} \gamma^\mu \gamma_5 p + \dots$$

$$\mathcal{L}_{\Delta B=1} = \beta c_1 \bar{X}^c \left\{ n_R - \frac{i}{f_\pi} \left[\frac{p_R}{\sqrt{2}} \pi^- + \frac{n_R}{2} (\sqrt{3}\eta - \pi^0) \right] \right\} \quad ; \quad (c_1 \equiv 1/M^2)$$

$D = 0.80, F = 0.47, \beta = 0.012 \pm 0.0026 \text{ GeV}^3$ (lattice, e.g., Aoki et al., 2008), $f_\pi \simeq 92.2 \text{ MeV}$

- $p \rightarrow X\pi^+$ (solid)

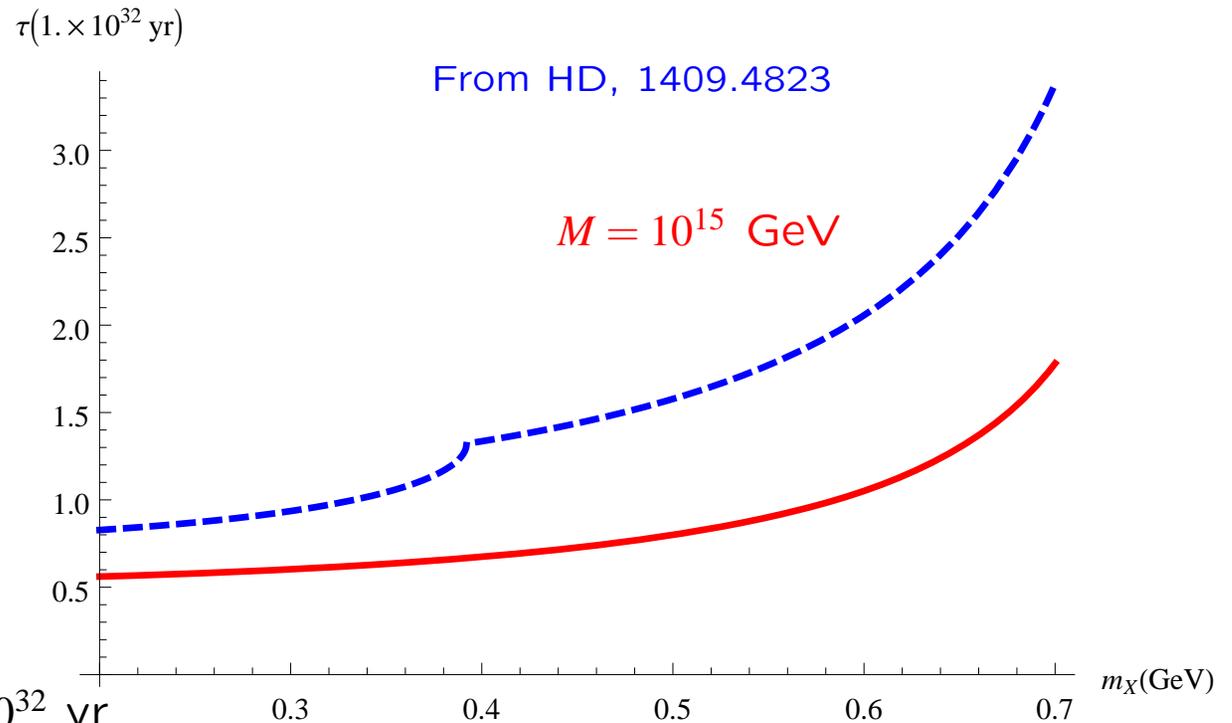
- $n \rightarrow X\pi^0, \eta$ (dashed)

- Discontinuity: η threshold

- Current bounds:

- $\tau(p \rightarrow \pi^+\bar{\nu}) > 1.6 \times 10^{31}$ yr
Soudan 2 Collaboration, 2000

- $\tau[n \rightarrow \pi^0(\eta)\bar{\nu}] > 1.12(1.58) \times 10^{32}$ yr
McGrew et al., 1999



- X final states: different kinematics, bounds not directly applicable

Example: $p \rightarrow \pi^+\bar{\nu}$, with $p_{\pi^+} \simeq 459$ MeV

- Simulations: π^+ exiting Fe nucleus on average lose $\sim 1/2$ their momentum

- Soudan 2 cut on π^+ momentum $p_{\pi^+} \in [140, 420]$ MeV

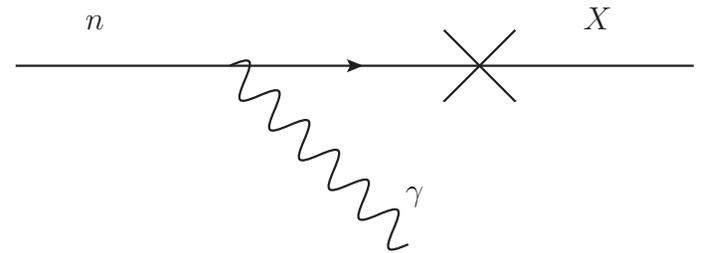
- $p \rightarrow X\pi^+$ with $m_X = 600$ MeV $\Rightarrow p_{\pi^+} \simeq 251 < 2 \times 140$ MeV

- Neutron decay $n \rightarrow X\gamma$ possible through n magnetic moment

- For $m_N - m_X < m_\pi$, $n \rightarrow X\gamma$ dominates

- PDG: $\tau(n \rightarrow \nu\gamma) > 2.8 \times 10^{31}$ yr

$\Rightarrow M \gtrsim 10^{14}$ GeV (softer γ can affect efficiency)



- X charged under $U(1)_d$, mediated by light Z_d may yield $n \rightarrow XZ_d$

- $U(1)_d$ coupling g_d ; kinetic mixing with photon or mass mixing with Z

- Possibly motivated by DM models or $g_\mu - 2$ data

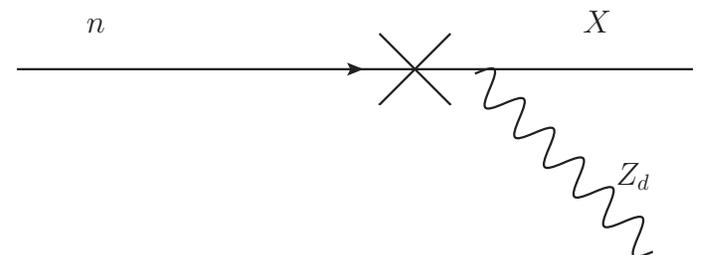
[Arkani-Hamed, Finkbeiner, Slatyer, Weiner, 2009](#); [Pospelov 2009](#)

- $Q_d = 1$, $\alpha_d = \alpha$, $M = 10^{15}$ GeV, $m_X = 700$ MeV, and $m_{Z_d} = 50$ MeV

$\Rightarrow \tau(n \rightarrow X\pi^0) \simeq 3.4 \times 10^{32}$ yr ; $\tau(n \rightarrow XZ_d) \simeq 1.7 \times 10^{33}$ yr

- $\text{Br}(Z_d \rightarrow e^+e^-) \sim 1 \Rightarrow$ dilepton resonance

$\tau(n \rightarrow e^+e^-\nu) > 2.6 \times 10^{32}$ yr



Case (II): $m_X \gtrsim 100 \text{ GeV}$

- Direct baryogenesis via Majorana neutrinos: *Dexiogenesis*

HD, Zhang, 1504.07244

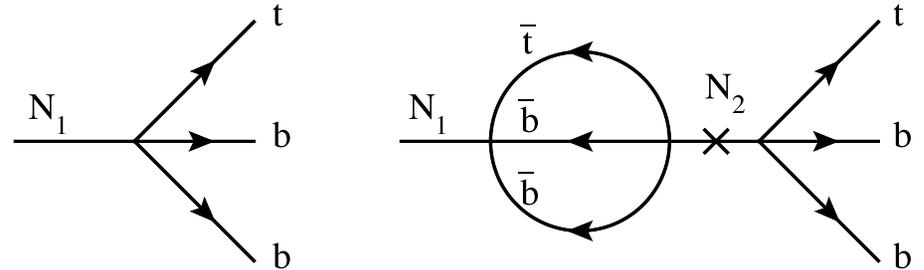
(dexios: Greek for right hand)

- Seesaw mechanism with Yukawa couplings $y_N \lesssim 10^{-6}$

- Baryon number violating effective operator $\lambda_a \frac{[\bar{N}_a^c P_R b][\bar{t}^c P_R b]}{\Lambda^2}$

Minimal choice $a = 1, 2$ for realistic seesaw; $\lambda_a \in \mathbb{C}$

- Baryogenesis at 2-loop
- Non-thermal production of N_1
- N_1 decaying out-of-equilibrium



- Baryon asymmetry:

$$\varepsilon = \frac{\text{Im}(\lambda_1^2 \lambda_2^{*2})}{3072 \pi^3 |\lambda_1|^2} \left(\frac{M_1}{\Lambda} \right)^4 \frac{M_1 M_2}{(M_2^2 - M_1^2)}$$

$n_B/s \sim 10^{-10}$

- $\varepsilon \gtrsim 10^{-8} \Rightarrow M_1/\Lambda \gtrsim 0.1 \quad (T_{\text{reheat}} \sim 1 \text{ GeV})$

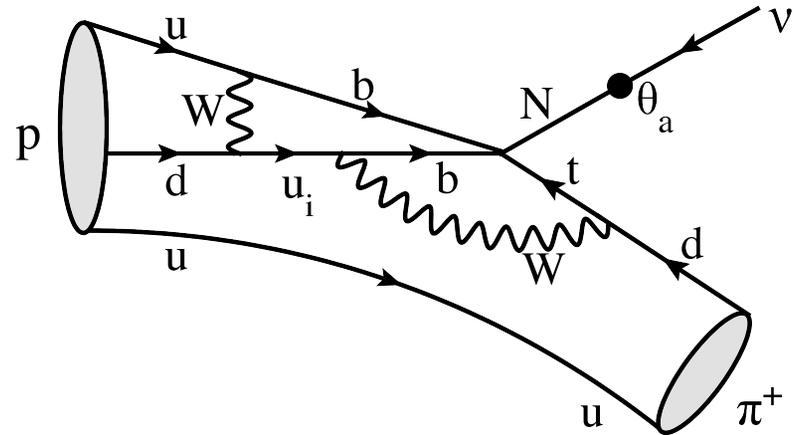
- Important constraints from nucleon decay

- $\Gamma(p \rightarrow \pi^+ \nu) = \frac{(1+g_A)^2 \alpha_h^2 m_p}{32\pi f_\pi^2} |\xi|^2$

- $g_A = 1.27$, $f_\pi = 131 \text{ MeV}$, $\alpha_h \approx -0.01125 \text{ GeV}^3$ (lattice, [Aoki et al., 2006](#))

- $\xi \approx \frac{\Lambda_{qcd} G_F^2 m_t m_b^2 V_{td}^2 V_{ub}^* V_{tb}^*}{(16\pi^2)^2 \Lambda^2} \lambda_a \theta_a$

- θ_a : N_a - ν_L mixing



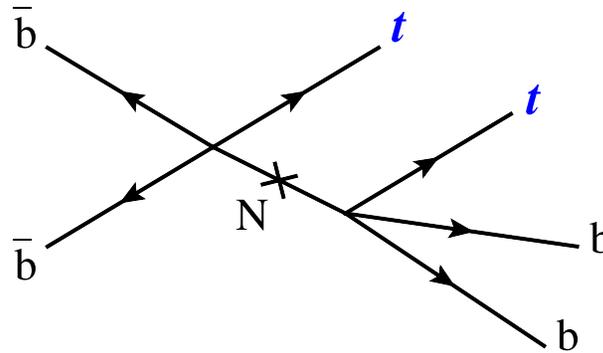
$$\tau(p \rightarrow \pi^+ \nu) \approx 2.5 \times 10^{32} \text{ yr} \left(\frac{\Lambda / \sqrt{\lambda_a}}{1.5 \text{ TeV}} \right)^4 \left(\frac{\theta_a}{10^{-6}} \right)^{-2}$$

- Within reach of proposed nucleon decay experiments
- Experimental bound: $\tau(p \rightarrow \pi^+ \nu) > 1.6 \times 10^{31} \text{ yr}$ [PDG](#)

- LHC signature: same sign tops

$\sigma(pp \rightarrow tN \rightarrow ttbb)$				
m_N	200 GeV	500 GeV	800 GeV	1 TeV
$\sqrt{s} = 13 \text{ TeV}$	0.34 fb	0.16 fb	$8 \times 10^{-2} \text{ fb}$	$5 \times 10^{-2} \text{ fb}$

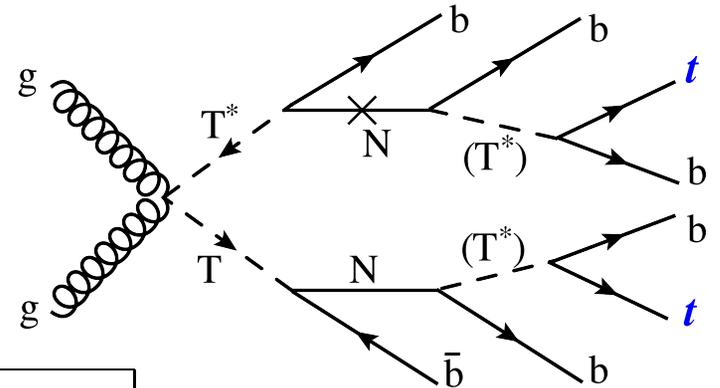
$$\Lambda/\sqrt{\lambda_a} = 1.5 \text{ TeV}$$



- Simple UV completion: color triplet scalar T

$$\mathcal{L}_{\text{UV}} = f_a \bar{N}_a^c P_R b T + f' \bar{t}^c P_R b T^* + M_T^2 |T|^2$$

- $\lambda_a/\Lambda^2 \equiv f_a f' / M_T^2$
- Possibility: $T \in \mathbf{5}'_H$ of SU(5)



$\sigma(pp \rightarrow TT^*)$				
m_T	1.5 TeV	2 TeV	5 TeV	10 TeV
$\sqrt{s} = 13 \text{ TeV}$	0.16 fb	0.01 fb	—	—
$\sqrt{s} = 100 \text{ TeV}$	384 fb	92 fb	0.54 fb	$4 \times 10^{-3} \text{ fb}$

Concluding Remarks

- Dark state X (DM or Majorana N) may have B -violating interactions
- Nucleons decay into sub-GeV X could probe the dark sector
 - Kinematically distinct from ν final state
 - Independent of local abundance
 - Does not require any appreciable X coupling to SM
- For weak scale Majorana N , direct baryogenesis: *dexiogenesis*
 - Mediated by $(Ntbb)_R/\Lambda^2$ with $\Lambda \gtrsim 1$ TeV
 - Distinct LHC signature: same sign tops
 - Concomitant nucleon decay signals, potentially accessible to experiments
 - Simple UV completion: a color triplet scalar (possible GUT connection)