

# Neutrinoless Double Beta Decay: Theory



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SSI 2015

17/08/15

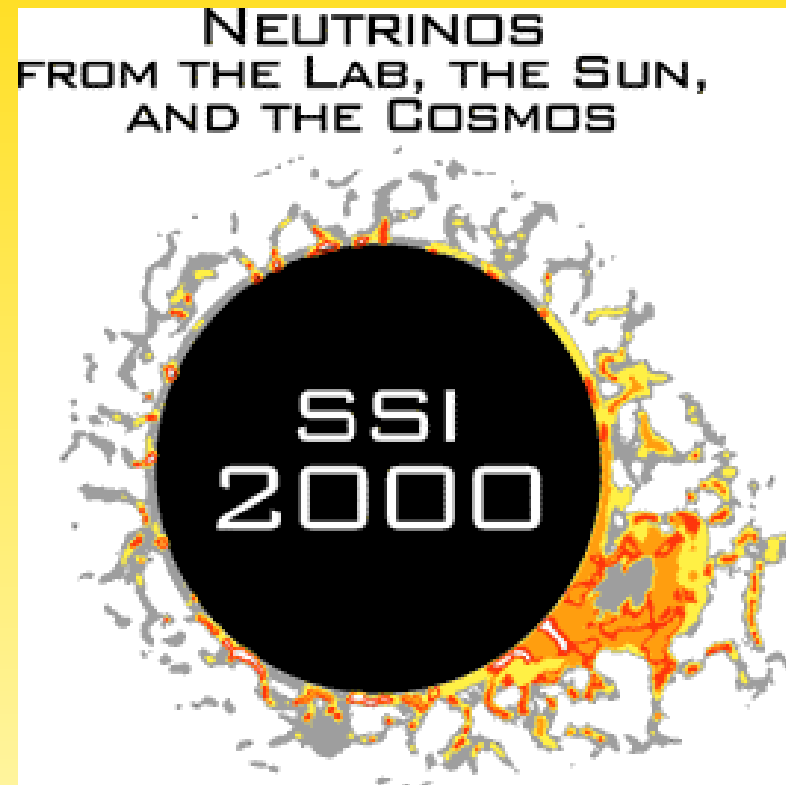
$$m_\nu = m_L - m_D^T M_R^{-1} m_D$$

**MANITOP**

Massive Neutrinos: Investigating their  
Theoretical Origin and Phenomenology



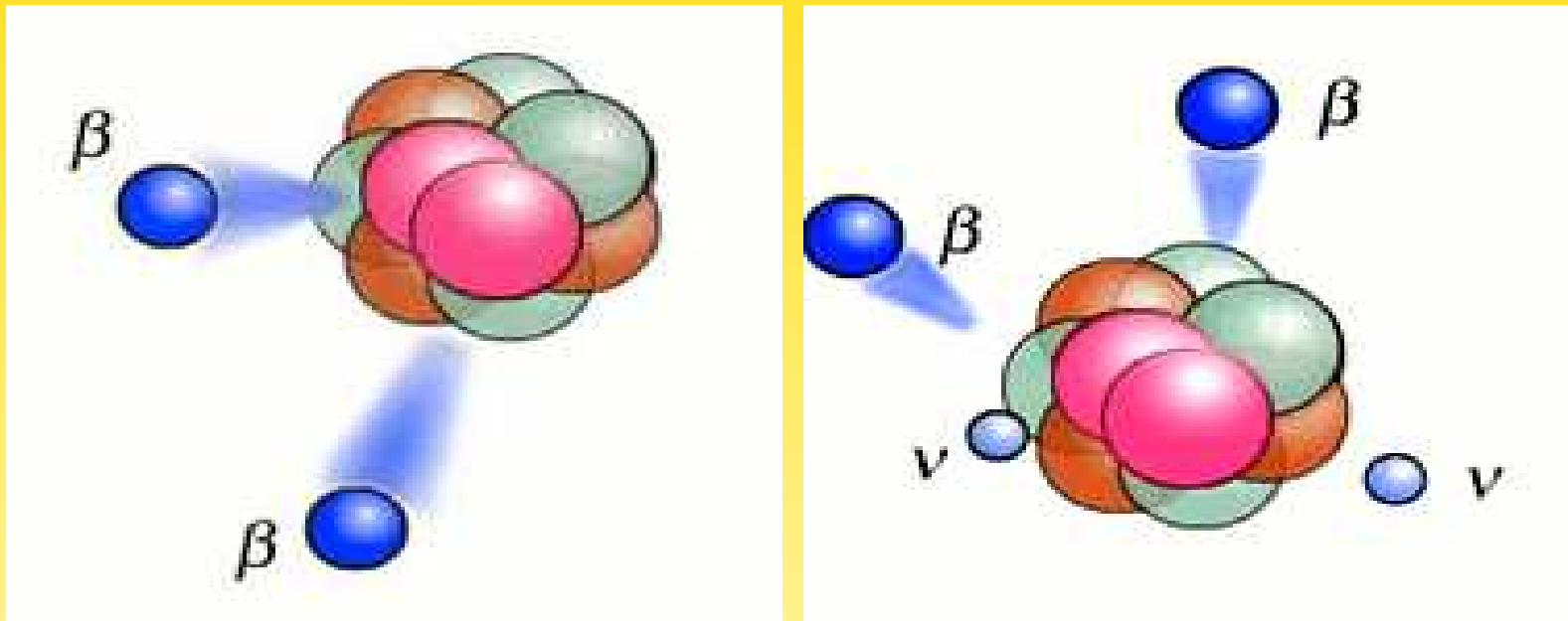
Neutrinos...



...still interesting...

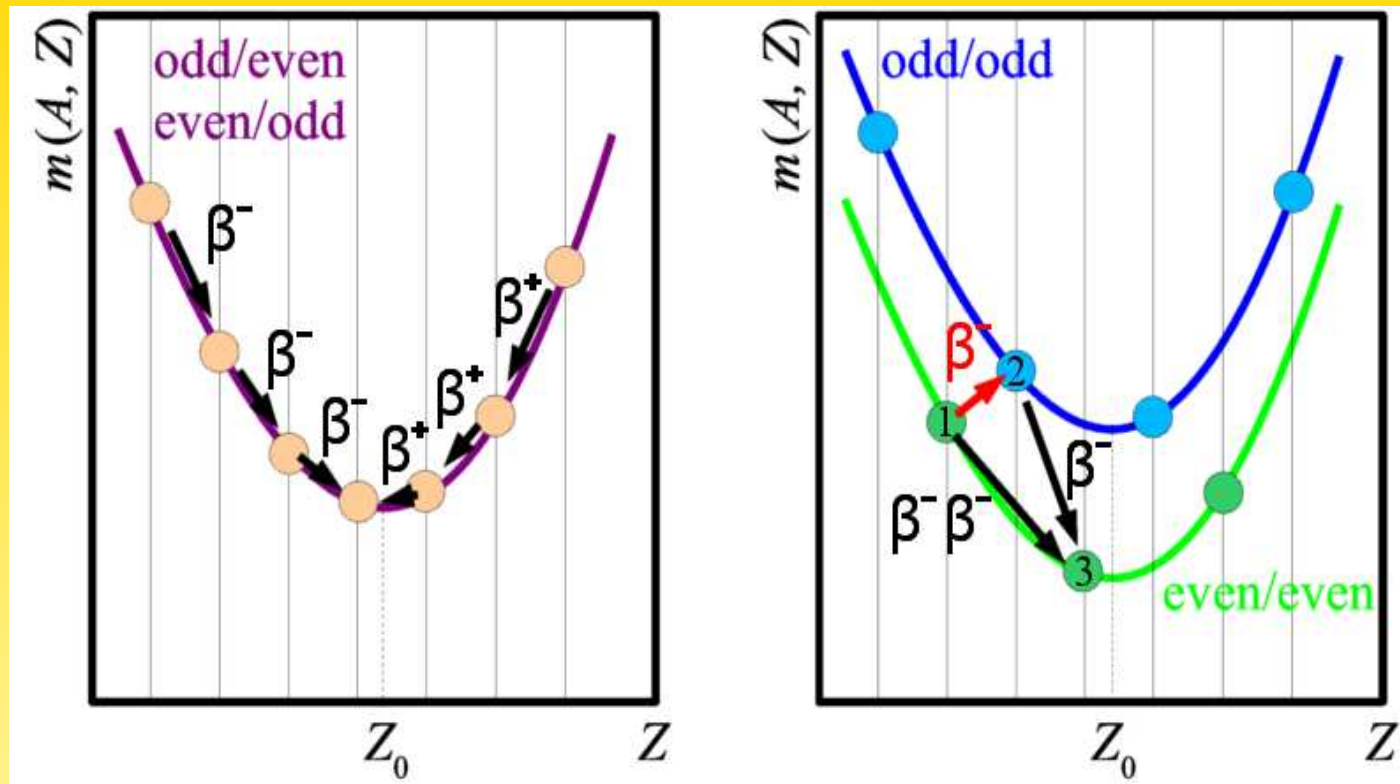
## What is Neutrinoless Double Beta Decay?

$$(A, Z) \rightarrow (A, Z + 2) + 2 e^{-} \quad (0\nu\beta\beta)$$



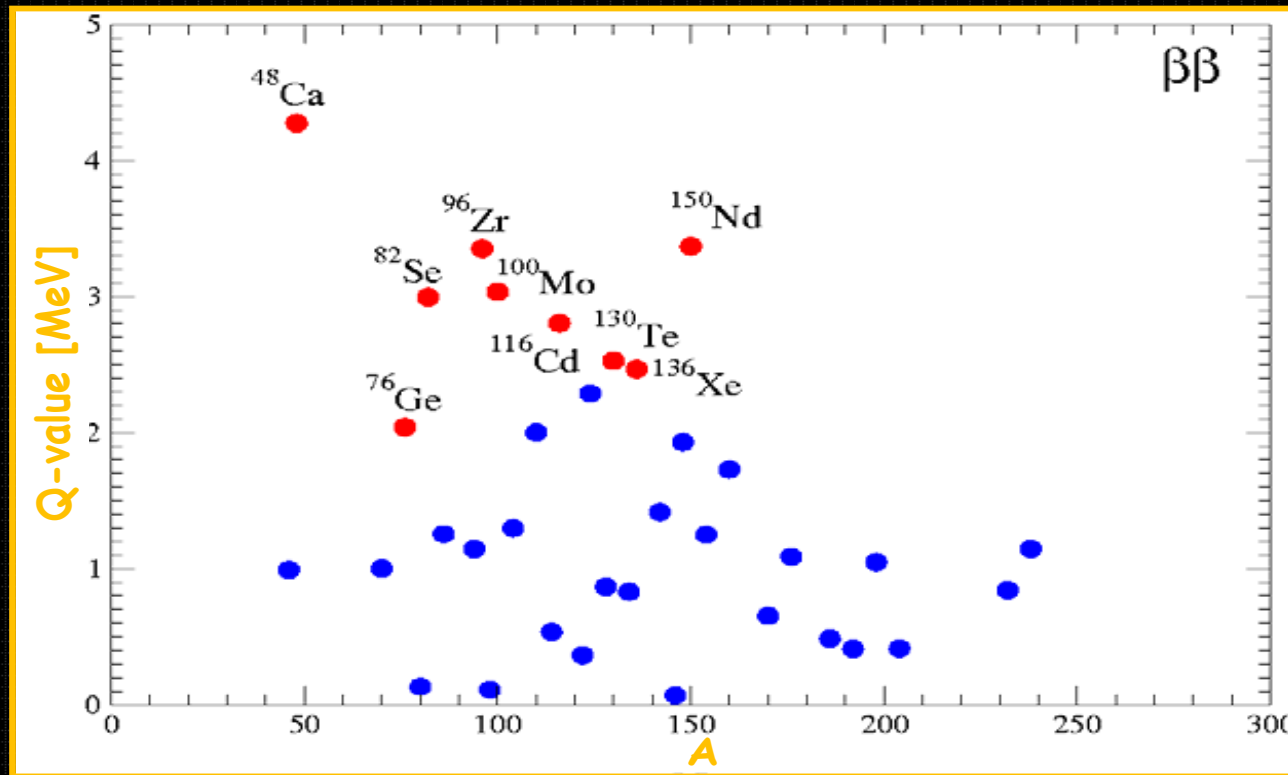
- second order in weak interaction,  $\Gamma \propto G_F^4 Q^5 \Rightarrow$  rare!
- not to be confused with  $(A, Z) \rightarrow (A, Z + 2) + 2 e^{-} + 2 \bar{\nu}_e \quad (2\nu\beta\beta)$   
( $\Gamma \propto G_F^4 Q^{11}$ , but occurs more often...)

Need to forbid single  $\beta$  decay:



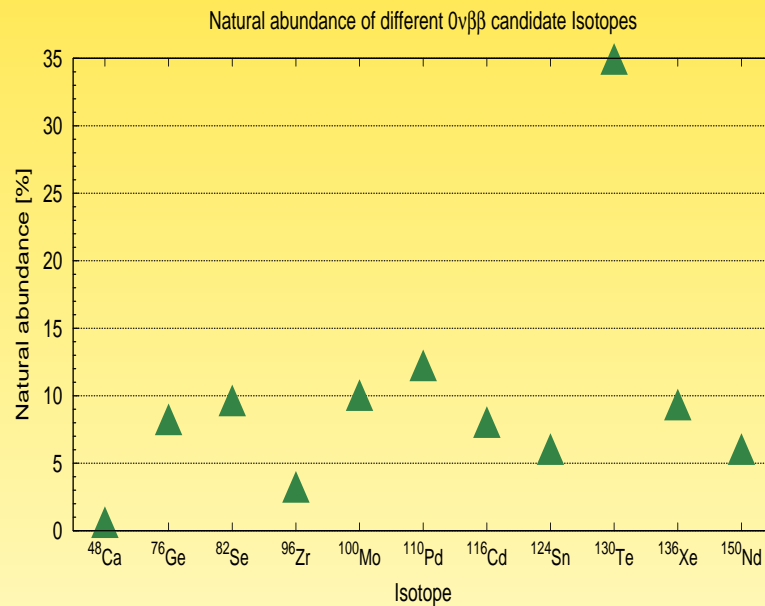
- $E_{\text{Bindung}} = E_{\text{Volumen}} - E_{\text{Oberfläche}} - E_{\text{Coulomb}} - E_{\text{Symmetrie}} \pm E_{\text{Paarbildung}}$
- $\Rightarrow$  even/even  $\rightarrow$  even/even
- either direct ( $0\nu\beta\beta$ ) or two simultaneous decays with virtual (energetically forbidden) intermediate state ( $2\nu\beta\beta$ )

How many nuclei in this condition?

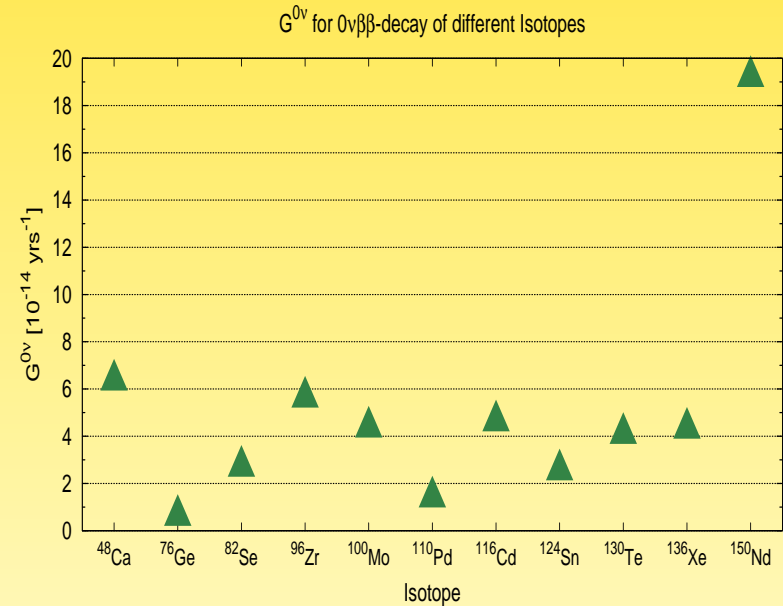


Slide by A. Giuliani, see lecture by Lisa Kaufman

- 35 candidate isotopes
- 10 are interesting:  $^{48}\text{Ca}$ ,  $^{76}\text{Ge}$ ,  $^{82}\text{Se}$ ,  $^{96}\text{Zr}$ ,  $^{100}\text{Mo}$ ,  $^{116}\text{Cd}$ ,  $^{124}\text{Sn}$ ,  $^{130}\text{Te}$ ,  $^{136}\text{Xe}$ ,  $^{150}\text{Nd}$
- $Q$ -value vs. natural abundance vs. reasonably priced enrichment vs. association with a well controlled experimental technique vs. ...  
 $\Rightarrow$  no superisotope



$$T_{1/2}^{0\nu} \propto 1/a$$



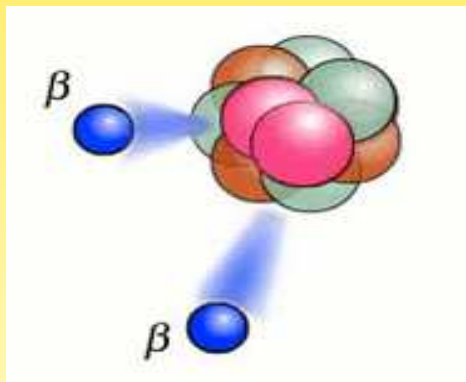
$$T_{1/2}^{0\nu} \propto Q^{-5}$$

## Neutrinoless Double Beta Decay

$$(A, Z) \rightarrow (A, Z + 2) + 2 e^{-} \quad (0\nu\beta\beta) \Rightarrow \text{Lepton Number Violation}$$

- **Standard Interpretation** (neutrino physics)
- **Non-Standard Interpretations** (BSM  $\neq$  neutrino physics)

Int. J. Mod. Phys. **E20**, 1833 (2011) [1106.1334]; J. Phys. **G39**, 124008  
(2012) [1206.2560]; 1507.00170



## Why should we probe Lepton Number Violation?

- $L$  and  $B$  accidentally conserved in SM
- effective theory:  $\mathcal{L} = \mathcal{L}_{\text{SM}} + \frac{1}{\Lambda} \mathcal{L}_{\text{LNV}} + \frac{1}{\Lambda^2} \mathcal{L}_{\text{LFV, BNV, LNV}} + \dots$
- baryogenesis:  $B$  is violated
- $B, L$  often connected in GUTs
- GUTs have seesaw and Majorana neutrinos
- (chiral anomalies:  $\partial_\mu J_{B,L}^\mu = c G_{\mu\nu} \tilde{G}^{\mu\nu} \neq 0$  with  $J_\mu^B = \sum \bar{q}_i \gamma_\mu q_i$  and  $J_\mu^L = \sum \bar{l}_i \gamma_\mu l_i$ )

⇒ Lepton Number Violation as important as Baryon Number Violation

$0\nu\beta\beta$  is **NOT** a neutrino mass experiment!!



## Upcoming/running experiments: exciting time!!

best limit was from 2001, improved 2012

Name	Isotope	Source = Detector; calorimetric with			Source $\neq$ Detector topology
		high $\Delta E$	low $\Delta E$	topology	
AMoRE	$^{100}\text{Mo}$	✓	–	–	–
CANDLES	$^{48}\text{Ca}$	–	✓	–	–
COBRA	$^{116}\text{Cd}$ (and $^{130}\text{Te}$ )	–	–	✓	–
CUORE	$^{130}\text{Te}$	✓	–	–	–
DCBA/MTD	$^{82}\text{Se}$ / $^{150}\text{Nd}$	–	–	–	✓
EXO	$^{136}\text{Xe}$	–	–	✓	–
GERDA	$^{76}\text{Ge}$	✓	–	–	–
CUPID	$^{82}\text{Se}$ / $^{100}\text{Mo}$ / $^{116}\text{Cd}$ / $^{130}\text{Te}$	✓	–	–	–
KamLAND-Zen	$^{136}\text{Xe}$	–	✓	–	–
LUCIFER	$^{82}\text{Se}$ / $^{100}\text{Mo}$ / $^{130}\text{Te}$	✓	–	–	–
LUMINEU	$^{100}\text{Mo}$	✓	–	–	–
MAJORANA	$^{76}\text{Ge}$	✓	–	–	–
MOON	$^{82}\text{Se}$ / $^{100}\text{Mo}$ / $^{150}\text{Nd}$	–	–	–	✓
NEXT	$^{136}\text{Xe}$	–	–	✓	–
SNO+	$^{130}\text{Te}$	–	✓	–	–
SuperNEMO	$^{82}\text{Se}$ / $^{150}\text{Nd}$	–	–	–	✓
XMASS	$^{136}\text{Xe}$	–	✓	–	–

see lecture by Lisa Kaufman

## Interpretation of Experiments

Master formula:

$$\Gamma^{0\nu} = G_x(Q, Z) |\mathcal{M}_x(A, Z) \eta_x|^2$$

- $G_x(Q, Z)$ : phase space factor
- $\mathcal{M}_x(A, Z)$ : nuclear physics
- $\eta_x$ : particle physics

## Interpretation of Experiments

Master formula:

$$\Gamma^{0\nu} = G_x(Q, Z) |\mathcal{M}_x(A, Z) \eta_x|^2$$

- $G_x(Q, Z)$ : phase space factor; **calculable**
- $\mathcal{M}_x(A, Z)$ : nuclear physics; **problematic**
- $\eta_x$ : particle physics; **interesting**

## Interpretation of Experiments

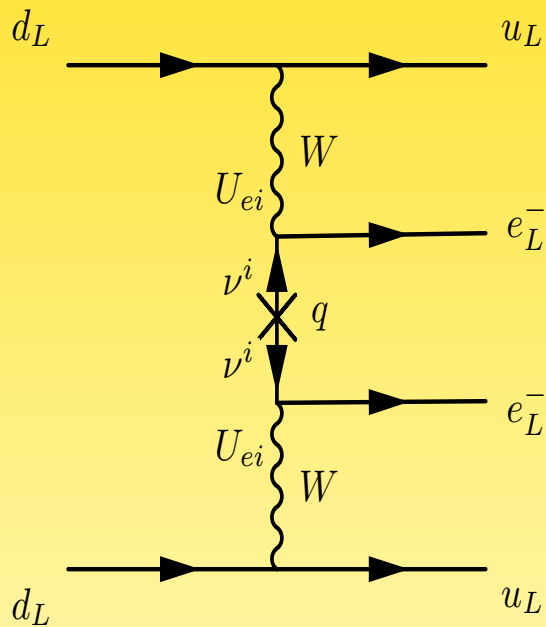
Master formula:

$$\Gamma^{0\nu} = G_x(Q, Z) |\mathcal{M}_x(A, Z) \eta_x|^2$$

- $G_x(Q, Z)$ : phase space factor, typically  $Q^5$
- $\mathcal{M}_x(A, Z)$ : nuclear physics; factor 2-3 uncertainty
- $\eta_x$ : particle physics; many possibilities

## Standard Interpretation

Neutrinoless Double Beta Decay is mediated by light and massive Majorana neutrinos (the ones which oscillate) and all other mechanisms potentially leading to  $0\nu\beta\beta$  give negligible or no contribution



prediction of  $(100 - \epsilon)\%$  of all neutrino mass mechanisms

(lecture by Andre De Gouvea)

## Paths to Neutrino Mass

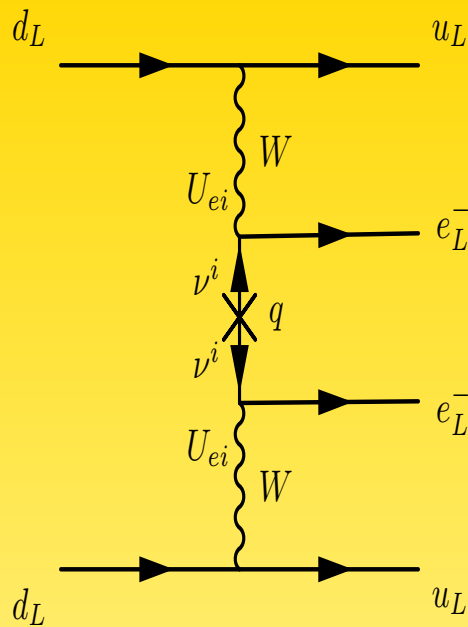
approach	ingredient	quantum number of messenger	$\mathcal{L}$	$m_\nu$	scale
“SM” (Dirac mass)	RH $\nu$	$N_R \sim (1, 0)$	$h \overline{N}_R \Phi L$	$h v$	$h = \mathcal{O}(10^{-12})$
“effective” (dim 5 operator)	new scale + LNV	–	$h \overline{L}^c \Phi \Phi L$	$\frac{h v^2}{\Lambda}$	$\Lambda = 10^{14}$ GeV
“direct” (type II seesaw)	Higgs triplet + LNV	$\Delta \sim (3, -2)$	$h \overline{L}^c \Delta L + \mu \Phi \Phi \Delta$	$h v_T$	$\Lambda = \frac{1}{h \mu} M_\Delta^2$
“indirect 1” (type I seesaw)	RH $\nu$ + LNV	$N_R \sim (1, 0)$	$h \overline{N}_R \Phi L + \overline{N}_R M_R N_R^c$	$\frac{(h v)^2}{M_R}$	$\Lambda = \frac{1}{h} M_R$
“indirect 2” (type III seesaw)	fermion triplets + LNV	$\Sigma \sim (3, 0)$	$h \overline{\Sigma} L \Phi + \text{Tr} \overline{\Sigma} M_\Sigma \Sigma$	$\frac{(h v)^2}{M_\Sigma}$	$\Lambda = \frac{1}{h} M_\Sigma$

plus seesaw variants (linear, double, inverse, . . .)

plus radiative mechanisms

plus extra dimensions

plusplus



- $U_{ei}^2$  from charged current
- $m_i/E_i$  from spin-flip and *if neutrinos are Majorana particles*

amplitude proportional to coherent sum (“effective mass”)

$$|m_{ee}| = \left| \sum U_{ei}^2 m_i \right|$$

$m/E \simeq \text{eV}/100 \text{ MeV}$  is tiny: only  $N_A$  can save the day!

## Dirac vs. Majorana

in  $V - A$  theories: observable difference always suppressed by  $(m/E)^2$

- In terms of degrees of freedom (helicity and particle/antiparticle):  
 $\nu_D = (\nu_\uparrow, \nu_\downarrow, \bar{\nu}_\uparrow, \bar{\nu}_\downarrow)$  versus  $\nu_M = (\nu_\uparrow, \nu_\downarrow)$
- weak interactions act on chirality (left-/right-handed)
- chirality is not a good quantum number ("spin flip"):  $L = \downarrow + \frac{m}{E} \uparrow$
- Dirac:
  - what we produce from a  $W^-$  is  $\ell^- (\bar{\nu}_\uparrow + \frac{m}{E} \bar{\nu}_\downarrow)$
  - the  $\bar{\nu}_\downarrow$  CANNOT interact with another  $W^-$  to generate another  $\ell^-$ :  
 $\Delta L = 0$
- Majorana:
  - what we produce from a  $W^-$  is  $\ell^- (\nu_\uparrow + \frac{m}{E} \nu_\downarrow)$
  - the  $\nu_\downarrow$  CAN interact with another  $W^-$  to generate another  $\ell^-$ :  $\Delta L = 2$   
 $\Rightarrow$  amplitude  $\propto (m/E) \Rightarrow$  probability  $\propto (m/E)^2$



## Dirac vs. Majorana

- Z-decay:

$$\frac{\Gamma(Z \rightarrow \nu_D \nu_D)}{\Gamma(Z \rightarrow \nu_M \nu_M)} \simeq 1 - 3 \frac{m_\nu^2}{m_Z^2}$$

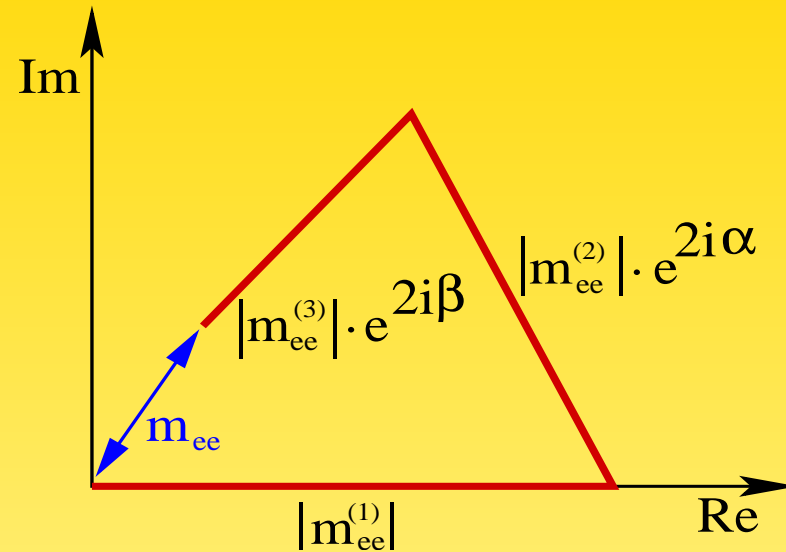
- Meson decays

$$\text{BR}(K^+ \rightarrow \pi^- e^+ \mu^+) \propto |m_{e\mu}|^2 = \left| \sum U_{ei} U_{\mu i} m_i \right|^2 \sim 10^{-30} \left( \frac{|m_{e\mu}|}{\text{eV}} \right)^2$$

- neutrino-antineutrino oscillations

$$P(\nu_\alpha \rightarrow \bar{\nu}_\beta) = \frac{1}{E^2} \left| \sum_{i,j} U_{\alpha j} U_{\beta j} U_{\alpha i}^* U_{\beta i}^* m_i m_j e^{-i(E_j - E_i)t} \right|$$

## The effective mass



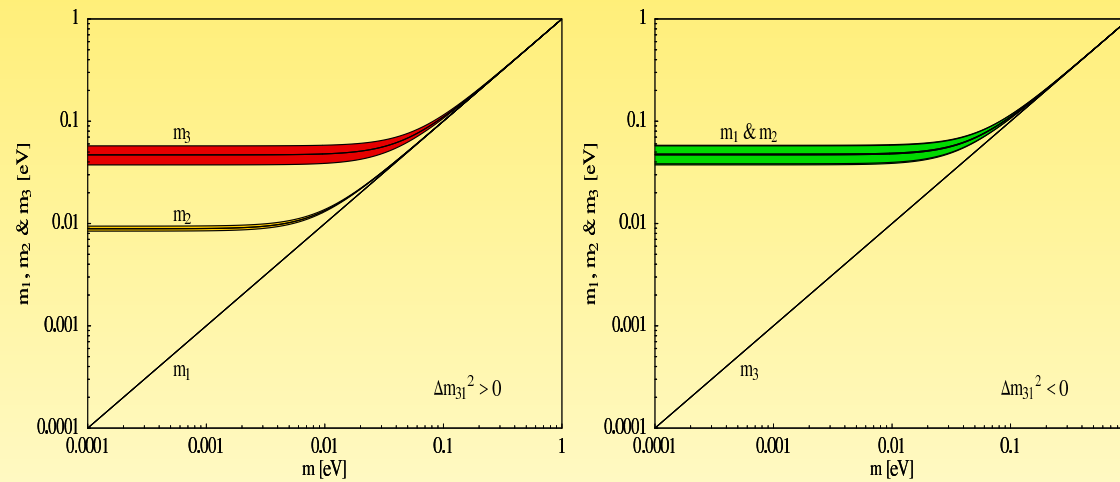
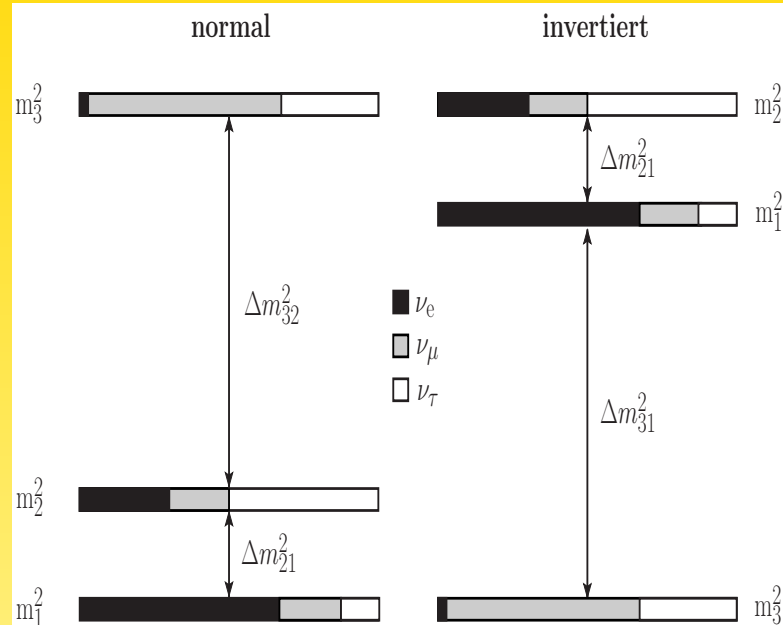
Amplitude proportional to coherent sum (“effective mass”):

$$|m_{ee}| \equiv \left| \sum U_{ei}^2 m_i \right| = \left| |U_{e1}|^2 m_1 + |U_{e2}|^2 m_2 e^{2i\alpha} + |U_{e3}|^2 m_3 e^{2i\beta} \right|$$

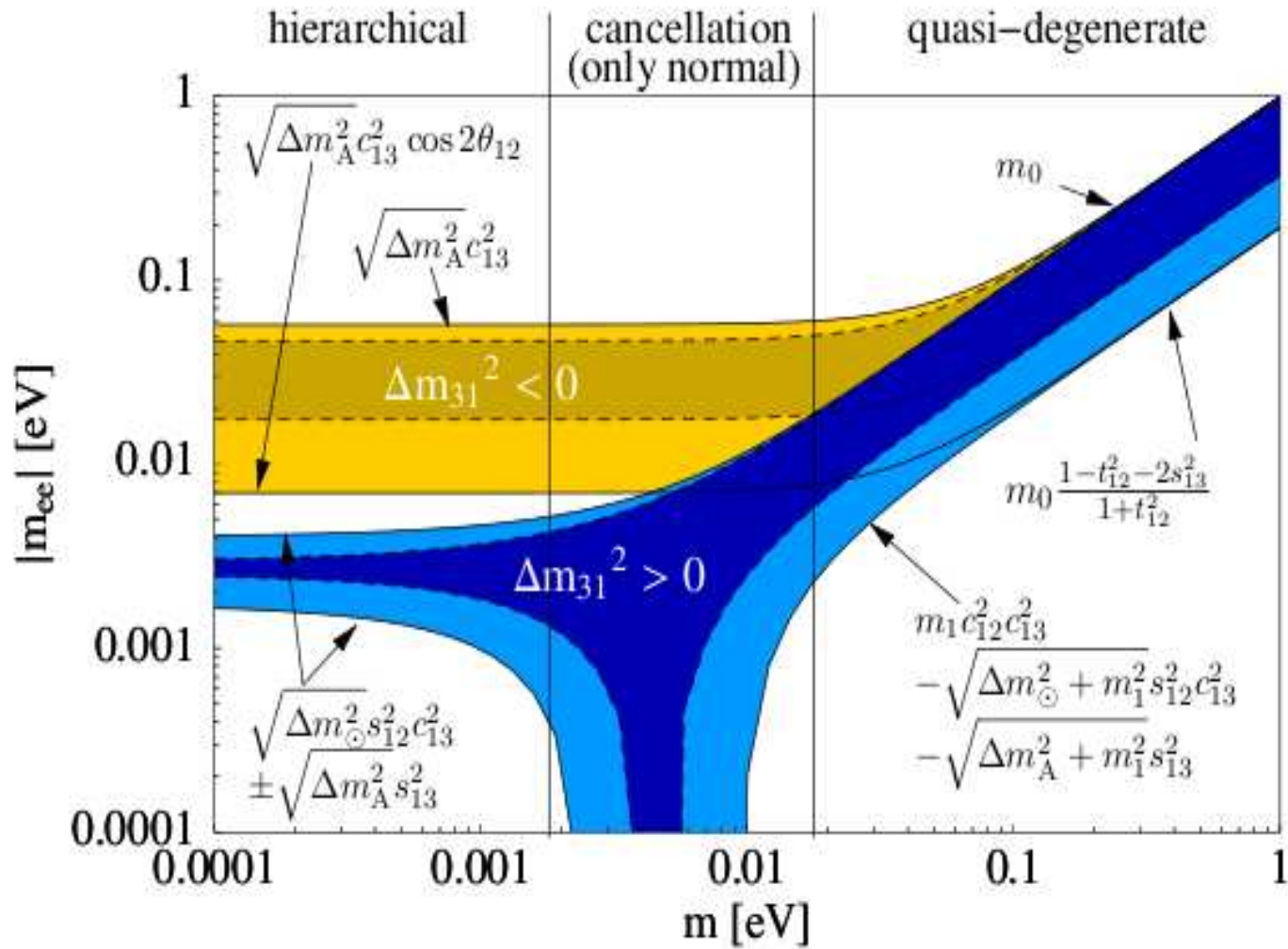
$$= f(\theta_{12}, |U_{e3}|, m_i, \text{sgn}(\Delta m_A^2), \alpha, \beta)$$

7 out of 9 parameters of neutrino physics!

Fix known things, vary unknown things...

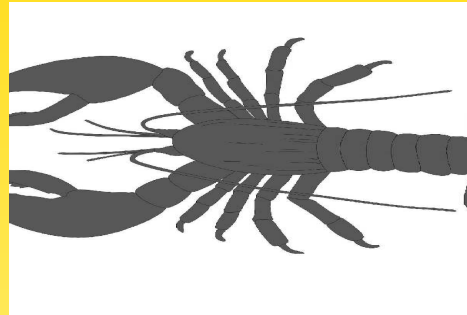
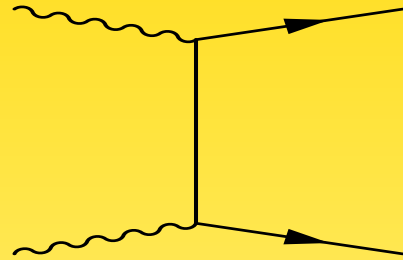


## The usual plot

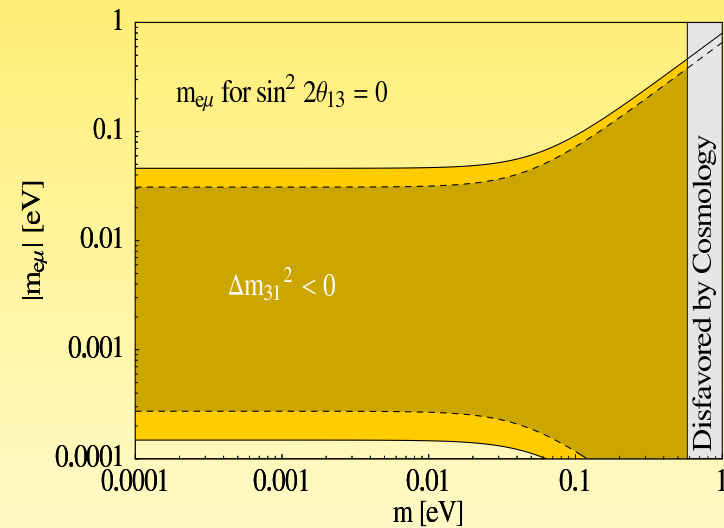
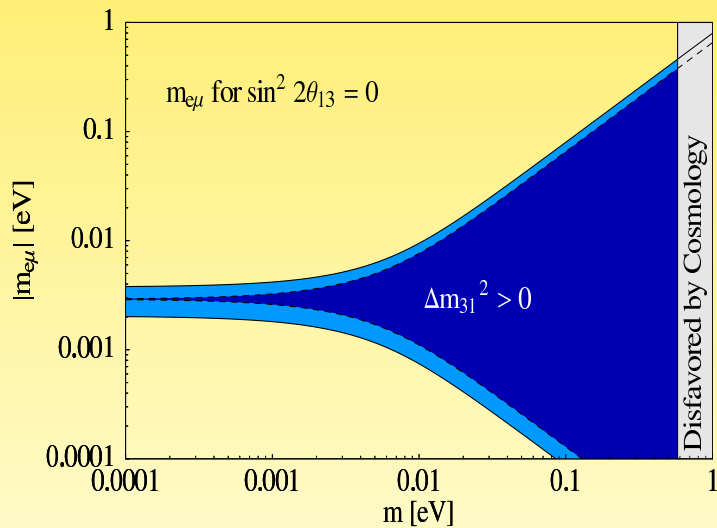


# Alternative processes

The lobster:

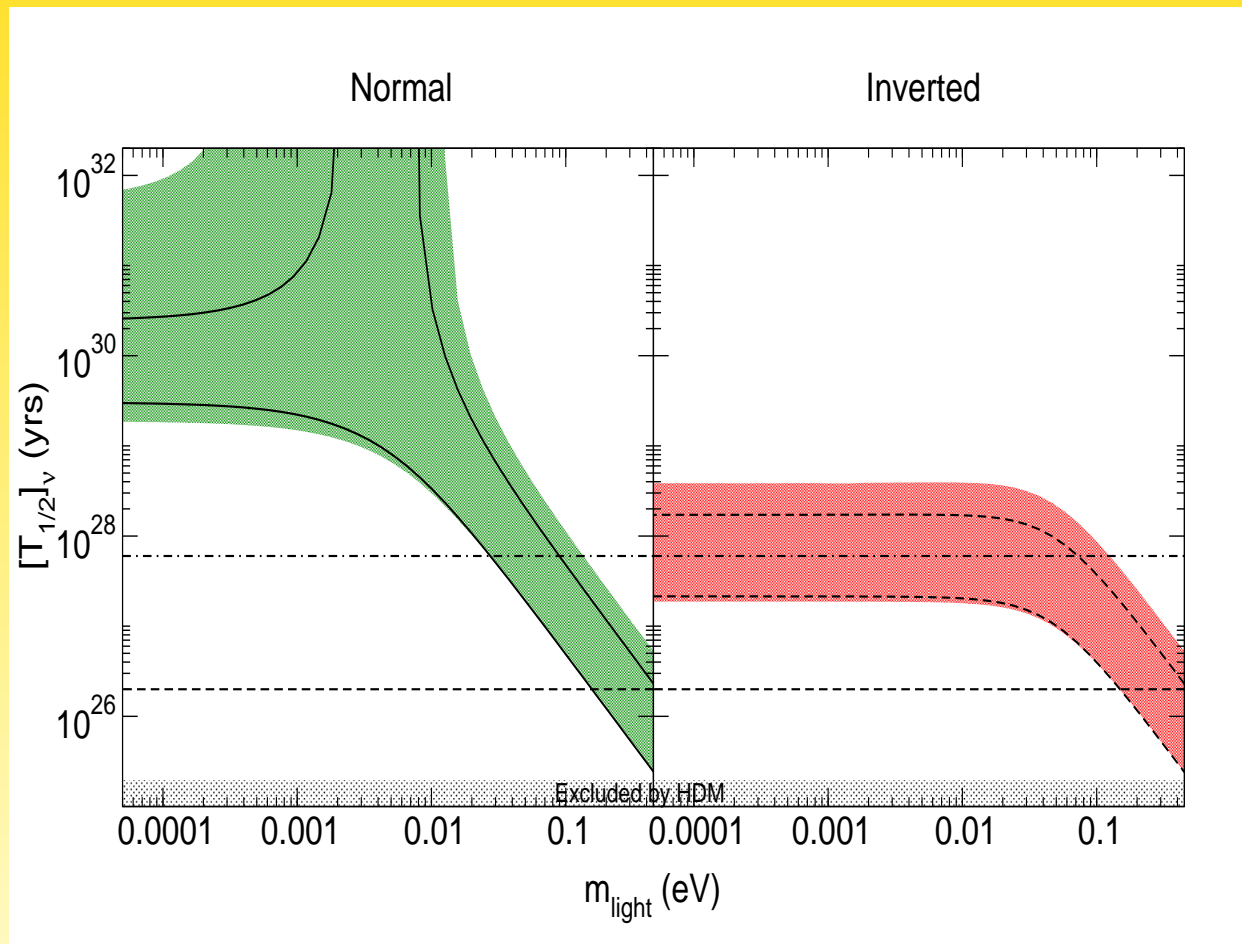


$$\text{BR}(K^+ \rightarrow \pi^- e^+ \mu^+) \propto |m_{e\mu}|^2 = \left| \sum U_{ei} U_{\mu i} m_i \right|^2 \sim 10^{-30} \left( \frac{|m_{e\mu}|}{\text{eV}} \right)^2$$



# The usual plot

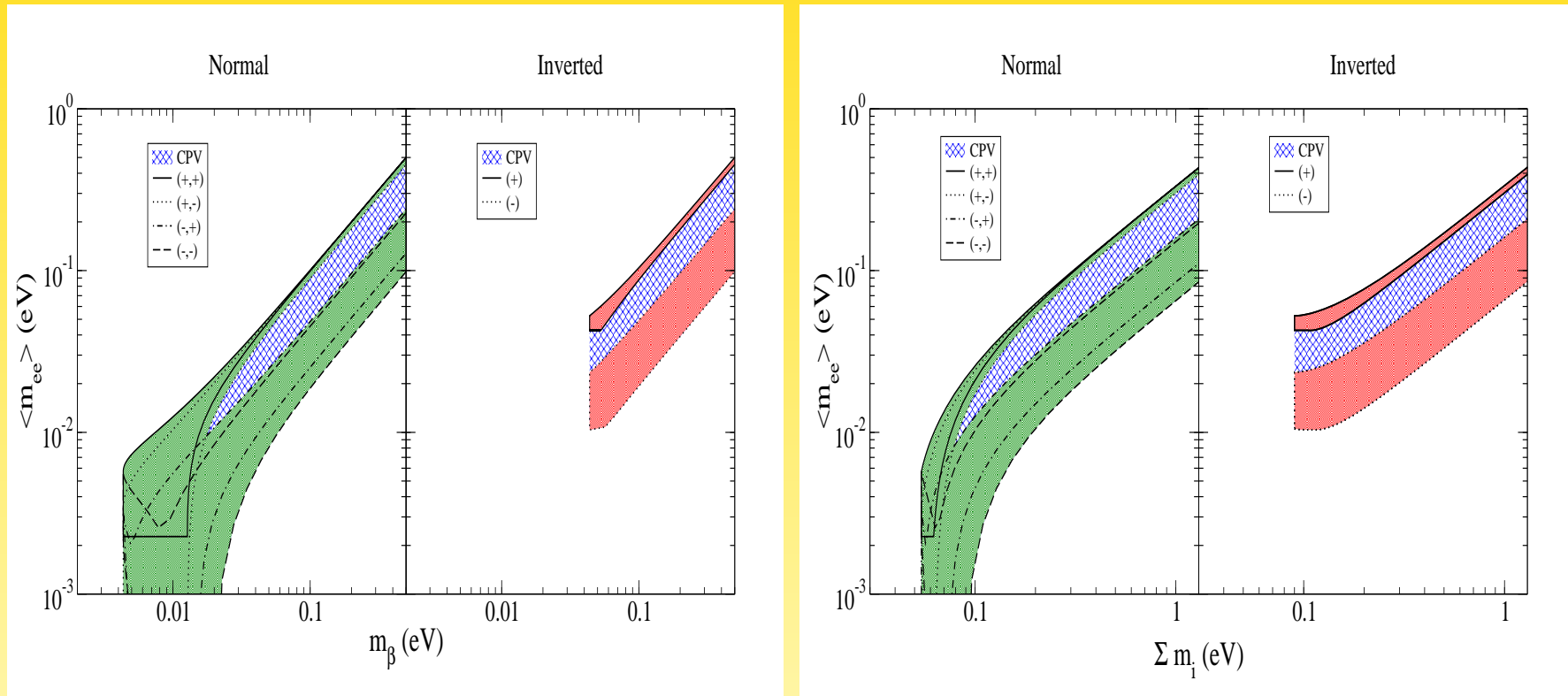
(life-time instead of  $|m_{ee}|$ )



## Which mass ordering with which life-time?

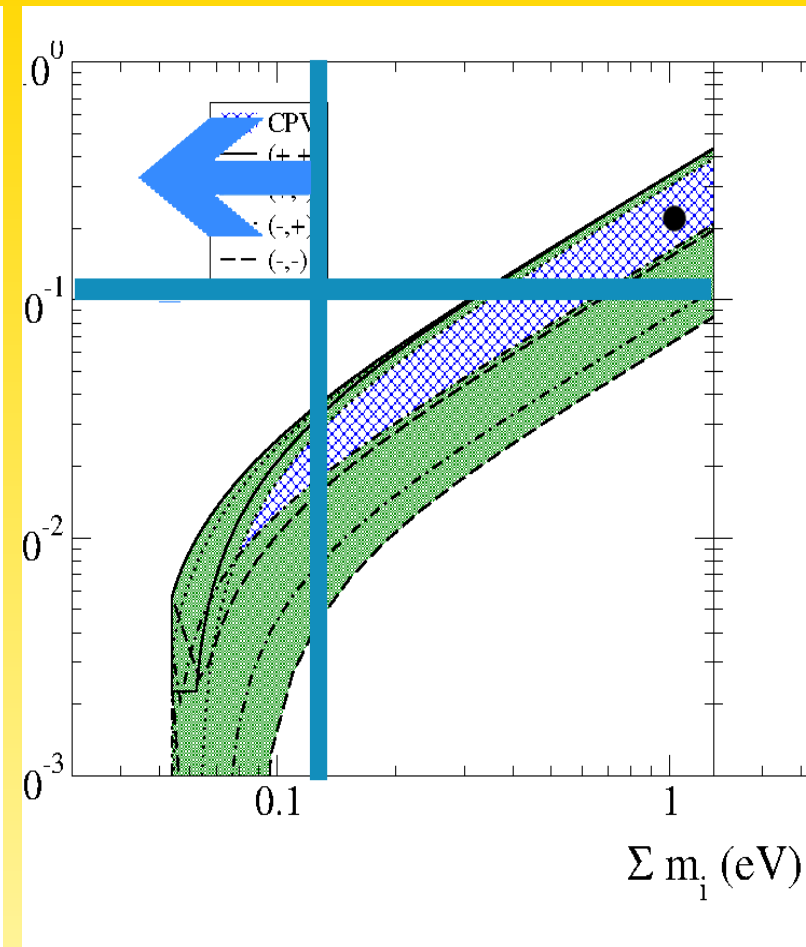
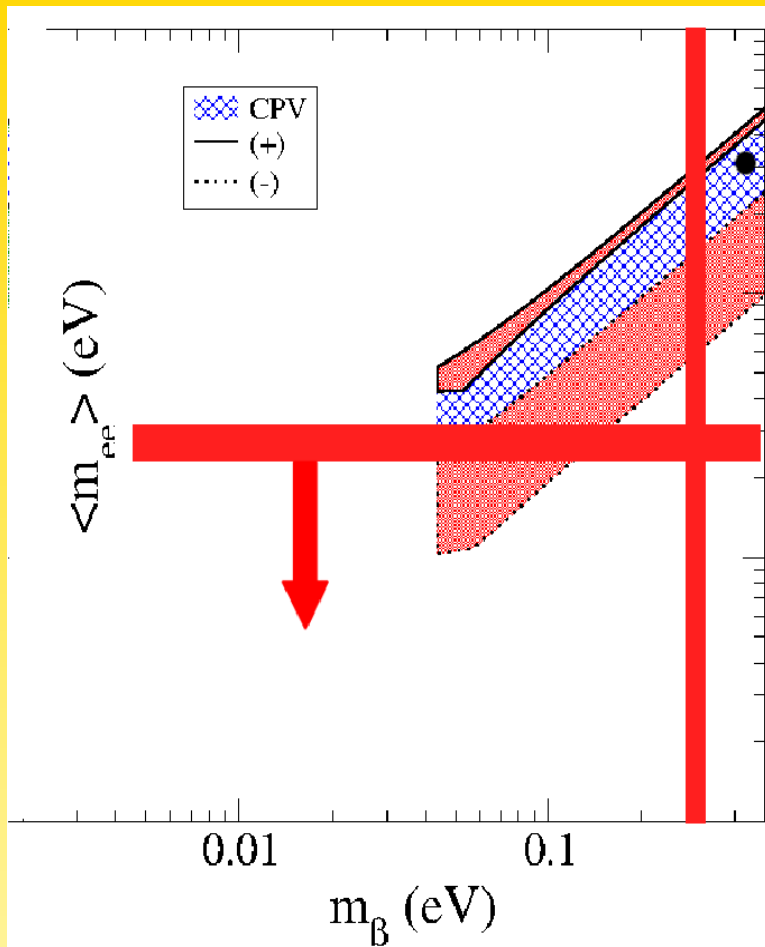
	$\Sigma$	$m_\beta$	$ m_{ee} $
NH	$\sqrt{\Delta m_A^2}$ $\simeq 0.05 \text{ eV}$	$\sqrt{\Delta m_\odot^2 +  U_{e3} ^2 \Delta m_A^2}$ $\simeq 0.01 \text{ eV}$	$\left  \sqrt{\Delta m_\odot^2 +  U_{e3} ^2 \Delta m_A^2} e^{2i(\alpha-\beta)} \right $ $\sim 0.003 \text{ eV} \Rightarrow T_{1/2}^{0\nu} \gtrsim 10^{28-29} \text{ yrs}$
IH	$2\sqrt{\Delta m_A^2}$ $\simeq 0.1 \text{ eV}$	$\sqrt{\Delta m_A^2}$ $\simeq 0.05 \text{ eV}$	$\sqrt{\Delta m_A^2} \sqrt{1 - \sin^2 2\theta_{12} \sin^2 \alpha}$ $\sim 0.03 \text{ eV} \Rightarrow T_{1/2}^{0\nu} \gtrsim 10^{26-27} \text{ yrs}$
QD	$3m_0$	$m_0$	$m_0 \sqrt{1 - \sin^2 2\theta_{12} \sin^2 \alpha}$ $\gtrsim 0.1 \text{ eV} \Rightarrow T_{1/2}^{0\nu} \gtrsim 10^{25-26} \text{ yrs}$

## Plot against other observables



**Complementarity** of  $|m_{ee}| = U_{ei}^2 m_i$  ,  $m_\beta = \sqrt{|U_{ei}|^2 m_i^2}$  and  $\Sigma = \sum m_i$





CP violation!

Dirac neutrinos!

something else does  $0\nu\beta\beta$ !

## Neutrino Mass

$$m(\text{heaviest}) \gtrsim \sqrt{|m_3^2 - m_1^2|} \simeq 0.05 \text{ eV}$$

3 **complementary** methods to measure neutrino mass:

Method	observable	now [eV]	near [eV]	far [eV]	pro	con
Kurie	$\sqrt{\sum  U_{ei} ^2 m_i^2}$	2.3	0.2	0.1	model-indep.; theo. clean	final?; worst
Cosmo.	$\sum m_i$	0.7	0.3	0.05	best; NH/IH	systemat.; model-dep.
$0\nu\beta\beta$	$ \sum U_{ei}^2 m_i $	0.3	0.1	0.05	fundament.; NH/IH	model-dep.; theo. dirty

see lectures by Alessandro Melchiorri and Joe Formaggio

## Recent Results

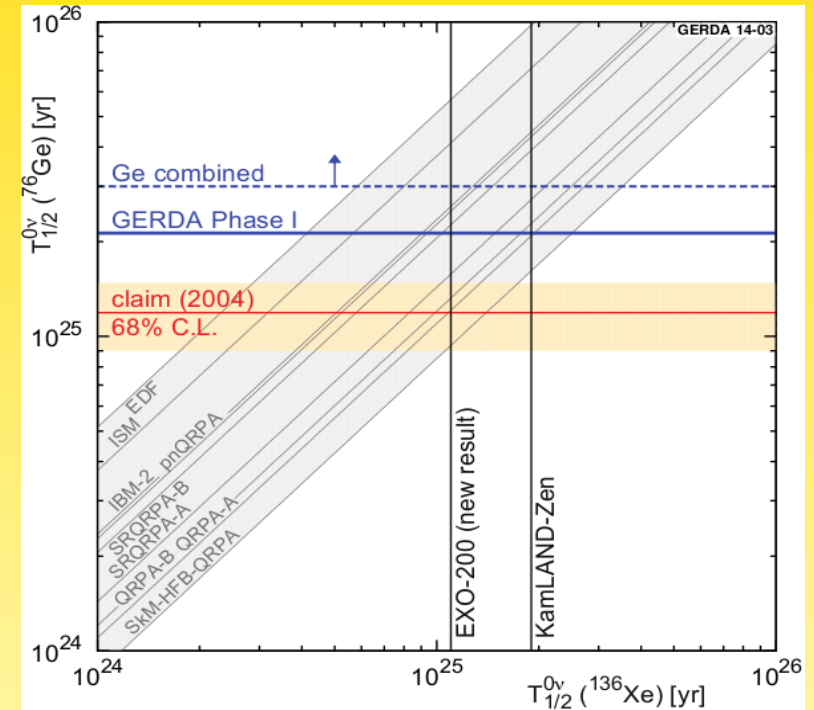
- $^{76}\text{Ge}$ :
  - GERDA:  $T_{1/2} > 2.1 \times 10^{25}$  yrs
  - GERDA + IGEX + HDM:  $T_{1/2} > 3.0 \times 10^{25}$  yrs
- $^{136}\text{Xe}$ :
  - EXO-200:  $T_{1/2} > 1.1 \times 10^{25}$  yrs (first run with less exposure:  $T_{1/2} > 1.6 \times 10^{25}$  yrs. . .)
  - KamLAND-Zen:  $T_{1/2} > 2.6 \times 10^{25}$  yrs

Xe-limit is stronger than Ge-limit when:

$$T_{\text{Xe}} > T_{\text{Ge}} \frac{G_{\text{Ge}}}{G_{\text{Xe}}} \left| \frac{\mathcal{M}_{\text{Ge}}}{\mathcal{M}_{\text{Xe}}} \right|^2 \text{ yrs}$$

## Current Limits on $|m_{ee}|$

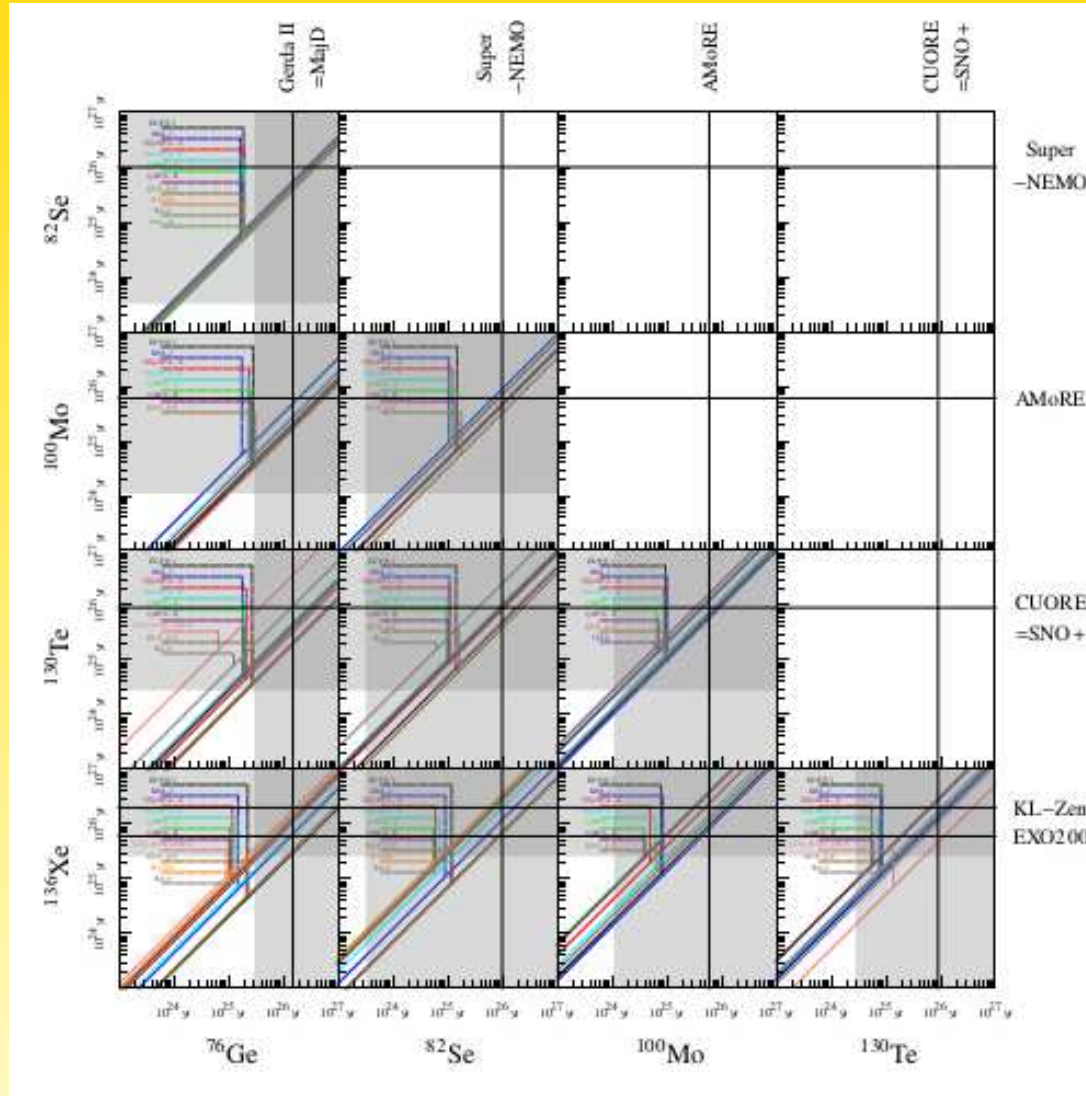
NME	$^{76}\text{Ge}$		$^{136}\text{Xe}$	
	GERDA	comb	KLZ	comb
EDF(U)	0.32	0.27	0.13	—
ISM(U)	0.52	0.44	0.24	—
IBM-2	0.27	0.23	0.16	—
pnQRPA(U)	0.28	0.24	0.17	—
SRQRPA-A	0.31	0.26	0.23	—
QRPA-A	0.28	0.24	0.25	—
SkM-HFB-QRPA	0.29	0.24	0.28	—



GERDA

Bhupal Dev, Goswami, Mitra,  
W.R., Phys. Rev. **D88**

# Generalization...

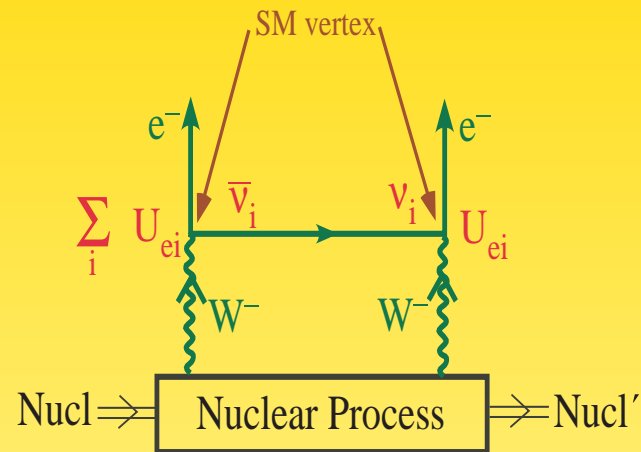


## From life-time to particle physics: Nuclear Matrix Elements



*Dark Lord Of The Sith by AG 88, 1999 Star Wars: Ord Mantell. [www.starwars.priv.pl](http://www.starwars.priv.pl)*

## From life-time to particle physics: Nuclear Matrix Elements



- 2 point-like Fermi vertices; “long-range” neutrino exchange; momentum  $q \simeq 1/r \simeq 0.1$  GeV
- NME  $\leftrightarrow$  overlap of decaying nucleons. . .
- different approaches (QRPA, NSM, IBM, GCM, pHFB) imply uncertainty
- plus uncertainty due to model details

typical model for NME: set of single particle states with a number of possible wave function configurations; obtained by solving Dirac equation in a mean background field; interaction known, treatment of fields differs

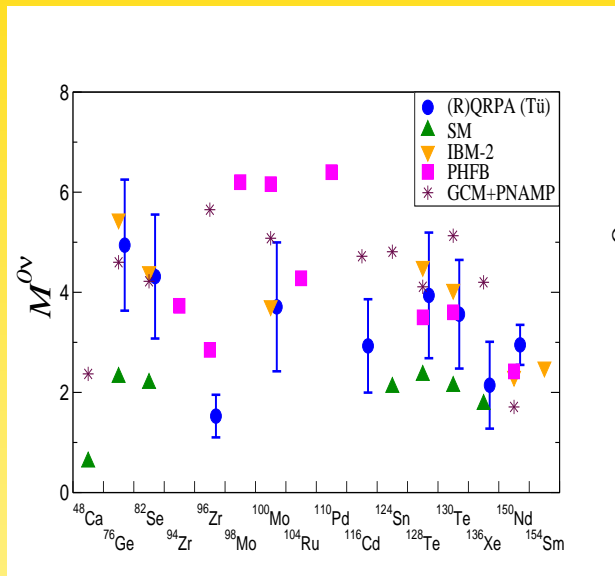
- Quasi-particle Random Phase Approximation (QRPA) (many single particle states, few configurations)
- Nuclear Shell Model (NSM) (many configurations, few single particle states)
- Interacting Boson Model (IBM) (many single particle states, few configurations)
- Generating Coordinate Method (GCM) (many single particle states, few configurations)
- projected Hartree-Fock-Bogoliubov model (pHFB)

tends to overestimate NMEs

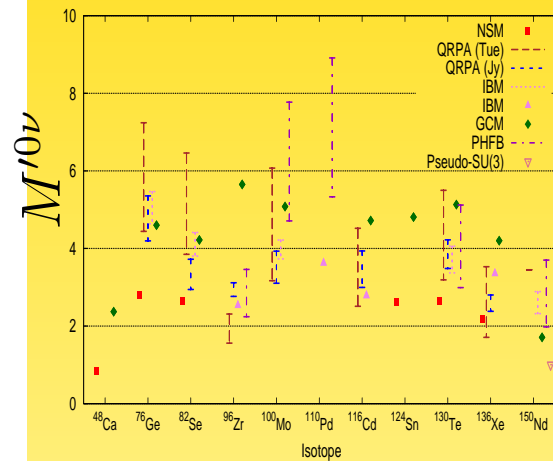
tends to underestimate NMEs



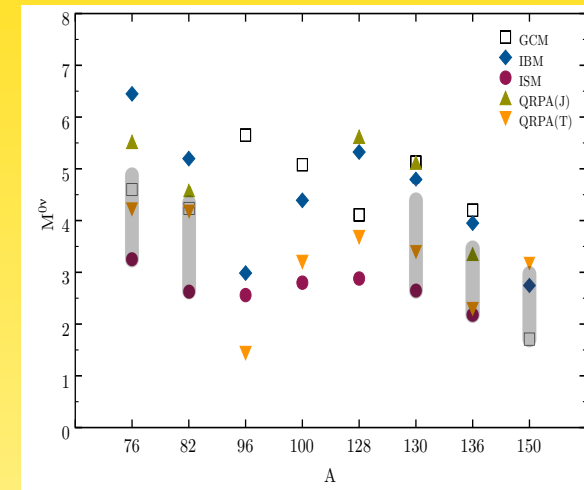
# From life-time to particle physics: Nuclear Matrix Elements



Faessler, 1104.3700

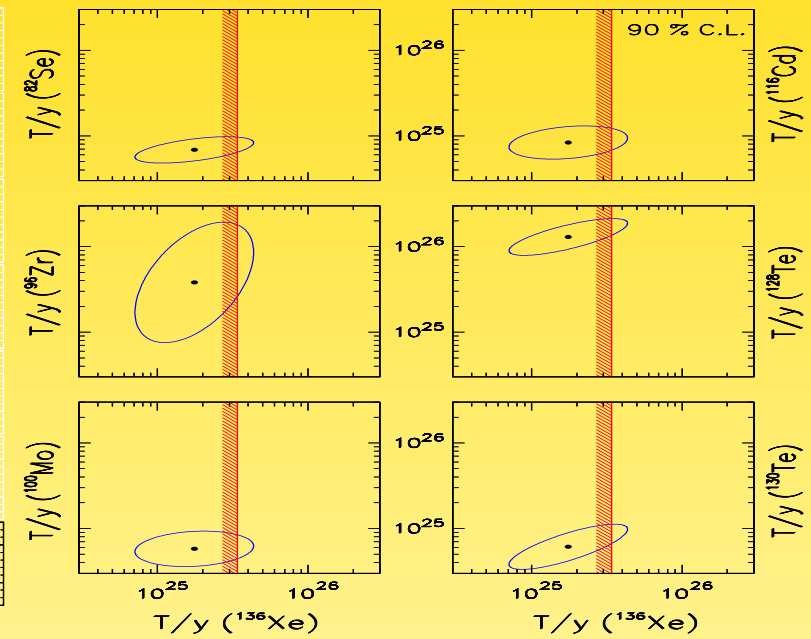
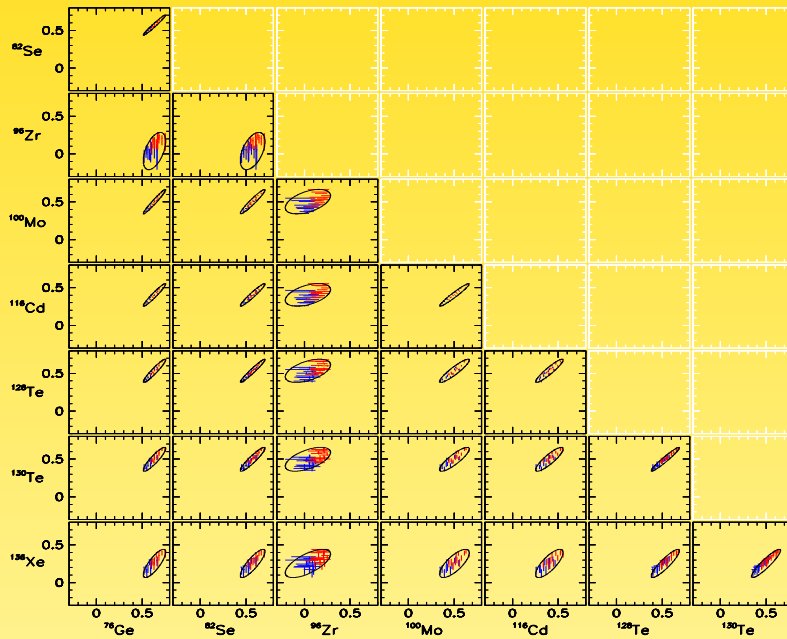


Dueck, W.R., Zuber, PRD **83**



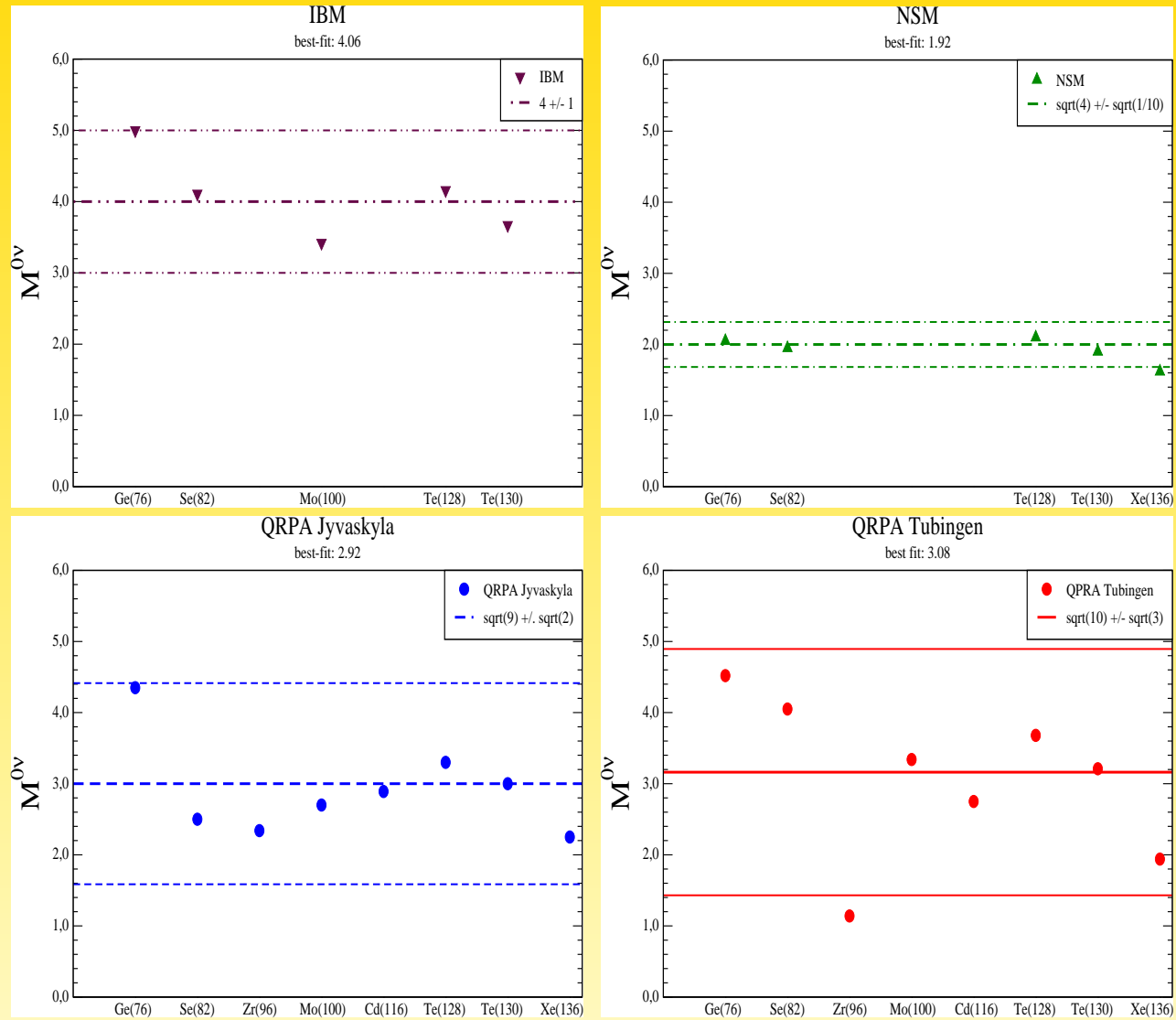
Gomez-Cadenas *et al.*, 1109.5515

to better estimate error range: correlations need to be understood



Faessler, Fogli *et al.*, PRD 79+87

# NMEs are order one numbers :-)



## Experimental Aspect

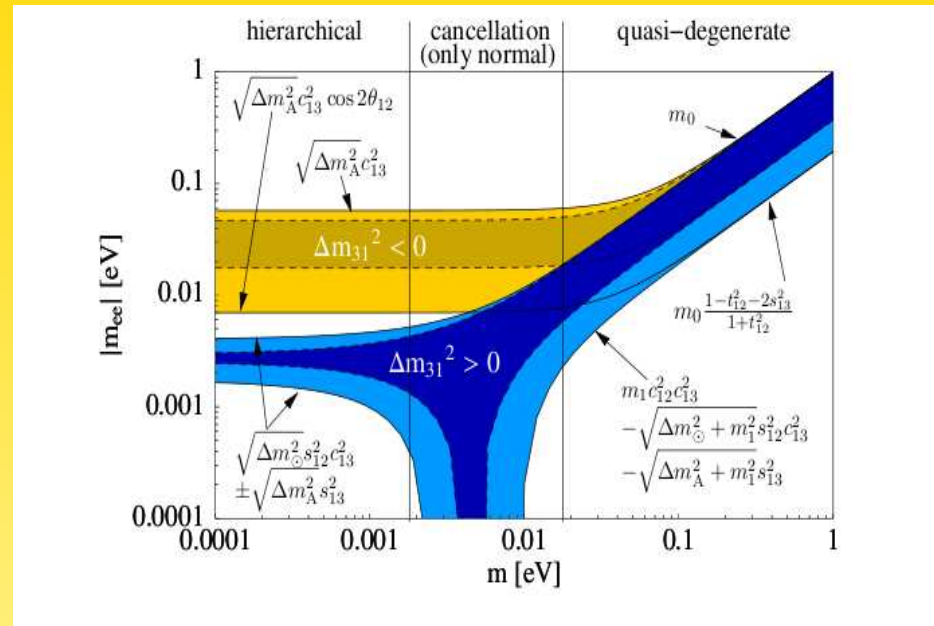
$$(T_{1/2}^{0\nu})^{-1} \propto \begin{cases} a M \epsilon t & \text{without background} \\ a \epsilon \sqrt{\frac{M t}{B \Delta E}} & \text{with background} \end{cases}$$

with

- $B$  is background index in counts/(keV kg yr)
- $\Delta E$  is energy resolution
- $\epsilon$  is efficiency
- $(T_{1/2}^{0\nu})^{-1} \propto (\text{particle physics})^2$

*Note: factor 2 in particle physics is combined factor of 16 in  $M \times t \times B \times \Delta E$*

## Inverted Ordering



Nature provides 2 scales:

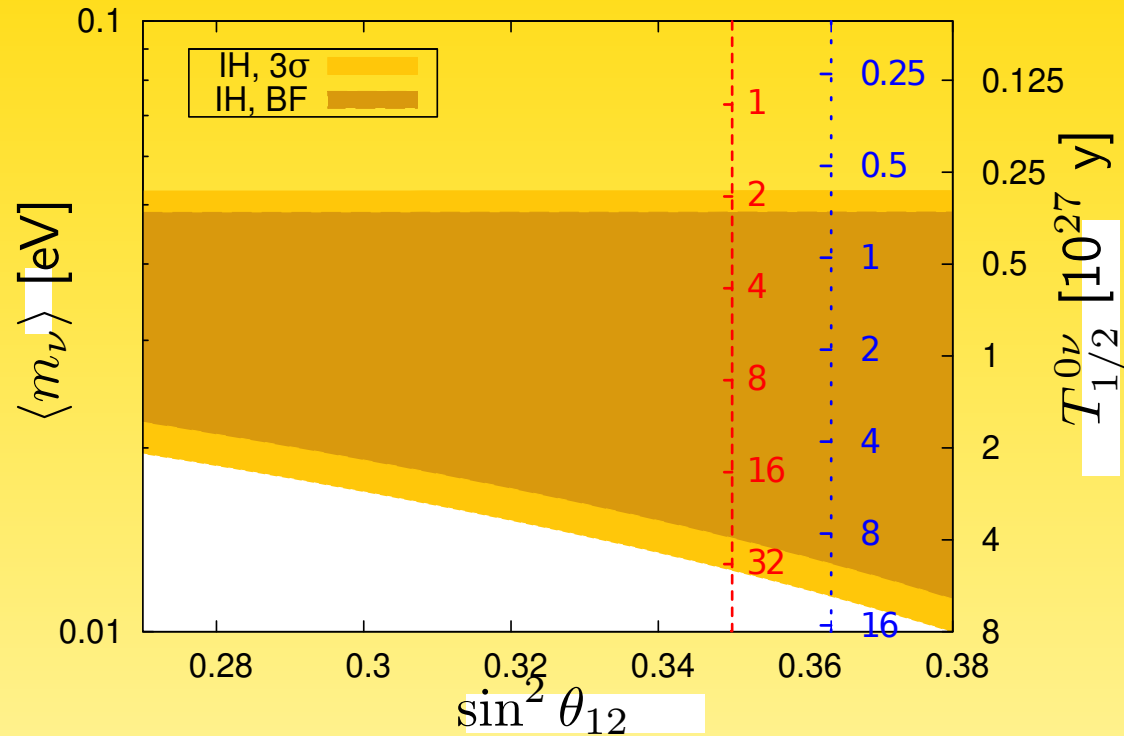
$$|m_{ee}|_{\max}^{\text{IH}} \simeq c_{13}^2 \sqrt{\Delta m_A^2} \quad \text{and} \quad |m_{ee}|_{\min}^{\text{IH}} \simeq c_{13}^2 \sqrt{\Delta m_A^2} \cos 2\theta_{12}$$

requires  $\mathcal{O}(10^{26} \dots 10^{27})$  yrs

is the lower limit  $|m_{ee}|_{\min}^{\text{IH}}$  fixed?

# Inverted Hierarchy

$m_3 = 0.001 \text{ eV}$

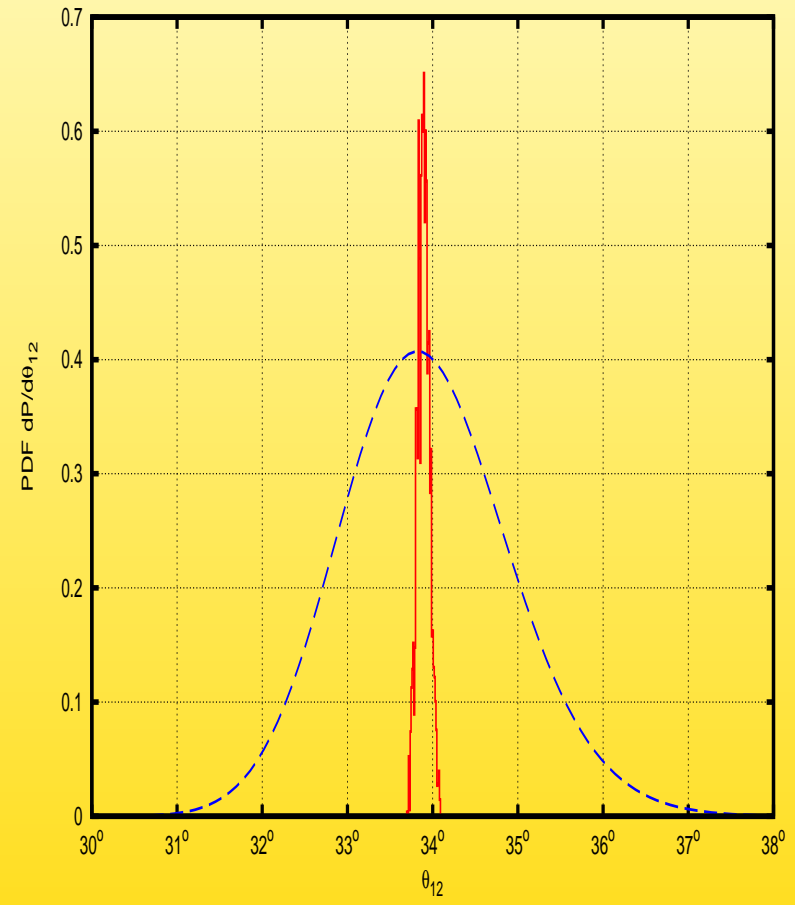
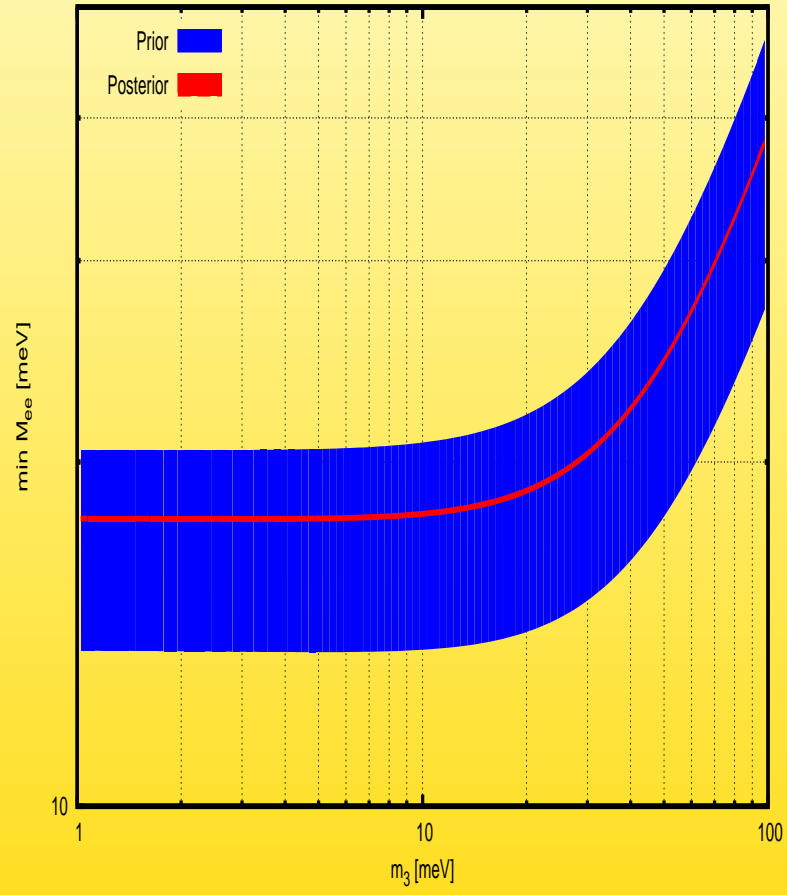


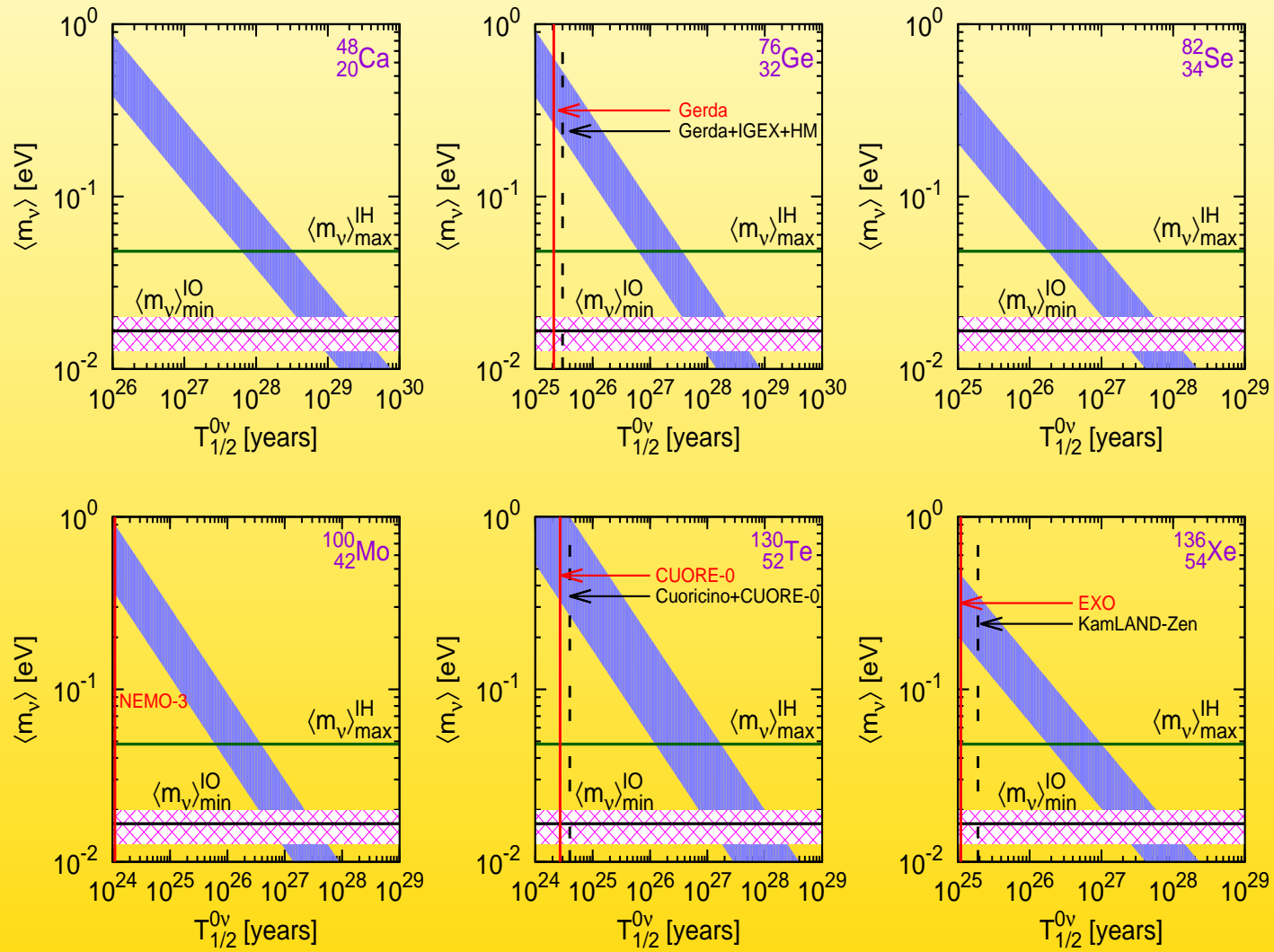
Current  $3\sigma$  range of  $\sin^2 \theta_{12}$  gives factor of  $\sim 2$  uncertainty for  $|m_{ee}|_{\min}^{\text{IH}}$

$\Rightarrow$  combined factor of  $\sim 16$  in  $M \times t \times B \times \Delta E$

$\Rightarrow$  need precision determination of  $\theta_{12}$ !  $\leftrightarrow$  JUNO

Dueck, W.R., Zuber, PRD **83**; Ge, W.R., 1507.05514







## With $0\nu\beta\beta$ one can

- test lepton number violation
- test Majorana nature of neutrinos
- probe neutrino mass scale
- extract Majorana phase
- constrain inverted ordering

conceptually, it would increase our believe in

- GUTs
- seesaw mechanism
- leptogenesis

## Sterile Neutrinos??

- LSND/MiniBooNE/gallium
- cosmology
- BBN
- $r$ -process nucleosynthesis in Supernovae
- reactor anomaly

	$\Delta m_{41}^2 [\text{eV}^2]$	$ U_{e4} $	$ U_{\mu 4} $	$\Delta m_{51}^2 [\text{eV}^2]$	$ U_{e5} $	$ U_{\mu 5} $
3+2/2+3	0.47	0.128	0.165	0.87	0.138	0.148
1+3+1	0.47	0.129	0.154	0.87	0.142	0.163

$$\text{or } \Delta m_{41}^2 = 1.78 \text{ eV}^2 \text{ and } |U_{e4}|^2 = 0.151$$

Kopp, Maltoni, Schwetz, 1103.4570

see lectures by William Louis and Karsten Heeger

## Sterile Neutrinos and $0\nu\beta\beta$

- recall:  $|m_{ee}|_{\text{NH}}^{\text{act}}$  can vanish and  $|m_{ee}|_{\text{IH}}^{\text{act}} \sim 0.03 \text{ eV}$  cannot vanish
- $|m_{ee}| = \left| \underbrace{|U_{e1}|^2 m_1 + |U_{e2}|^2 m_2 e^{2i\alpha} + |U_{e3}|^2 m_3 e^{2i\beta}}_{m_{ee}^{\text{act}}} + \underbrace{|U_{e4}|^2 m_4 e^{2i\Phi_1}}_{m_{ee}^{\text{st}}} \right|$
- sterile contribution to  $0\nu\beta\beta$  (assuming 1+3):

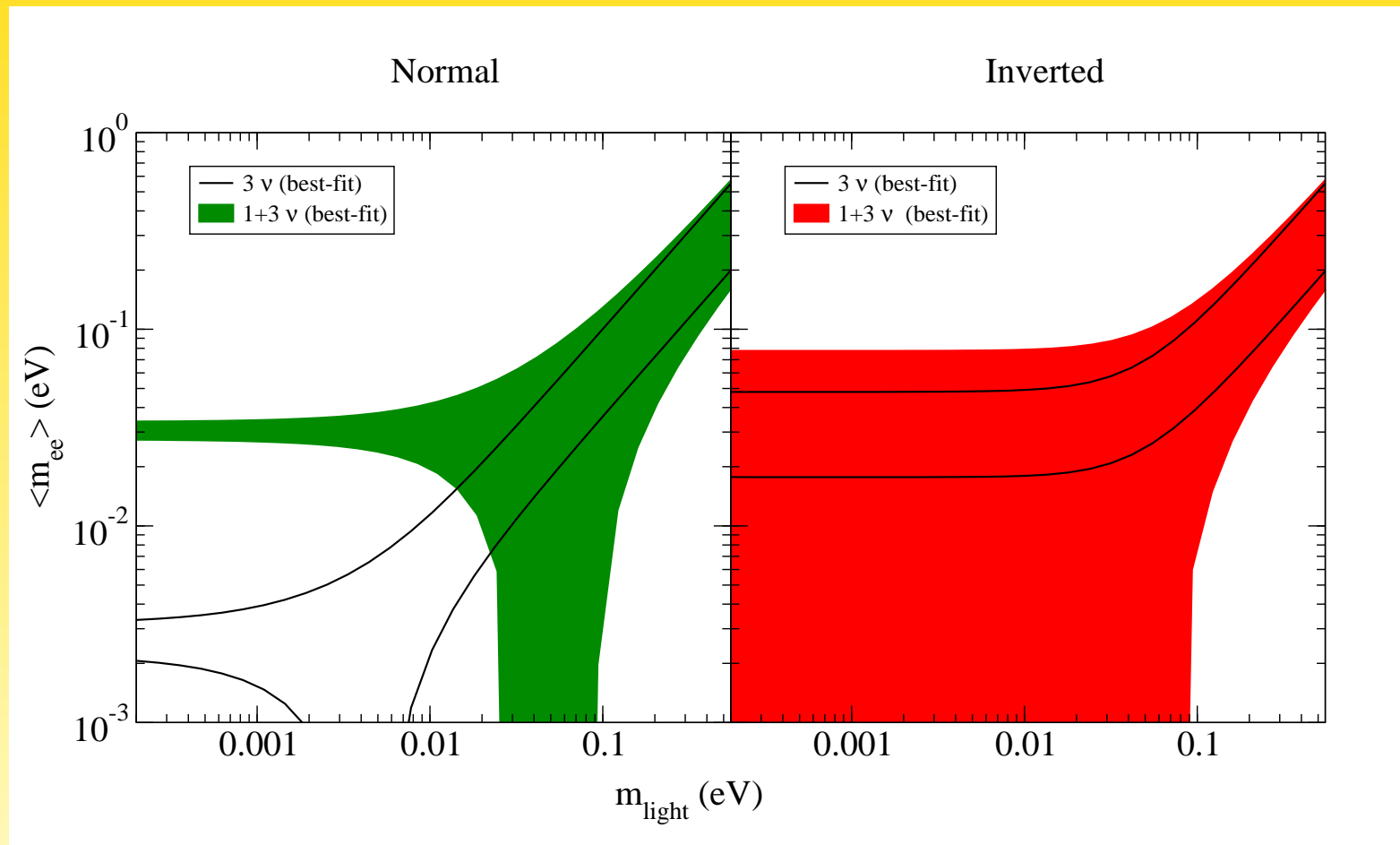
$$|m_{ee}|^{\text{st}} \simeq \sqrt{\Delta m_{\text{st}}^2} |U_{e4}|^2 \begin{cases} \gg |m_{ee}|_{\text{NH}}^{\text{act}} \\ \simeq |m_{ee}|_{\text{IH}}^{\text{act}} \end{cases}$$

$\Rightarrow |m_{ee}|_{\text{NH}}$  cannot vanish and  $|m_{ee}|_{\text{IH}}$  can vanish!

$\Rightarrow$  usual phenomenology gets completely turned around!

Barry, W.R., Zhang, JHEP **1107**; Giunti *et al.*, PRD **87**; Girardi, Meroni, Petcov, JHEP **1311**; Giunti, Zavanin, 1505.00978

Usual plot gets completely turned around!



## Interpretation of Neutrino-less Double Beta Decay

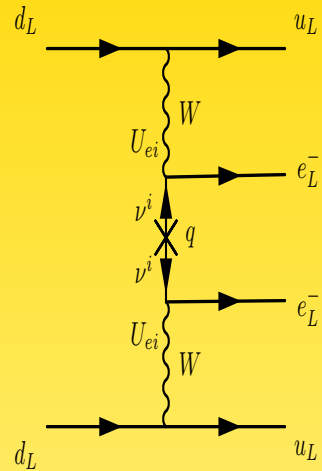
- **Standard Interpretation:**

Neutrinoless Double Beta Decay is mediated by light and massive Majorana neutrinos (the ones which oscillate) and all other mechanisms potentially leading to  $0\nu\beta\beta$  give negligible or no contribution

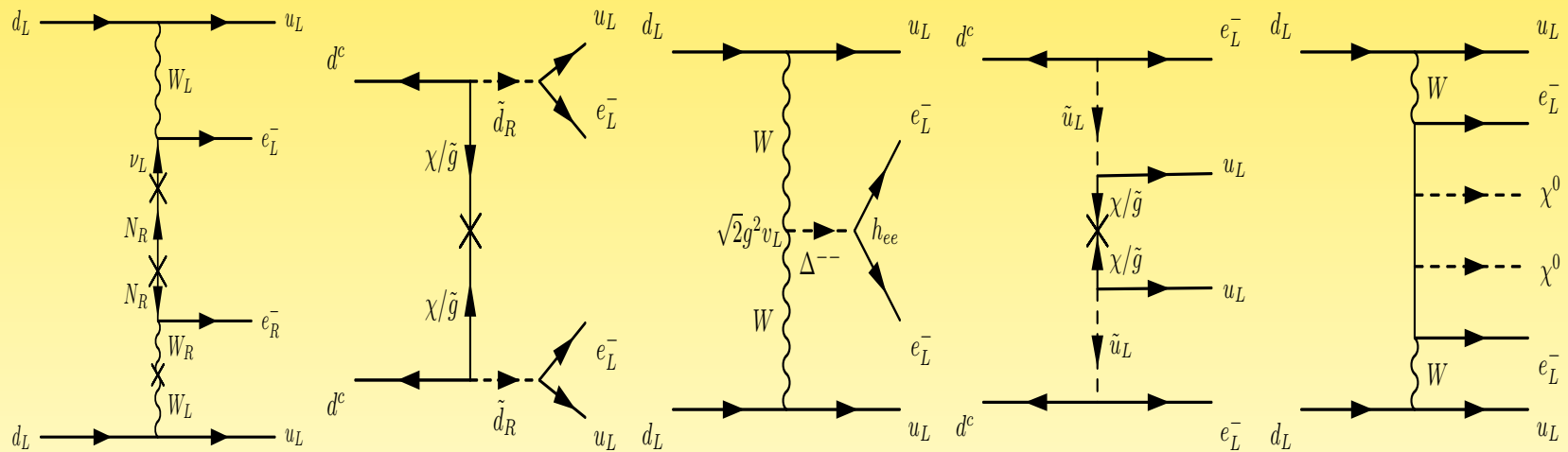
- **Non-Standard Interpretations:**

There is at least one other mechanism leading to Neutrinoless Double Beta Decay and its contribution is at least of the same order as the light neutrino exchange mechanism

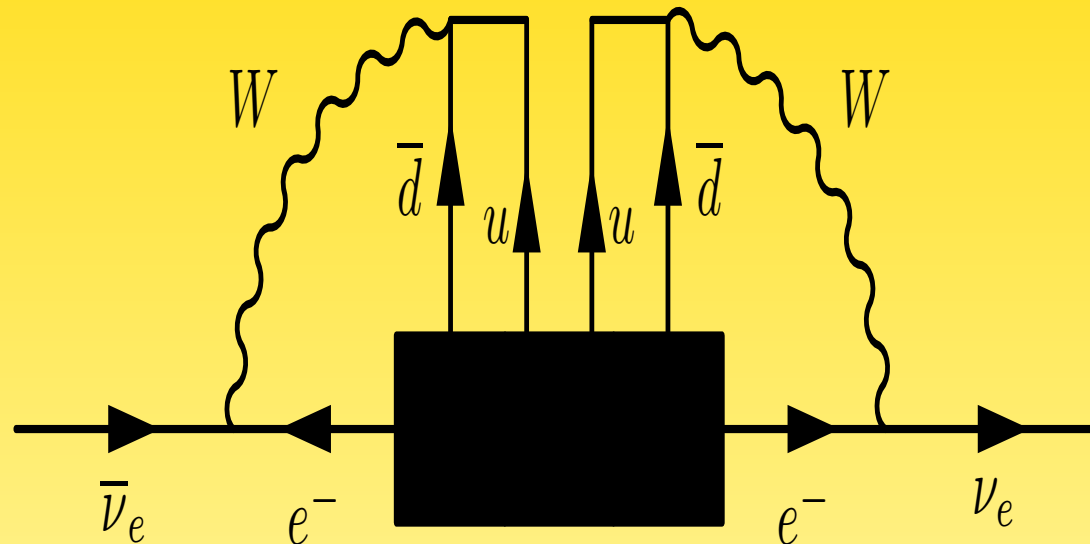
• **Standard Interpretation:**



• **Non-Standard Interpretations:**



Schechter-Valle theorem: observation of  $0\nu\beta\beta$  implies Majorana neutrinos



is 4 loop diagram:  $m_{\nu}^{\text{BB}} \sim \frac{G_F^2}{(16\pi^2)^4} \text{MeV}^5 \sim 10^{-25} \text{eV}$

explicit calculation: [Duerr, Lindner, Merle, 1105.0901](#)

mechanism	physics parameter	current limit	test
light neutrino exchange	$ U_{ei}^2 m_i $	0.5 eV	oscillations, cosmology, neutrino mass
heavy neutrino exchange	$ \frac{S_{ei}^2}{M_i} $	$2 \times 10^{-8} \text{ GeV}^{-1}$	LFV, collider
heavy neutrino and RHC	$ \frac{V_{ei}^2}{M_i M_{WR}^4} $	$4 \times 10^{-16} \text{ GeV}^{-5}$	flavor, collider
Higgs triplet and RHC	$ \frac{(M_R)_{ee}}{m_{\Delta_R}^2 M_{WR}^4} $	$10^{-15} \text{ GeV}^{-1}$	flavor, collider $e^-$ distribution
$\lambda$ -mechanism with RHC	$ \frac{U_{ei} \tilde{S}_{ei}}{M_{WR}^2} $	$1.4 \times 10^{-10} \text{ GeV}^{-2}$	flavor, collider, $e^-$ distribution
$\eta$ -mechanism with RHC	$\tan \zeta  U_{ei} \tilde{S}_{ei} $	$6 \times 10^{-9}$	flavor, collider, $e^-$ distribution
short-range $\mathcal{R}$	$\frac{ \lambda'_{111} ^2}{\Lambda_{\text{SUSY}}^5}$ $\Lambda_{\text{SUSY}} = f(m_{\tilde{g}}, m_{\tilde{u}_L}, m_{\tilde{d}_R}, m_{\chi_i})$	$7 \times 10^{-18} \text{ GeV}^{-5}$	collider, flavor
long-range $\mathcal{R}$	$ \sin 2\theta^b \lambda'_{131} \lambda'_{113} \left( \frac{1}{m_{b_1}^2} - \frac{1}{m_{b_2}^2} \right) $ $\sim \frac{G_F}{q} m_b \frac{ \lambda'_{131} \lambda'_{113} }{\Lambda_{\text{SUSY}}^3}$	$2 \times 10^{-13} \text{ GeV}^{-2}$ $1 \times 10^{-14} \text{ GeV}^{-3}$	flavor, collider
Majorons	$ \langle g_\chi \rangle $ or $ \langle g_\chi \rangle ^2$	$10^{-4} \dots 1$	spectrum, cosmology



## Energy Scale:

Note: *standard amplitude* for light Majorana neutrino exchange:

$$\mathcal{A}_l \simeq G_F^2 \frac{|m_{ee}|}{q^2} \simeq 7 \times 10^{-18} \left( \frac{|m_{ee}|}{0.5 \text{ eV}} \right) \text{ GeV}^{-5} \simeq 2.7 \text{ TeV}^{-5}$$

if new heavy particles are exchanged:

$$\mathcal{A}_h \simeq \frac{c}{M^5}$$

$\Rightarrow$  for  $0\nu\beta\beta$  holds:

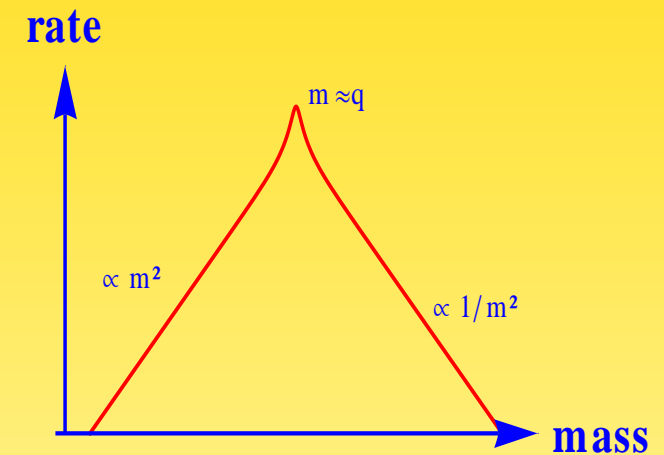
$$1 \text{ eV} = 1 \text{ TeV}$$

$\Rightarrow$  Phenomenology in colliders, LFV

## Examples

- Fermions (and no RHC):

$$\mathcal{A} \propto \frac{m}{q^2 - m^2} \rightarrow \begin{cases} m & \text{for } q^2 \gg m^2 \\ \frac{1}{m} & \text{for } q^2 \ll m^2 \end{cases}$$

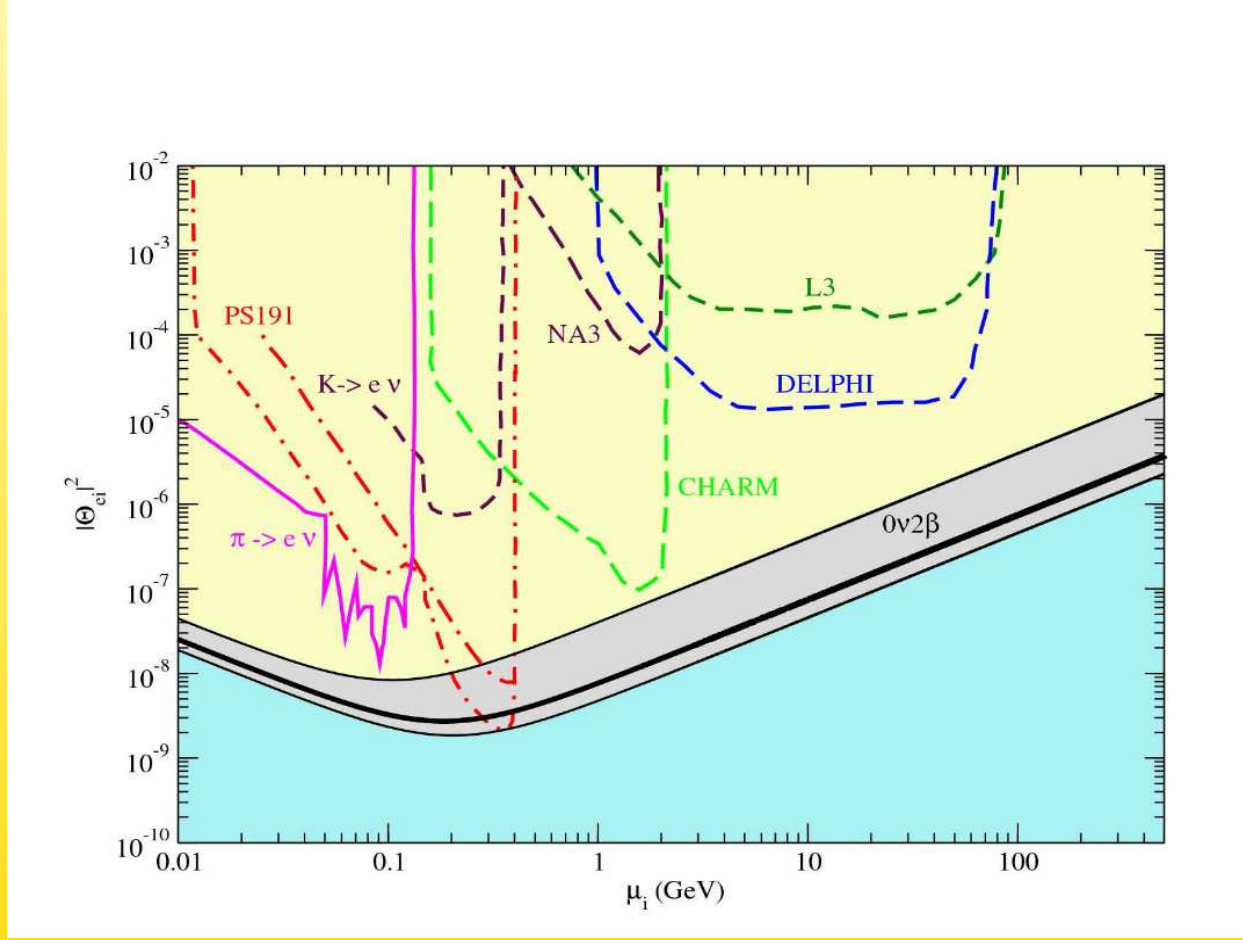


Note: maximum  $\mathcal{A}$  corresponds to  $m \simeq \langle q \rangle$ : limits on  $\mathcal{O}(m_K)$  Majorana neutrinos from  $K^+ \rightarrow \pi^- \mu^+ e^+$

- heavy scalar/vector:

$$\mathcal{A} \propto \frac{1}{q^2 - m^2} \rightarrow \frac{1}{m^2}$$

Mitra, Senjanovic, Vissani, 1108.0004



Heavy neutrinos

$$T^{0\nu}(1 \text{ eV}) = T^{0\nu}(1 \text{ TeV})$$

- RPV Supersymmetry
- left-right symmetry
- heavy neutrinos
- color octets
- leptoquarks
- effective operators
- extra dimensions
- ...

⇒ need to solve the inverse problem...

## Distinguishing Mechanisms

### The inverse problem of $0\nu\beta\beta$

- 1.) Other observables (LHC, LFV, KATRIN, cosmology, ...)
- 2.) Decay products (individual  $e^-$  energies, angular correlations, spectrum, ...)
- 3.) Nuclear physics (multi-isotope,  $0\nu\text{ECEC}$ ,  $0\nu\beta^+\beta^+$ , ...)

## 1.) Example Left-Right Symmetry

very simple extension of SM gauge group to  $SU(2)_L \times SU(2)_R \times U(1)_{B-L}$

usual particle content:

$$L_{Li} = \begin{pmatrix} \nu'_L \\ \ell_L \end{pmatrix}_i \sim (\mathbf{2}, \mathbf{1}, -1), \quad L_{Ri} = \begin{pmatrix} \nu_R \\ \ell_R \end{pmatrix}_i \sim (\mathbf{1}, \mathbf{2}, -1)$$

$$Q_{Li} = \begin{pmatrix} u_L \\ d_L \end{pmatrix}_i \sim (\mathbf{2}, \mathbf{1}, \frac{1}{3}), \quad Q_{Ri} = \begin{pmatrix} u_R \\ d_R \end{pmatrix}_i \sim (\mathbf{1}, \mathbf{2}, \frac{1}{3})$$

for symmetry breaking:

$$\Delta_L \equiv \begin{pmatrix} \delta_L^+/\sqrt{2} & \delta_L^{++} \\ \delta_L^0 & -\delta_L^+/\sqrt{2} \end{pmatrix} \sim (\mathbf{3}, \mathbf{1}, \mathbf{2}), \quad \Delta_R \equiv \begin{pmatrix} \delta_R^+/\sqrt{2} & \delta_R^{++} \\ \delta_R^0 & -\delta_R^+/\sqrt{2} \end{pmatrix} \sim (\mathbf{1}, \mathbf{3}, \mathbf{2})$$

$$\phi \equiv \begin{pmatrix} \phi_1^0 & \phi_2^+ \\ \phi_1^- & \phi_2^0 \end{pmatrix} \sim (\mathbf{2}, \mathbf{2}, \mathbf{0})$$

## Left-right Symmetry

- very rich Higgs sector (13 extra scalars)

- rich gauge boson sector ( $Z', M_{W_R^\pm}$ ) with

$$M_{Z'} = \sqrt{\frac{2}{1-\tan^2 \theta_W}} M_{W_R} \simeq 1.7 M_{W_R} \gtrsim 4.3 \text{ TeV}$$

- 'sterile' neutrinos  $\nu_R$

- type I + type II seesaw for neutrino mass

- right-handed currents with strength  $G_F \left(\frac{g_R}{g_L}\right)^2 \left(\frac{m_W}{M_{W_R}}\right)^2$

- $m_\nu \propto 1/M_{W_R}$ : maximal parity violation  $\leftrightarrow$  smallness of neutrino mass

*(Note: in case of modified symmetry breaking  $g_L \neq g_R$  and  $M_{Z'} < M_{W_R}$  possible. . .)*

## Left-right Symmetry

6 neutrinos with flavor states  $n'_L$  and mass states  $n_L = (\nu_L, N_R^c)^T$

$$n'_L = \begin{pmatrix} \nu'_L \\ \nu_R^c \end{pmatrix} = \begin{pmatrix} K_L \\ K_R \end{pmatrix} n_L = \begin{pmatrix} U & S \\ T & V \end{pmatrix} \begin{pmatrix} \nu_L \\ N_R^c \end{pmatrix}$$

Right-handed currents:

$$\mathcal{L}_{CC}^{\text{lep}} = \frac{g}{\sqrt{2}} \left[ \bar{\ell}_L \gamma^\mu K_L n_L (W_{1\mu}^- + \xi e^{i\alpha} W_{2\mu}^-) + \bar{\ell}_R \gamma^\mu K_R n_L^c (-\xi e^{-i\alpha} W_{1\mu}^- + W_{2\mu}^-) \right]$$

( $K_L$  and  $K_R$  are  $3 \times 6$  mixing matrices)

plus: gauge boson mixing

$$\begin{pmatrix} W_L^\pm \\ W_R^\pm \end{pmatrix} = \begin{pmatrix} \cos \xi & \sin \xi e^{i\alpha} \\ -\sin \xi e^{-i\alpha} & \cos \xi \end{pmatrix} \begin{pmatrix} W_1^\pm \\ W_2^\pm \end{pmatrix}$$



## Connection to Neutrinos

Majorana mass matrices  $M_L = f_L v_L$  from  $\langle \Delta_L \rangle$  and  $M_R = f_R v_R$  from  $\langle \Delta_R \rangle$   
(with  $f_L = f_R = f$ )

$$\mathcal{L}_{\text{mass}}^\nu = -\frac{1}{2} \begin{pmatrix} \overline{\nu'_L} & \overline{\nu'_R} \end{pmatrix} \begin{pmatrix} M_L & M_D \\ M_D^T & M_R \end{pmatrix} \begin{pmatrix} \nu'_L \\ \nu'_R \end{pmatrix} \Rightarrow m_\nu = M_L - M_D M_R^{-1} M_D^T$$

useful special cases

- (i) type I dominance:  $m_\nu = M_D M_R^{-1} M_D^T = M_D f_R^{-1} / v_R M_D^T$
- (ii) type II dominance:  $m_\nu = f_L v_L$

for case (i): mixing of light neutrinos with heavy neutrinos of order

$$|S_{\alpha i}| \simeq |T_{\alpha i}^T| \simeq \sqrt{\frac{m_\nu}{M_i}} \lesssim 10^{-7} \left( \frac{\text{TeV}}{M_i} \right)^{1/2}$$

small (or enhanced up to  $10^{-2}$  by cancellations)

## Right-handed Currents in Double Beta Decay

$$(A, Z) \rightarrow (A, Z + 2) + 2e^-$$

$$\mathcal{L}_{CC}^{\text{lep}} = \frac{g}{\sqrt{2}} \sum_{i=1}^3 \left[ \bar{e}_L \gamma^\mu (U_{ei} \nu_{Li} + S_{ei} N_{Ri}^c) (W_{1\mu}^- + \xi e^{i\alpha} W_{2\mu}^-) \right. \\ \left. + \bar{e}_R \gamma^\mu (T_{ei}^* \nu_{Li}^c + V_{ei}^* N_{Ri}) (-\xi e^{-i\alpha} W_{1\mu}^- + W_{2\mu}^-) \right]$$

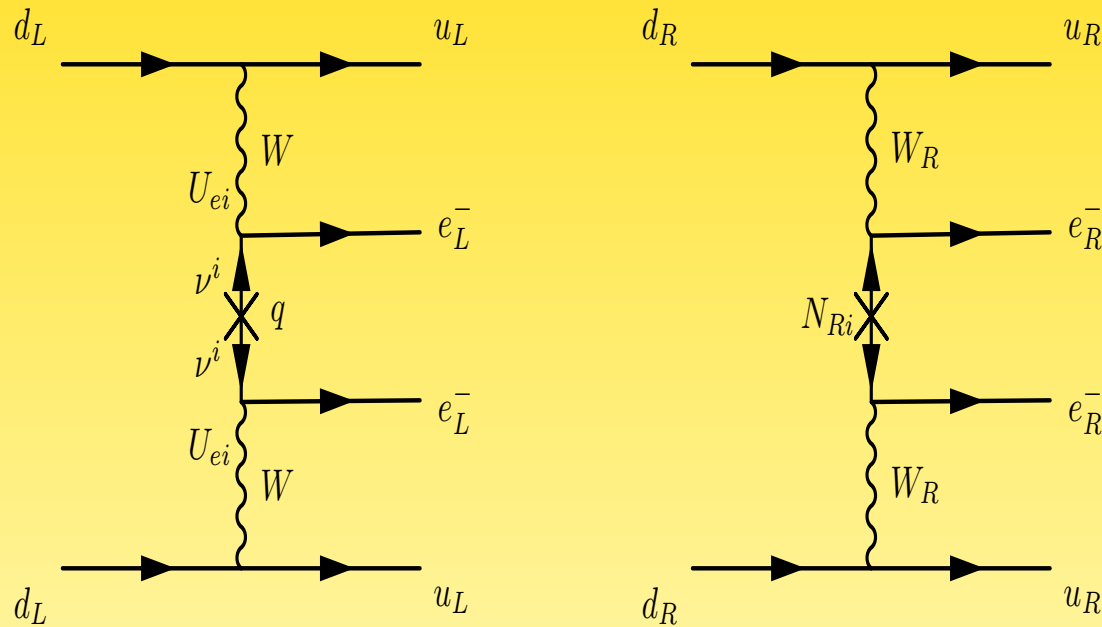
$$\mathcal{L}_Y^\ell = -\bar{L}_L'^c i\sigma_2 \Delta_L f_L L'_L - \bar{L}_R'^c i\sigma_2 \Delta_R f_R L'_R$$

classify diagrams:

- mass dependent diagrams (same helicity of electrons)
- triplet exchange diagrams (same helicity of electrons)
- momentum dependent diagrams (different helicity of electrons)

## Mass Dependent Diagrams

electrons either both left- or right-handed, leading diagrams:

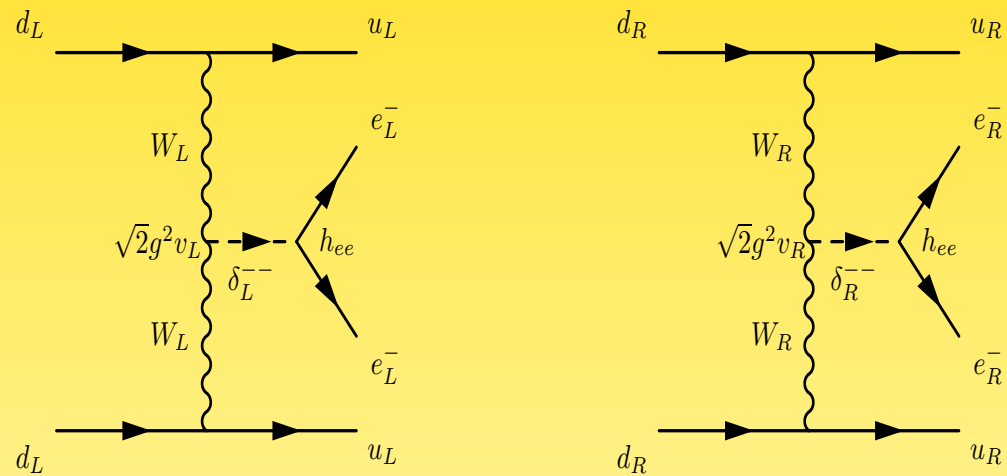


$$A_\nu \simeq G_F^2 \frac{m_{ee}}{q^2}$$

$$A_{N_R}^R \simeq G_F^2 \left( \frac{m_{W_L}}{m_{W_R}} \right)^4 \sum_i V_{ei}^{*2} \frac{1}{M_i}$$

## Triplet Exchange Diagrams

leading diagrams:



$$\mathcal{A}_{\delta_L} \simeq G_F^2 \frac{h_{ee} v_L}{m_{\delta_L}^2} \quad \mathcal{A}_{\delta_R} \simeq G_F^2 \left( \frac{m_{W_L}}{m_{W_R}} \right)^4 \sum_i \frac{V_{ei}^2 M_i}{m_{\delta_R}^2}$$

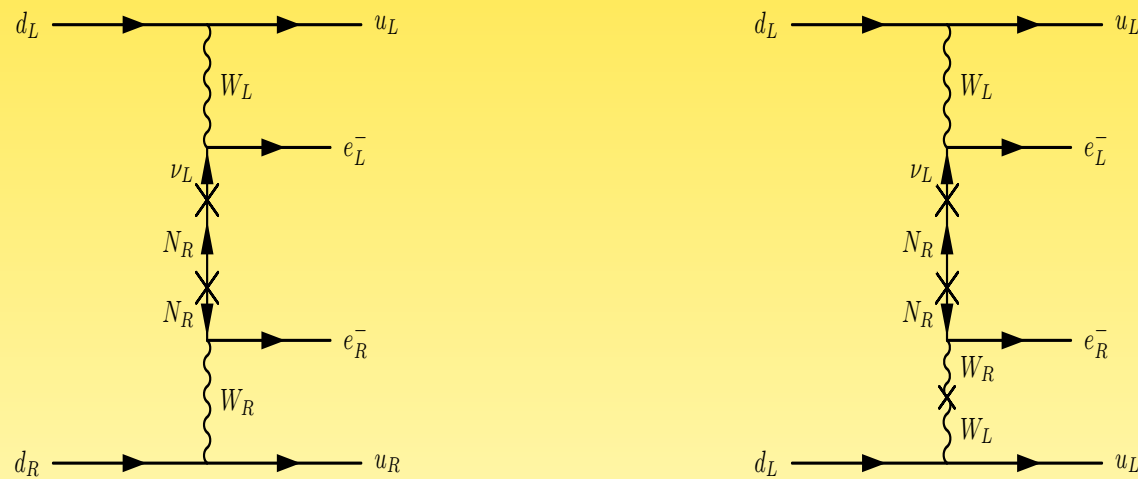
(negligible)

## Momentum Dependent Diagrams

electrons with opposite helicity

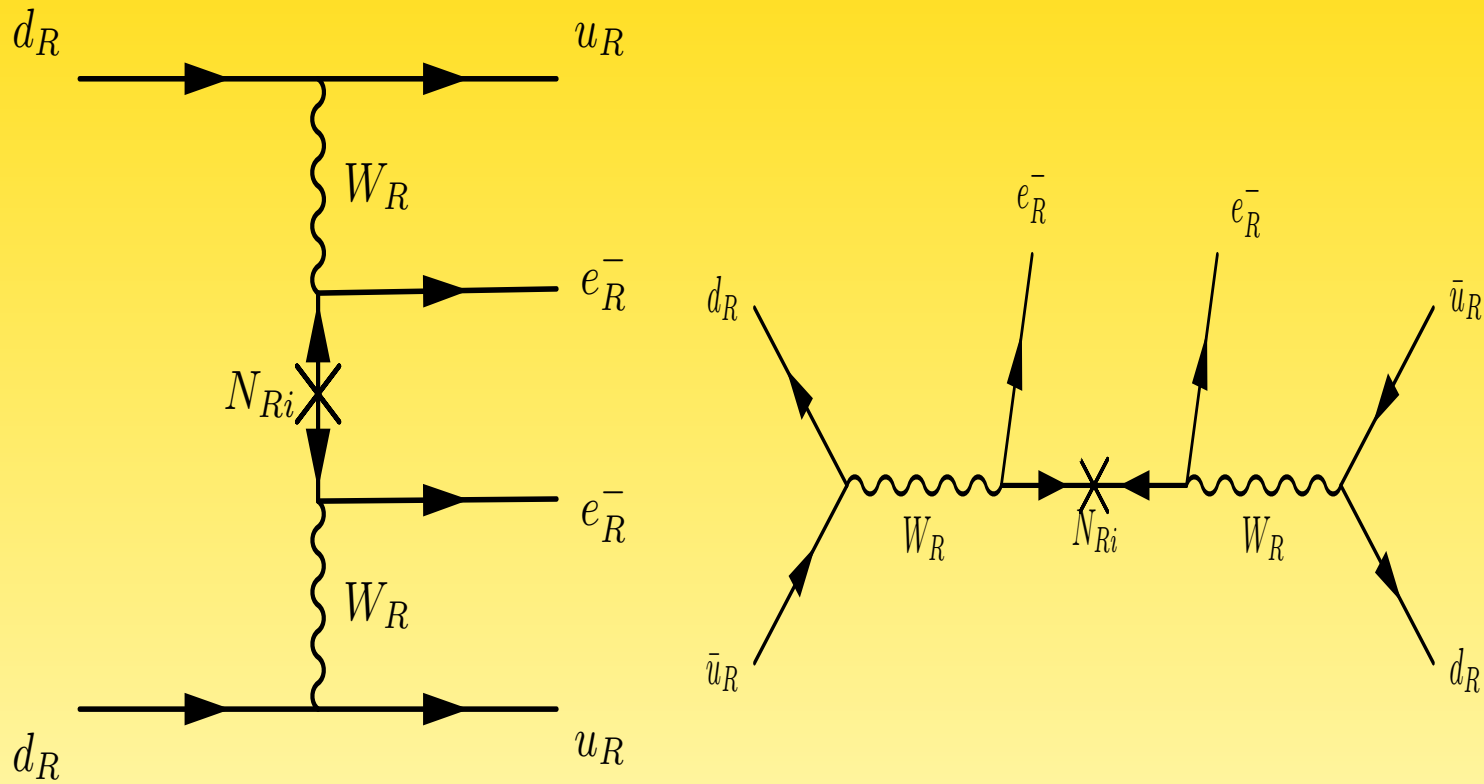
$$\mathcal{A}_{LR} \simeq G_F^2 \left( \frac{m_{W_L}^2}{m_{W_R}^2} + \tan \xi \right) \sum_i \left( U_{ei} T_{ei}^* \frac{1}{q} - S_{ei} V_{ei}^* \frac{q}{M_i^2} \right)$$

leading diagrams (long range):

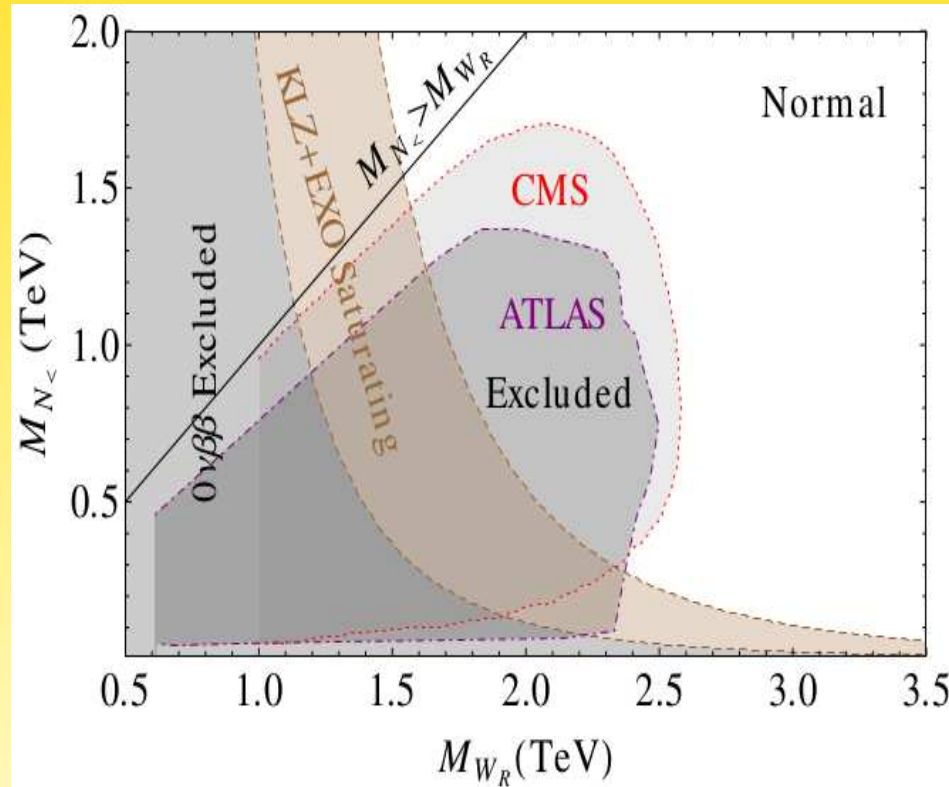
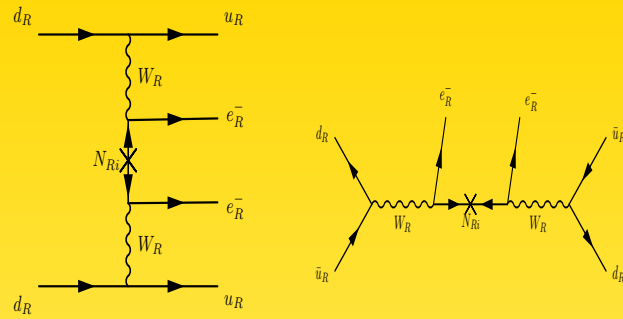


$$\mathcal{A}_\lambda \simeq G_F^2 \left( \frac{m_{W_L}}{m_{W_R}} \right)^2 \sum_i U_{ei} T_{ei}^* \frac{1}{q} \quad \mathcal{A}_\eta \simeq G_F^2 \tan \xi \sum_i U_{ei} T_{ei}^* \frac{1}{q}$$

## Left-right symmetry



Senjanovic, Keung, 1983; Tello *et al.*; Nemevsek *et al.*



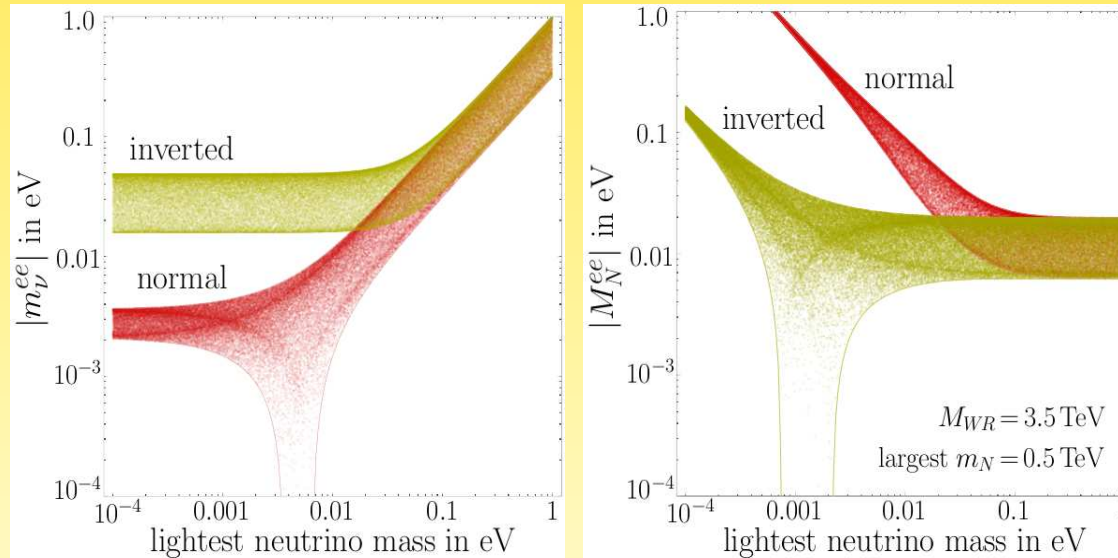
Bhupal Dev, Goswami, Mitra, W.R.

## Type II dominance (Tello *et al.*, 1011.3522)

$$m_\nu = m_L - m_D M_R^{-1} m_D^T = v_L f - \frac{v^2}{v_R} Y_D f^{-1} Y_D^T \xrightarrow{*} v_L f$$

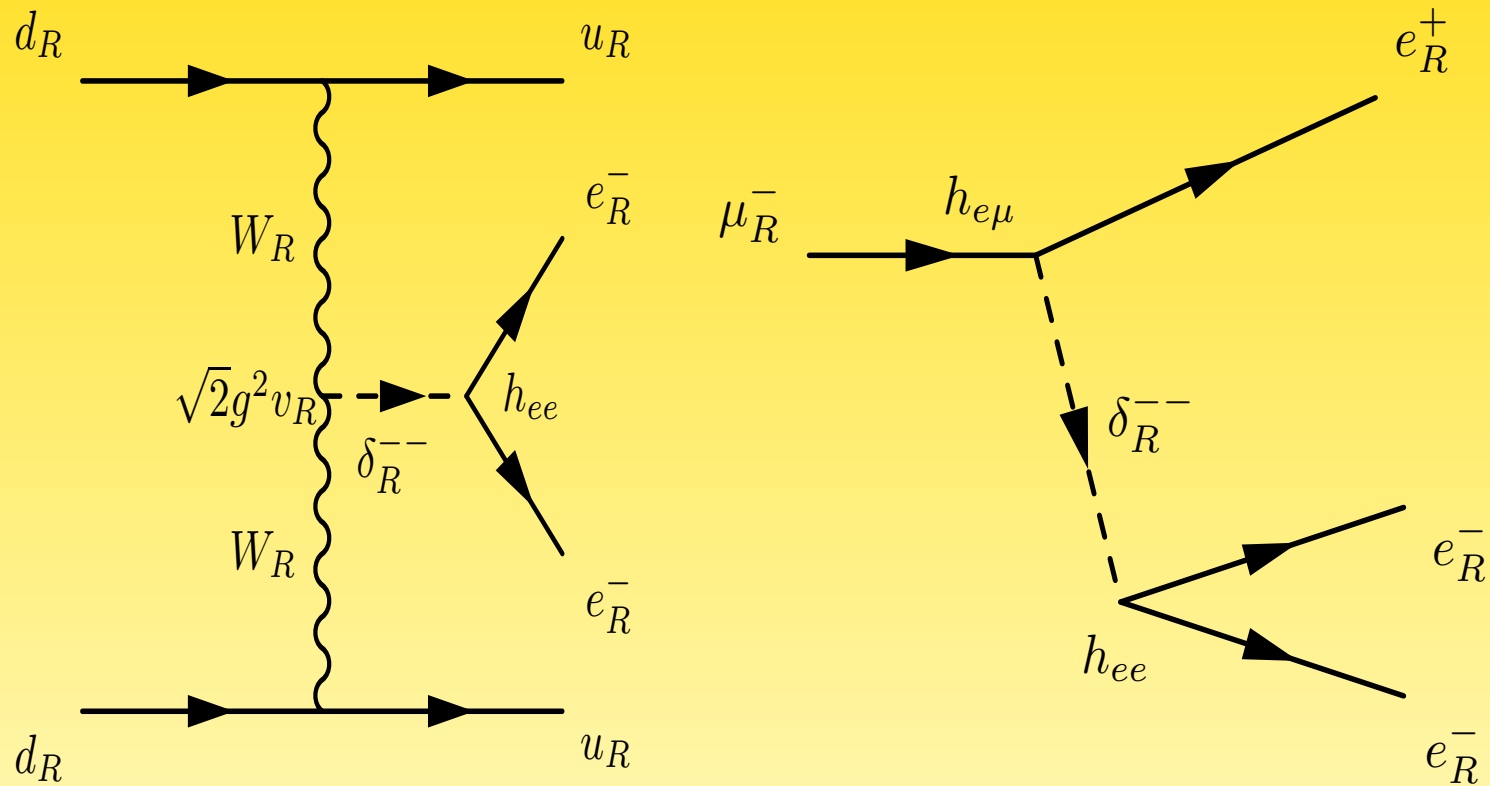
$\Rightarrow m_\nu$  fixes  $M_R$  and exchange of  $N_R$  with  $W_R$  fixed in terms of PMNS:

$$\Rightarrow \mathcal{A}_{N_R} \simeq G_F^2 \left( \frac{m_W}{M_{W_R}} \right)^4 \sum \frac{V_{ei}^2}{M_i} \propto \sum \frac{U_{ei}^2}{m_i}$$

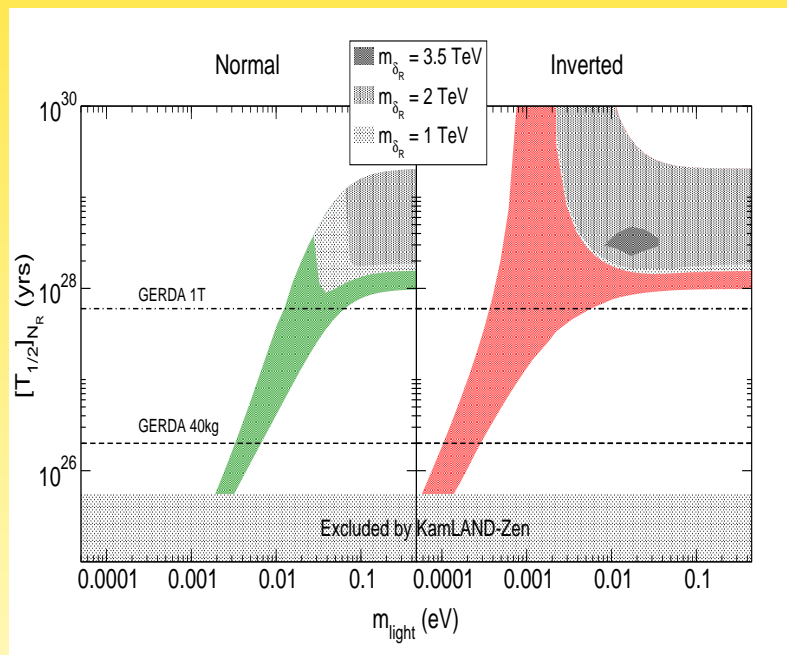
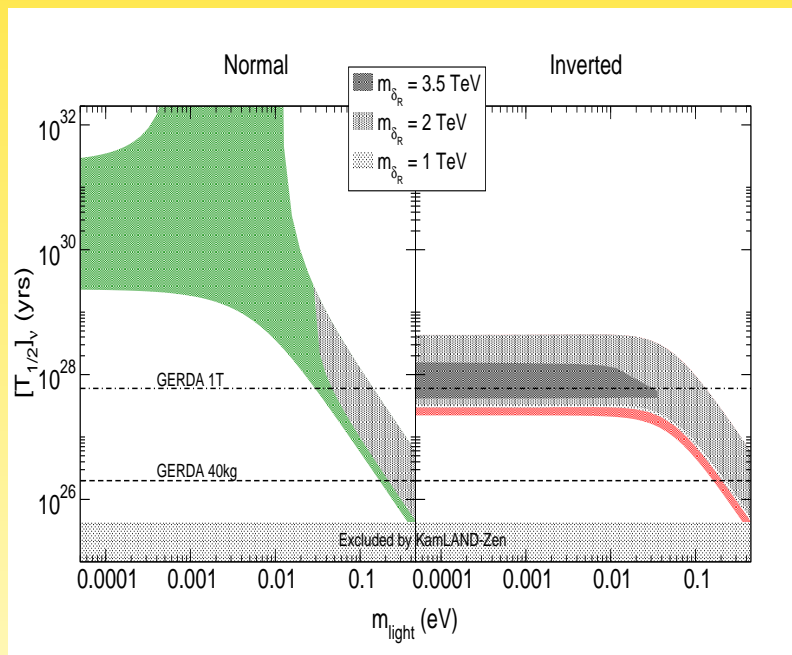
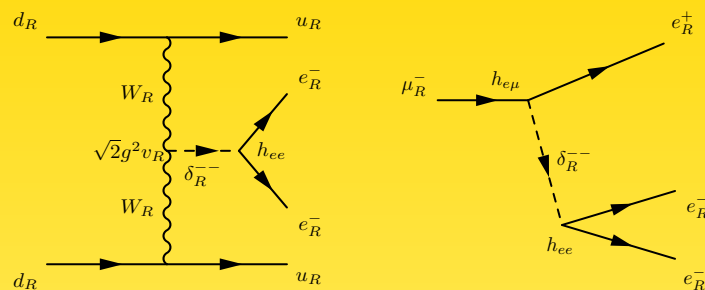




## Constraints from Lepton Flavor Violation

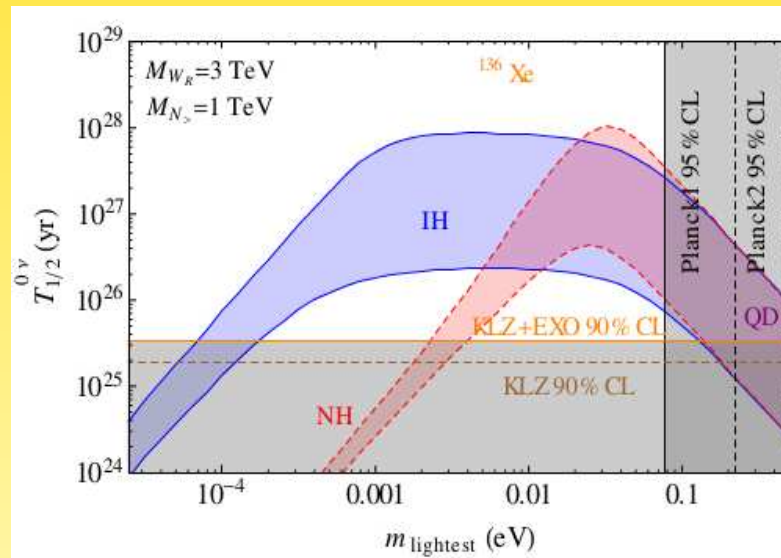
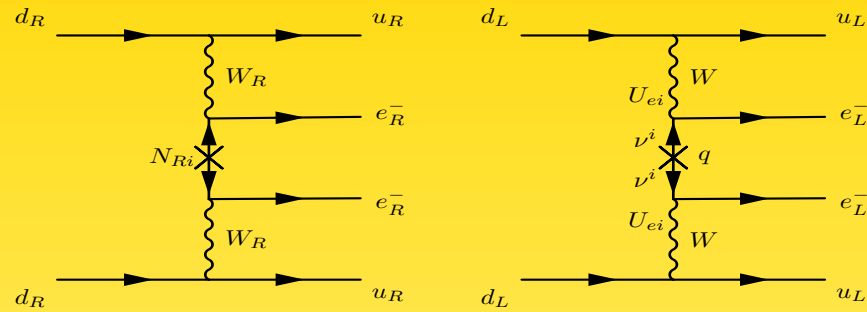


# Constraints from Lepton Flavor Violation



Barry, W.R., JHEP 1309

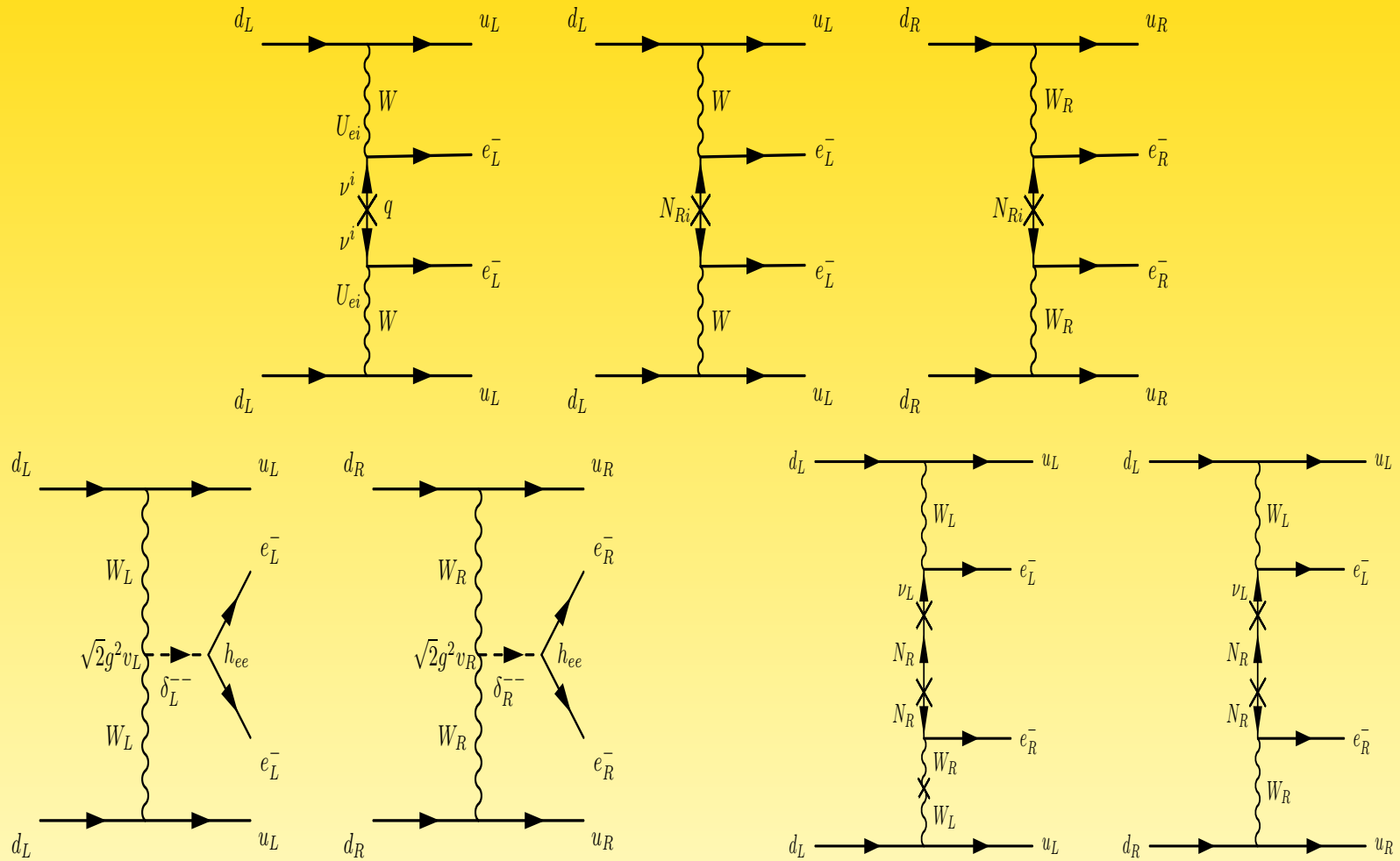
## Adding diagrams



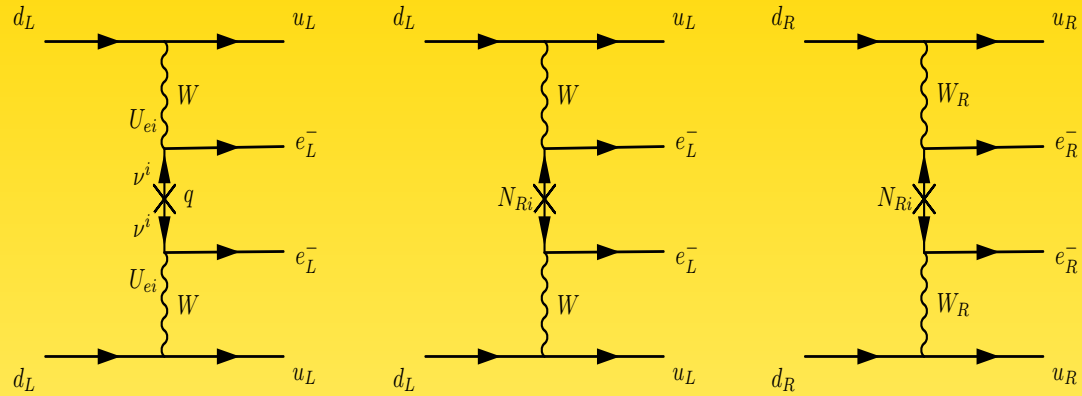
$\Rightarrow$  lower bound on  $m(\text{lightest}) \gtrsim \text{meV}$

Bhupal Dev, Goswami, Mitra, W.R., Phys. Rev. **D88**

# Left-right symmetry



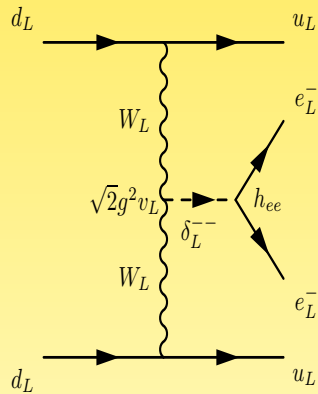
# Left-right symmetry



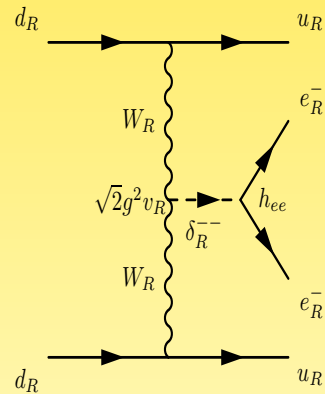
$$U_{ei}^2 m_i$$

$$\frac{S_{ei}^2}{M_i}$$

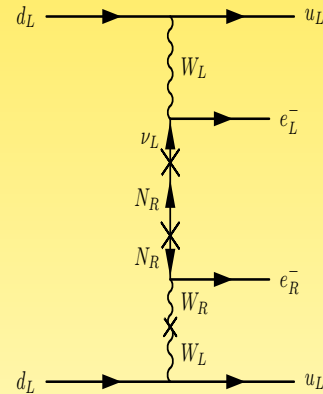
$$\frac{V_{ei}^2}{M_{W_R}^4 M_i}$$



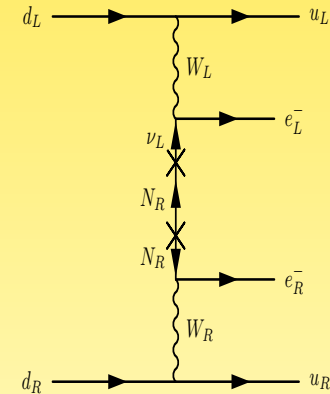
$$\frac{U_{ei}^2 m_i}{M_{\Delta_L}^2}$$



$$\frac{V_{ei}^2 M_i}{M_{W_R}^4 M_{\Delta_R}^2}$$

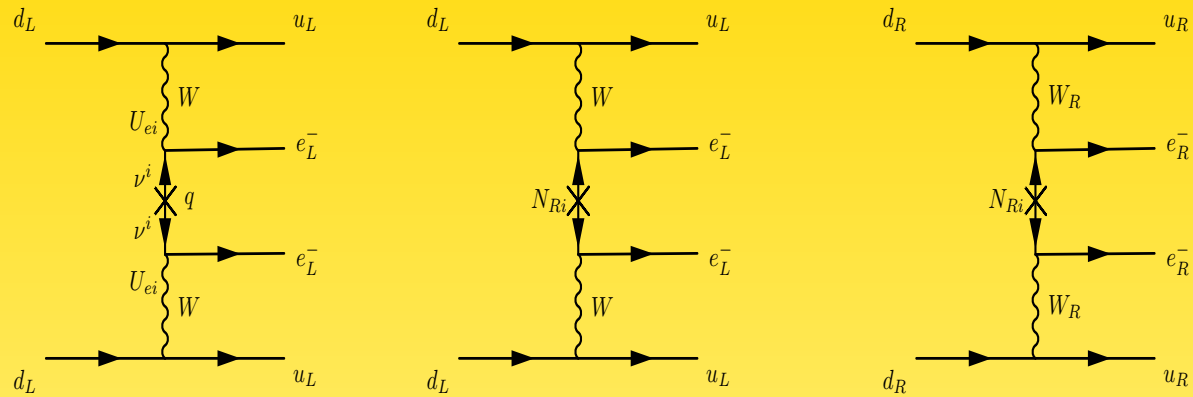


$$U_{ei} T_{ei} \tan \zeta$$



$$\frac{U_{ei} T_{ei}}{M_{W_R}^2}$$

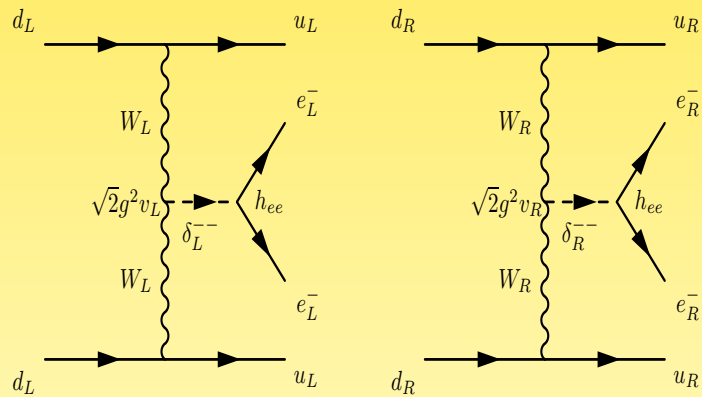
# Left-right symmetry



$0.4 \text{ eV}$

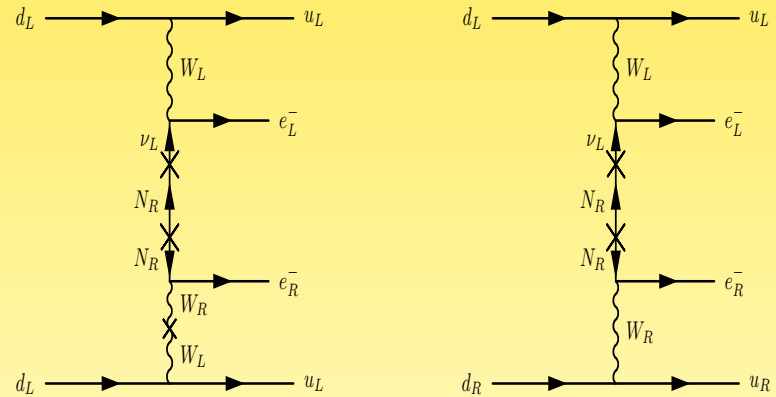
$2 \times 10^{-8} \text{ GeV}^{-1}$

$4 \times 10^{-16} \text{ GeV}^{-5}$



—

$10^{-15} \text{ GeV}^{-5}$

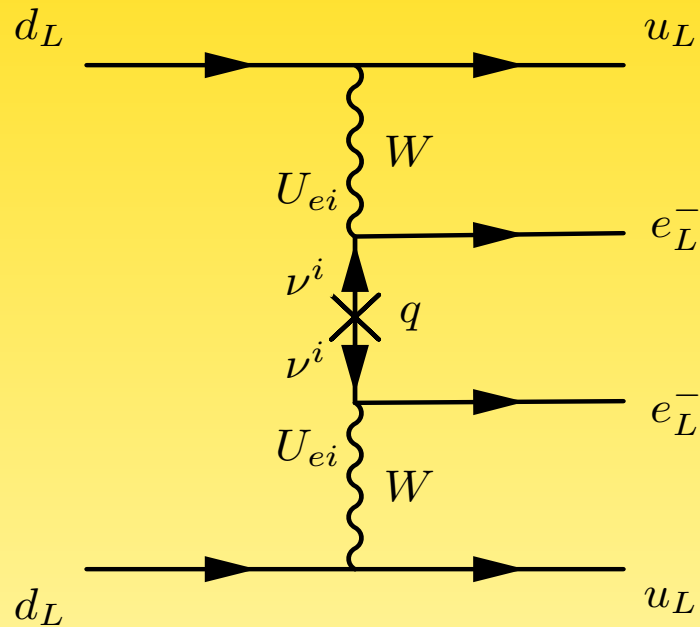


$6 \times 10^{-9}$

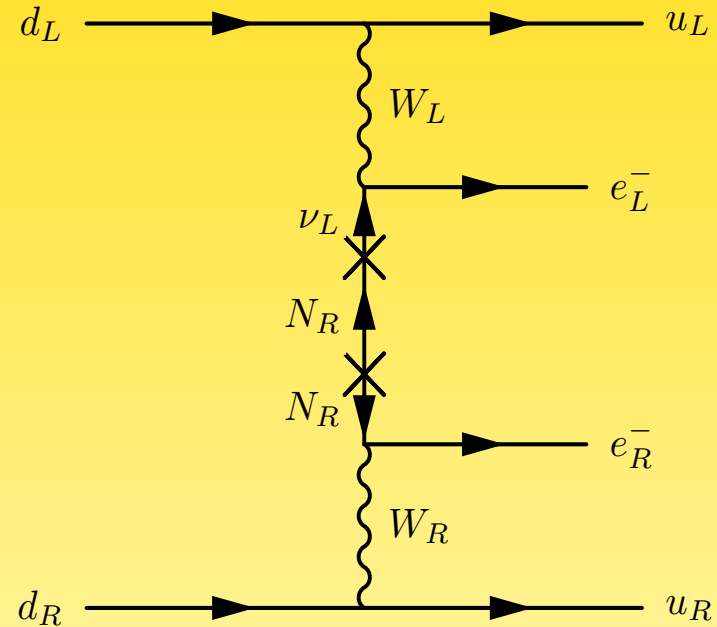
$1.4 \times 10^{-10} \text{ GeV}^{-2}$

## 2.) Distinguishing via decay products

Consider standard plus  $\lambda$ -mechanism



$$\frac{d\Gamma}{dE_1 dE_2 d\cos\theta} \propto (1 - \beta_1 \beta_2 \cos\theta)$$



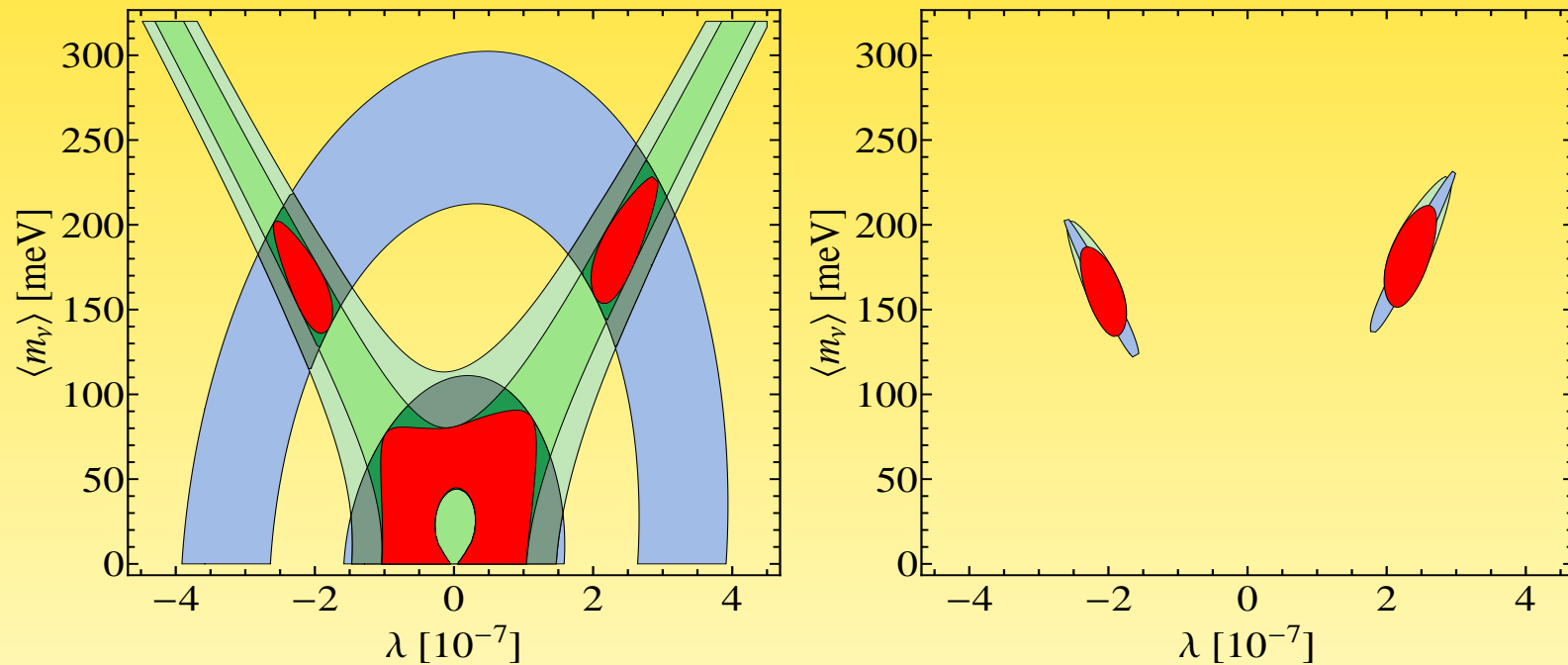
$$\frac{d\Gamma}{dE_1 dE_2 d\cos\theta} \propto (E_1 - E_2)^2 (1 + \beta_1 \beta_2 \cos\theta)$$

Arnold *et al.*, 1005.1241

## 2.) Distinguishing via decay products

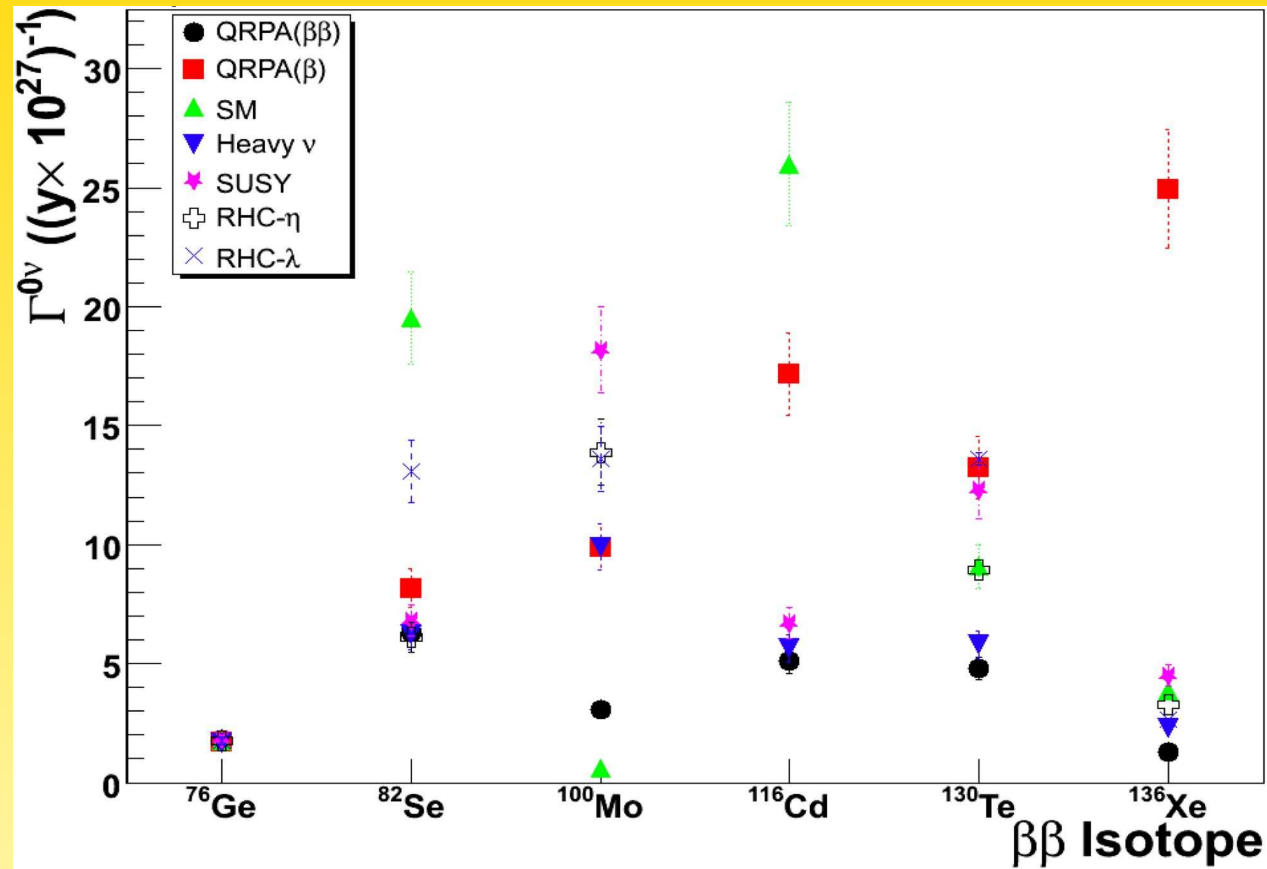
Defining asymmetries

$$A_\theta = (N_+ - N_-)/(N_+ + N_-) \text{ and } A_E = (N_{>} - N_{<})/(N_{>} + N_{<})$$





### 3.) Distinguishing via nuclear physics

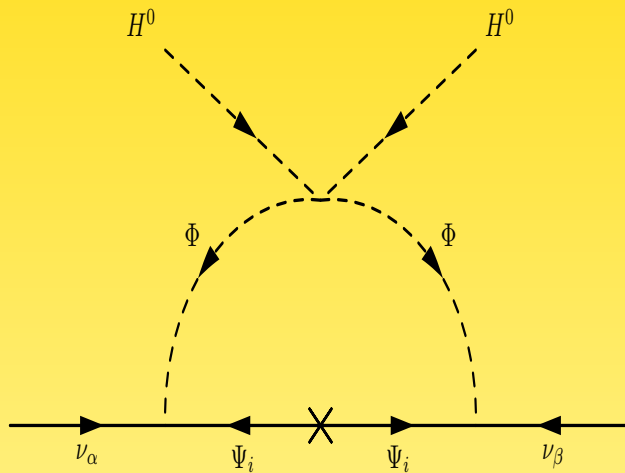


Gehman, Elliott, hep-ph/0701099

3 to 4 isotopes necessary to disentangle mechanism

## Direct versus indirect contribution

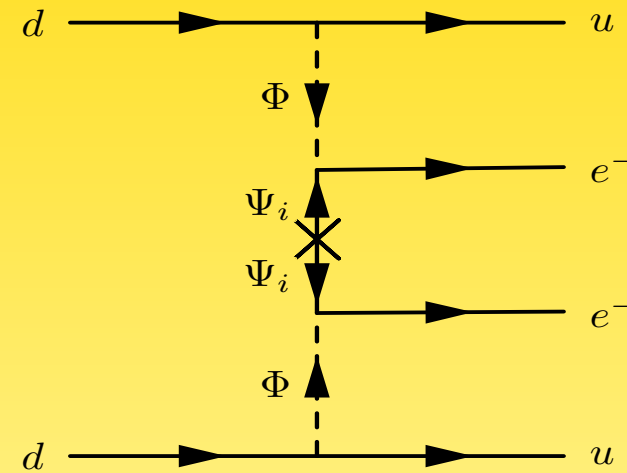
Example: introduce  $\Psi_i = (8, 1, 0)$  and  $\Phi = (8, 2, \frac{1}{2})$



1-loop  $m_\nu$

indirect contribution to  $0\nu\beta\beta$ :

$$\mathcal{A}_1 \simeq G_F^2 \frac{|m_{ee}|}{q^2}$$



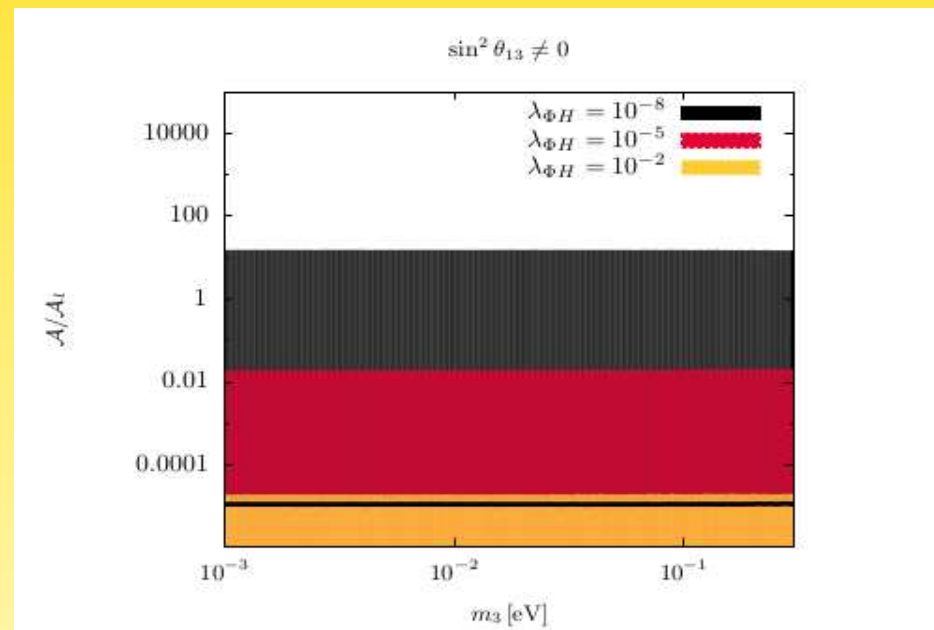
direct contribution to  $0\nu\beta\beta$ :

$$\mathcal{A} \simeq c_{ud}^2 \frac{y_{e\alpha}^2}{M_{\Psi_i} M_\Phi^4}$$

Choubey, Dürr, Mitra, W.R., JHEP **1205**

## Direct versus indirect contribution

new contribution can dominate over standard one:



## Do Dirac neutrinos imply that there is no Lepton Number Violation?

possible to construct  $U(1)_{B-L}$  model with  $\chi \sim -2$  and  $\phi \sim 4$

$$\mathcal{L} = \left( y_{\alpha\beta} \bar{L}_\alpha H \nu_{R,\beta} + \kappa_{\alpha\beta} \chi \bar{\nu}_{R,\alpha} \nu_{R,\beta}^c + h.c. \right) + \sum_{X=H,\phi,\chi} (\mu_X^2 |X|^2 + \lambda_X |X|^4) \\ + \lambda_{H\phi} |H|^2 |\phi|^2 + \lambda_{H\chi} |H|^2 |\chi|^2 + \lambda_{\chi\phi} |\chi|^2 |\phi|^2 - (\mu\phi\chi^2 + h.c.)$$

break it by allowing only scalar  $\phi$  to obtain VEV:

$\Rightarrow$  neutrinos are Dirac particles, and Lepton Number violated by 4 units!

$\Rightarrow$  neutrinoless double beta decay forbidden...

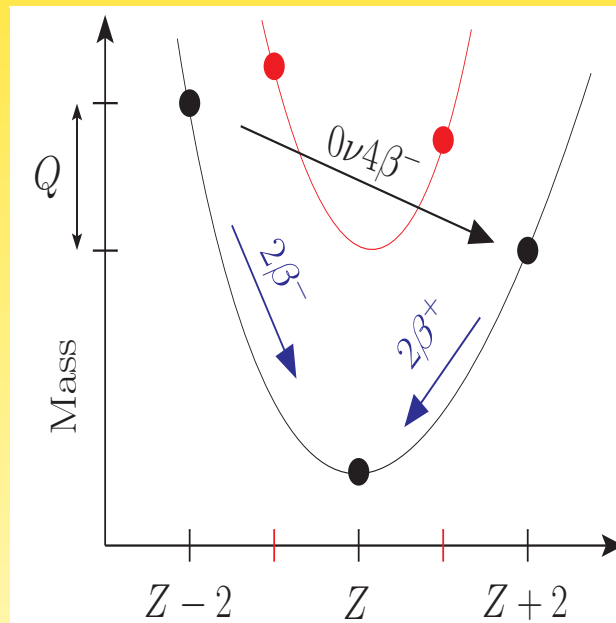
Heeck, W.R., EPL 103

## Do Dirac neutrinos mean there is no Lepton Number Violation?

Model based on gauged  $B - L$ , broken by 4 units

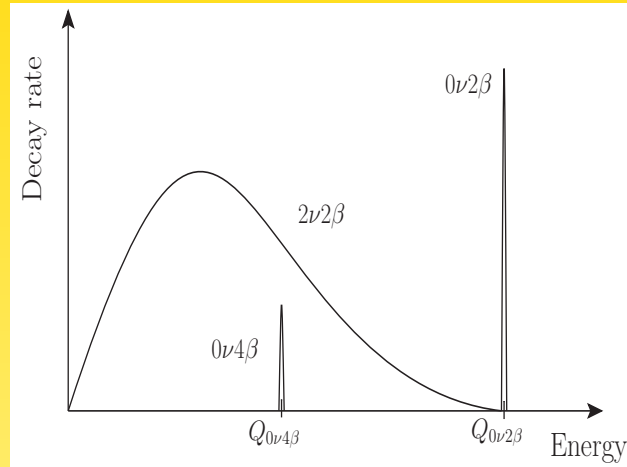
$\Rightarrow$  Neutrinos are Dirac particles,  $\Delta L = 2$  forbidden, but  $\Delta L = 4$  allowed...

$\Rightarrow$  observable: neutrinoless quadruple beta decay  $(A, Z) \rightarrow (A, Z + 4) + 4e^-$



Heeck, W.R., EPL 103

## Candidates for neutrinoless quadruple beta decay



	$Q_{0\nu 4\beta}$	Other decays	NA
${}^{96}_{40}\text{Zr} \rightarrow {}^{96}_{44}\text{Ru}$	0.629	$\tau_{1/2}^{2\nu 2\beta} \simeq 2 \times 10^{19}$	2.8
${}^{136}_{54}\text{Xe} \rightarrow {}^{136}_{58}\text{Ce}$	0.044	$\tau_{1/2}^{2\nu 2\beta} \simeq 2 \times 10^{21}$	8.9
${}^{150}_{60}\text{Nd} \rightarrow {}^{150}_{64}\text{Gd}$	2.079	$\tau_{1/2}^{2\nu 2\beta} \simeq 7 \times 10^{18}$	5.6

Heeck, W.R., EPL **103**

## Summary

**Chi l'ha visto ?**



Ettore Majorana, ordinario di fisica teorica all'Università di Napoli, è misteriosamente scomparso dagli ultimi di marzo. Di anni 31, alto metri 1,70, snello, con capelli neri, occhi scuri, una lunga cicatrice sul dorso di una mano. Chi ne sapesse qualcosa è pregato di scrivere al R. P. E. Maria-necci, Viale Regina Margherita 66 - Roma.

## Mass terms

$$\mathcal{L} = m_\nu \bar{\nu}_L \nu_R + h.c. = m_\nu (\bar{\nu}_L \nu_R + \bar{\nu}_R \nu_L)$$

both chiralities a must for mass term!

Possibilities for  $\bar{L}R$ :

(i)  $\nu_R$  independent of  $\nu_L$ : **Dirac particle**

(ii)  $\nu_R = (\nu_L)^c$ : **Majorana particle**

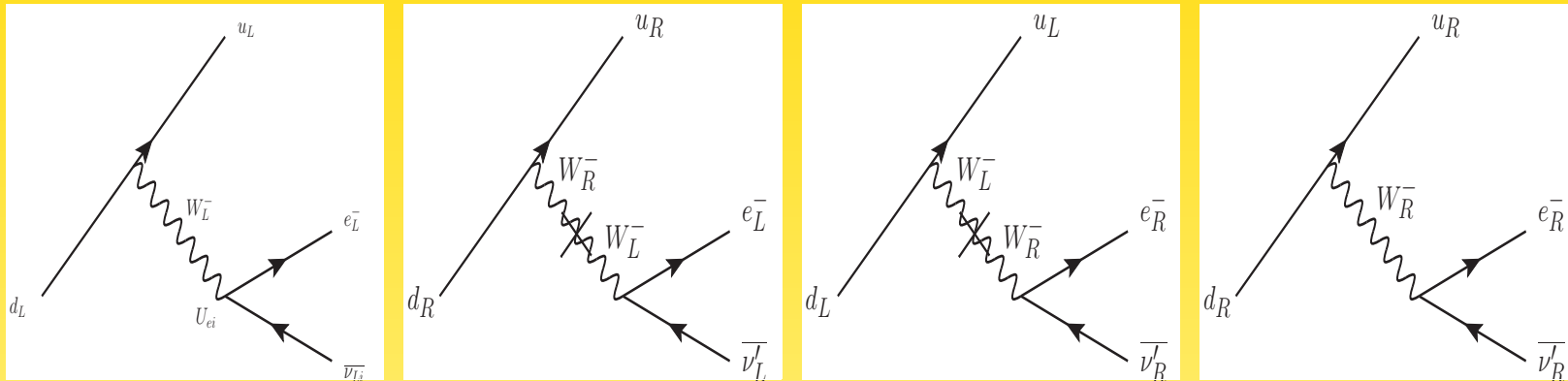
$$\Rightarrow \nu^c = (\nu_L + \nu_R)^c = (\nu_L)^c + (\nu_R)^c = \nu_R + \nu_L = \nu : \nu^c = \nu$$

$\Rightarrow$  Majorana fermion is identical to its antiparticle, a truly neutral particle

$\Rightarrow$  all additive quantum numbers ( $Q, L, B, \dots$ ) are zero



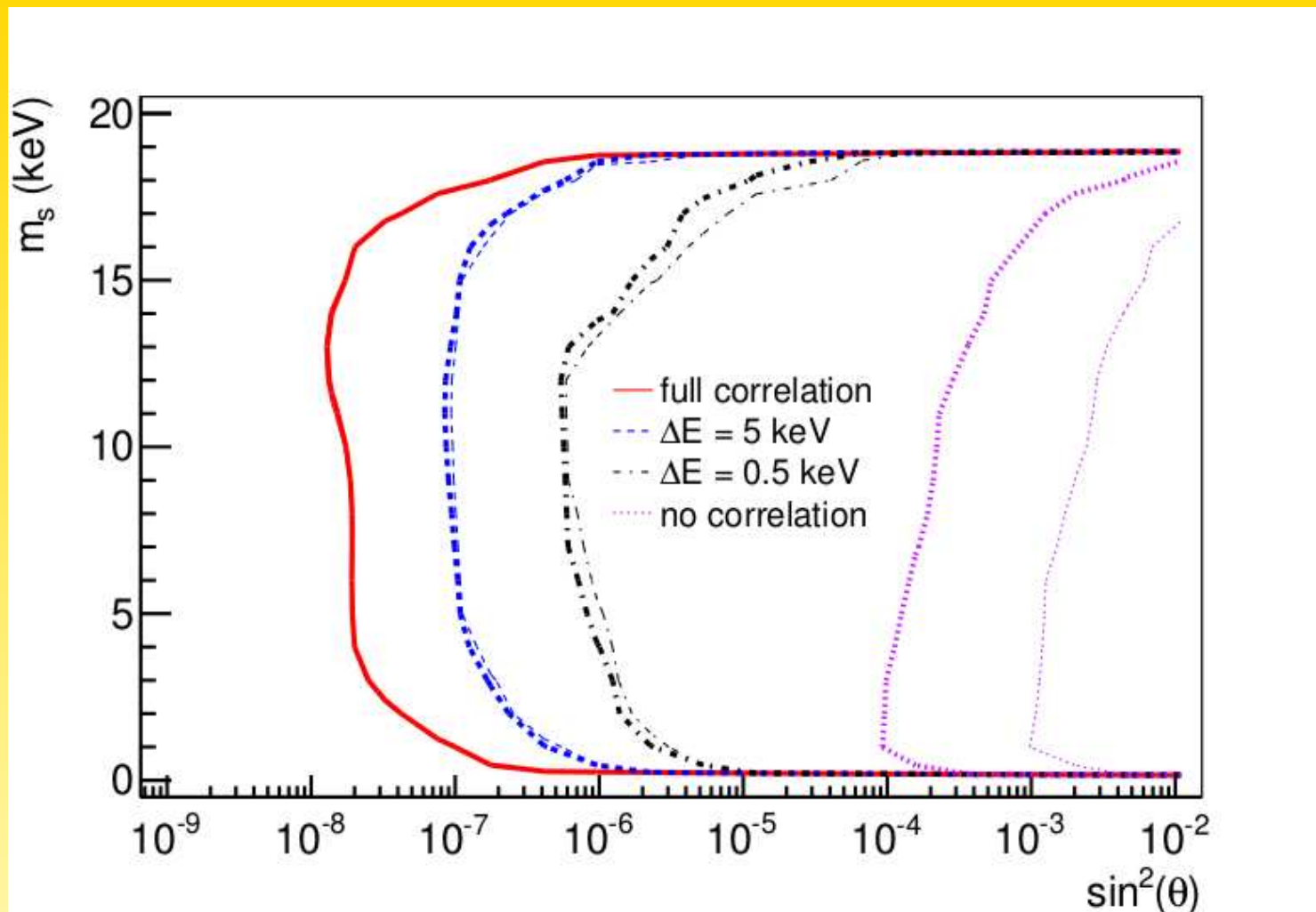
## KATRIN and right-handed currents



- left-handed contribution
- right-handed contribution
- interference contribution

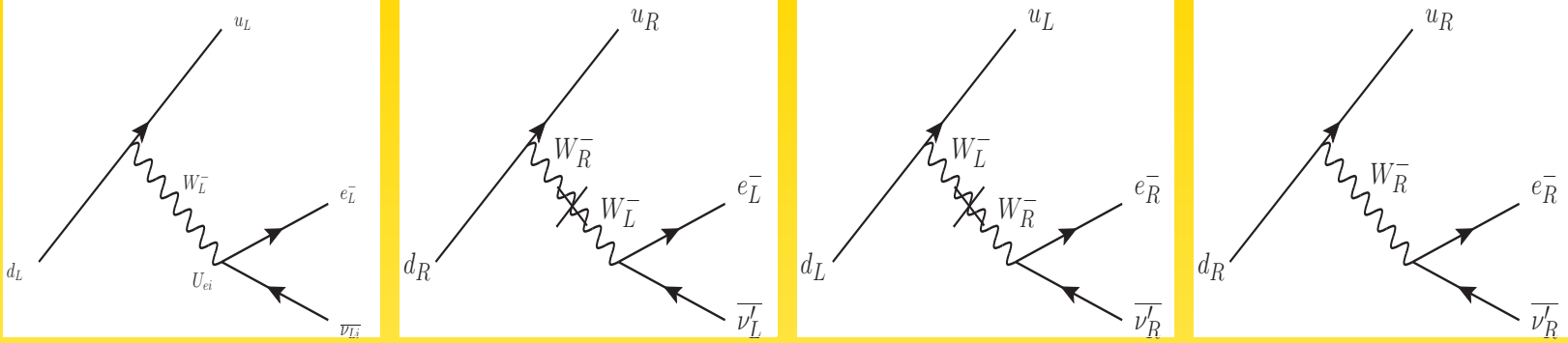
Neutrino masses up to  $m = 18.6$  keV testable  $\Leftrightarrow$  Warm Dark Matter!

(see lecture by Kevork Abazajian)



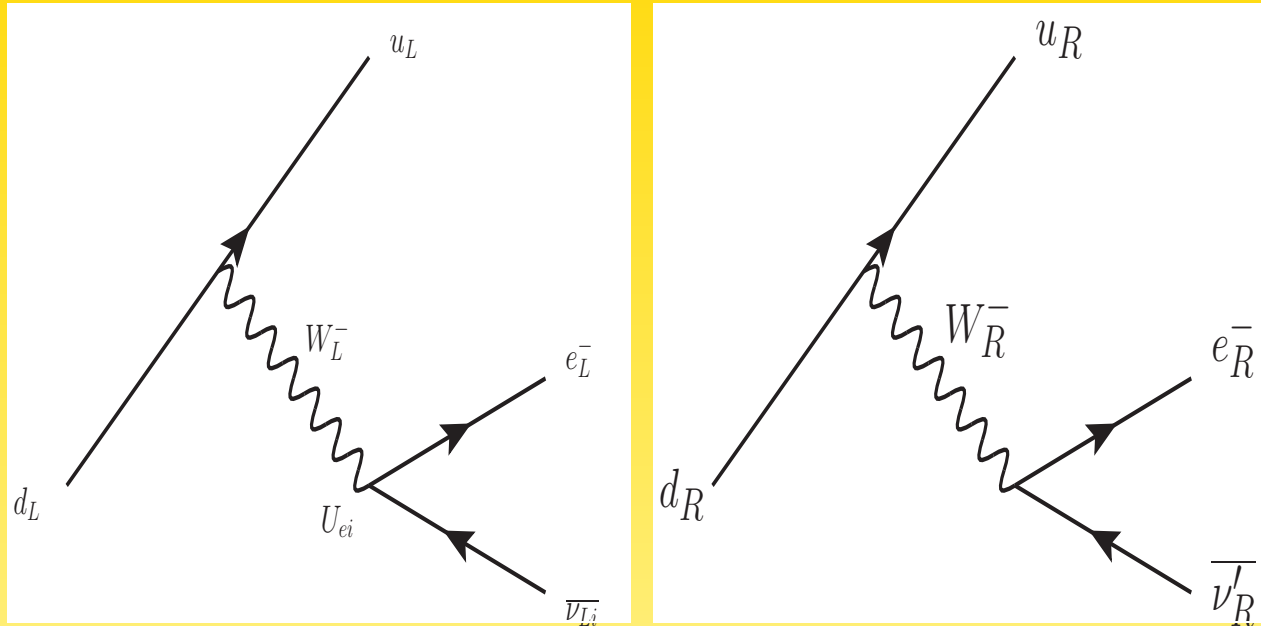
$\Rightarrow$  mixing down to  $10^{-7}$  in reach!?

Mertens *et al.*, 1409.0920



$$\begin{aligned}
& \left( \frac{d\Gamma}{dE} \right)_{LL} = K' (E + m_e) p_e X [1 + 2C \tan \xi] \\
& \times \left[ |U_{ei}|^2 \sqrt{X^2 - m_i^2} \Theta(X - m_i) + |S_{ei}|^2 \sqrt{X^2 - M_i^2} \Theta(X - M_i) \right] \\
& \left( \frac{d\Gamma}{dE} \right)_{RR} \simeq K' (E + m_e) p_e X \left[ \frac{m_{W_L}^4}{m_{W_R}^4} + \tan^2 \xi + 2C \frac{m_{W_L}^2}{m_{W_R}^2} \tan \xi \right] \\
& \times |V_{ei}|^2 \sqrt{X^2 - M_i^2} \Theta(X - M_i) \\
& \left( \frac{d\Gamma}{dE} \right)_{LR} = -2K' m_e p_e \text{Re} \left\{ \left[ \left( \frac{m_{W_L}}{m_{W_R}} \right)^2 + C \tan \xi \right] \right. \\
& \times \left. \left[ U_{ei} T_{ei} m_i \sqrt{X^2 - m_i^2} \Theta(X - m_i) + S_{ei} V_{ei} M_i \sqrt{X^2 - M_i^2} \Theta(X - M_i) \right] \right\} \\
& \text{with } X = E_0 - E
\end{aligned}$$

Focus for simplicity on



total contribution of keV neutrino with mass  $M$  to beta decay:

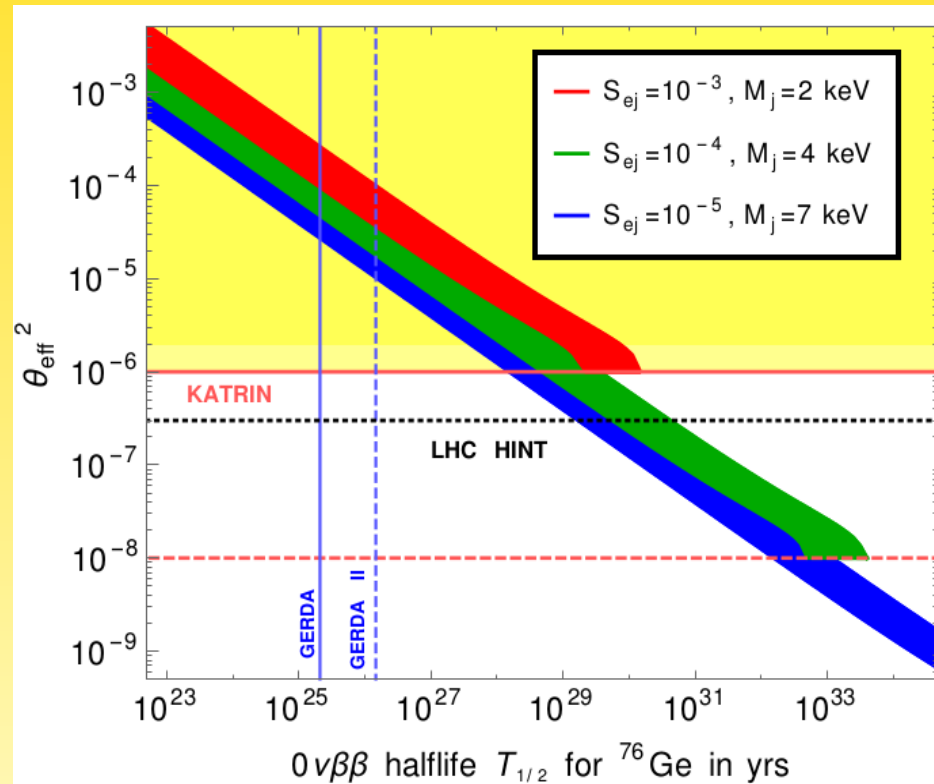
$$\theta_{\text{eff}}^2 \simeq |S_{ej}|^2 + 1.1 \times 10^{-6} |V_{ej}|^2 \left( \frac{2.5 \text{ TeV}}{m_{W_R}} \right)^4$$

and note that  $M$  does  $0\nu\beta\beta$  with amplitude  $\propto |V_{ej}|^2 (m_W/m_{W_R})^4 M$

$\Rightarrow$  connection to  $0\nu\beta\beta$  constraints!

connection to  $0\nu\beta\beta$  constraints:

$$\theta_{\text{eff}}^2 = |S_{ej}|^2 + \frac{m_e}{M_j} \left[ |\mathcal{M}_\nu^{0\nu}|^{-2} (G_{01}^{0\nu})^{-1} (T_{1/2}^{0\nu})^{-1} - |S_{ej}^2 M_j / m_e|^2 \right]^{\frac{1}{2}}$$

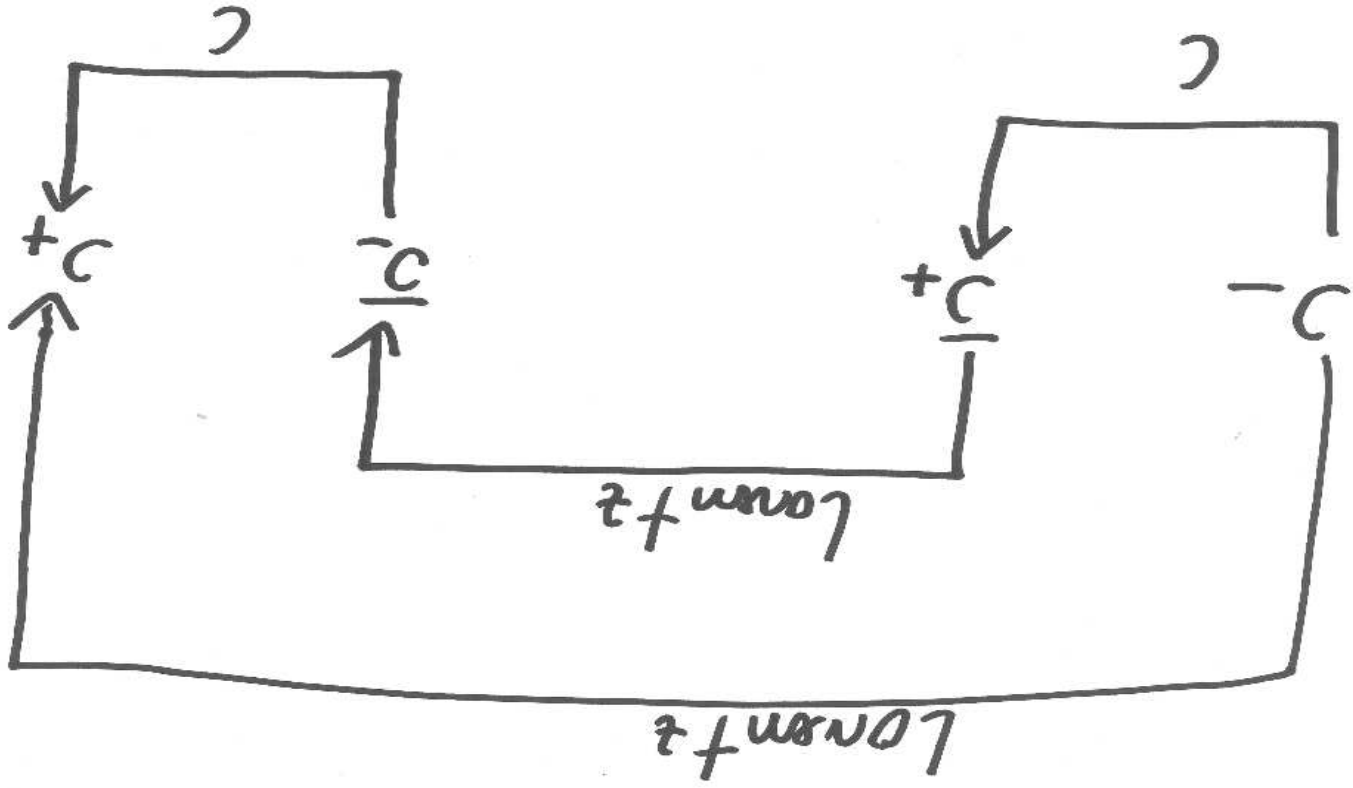


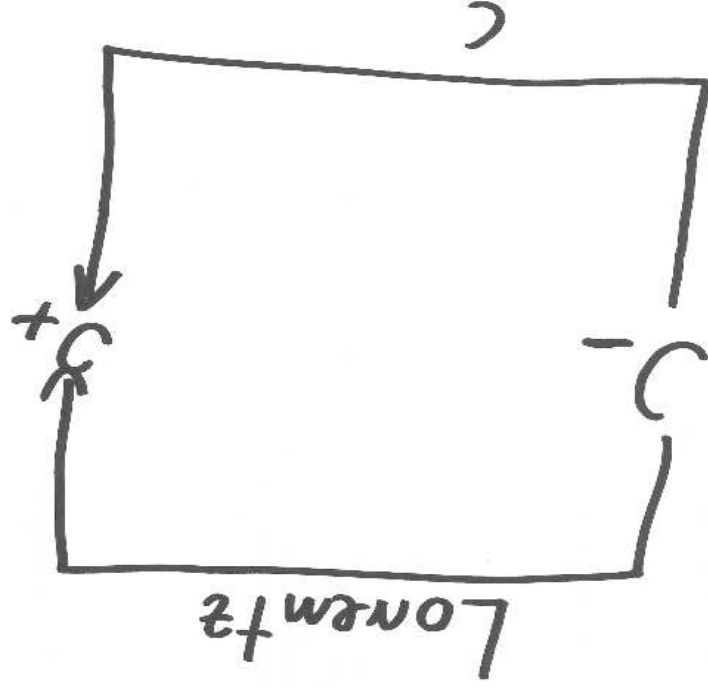
Barry, Heeck, W.R., JHEP 1407

## How the additional interactions save the day

- double beta decay without RHC:  $\theta^2 M = 7 \times 10^{-10} \text{ keV} = 70 \mu\text{eV}$
- double beta decay with RHC:  $(m_{W_L}/m_{W_R})^4 |V_{ei}|^2 M = 8 \text{ meV}$
- decay:  $\frac{\Gamma_{\text{RHC}}(N_j \rightarrow \bar{\nu}\gamma)}{\Gamma_{\text{SM}}(N_j \rightarrow \nu\gamma)} \simeq \frac{m_{W_L}^4 |S_{ei}|^2}{m_{W_R}^4 |T_{ei}|^2} \simeq \frac{m_{W_L}^4}{m_{W_R}^4}$
- beta decay:  $\theta_{\text{eff}}^2 \simeq |S_{ej}|^2 + 1.1 \times 10^{-6} |V_{ej}|^2 \left(\frac{2.5 \text{ TeV}}{m_{W_R}}\right)^4 > |S_{ej}|^2$

Dirac particles:  $v \neq v^c$  helicity states  $\pm$ : 4 d.o.f.





Majorana particles:  $\nu = \nu^c$  helicity states  $\pm$ : 2 d.o.f.



Mass term for Majorana particles:

$$\mathcal{L}_M = \frac{1}{2} \bar{\nu} m_\nu \nu = \frac{1}{2} \overline{\nu_L + (\nu_L)^c} m_\nu (\nu_L + (\nu_L)^c) = \frac{1}{2} \overline{\nu_L} m_\nu \nu_L^c + h.c.$$

- Majorana mass term
- $\mathcal{L}_M \propto \nu^\dagger \nu^* \Rightarrow$  NOT invariant under  $\nu \rightarrow e^{i\alpha} \nu$   
 $\Rightarrow$  breaks Lepton Number by 2 units

## Dirac vs. Majorana

in  $V - A$  theories: observable difference always suppressed by  $(m/E)^2$

- suppose beam from  $\pi^+$  decays:  $\pi^+ \rightarrow \mu^+ \nu_\mu$  (left-handed)
- can we observe  $\bar{\nu}_\mu + p \rightarrow n + \mu^+$  ? (right-handed)
- chirality is not a good quantum number: “spin flip”
- emitted  $\nu_\mu$  (negative helicity) is not purely left-handed:

$$u_\downarrow(p) = u_L^{(m=0)}(p) + \frac{m}{2E} u_R^{(m=0)}(-p)$$

- $P_R u_\downarrow \neq 0$  and  $\mu^+$  can be produced if  $m \neq 0$  and “ $u \propto v$ ” ( $\leftrightarrow$  Majorana!)

$\Rightarrow$  amplitude  $\propto (m/E) \Rightarrow$  probability  $\propto (m/E)^2$

$\Rightarrow$  only  $N_A$  can save the day!

plus uncertainty due to model details:

- Short range correlations:
  - Jastrow function
  - Unitary Correlation Operator method
  - Coupled Cluster method
- model parameters
  - correlated: e.g.,  $g_A$ ,  $g_{pp}$ , SRC
  - uncorrelated: e.g., model space of single particle states
  - some include errors, some don't...

$$\mathcal{M}^{0\nu} = \left( \frac{g_A}{1.25} \right)^2 \left( \mathcal{M}_{\text{GT}}^{0\nu} - \frac{g_V^2}{g_A^2} \mathcal{M}_{\text{F}}^{0\nu} \right)$$

with

$$\mathcal{M}_{\text{GT}}^{0\nu} = \langle f | \sum_{lk} \sigma_l \sigma_k \tau_l^- \tau_k^- H_{\text{GT}}(r_{lk}, E_a) | i \rangle$$

$$\mathcal{M}_{\text{F}}^{0\nu} = \langle f | \sum_{lk} \tau_l^- \tau_k^- H_{\text{F}}(r_{lk}, E_a) | i \rangle$$

- $r_{lk} \simeq 1/p \simeq 1/(0.1 \text{ GeV})$  distance between the two decaying neutrons
- $E_a$  average energy
- *sum over all multipolarities!*
- 'neutrino potential'  $H_{\text{GT,F}}(r_{lk}, E_a)$  integrates over the virtual neutrino momenta
- (if heavy particles exchanged: nucleon structure:  $g_A = g_A(0)/(1 - q^2/M_A^2)^2$ )

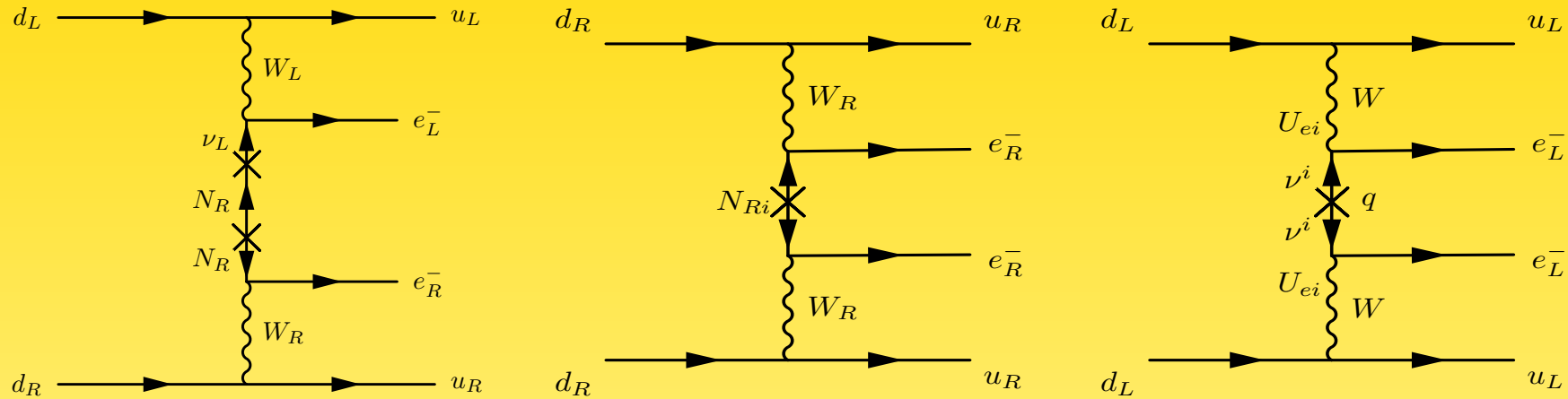
The  $2\nu\beta\beta$  matrix elements can be written as

$$\mathcal{M}_{\text{GT}}^{2\nu} = \sum_n \frac{\langle f | \sum_a \sigma_a \tau_a^- | n \rangle \langle n | \sum_b \sigma_b \tau_b^- | i \rangle}{E_n - (M_i - M_f)/2}$$

$$\mathcal{M}_{\text{F}}^{2\nu} = \sum_n \frac{\langle f | \sum_a \tau_a^- | n \rangle \langle n | \sum_b \tau_b^- | i \rangle}{E_n - (M_i - M_f)/2}$$

- no direct connection to  $0\nu\beta\beta$ ...
- *sum over  $1^+$  states* (low momentum transfer)
- adjust some parameters to reproduce  $2\nu\beta\beta$ -rates
- $2\nu\beta\beta$  observed in  $^{48}\text{Ca}$ ,  $^{76}\text{Ge}$ ,  $^{82}\text{Se}$ ,  $^{96}\text{Zr}$ ,  $^{100}\text{Mo}$  (plus exc. state),  $^{116}\text{Cd}$ ,  $^{128}\text{Te}$ ,  $^{130}\text{Te}$ ,  $^{136}\text{Xe}$ ,  $^{150}\text{Nd}$  (plus exc. state) with half-lives from  $10^{18}$  to  $10^{24}$  yrs
- can partly be tested with forward angle (= low momentum transfer)  $(p, n)$  charge exchange reactions

## Mixed Diagrams can dominate



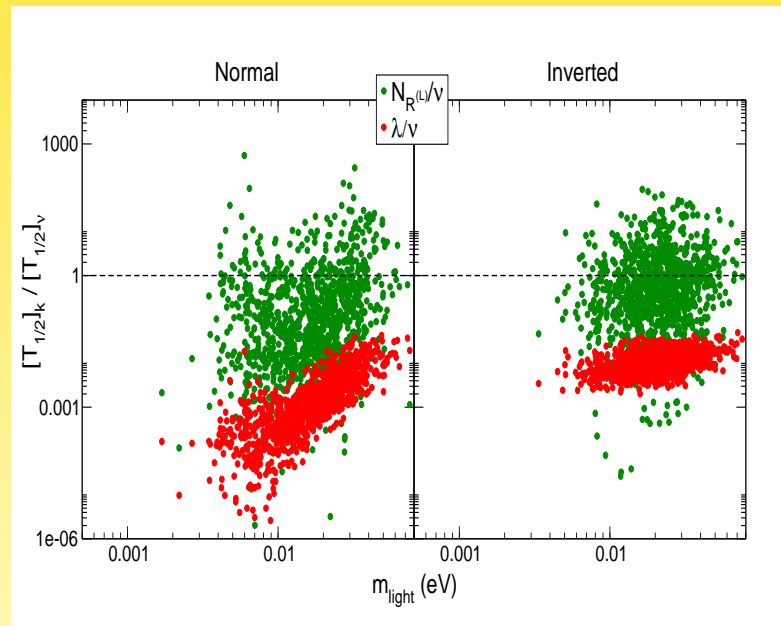
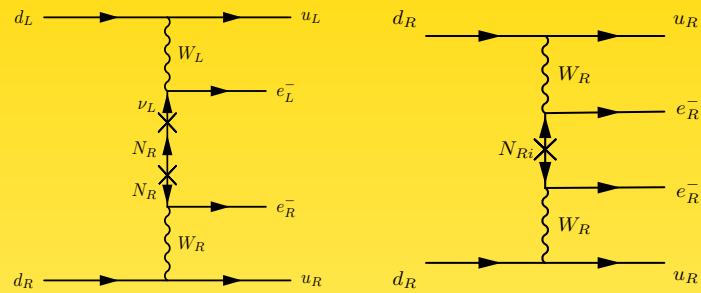
$$A_\lambda \sim \left( \frac{m_W}{M_{W_R}} \right)^2 \frac{UT}{q} \quad A_{N_R} \sim \left( \frac{m_W}{M_{W_R}} \right)^4 \frac{V^2}{M_R} \quad A_\nu \sim U^2 \frac{m_i}{q^2}$$

with  $T \simeq \sqrt{\frac{m_\nu}{M_R}} \sim 10^{-7}$  (or huge enhancements up to  $10^{-2}$ )

$$\Rightarrow \frac{A_\lambda}{A_{N_R}} \simeq \frac{M_R}{q} \left( \frac{M_{W_R}}{m_W} \right)^2 T \simeq 10^5 T$$

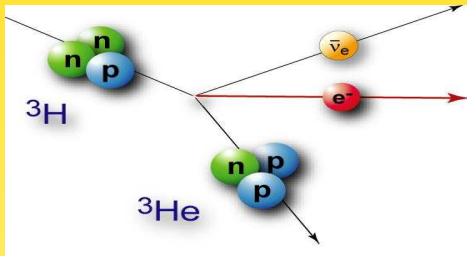
Barry, W.R., JHEP **1309**

# Mixed Diagrams can dominate



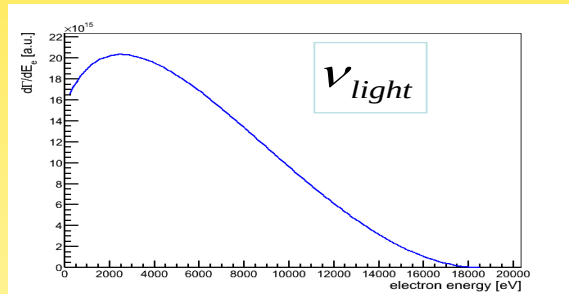
Barry, W.R., 1303.6324

# Imprint of keV neutrinos on $\beta$ -spectrum

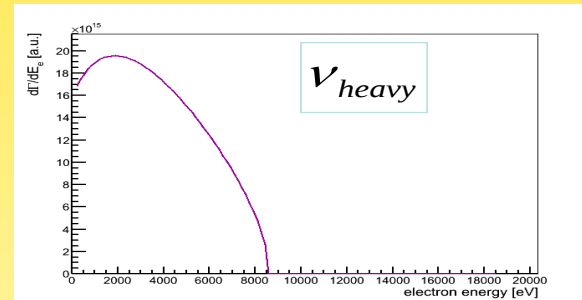


$$\begin{pmatrix} \nu_e \\ \nu_s \end{pmatrix} = \begin{pmatrix} \cos \theta & -\sin \theta \\ \sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_{light} \\ \nu_{heavy} \end{pmatrix}$$

$\cos^2(\theta)$

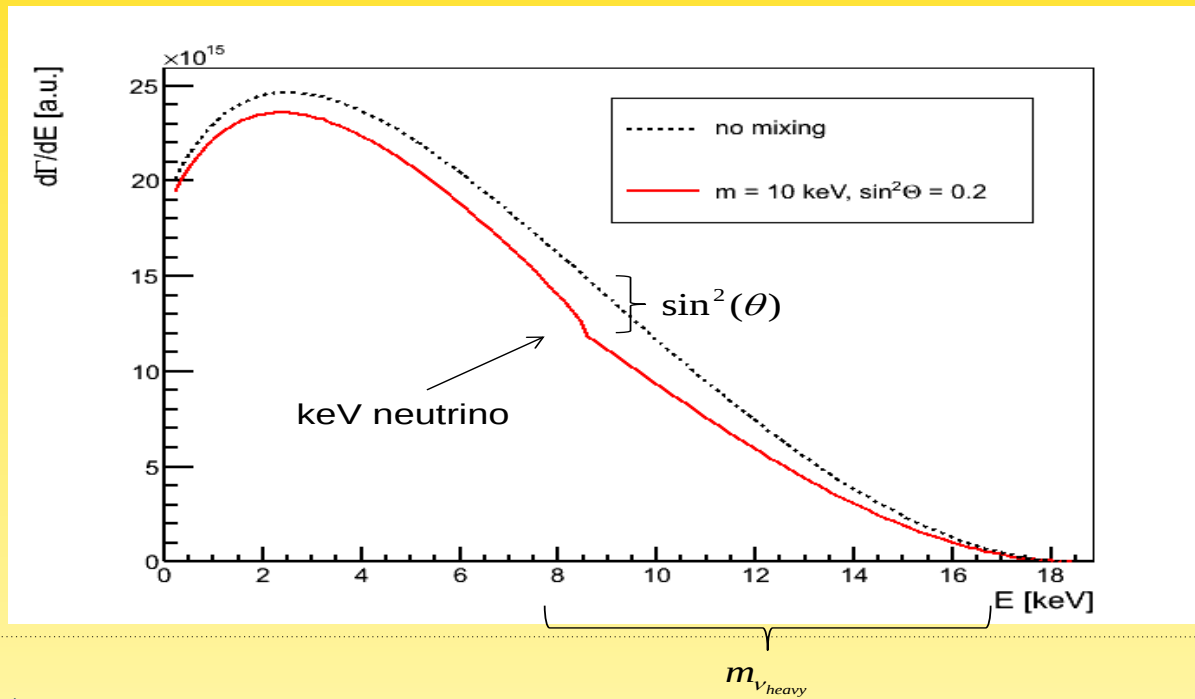


+  $\sin^2(\theta)$





# Imprint of keV neutrinos on $\beta$ -spectrum



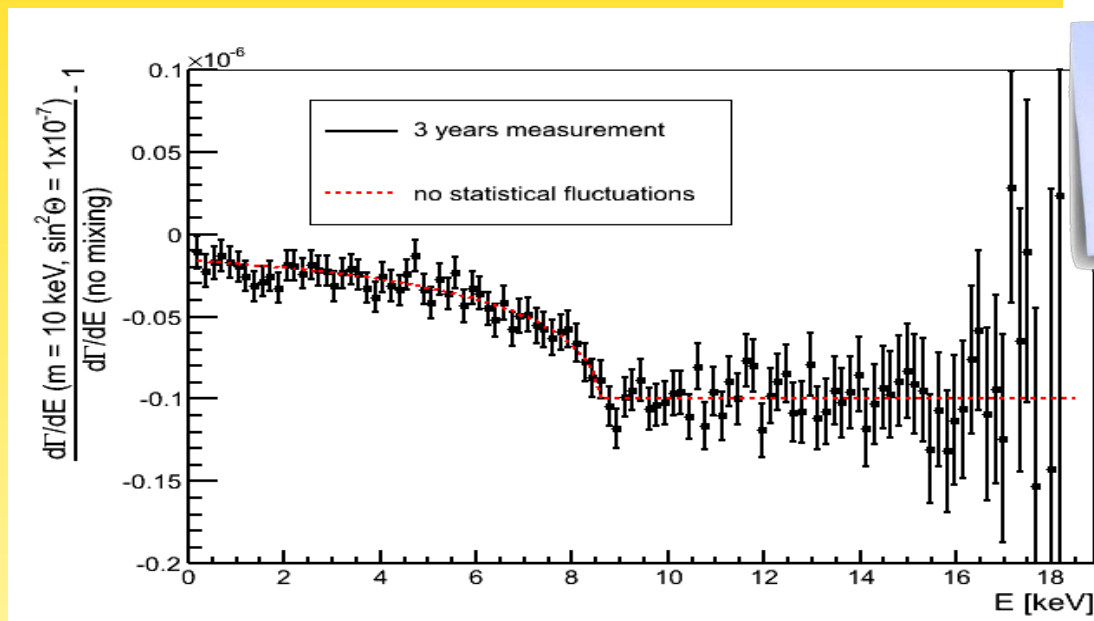
Susanne Mertens

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Mertens *et al.*, 1409.0920

# Imprint of keV neutrinos on $\beta$ -spectrum

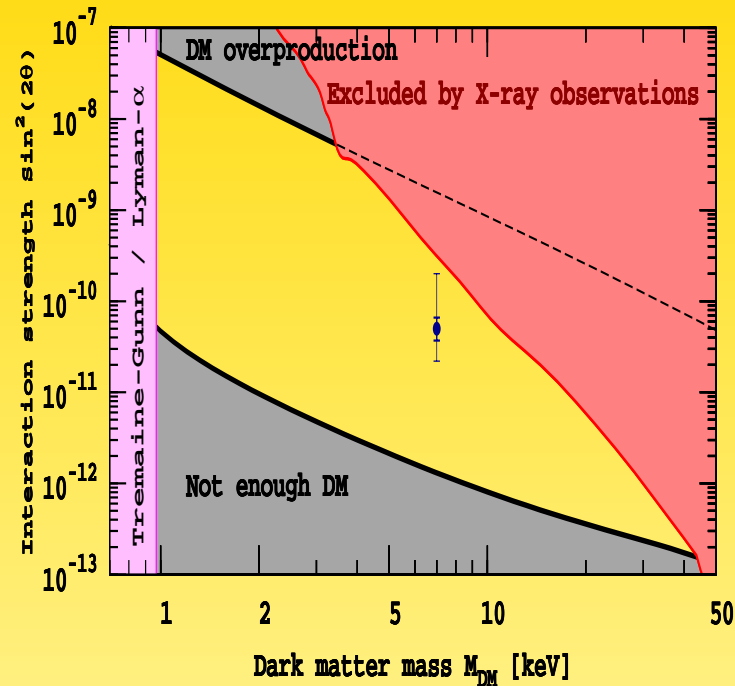


Susanne Mertens

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Mertens *et al.*, 1409.0920



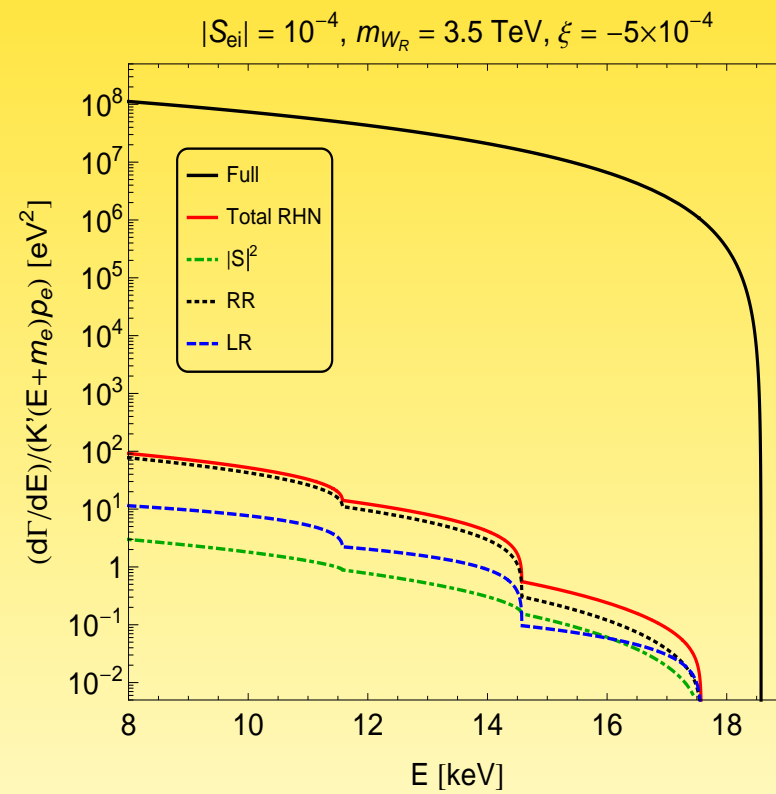
⇒ mixing down to  $10^{-7}$  in reach!?

hard for KATRIN to see something...

...can increase signal with additional interactions, e.g. right-handed currents

in total one has

$$\frac{d\Gamma}{dE} = \left(\frac{d\Gamma}{dE}\right)_{LL} + \left(\frac{d\Gamma}{dE}\right)_{RR} + \left(\frac{d\Gamma}{dE}\right)_{LR}$$

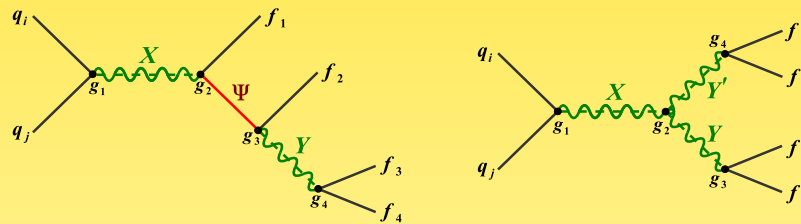


# Observation of LNV at LHC implies washout effects in early Universe!

Example TeV-scale  $W_R$ : leading to washout  $e_R^\pm e_R^\pm \rightarrow W_R^\pm W_R^\pm$  and  $e_R^\pm W_R^\mp \rightarrow e_R^\mp W_R^\pm$ . Further,  $e_R^\pm W_R^\mp \rightarrow e_R^\mp W_R^\pm$  stays long in equilibrium

(Frere, Hambye, Vertongen; Bhupal Dev, Lee, Mohapatra; U. Sarkar *et al.*)

More model-independent (Deppisch, Harz, Hirsch):

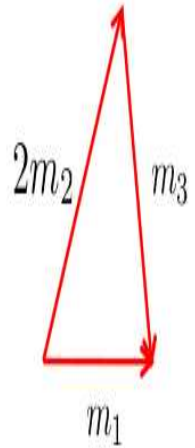


$$\text{washout: } \log_{10} \frac{\Gamma_W(qq \rightarrow \ell^+ \ell^+ qq)}{H} \gtrsim 6.9 + 0.6 \left( \frac{M_X}{\text{TeV}} - 1 \right) + \log_{10} \frac{\sigma_{\text{LHC}}}{\text{fb}}$$

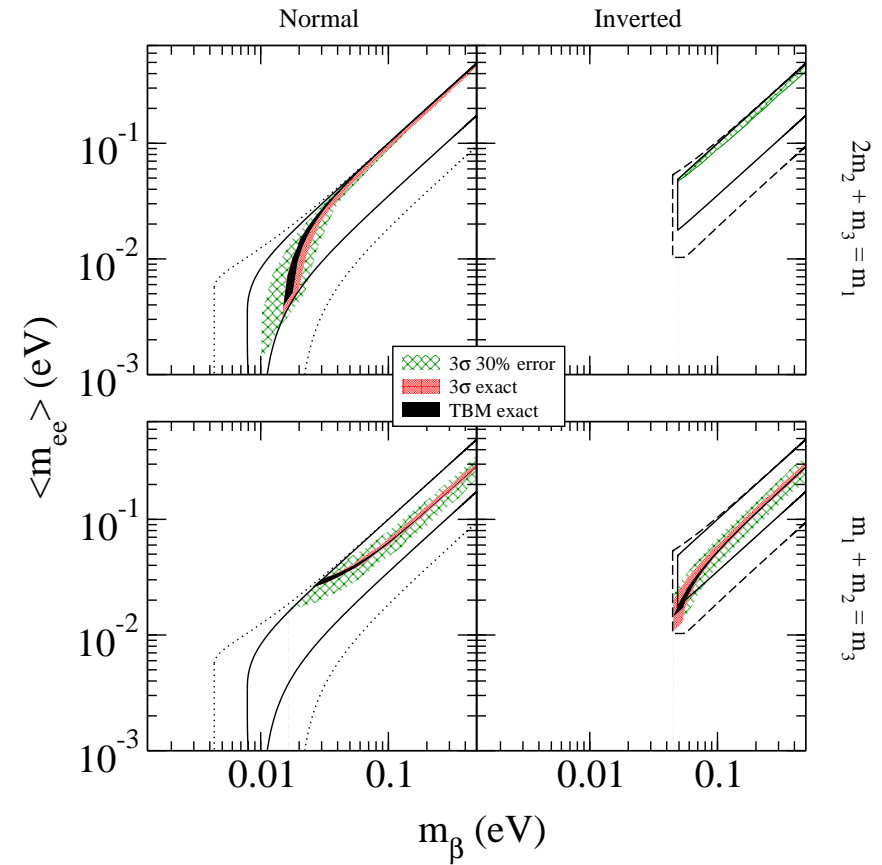
(TeV- $0\nu\beta\beta$ , LFV and  $Y_B$ : Deppisch, Harz, Huang, Hirsch, Päs)

$\leftrightarrow$  post-Sphaleron mechanisms,  $\tau$  flavor effects,...

# Flavor Symmetry Models: sum-rules



Sum-rule	Flavour symmetry
$2m_2 + m_3 = m_1$	$A_4, T', (S_4)$
$m_1 + m_2 = m_3$	$S_4, (A_4)$
$\frac{2}{m_2} + \frac{1}{m_3} = \frac{1}{m_1}$	$A_4, T'$
$\frac{1}{m_1} + \frac{1}{m_2} = \frac{1}{m_3}$	$S_4$



constraints on masses and Majorana phases

Barry, W.R., Nucl. Phys. **B842**

## Predictions of $SO(10)$ theories

Yukawa structure of  $SO(10)$  models depends on Higgs representations

$$10_H (\leftrightarrow H), \overline{126}_H (\leftrightarrow F), 120_H (\leftrightarrow G)$$

Gives relation for mass matrices:

$$m_{\text{up}} \propto r(H + sF + it_u G)$$

$$m_{\text{down}} \propto H + F + iG$$

$$m_D \propto r(H - 3sF + it_D G)$$

$$m_\ell \propto H - 3F + it_l G$$

$$M_R \propto r_R^{-1} F$$

Numerical fit including RG, Higgs,  $\theta_{13}$

$$10_H + \overline{126}_H: 19 \text{ free parameters}$$

$$10_H + \overline{126}_H + 120_H: 18 \text{ free parameters}$$

$$20 \text{ (19) observables to be fitted}$$

## Predictions of $SO(10)$ theories

Model	Fit	$ m_{ee} $ [meV]	$m_0$ [meV]	$M_3$ [GeV]	$\chi^2$
$10_H + \overline{126}_H$	NH	0.49	2.40	$3.6 \times 10^{12}$	23.0
$10_H + \overline{126}_H + SS$	NH	0.44	6.83	$1.1 \times 10^{12}$	3.29
$10_H + \overline{126}_H + 120_H$	NH	2.87	1.54	$9.9 \times 10^{14}$	11.2
$10_H + \overline{126}_H + 120_H + SS$	NH	0.78	3.17	$4.2 \times 10^{13}$	$6.9 \times 10^{-6}$
$10_H + \overline{126}_H + 120_H$	IH	35.52	30.2	$1.1 \times 10^{13}$	13.3
$10_H + \overline{126}_H + 120_H + SS$	IH	24.22	12.0	$1.2 \times 10^{13}$	0.6

Dueck, W.R., JHEP **1309**