

QCD, Strings and AdS/CFT, an Introduction

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LOW-x WORKSHOP,

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- Gauge/Gravity Correspondence
Gauge Fields/String: “Duality” and “Holography”
- AdS/CFT Correspondence for S^4 QCD
Extensions beyond CFT
- A case study: AdS/CFT and QGP Hydrodynamics
Emergence of a 5d Black Hole Geometry
- AdS/CFT and QGP Dynamics
Quasi-Perfect Fluidity, Viscosity, Thermalization,...
- Conclusions and Prospects
Progress and Open Problems of the QCD/String Connection

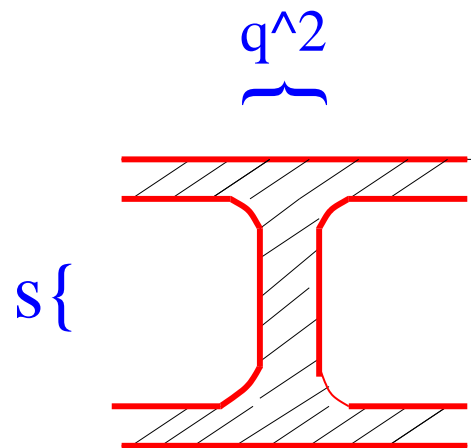
Strong Interactions and Strings

Strong Interactions History

Strings \nleftrightarrow QCD \leftrightarrow Strings

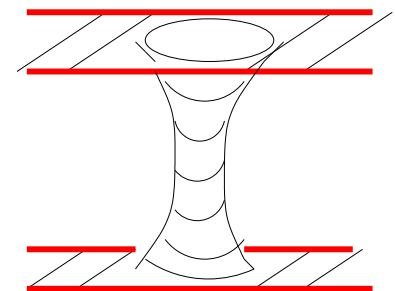
1968 \Rightarrow 1974 \Rightarrow 1998 \Rightarrow 2007 ...

1968



Veneziano Amplitude

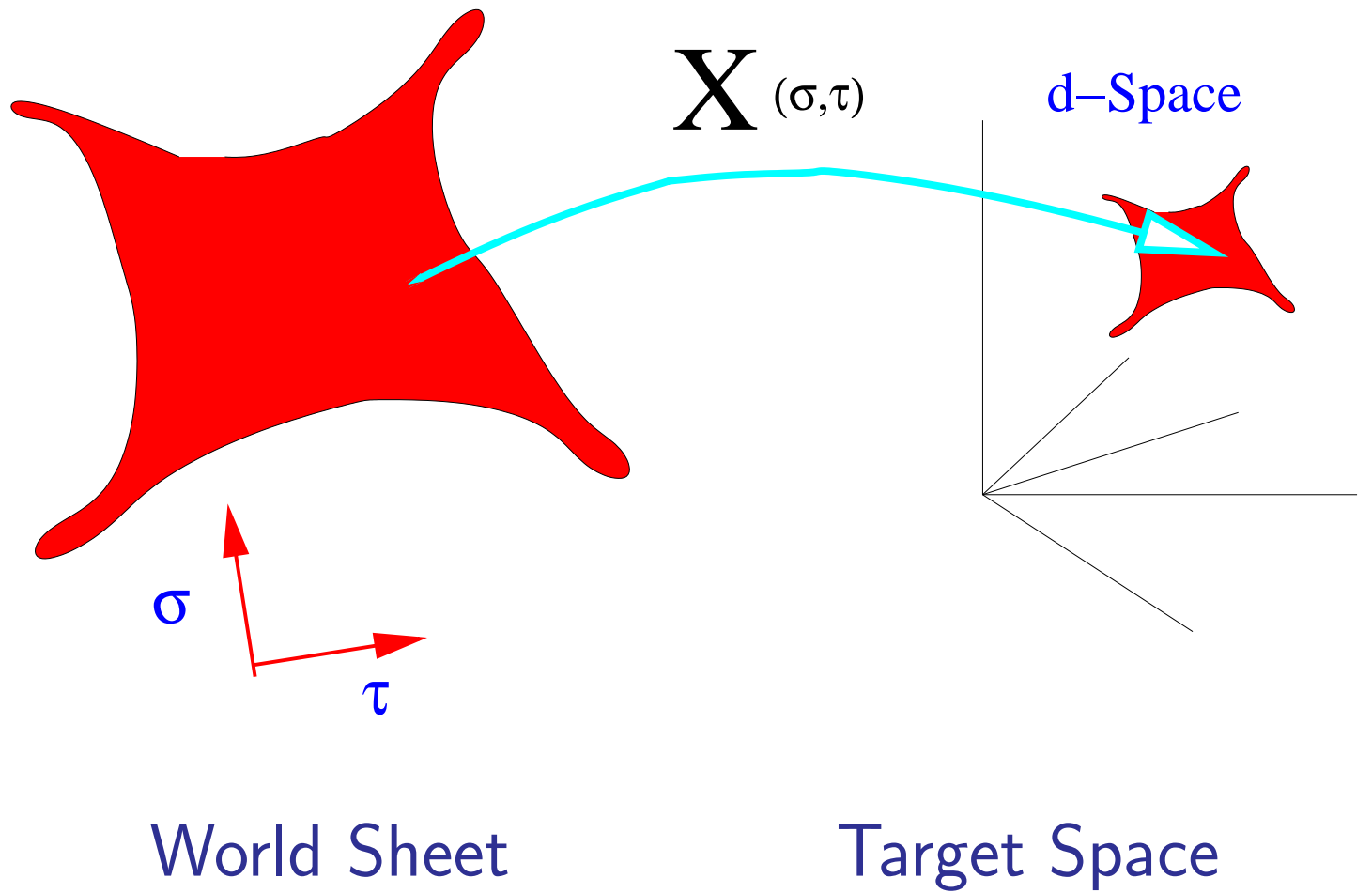
$$A_R(s, q^2)$$



Shapiro-Virasoro Amplitude

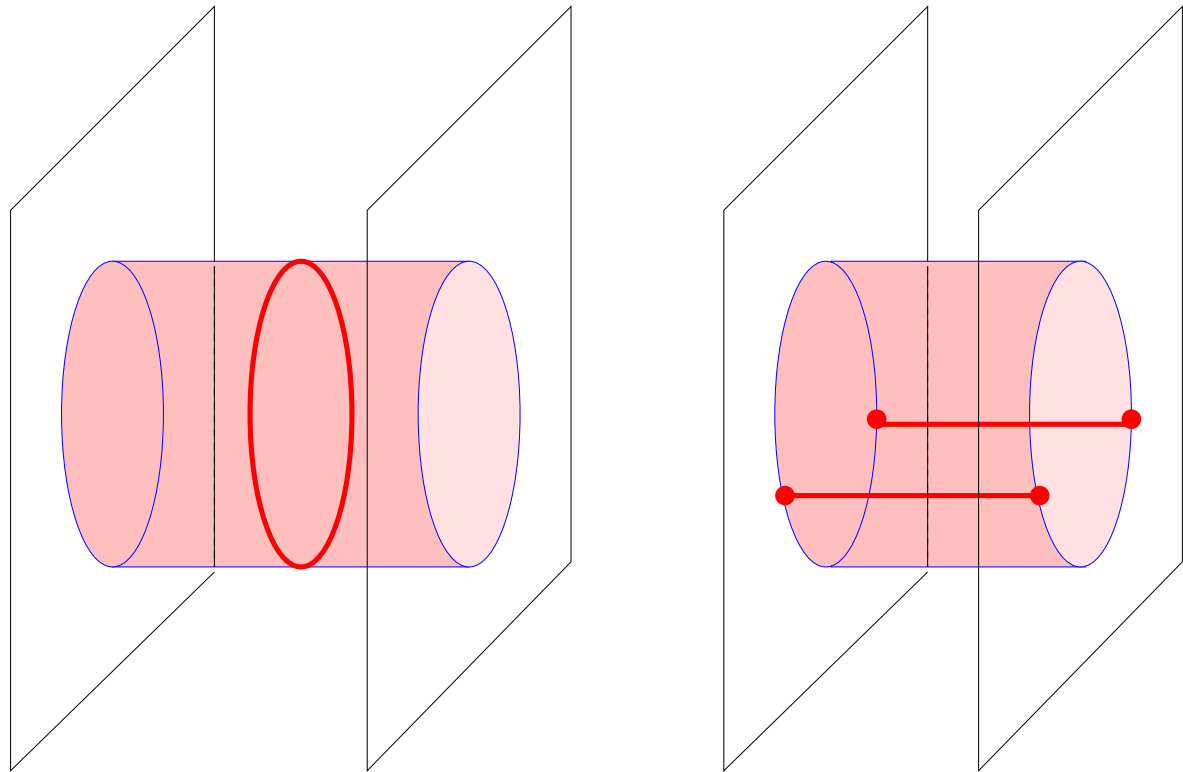
$$A_P(s, q^2)$$

String Apparatus



The Gauge-Gravity Correspondence

“Duality”: Open String \Leftrightarrow Closed String



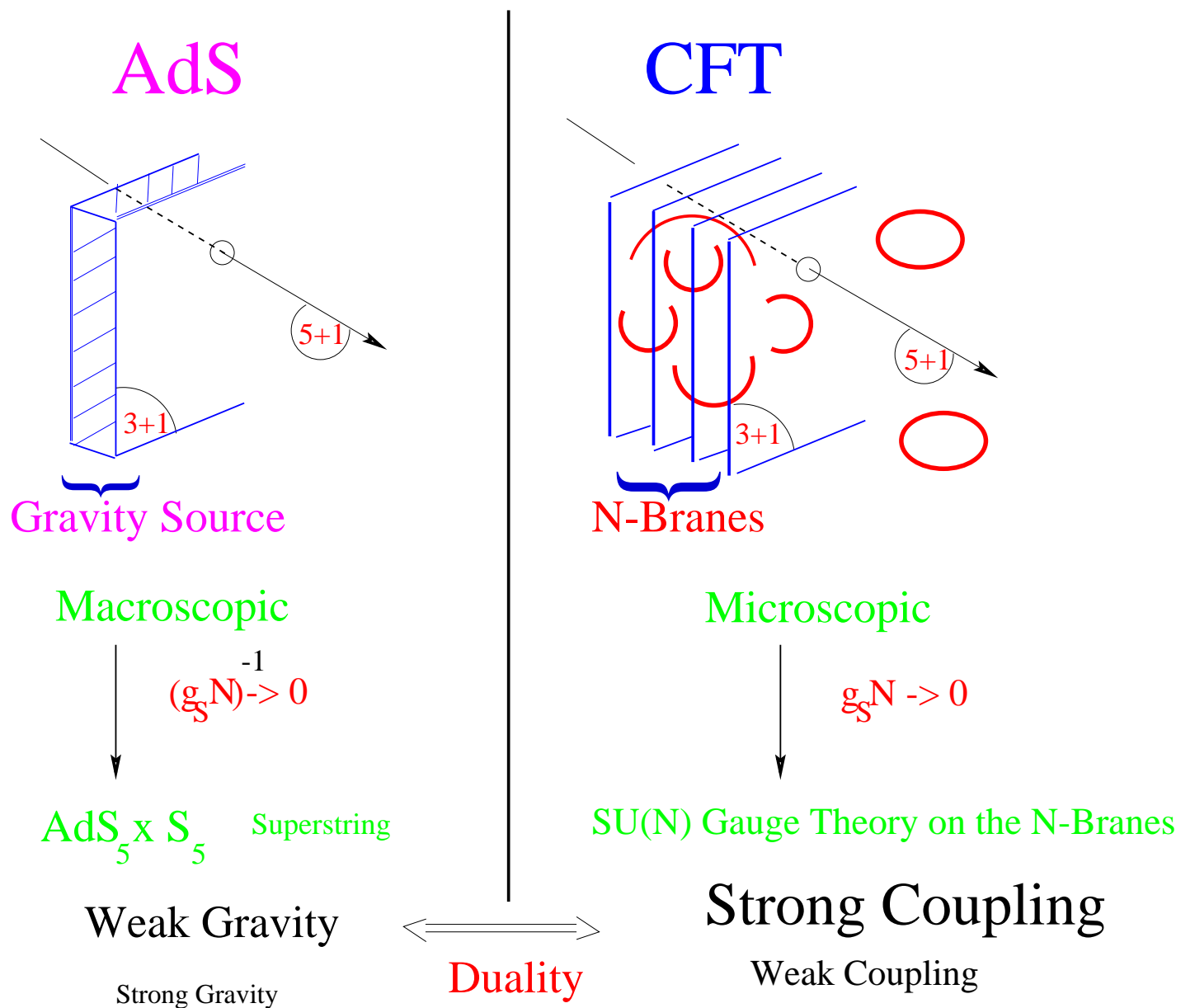
Schomerus, 2006

Closed String \Leftrightarrow *1 – loop Open String*
D – Brane “Universe” \Rightarrow *Open String Ending*
Gravity \Leftrightarrow *Gauge*
Large/Small Distance \Rightarrow *Gravity/Gauge Correspondence*

AdS/CFT Correspondence

J. Maldacena

1998



DUALITY

- D_3 -brane Solution of Supergravity: Horowitz, Strominger, 1991

$$ds^2 = f^{-1/2}(-dt^2 + \sum_1^3 dx_i^2) + f^{1/2}(dr^2 + r^2 d\Omega_5)$$

“Physical” Brane + Extra-Dimensions

$$f = 1 + \frac{R^4}{r^4} ; R^4 = 4\pi\alpha'^2 g_{YM}^2 N_c$$

- “Maldacena limit”:

$$\frac{\alpha'(\rightarrow 0)}{r(\rightarrow 0)} \rightarrow z, R \text{ fixed} \Rightarrow g_{YM}^2 N_c \rightarrow \infty$$

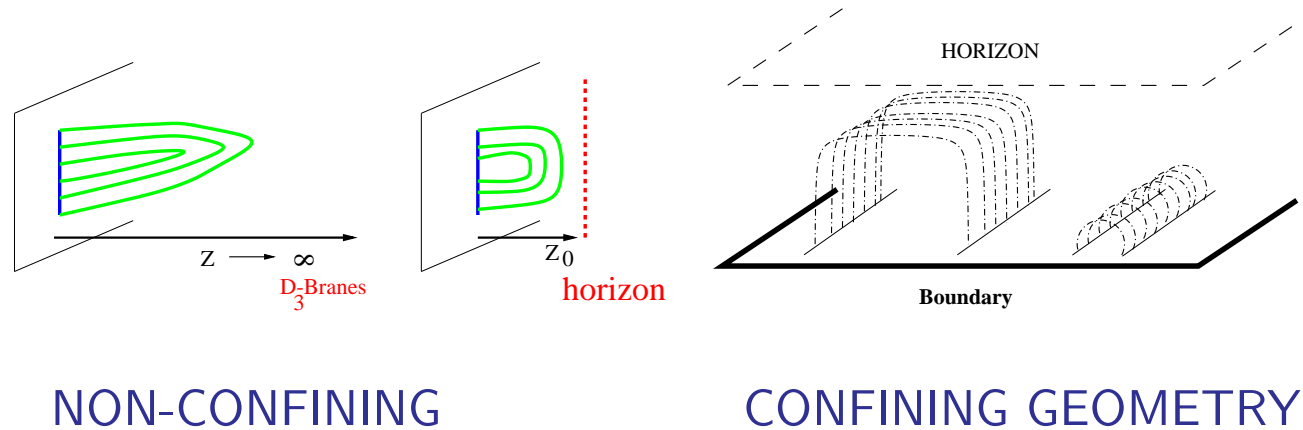
Strong coupling limit

$$ds^2 = \frac{1}{R^2 z^2}(-dt^2 + \sum_{1-3} dx_i^2 + dz^2) + R^2 d\Omega_5$$

Background Structure: $AdS_5 \times S_5$ (same R^2)

HOLOGRAPHY

Exemple of $\langle Wilson Lines \rangle$



NON-CONFINING

CONFINING GEOMETRY

$$\langle e^{iP \int_C \vec{A} \cdot d\vec{l}} \rangle = \int_{\Sigma} e^{-\frac{Area(\Sigma)}{\alpha'}} \approx e^{-\frac{Min.Area}{\alpha'}} \times Fluctuations$$

- Conformal case : AdS_5/CFT_4

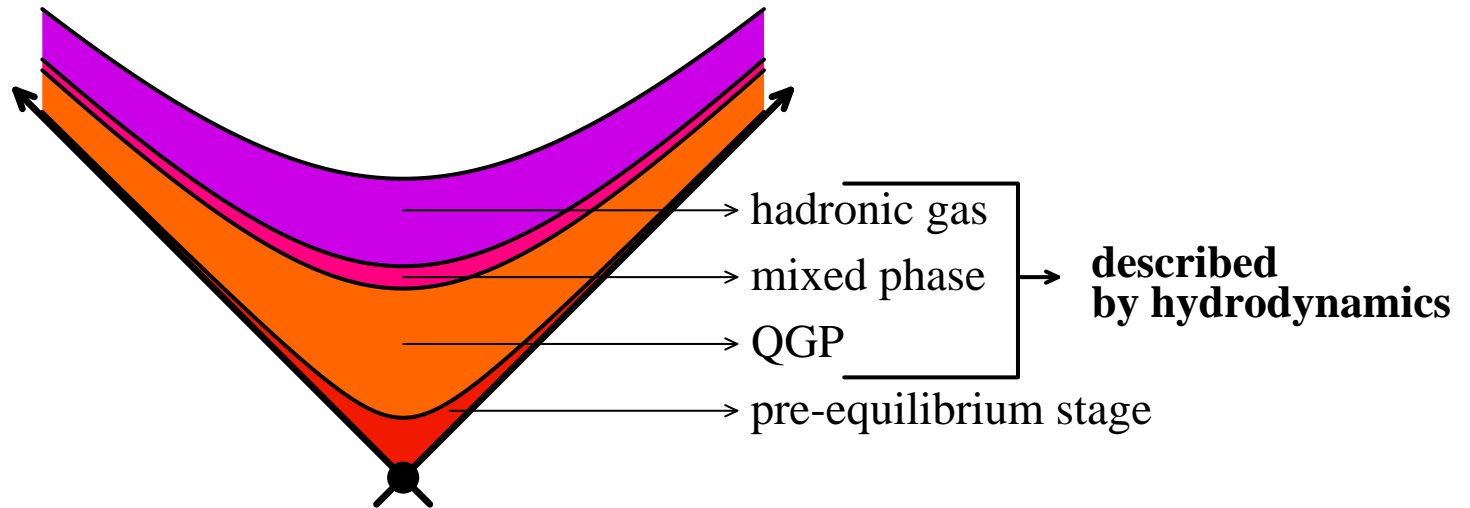
$$\langle Wilson Lines \rangle = e^{T*V(L)} \sim e^{T \times 1/L}$$

- Confining case : ex: Horizon at z_h (Witten 1998)

$$\langle Wilson Lines \rangle = e^{T*V(L)} \sim e^{T \times L (\times z_h^2)}$$

2007

A case Study: QGP formation and Relativistic Hydrodynamics



“Abstracted” from Experiments:

- Kinematic Landscape

$$\tau = \sqrt{x_0^2 - x_1^2} ; \eta = \frac{1}{2} \log \frac{x_0 + x_1}{x_0 - x_1} ; x_T = \{x_2, x_3\}$$

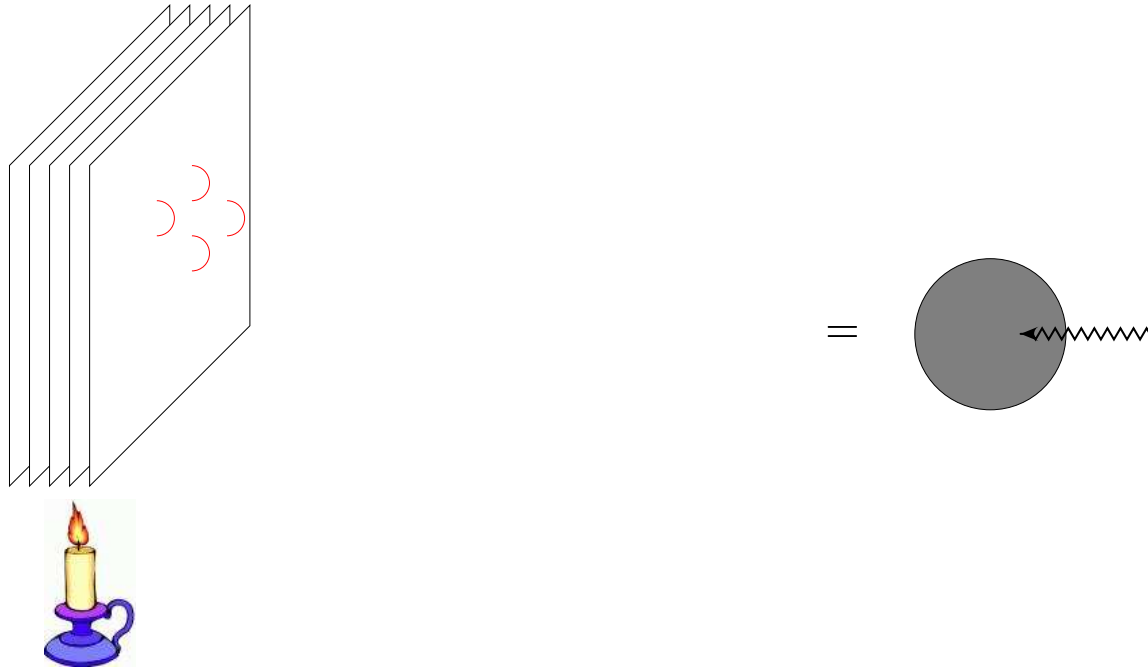
- QGP: (Almost) Perfect fluid behaviour (small Viscosity)
- Fast QGP Formation

Strong Coupling QGP: What Strings can teach us?

Viscosity on the light of duality

Consider a graviton that falls on this stack of N D3-branes
 Will be absorbed by the D3 branes.

The process of absorption can be looked at from two different perspectives:



Absorption by D3 branes (\sim viscosity) = absorption by black hole

$$\sigma_{abs}(\omega) \propto \int d^4x \frac{e^{i\omega t}}{\omega} \langle [T_{x_2x_3}(x), T_{x_2x_3}(0)] \rangle \Rightarrow \frac{\eta}{s} \equiv \frac{\sigma_{abs}(0)/(16\pi G)}{A/(4G)} = \frac{1}{4\pi}$$

Kovtun, Policastro, Son, Starinets (2001-...)

Tool: Holographic Renormalization

K. Skenderis (2002)

- Using Fefferman-Graham Coordinates:

$$ds^2 = \frac{g_{\mu\nu}(z) dx^\mu dx^\nu + dz^2}{z^2}$$

- $4d \Leftrightarrow 5d$ metric:

$$g_{\mu\nu}(z) = g_{\mu\nu}^{(0)} (= \eta_{\mu\nu}) + z^2 g_{\mu\nu}^{(2)} (= 0) + z^4 g_{\mu\nu}^{(4)} (\propto \langle T_{\mu\nu} \rangle) + \dots$$

+ ...: derived from Einstein Eqs.

- $4d$ Constraints

$$T^\mu{}_\mu = 0 \quad ; \quad \mathcal{D}_\nu T^{\mu\nu} = 0$$

EMERGENCE of the 5d BLACK HOLE

Balasubramanian, de Boer, Minic; Myers; Janik, R.P.

- 4d Perfect Fluid “on the brane”

$$\langle T_{\mu\nu} \rangle \propto g_{\mu\nu}^{(4)} = \begin{pmatrix} 3/z_0^4 = \epsilon & 0 & 0 & 0 \\ 0 & 1/z_0^4 = p_1 & 0 & 0 \\ 0 & 0 & 1/z_0^4 = p_2 & 0 \\ 0 & 0 & 0 & 1/z_0^4 = p_3 \end{pmatrix}$$

- Holographic Renormalisation (Resummed)

$$ds^2 = -\frac{(1 - z^4/z_0^4)^2}{(1 + z^4/z_0^4)z^2} dt^2 + (1 + z^4/z_0^4) \frac{dx^2}{z^2} + \frac{dz^2}{z^2}$$

- \Rightarrow 5d Black Brane with horizon at $z_0 \sim T_0^{-3}$

$$ds^2 = -\frac{1 - \tilde{z}^4/\tilde{z}_0^4}{\tilde{z}^2} dt^2 + \frac{dx^2}{\tilde{z}^2} + \frac{1}{1 - \tilde{z}^4/\tilde{z}_0^4} \frac{d\tilde{z}^2}{\tilde{z}^2}$$

$$z \rightarrow \tilde{z} = z / \sqrt{1 + \frac{z^4}{z_0^4}}$$

AdS/CFT: From Statics to Dynamics

R.Janik, RP (2005)

- Constraint Equations

$$\begin{aligned} T^\mu{}_\mu &\equiv -T_{\tau\tau} + \frac{1}{\tau^2}T_{yy} + 2T_{xx} = 0 \\ D_\nu T^{\mu\nu} &\equiv \tau \frac{d}{d\tau} T_{\tau\tau} + T_{\tau\tau} + \frac{1}{\tau^2}T_{yy} = 0 \end{aligned}$$

- Boost-invariant Tensor

$$T_{\mu\nu} = \begin{pmatrix} f(\tau) & 0 & 0 & 0 \\ 0 & -\tau^3 \frac{d}{d\tau} f(\tau) - \tau^2 f(\tau) & 0 & 0 \\ 0 & 0 & f(\tau) + \frac{1}{2}\tau \frac{d}{d\tau} f(\tau) & 0 \\ 0 & 0 & 0 & \dots \end{pmatrix}$$

- 1-parameter family of proper-time evolution

$$f(\tau) \propto \tau^{-s} : T_{\mu\nu} t^\mu t^\nu \geq 0 \Rightarrow 0 < s < 4$$

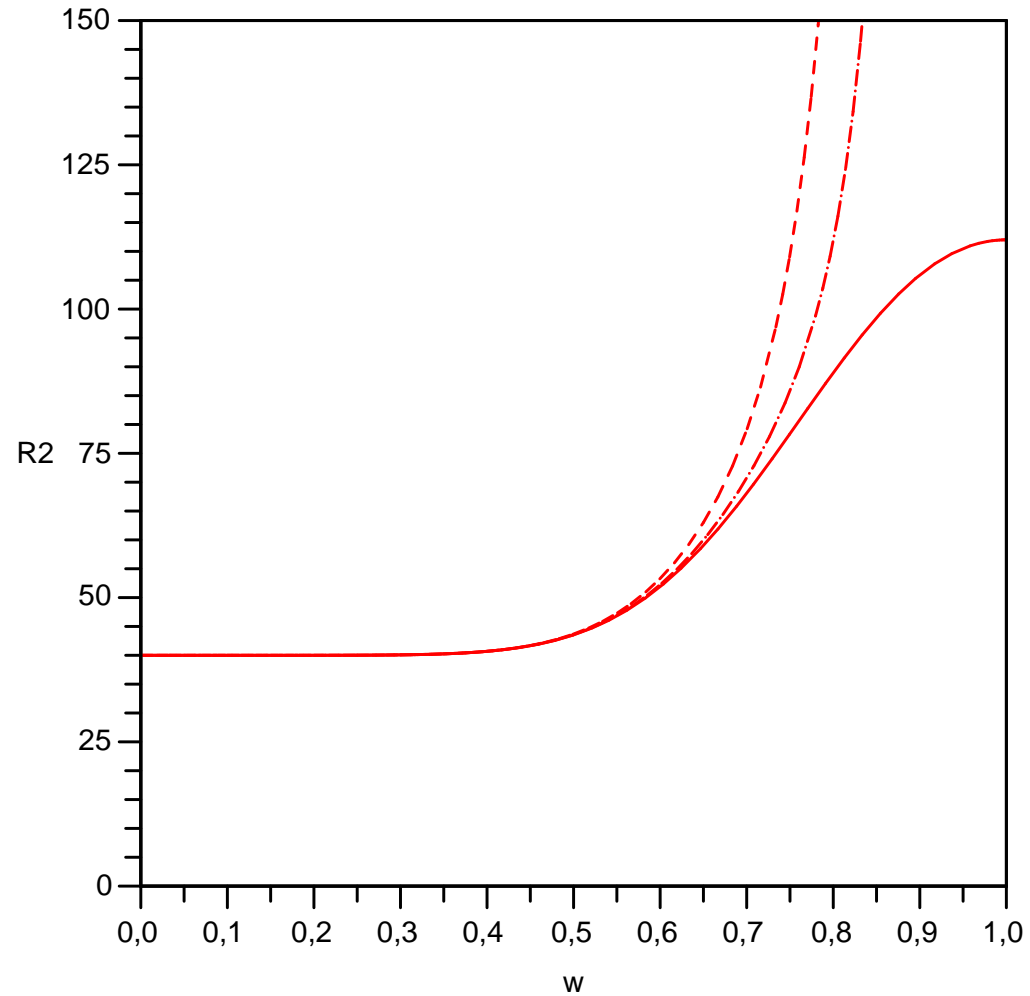
$$f(\tau) \propto \tau^{-\frac{4}{3}} : \text{Perfect Fluid}$$

$$f(\tau) \propto \tau^{-1} : \text{Free streaming}$$

$$f(\tau) \propto \tau^{-0} : \text{“Full Anisotropy” } \epsilon = p_\perp = -p_L$$

AdS/CFT \Rightarrow Perfect Fluid at large τ

$$\mathfrak{R}^2 = R^{\mu\nu\alpha\beta} R_{\mu\nu\alpha\beta}$$



$$s = \frac{4}{3} \pm .1$$

A nonsingular background selects a moving Black Hole geometry corresponding to the perfect fluid at large proper-times

Dual of a Perfect fluid in S^4 QCD “heavy ion experiment” : a Moving Black Hole

$$v = \frac{z}{\tau^{1/3}}$$

- Asymptotic metric

$$ds^2 = \frac{1}{z^2} \left[-\frac{\left(1 - \frac{e_0}{3} \frac{z^4}{\tau^{4/3}}\right)^2}{1 + \frac{e_0}{3} \frac{z^4}{\tau^{4/3}}} d\tau^2 + \left(1 + \frac{e_0}{3} \frac{z^4}{\tau^{4/3}}\right) (\tau^2 dy^2 + dx_{\perp}^2) \right] + \frac{dz^2}{z^2}$$

- Black Hole off in the 5th dimension

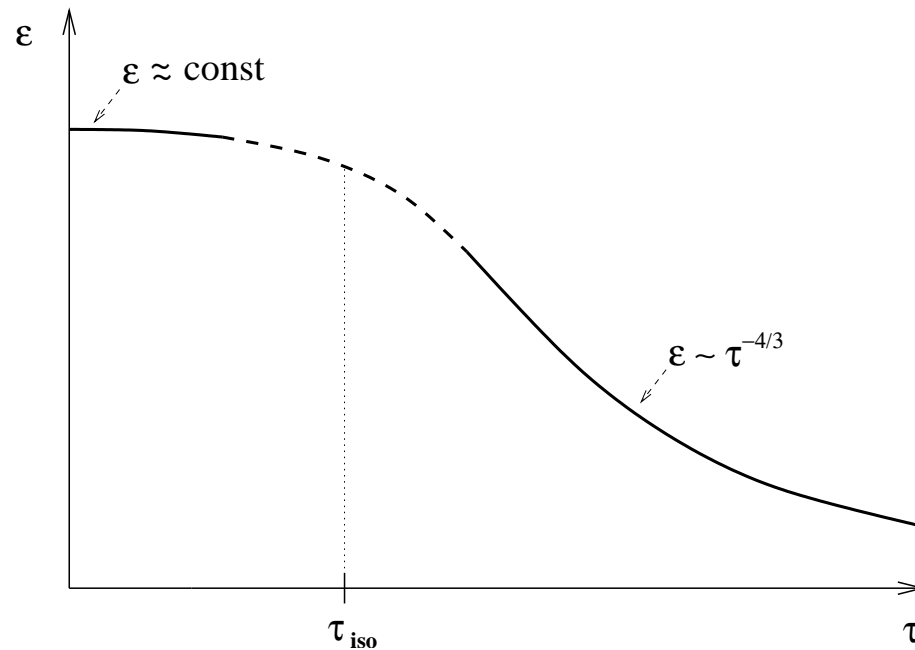
$$\text{Horizon : } z_h = \left(\frac{3}{e_0}\right)^{\frac{1}{4}} \cdot \tau^{\frac{1}{3}} .$$

$$\text{Temperature : } T(\tau) \sim \frac{1}{z_h} \sim \tau^{-\frac{1}{3}}$$

$$\text{Entropy : } S(\tau) \sim \text{Area} \sim \tau \cdot \frac{1}{z_h^3} \sim \text{const}$$

AdS/CFT: Anisotropy at small τ

Kovchegov, Taliotis arXiv:0705.1234



Evaluation of τ_{iso}

$$\text{Matching : } z_h^{late}(\tau) = \left(\frac{3}{e_0}\right)^{\frac{1}{4}} \equiv z_h^{early}(\tau) = \tau$$

$$\text{Isotropization : } \tau_{iso} = \left(\frac{3N_c^2}{2\pi^2 e_0}\right)^{3/8}$$

$$\text{Evaluation : } \epsilon(\tau) = e_0 \tau^{4/3}|_{\tau=.6} \sim 15 \text{ GeV fermi}^{-3} \Rightarrow \tau_{iso} \sim .3 \text{ fermi}$$

Conclusions

In progress:

- Gauge-Gravity Correspondence
A promising way towards QCD at strong couplings
- Results on AdS/CFT \rightarrow S^4 QCD Hydrodynamics
Perfect Fluid, Viscosity, Isotropization
- Other studies
Jet Quenching, Quark Dragging, and many others...

In outlook:

- Can we go beyond Boost Invariance?
From Bjorken to Landau to “real” Hydrodynamics?
- Can we follow the flow from Ions to Hadrons?
Initial and Final conditions for Hydrodynamics
- From S^4 QCD to S^0 QCD Hydrodynamics ?
Can we construct the “Dual” of the QGP?

Some references...

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EXTRA SLIDES

Initial Conditions: Shock Waves (1)

- One Initial Shock Wave

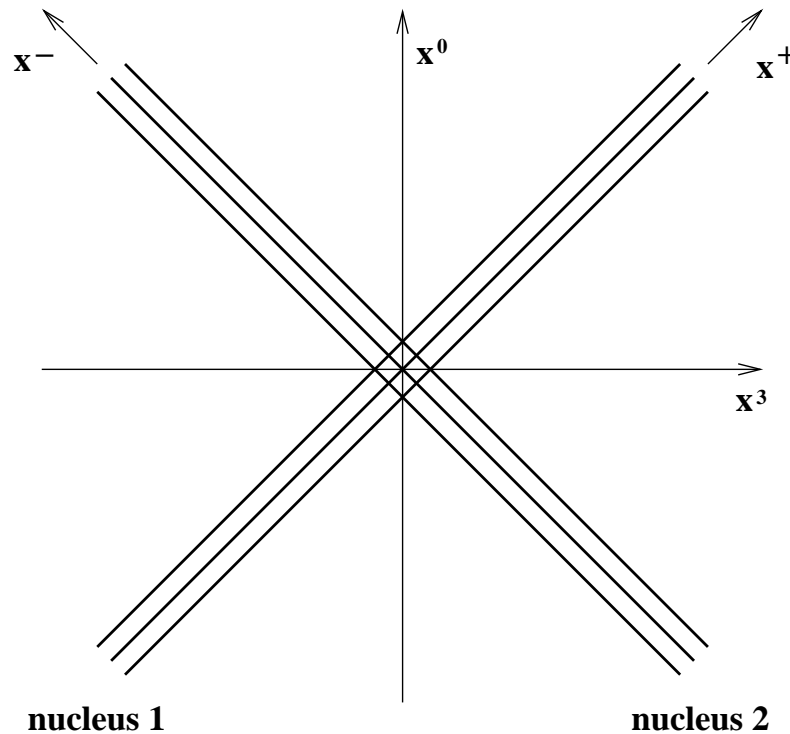
Janik, R.P. (2005)

$$ds^2 = \frac{-2dx^+dx^- + \mu_1 z^4 \delta(x^-) dx^{-2} + d\mathbf{x}_\perp^2 + dz^2}{z^2}$$

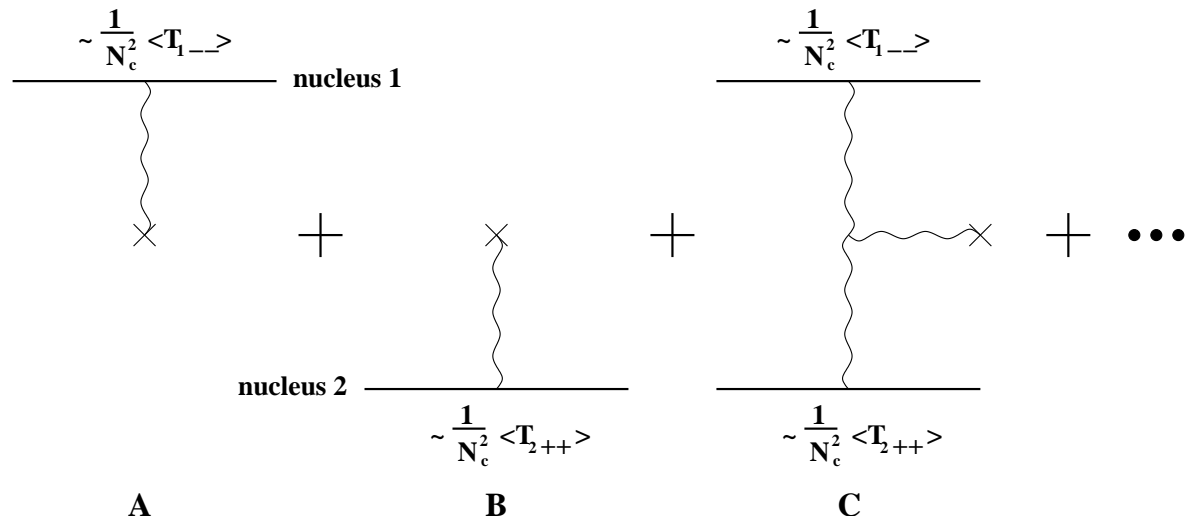
- Shock-wave collisions

Grumiller, Romatschke (2008)

Albacete, Kovchegov, Taliotis (2008)



Initial Conditions: Shock Waves (2)

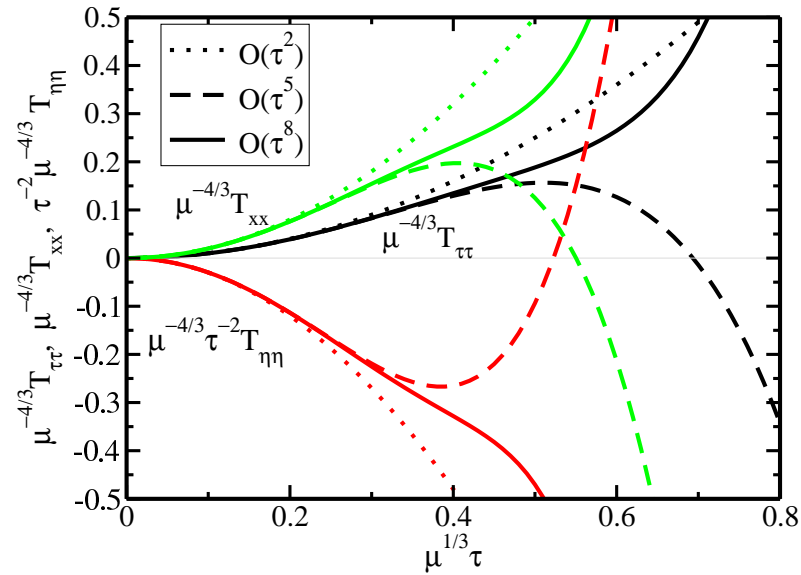


$$ds^2 = \frac{L^2}{z^2} \left\{ -2 dx^+ dx^- + dx_\perp^2 + dz^2 + \frac{2\pi^2}{N_c^2} \langle T_{1--}(x^-) \rangle z^4 dx^{-2} + \frac{2\pi^2}{N_c^2} \langle T_{2++}(x^+) \rangle z^4 dx^{+2} \right. \\
 \left. + \text{higher order graviton exchanges} \right\} \quad (1)$$

From: Albacete, Kovchegov, Taliotis

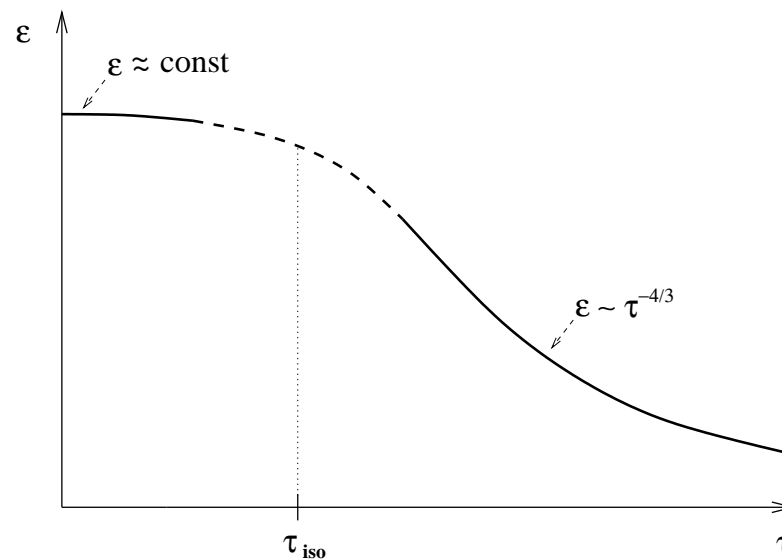
Initial Conditions: Shock Waves (3)

- Strong Coupling: Full Stopping \sim Landau



From: Grumiller, Romatschke

- Weak Coupling



Calculation of the Moving Duals

- Boost-Invariant 5-d F-G metric:

$$ds^2 = \frac{-e^{a(\tau,z)} d\tau^2 + \tau^2 e^{b(\tau,z)} dy^2 + e^{c(\tau,z)} dx_{\perp}^2}{z^2} + \frac{dz^2}{z^2}$$

- Scaling : $v = \frac{z}{\tau^{s/4}}$

$$[a(\tau, z), b(\tau, z), c(\tau, z)] = [a(v), b(v), c(v)] + \mathcal{O}\left(\frac{1}{\tau^{\#}}\right)$$

$$v(2a'(v)c'(v) + a'(v)b'(v) + 2b'(v)c'(v)) - 6a'(v) - 6b'(v) - 12c'(v) + vc'(v)^2 = 0$$

$$3vc'(v)^2 + vb'(v)^2 + 2vb''(v) + 4vc''(v) - 6b'(v) - 12c'(v) + 2vb'(v)c'(v) = 0$$

$$2vsb''(v) + 2sb'(v) + 8a'(v) - vsa'(v)b'(v) - 8b'(v) + vsb'(v)^2 +$$

$$4vsc''(v) + 4sc'(v) - 2vsa'(v)c'(v) + 2vsc'(v)^2 = 0 .$$

- Asymptotic Solution

$$a(v) = A(v) - 2m(v)$$

$$b(v) = A(v) + (2s - 2)m(v)$$

$$c(v) = A(v) + (2 - s)m(v)$$

Free Streaming and general cases

$$v = \frac{z}{\tau^{1/4}}$$

- Asymptotic metric

$$ds^2 = \frac{\left(-\left(1 + \frac{v^4}{\sqrt{8}}\right) \frac{1-2\sqrt{2}}{2} \left(1 - \frac{v^4}{\sqrt{8}}\right) \frac{1+2\sqrt{2}}{2} dt^2 + \left(1 + \frac{v^4}{\sqrt{8}}\right)^{\frac{1}{2}} \left(1 - \frac{v^4}{\sqrt{8}}\right)^{\frac{1}{2}} \tau^2 dy^2 + \right.}{z^2} + \frac{\left. + \left(1 + \frac{v^4}{\sqrt{8}}\right) \frac{1+\sqrt{2}}{2} \left(1 - \frac{v^4}{\sqrt{8}}\right) \frac{1-\sqrt{2}}{2} dx_{\perp}^2 \right)}{z^2} + \frac{dz^2}{z^2}$$

- True Singularities ?

Ricci scalar:

$$R = -20 + \mathcal{O}\left(\frac{1}{\tau^{\#}}\right).$$

Riemann tensor squared:

$$\mathfrak{R}^2 = R^{\mu\nu\alpha\beta} R_{\mu\nu\alpha\beta}$$

AdS/CFT: Selection of the Perfect Fluid

- Curvature invariant: $\mathfrak{R}^2 = R^{\mu\nu\alpha\beta} R_{\mu\nu\alpha\beta}$

$$\mathfrak{R}^2 = \frac{4}{(1 - \Delta(s)^2 v^8)^4} \cdot \left[\begin{aligned} &10 \Delta(s)^8 v^{32} - 88 \Delta(s)^6 v^{24} + 42 v^{24} s^2 \Delta(s)^4 + \\ &+ 112 v^{24} \Delta(s)^4 - 112 v^{24} \Delta(s)^4 s + 36 v^{20} s^3 \Delta(s)^2 - 72 v^{20} s^2 \Delta(s)^2 + \\ &+ 828 \Delta(s)^4 v^{16} + 288 v^{16} \Delta(s)^2 s - 288 v^{16} \Delta(s)^2 - 108 v^{16} s^2 \Delta(s)^2 + \\ &- 136 v^{16} s^3 + 27 v^{16} s^4 - 320 v^{16} s + 160 v^{16} + 296 v^{16} s^2 + 36 v^{12} s^3 + \\ &- 72 v^{12} s^2 - 88 \Delta(s)^2 v^8 + 42 v^8 s^2 + 112 v^8 - 112 v^8 s + 10 \end{aligned} \right] + \mathcal{O}\left(\frac{1}{r^\#}\right)$$

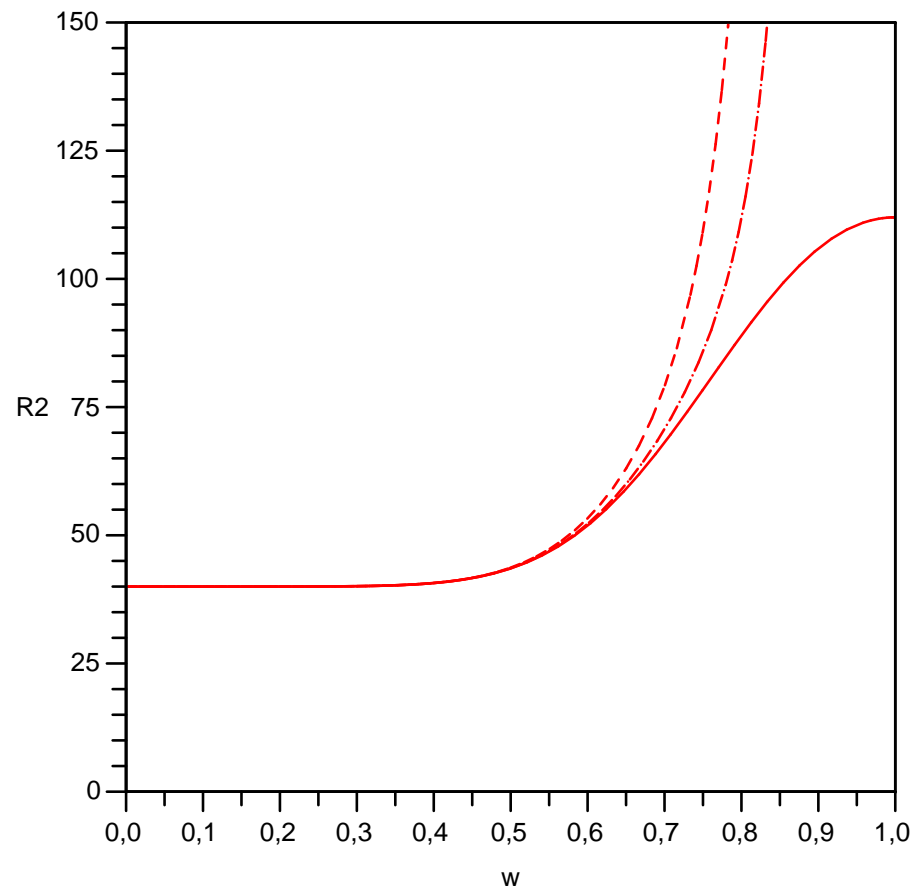
- \mathfrak{R}^2 for $s = \frac{4}{3}$:

$$\mathfrak{R}^2_{\text{perfect fluid}} = \frac{8(5w^{16} + 20w^{12} + 174w^8 + 20w^4 + 5)}{(1 + w^4)^4}$$

where $w = v/\Delta\left(\frac{4}{3}\right)^{\frac{1}{4}} \equiv \sqrt[4]{3} v$.

AdS/CFT: Selection of the Perfect Fluid

\mathcal{R}^2 for $s = \frac{4}{3} \pm .1$



Thermalization response-time of a perfect fluid

- Quasi-normal scalar modes of a static Black Hole in Fefferman-Graham coordinates

$$\Delta\phi \equiv \frac{1}{\sqrt{-g}} \partial_i (\sqrt{-g} g^{ij} \partial_j \phi) = 0$$

$$-\frac{1}{z^3} \frac{\left(1 - \frac{z^4}{z_0^4}\right)^2}{\left(1 + \frac{z^4}{z_0^4}\right)} \partial_t^2 \phi(t, z) + \partial_z \left(\frac{1}{z^3} \left(1 - \frac{z^8}{z_0^8}\right) \partial_z \phi(t, z) \right) = 0.$$

- Separation of variables $\phi(t, z) = e^{i\omega t} \phi(z)$

$$\phi'' + \frac{1 - \tilde{z}^2}{\tilde{z}(1 - \tilde{z})(2 - \tilde{z})} \phi' + \left(\frac{\omega}{\pi T}\right)^2 \frac{1}{4\tilde{z}(1 - \tilde{z})(2 - \tilde{z})} \phi = 0$$

- Dominant Decay Time

$$\frac{\omega_c}{\pi T} \sim 3.1194 - 2.74667 i$$

- A note on Viscosity vs. Q-n modes

Thermalization response-time of a perfect fluid

- Quasinormal scalar modes for the boost-invariant Black Hole geometry

$$\Delta\phi \equiv \frac{1}{\sqrt{-g}} \partial_i (\sqrt{-g} g^{ij} \partial_j \phi) = 0$$

$$\left[\partial_z \rightarrow \tau^{-\frac{1}{3}} \partial_v ; \partial_\tau \rightarrow \partial_\tau - \frac{1}{3} \tau^{-\frac{4}{3}} \partial_v \right]$$

$$-\frac{1}{v^3} \frac{(1+v^4)^2}{1-v^4} \partial_\tau^2 \phi(\tau, v) + \tau^{-\frac{2}{3}} \partial_v \left(\frac{1}{v^3} (1-v^8) \partial_v \phi(\tau, v) \right) = 0$$

- Separation of variables $\phi(\tau, v) = f(\tau)\phi(v)$

$$\partial_\tau^2 f(\tau) = -\omega^2 \tau^{-\frac{2}{3}} f(\tau) \Rightarrow f(\tau) = \sqrt{\tau} J_{\pm\frac{3}{4}} \left(\frac{3}{2} \omega \tau^{\frac{2}{3}} \right) \sim \tau^{\frac{1}{6}} e^{\frac{3}{2} i \omega \tau^{\frac{2}{3}}}$$

$$\partial_v \left(\frac{1}{v^3} (1-v^8) \partial_v \phi(v) \right) + \omega^2 \frac{1}{v^3} \frac{(1+v^4)^2}{1-v^4} \phi(v) = 0$$

- Dominant Decay Proper-Time

$$\frac{\omega_c}{\pi T} \sim 3.1194 - 2.74667 i$$