

# The muon $(g - 2)$ in MSSM: precision corrections and novel phenomenological scenario

Hyejung Stöckinger-Kim

TU Dresden

with

M. Bach, H. Fagnoli, C. Gneidiger, J. Park, S. Paßehr, D. Stöckinger

HEPKIT Karlsruhe, 6. Oct. 2015

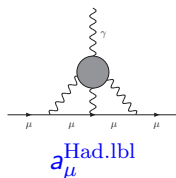
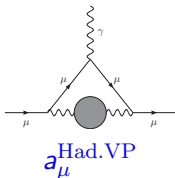
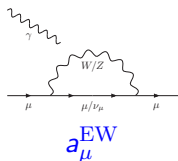
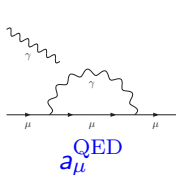
Current status of  $a_\mu^{\text{SM}}$  and  $a_\mu^{\text{Exp}}$

$(g-2)_\mu$  in MSSM

MSSM loop corrections

MSSM:  $\tan \beta \rightarrow \infty$

Summary

$a_\mu^{\text{SM}}$ 

- QED 5-loop calculation completed. [Kinoshita et al '12]
- Convergence of hadronic contributions.
- New BESIII results [1507:08188]
- Precision improvement in  $a_\mu^{\text{EW}}$ : uncertainty caused by  $M_H$  eliminated  $\Rightarrow a_\mu^{\text{EW}} = (154 \pm 1 \pm 2) \times 10^{-11}$  [Czarnecki, Krause, Marciano]  
 $a_\mu^{\text{EW}} = (153.6 \pm 1.0) \times 10^{-11}$  [Gnendiger, Stöckinger, S-K '13]  
 with  $M_H = 125.6 \pm 1.5 \text{ GeV}$

$$a_{\mu}^{\text{Exp}}$$

- The latest result from BNL

$$a_{\mu}^{\text{E821}} = (11659208.9 \pm 6.3) \times 10^{-10}$$

- New experiment at Fermilab E989:

$$0.54 \rightarrow 0.14 \text{ ppm.}$$

Current uncertainty in  $a_{\mu}^{\text{SM}}$ : 0.42 ppm

[Bennett et al. '06]



Current status of  $a_{\mu}^{\text{Exp}} - a_{\mu}^{\text{SM}}$ 

$$\Delta a_{\mu}(\text{E821} - \text{SM}) = \begin{cases} (28.7 \pm 8.0) \times 10^{-10} & [\text{Davier et al.}] \\ (26.1 \pm 8.0) \times 10^{-10} & [\text{Hagiwara et al.}] \end{cases}$$

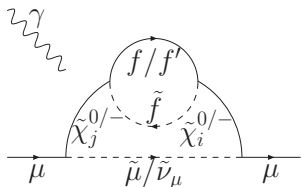
$\sim 3...4 \sigma$

- This deviation motivates New Physics: SUSY can explain it.
- $a_{\mu}$  is an important constraint on SUSY
- Still, large contributions possible, e.g. if sleptons  $\ll$  squarks (non-traditional models)

[Endo, Hamaguchi, Iwamoto, Yanagida, D.P. Roy, et al]

$a_\mu$  in MSSM

## 1. MSSM loop corrections:

 $f\tilde{f}$  two-loop

## 2. Radiative muon mass generation:

$$\tan \beta \equiv \frac{v_u}{v_d} \Big|_{v_d=0} \rightarrow \infty$$

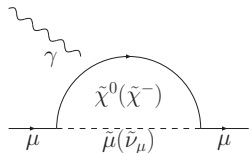
## MSSM loop corrections

# MSSM loop corrections

## 1. MSSM one-loop corrections

- Parameter dependence:  
 $\mu, M_1, M_2, M_L, M_E, \tan \beta$
- For the equal mass case:

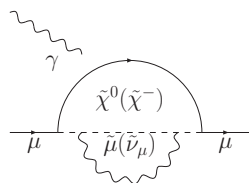
$$a_{\mu}^{\text{SUSY,1L}} \approx 13 \times 10^{-10} \text{sgn}(\mu) \tan \beta \left( \frac{100 \text{ GeV}}{M_{\text{SUSY}}} \right)^2$$



## 2. MSSM photonic two-loop corrections

[v.Weitershausen, Schäfer, Stöckinger, S-K '10]

- $\sim \frac{\alpha}{4\pi} \log \frac{m_{\mu}}{m_{\tilde{\mu}}(m_{\tilde{\nu}_{\mu}})} \times a_{\mu}^{\text{SUSY,1L}} \chi^0(\chi^-)$
- $-(7 \dots 9) \%$  for SUSY masses between 100 - 1000 GeV



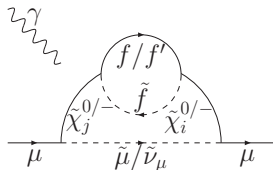


# MSSM loop corrections

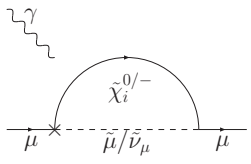
## 3. MSSM $f\tilde{f}$ two-loop corrections

[Fagnoli, Gnendiger, Paßehr, Stöckinger, S-K '13]

- Maximum complexity:  
5 heavy scales + 2 light scales
- Additional parameter dependence:  
 $M_{Q_i}, M_{U_i}, M_{D_i}, M_{L_i}, M_{E_i}, i \in \{1, 2, 3\}$
- solve the one-loop ambiguity caused by  $\alpha \Rightarrow$
- Non-decoupling behaviour  $\Rightarrow$



$\alpha$  ambiguity solved



$$\begin{aligned}
 &= a_\mu^{1\text{L}} \times \left( \frac{\delta(e^2/s_W^2)}{e^2/s_W^2} + \dots \right) \\
 &= a_\mu^{1\text{L}} \times \left( \Delta\alpha_{f,\tilde{f}} - \frac{c_W^2}{s_W^2} \Delta\rho_{f,\tilde{f}} + \dots \right)
 \end{aligned}$$

1-loop ambiguity

fixed by full 2-L  $f\tilde{f}$ 

$$a_\mu^{1\text{L}}|\alpha(0) = 29.4 \times 10^{-10}$$

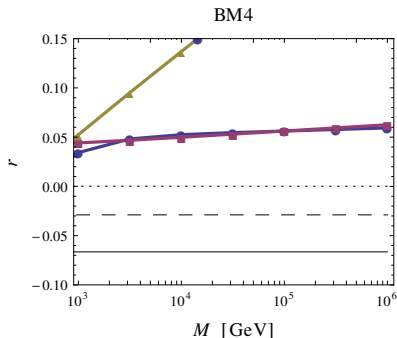
$$a_\mu^{1\text{L}}|\alpha(M_Z) = 31.6 \times 10^{-10}$$

$$a_\mu^{1\text{L}}|\alpha(G_F) = 30.5 \times 10^{-10}$$

$$\Rightarrow a_\mu^{1\text{L}+2\text{L}f\tilde{f}} = 32.2 \times 10^{-10}$$

Difference from  $\Delta\alpha$  and  $\Delta\rho$ : 2-loop  $f\tilde{f}$  terms.

# $f\tilde{f}$ -loop corrections: Non-decoupling behaviour



- $M_2, m_{\tilde{\mu}_L} \gg M_1, m_{\tilde{\mu}_R}$
- $M_1 = 140\text{GeV}$
- $m_{\tilde{\mu}_R} = 200\text{GeV}$
- $M_2 = m_{\tilde{\mu}_L} = 2000\text{GeV}$
- $\mu = -160, \tan \beta = 40$
- $\mathcal{O}(10\dots30\%)$

	$M_{U3, D3, Q3, E3, L3}$
	$M_{U, D, Q}$
	$M_{Q3}; M_{U3} = 1 \text{ TeV}$
	$(\tan \beta)^2$
	photonic
	$2L(a)$

# Leading logarithmic approximation

[<http://iktp.tu-dresden.de/index.php?id=theory-software>]

$$a_\mu^{\text{1LSUSY, M.I.}} = a_\mu^{\text{1L}}(\tilde{W}-\tilde{H}, \tilde{\nu}_\mu) + a_\mu^{\text{1L}}(\tilde{W}-\tilde{H}, \tilde{\mu}_L) + a_\mu^{\text{1L}}(\tilde{B}-\tilde{H}, \tilde{\mu}_L) \\ + a_\mu^{\text{1L}}(\tilde{B}-\tilde{H}, \tilde{\mu}_R) + a_\mu^{\text{1L}}(\tilde{B}, \tilde{\mu}_L-\tilde{\mu}_R)$$

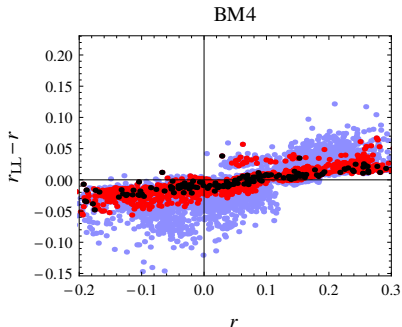
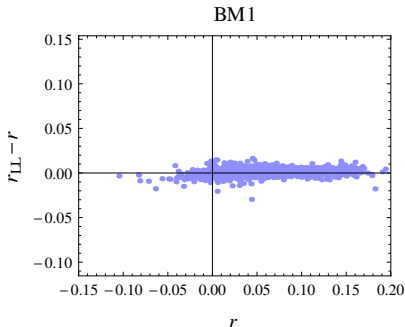
$$a_\mu^{\text{1L}}(\tilde{W}/\tilde{B}, \dots) \propto g_{2/1}^2 \mu \tan \beta$$

$$\begin{aligned} \tilde{a}_\mu^{\text{2L, f\tilde{f}LL}} &= a_\mu^{\text{1L}}(\tilde{W}-\tilde{H}, \tilde{\nu}_\mu) \left( \Delta_{g_2} + \Delta_{\tilde{H}} + \Delta_{\tilde{W}\tilde{H}} + \Delta_{t\beta} + 0.015 \right), \\ &+ a_\mu^{\text{1L}}(\tilde{W}-\tilde{H}, \tilde{\mu}_L) \left( \Delta_{g_2} + \Delta_{\tilde{H}} + \Delta_{\tilde{W}\tilde{H}} + \Delta_{t\beta} + 0.015 \right), \\ &+ a_\mu^{\text{1L}}(\tilde{B}-\tilde{H}, \tilde{\mu}_L) \left( \Delta_{g_1} + \Delta_{\tilde{H}} + \Delta_{\tilde{B}\tilde{H}} + \Delta_{t\beta} + 0.015 \right), \\ &+ a_\mu^{\text{1L}}(\tilde{B}-\tilde{H}, \tilde{\mu}_R) \left( \Delta_{g_1} + \Delta_{\tilde{H}} + \Delta_{\tilde{B}\tilde{H}} + \Delta_{t\beta} + 0.04 \right), \\ &+ a_\mu^{\text{1L}}(\tilde{B}, \tilde{\mu}_L-\tilde{\mu}_R) \left( \Delta_{g_1} + \Delta_{t\beta} + 0.03 \right) \end{aligned}$$

$$\begin{aligned} \Delta_{\tilde{H}} &= \frac{1}{16\pi^2} \frac{1}{2} \left( 3y_t^2 \log \frac{M_{U_3}}{m_{\text{SUSY}}} + 3y_b^2 \log \frac{M_{D_3}}{m_{\text{SUSY}}} \right. \\ &\left. + 3(y_t^2 + y_b^2) \log \frac{M_{Q_3}}{m_{\text{SUSY}}} + y_\tau^2 \log \frac{M_{E_3}}{m_{\text{SUSY}}} + y_\tau^2 \log \frac{M_{L_3}}{m_{\text{SUSY}}} \right) \end{aligned}$$

# Leading logarithmic approximation

[PLB 726 (2013)][arXiv:1309.0980][Fagnoli, Gwendiger, Paßehr, Stöckinger, S-K '13]



[GeV]	BM1	BM4
$\mu$	[100,200]	[-200,-100]
$M_1$	[100,200]	[100,200]
$M_2$	[200,400]	[1000,3000]
$M_E$	[200,500]	[100,300]
$M_L$	[200,500]	[1000,3000]

$$r \equiv a_\mu^{2L, \tilde{f}\tilde{f}} / a_\mu^{1L}, \quad r_{LL} \equiv a_\mu^{2L, \tilde{f}\tilde{f}, LL} / a_\mu^{1L},$$

red/black  $\min(M_{SUSY}) > 150/180\text{GeV}$

invalid when  $\tan\beta$  or  $M_{SUSY}$  is very small,

$M_{SUSY} > 200\text{GeV}$

MSSM: Radiative muon mass generation

$$\tan \beta \rightarrow \infty$$

Novel phenomenological approach

Radiative muon mass generation:  $\tan \beta \rightarrow \infty$ 

$$m_\mu^{\text{tree}} = y_\mu v_d \Rightarrow 0$$

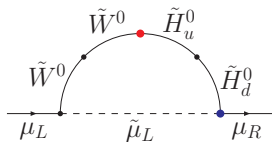
- $v_d \rightarrow 0$ ,  $\tan \beta \equiv \frac{v_u}{v_d} \rightarrow \infty$ ,  
 $m_\mu$  generated via coupling to  $v_u$

[Dobrescu, Fox '10][Altmannshofer, Straub '10]

- $m_\mu \equiv \cancel{y}_\mu v_d + y_\mu v_u \Delta_\mu^{\text{red}}$ :  
 $y_\mu$  obtained from one-loop self energy.

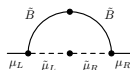
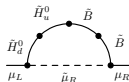
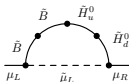
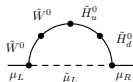
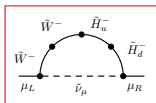
- $a_\mu^{\text{SUSY}} = \frac{y_\mu v_u}{m_\mu} a_\mu^{\text{red}}$

- $a_\mu^{\text{SUSY}} \propto y_\mu$  and  $m_\mu \propto y_\mu$



$$\Rightarrow a_\mu^{\text{SUSY}} = \frac{a_\mu^{\text{red}}}{\Delta_\mu^{\text{red}}}$$

[1504.05500][Bach, Park, Stöckinger, S-K]

The limit of  $\tan \beta \rightarrow \infty$ 

$$\begin{aligned}
 a_\mu^{\text{red}} &= a_\mu^{\text{red}}(\tilde{W}\tilde{H}\tilde{\nu}) + a_\mu^{\text{red}}(\tilde{W}\tilde{H}\tilde{\mu}_L) + a_\mu^{\text{red}}(\tilde{B}\tilde{H}\tilde{\mu}_L) + a_\mu^{\text{red}}(\tilde{B}\tilde{H}\tilde{\mu}_R) + a_\mu^{\text{red}}(\tilde{B}\tilde{\mu}_L\tilde{\mu}_R) \\
 \Delta_\mu^{\text{red}} &= \Delta_\mu^{\text{red}}(\tilde{W}\tilde{H}\tilde{\nu}) + \Delta_\mu^{\text{red}}(\tilde{W}\tilde{H}\tilde{\mu}_L) + \Delta_\mu^{\text{red}}(\tilde{B}\tilde{H}\tilde{\mu}_L) + \Delta_\mu^{\text{red}}(\tilde{B}\tilde{H}\tilde{\mu}_R) + \Delta_\mu^{\text{red}}(\tilde{B}\tilde{\mu}_L\tilde{\mu}_R)
 \end{aligned}$$

- $a_\mu^{\text{SUSY}}$  sign depends on the mass ratios.
- $\text{sgn}(\mu)$  and  $\tan \beta$  dependence disappears.
- $a_\mu^{\text{red}}(\tilde{W}\tilde{H}\tilde{\nu})$  and  $\Delta_\mu^{\text{red}}(\tilde{W}\tilde{H}\tilde{\nu})$  have opposite signs.
- For the equal mass case,  $a_\mu^{\text{red}}(\tilde{W}\tilde{H}\tilde{\nu})$  and  $\Delta_\mu^{\text{red}}(\tilde{W}\tilde{H}\tilde{\nu})$  dominate  
 $\implies$  negative  $a_\mu^{\text{SUSY}}$

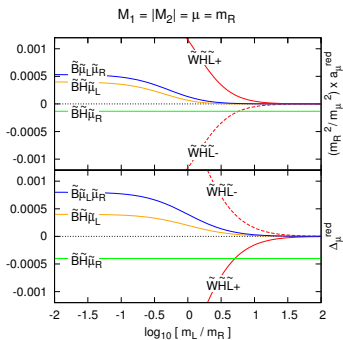
$$a_\mu^{\text{SUSY}} = \frac{a_\mu^{\text{red}}}{\Delta_\mu^{\text{red}}}$$

$$a_\mu^{\text{SUSY}}$$

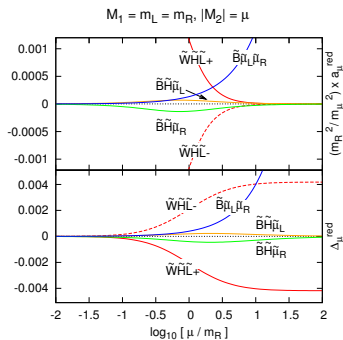
$$\approx -72 \times 10^{-10} \left( \frac{1\text{TeV}}{M_{\text{SUSY}}} \right)^2$$



However!  $a_\mu^{\text{SUSY}}$  can be positive,  
if any of  $\tilde{B}\tilde{H}\tilde{\mu}_L$ ,  $\tilde{B}\tilde{H}\tilde{\mu}_R$ , and  $\tilde{B}\tilde{\mu}_L\tilde{\mu}_R$  dominates.



“ $\mu_R$ -dominance”



“large  $\mu$ -limit”

The sign of  $a_\mu^{\text{SUSY}}$ 

At the equal mass limit,

$$a_\mu^{\text{SUSY}} \approx -72 \times 10^{-10} \left( \frac{1 \text{ TeV}}{M_{\text{SUSY}}} \right)^2$$

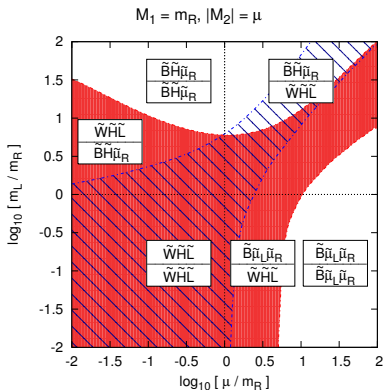
Generally,

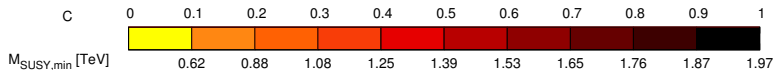
$$a_\mu^{\text{NP}} = C_{\text{NP}} \frac{m_\mu^2}{M_{\text{NP}}^2},$$

$$C_{\text{NP}} = \mathcal{O}(\delta m_\mu^{\text{NP}} / m_\mu)$$

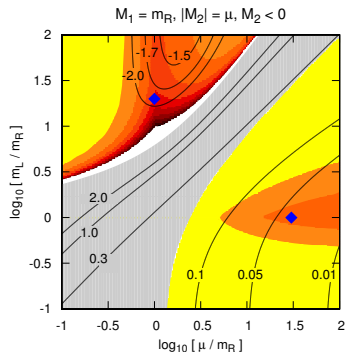
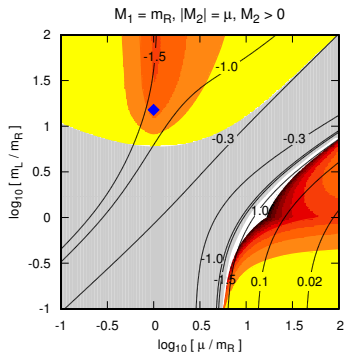
(model dependent)

What can be the  
 $C$ -value/ $a_\mu^{\text{SUSY}}$  for the given  
parameter ratio space?

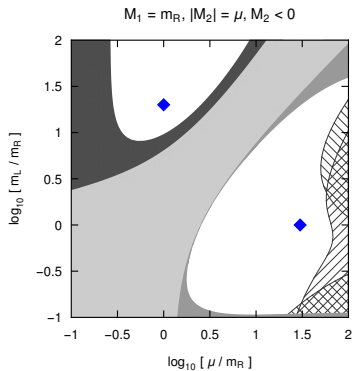
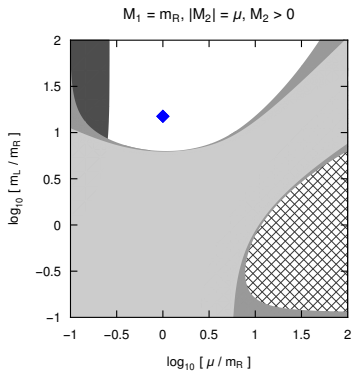


$\tan \beta \rightarrow \infty$ : C-value and  $y_{\mu}$ 

$M_{\text{SUSY,min}}$  evaluated using  $a_{\mu}^{\text{SUSY}} = 28.7 \times 10^{-10}$ .



## Further constraints



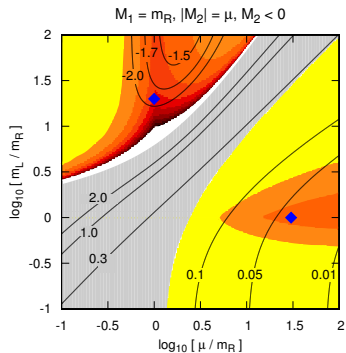
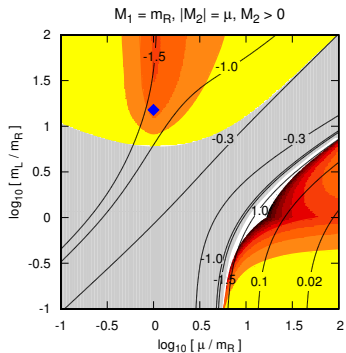
vacuum unstable (Endo et al.)

vacuum unstable (one-loop)

 $|y_{\mu}| > \sqrt{4\pi}$ 

charged particle &lt; 100 GeV

 $\hat{a}_{\mu}^{\text{SUSY}} < 0$

$\tan \beta \rightarrow \infty$ : Benchmark points

[TeV]	$\mu$	$M_1$	$M_2$	$m_L$	$m_R$	$a_{\mu}/10^{-10}$
1	1	1	1	15	1	26.4
1.3	1.3	-1.3	26	1.3	1	29.0
30	1	-30	1	1	1	28.0

$a_\mu^{\text{SUSY}}$  simple approximation

MSSM one-loop approximation with finite  $\tan\beta$ :

$$a_\mu^{\text{SUSY,1L}} \approx 13 \times 10^{-10} \text{sgn}(\mu) \tan\beta \left( \frac{100 \text{ GeV}}{M_{\text{SUSY}}} \right)^2$$

MSSM in the limit of  $\tan\beta \rightarrow \infty$ :

1. The equal mass case,

$$a_\mu^{\text{SUSY}} \approx -72 \times 10^{-10} \left( \frac{1 \text{ TeV}}{M_{\text{SUSY}}} \right)^2$$

2.  $|\mu| \gg |M_1| = m_L = m_R \equiv M_{\text{SUSY}}$ ,  $m_L \gg |\mu| = |M_1| = m_R \equiv M_{\text{SUSY}}$

$$a_\mu^{\text{SUSY}} \approx 37 \times 10^{-10} \left( \frac{1 \text{ TeV}}{M_{\text{SUSY}}} \right)^2$$

# GM2Calc

all known 2-loop results taken into account

[\[https://gm2calc.hepforge.org/\]](https://gm2calc.hepforge.org/)

- [Home](#)
- [Downloads](#)
- [Documentation](#)
- [Online](#)
- [ChangeLog](#)
- [Bugs](#)
- [Development](#)
- [Contact](#)

## GM2Calc

This is the home page of the GM2Calc project. GM2Calc is a C++ program which calculates the anomalous magnetic moment of the muon in the MSSM at 1-loop and leading 2-loop order.

### Current version

The current release is **0.2.16** (released 05 October 2015).  
Download: [GM2Calc-0.2.16.tar.gz](#)

### Authors

- Peter Athron
- Markus Bach
- Helvecio Fargnoli
- Christoph Gnendiger
- Robert Greifenhagen
- Jae-hyeon Park
- Sebastian Pässehr
- Dominik Stöckinger
- Hyejung Stöckinger-Kim
- [Alexander Voigt](#)

### References

## Summary

- $3\sigma$  deviation in  $a_\mu$  motivates New Physics.
- MSSM  $f\tilde{f}$  two-loop corrections solve the  $\alpha$  ambiguity in one-loop result and show non-decoupling behaviour  $\mathcal{O}(10 \dots 30\%)$ .
- Radiative muon mass generation in the limit of  $\tan\beta \rightarrow \infty$ 
  - Relevant parameter regions different from the usual MSSM
  - TeV-scale SUSY masses can explain  $a_\mu$ .
  - “large  $\mu$ -limit” and “ $\mu_R$ -dominance” regions:

$$a_\mu^{\text{SUSY}} \approx 37 \times 10^{-10} \left( \frac{1 \text{ TeV}}{M_{\text{SUSY}}} \right)^2$$

- <https://gm2calc.hepforge.org/>

$a_\mu^{\text{THDM},2\text{L}}$  coming soon!