Lecture II: Effective Field Theory and Supersymmetry



Beyond the Higgs discovery

 Higgs properties are amazingly consistent with all main compelling underlying theories (except higgsless ones!) Some parameter space of BSM theories was eventually excluded.



CPNSH workshop CERN 2006-009



Beyond the Higgs discovery

 Higgs properties are amazingly consistent with all main compelling underlying theories (except higgsless ones!) Some parameter space of BSM theories was eventually excluded.



Present Status



Beyond the Higgs discovery

 Higgs properties are amazingly consistent with all main compelling underlying theories (except higgsless ones!) Some parameter space of BSM theories was eventually excluded.



NEXT

Remarks on the fine-tuning problem

- Actually the problem cannot be strictly formulated in the strict context of the Standard Model – the Higgs mass is not calculable
- However the this problem is related to yet unknown mechanism of underlying theory where Higgs mass is calculable! In this BSM theory Higgs mass should not have tremendous fine-tuing.
- There is no hint yet about such a mechanism and this is the main source of our worries about fine-tuning



Effective Field Theory useful reviews

- J. Polchinski "Effective field theory and the Fermi surface" hep-th/9210046
- A. V. Manohar "Effective field theories" hep-ph/9606222
- I. Z. Rothstein, "TASI lectures on effective field theories" hep-ph/0308266
- D. B. Kaplan "Five lectures on effective field theory" nucl-th/0510023
- B. Gripaios "Lectures on Effective Field Theory" arXiv:1506.05039



Once we go BSM, it seems like an **infinity of possibilities** opens up - we could write down any Lagrangian we like. Fortunately, we have good starting point, since we know that **we must reproduce the SM in some limit!**



- Once we go BSM, it seems like an **infinity of possibilities** opens up we could write down any Lagrangian we like. Fortunately, we have good starting point, since we know that **we must reproduce the SM in some limit!**
- Let us suppose that any new physics is rather heavy. This is indicated experimentally by the fact that observed deviations from the SM are small (not the only possibility).

With this assumption (heavy new particles) in hand, we can analyse physics BSM using methods of EFT.



- Once we go BSM, it seems like an infinity of possibilities opens up we could write down any Lagrangian we like. Fortunately, we have good starting point, since we know that we must reproduce the SM in some limit!
- Let us suppose that any new physics is rather heavy. This is indicated experimentally by the fact that observed deviations from the SM are small (not the only possibility).

With this assumption (heavy new particles) in hand, we can analyse physics BSM using methods of EFT.

- We/start with the renormalizable SM, and consider only energies and momenta well below the weak scale ~ 100 GeV.
 - We can never produce W,Z or Higgs bosons on-shell and so we can simply do the path integral with respect to these fields (`integrate them out'). At tree-level, this just corresponds to replacing the fields using their classical equations of motion, and expanding and expanding

$$\frac{-1}{q^2 - m_W^2} = \frac{1}{m_W^2} + \frac{q^2}{m_W^4} + .$$



- Once we go BSM, it seems like an infinity of possibilities opens up we could write down any Lagrangian we like. Fortunately, we have good starting point, since we know that we must reproduce the SM in some limit!
- Let us suppose that any new physics is rather heavy. This is indicated experimentally by the fact that observed deviations from the SM are small (not the only possibility).

With this assumption (heavy new particles) in hand, we can analyse physics BSM using methods of EFT.

- We/start with the renormalizable SM, and consider only energies and momenta well below the weak scale ~ 100 GeV.
 - We can never produce W,Z or Higgs bosons on-shell and so we can simply do the path integral with respect to these fields (`integrate them out'). At tree-level, this just corresponds to replacing the fields using their classical equations of motion, and expanding and expanding

$$\frac{-1}{q^2 - m_W^2} = \frac{1}{m_W^2} + \frac{q^2}{m_W^4} + \dots$$

Expansion breaks down for momenta ~ m_wand theory is naturally equipped with a cut-off scale.



Beyond The Standard Model

Since we are only interested in low energies and momenta, we can expand in powers of the space-time derivatives (and the fields) to obtain an infinite series of **local** operators.



- Since we are only interested in low energies and momenta, we can expand in powers of the space-time derivatives (and the fields) to obtain an infinite series of **local** operators.
- we can simply cut off our loop integrals there and never have to worry about divergences. We call such a theory an EFT. The rules for making an EFT are exactly the same as those for making a QFT, except that we no longer insist on renormalizability.



- Since we are only interested in low energies and momenta, we can expand in powers of the space-time derivatives (and the fields) to obtain an infinite series of local operators.
- we can simply cut off our loop integrals there and never have to worry about divergences. We call such a theory an EFT. The rules for making an EFT are exactly the same as those for making a QFT, except that we no longer insist on renormalizability.
- Instead, we specify the fields and the symmetries, write down all the possible operators, and accept that the theory will come equipped with a cut-off Λ beyond which the expansion breaks down.



- Since we are only interested in low energies and momenta, we can expand in powers of the space-time derivatives (and the fields) to obtain an infinite series of local operators.
- we can simply cut off our loop integrals there and never have to worry about divergences. We call such a theory an EFT. The rules for making an EFT are exactly the same as those for making a QFT, except that we no longer insist on renormalizability.
- Instead, we specify the fields and the symmetries, write down all the possible operators, and accept that the theory will come equipped with a cut-off Λ beyond which the expansion breaks down.
- So, let us imagine that the SM itself is really just an effective, low-energy description of some more complete BSM theory. Thus, the fields and the (gauge) symmetries of the theory are exactly the same as in the SM, but we no longer insist on renormalizability.



SM as an Effective Field Theory

- For operators up to dimension 4, we simply recover the SM. But at dimensions higher than 4, we obtain new operators, with new physical effects.
- As a striking example of these, we expect that the accidental baryon and lepton number symmetries of the SM will be violated at some order in the expansion, and protons will decay!
- We don't know what the BSM theory need to write down all possible operators - infinitely many! Predictivity is lost?! (infinitely many measurements to fix all the coeff).
- No! Once we truncate the theory at a given order in the operator/momentum expansion - the number of coefficients is finite - can make predictions



• the natural size of coefficients is typically just an O(1) in units of Λ from dimensional analysis



- the natural size of coefficients is typically just an O(1) in units of Λ from dimensional analysis
- operators of a given dimension form a vector space, we should choose a basis to remove operators
 - a) equal up to a total divergence
 - b) operators that differ by terms vanishing for equations of motion (to be removed by a field redefinition in the path integral).



- the natural size of coefficients is typically just an O(1) in units of Λ from dimensional analysis
- operators of a given dimension form a vector space, we should choose a basis to remove operators
 - a) equal up to a total divergence
 - b) operators that differ by terms vanishing for equations of motion (to be removed by a field redefinition in the path integral).
 - **loop effects:** not obvious how to insert these operators into loops, and integrate over all loop momenta up to the cut-off Λ .
 - One can show, that expanded in powers of the external momenta they generate corrections to lower dimensional operators.
 - This suggests regularization scheme in which these corrections are already taken into account. The `right' scheme is turned to be DIM REG (A don't appear in the numerators of the loop amplitudes).



- the natural size of coefficients is typically just an O(1) in units of Λ from dimensional analysis
- operators of a given dimension form a vector space, we should choose a basis to remove operators
 - a) equal up to a total divergence
 - b) operators that differ by terms vanishing for equations of motion (to be removed by a field redefinition in the path integral).
- Ioop effects: not obvious how to insert these operators into loops, and integrate over all loop momenta up to the cut-off Λ .
 - One can show, that expanded in powers of the external momenta they generate corrections to lower dimensional operators.
 - This suggests regularization scheme in which these corrections are already taken into account. The `right' scheme is turned to be DIM REG (A don't appear in the numerators of the loop amplitudes).
- If EFT make sense, why did we ever insist on renormalizability of SM? Actually, it can now be thought of as a special case of a non-renormalizable theory, in which Λ to be very large.
 - DIM>4 operators become completely negligible (`irrelevant')
 - DIM=4 operators stay the same (`marginal')
 - DIM <4 dominate (and are called `relevant') actually has problem since $m \sim \Lambda$ (from din analysis) so theory needs dynamical mechanism or tuning



D=0: the cosmological constant

- adds an arbitrary constant, to the Lagrangian; no dependence on any fields & derivatives, can be interpreted as the energy density of the vacuum
- the vacuum energy is measurable is equivalent to including of Einstein's "cosmological constant" $\rho_{cc} \sim (10^{-3} \text{ev})^4$ into the gravitational field equation
 - good news, on one hand Universe is observed to accelerate
 - bad news, on the other hand the size of this operator coefficient Λ⁴: for Planck scale we need (10¹⁹ GeV/10⁻³eV)⁴ = (10³¹)⁴=10¹²⁴ tuning! for SUSY scale we need (10³ GeV/10⁻³eV)⁴ = (10¹⁵)⁴=10⁶⁰ tuning! many attempts no satisfactory dynamical solution has been suggested
 an alternative is to argue that we live in a multiverse in which the constant takes many different values in different corners, and we happen to live in one which is conducive to life (Weinberg, 1988)



D=2: the Higgs mass parameter

the SM is the Higgs mass parameter, the natural size is Λ , while we measure v ~ 100 GeV -> two options: a) the natural cut-off of the SM is not far above the weak scale (LHC will tell); b) the cut-off is much larger, and the weak scale is tuned (anthropics etc)

D=4: marginal operators – renormalisable SM – discussed at previous lecture



D=5: neutrino masses and mixings there is precisely one (exercise) operator dimensionless 3x3 matrix in flavour space
this operator violates the individual and total lepton numbers
it gives masses to neutrinos after EWSB, just as we observe
given the observed Δm² = 10⁻³ eV² for neutrinos, Λ ~ 10¹⁴ GeV
one could argue that while neutrino masses are evidence for physics BSM
Alternatively one can add three v^c, singlets under SU(3)x SU(2) xU(1) for each SM family replacing D=5 operator renormalizable Yukawa term λ^v lH^c v^c (Dirac mass term after EWSB) and/or

 $\mathbf{m}^{\mathbf{v}}\mathbf{v}^{\mathbf{c}}\mathbf{v}^{\mathbf{c}}$ (Majorana mass term)

(exercise: how λ^{ll} is related to λ^{ν} and \mathbf{m}^{ν} ?) neutrino mass eigenstates in this renormalizable model need not be heavy, but very weakly coupled to SM states! one can **redefine SM** to include these terms (recall yesterday remark from Dima Kazakov)



D=6: mbaryon-number violation many operators appear, including baryon and lepton number violating ones $\frac{qqq\ell}{\Lambda^2} \quad \text{and} \quad \frac{u^c u^c d^c e^c}{\Lambda^2}$

(exercise: check these are invariants) cause the proton decay $p \rightarrow e^+ \pi^0$.

 $\Lambda > 10^{15}$ GeV comes the exp bounds on the proton lifetime, $\tau^{p} > 10^{33}$ yr: new physics either respects baryon or lepton number, or is a long way away



$$\Gamma(p \to \pi^0 e^+) \propto \frac{M_p^5}{\Lambda^4}$$

Operators that give corrections to FCNC are highly suppressed in the SM e.g. $(s^c d)(d^c s)/\Lambda^2$ contributes to Kaon mixing, Λ > 10⁸ GeV



Grand Unification

- The basic idea is that the Standard model gauge group SU(3)xSU(2)xU(1) is a subgroup of a larger gauge symmetry group
 The simplest is CU(5)
- The simplest is SU(5)
 - Another example is SO(10): SU(5)xU(1) \subset SO(10) comes with RH neutrinos!



SU(5)

SU(3) has 3²-1=8 generators, they correspond to
 the 8 gluons
 The guarks are in the fundamental representation of

SU(3)

SU(5) has 5²-1=24 generators, which means that
we have 24 gauge bosons
- 8 gluons and 4 electroweak bosons
- so we get 12 new gauge bosons





Generators of SU(5)





SU(5)

The right handed down type quarks and left-handed leptons form a 5 representation of SU(5)
 The rest forms a 10 representation

$$\begin{pmatrix} d \\ d \\ d \\ e^{c} \\ \bar{\nu}_{e} \end{pmatrix} \begin{pmatrix} 0 & u^{c} & -u^{c} & -u & -d \\ u^{c} & 0 & u^{c} & -u & -d \\ u^{c} & -u^{c} & 0 & -u & -d \\ u & u & u & 0 & -e^{c} \\ d & d & d & e^{c} & 0 \end{pmatrix}$$

Simplest rep:
$$\overline{5} = (\overline{3}, 1)_{+2/3} \oplus (1, 2)_{-1} \qquad 10 = (\overline{3}, 1)_{-4/3} \oplus (3, 2)_{+1/3} \oplus (1, 1)_{+2}$$



Beyond The Standard Model

Grand Unified Theories

- In this model there are two stages of symmetry breaking
- At the GUT scale the SU(5) symmetry is broken and the X and Y bosons get masses
- At the electroweak scale the SU(2)xU(1) symmetry is broken as before
- Problems with this theory
 The couplings don't meet at the GUT scale
 Proton decay



Proton Decay

 Since in Grand Unified theories we have the X/Y bosons which couple quarks and leptons, they predict the decay of the proton





The hint about GUT scale and couplings unification

We ignore threshold corrections and assume desert! Then 1-loop RGEs for SU(N):

$$\frac{1}{g_i^2(\mu)} = \frac{1}{g_i^2(Q)} + b_i \log\left(Q^2/\mu^2\right) \quad b_N = \frac{1}{(4\pi)^2} \left[-\frac{11}{3}N + \frac{4}{3}n_g\right]$$
$$b_1 = \frac{1}{4\pi^2} \quad b_2 = -\frac{5}{24\pi^2} \quad b_3 = -\frac{7}{16\pi^2}$$



Beyond The Standard Model

The hint about GUT scale and couplings unification



There is a clear hing about couplings unification
Couplings do not unify exactly
GUT scale can be roughly estimated to be in the 10¹⁴ -10¹⁷ GeV range



Hints on Supersymmetry



From Hitoshi Murayama, 2007 Once upon a time, there was a hierarchy problem...

- At the end of 19th century: a "crisis" about electron
 - Like charges repel: hard to keep electric charge in a small pack
 - Electron is point-like
- Need a lot of energy to keep it small!

 $\Delta m_e c^2 \sim \frac{e^2}{r_e} \sim \text{GeV} \frac{10^{-17} \text{cm}}{r_e}$ Correction $\Delta m_e c^2 > m_e c^2$ for $r_e < 10^{-13} \text{cm}$ Breakdown of theory of electromagnetism \Rightarrow Can't discuss physics below 10^{-13}cm



Beyond The Standard Model

Anti-Matter Comes to Rescue by Doubling of #Particles

Electron creates a force to repel itself Vacuum bubble of matter anti-matter creation/annihilation Electron annihilates the positron in the bubble \Rightarrow only 10% of mass even for Planck-size $r_e \sim 10^{-33}$ cm



 $\Delta m_e \sim m_e \frac{\alpha}{4\pi} \log(m_e r_e)$



Higgs repels itself, too

Just like electron repelling itself H because of its charge, Higgs boson also repels itself Requires a lot of energy to contain itself in its point-like size! Breakdown of theory of weak force

Can't get started!

$$\int_{H} \frac{H}{\Delta m_{H}^{2}c^{4}} \sim \left(\frac{\hbar c}{r_{H}}\right)^{2}$$



History repeats itself?

- Oouble #particles
 again ⇒
- superpartners
 "Vacuum bubbles" of superpartners cancel the energy required to contain Higgs boson in itself
- Standard Model made consistent with whatever physics at shorter distances



 $\Delta m_H^2 \sim \frac{\alpha}{4\pi} m_{SUSY}^2 \log(m_H r_H)$



Supersymmetry



Supersymmetry (SUSY)

boson-fermion symmetry aimed to unify all forces in nature $Q|BOSON\rangle = |FERMION\rangle, \quad Q|FERMION\rangle = |BOSON\rangle$

extends Poincare algebra to Super-Poincare Algebra:

the most general set of space-time symmetries! (1971-74)

$$\{f,f\}=0, \ \ [B,B]=0, \ \ \{Q_{lpha},ar{Q}_{eta}\}=2\gamma^{\mu}_{lphaeta}P_{\mu}$$

Golfand and Likhtman'71; Ramond'71; Neveu,Schwarz'71; Volkov and Akulov'73; Wess and Zumino'74





Supersymmetry (SUSY)

boson-fermion symmetry aimed to unify all forces in nature $Q|BOSON\rangle = |FERMION\rangle, \quad Q|FERMION\rangle = |BOSON\rangle$

extends Poincare algebra to Super-Poincare Algebra:

the most general set of space-time symmetries! (1971-74)

$$\{f,f\}=0, ~~[B,B]=0, ~~\{Q_{lpha},ar{Q}_{eta}\}=2\gamma^{\mu}_{lphaeta}P_{\mu}$$

Golfand and Likhtman'71; Ramond'71; Neveu,Schwarz'71; Volkov and Akulov'73; Wess and Zumino'74





"Beyond the Standard Model"

SUSY principles

boson-fermion symmetry aimed to unify all forces in nature $Q|BOSON\rangle = |FERMION\rangle, Q|FERMION\rangle = |BOSON\rangle$

extends Poincare algebra to Super-Poincare Algebra: the most general set of space-time symmetries! (1971-74)

 $\{f,f\}=0, \ \ [B,B]=0, \ \ \{Q_{lpha},ar{Q}_{eta}\}=2\gamma^{\mu}_{lphaeta}P_{\mu}$

Golfand and Likhtman'71; Ramond'71; Neveu, Schwarz'71; Volkov and Akulov'73; Wess and Zumino'74





SUSY principles

boson-fermion symmetry aimed to unify all forces in nature $Q|BOSON\rangle = |FERMION\rangle, Q|FERMION\rangle = |BOSON\rangle$

extends Poincare algebra to Super-Poincare Algebra: the most general set of space-time symmetries! (1971-74)

 $\{f,f\}=0, \ \ [B,B]=0, \ \ \{Q_{lpha},ar{Q}_{eta}\}=2\gamma^{\mu}_{lphaeta}P_{\mu}$

Golfand and Likhtman'71; Ramond'71; Neveu, Schwarz'71; Volkov and Akulov'73; Wess and Zumino'74



R-parity guarantees Lightest SUSY particle (LSP) is stable - DM candidate!



We are still inspired by this beauty ...





We are still inspired by this beauty ... after more than 30 year unsuccessful searches ...





Beauty of SUSY

 $1/\alpha_i$

- Provides good DM candidate LSP
- CP violation can be incorporated baryogenesis via leptogenesis
- Radiative EWSB
- Solves fine-tuning problem
- Provides gauge coupling unification
- local supersymmetry requires spin 2 boson – graviton!
- allows to introduce fermions into string theories





Beauty of SUSY

 $1/\alpha_i$

- Provides good DM candidate LSP
- CP violation can be incorporated baryogenesis via leptogenesis
- Radiative EWSB
- Solves fine-tuning problem
- Provides gauge coupling unification
- local supersymmetry requires spin 2 boson – graviton!
- allows to introduce fermions into string theories



But the real beauty of SUSY is that It was not deliberately designed to solve the SM problems!



SUSY breaking and mSUGRA scenario

SUSY is not observed \Rightarrow must be broken



SUSY breaking and mSUGRA scenario

SUSY is not observed \Rightarrow must be broken



▶ B - parameter – usually expressed via $\tan \beta$

 $\blacktriangleright \Rightarrow \textbf{mSUGRA parameters: } m_0, m_{1/2}, A_0, \tan\beta, sign(\mu)$

Limits from LHC for mSUGRA scenario







Limits from LHC for mSUGRA scenario





SUSY, where are you?!

Summary of CMS SUSY Results* in SMS framework

ICHEP 2014





Evolution of neutralino relic density





Evolution of neutralino relic density





Neutralino relic density in mSUGRA

most of the parameter space is ruled out! $\Omega h^2 \gg 1$ special regions with high σ_A are required to get $0.094 < \Omega h^2 < 0.129$



Neutralino relic density in mSUGRA

most of the parameter space is ruled out! $\Omega h^2 \gg 1$ special regions with high σ_A are required to get $0.094 < \Omega h^2 < 0.129$



$$b \rightarrow s\gamma, (g-2)_{\mu}/2, B_{S} \rightarrow \mu^{+}\mu^{-} \text{ constraints}$$

$$b \rightarrow s\gamma; BF(b \rightarrow s\gamma) = (3.55 \pm 0.26) \times 10^{-4} (BELLE.CLEO \text{ and } ALEPH]$$
Theory:
$$(3.15 \pm 0.23) \times 10^{-4} \text{ Misiak, Steinhauser '06}$$

$$2.85 \times 10^{-4} \leq Br(b \rightarrow s\gamma) \leq 4.24 \times 10^{-4} (95\% \text{ CL incl } 10\% \text{ theory}) \quad \psi^{-}_{W} \quad \varphi^{-}_{\overline{q}} = b$$

$$2.85 \times 10^{-4} \leq Br(b \rightarrow s\gamma) \leq 4.24 \times 10^{-4} (95\% \text{ CL incl } 10\% \text{ theory}) \quad \psi^{-}_{W} \quad BR(b \rightarrow s\gamma)|_{\chi^{\pm}} \propto \mu A_{t} \tan \beta$$

$$(g-2)_{\mu}/2 \text{ results} \quad (g-2)_{\mu}/2 \text{ results} \quad (f \text{ decay } data \Delta a_{\mu} = (12.4 \pm 8.3) \times 10^{-10} \text{ (Davier et al.)} \quad \Phi^{-}_{a} \text{ dased on e'e' data} \quad e'e' \text{ data are more to be trusted since they offer a direct determination of the hadronic vacuum polarization $\sim 3\sigma \implies \text{ second generation of slepton are relatively light!}$

$$BF(B_{s} \rightarrow \mu^{+}\mu^{-}) < 1.0 \times 10^{-7} (CDF), (SM: 3.4 \times 10^{-9}) \quad amplitude for H-mediated decay grows as targ^{3} (l) \Rightarrow relevant to high \tan \beta \text{ scenario} [Babu, Kolda; Dedes, Dreiner, Nierste; Arnowitt, Dutta, Tanaka; Mizukoshi, Tata, Wang]$$$$

Alexander Belyaev



Pre LHC mSUGRA $\chi^2 = \chi^2_{\delta a_{\mu}} + \chi^2_{\Omega h^2} + \chi^2_{b \rightarrow s \gamma}$

 Δa_{μ} favors light second generation sleptons, while $BF(b \rightarrow s\gamma)$ prefers heavy third generation: hard to realize in mSUGRA model.





analysis

Implications of LHC search for SUSY fits

Buchmueller, Cavanaugh, De Roeck, Dolan, Ellis, Flaecher, Heinemeyer, Isidori, Marrouche, Martinez, Santos, Olive, Rogerson, Ronga, de Vries, Weiglein,

Global frequentist fits to the CMSSM using the MasterCode framework





The EW measure of Fine Tuning

 $\mathcal{L}_{\text{MSSM}} = \mu \tilde{H}_{u} \tilde{H}_{d} + \text{h.c.} + (m_{H_{u}}^{2} + |\mu|^{2}) |H_{u}|^{2} + (m_{H_{d}}^{2} + |\mu|^{2}) |H_{d}|^{2} + \dots$

The EW measure requires that there be no large/unnatural cancellations in deriving m_{τ} from the weak scale scalar potential:

$$\frac{m_Z^2}{2} = \frac{(m_{H_d}^2 + \Sigma_d^d) - (m_{H_u}^2 + \Sigma_u^u) \tan^2 \beta}{(\tan^2 \beta - 1)} - \mu^2 \simeq -m_{H_u}^2 - \mu^2$$

using fine-tuning definition which became standard Ellis, Enqvist, Nanopoulos, Zwirner '86; Barbieri, Giudice '88

$$\Delta_{FT} = max[c_i], \quad c_i = \left|\frac{\partial \ln m_Z^2}{\partial \ln p_i}\right| = \left|\frac{p_i}{m_Z^2}\frac{\partial m_Z^2}{\partial p_i}\right|$$

one finds $\Delta_{FT} \simeq \Delta_{EW}$ which requires $\begin{vmatrix} \mu^2 \\ \mu^2 \end{vmatrix} \simeq M_Z^2 \\ m_{H_u}^2 \end{vmatrix} \simeq M_Z^2$



The last one is GUT model-dependent, so we consider the value $|\mu^2|$ as a measure of the minimal fine-tuning



"Compressed Higgsino" Scenario (CHS)

chargino-neutralino mass matrices



 M_2 real, $M_1 = |M_1|e^{-\Phi_1}$, $\mu = |\mu|e^{i\Phi_{\mu}}$

- Case of $\mu \leftrightarrow M1$, M2: $\chi^0_{1,2}$ and χ^{\pm} become quasi-degenerate and acquire large higgsino component. This provides a naturally low DM relic density via gaugino annihilation and co-annihilation processes into SM V's and H
- This is the case of relatively light higgsinos-electroweakinos compared to the other SUSY particles.
- This scenario is not just motivated by its simplicity, but also by the lack of evidence for SUSY to date, indicating that a weak scale SUSY spectrum is likely non-universal



- The most challenging case takes place when only $\chi^0_{1,2}$ and χ^{\pm} are accessible at the LHC, and the mass gap between them is not enough for any leptonic signature as happen in FFP scenario.
- The only way to probe FFP is a mono-jet signature [Where the Sidewalk Ends? ... Alves, Izaguirre,Wacker '11], which has been used in studies on compressed SUSY spectra, e.g. Dreiner,Kramer,Tattersall '12; Han,Kobakhidze,Liu,Saavedra,Wu'13; Han,Kribs,Martin,Menon '14



NE

- The most challenging case takes place when only $\chi^0_{1,2}$ and χ^{\pm} are accessible at the LHC, and the mass gap between them is not enough for any leptonic signature as happen in FFP scenario.
- The only way to probe FFP is a mono-jet signature [Where the Sidewalk Ends? ... Alves, Izaguirre,Wacker '11], which has been used in studies on compressed SUSY spectra, e.g. Dreiner,Kramer,Tattersall '12; Han,Kobakhidze,Liu,Saavedra,Wu'13; Han,Kribs,Martin,Menon '14



NE

- The most challenging case takes place when only $\chi^0_{1,2}$ and χ^{\pm} are accessible at the LHC, and the mass gap between them is not enough for any leptonic signature as happen in FFP scenario.
- The only way to probe FFP is a mono-jet signature [Where the Sidewalk Ends? ... Alves, Izaguirre,Wacker '11], which has been used in studies on compressed SUSY spectra, e.g. Dreiner,Kramer,Tattersall '12; Han,Kobakhidze,Liu,Saavedra,Wu'13; Han,Kribs,Martin,Menon '14





- The most challenging case takes place when only $\chi^0_{1,2}$ and χ^{\pm} are accessible at the LHC, and the mass gap between them is not enough for any leptonic signature as happen in FFP scenario.
- The only way to probe FFP is a mono-jet signature [Where the Sidewalk Ends? ... Alves, Izaguirre,Wacker '11], which has been used in studies on compressed SUSY spectra, e.g. Dreiner,Kramer,Tattersall '12; Han,Kobakhidze,Liu,Saavedra,Wu'13; Han,Kribs,Martin,Menon '14



Alexander Belyaev

NE

S/B VS

Signal significance



	$Z(\nu\bar{\nu})j$	$W(\ell\nu)j$	$\mu = 93 \text{ GeV}$	$\mu = 500 \text{ GeV}$
$p_{jet}^T > 50 \text{ GeV}, \eta_{jet} < 5$	6.4 E+7	2.9 E+8	2.6 E+5	948
Veto $p_{e^{\pm},\mu^{\pm}/\tau^{\pm}}^{T} > 10/20 \text{ GeV}$	6.2 E+7	1.2 E+8	2.5 E+5	921
$p_j^T > 500 \text{ GeV}$	2.5 E+4	2.0 E+4	1051	32
$p_j^T = E_T > 500 \text{ GeV}$	1.5 E+4	4.1 E+3	747	27
$p_j^T = E_T > 1000 \text{ GeV}$	315 (375)	65 (32)	21 (31)	2 (2)
$p_j^T = E_T > 1500 \text{ GeV}$	18 (20)	2(1)	1 (2)	0 (0)
$p_j^T = \not\!\! E_T > 2000 \text{ GeV}$	1 (1)	0 (0)	0(1)	0 (0)

- tension between S/B and signal significance
- S/B pushes E^{miss} cut up towards an acceptable systematic
- significance requires comparatively low (below 500 GeV) E^{miss} cut



LHC/DM direct detection sensitivity to CHS



"Uncovering Natural Supersymmetry via the interplay between the LHC and Direct Dark Matter Detection", Barducci, AB, Bharucha, Porod, Sanz, arXiv:1504.02472 (JHEP)

 SUSY, at least DM, can be around the corner (100 GeV), it is just very hard to detect it!

Alexander Belyaev

NEXT

"Beyond the Standard Model"

Your question: "Can experimentally rule out SUSY in general and e.g. cMSSM in particular?"



Your question: "Can experimentally rule out SUSY in general and e.g. cMSSM in particular?" The Answer is: NO! SUSY can be either discovered or abandoned!

