



lisa pathfinder

# LISA Pathfinder

## a status report

M Hewitson, AEI Hannover  
for the LPF Team

eLISA Cosmology Workshop, CERN, April 14th 2015



# Observing GWs from space



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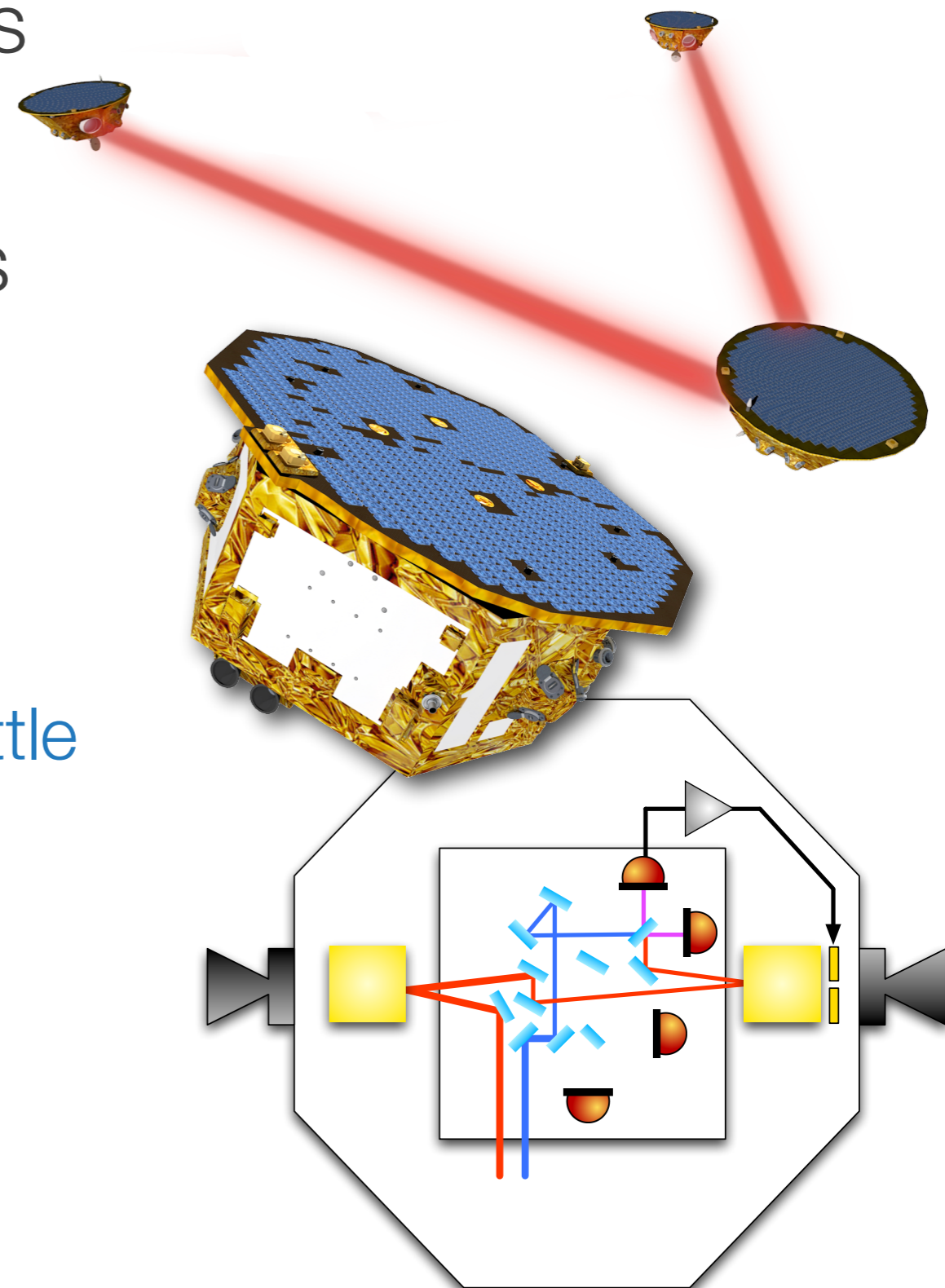
- Push down to much lower frequencies: mHz
  - expect rich science from, e.g., super-massive BH binaries out to high redshift, EMRIs
- Fractional length stability of ground-based instruments is difficult to achieve at these frequencies
  - temperature and structural fluctuations
- Increase arm-length to improve strain sensitivity
  - go to space

# LISA Pathfinder



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- Devised to test key technologies for LISA-like space-based gravitational wave observatories
- The idea was to shrink a LISA arm into a single space-craft
  - tests length stability but has very little sensitivity to strains in space-time
  - allows us to test most of the technology needed for a LISA-like mission



# What do we test?



- Technology demonstrator for a GW Observatory in space (LISA-like design):
  - micro-Newton propulsion
  - Gravitational Reference Sensor
  - Interferometric techniques
  - Drag-free control
- LPF will be placed in orbit around L1
- Requirements relaxed compared to a typical space-borne GW observatory to make testing feasible:
  - 1 order of magnitude in differential acceleration
  - 1 order of magnitude higher frequency
- Two science payloads:
  - LISA Technology Package (LTP)
  - ST7 (NASA provided payload)

# What do we test?



- Technology demonstrator for a GW Observatory in space (LISA-like design):

- micro-
- Gravitational
- Interferometry
- Drag-free

measurement of  
residual differential acceleration of  
 $30 \text{ fm s}^{-2} / \sqrt{\text{Hz}}$  at 1 mHz  
between two free-falling test-masses  
(pm accuracy position measurement)

- LPF will
- Require

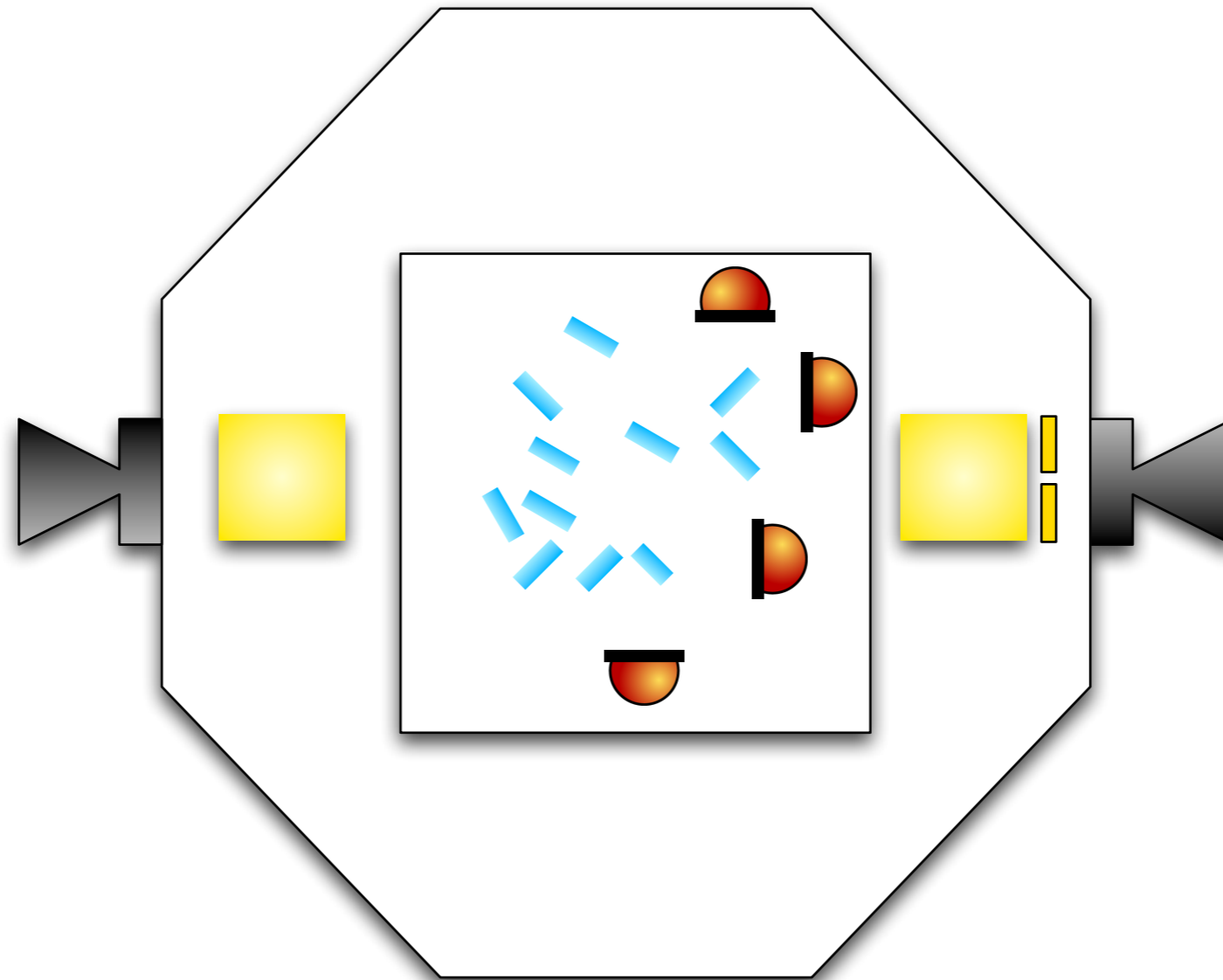
observatory to make testing feasible:

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# Measurement concept



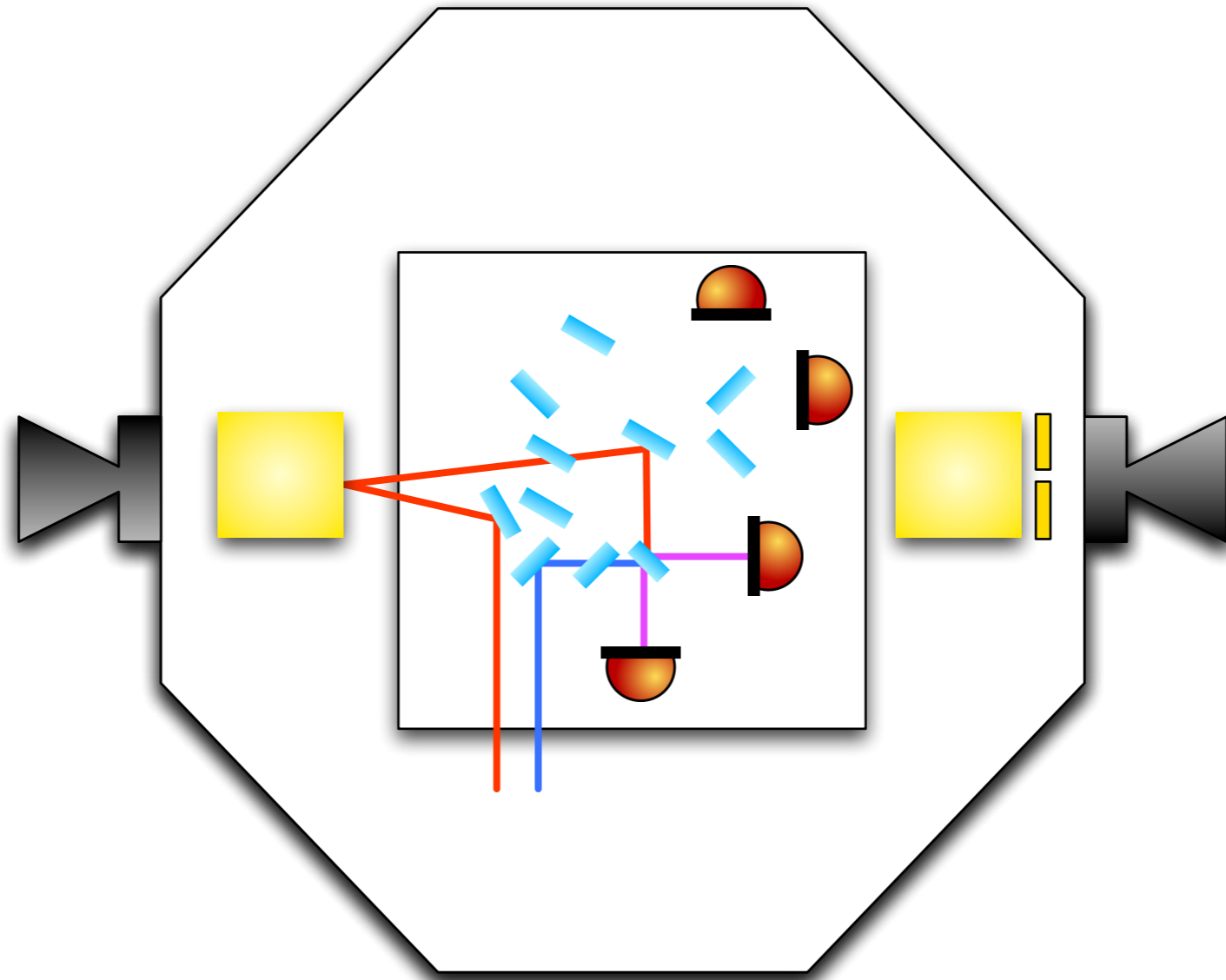
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# Measurement concept



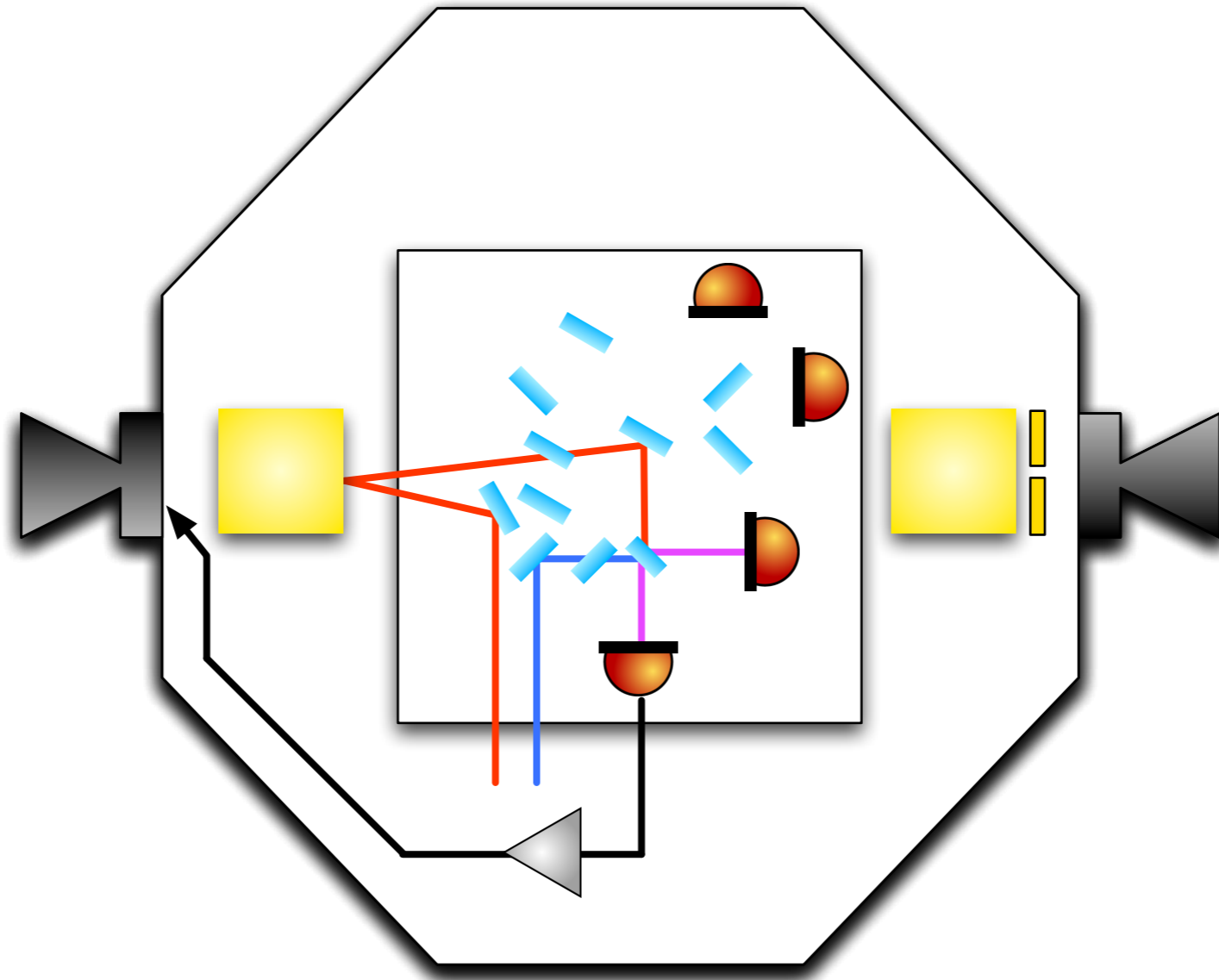
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# Measurement concept



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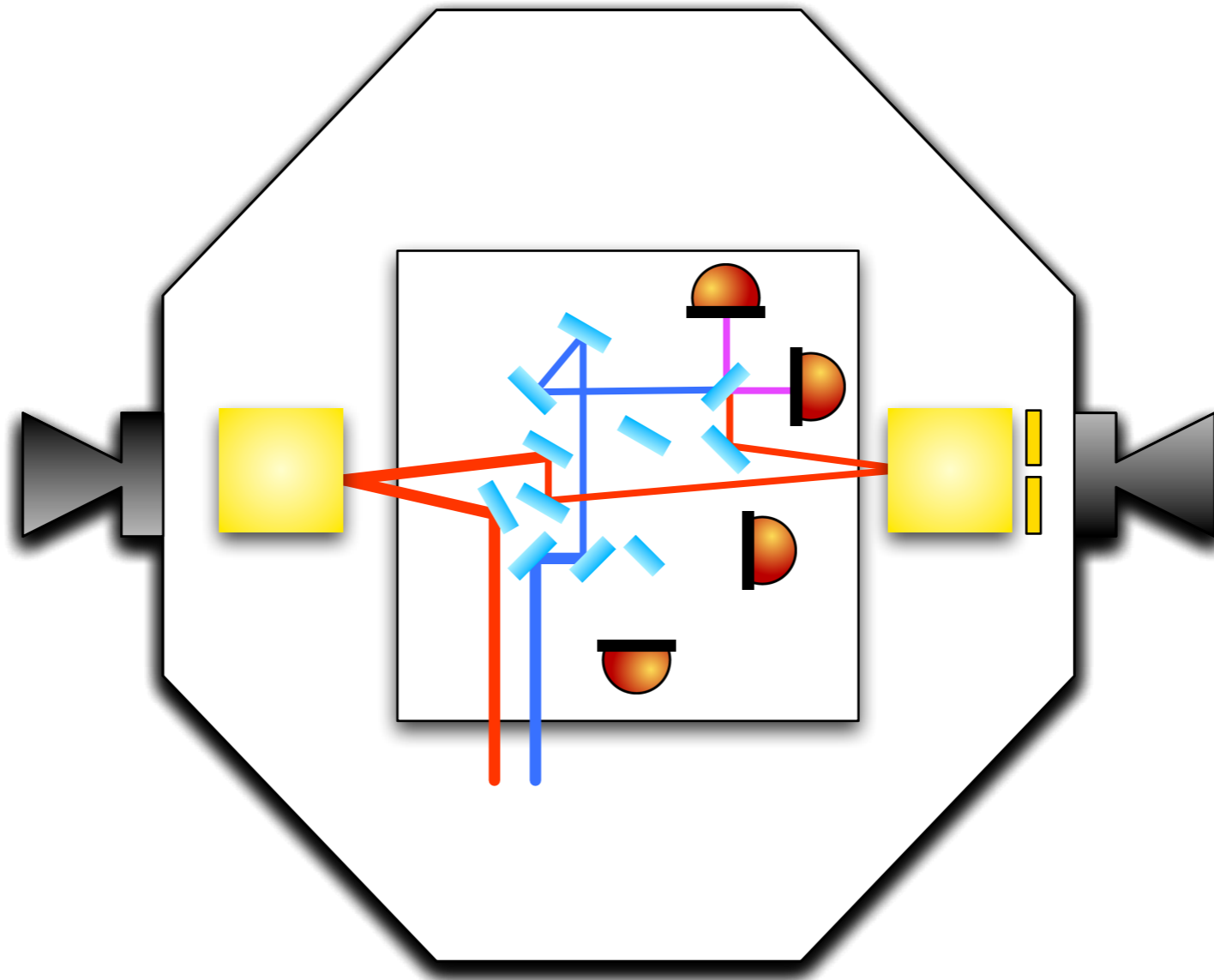




# Measurement concept



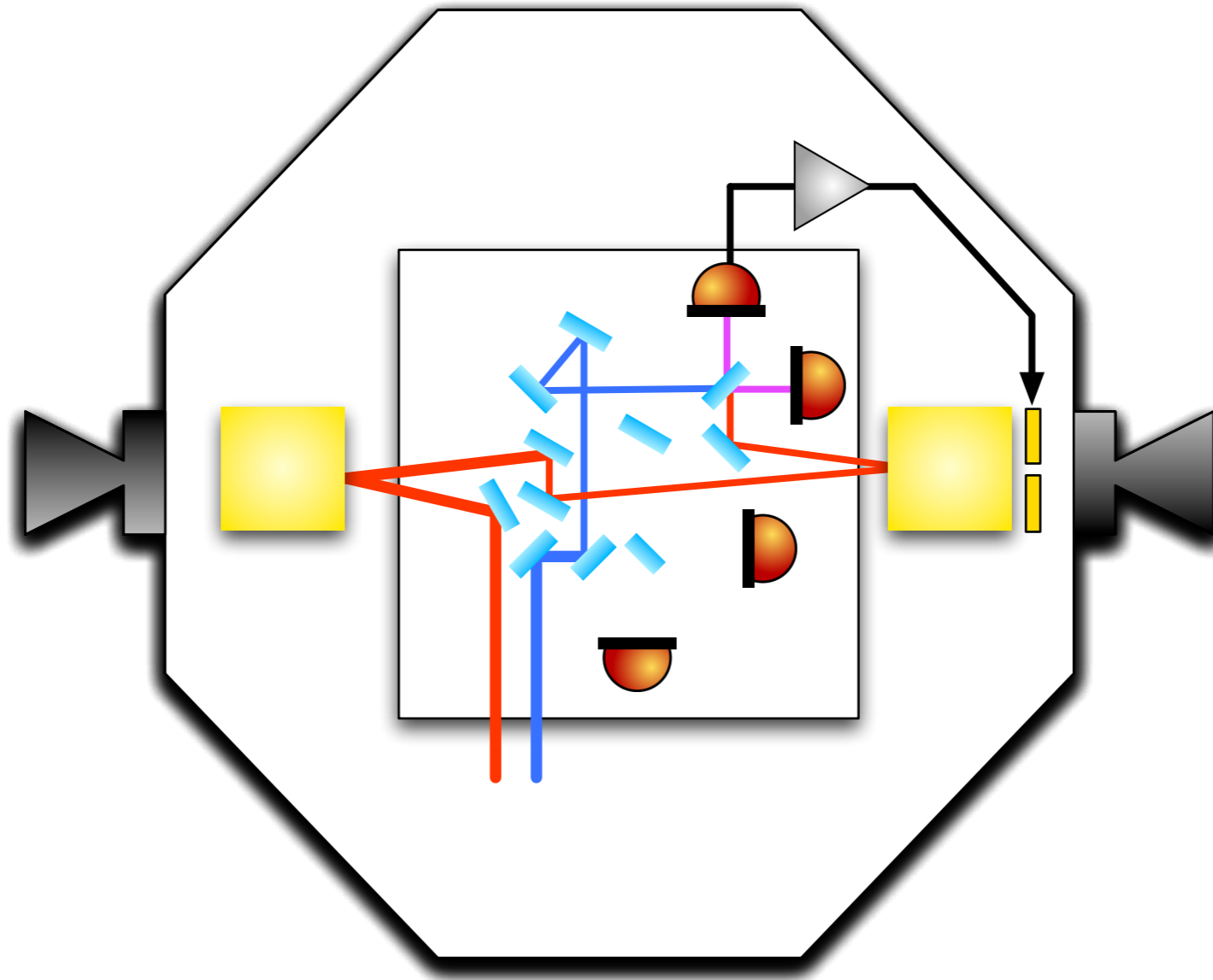
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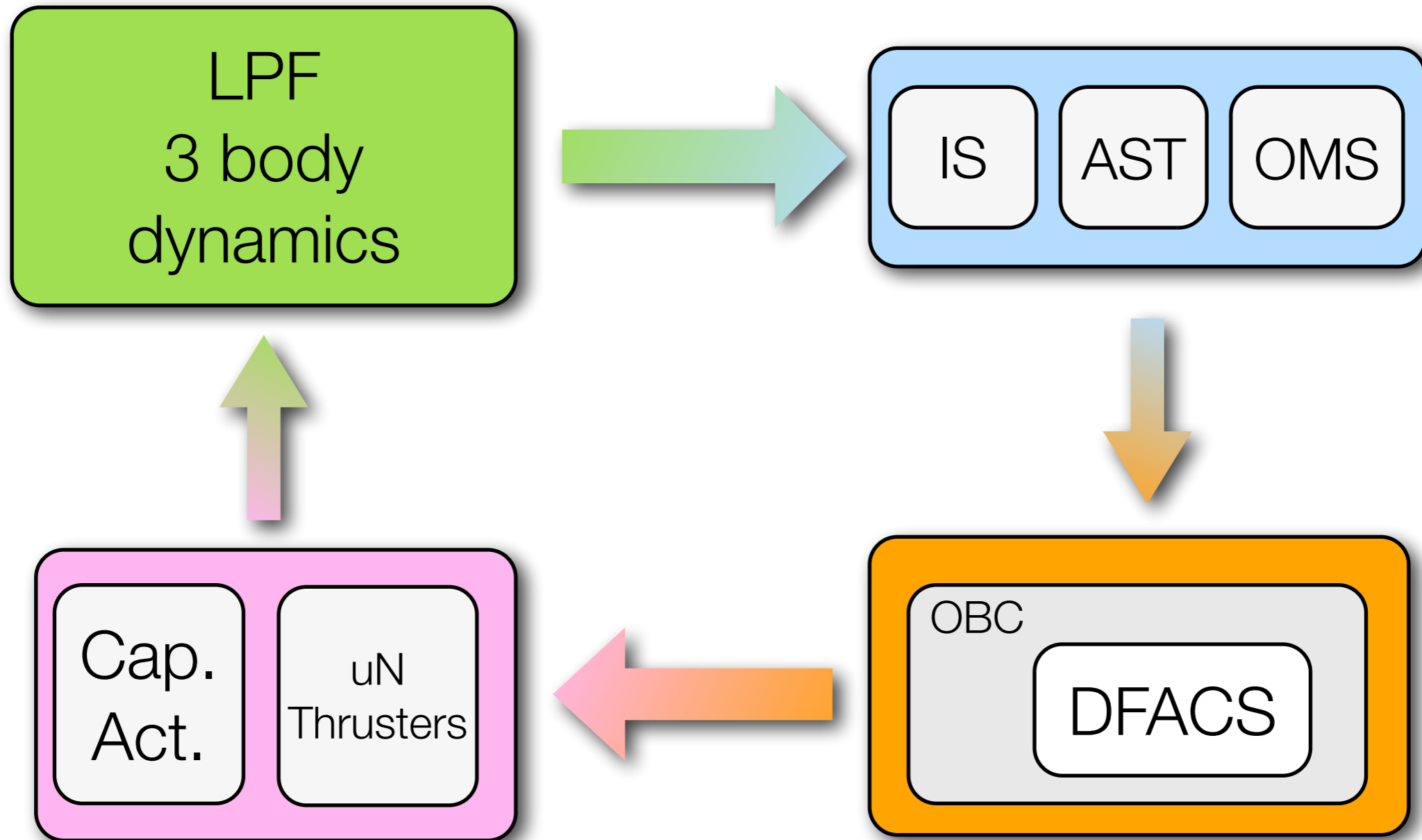
# Measurement concept



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# System level

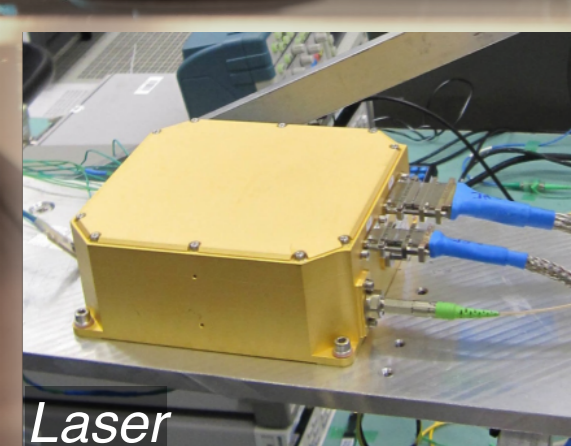


# Optical Metrology System

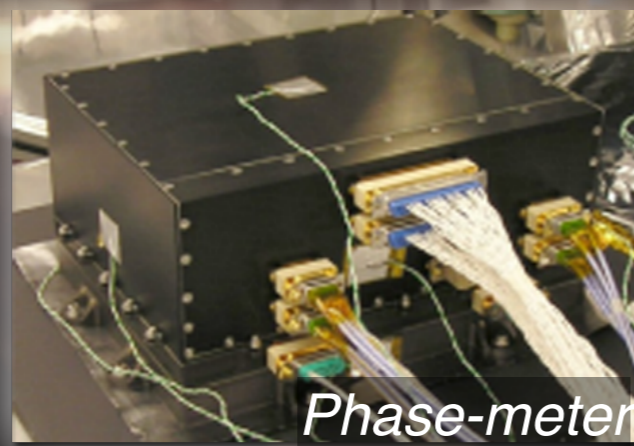


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- 2W laser
- Multi-channel Phasemeter
- 4 interferometers bonded on a Zerodur baseplate
  - 2 measure relative TM/SC positions
  - 1 measures frequency noise of laser
  - 1 measures optical path length fluctuations
- Data Management Unit



Laser



Phase-meter



Data Management Unit

X1

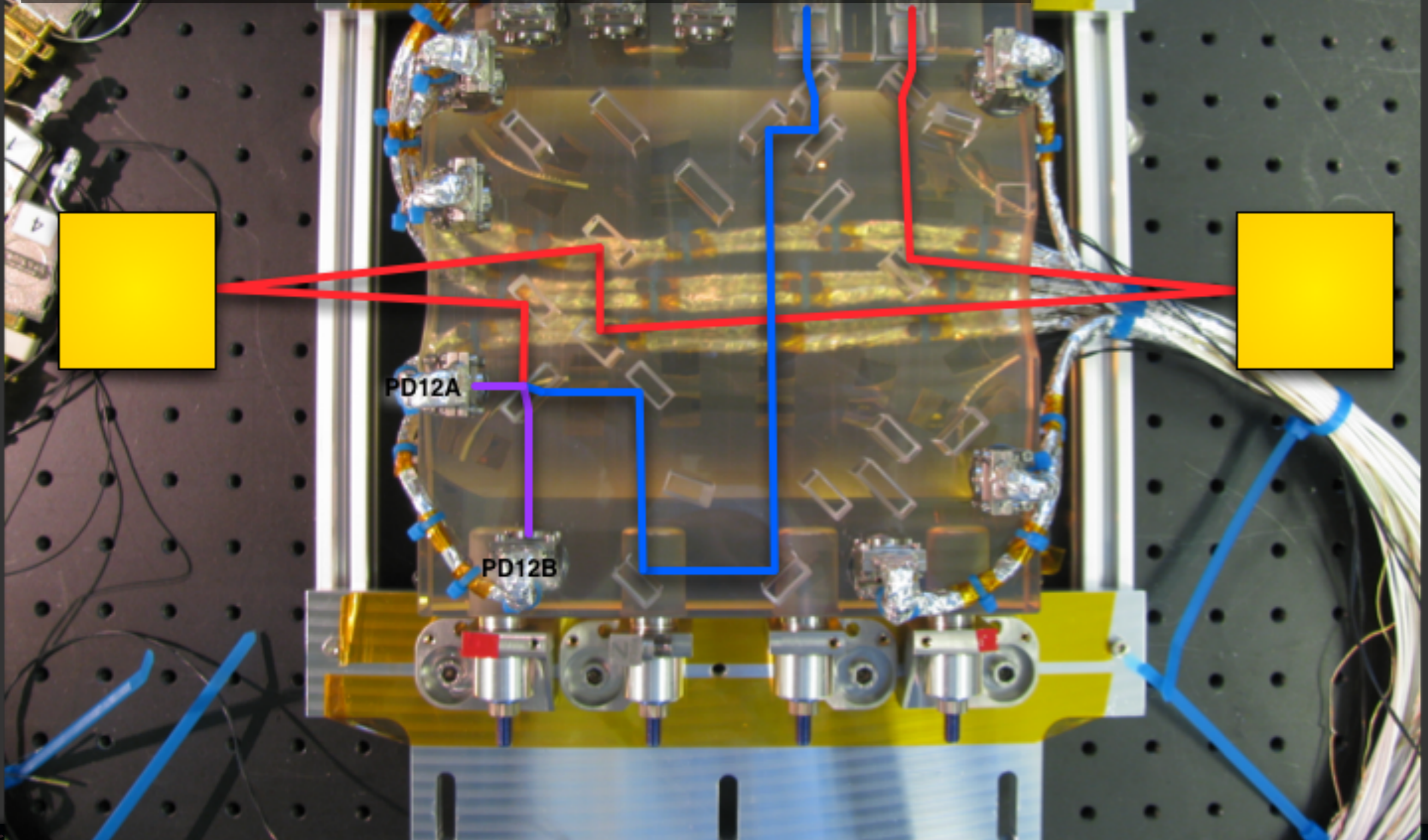
PD1A PD1B

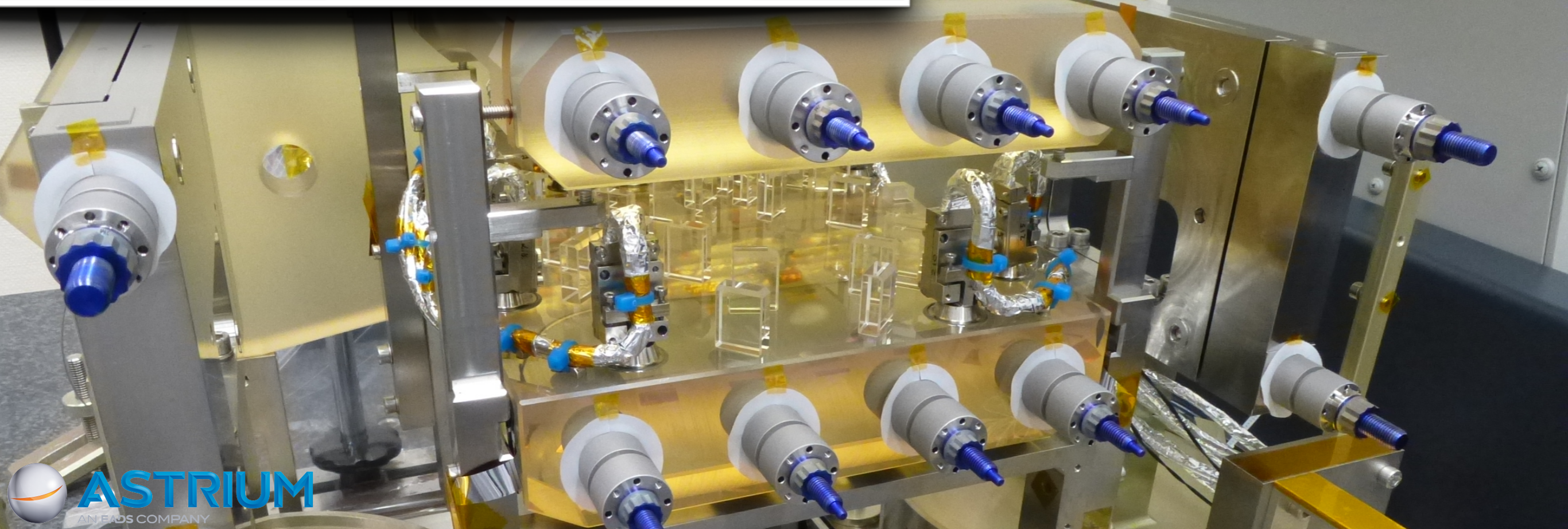
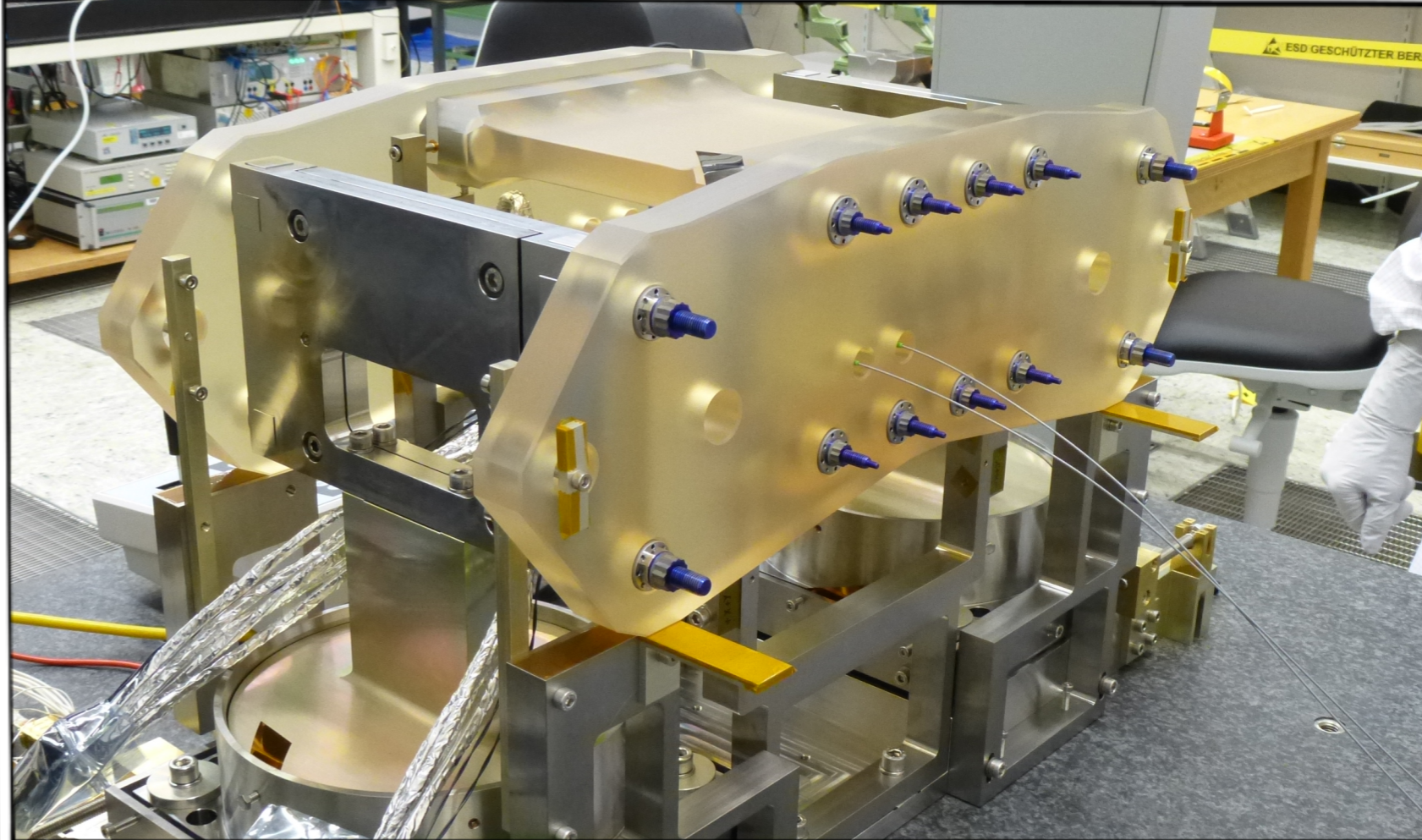
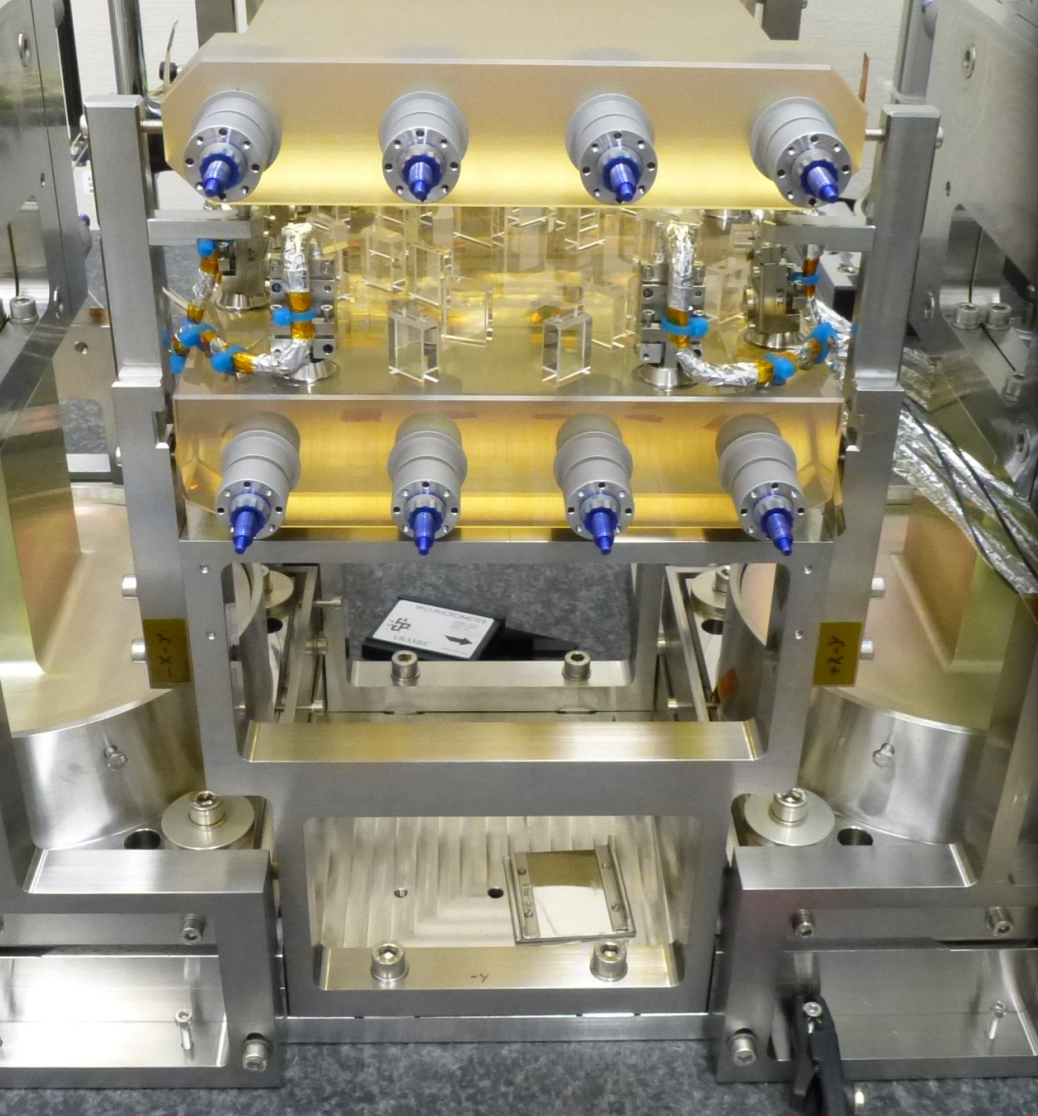


X1 measures distance between SC (bench) and TM1

X12 measures differential distance between two TMs

X12



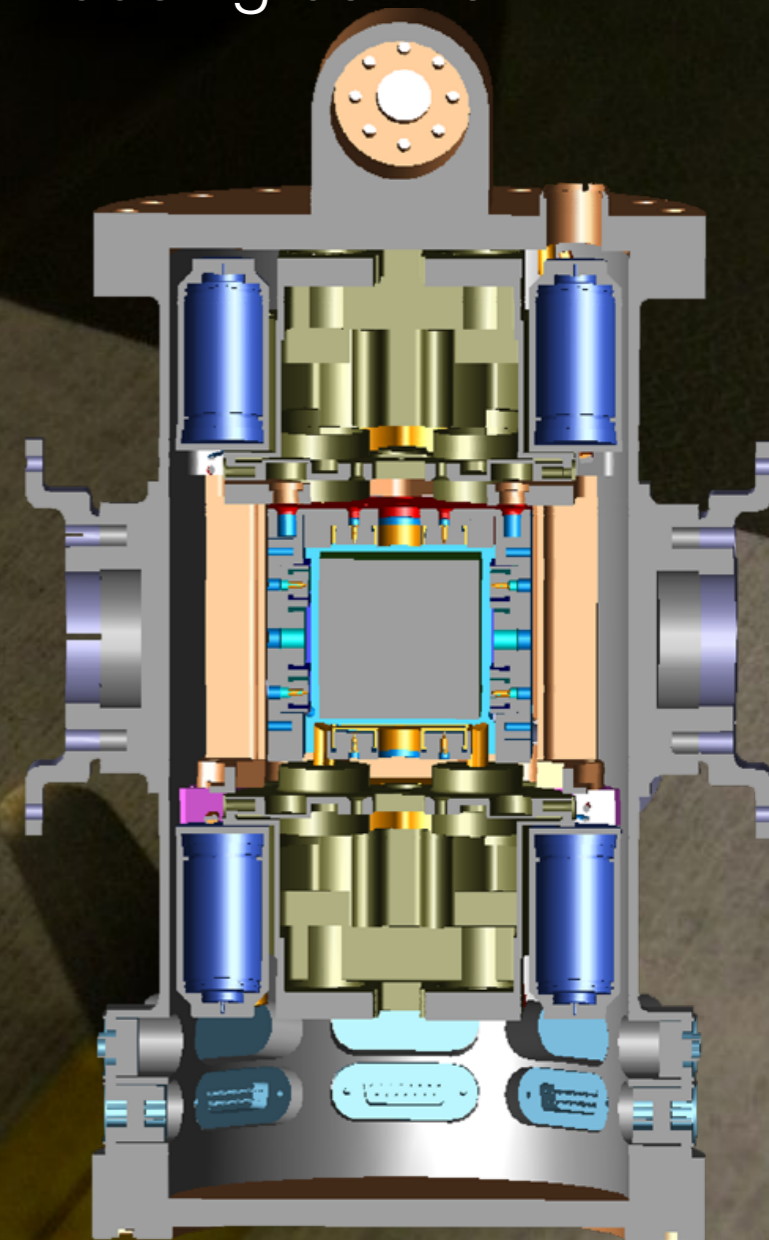


# Inertial Sensor Subsystem

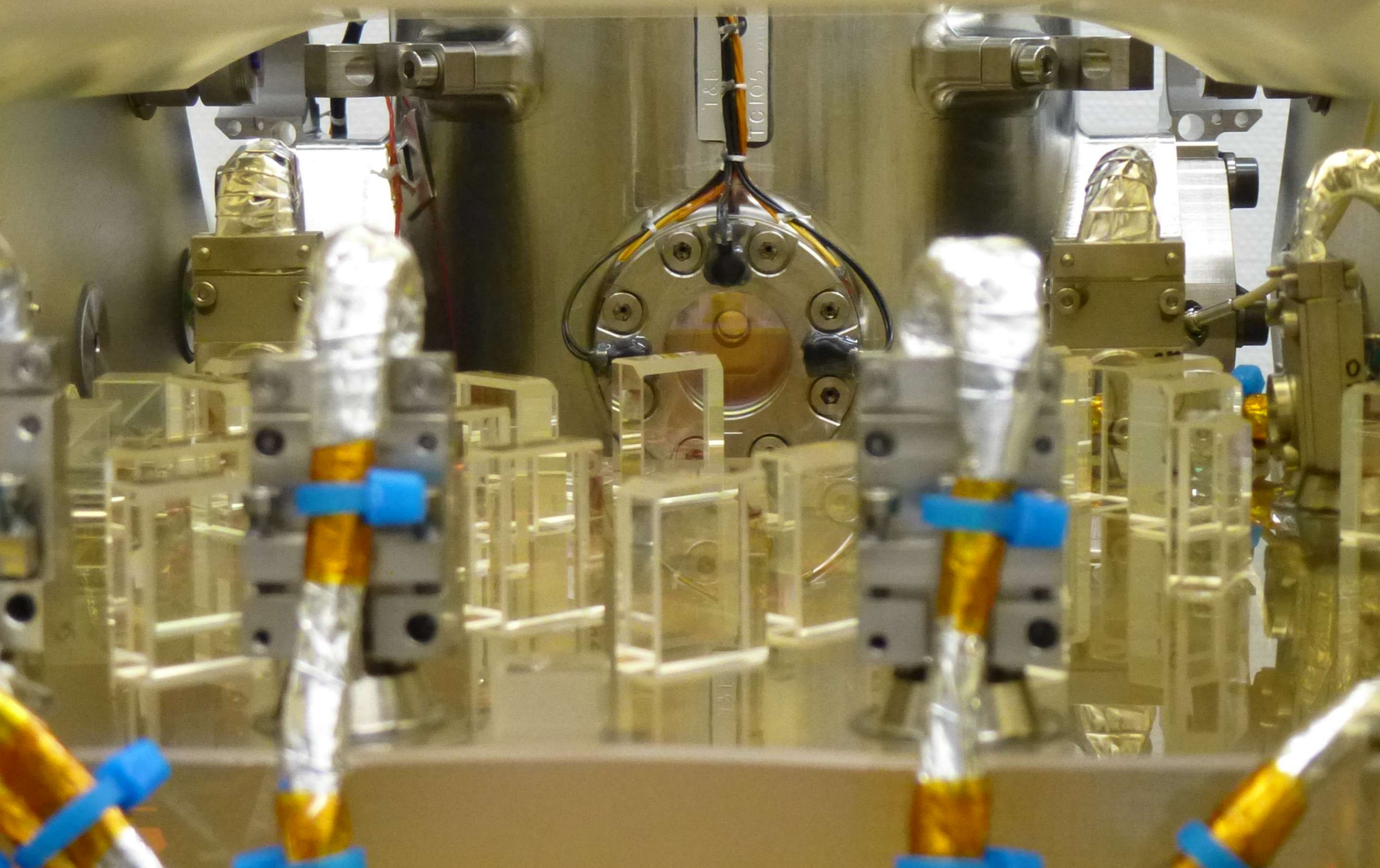


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- Test masses are surrounded by electrodes which provide 6 degree-of-freedom sensing (and also actuation)
  - forms a differential capacitance meter
- Can sense absolute position of TM within 200 $\mu$ m of the housing centre with an accuracy of nanometers
- System includes the grabbing and release mechanism
- Has two modes of operation:
  - High Resolution mode
    - low sensing noise 1.8 nm/Hz
    - Sensing range +/- 200  $\mu$ m
    - Actuation authority 5 nN
  - Wide Range mode
    - Relatively high sensing noise
    - Sensing range 4 mm (all motions)
    - Actuation authority 8  $\mu$ N



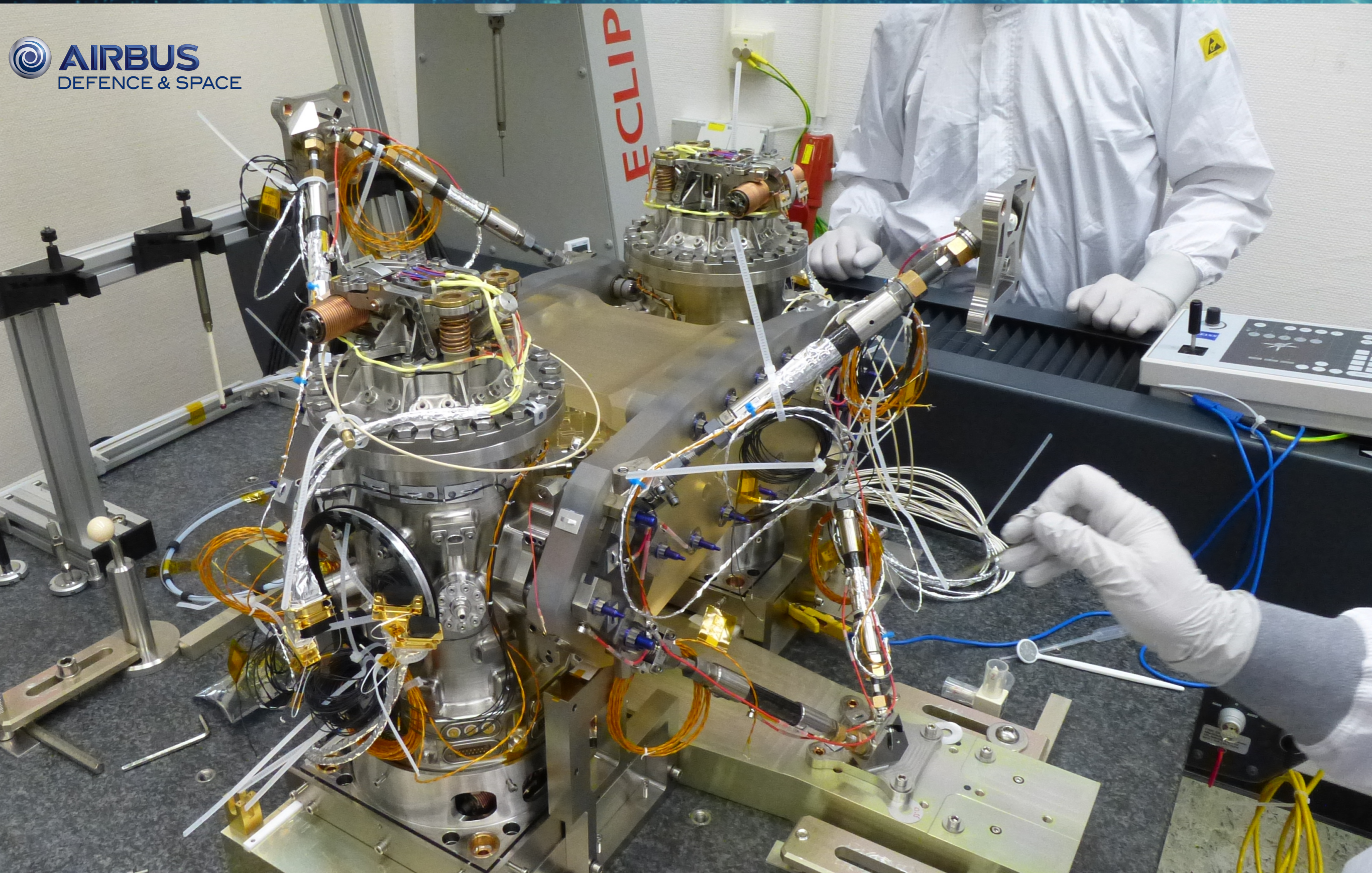




# LTP Integration



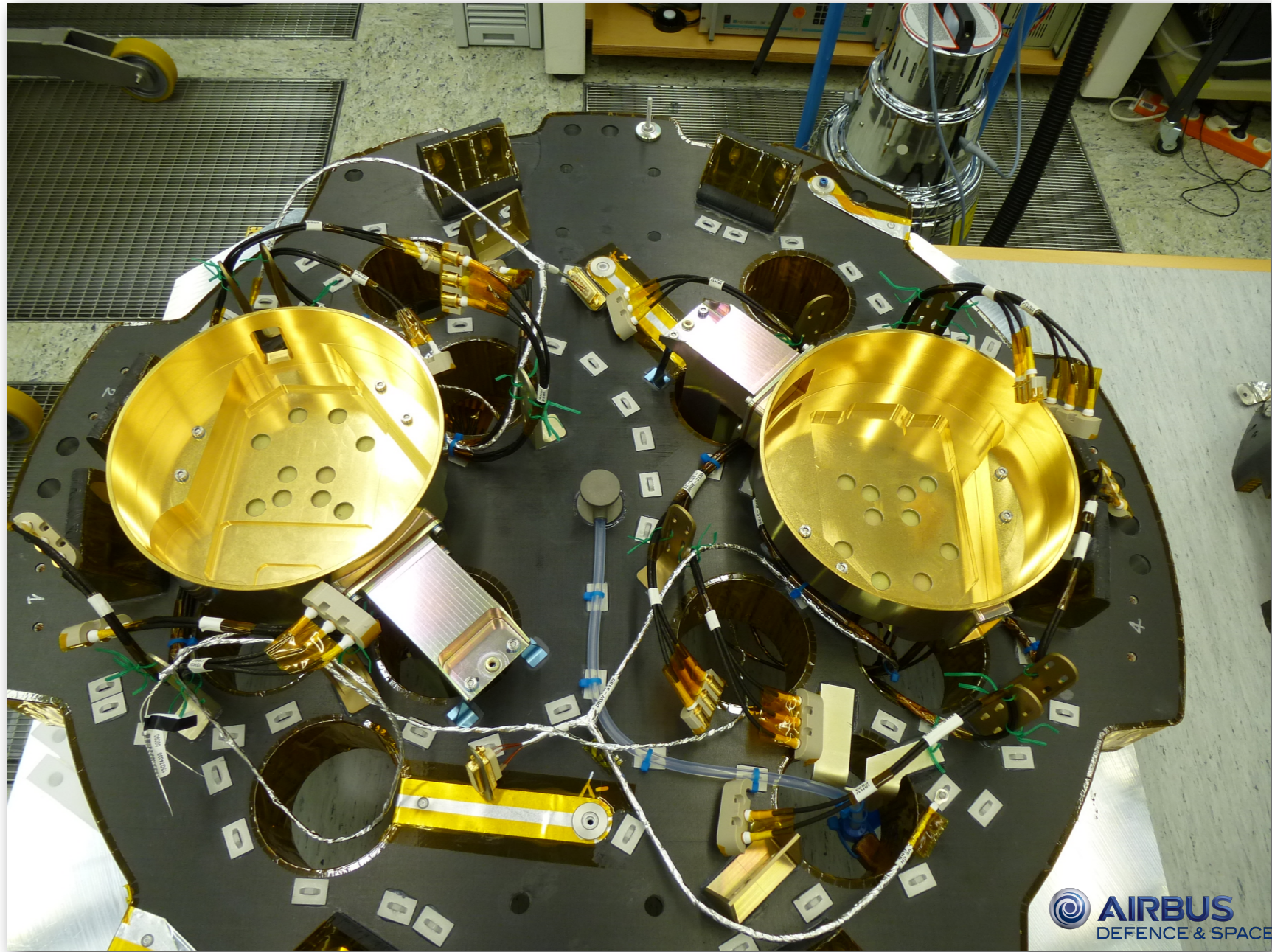
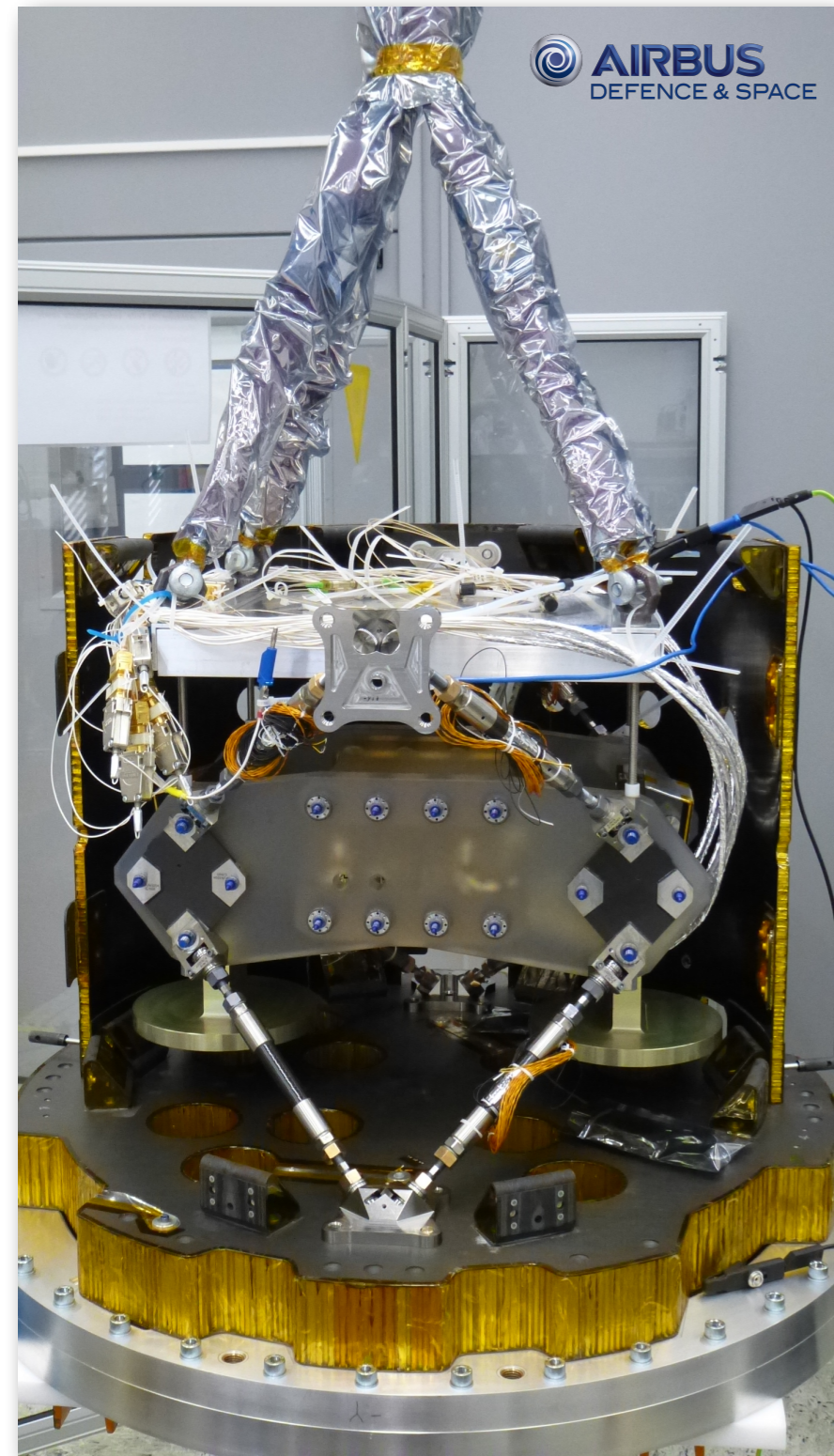
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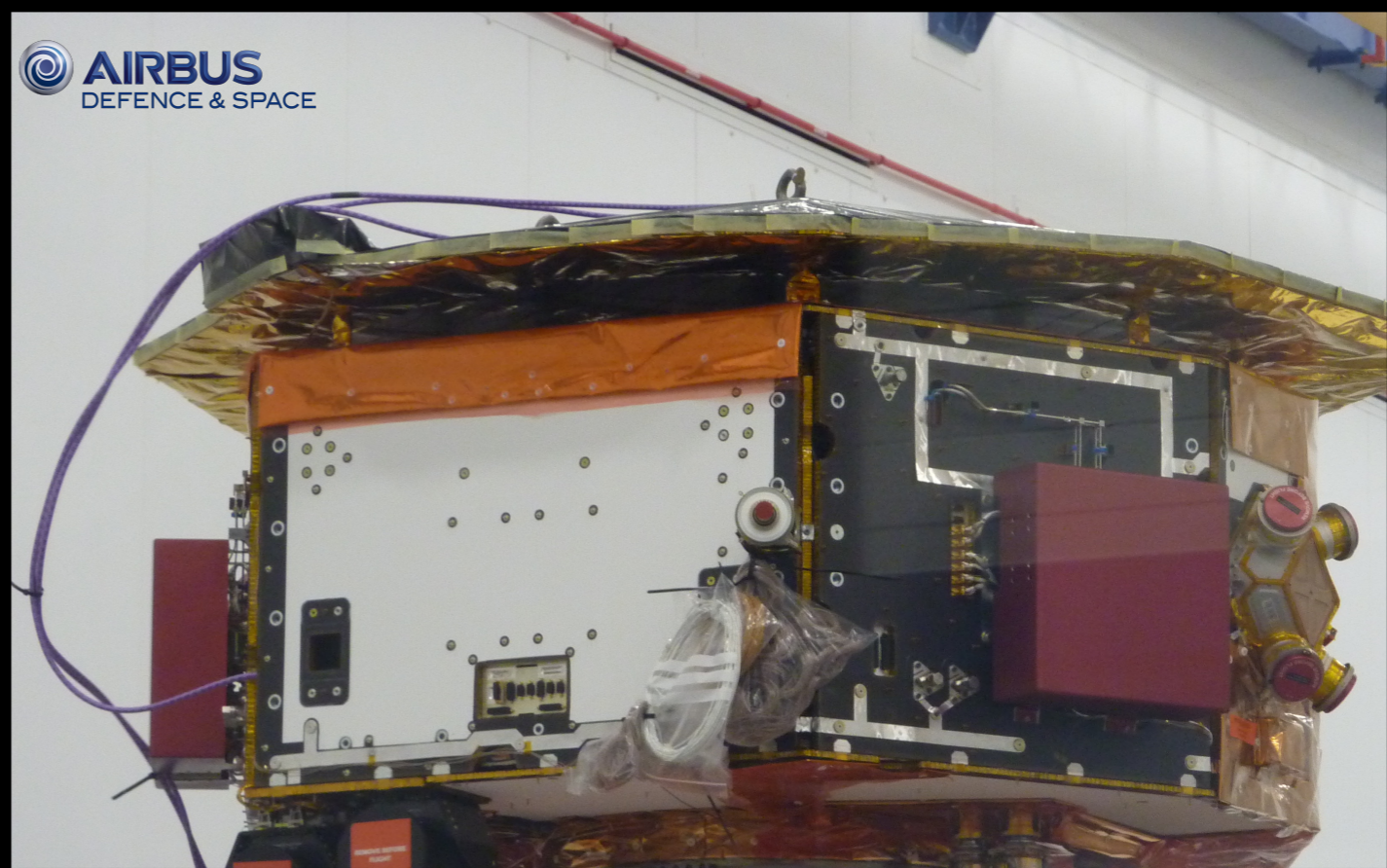
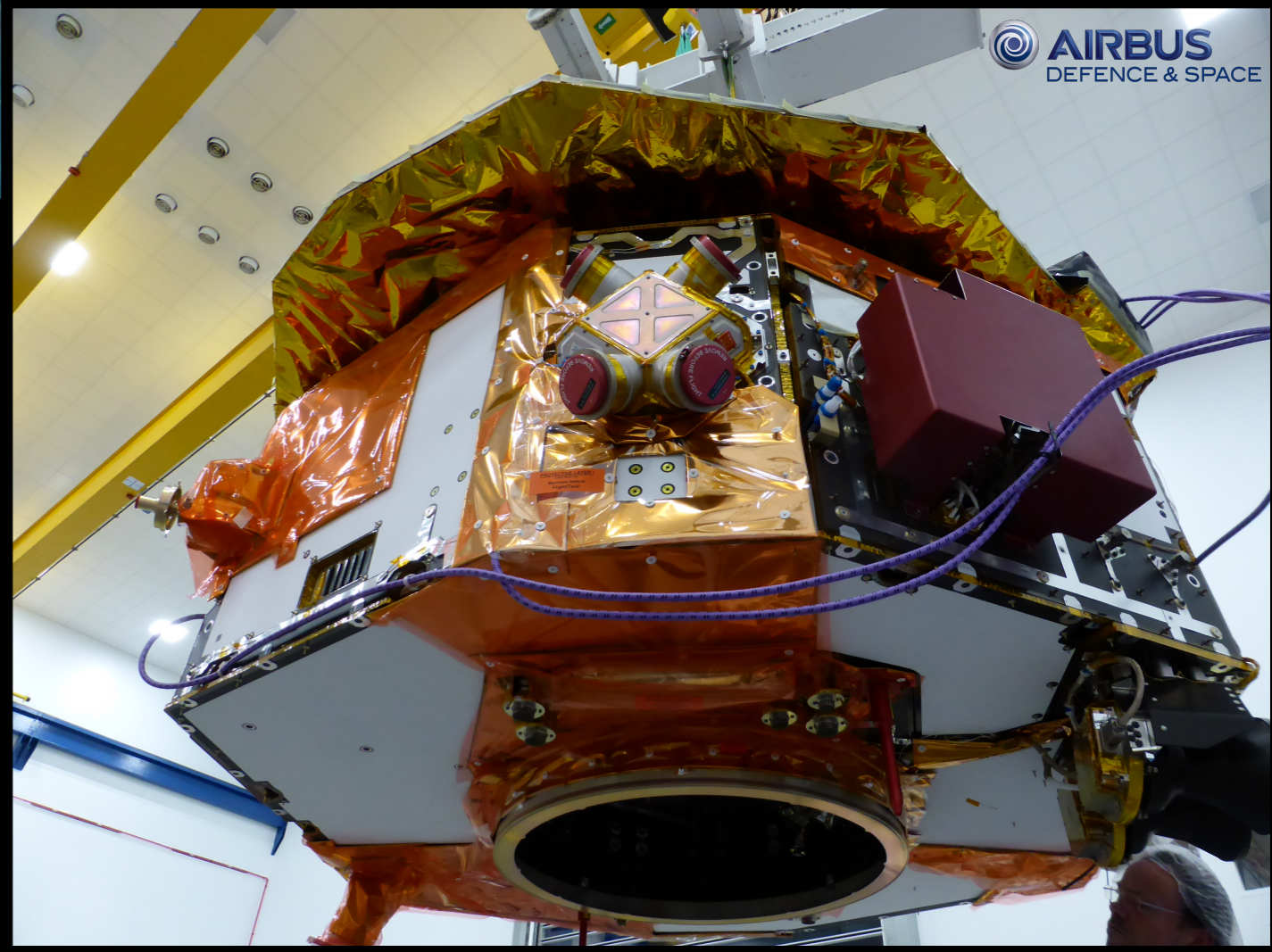
# Integration of LCA



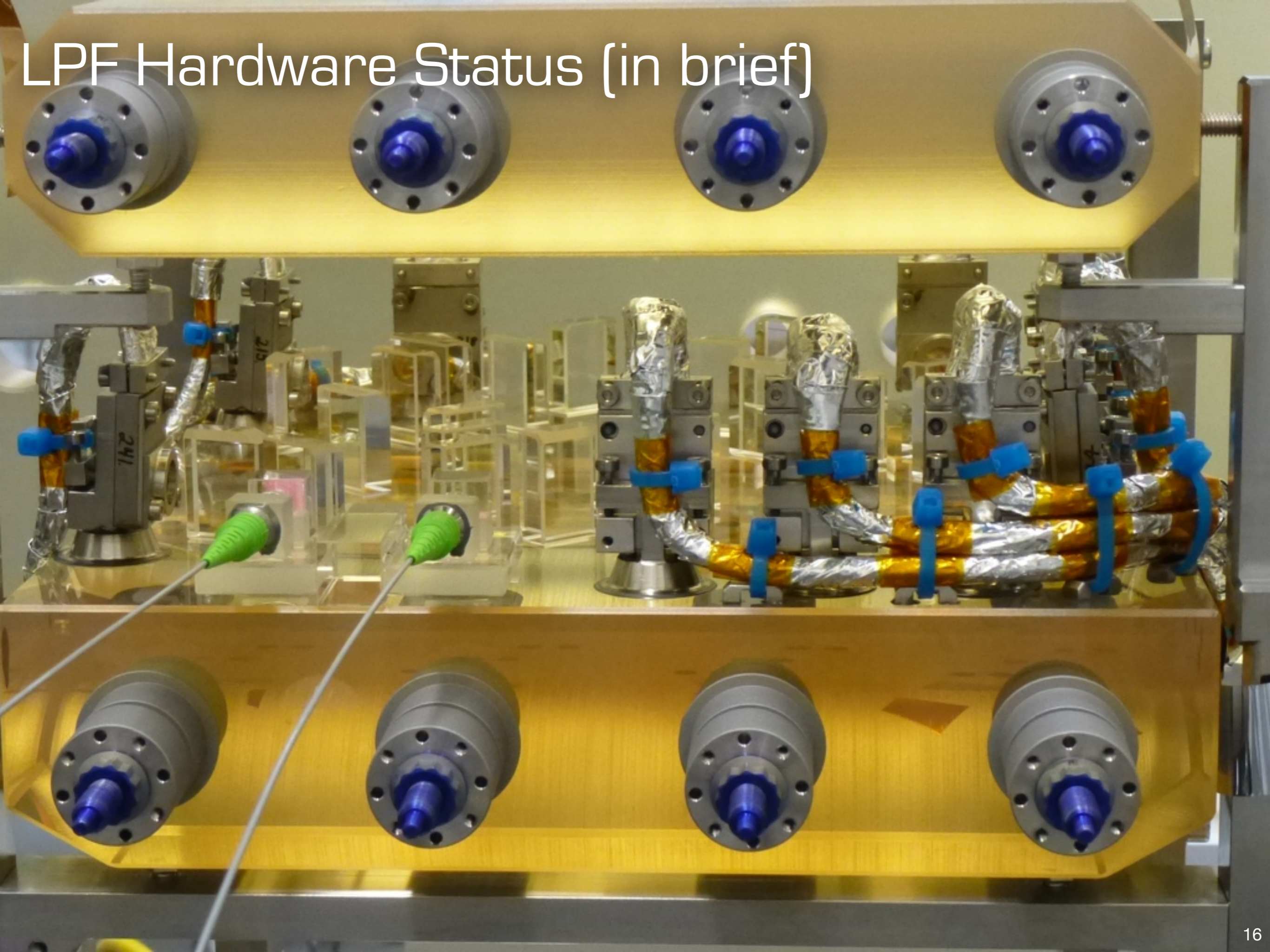
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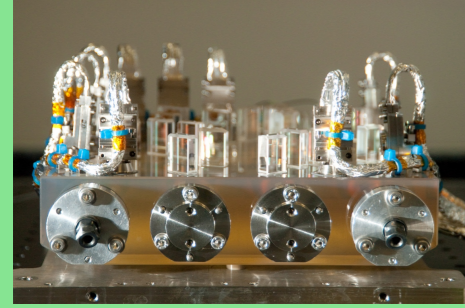
# Spacecraft



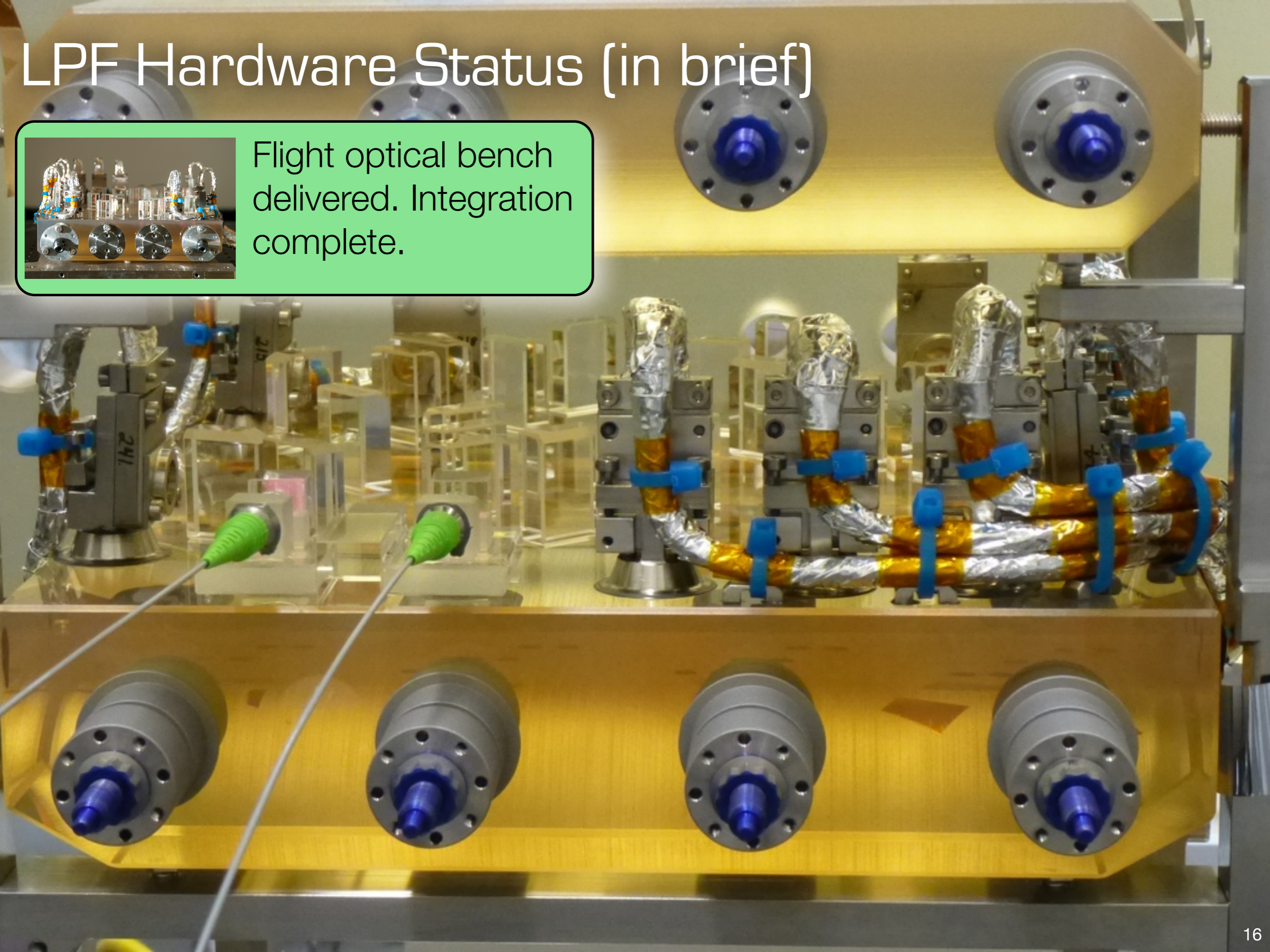
# LPF Hardware Status (in brief)



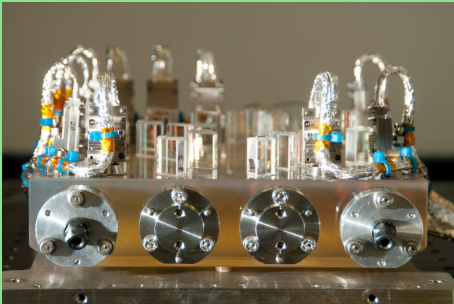
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Flight optical bench delivered. Integration complete.



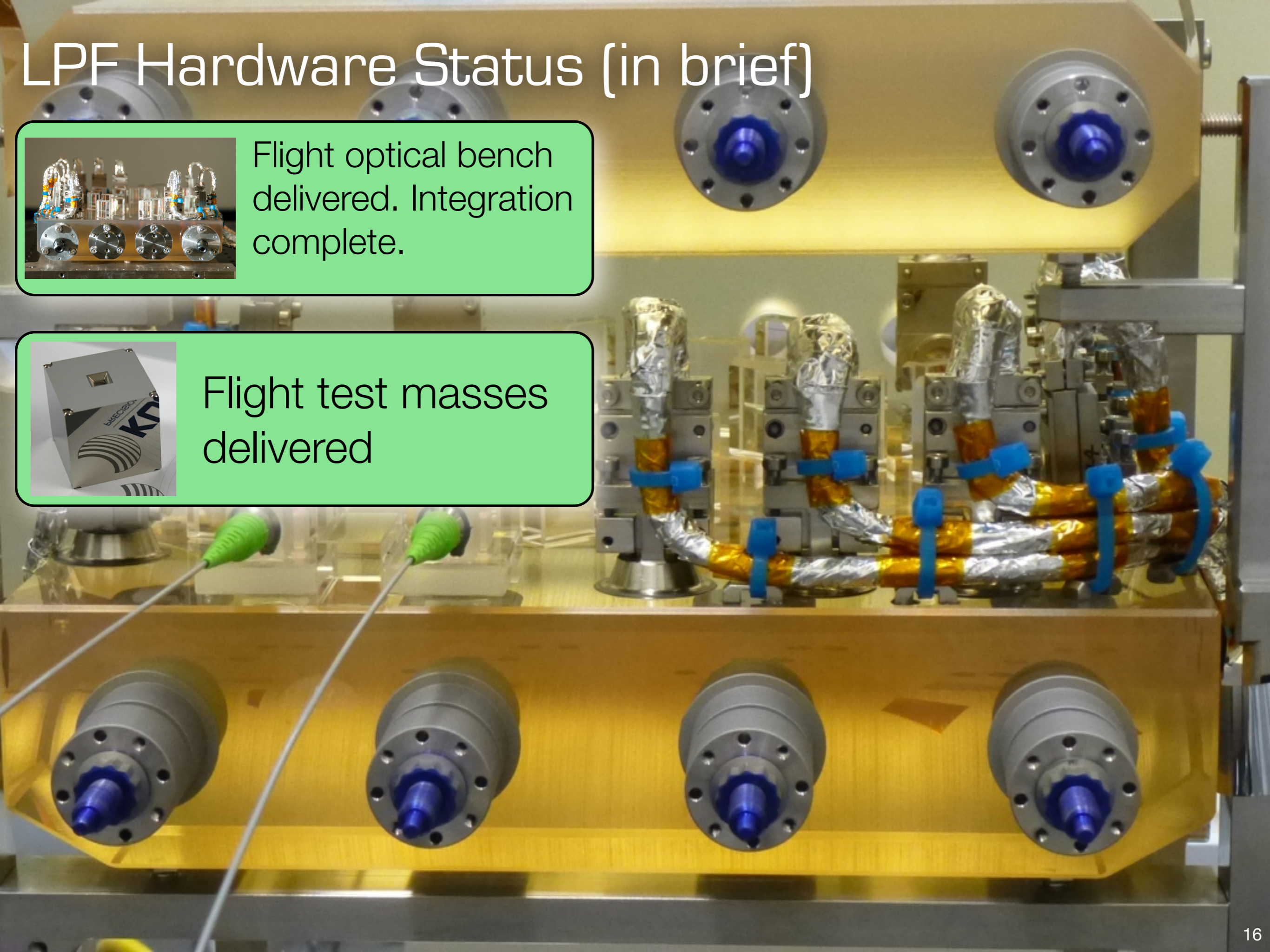
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Flight test masses delivered



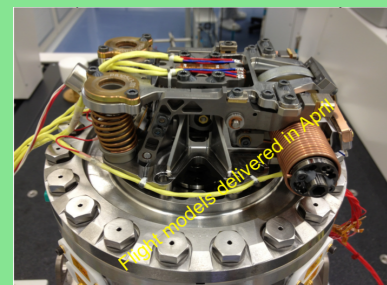
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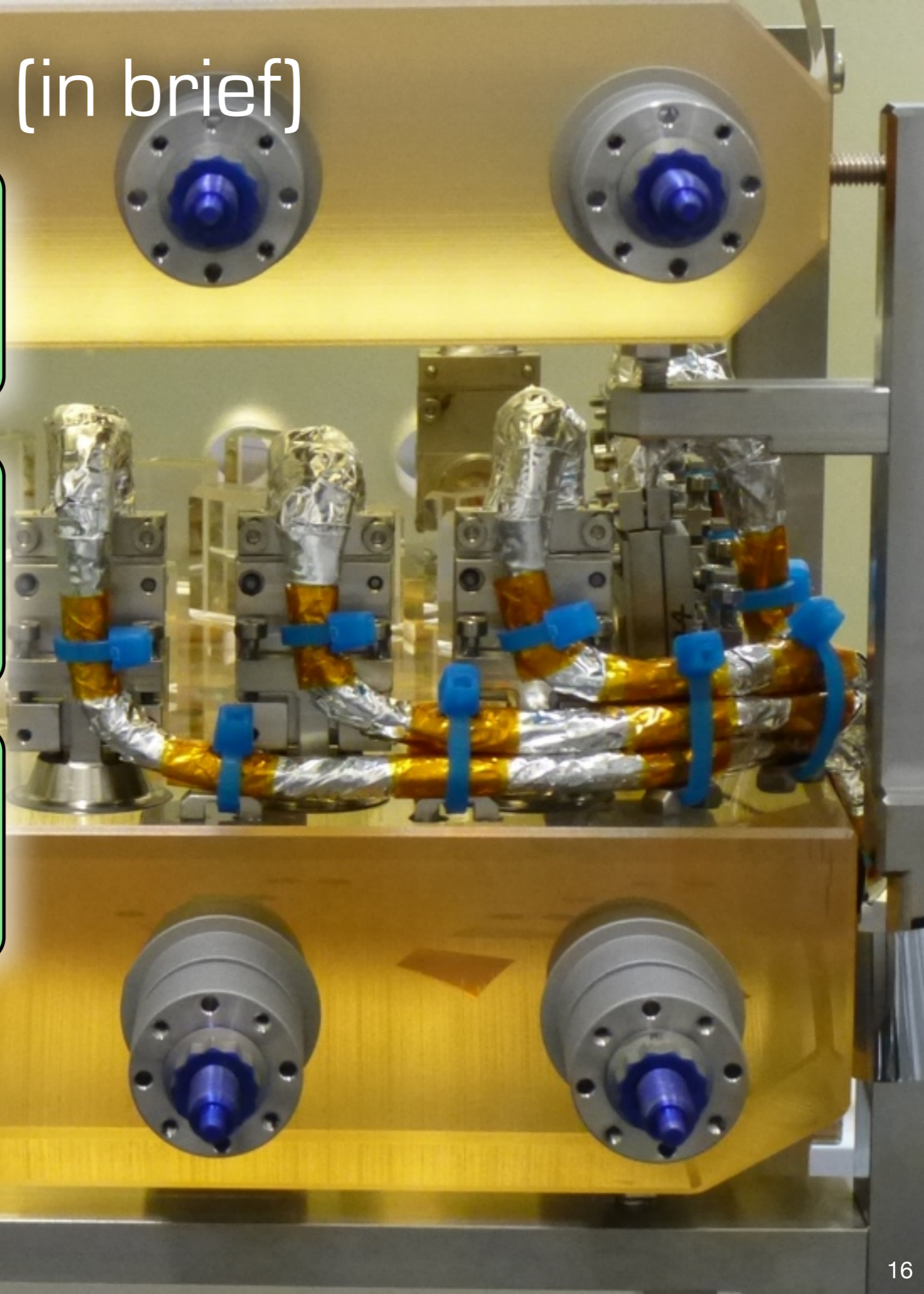
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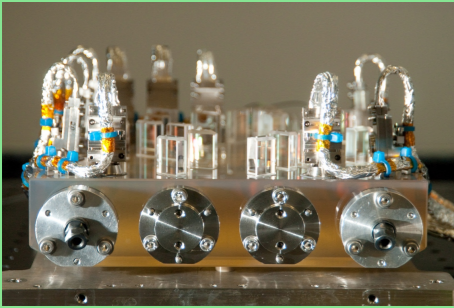


Caging mechanism delivered





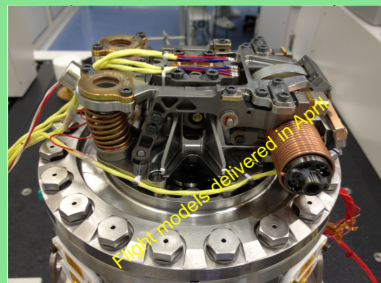
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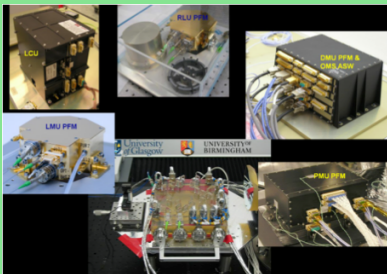
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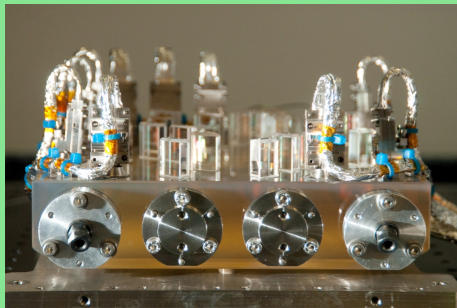
Caging mechanism delivered



Laser and electronics on SC



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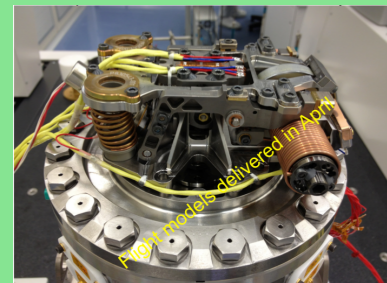
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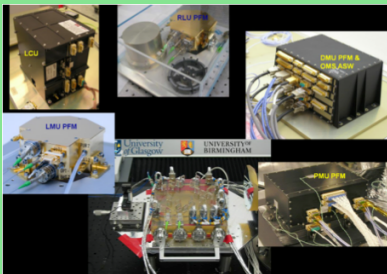
SC Module is ready to integrate payload.



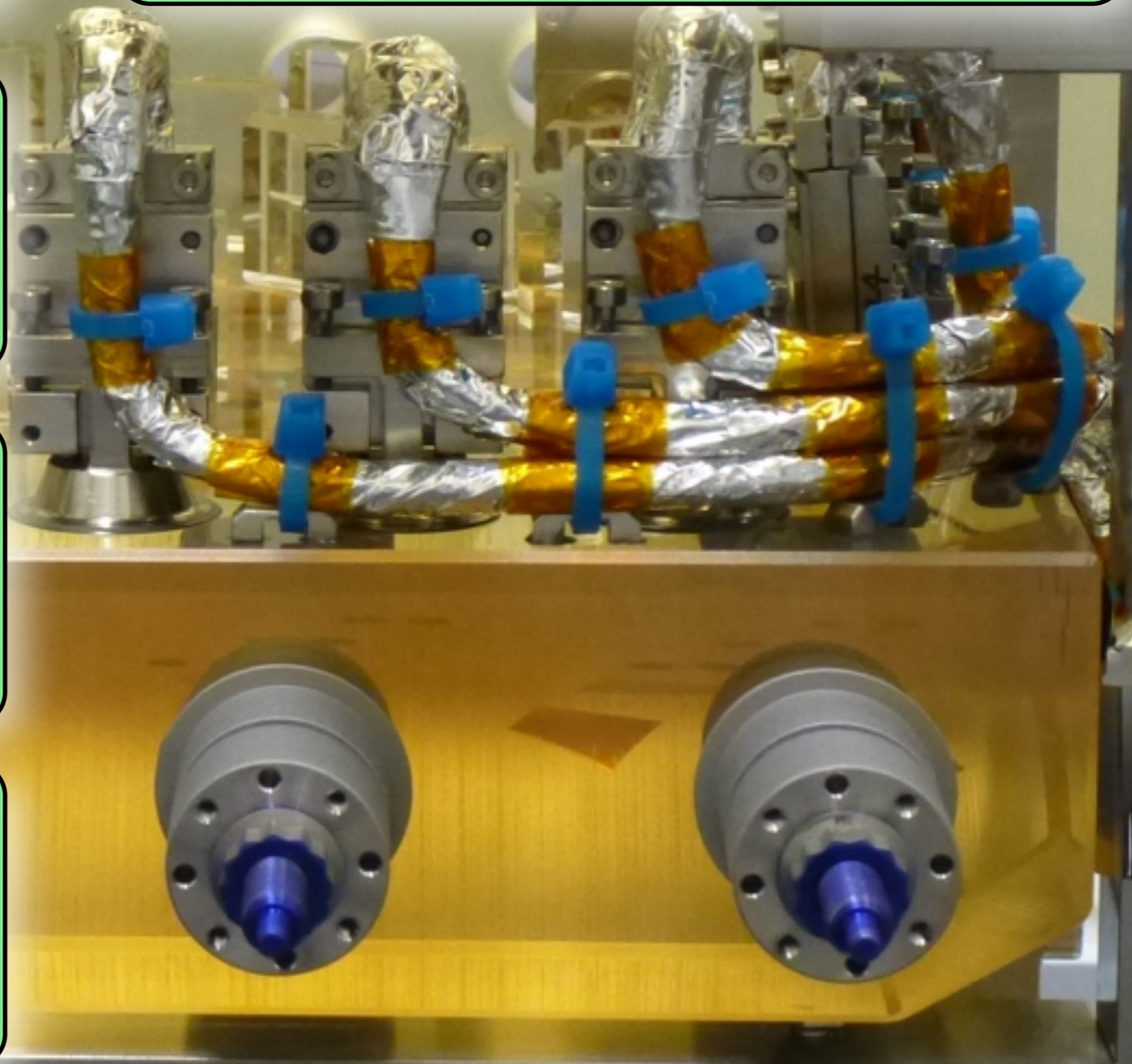
Flight test masses delivered



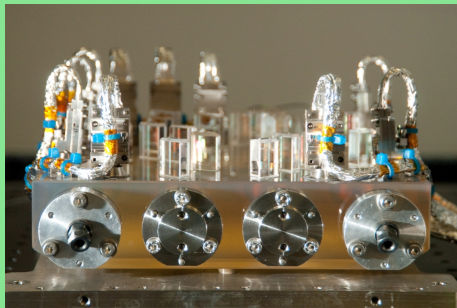
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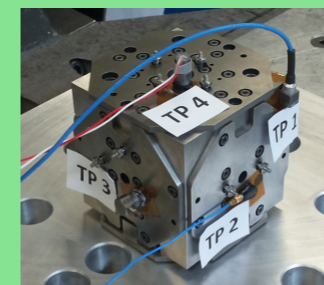
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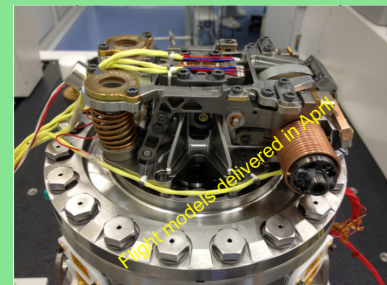
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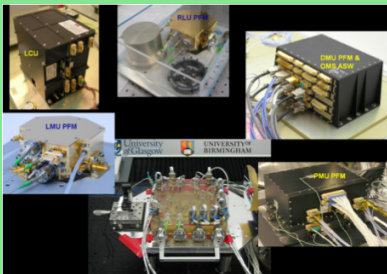
Flight test masses delivered



Redesigned electrode housing - both flight units delivered. Integration complete.

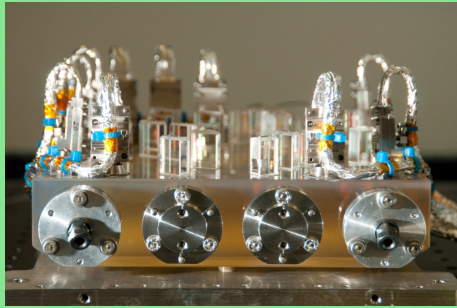


Caging mechanism delivered

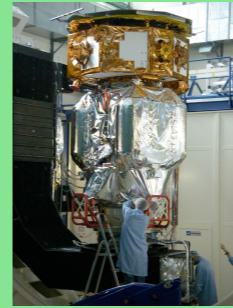


Laser and electronics on SC

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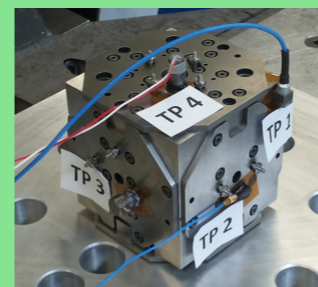
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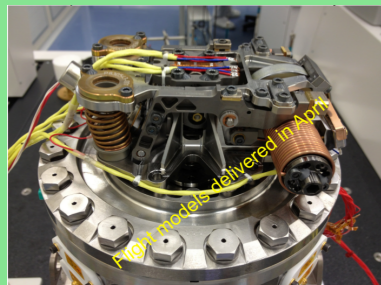
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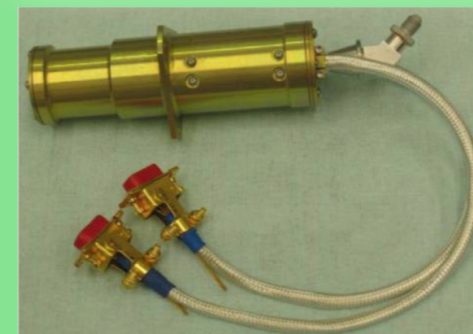
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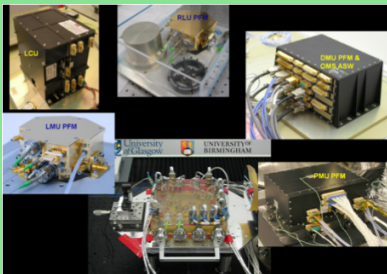
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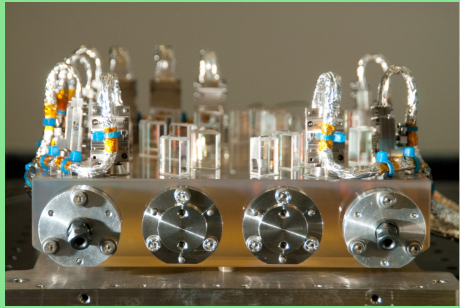


Cold Gas thrusters: off the shelf, flying on Gaia.



Laser and electronics on SC

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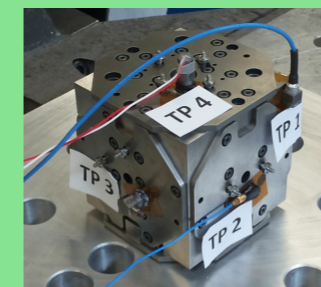
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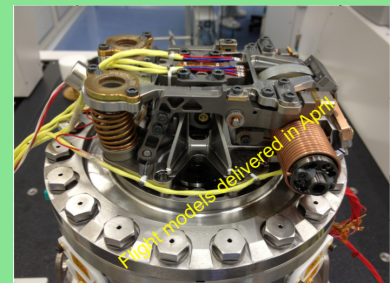
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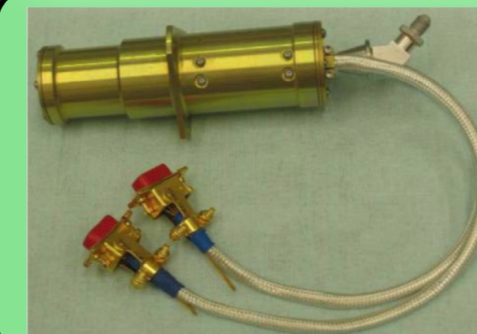
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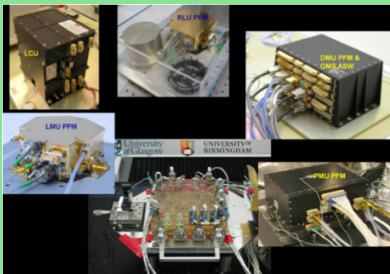
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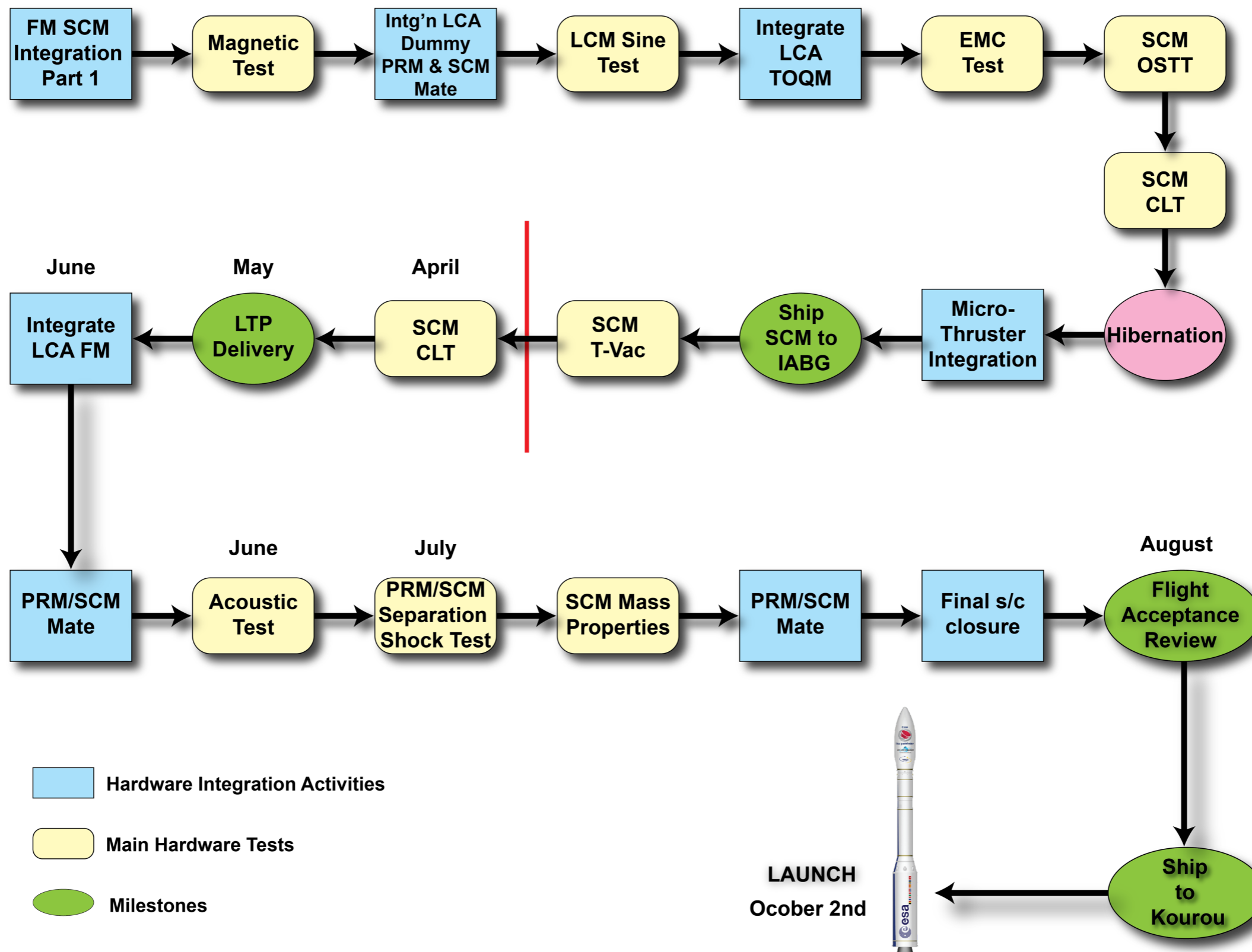


Laser and electronics on SC



Well on-track for Oct 2015 launch

# Path to launch



# Operations



Launch

Launch, LEOP, Transfer, Separation, De-spin	Commissioning	LTP Science Ops	DRS Commissioning	DRS Operations
60 days	14 days	3 months	10 days	3 months

IOCR

	Day 1	Day 2	Day 3	Day 4
H1	Noise Run	Discharge	Noise Run	Discharge
H2		Working Point		Stray Potentials
H3	Sys ID			
H4				
H5				

# What's on the menu?



Measurement of Parasitic Voltages

Measurement of differential acceleration noise on LISA Pathfinder

Measurement of cross-talks in the system

Measurement of LTP dynamical coefficients by system identification

Analysis of Data from the Radiation Monitor on LISA Pathfinder

Thermal experiments on board the LTP

Magnetic experiments on board the LTP

OPD noise investigations for LTP

Laser frequency noise characterisation for LTP

Laser Amplitude Noise Characterisation for LTP

The Drift Mode for LISA Pathfinder

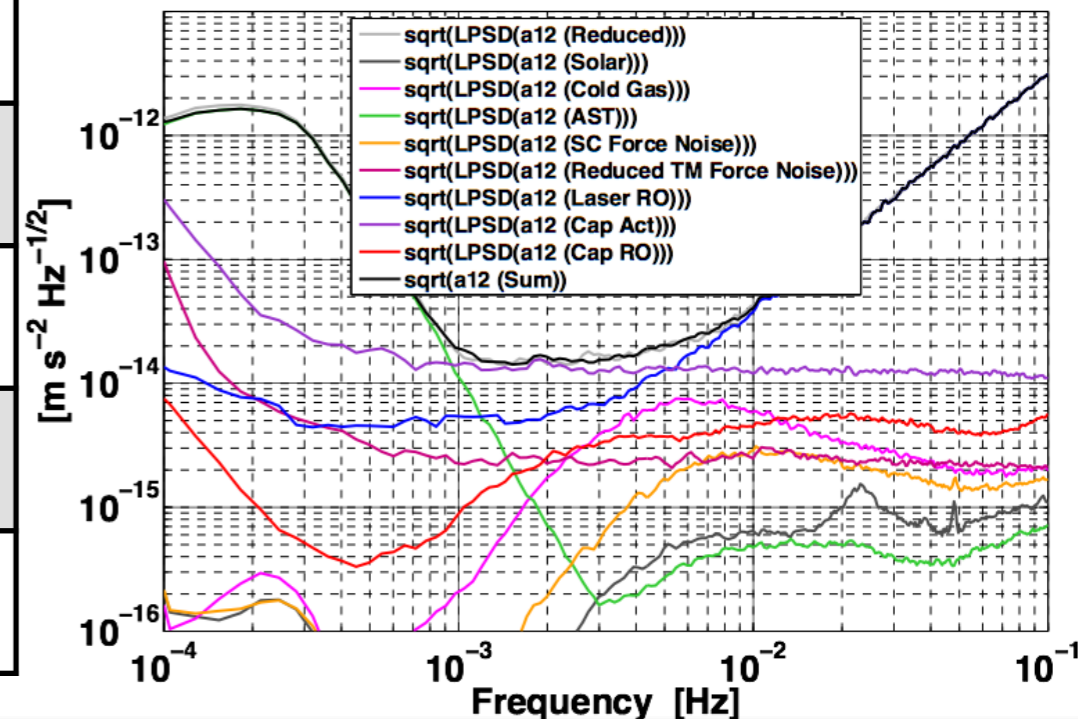
Modulate TM2 z electrodes with compensation voltage -100mV

Modulate TM2 y electrodes with compensation voltage -100mV

Modulate TM2 x electrodes with compensation voltage -100mV

...

LTPDA 2.7.clev (R2013b Pre-release)  
2013-08-22 12:08:28.016 UTC  
ltpda: ec16324  
ose52\_breakdown





# Starting science operations



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- The first two weeks are all about gathering information and gaining experience
- This is our first interaction with the system
- Focus on:
  - noise runs
  - first tests of signal injection (system identification)
  - getting a handle on the charge rate and discharging

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	Hour																								
	0	1	2	3	4	5	6	7	8	9	10	11	12	13	14	15	16	17	18	19	20	21	22	23	
1	Noise run in Sci 1.2																								
2	CE1	CE2	Noise run in Sci 1.2																						
3	CE1	CE2	Sys ID (low amp)										Noise run in Sci 1.2												
4	CE1	CE2	Working point scan (x,y,z), both TMs																						
5	CE1	CE2	Cross-talk investigations, low amplitude																						
6	CE1	CE2	Noise run in Sci 1.2																						
7	Station Keeping											Transition Acc3 -> Sci 1.2							FD1	FD2					

# Estimating Residual Differential Acceleration



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- Understanding the purity of the free-fall we achieve, and what limits it, requires us to assess the residual forces acting on the TMs
  - what's left when we subtract the forces we can account for



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- We compute the relative acceleration of the two TMs based on the observed relative position

$$g_{\text{res}} = x_{12}^{\ddot{}}$$



- Understanding the purity of the free-fall we achieve, and what limits it, requires us to assess the residual forces acting on the TMs
  - what's left when we subtract the forces we can account for
- We compute the relative acceleration of the two TMs based on the observed relative position

$$g_{\text{res}} = \ddot{x}_{12}$$

- Try to account for the contributions of  $g_{\text{res}}$  that we know

- applied control forces
- couplings due to force gradients

$$g_{\text{res}} = \ddot{x}_{12}$$

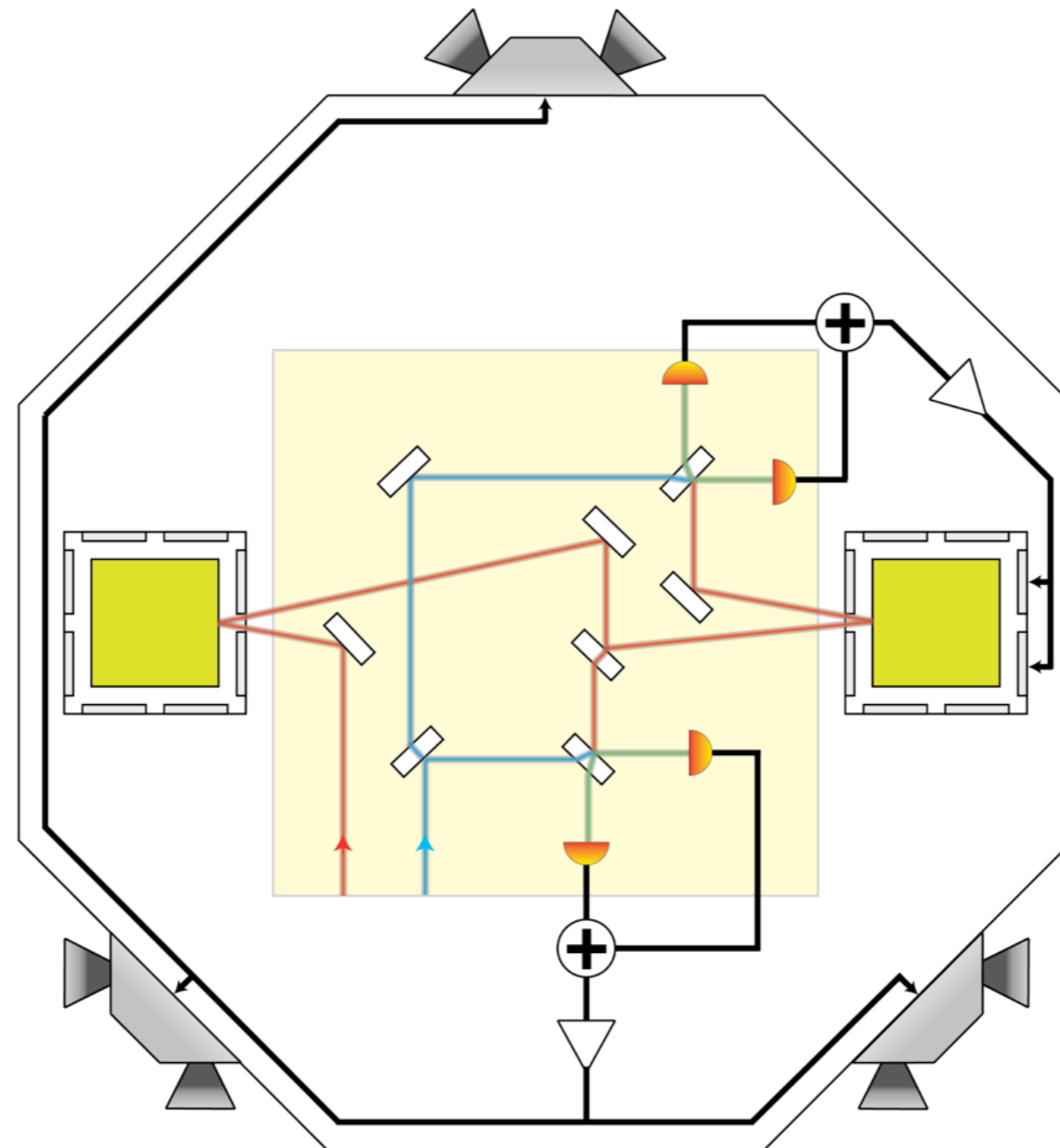
→  $-g_{\text{control}}$

→  $-\omega_{\Delta}^2 x_1 - \omega_2^2 x_{12}$

# System Identification



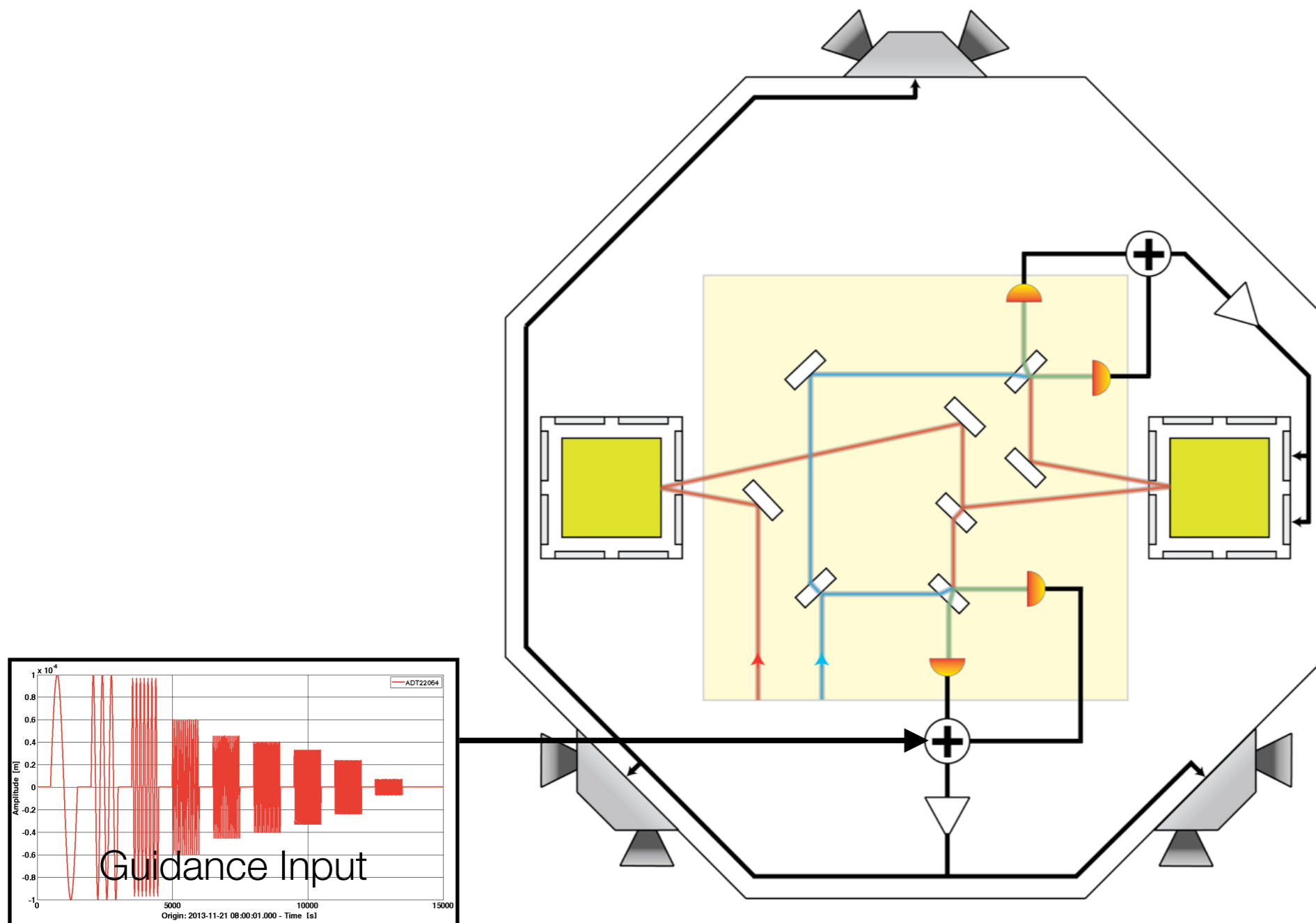
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# System Identification



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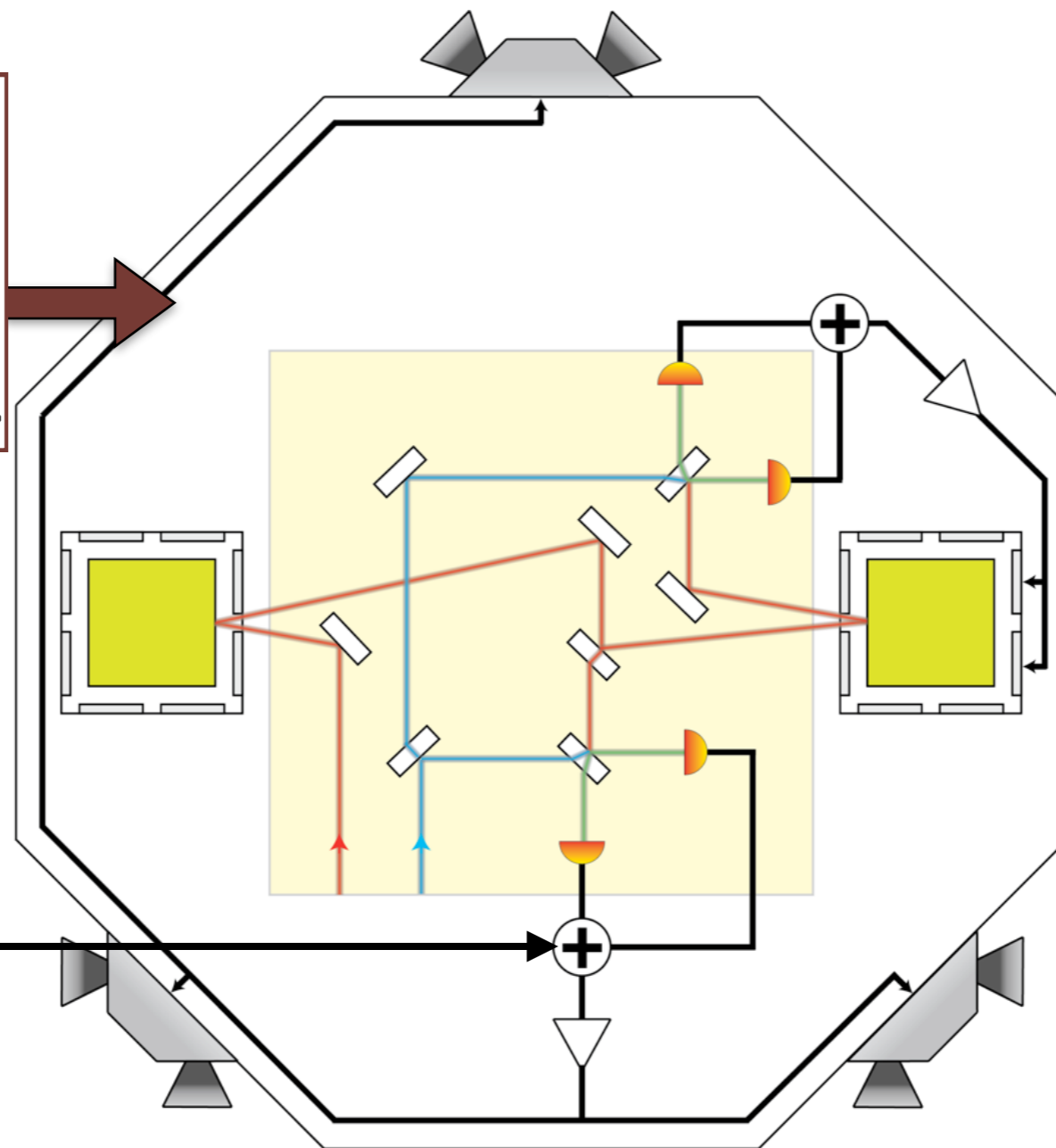
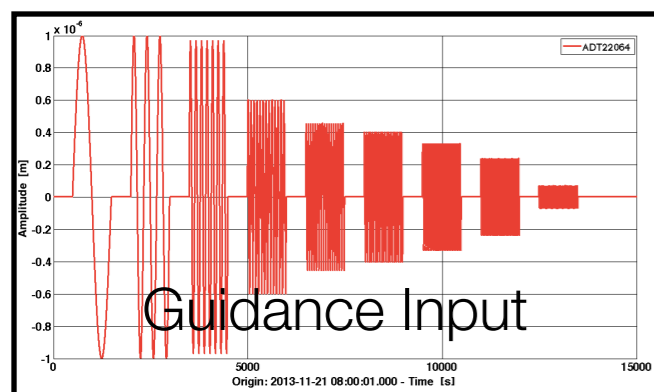
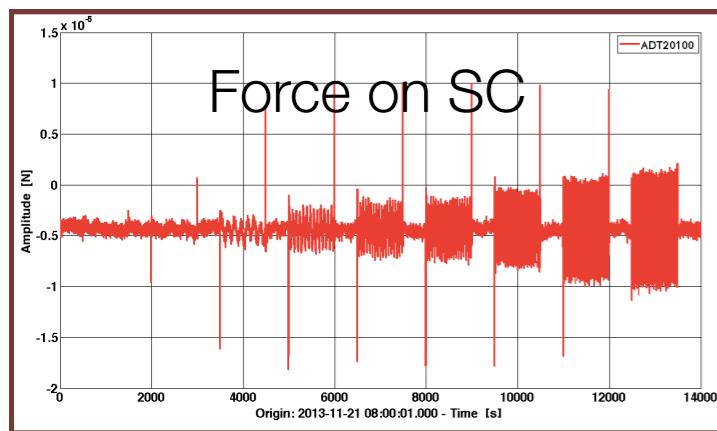




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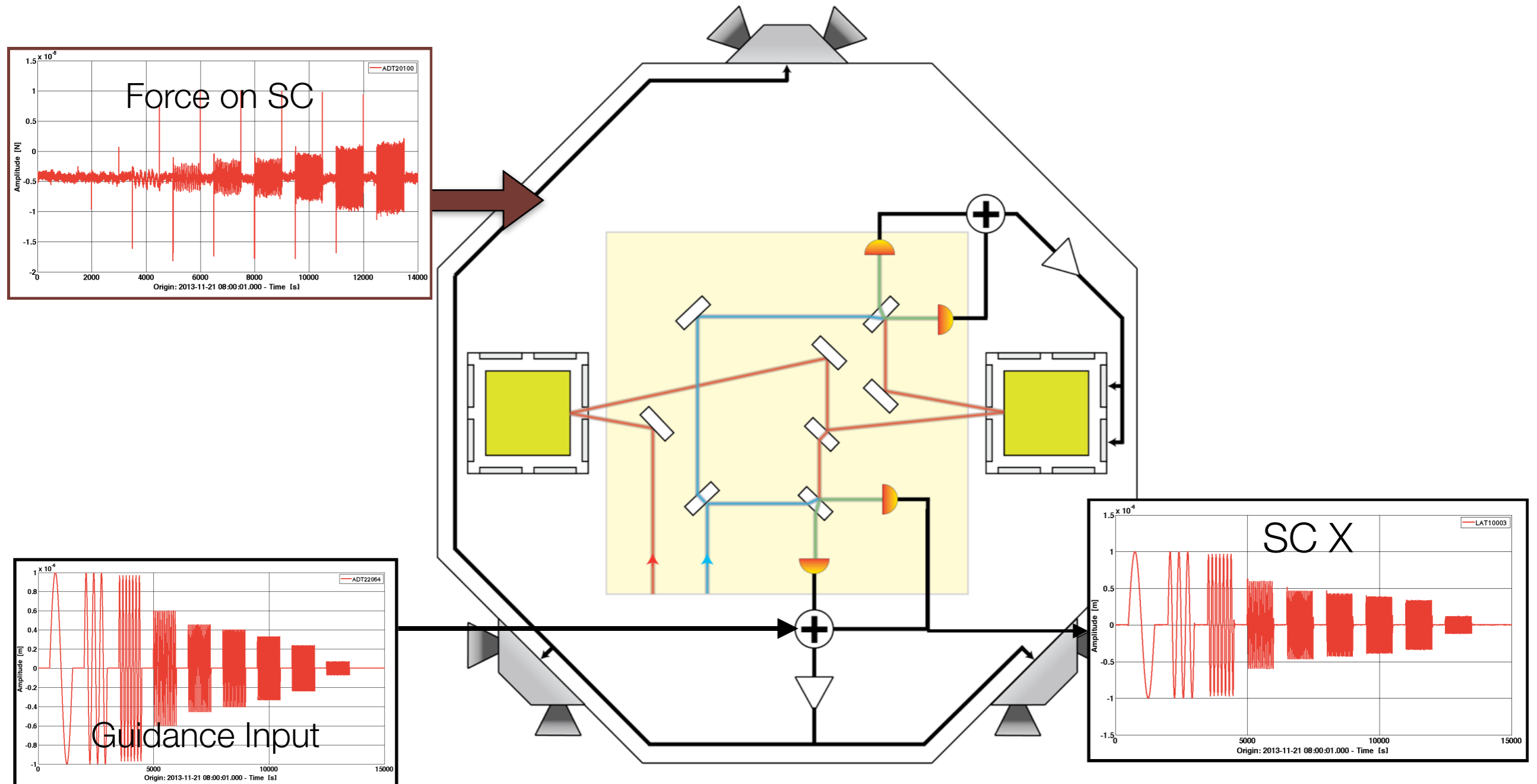
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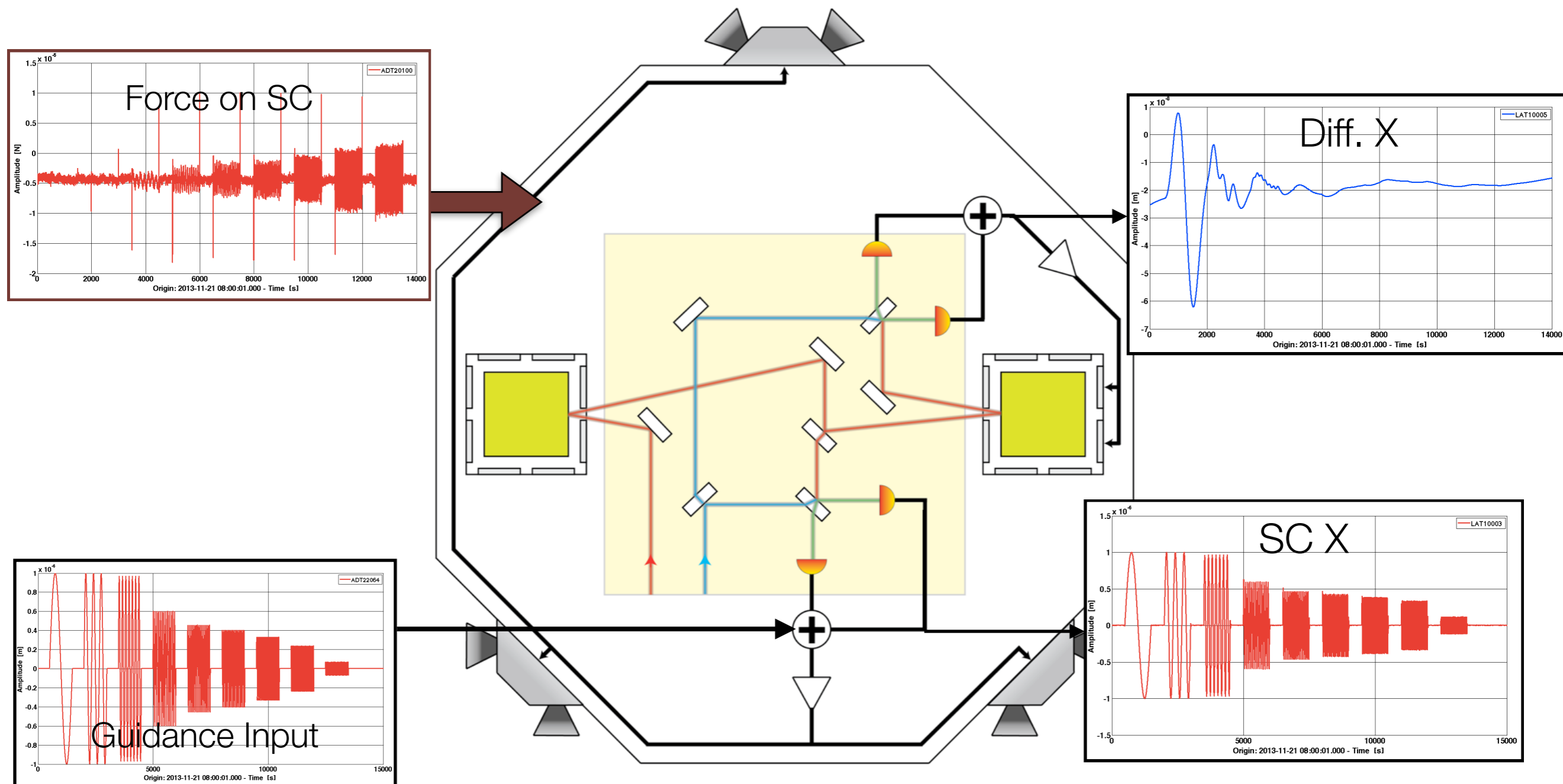
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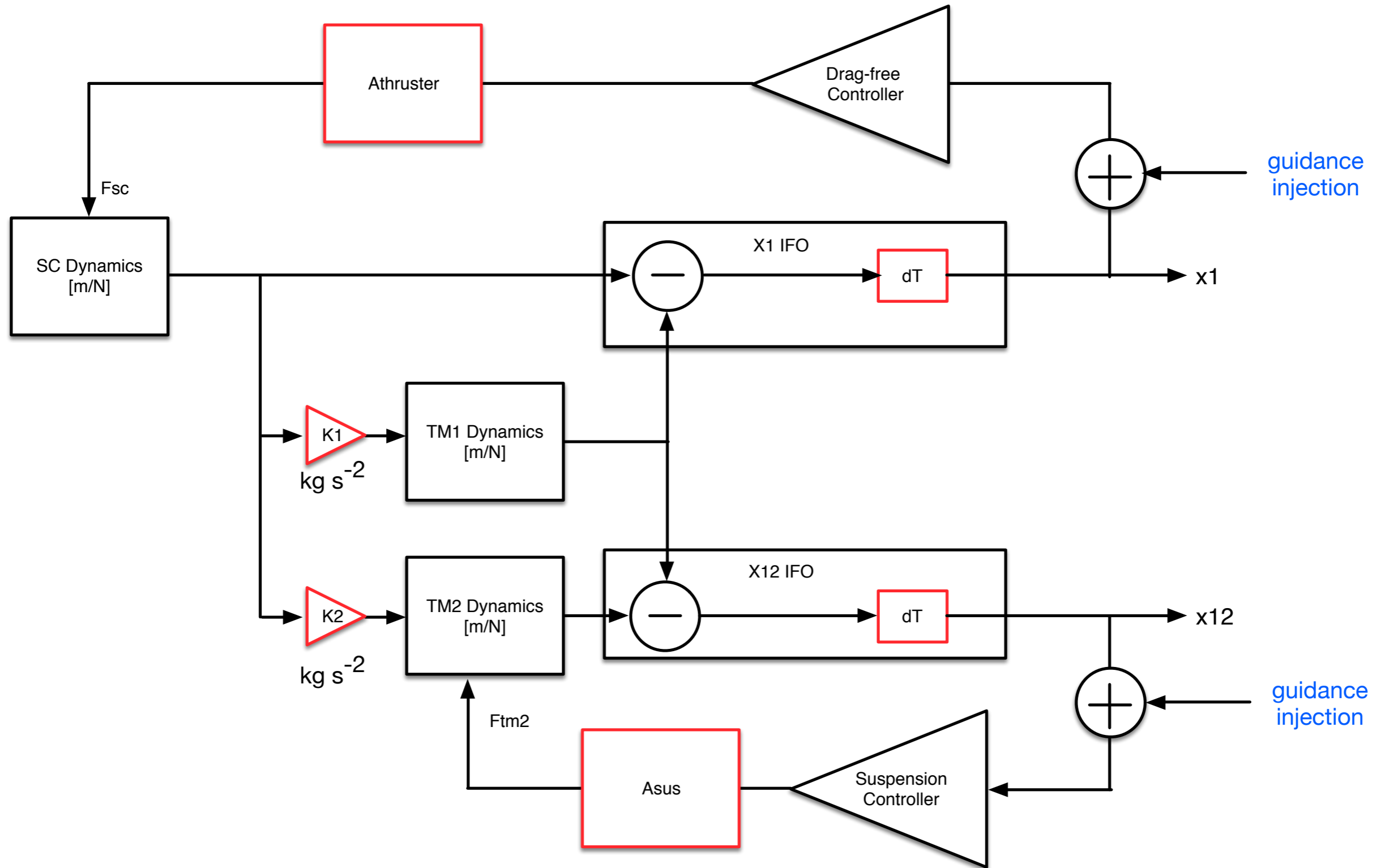
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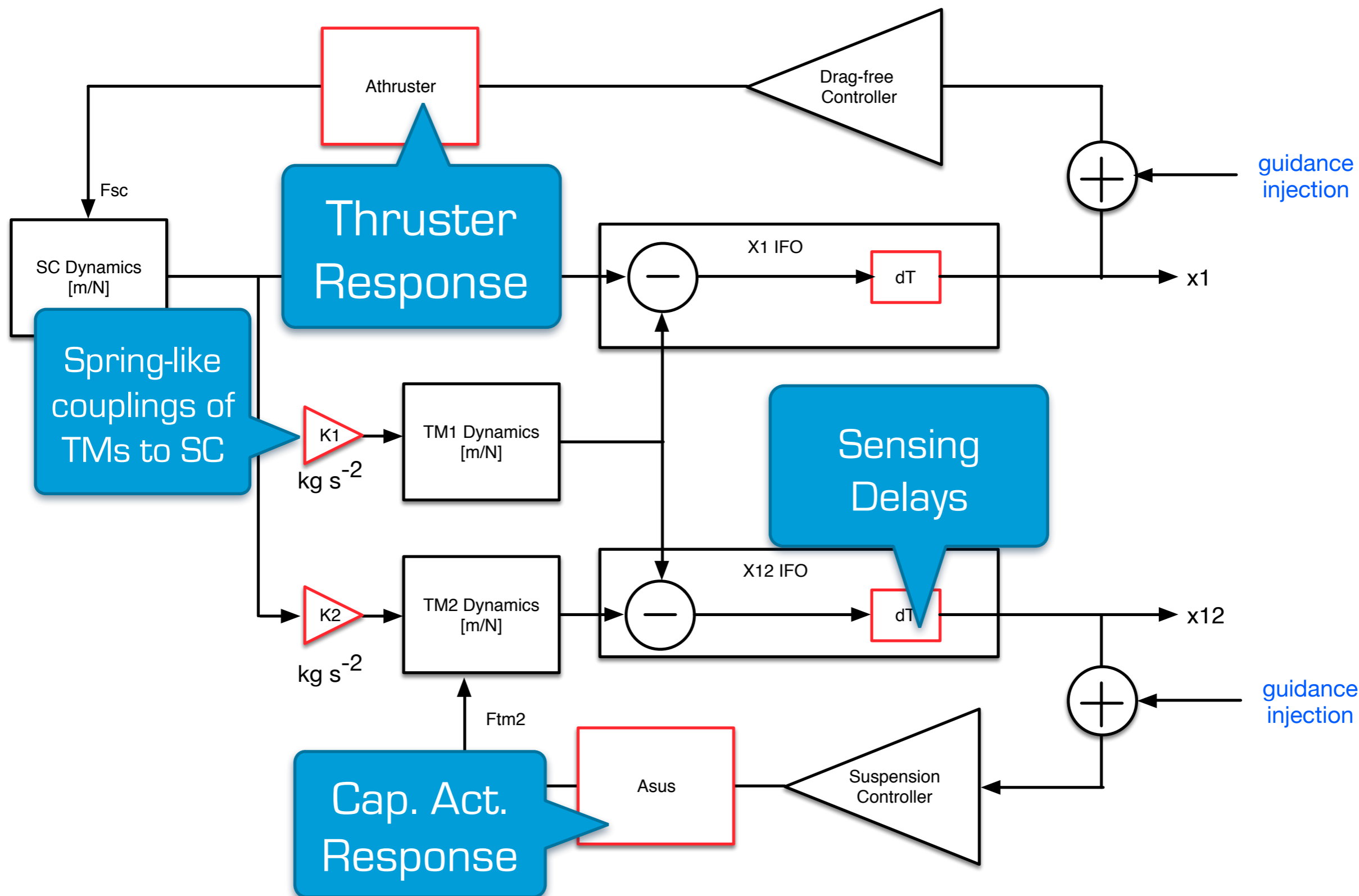
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# Parameter estimation

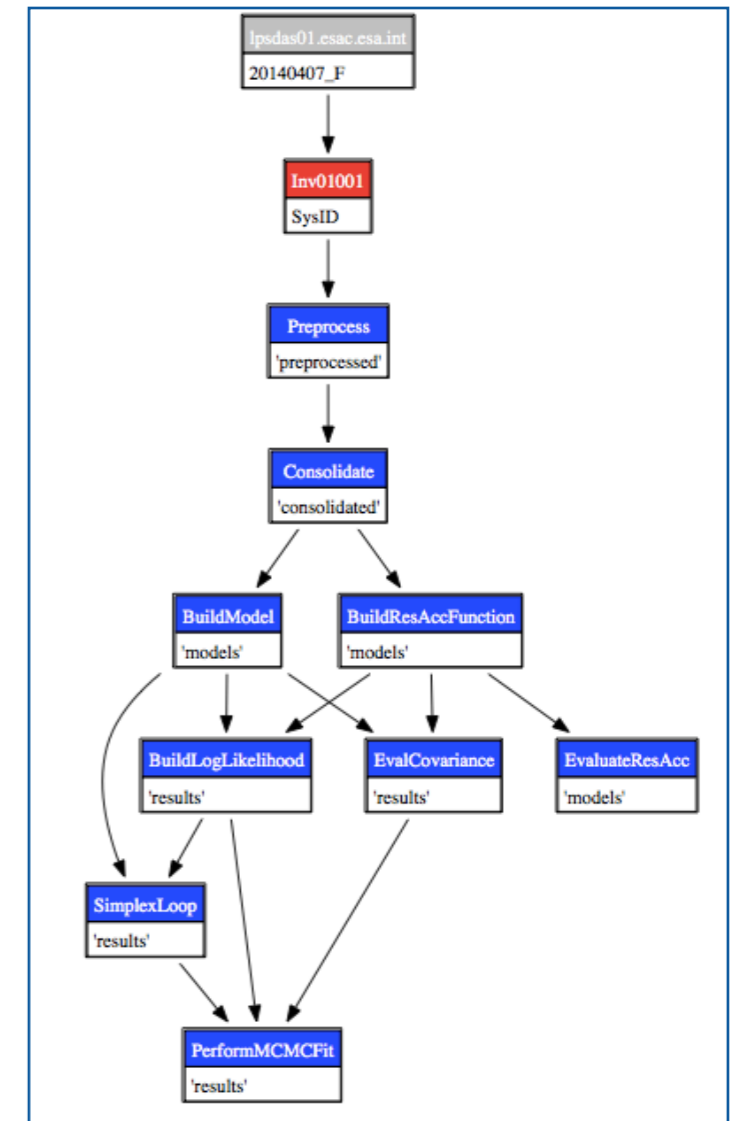


# Parameter estimation





- Follows the same form as for estimating residual differential acceleration
- But now the coefficients in the model are fit so that the linear combination of terms fit the observation
- When a good fit is found, the residuals contain no trace of the injected signals



Fit

$$\text{observation} = x_{12}''[k]$$

to

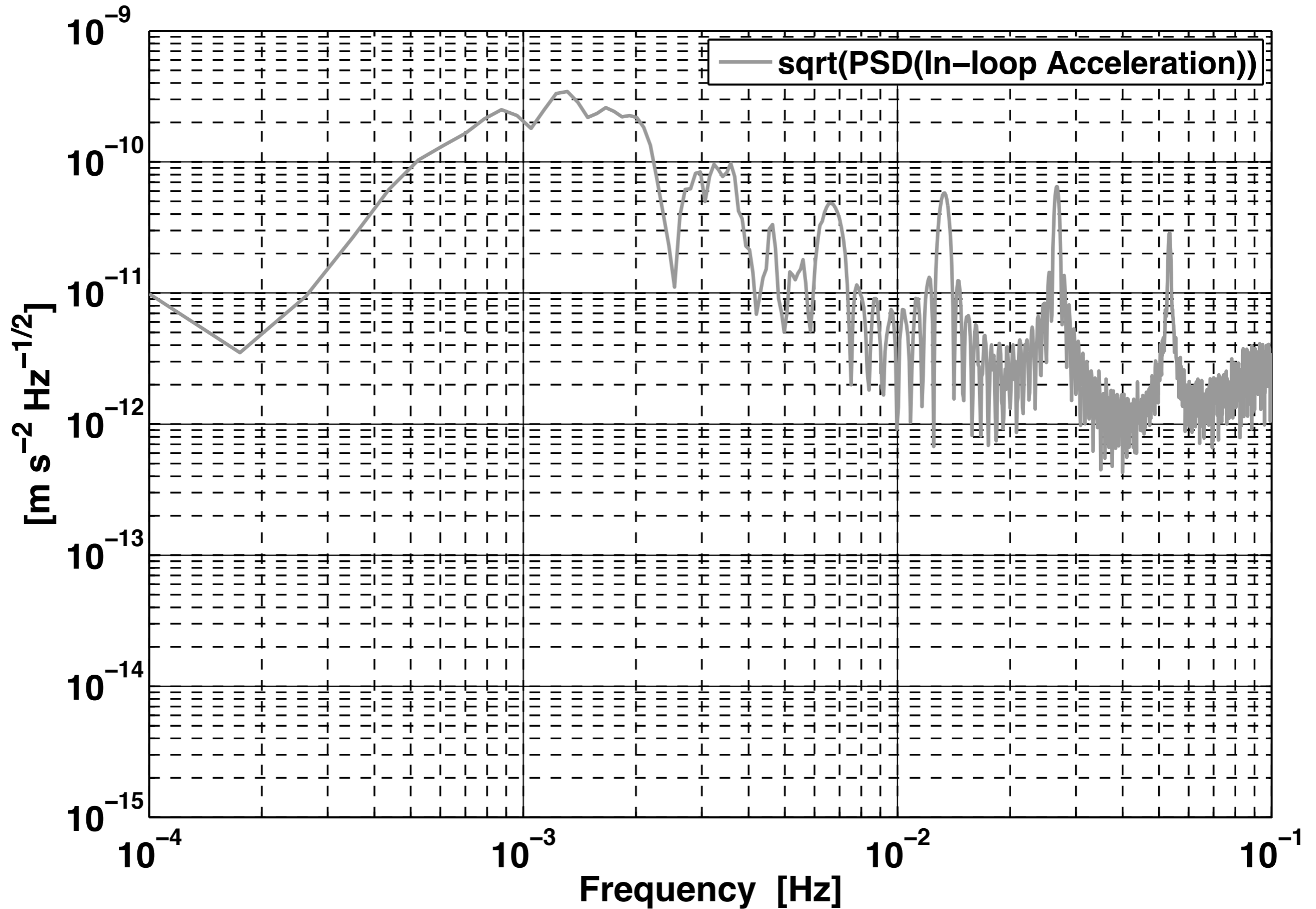
$$\begin{aligned} \text{model} = & -A_{\text{sus}} F(g_{\text{control}}[k], \Delta T) \\ & -(\omega_2^2 - \omega_1^2)x_1[k] \\ & -\omega_2^2 x_{12}[k] \end{aligned}$$

# Residuals



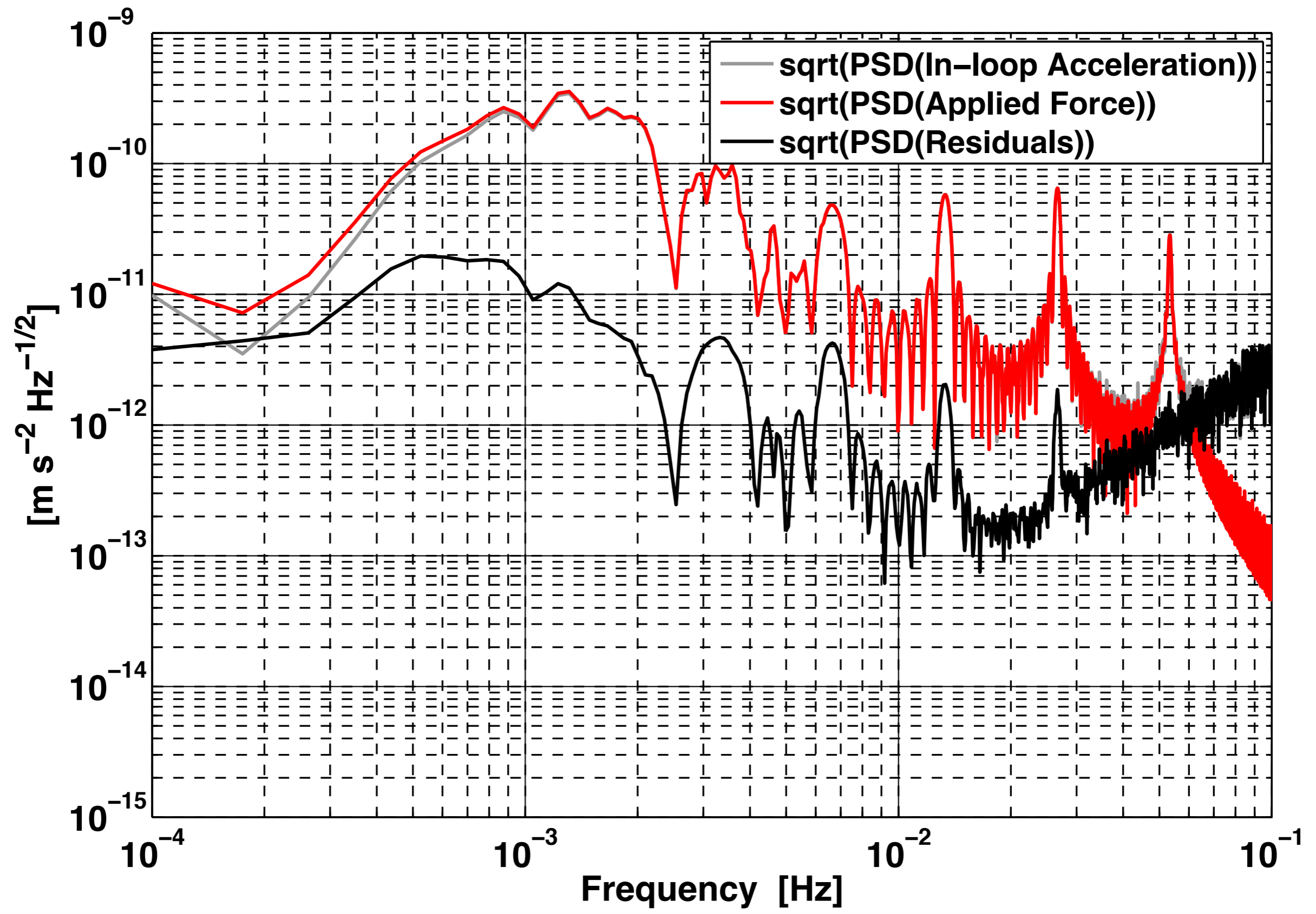
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# Residuals

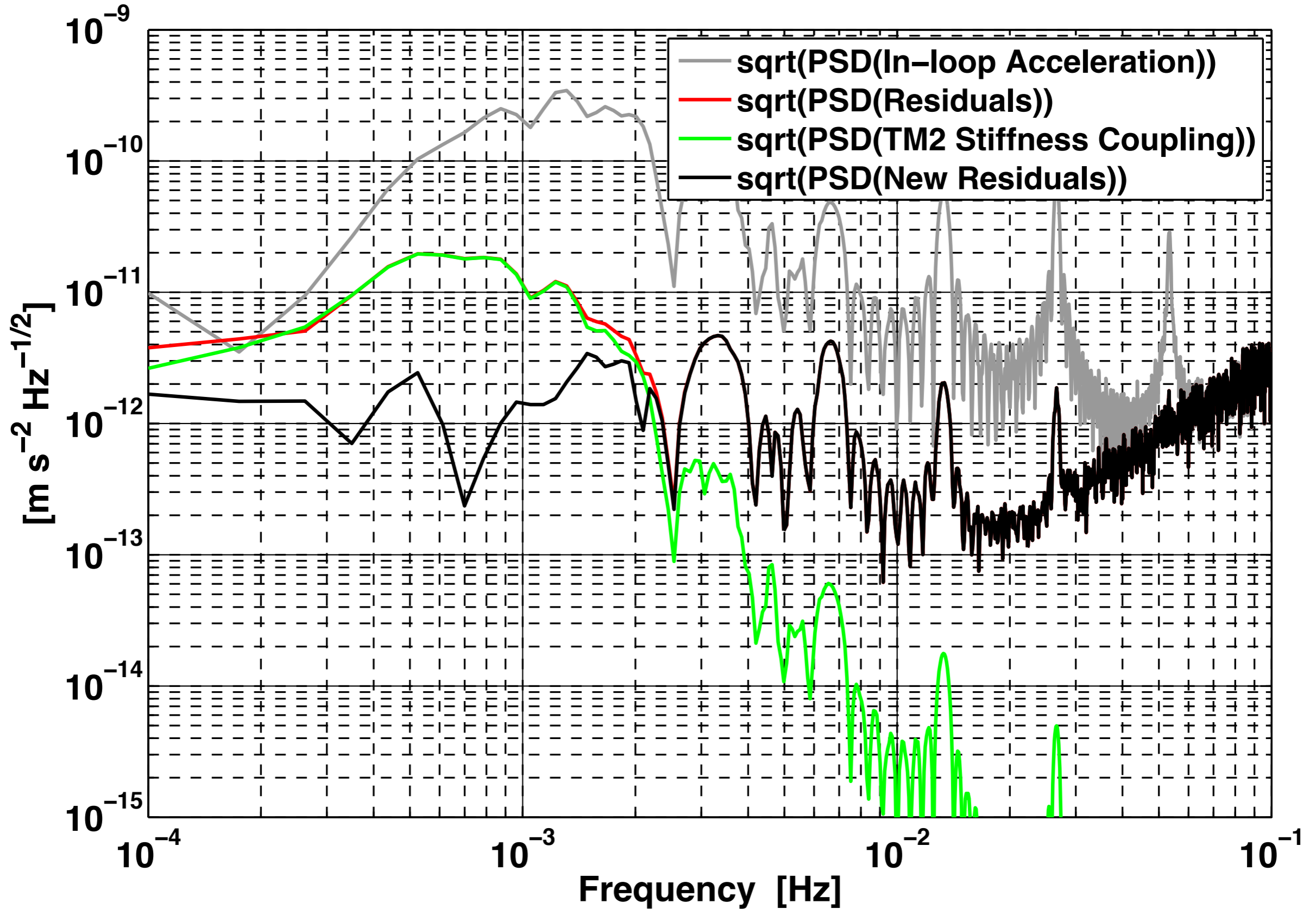




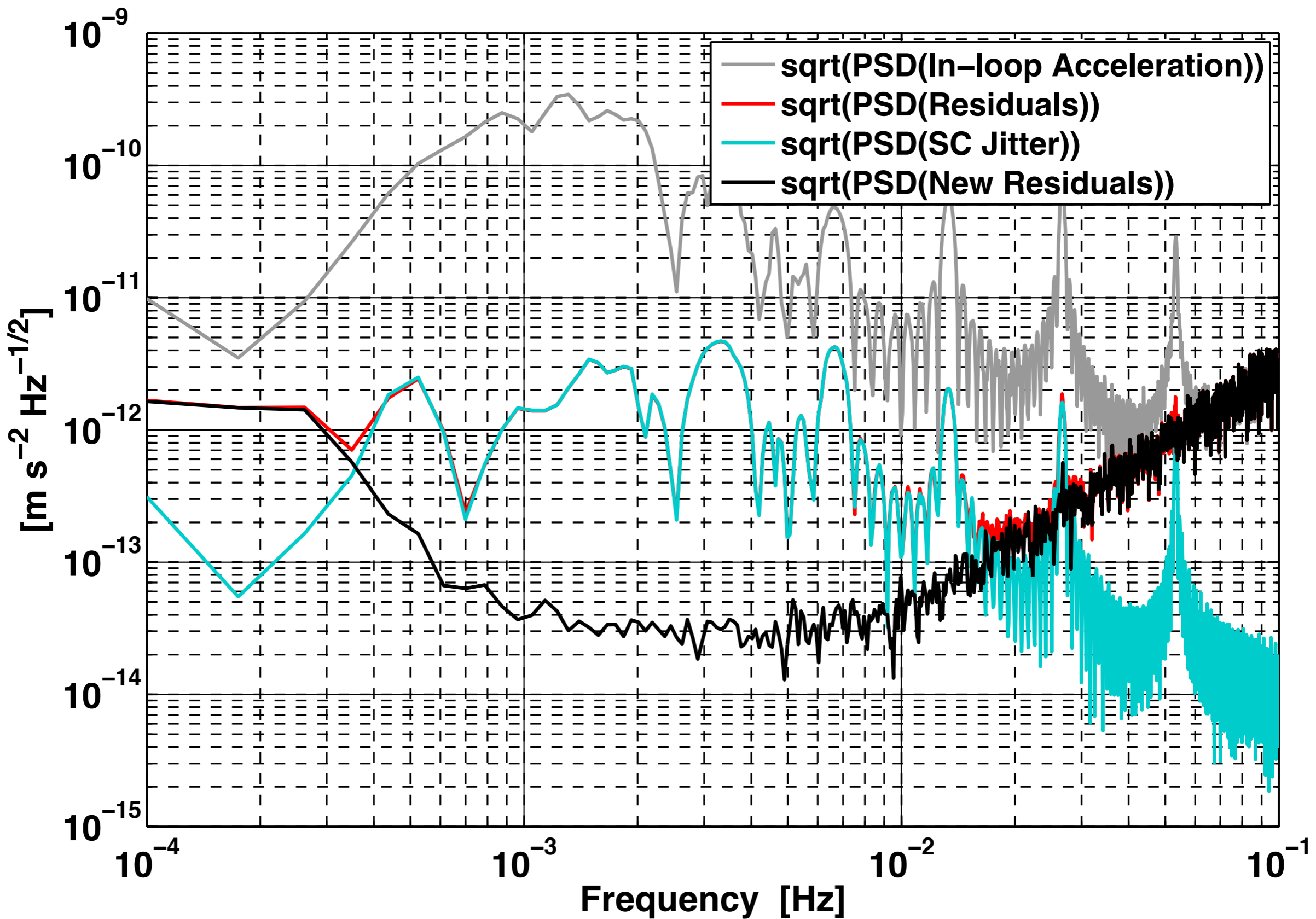
# Residuals



# Residuals



# Residuals

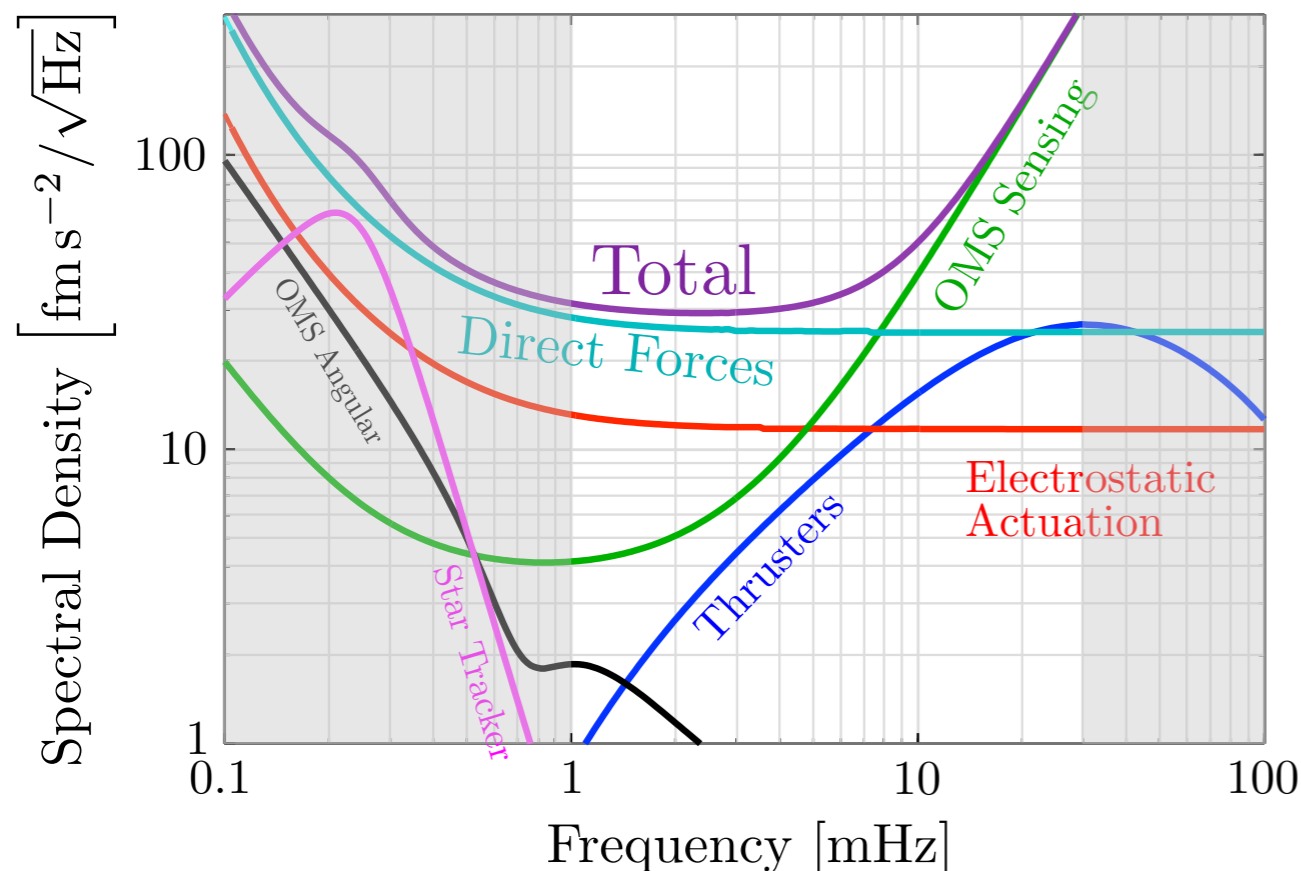


# Noise Budget

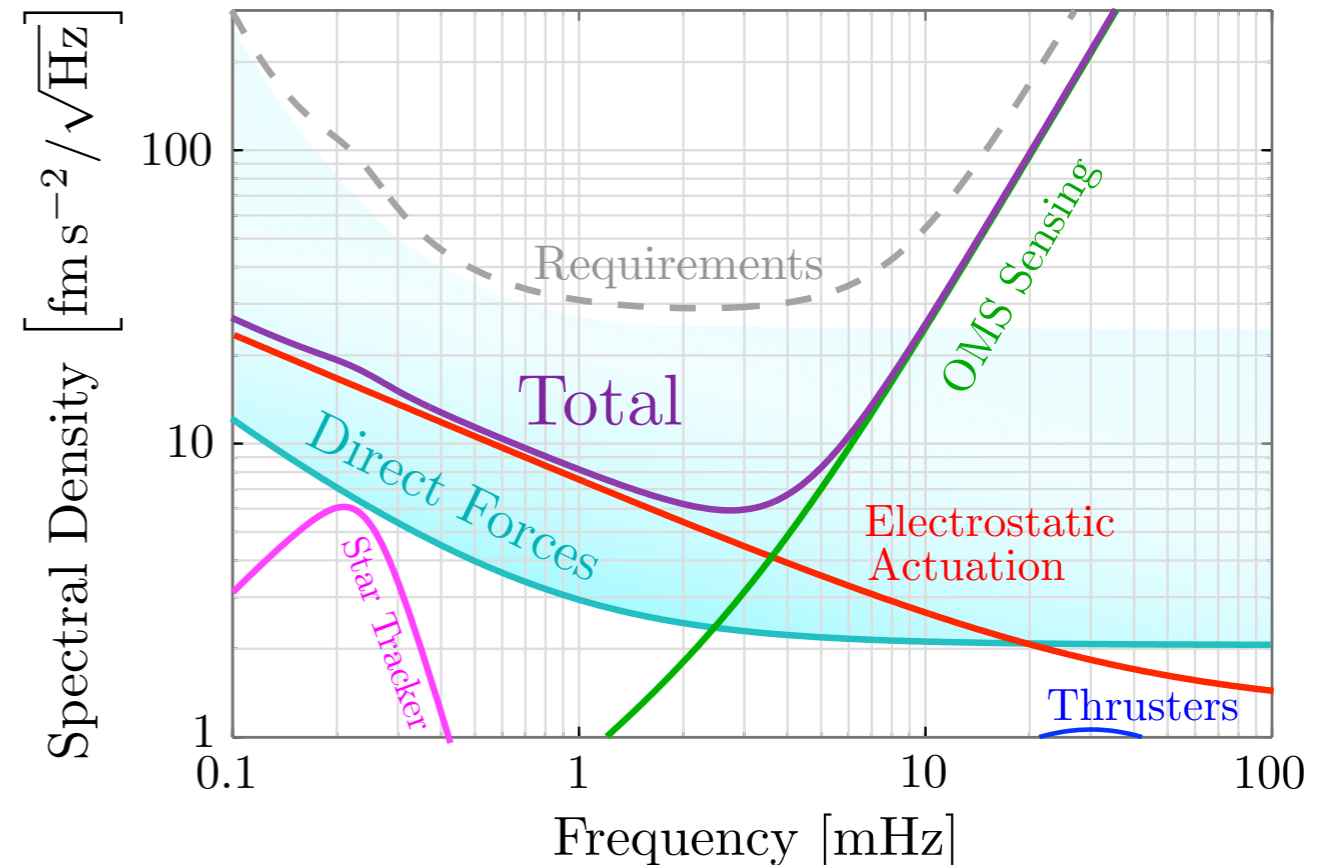


- How does our observed residual differential acceleration differ from what we expect?
- Why does it differ?
  - this drives the next activities to be performed

Requirements Noise Breakdown



Current Best Estimate



# Preparing for operations



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# Preparing for operations



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- Mock Data Challenges
  - kick-off development of tools
  - learn about system modelling

# Preparing for operations



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- Mock Data Challenges
  - kick-off development of tools
  - learn about system modelling
- STOC Simulations
  - more operational like simulations of segments of the science operations timeline



- Mock Data Challenges
  - kick-off development of tools
  - learn about system modelling
- STOC Simulations
  - more operational like simulations of segments of the science operations timeline
- SOVT (System Operations Validation Test)
  - formal ESA test campaign to validate the ground segment
  - Many pre-SOVT runs to prepare for this



# Preparing for the SOVT

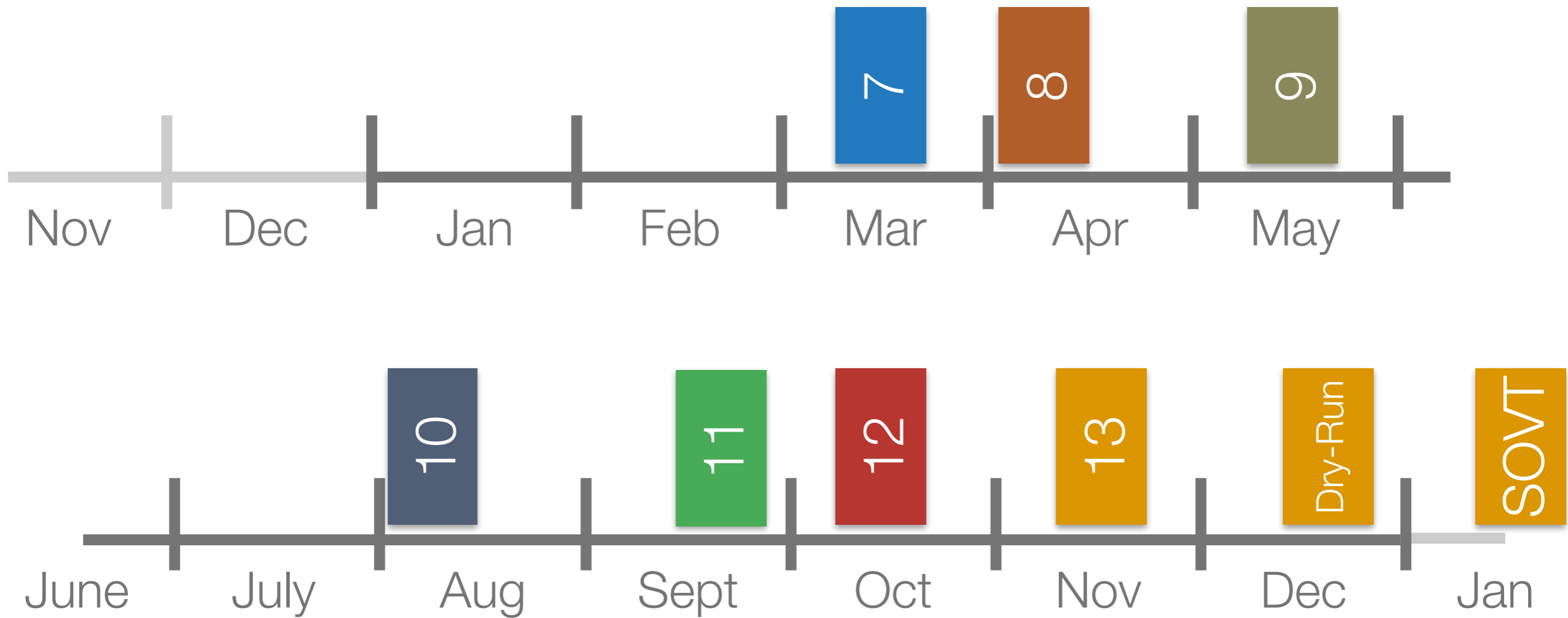


- We ran 14 pre-SOVT tests
- The first 6 of these were for developing the ground segment, commanding and analysis
- The last 8 have targeted the testing of specific investigations with the aim of covering as much of the mission time-line as we can

# Run Calendar



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- Time-line was executed on the Real-time Test Bed (RTB)
  - includes hardware IS FEE, OBC, DMU
  - SCOEes for dynamics (RTS), OMS, thrusters, etc

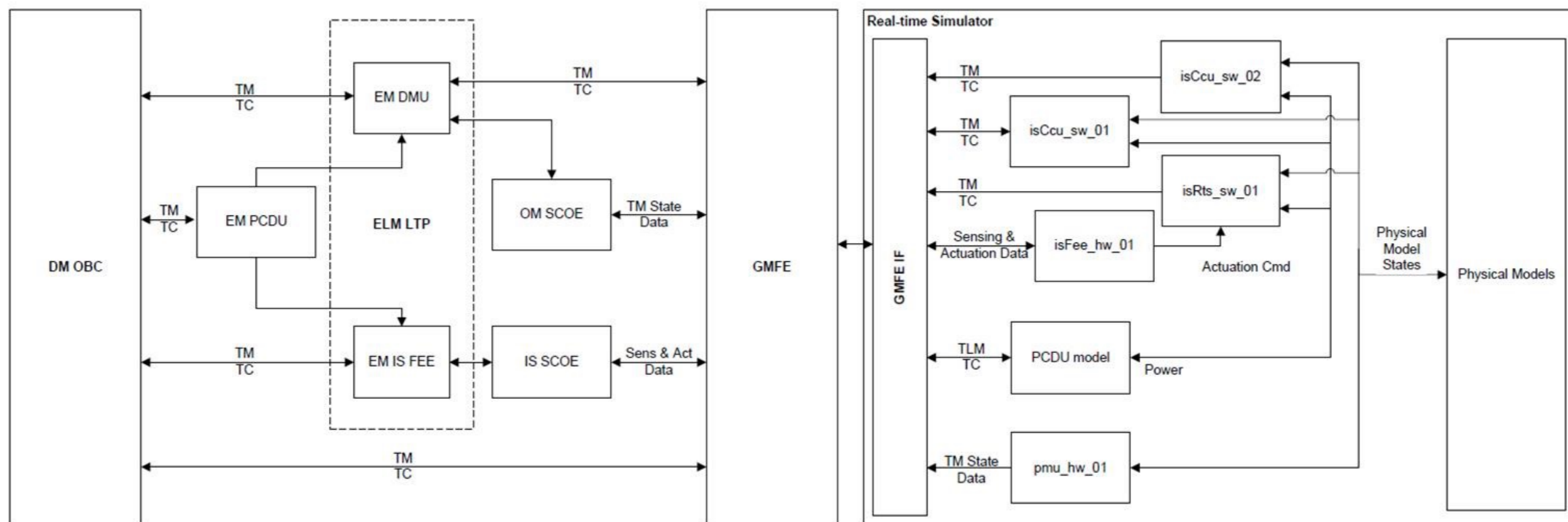


Figure 3-2: Set-up in LTP & AOC RTB configuration

# Run Constraints



	Day 1	Day 2	Day 3	Day 4	Day 5				
0		LTP Ops	LTP Ops	LTP Ops	DRS Ops				
1									
2									
3									
4									
5									
6									
7	RTB Powerup & Handover to MOC	LTP Ops	LTP Ops	LTP Ops	DRS Ops				
8									
9									
10	LTP Ops					LTP Ops	LTP Ops	Handback to MOC SCOE Reset Handover to DRS	DRS Ops
11									
12									
13									
14									
15									
16									
17		MOC Maintenance							
18		LTP Ops	LTP Ops	DRS Ops	Handback to Airbus				
19									
20									
21									
22									
23									



## DOY019: 19/01/15, (08:00 – 08:00)

07:30	–	09:00	RTB Power up	ASU	[90']
09:00	–	09:00	Handover to MOC (ACC3)	ASU -> MOC	[0']
09:00	–	09:14	Load SDM19 (OMS @ 10Hz)	con_00017_load__:V1	[14']
10:30	–	15:30	Transition from ACC3 to SCI1_2	con_dfa_a3_s12__:V1	[300']
15:30	–	15:31	Enable Sigma-Delta Loop	con_dfa_sdl_on__:V1	[1']
15:31	–	16:31	Charge measurement, TM1, inject on theta, 3V at 3mHz with +-+-	inv04014:V001	[60']
16:31	–	17:31	Charge measurement, TM2, inject on theta, 3V at 3mHz with +-+-	inv04024:V001	[60']
17:31	–	22:45	Acceleration Noise Measurement	inv00001	[314']
22:45	–	01:45	Drag-free Injections	inv01101:V001	[180']
01:45	–	04:45	Suspension Injections	inv01102:V001	[180']
04:45	–	05:00	Match Stiffness	con_dfa_c1_s12MS:V1	[15']
05:00	–	06:00	Acceleration Noise Measurement	inv00001	[60']
> 06:00	–	09:00	Suspension Injections	inv01102:V001	[180']

- charge measurement
  - assess initial charge (modelled)
  - tests modulated dc voltages on IS FEE (both)
- system identification
  - allows us to measure representative system delays
  - IS FEE (actuator) gains

<span style="color: green;">■</span>	Configuration blocks
<span style="color: blue;">■</span>	Investigations



**DOY020: 20/01/15, (08:00 – 08:00)**

09:00	–	12:20	Swap TMs	con_acc3sci12itm:V1	[200']
12:20	–	16:00	Acceleration Noise Measurement	inv00001	[220']
16:00	–	16:50	Go back to Acc 3	con_dfxx_acc3__:V1	[50']
16:50	–	16:55	Unload SDM19	unknown	[5']
16:55	–	17:00	Load SDM17 (OMS @ 1Hz)	unknown	[5']
17:00	–	17:00	Handback to MOC		[0']
17:00	–	17:00	Transition to Acc3		[0']
17:00	–	18:00	SCOE Reset		[60']
18:00	–	18:00	Handback to STOC		[0']
18:00	–	18:05	Unload SDM17	unknown	[5']
18:05	–	18:19	Load SDM19	unknown	[14']
18:19	–	21:34	Acceleration Noise Measurement in Acc3	inv00001	[195']
21:34	–	03:34	Transition back to Sci 1.2	con_acc3_sci12__:V1	[360']
03:34	–	03:35	Enable Sigma-Delta Loop	con_dfa_sdl_on__:V1	[1']
03:35	–	07:35	phi1 guidance injections	inv01168:V002	[240']
> 07:35	–	11:35	phi2 guidance injections	inv01174:V002	[240']

- Swap TMs
  - x-axis actuation on 1st test mass
  - noise of TM1 IS FEE
- Acc 3 noise run
  - wide range noise of IS FEEs
- phi guidance injections
  - delays and gains on other degrees of freedom
  - IS FEE cross-talk



- We had two teams of 4 taking shifts
- Additional personnel in STOC
  - Operations Scientist
  - Project Scientist
  - LTPDA Manager
  - LTPDA Developer
  - Additional local DA support
- External support
  - supporting teams at main LPF institutions

# SOVT photos



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# Life in the PISA room



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- Ground segment aspects worked well
  - operational environment was suitable
  - interactions worked well (various daily telecons)
  - data flow performed as expected
  - analysis tools works as expected
- Science results
  - difficult to extract much ‘science’ from the tests on the hardware
    - glitches, test environment instabilities leading to restarts
  - validation of a (compressed) section of mission timeline



- Further operations simulations
  - full-scale STOC Simulation (~10 days)
  - MOC simulations (2 or 3 on the science ops)
  - Training: tools, operations
- Data Analysis
  - Preparation of:
    - investigations
    - pipelines
    - analysis procedures
- Observation of tests
  - Space-craft closed-loop tests
  - Short functional tests of sub-systems (in particular OMS)
  - System Validation Tests (run by MOC)



- We have about 172 days to go to launch
- Commissioning starts around beginning of December (~60 days after launch)
- LTP Operations should start in mid-December (~2 weeks after commissioning)
- LTP Operations will end around end of March (after about 90 days of science operations)

**Thank you**

**MBG**

