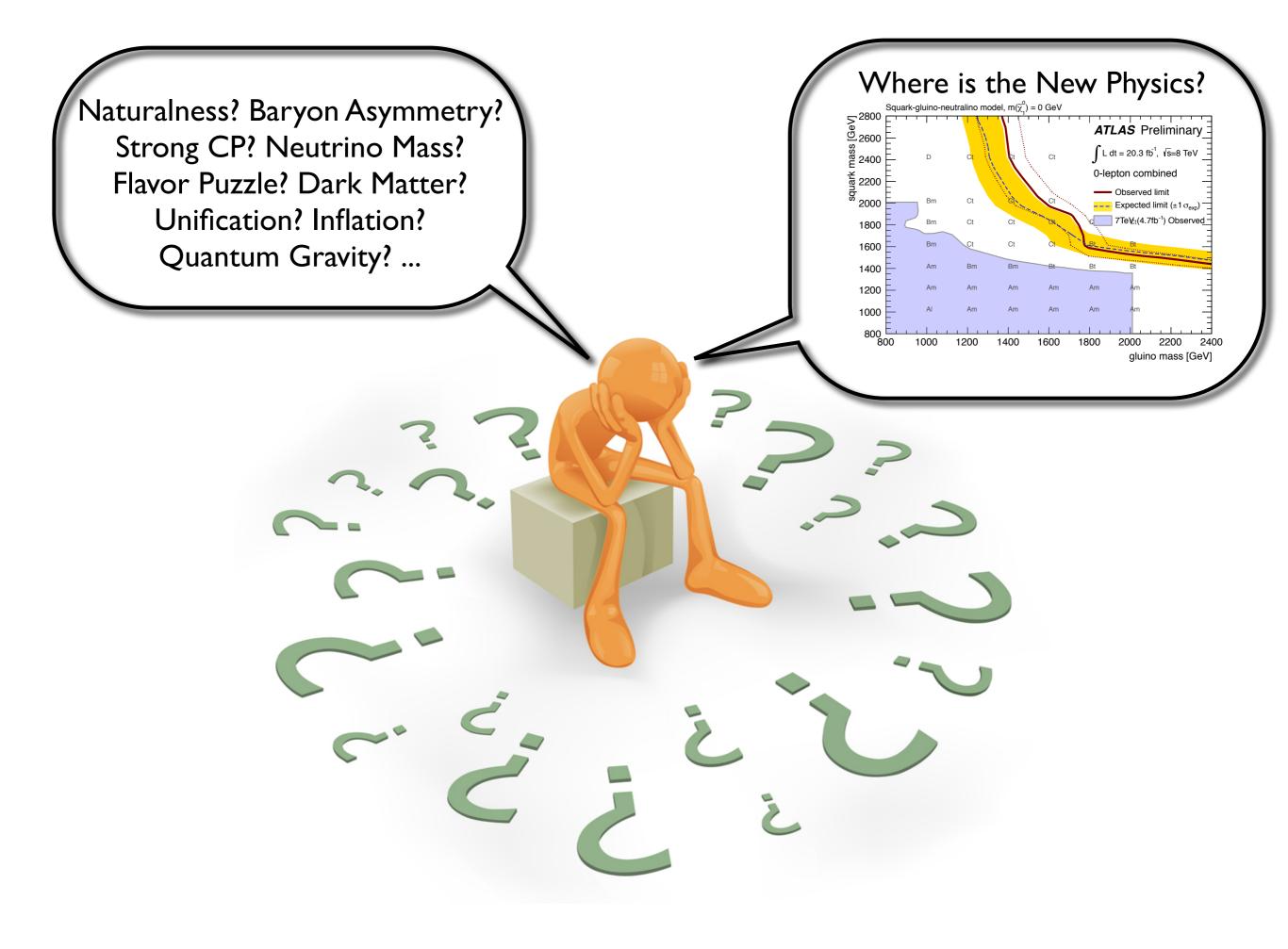
EDMs theoretical interest

Brian Batell CERN



FCC Working Group - Experiments with the CERN injectors February 13, 2015



Why Electric Dipole Moments?

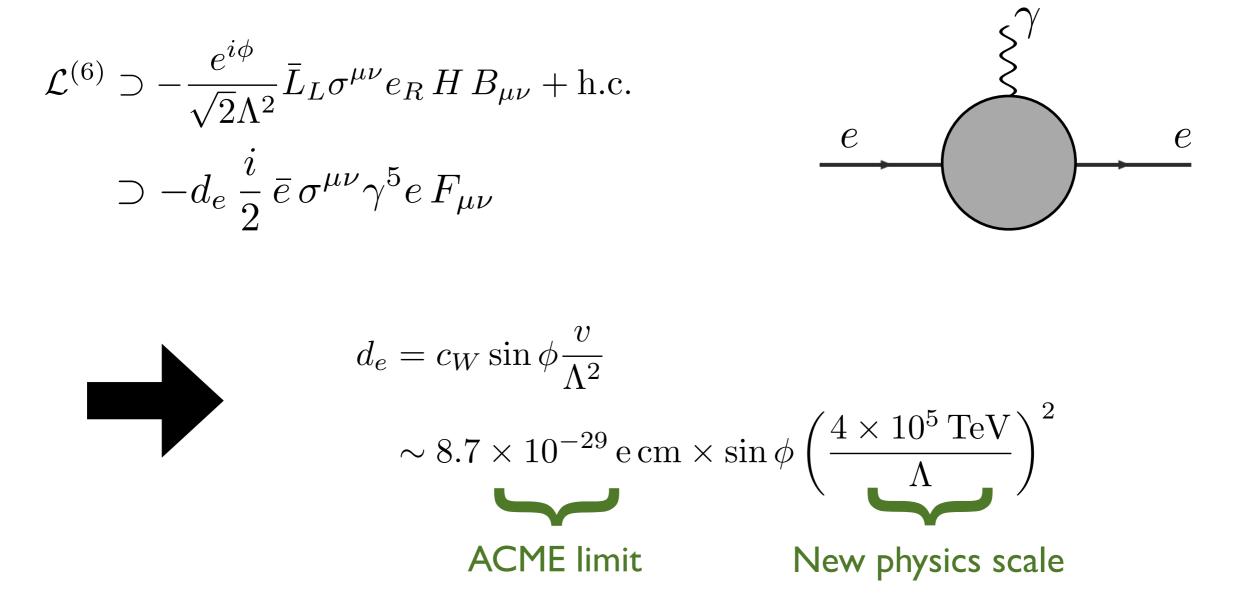
EDMs from CKM:

- CKM contribution to quark, electron EDMs contribution is tiny
 - Requires exchange of three generations of quarks to "see" CP violating phase ~ several loops, Jarlskog Invariant is small ~ 10⁻⁵
- In atoms, molecules, long distance contributions dominate, but still several orders of magnitude below experimental sensitivity

EDMs provide "background free" probe of New Physics

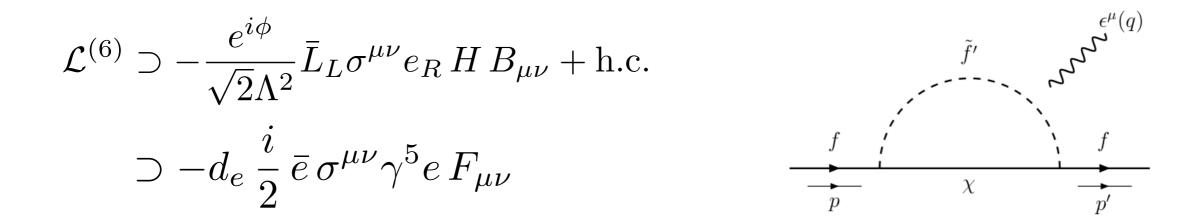
Why Electric Dipole Moments?

EDMs from New Physics:



EDMs indirectly probe scales much higher than direct searches at the LHC!

Why Electric Dipole Moments?

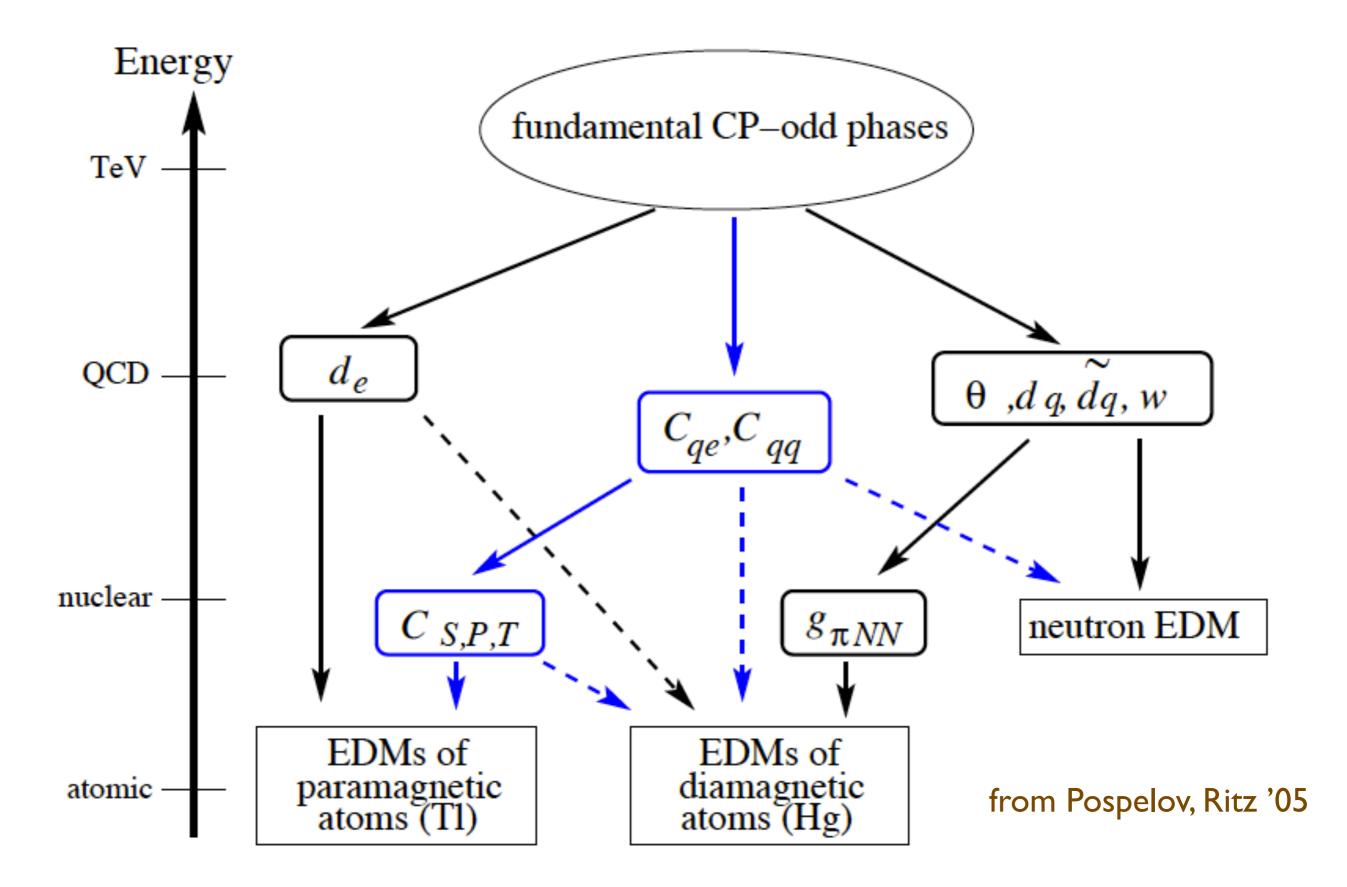


Perturbative, flavor-safe, I-loop new physics contribution:

$$\frac{1}{\Lambda^2} \sim y_e g' \frac{g^2}{16\pi^2} \frac{1}{M^2}$$

$$d_e \sim e \sin \phi \frac{g^2}{16\pi^2} \frac{m_e}{M^2}$$

$$\sim 8.7 \times 10^{-29} \text{ e cm} \times \sin \phi \left(\frac{20 \text{ TeV}}{M}\right)^2$$
ACME limit New physics scale



Current EDM limits and future experimental benchmarks

	System	Present 90% C.L.	Sensitivity goal ^b	Group
Paramagnetic $d_{\mathrm{para}}(d_e, C_{eeqq})$		Limit (e fm) ^a		
	Cs Tl	1.2×10^{-10} 9.5×10^{-12}		[169] [170]
	YbF ^d ThO ^d	10.5×10^{-15}	$10^{-15} \rightarrow 10^{-17}$	[152]
Nucleons $d_{\mathrm{n,p}}(ar{ heta}, d_q, \widetilde{d}_q, w, C_{qqqq})$	n n n n n n	2.7 × 10 ⁻¹³	$(1-3) \times 10^{-14}$ 4×10^{-15} 5×10^{-14} 5×10^{-15} 2×10^{-15} $10^{-14} - 10^{-15}$	[171] CryoEDM nEDM/SNS nEDM/PSI n2EDM/PSI nedm/FRM-II Munich TRIUMF
Diamagnetic $d_{ ext{dia}}(ar{ heta}, d_q, \widetilde{d}_q, w, C_{qqqq})$	p ¹⁹⁹ Hg ²²⁵ Ra ^{221/223} Rn ^{221/223} Rn ¹²⁹ Xe	$2.6 imes 10^{-16}$ $5.5 imes 10^{-14}$	$\begin{array}{c} 10^{-16} \\ (2.6-5)\times10^{-17} \\ (10-100)\times10^{-15} \\ 1.3\times10^{-14} \\ 2\times10^{-15} \end{array}$	srEDM [172] Argonne TRIUMF FRIB [173]

Engel et al. 2013

Complementarity: different systems probe different combinations of underlying CPV parameters

Strong CP problem

CPV QCD theta term:

$$\mathcal{L} \supset \frac{g_3^2}{32\pi^2} \,\overline{\theta} \, G^A_{\mu\nu} \widetilde{G}^{\mu\nu A} \qquad \overline{\theta} = \theta + \arg \det(Y_u Y_d)$$

Neutron EDM:

$$d_n \sim \frac{em_q^*}{m_N^2} \sim 10^{-17} e\,\mathrm{cm} \times \overline{\theta}$$

Experimental Bound:

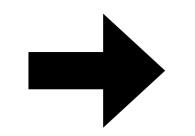
$$d_n^{\rm exp} \lesssim 10^{-26} \, e \, {\rm cm}$$

 $\overline{\theta} \lesssim 10^{-9}$

Strong CP Problem: Why so small?

Baryon Asymmetry of the Universe

- We observe matter, but not antimatter! $n_B \gg n_{\overline{B}}$
- Initial conditions? CMB anisotropies support inflation, early universe dominated by vacuum - matter produced via reheating



Need for baryogenesis!

Sakharov conditions:

- **1**. Baryon number violation
- **2.** C and CP violation
- **3.** Departure from thermal equilibrium

Baryon Asymmetry of the Universe

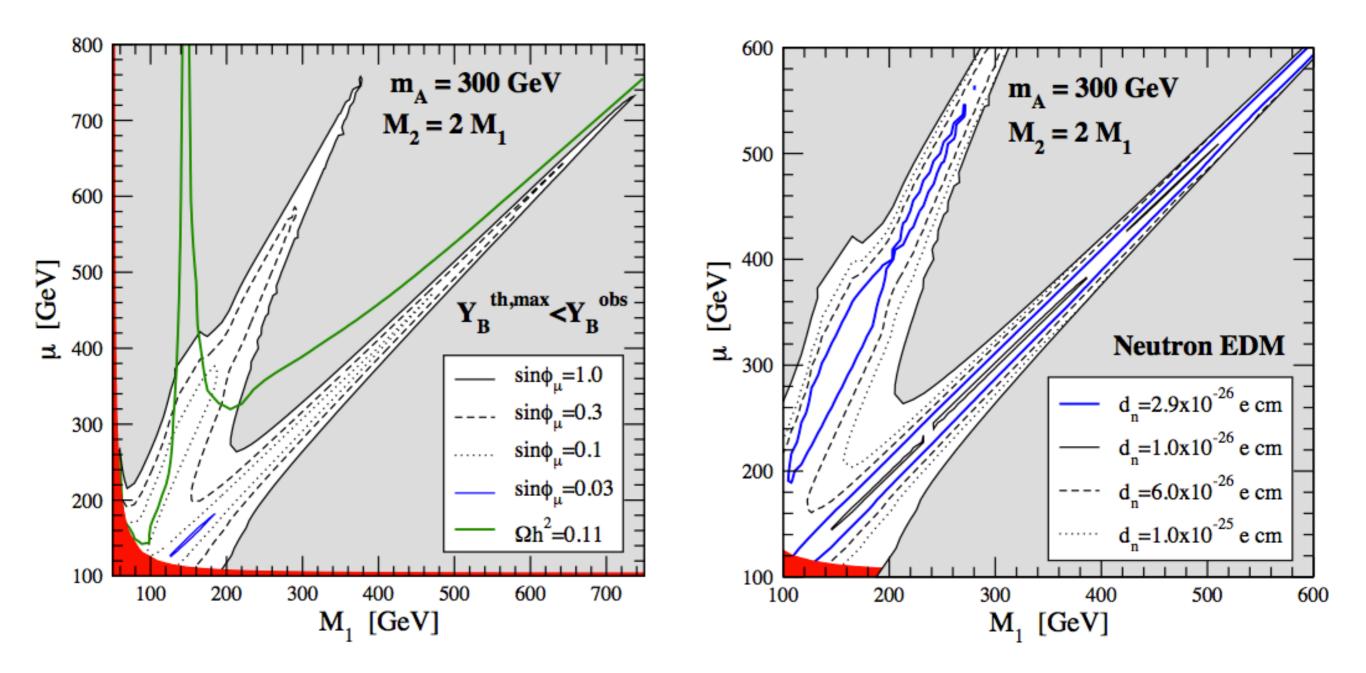
- In principle, Sakharov conditions satisfied in the Standard Model
 - Sphalerons, CKM phase, EW phase transition
- However, in practice EW baryogenesis doesn't work in the SM:
 - CP violation too small
 - 125 GeV Higgs absence of 1st order EW phase transition

New dynamics and sources of CP violation needed to account for the baryon asymmetry!

- Unfortunately we don't know the scale associated with this dynamics
 - GUT baryogenesis, Leptogenesis, Electroweak baryogenesis, ...

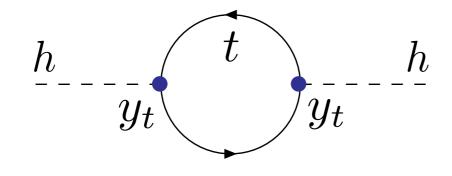
can lead to observable EDMs

EDMs probe baryogenesis in the MSSM



Cirigliano et al '09

Naturalness and the Weak Scale



$$\delta m_h^2 = -\frac{3y_t^2}{8\pi^2}\Lambda^2$$

 Λ ~ scale of new physics

Very naively we expect

$$|\delta m_h^2| \lesssim m_h^2 = (125 \,\mathrm{GeV})^2$$

$$\Lambda \lesssim 650 \, {\rm GeV}$$

or else we start to tune...

Suggests new dynamics at 100 GeV - 1 TeV!

Naturalness and the New Physics CP problem

- Theories addressing hierarchy problem scale typically introduces many new CP violating parameters
- Strong constraints from EDMS for ~ TeV scale SUSY

 $\phi \lesssim 0.001 - 0.1$

• SUSY baryogenesis prefers order one phases.

 $m_1 \tilde{B}\tilde{B} + m_2 \tilde{W}\tilde{W} + m_3 \tilde{g}\tilde{g}$ $+\mu \tilde{H}_u \tilde{H}_d + B\mu H_u H_d$ $\rightarrow 5 \text{ phases}$

Example: MSSM

 $m_Q^2 \tilde{Q}_L^{\dagger} \tilde{Q}_L + m_U^2 \tilde{U}_R^{\dagger} \tilde{U}_R + m_D^2 \tilde{D}_R^{\dagger} \tilde{D}_R$ $+ m_L^2 \tilde{L}_L^{\dagger} \tilde{L}_L + m_E^2 \tilde{E}_R^{\dagger} \tilde{E}_R$ $\rightarrow 15 \text{ flavor mixing angles } + 15 \text{ phases}$

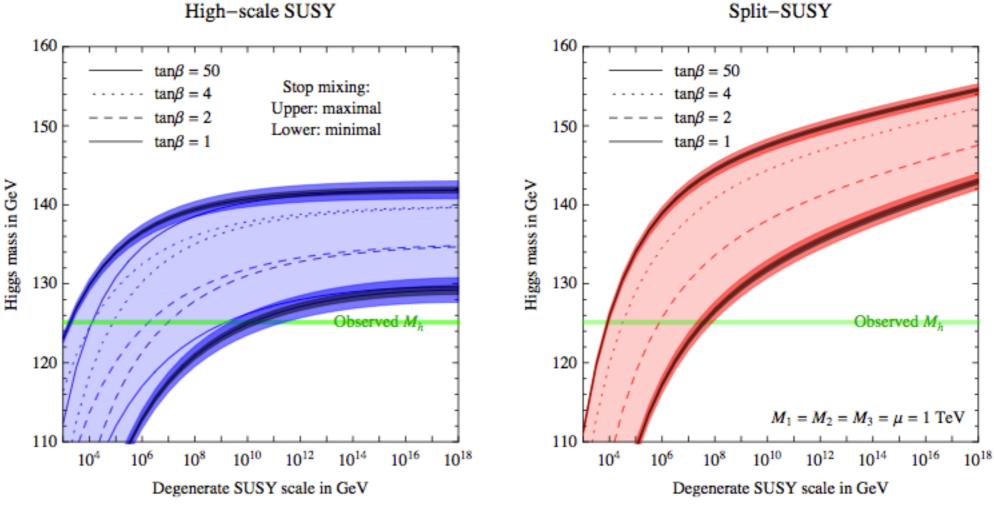
 $A_u H_u \tilde{Q}_L^{\dagger} \tilde{U}_R + A_d H_d \tilde{Q}_L^{\dagger} \tilde{D}_R + A_\ell H_d \tilde{L}_L^{\dagger} \tilde{E}_R$

→ 18 flavor mixing angles + 27 phases

2 phases can be rotated away... from talk by W. Altmanshoffer

Maybe SUSY is a bit heavier than expected?

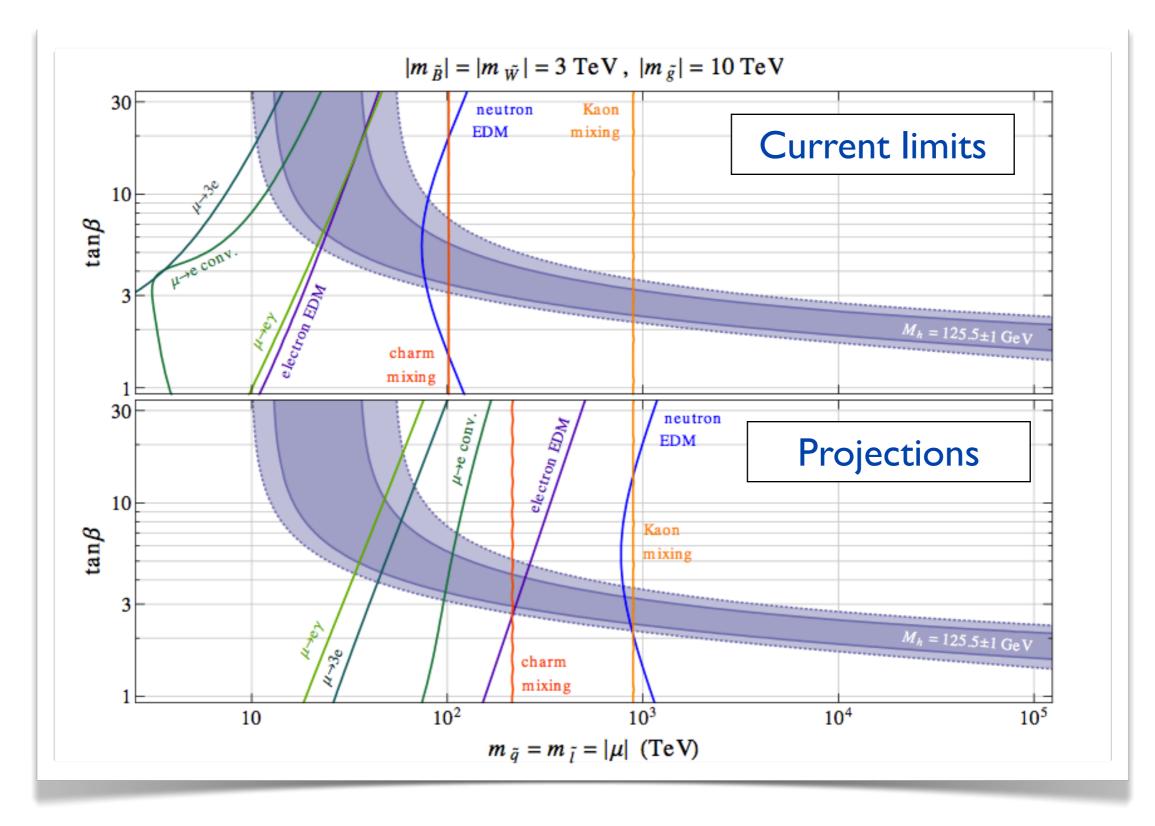
- No signs of superpartners in Run I
- SUSY flavor, CP problems ameliorated
- 125 GeV Higgs mass points towards heavy scalars



Split-SUSY

Bagnaschi et al. 2014

EDMs and other indirect tests probe heavy SUSY



Altmanshoffer, Harnik, Zupan

Summary

- We know there is new physics, but so far we haven't found it!
- With no direct evidence for new physics from the LHC, we need to pursue a broad program to test the SM and go beyond
- EDMs are a cruical component of this program!
 - Strong CP problem, Baryogenesis, Hierarchy problem are all good reasons to probe for new sources of CP violation
 - Allow indirect access to very high scales of order 1-1000 TeV!
- It is worthwhile to consider how CERN can contribute to this program!