

E-Science and research in astronomy, astrophysics and cosmology

NARIT

Utane Sawangwit



Outline

- Overview of the use of e-science in astronomy, astrophysics and cosmology
- Cosmology and structure evolutions and even larger facilities/surveys
- Review of NARIT facilities and need for e-science & Research and Education network





We are living during an age of very fast changes in the way to do Astronomy. Astronomy is one of the prime sciences of the human being

Astronomy goals through the centuries:

- Registering and recording of heaven events
- Developing of methodologies to keep and track such events
- **SHARING** the information and the knowledge
- Astronomy is one of the most internationally organized science endeavors (and of course, experimental particle physics)





- The technology (computers, communications, etc.) is changing the way to access to the information.
- Observational facilities based in astronomical satellites, "automated survey telescopes" and 48 meters class telescopes are available everywhere.
- Amazing facilities are planned in the coming decade
- The information gathered by these facilities (and their derived products) is organized in huge databases...
- and also, the services of such databases is being connected through computer grids **"Virtual Observatory"** is here now
- The observations are "there" in fact, we "don't need" telescopes to do high-quality observational astronomy

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For our countries with scarce resources:

- Opportunity to access to high-quality material.
- We can do good "**observational astrophysics**" to relatively very low costs: free scientific software based on LINUX are available through Internet.
- Training young astronomer in the use of these observations as an "experimental observatory" to apply for observing time in international observatories (ground and space).
 - Participating in the development of astronomical databases and grids and more...



National Astronomical Research Institut e of Thailand (Public Orrian..zation) Ministry o Sile ic is to et ic or v of Thailard

International Virtual Observatory Alliance

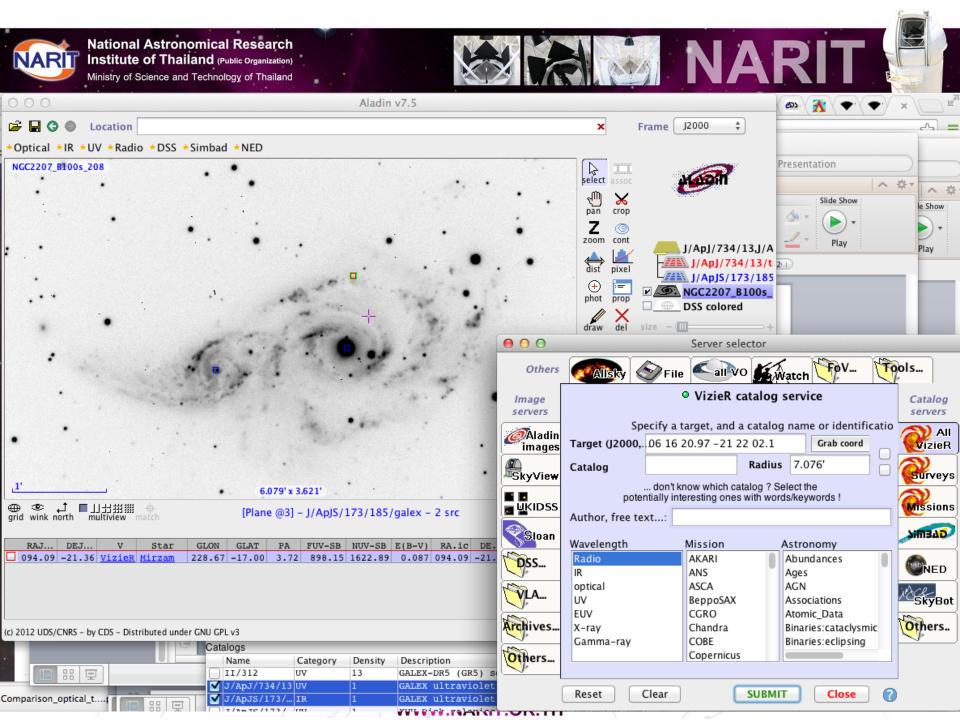
Member Organizations



- Argentine Virtual Observatory
- Armenian Virtual Observatory
- AstroGrid, United Kingdom

Concervatory AlfiaRce

- Australian Virtual Observatory
- Brazilian Virtual Observatory
- Chinese Virtual Observatory
- Canadian Virtual Observatory
- Chilean Virtual Observatory
- European Space Agency
- European Virtual Observatory
- German Astrophysical Virtual Observatory
- Hungarian Virtual Observatory
- Japanese Virtual Observatory
- Observatoire Virtuel France
- Russian Virtual Observatory
- South African Astroinformatics Alliance
- Spanish Virtual Observatory
- Italian Virtual Observatory
- Ukrainian Virtual Observatory
- Virtual Astronomical Observatory, USA
- Virtual Observatory India



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÷		Virtual Observatory Plot Tool Web Browser Catalog Tool	Chandra ROSAT All-Sky (MPE/MPG) SkyView (NASA/HEASARC) W3Browse (NASA/HEASARC)
		Analysis Command Log	NVSS (NRAO)
	•	Load Analysis Commands Clear Analysis Commands	4MASS (NRAO) SIRTF FLS/VLA (NRAO) VLA First (NRAO)
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		.png	SAO TDC (SAO)



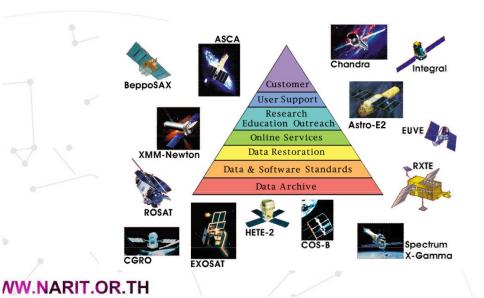
Web-based data archives

 Infrared missions (IRAS, ISO, Spitzer, Herschel, Planck, WISE) : IPAC http://www.ipac.caltech.edu

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Infrared Processing and Analysis Center Science and Data Center for Infrared Astronomy

- Highenergy: NASA HEASARC http://heasarc.gsfc.nasa.gov
 - Information and latest news about HEASARC Catalogs
 - Mission information
 - Search catalogs & retrieve data
 - Download analysis software
 - Access documentation
 - Astronomical Web site links
 - Public outreach & education





Web-based data archives

• Cosmic Microwave Background (CMB)

NASA's LAMBDA @ HEASARC Legacy Archive for Microwave Background Data Analysis "One Stop Shopping for CMB Researchers"

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Welcome to NASA's data center for Cosmic Microwave Background (CMB) research







Planck Data Products at IRSA

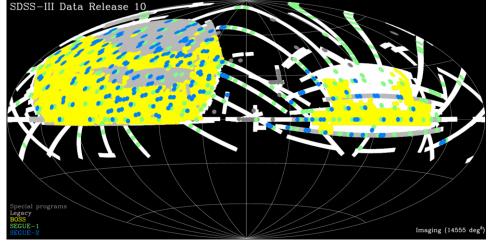
Planck is a European Space Agency (ESA) mission with significant NASA involvement in hardware and science, managed by the Jet Propulsion Laboratory, California Institute of Technology. Planck data products are provided by the Planck Consortium. Products are released by ESA through the archive at ESA Planck Legacy Archive (http://www.sciops.esa.int/index.php? project=planck&page=Planck_Legacy_Archive) and by NASA through the NASA/IPAC Infrared Science Archive (http://irsa.ipac.caltech.edu/), which is maintained at IPAC, California Institute of Technology, under contract with the National Aeronautics and Space Administration.

The links in the first two tables below lead directly to Planck data at IRSA.

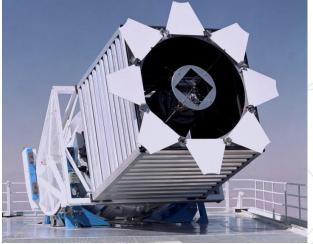
Data Release 1						
Product Download Page	Description	No. Of Files/Size				
All Sky Maps	Temperature and component separated maps.	72 files				
Catalog and Map Visualization	Visualization and enhanced access to the Planck All Sky Maps and the ERCSC	N/A				
Catalog Query	Access to the ERCSC through IRSA's Catalog Search service.	N/A				
Planck SZ Catalogs	Download the Planck Sunyaev Zeldovich Cluster sample.	1 file1				
Entire PCCS	Download the Planck Catalog of Compact Sources	1 file				
Planck Ancillary Data	Power spectra, masks and instrument parameters	78 files				
External Datasets	Multiwavelength maps useful for analysis of Planck data	38 files				
Software and Analysis Products	Useful software and products for analysis of Planck data	11 files				
Data Release 1 Explanatory Supplement	Documentation for Data Release 1	1 file				

The Sloan Digital Sky Survey

- 30 2k x 2k
- 6col. x 5 filt
- Drift scan mode
- FOV 2.5deg
- 200GB/ngt



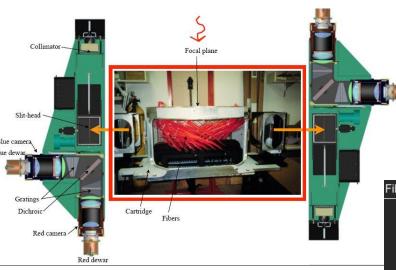
2.5m @ Apache Point Obs. since year 2000



Imaging data statistics

Total unique area covered	14,555 square degrees		
Total area of imaging (including overlaps)	31,637 square degrees (excluding supernovae runs)		
Individual image field size	1361x2048 pixels (0.0337 square degrees)		
Number of individual fields	938,046 (excluding supernovae runs)		
Number of catalog objects	1,231,051,050		
Number of unique detections	932,891,133		
	Total 469,053,874		
	Stars 260,562,744		
Number of unique, primary sources	Galaxies 208,478,448		
	Unknown 12,682		





BOSS spectrograph

- 1000 fibers
- 2 separate blue & red channel
- R ~ 2000
- Dark Energy experiment with BAO
- 1.5 millions LRGs
- 160,000high-z quasars (Ly-alpha forest)

SDSS Spectroscopic survey

SDSS spectrograph BOSS spectrograph

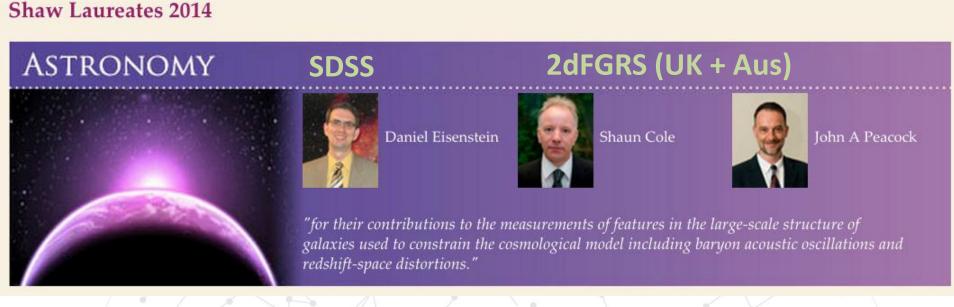
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ers per plate	640				1000			
	Category	Total	On good or marginal plates	Unique				
	All programs	1,843,200	1,768,960	1,629,129	Category Total Total 1,507,954 Stars 159,327 Galaxies 927,844 Quasars 182,009 Sky 144,503	144,968 859,322		
	Main galaxy targets	778,410	755,111	711,726				
mbers of spectra	LRG targets (excluding Main)	106,650	103,662	95,990				
	SEGUE-1 targets	250,422	242,008	220,851				
	SEGUE-2 targets	128,112	128,112	118,151	Unknown	101,550	89,003	
	Stars	600,967	577,157	521,990				
	Galaxies	952,740	921,007	860,836				
	Quasars	130,300	126,368	116,003				
	Skies	110,288	103,046	93,187				
	Unknown	48,905	41,382	37,113				
		1	-	1		O	10	



NARIT Many great results from SDSS

- 5800 peer-reviewed publication over 15 years
- These paper in turn have been cited a total 245,000 times





NARIT Multi-mission & multi-wavelength data archives

NED http://ned.ipac.caltech.edu/

CDS http://cds.u-strasbg.fr/







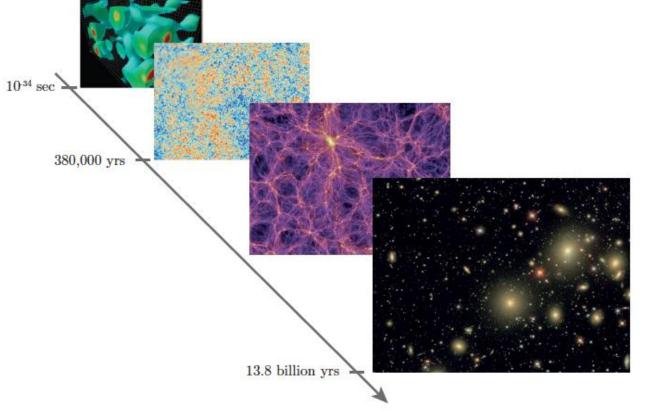
MAST https://archive.stsci.edu/



The Internet's Virtual Telescope

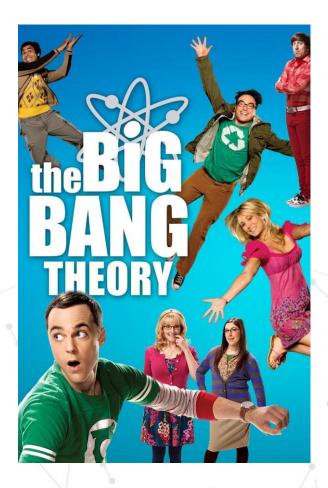


Cosmology and structures evolution





"It all started with a Big Bang"







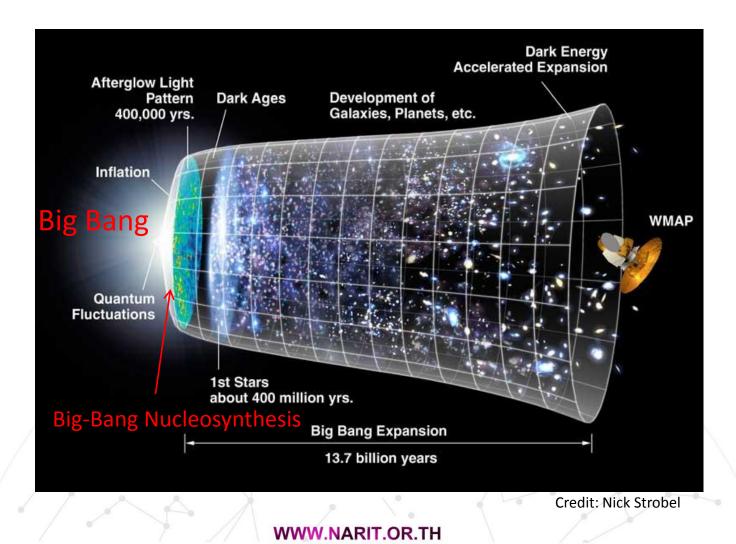
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- This implies that all the matter in the Universe was once (14 billion years ago) **very close together in an extremely hot dense state** of the so called 'Big Bang' theory (a term coined by Sir Fred Hoyle)
- Proposed by G. Lemaitre (1931) "Hypothesis of the primeval atom"
- The universe has since expanded and cooled down to temperature of **2.725 Kelvin**
- What we learned about the universe is now called **modern cosmology** which began with observations made by Edwin Hubble



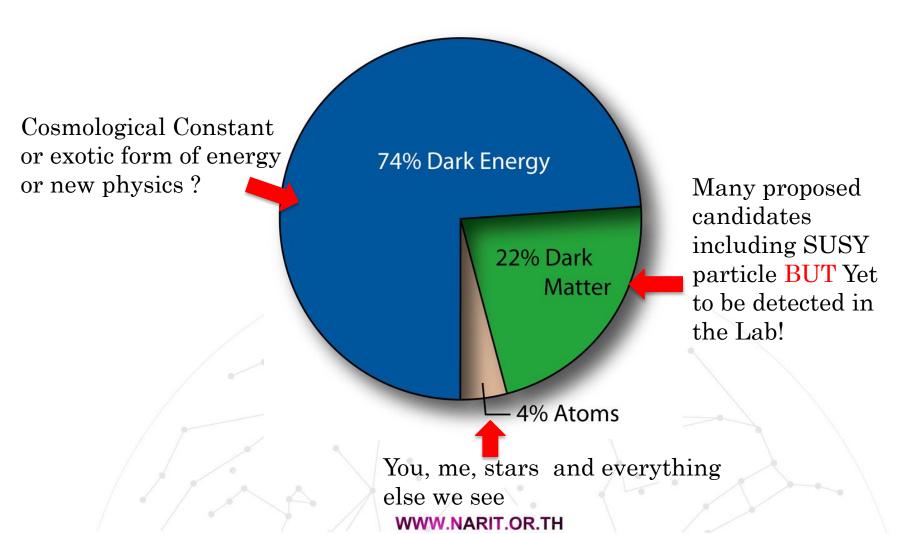
Quick overview of the past 14 billion years

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What the Universe is made of?



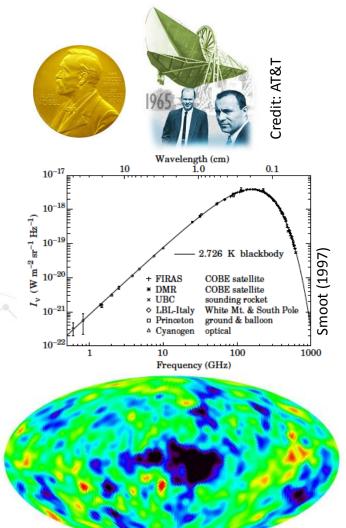


Cosmic Microwave Background (CMB)

Brief History

- Discovered by chance as the "Noise" by Penzias and Wilson (1964), T ~3K (c.f. ~5K prediction from BBN by Alpher & Herman 1948), Nobel prize 1978
- Interpret as the relic radiation left from the Hot Big Bang (Dicke, Peebles, Roll & Wilkinson 1965)
- 1990, FIRAS on the COBE mission measured the near-perfect blackbody spectrum
- 1992, DMR on the COBE mission measured for the first time the 10⁻⁵ primary fluctuation
- Late 1990s and early 2000s, the "precision" measurements for anisotropy begin in earnest (balloon, ground-based & Space)

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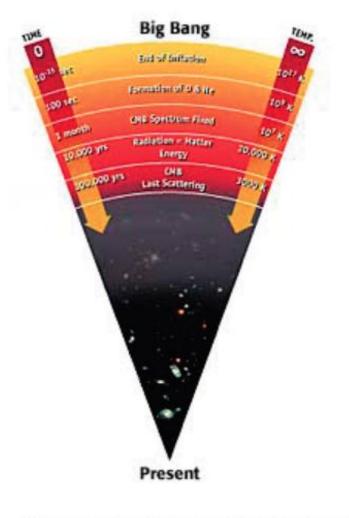
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Cosmic Michrowave Background Universe was "fire ball" 14 billion years ago (redshift~1000) Everywhere was plasma gas \rightarrow electrons scattered photons \rightarrow Everywhere was in "fog". photon Universe expanded, got cooler and less dense. \rightarrow Protons and electrons "decoupled" after 380000 years Universe became "Transparent" We can observe the last scattering surface toward all directions at the same distance



The last scattering surface: a snapshot of the early universe



Can only see surface of cloud where light was last scattered.

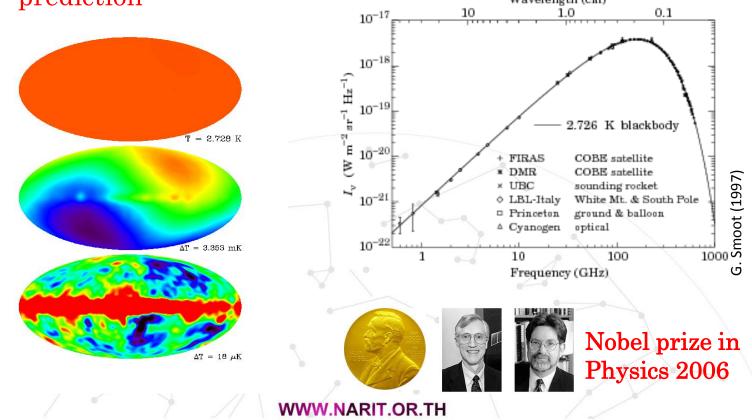
The Cosmic Microwave Background Radiation's "surfact of last scatter" is analogous to the light coming through the clouds to our eye on a cloudy day.





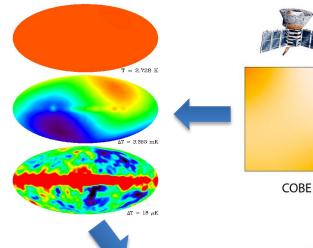
Golden Age of Observational Cosmology

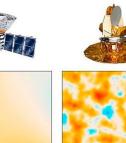
 Cosmic Microwave Background (CMB) Remnant of heat radiation left over from the hot Big Bang B.B. spectrum, T₀=2.725 K consistent with the BBN prediction



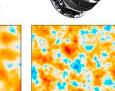


The Golden Age of Observational Cosmology

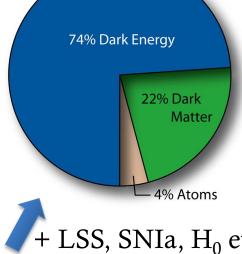


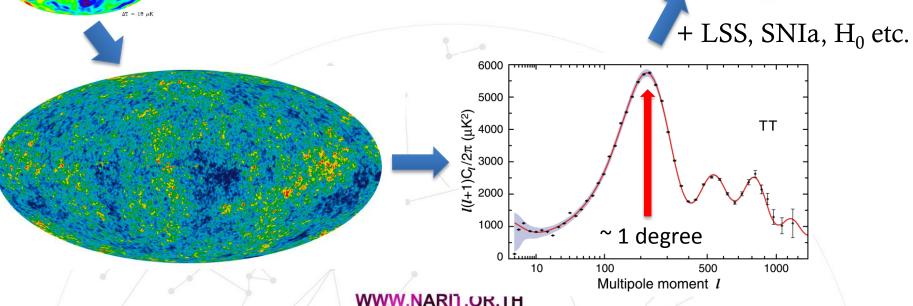






Planck

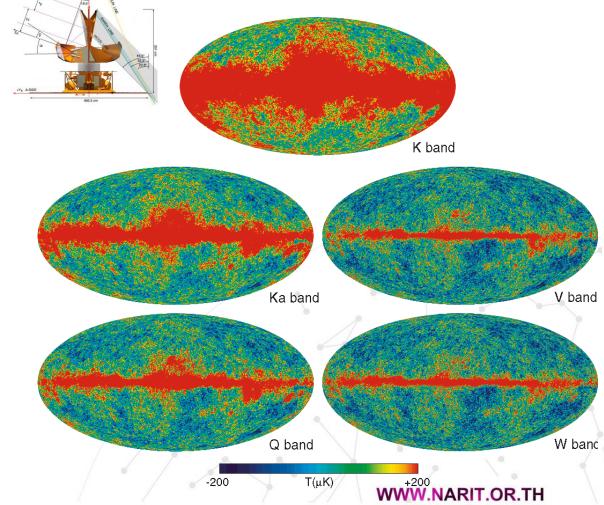


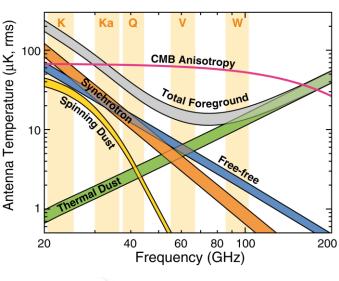


WMAP



The Wilkinson Microwave Anisotropy Probe (WMAP) mission

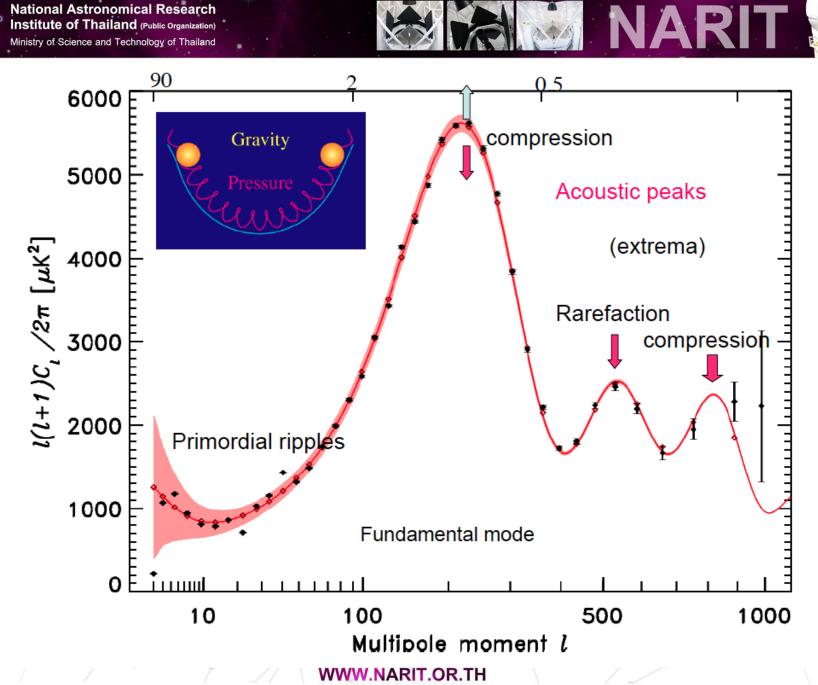




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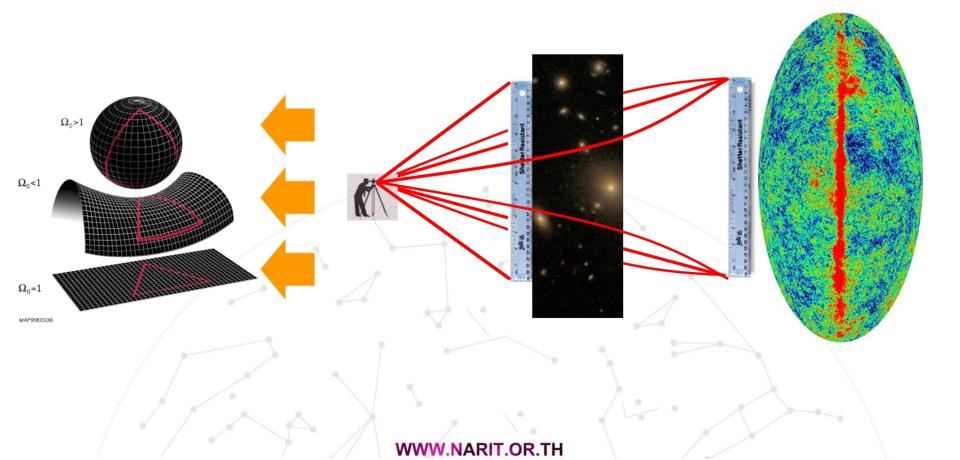
CMB + Foreground radiation NARIT

National Astronomical Research Institute of Thailand (Public Organization) Ministry of Science and Technology of Thailand

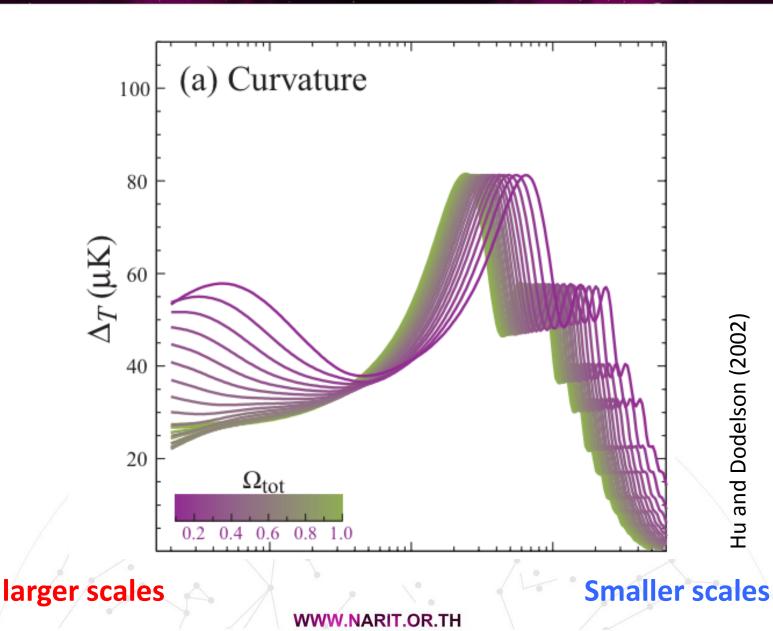




Measuring curvature from the CMB acoustic peak









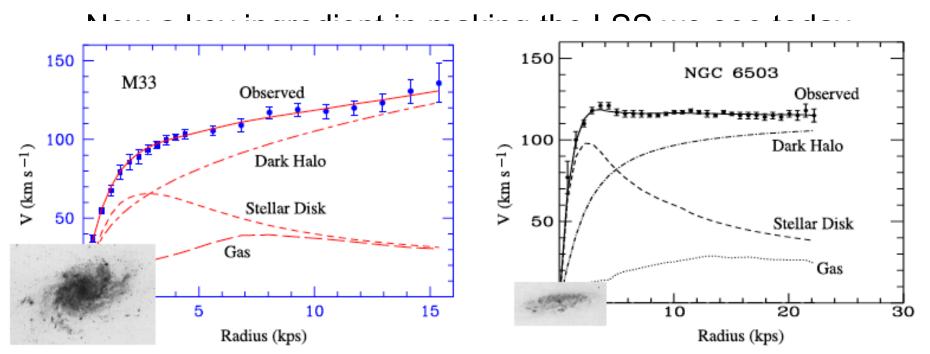
The Dark Universe





Dark Matter

- Proposed by Fritz Zwicky (1933) to explain the missing mass in the galaxy rotation curve
- The missing mass inferred from Mass-to-light ratio is about factor of 10 for galaxy and a few 100s for galaxy cluster



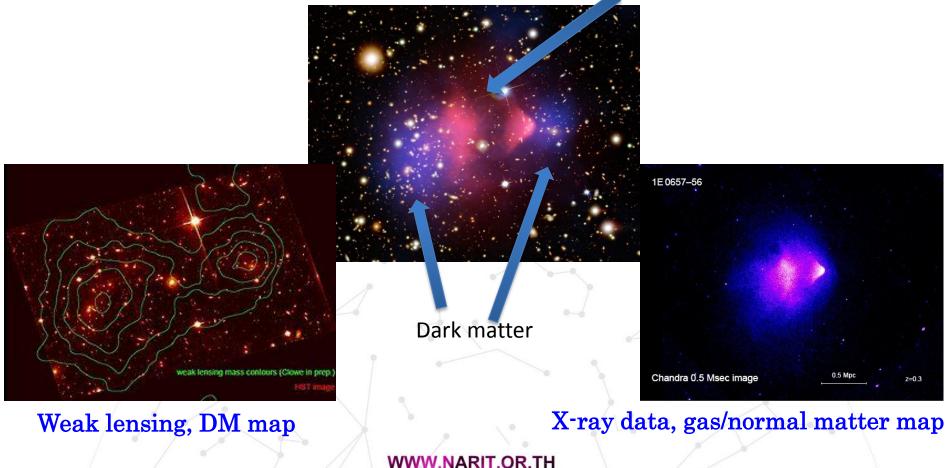




Smoking Gun for Dark Matter

Bullet Cluster (1E 0657-558) (Cowe et al.)

Normal matter

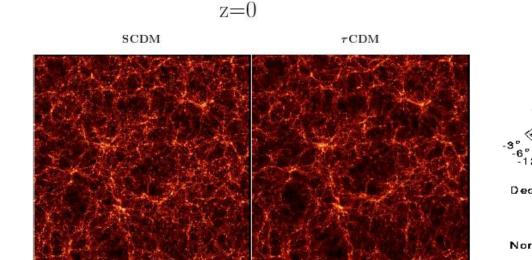






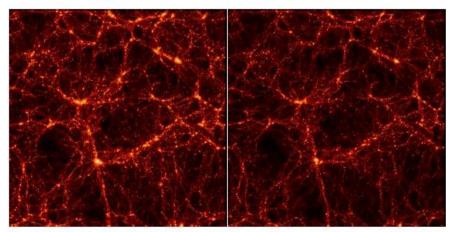
simulations

observations

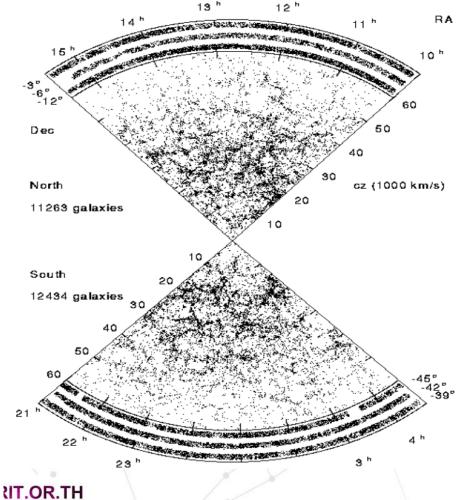


ACDM

OCDM



The VIRGO Collaboration 1996



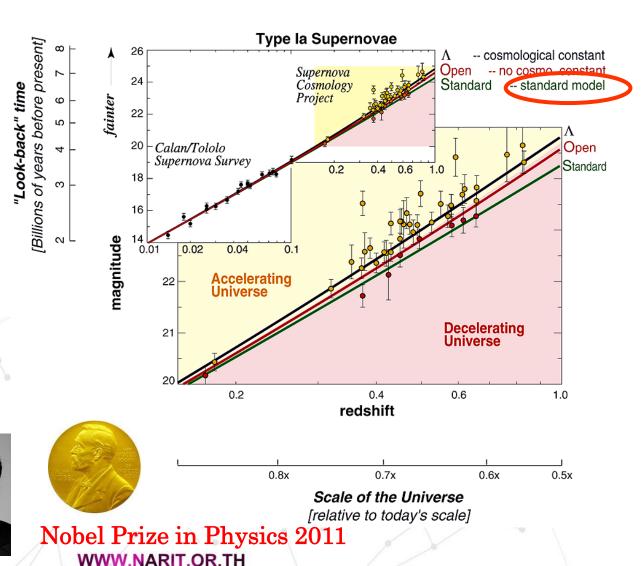


Accelerated expansion; Dark Energy or New Physics?

- In 1998, the first Hubble diagram using Type Ia supernovae was constructed
- It showed that expansion of the universe is accelerating!
- Result by Perlmutter et al quickly confirmed by Riess et al



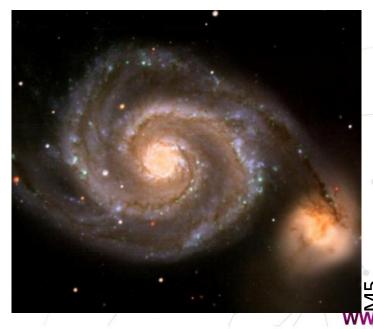


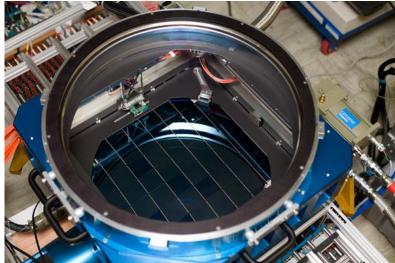




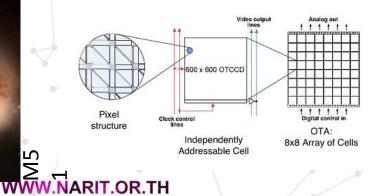
The Panoramic Survey Telescope and Rapid **Response System**







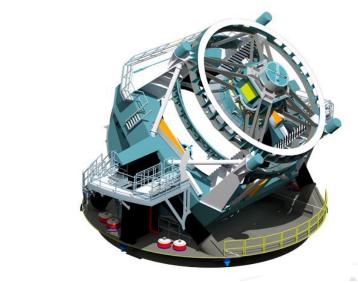
1.4 Gigapixels Camera (600 x 600) pixels x (64 x 64) arrays



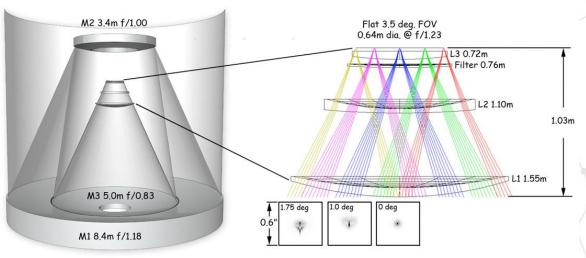
2GB /raw img ~ several TB/nt

Large Synoptic Survey Telescope

the widest, fastest, deepest eye of the new digital age



Construction started Aug.1, 2014 First light around 2022 Dark Matter + Dark Energy surveys





FOV 3.5deg3.2 Giga-pixels



Space Missions



GAIA launched December 2013 Aims to construct a 3D space catalog of Approx. 1 billions stars in the MW from L2

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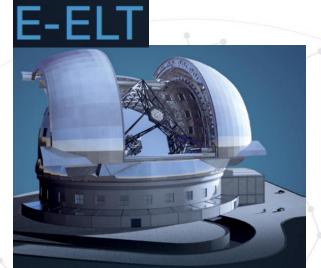


Euclid planned launch in 2020 To study the the origin of the accelerating Universe from L2 for 6 years



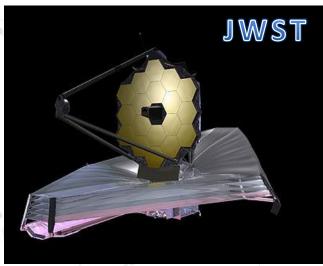
And many more to come ...

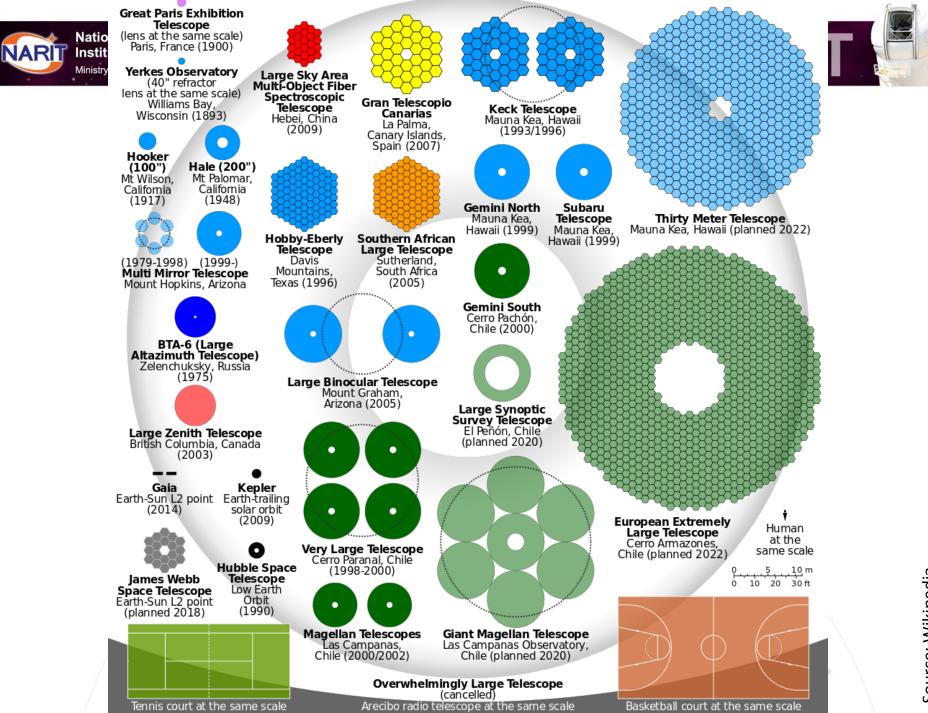




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Source: Wikipedia





National Astronomical Research Institute of Thailand

Established since 1 January 2009

under the Ministry of Science and Technology

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Key Sciences & Research Direction



1. Effects from Space to Earth and Mankind

2. Understanding the Physics of astronomical phenomena

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- 3. A Study of Planets and signs of lives outside the Solar System
- 4. The Understanding on the Origin of the Universe



Infrastructures of NARIT



1) The Thai National Observatory



2) Regional Observatories for the Public (in 5 provinces)

NARIT



3) The Astro Park



4) Thai Robotic Telescope Network (Thai Robonet)



The Thai National Telescope

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At Km. 44 on Doi Inthanon (the highest mountain in Thailand), Chiang Mai, Thailand, which is also renowned for the superb climate and tourist attraction. At 2,457 meters above the mean sea level WWW.NARIT.OR.TH





TNO inauguration by HRH Princess Maha Chakri Sirindhorn

on 22 January 2013

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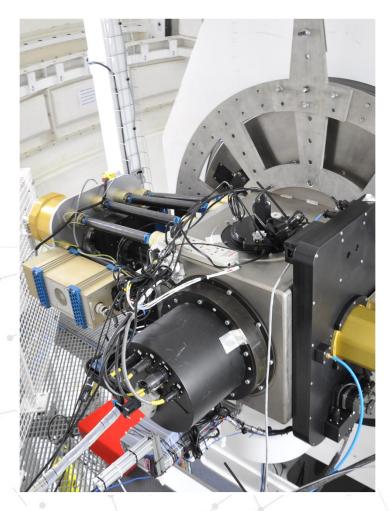
The 2.4-m Thai National Telescope (TNT)

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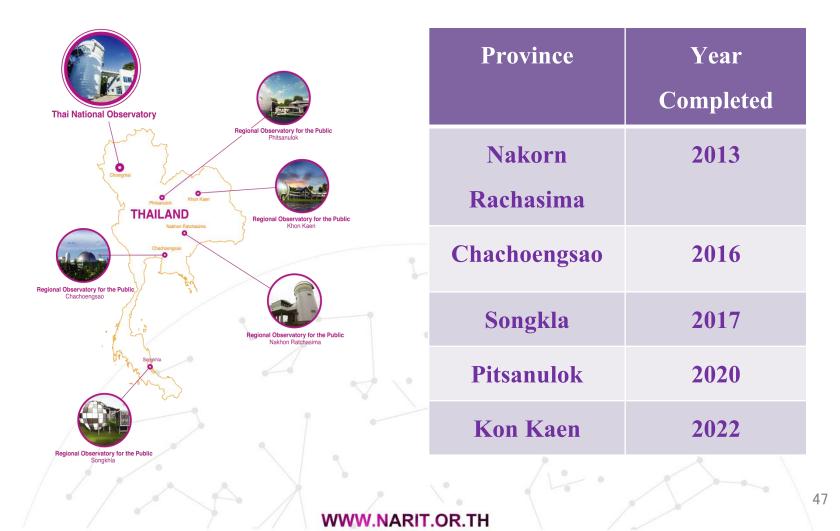
Major Instruments

- ULTRASPEC
- 4k x 4k CCD Camera
- Medium Resolution Spectrograph (MRES)





Regional Observatories for the Public







NARIT



Location: Don Kaew, Mae Rim, Chiang Mai (22 Acres)



Cerro Tololo

Inter-

American

Observatory,

Chile

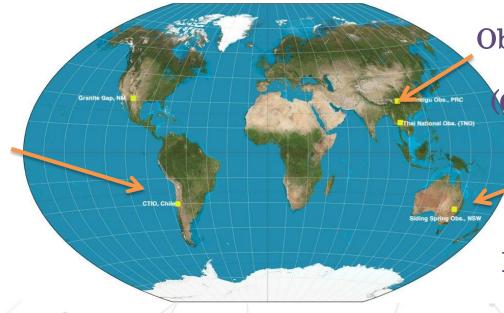
(Operated in

2013)



Thai Robotic Telescope Network

(Thai Robonet)



0.6-0.7 meter reflecting telescopes

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Gao Mei Gu Observatory, Lijiang, PR. China Operated in 2015)

Siding Spring Observatory, NSW, Australia



Thai Southern hemisphere telescope





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Location: CTIO, Chile Altitude : 2,207 meters above the sea level Telescope : 0.61 meter









Thai Robotic Telescope in Li Jiang, PR. Of China





NARIT

The 0.7-meter Reflecting

Telescope



Observatory at the Air Force Report Center, Doi Inthanon







National Astronomical Research Institute of Thailand (Public Organization) Ministry of Science and Technology of Thailand



NARIT High Performance Computer (HPC) cluster

NARIT







Development and test node "Pleione"

<u>Hardware</u>

- - Total 28 cores, RAM 160 GB (DDR3 RDIMM 1.6GHz)
 - 2x6 cores, 2.6GHz Intel Xeon E5-2630 v2
 - 2x8 cores, 2.0GHz Intel Xeon E5-2650
 - 1 CPU/node (due to Blade chassis power supply limit)
 - Storage14 TB (NAS)
 - Network switch Gigabit Ethernet

Performance

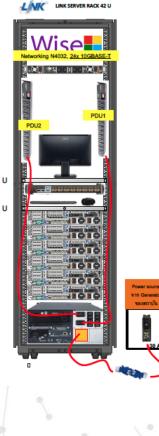
 R_{peak}= 380 GFLOPS, R_{max}=75-80% R_{peak} running Intel optimized High Performance Linpack (HPL)



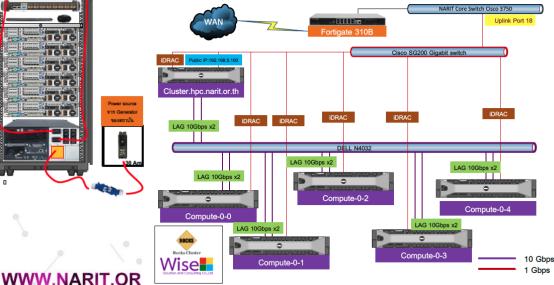
CASTOR

NARIT High Performance Computer (HPC) cluster

	NARIT HPC provide by DELL POWER EDGE R730
1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1 1	1.Dell Networking N4032, 24x 10GBASE-T Fixed Ports (1214 nn Port) 2x Power Supplies
	2.Cable Management
(REL)	3.DELL 18.5" HD Monitor with LED backligh
U	4.Fix Shelf
Tanuna	5.KVM Switch 8 Port -USB
u 1	6.USB Mouse/Keyboard
in the part of the second second second	7.Slide Shelf (เลื่อนเข้าออกใต้เพื่อให้ง่ายในการใช้งาน)
	8.DELL R730 = Computing Node 5
5	9.DELL R730 = Computing Node 4
	4 10.DELL R730 = Computing Node 3
3	11.DELL R730 = Computing Node 2
	2 12.DELL R730 = Computing Node 1
	13.DELL R730 = Management Node
	เว้นไว้ 2 U สำหรับ ติดตั้ง DELL R730 เพิ่ม 1 เครื่องในอนาคต
	14.CS SUA5000RMI5U APC Smart-UPS 5000VA 230V Rackmount/Tower 2



-Installed Feb-Mar 2015 -Testing and commissioning late Mar – April 2015 -Announced to Thai astronomical community during TNAM May 2015 - Open for community use in June/July 2015

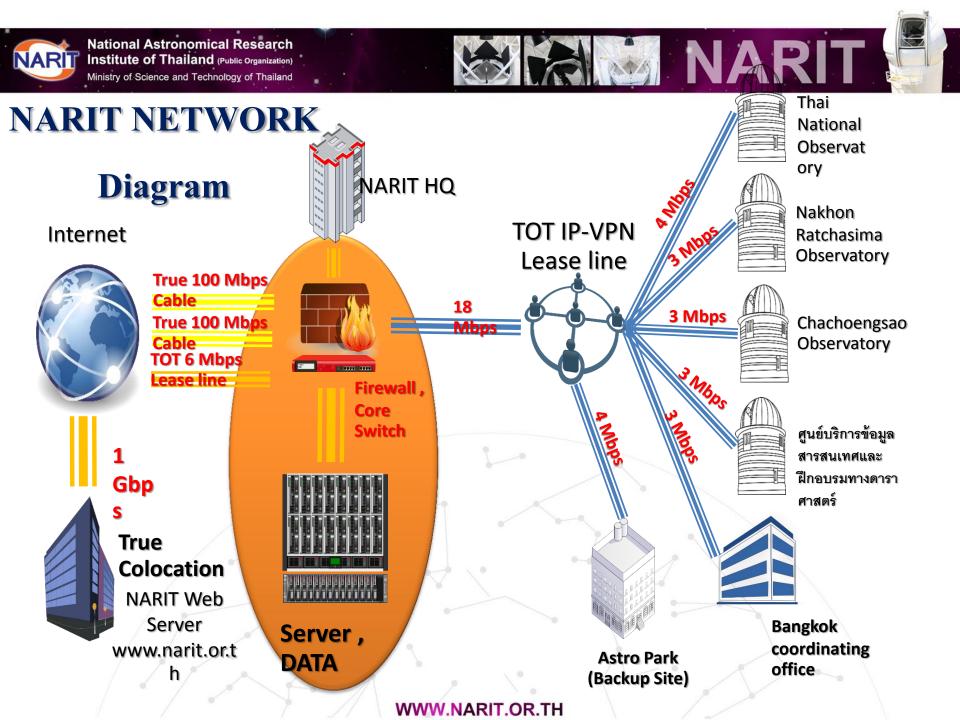




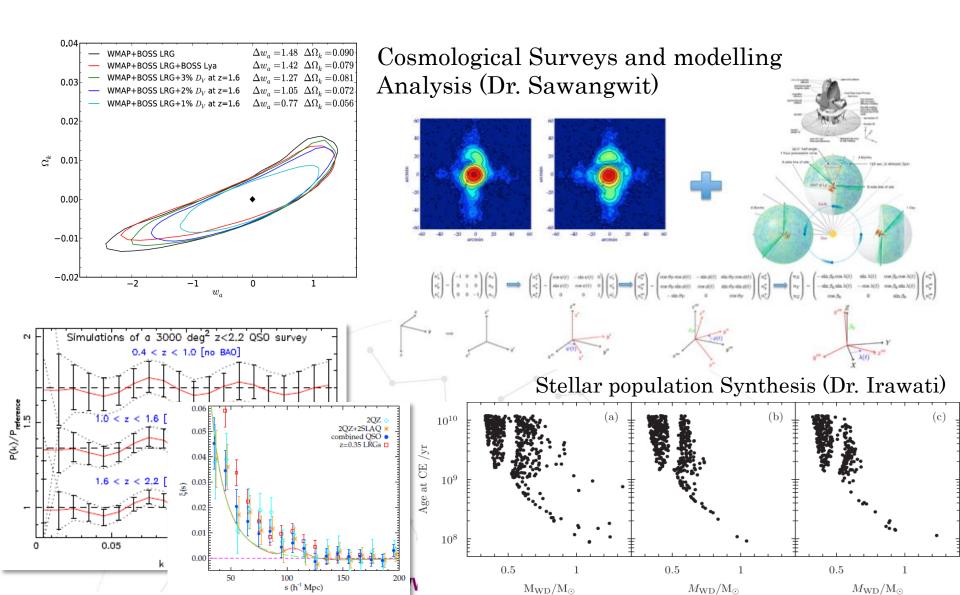
<u>Hardware</u>

- ♦ 1 Management node (12 cores, 2.4 GHz Intel Xeon E5-26xx v3, 32GB RAM)
- ♦ 5 compute nodes, rack servers
 - Total 80 cores (5 x 16 cores 2.6 GHz Intel Xeon E5-26xx v3)

- RAM 5 x 64 GB (320GB, 4GB per core) DDR4 RDIMM 2.13 GHz
- Dual-port 10 Gbps Ethernet, with teaming connections
- Each rack is compatible with up to 2 GPU cards upgrade
- ♦ Storage 7.2 TB SAS 10K rpm 6Gbps (RAID5, 2GB cache)
- ♦ Network Switch 10 Gigabit Ethernet
- ♦ Expected R_{max} = 2.26 TFLOPS

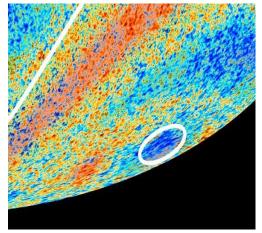






Na Ins Mini

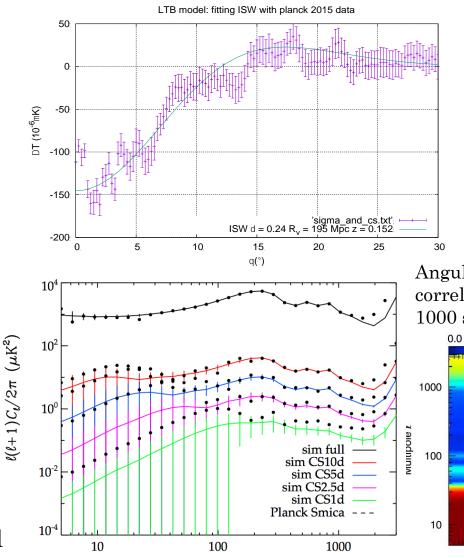
National Astronomical Research Institute of Thailand (Public Organization) Ministry of Science and Technology of Thailand



Planck CMB map



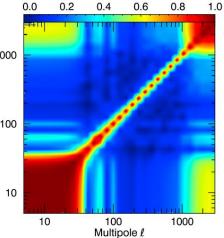
Example region of a void



Multipole l

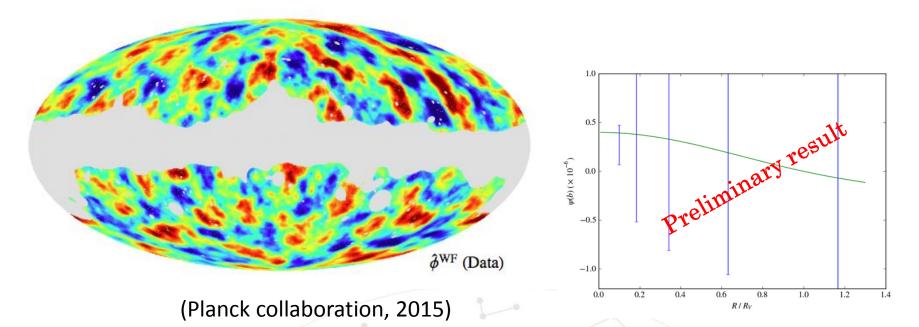
Temperature Profile fitted to a single Super-void model (Mr. Anut Sangkla, SUT)

Angular Power Spectrum & correlation matrix from 1000 simulations

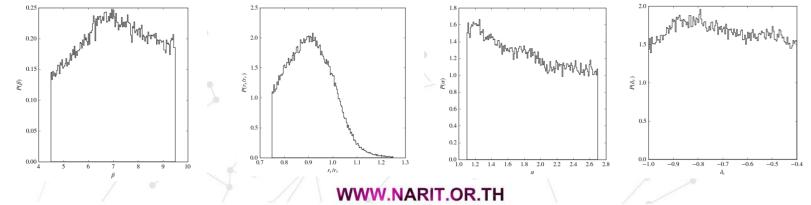




Gravitational lensing of the CMB by voids



Monte Carlo Markov Chain results (preliminary) running with NARIT HPC





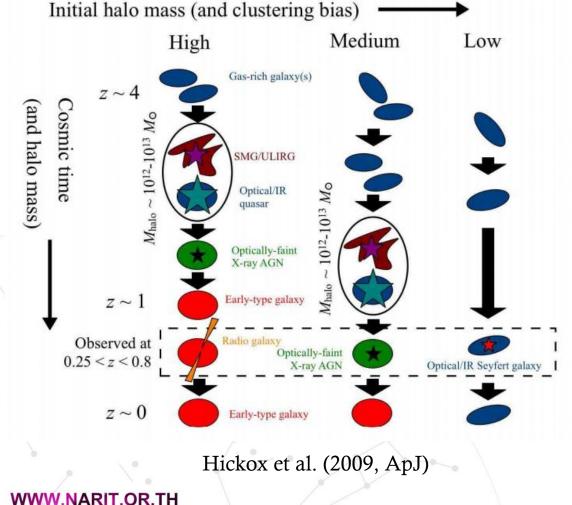
Galaxy evolution: AGN feedback

$SMBH-Host\ galaxy\ co\text{-}evolution$

* Stars/Early-type BCG N4889 Stars/Early-type non-BCG N3842 🖈 Stars/Late-type High 10^{10} Gas/Early-type BCG Gas/Early-type non-BCG A1836 Gas/Late-type $Z \sim$ Masers/Early-type IC145 (and halo mass Cosmic time Masers/Late-type $M_{\rm halo} \sim 10^{12} \cdot 10^{13} M_{\odot}$ 10^{9} $M_{BH} (M_{\odot})$ N4261 N3998 10⁸ 10⁷ $z \sim 1$ Observed at 0.25 < z < 0.810⁶ 80 100 200 300 400 60 σ (km/s) $z \sim 0$ McConnell +(2011, Nature)

Broad and simplified picture

MAI

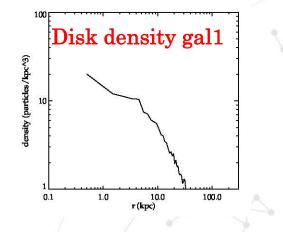


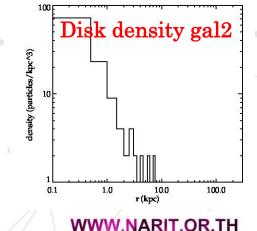


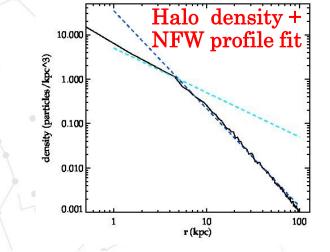
AGN feedback + interactions with IGM simulations



Images from HST (Keel et al, 2015, ApJ)



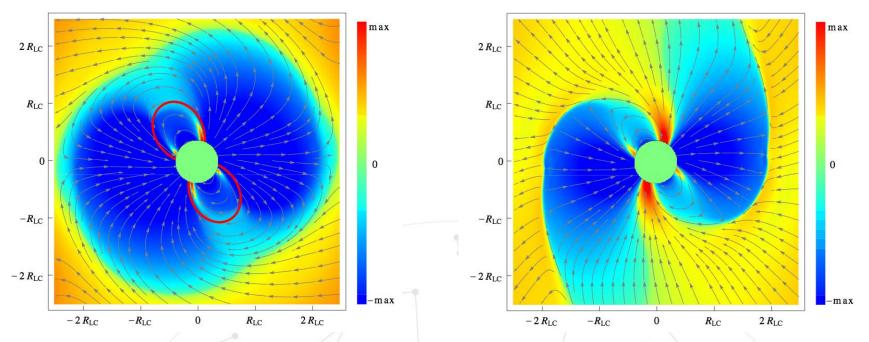






Intermittent and Mode-switching Pulsars

Jaroenjittichai, Phrudth (NARIT) Kramer, Michael (MPIfR, Germany)



Simulation of the magnetosphere in the off and on state (Li et al. 2012)

"Pulsar Search with PMPS dataset" search through 10TB of data to look For pulsar in the radio data

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