



# Benchmarking of the FLUKA code

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HSS06

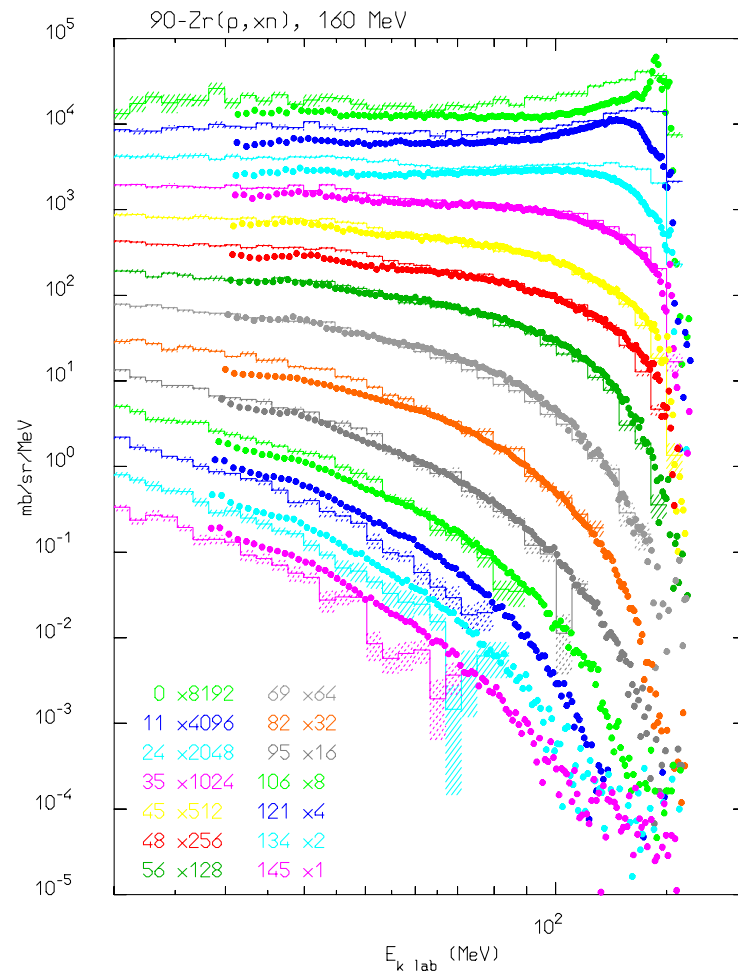


# ***Thin target***

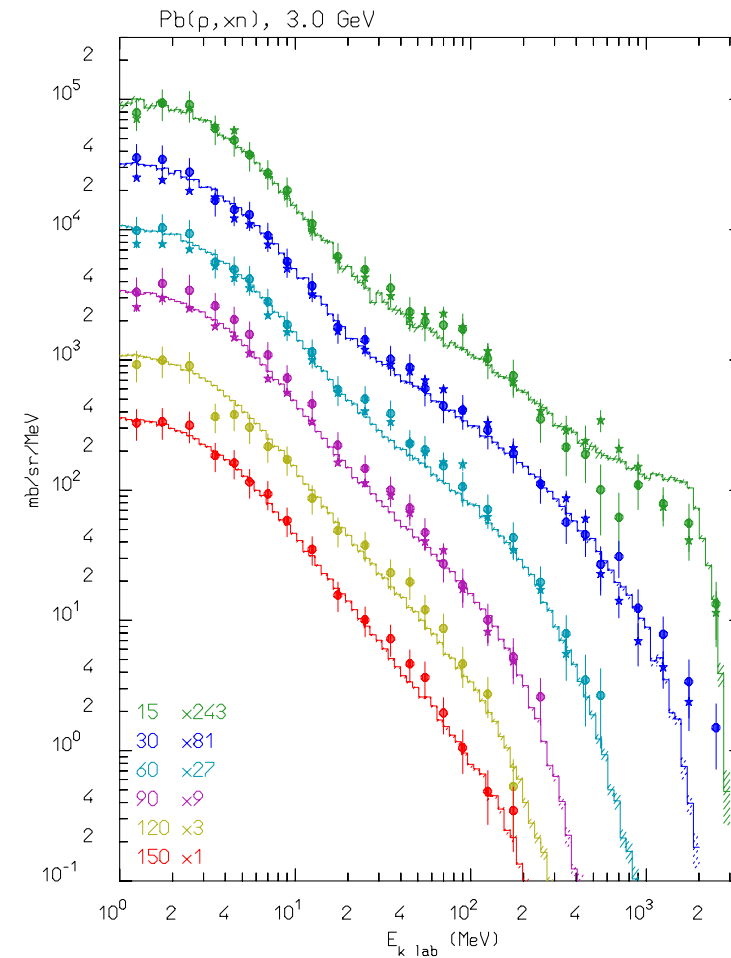
# Neutron production at low-intermediate energies

where most of the interactions occur..

160 MeV protons on Zr  
Phys. Rev. C41,2010 (1990)

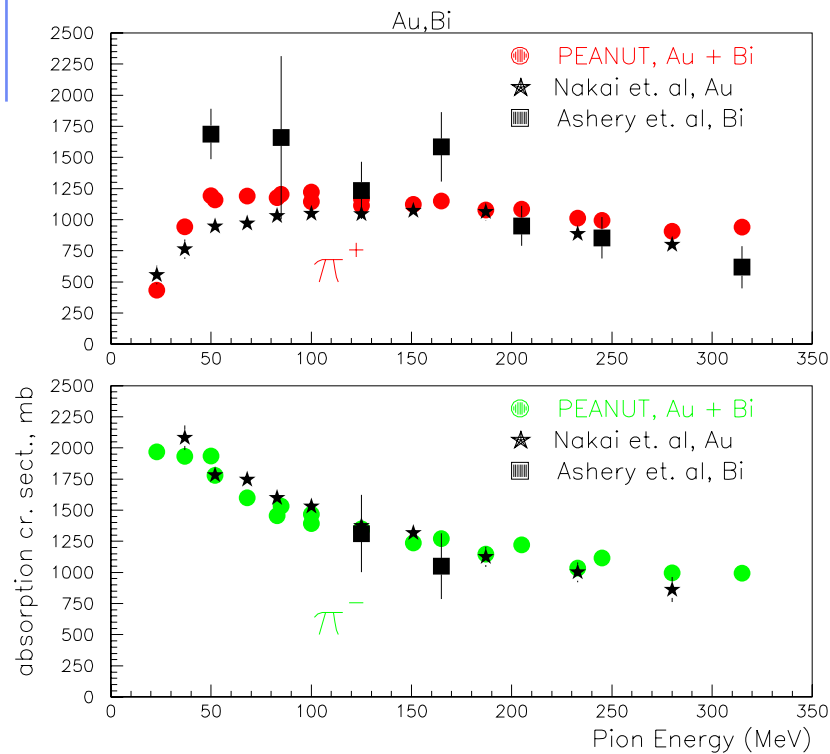


3 GeV p on Pb  
Nuc. Sci. Technol. 32, 827 (1995)



# Pion absorption

Pion absorption cross section  
on Gold and Bismuth in the  
 $\Delta$  resonance region  
(multibody absorption in  
PEANUT)

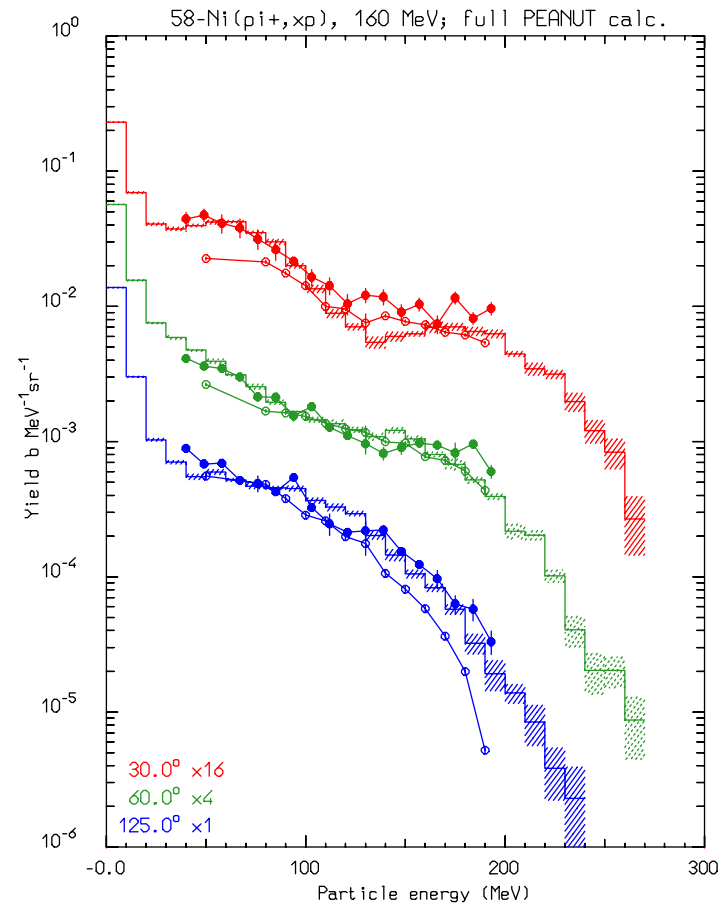


Emitted proton spectra at  
different angles , 160 MeV  $\pi^+$   
on Ni

Phys. Rev. C41,2215 (1990)

Phys. Rev. C24,211 (1981)

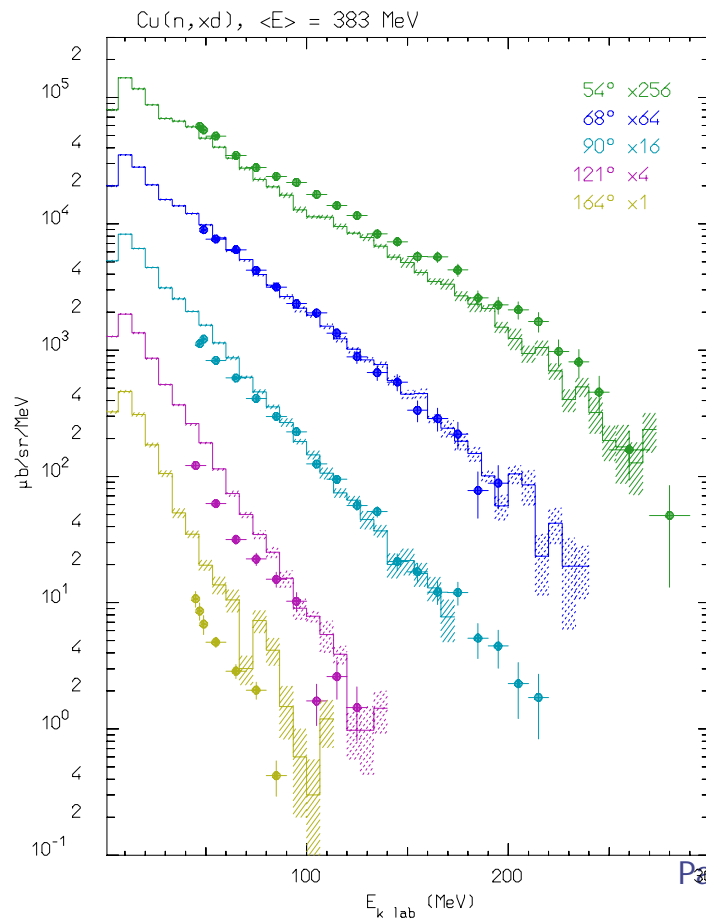
Proton spectra extend up to  
300 MeV



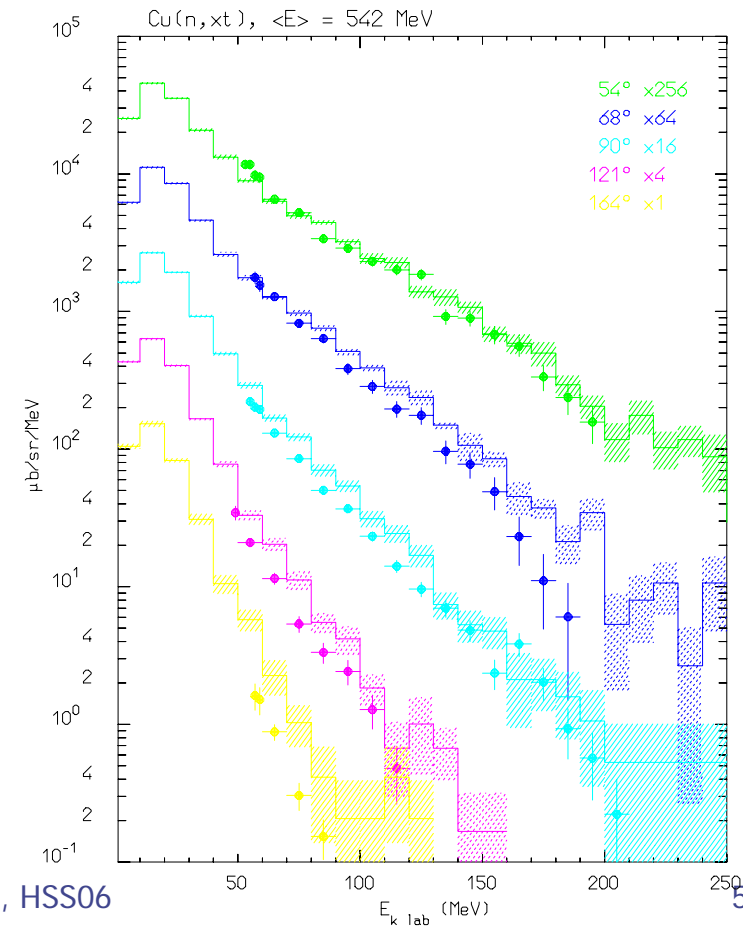
# Coalescence

High energy light fragments through the coalescence mechanism

double differential d production  
from 383 MeV neutrons on Copper  
Nucl. Phys. A510, 774 (1990)

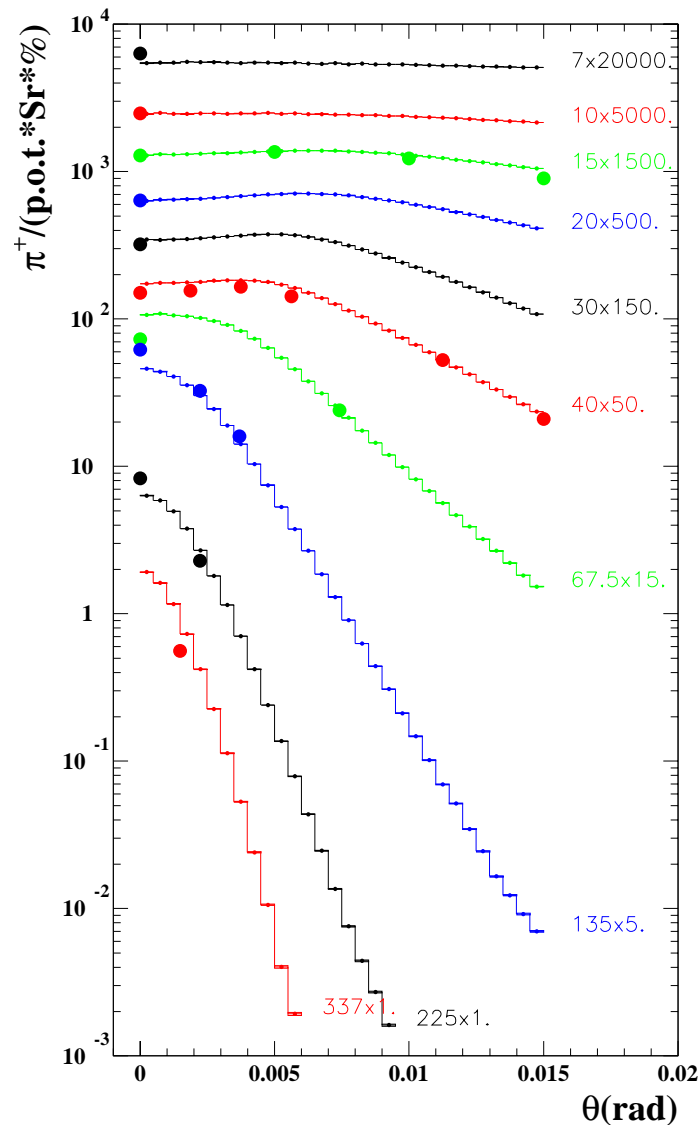


double differential t production  
from 542 MeV neutrons on Copper

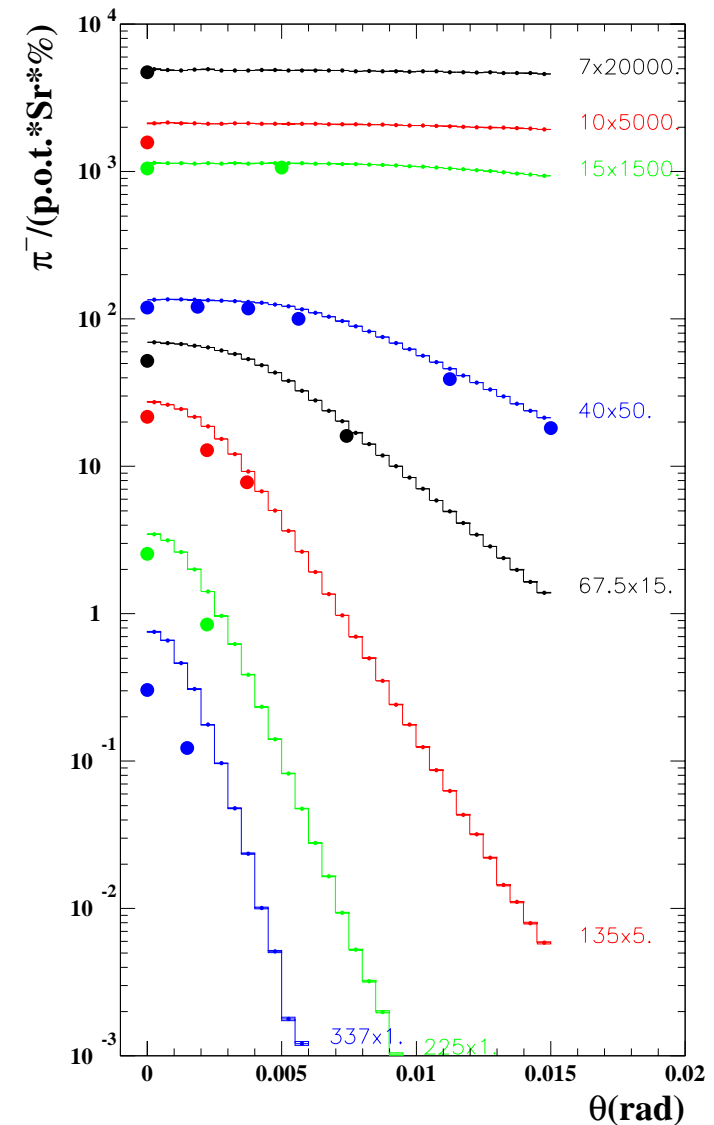


# Comparison with SPY data: 450 GeV p on Be

Double diff  
distribution for  
 $\pi^+/\pi^-$   
production from  
450 GeV/c p on  
Be  
H.W Atherton  
CERN 80-06 SPY :  
PLB 425, 208  
(1998)



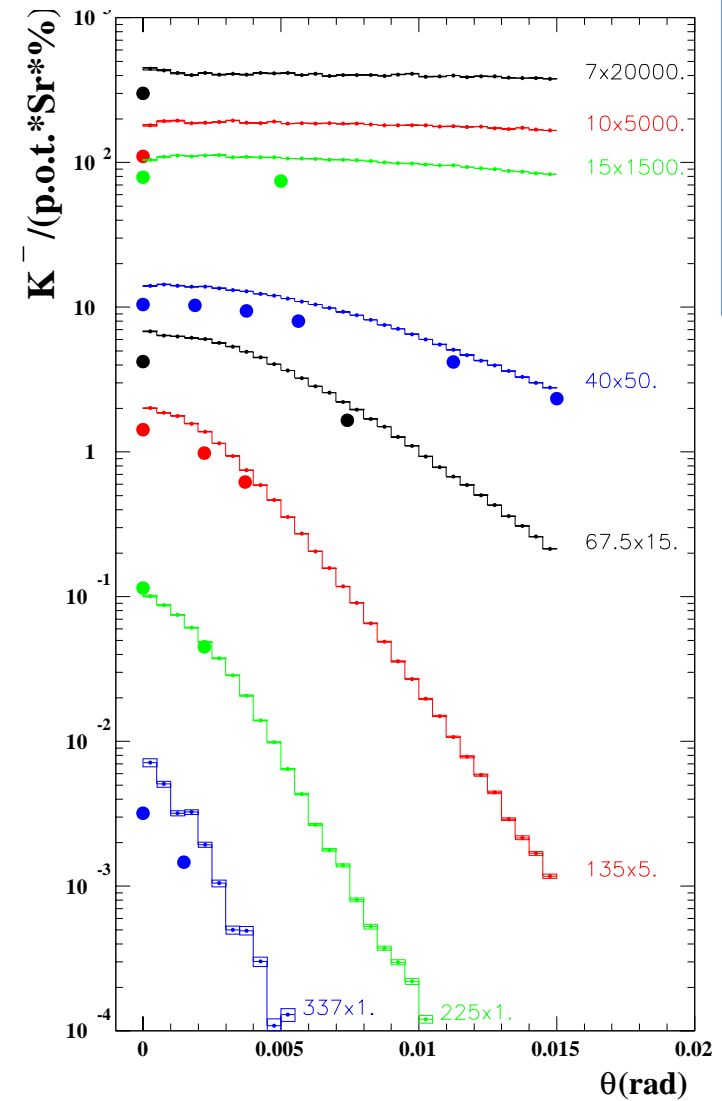
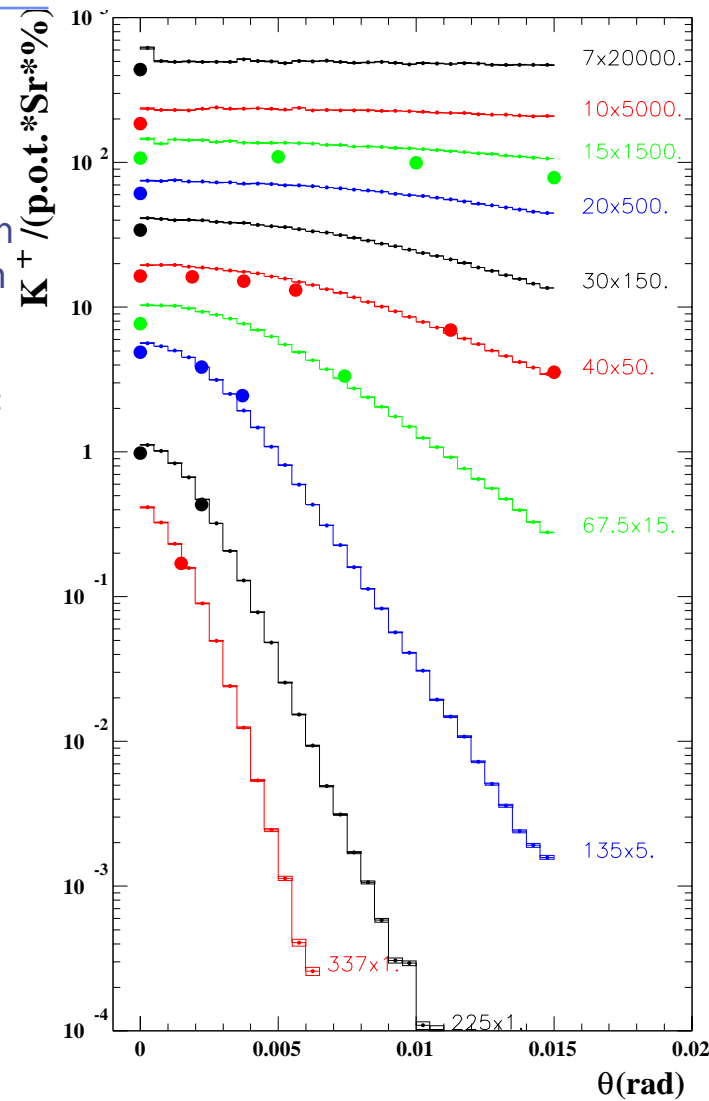
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# Comparison with SPY data: 450 GeV p on Be

Double diff  
distribution for  
 $K^+/K^-$   
production from  
450 GeV/c p on  
Be  
H.W Atherton  
CERN 80-06 SPY :  
PLB 425, 208  
(1998)



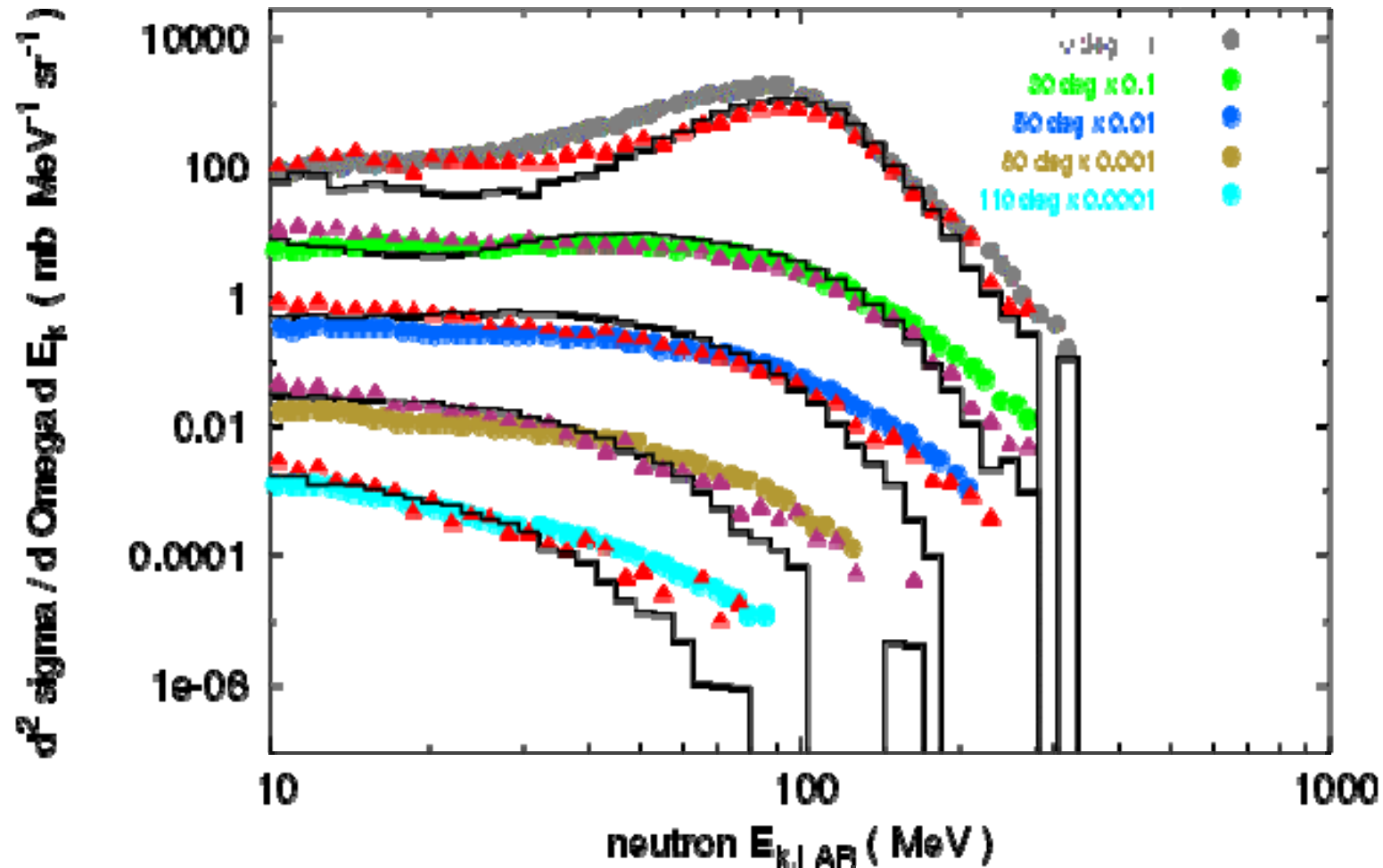


# ***Heavy Ions***



# QMD's: Neutron production cross section

Phys. Rev C64 034607 (2001) Ar + Cu 95 MeV/A all b



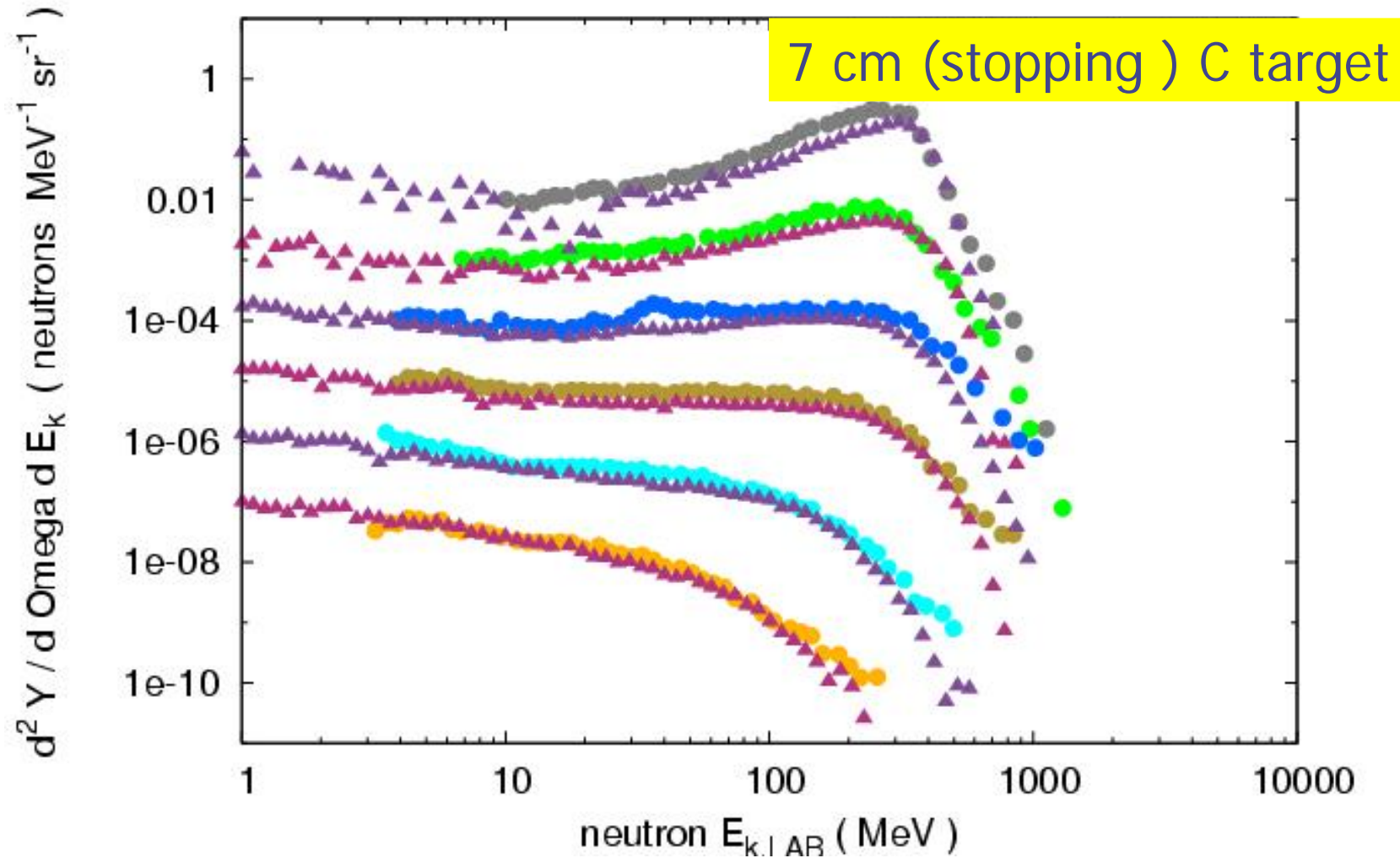
— RQMD + FLUKA

▲ QMD + FLUKA  
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● EXP data<sub>9</sub>

# The new QMD model: thick target example (full interface to FLUKA)

Phys. Rev. C62 044615 (2000) Ar + C 400 MeV/A all b



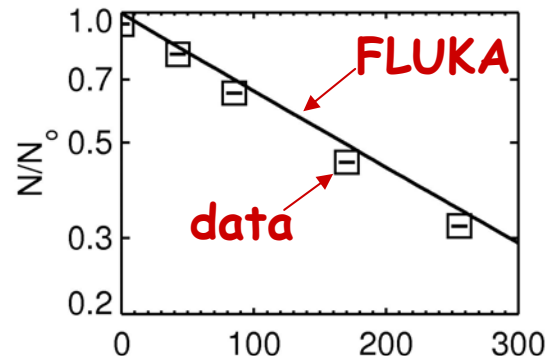
▲ QMD + FLUKA

● EXP data

# Fragmentation of therapeutic beams in a water target

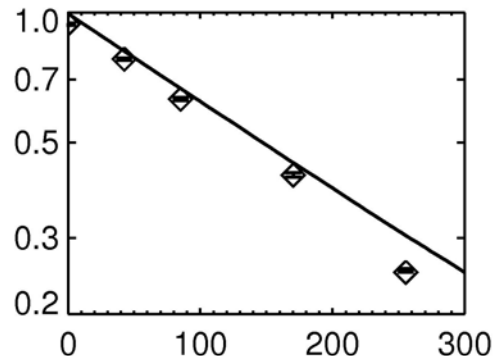
$^{12}\text{C}$

676 MeV/u



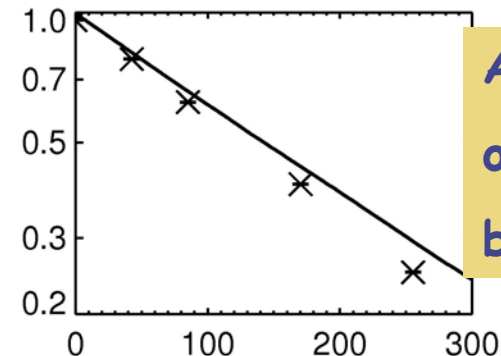
$^{14}\text{N}$

674 MeV/u



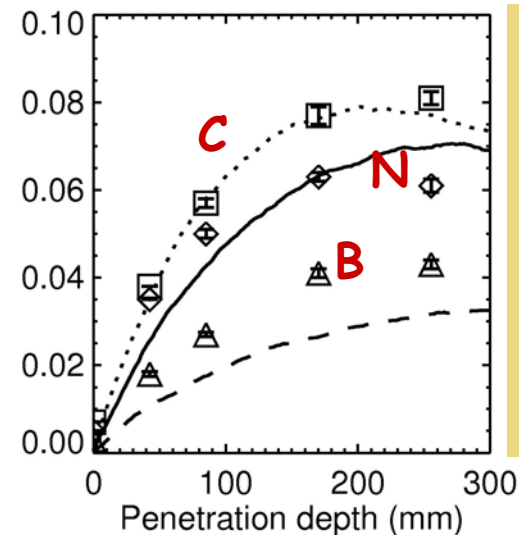
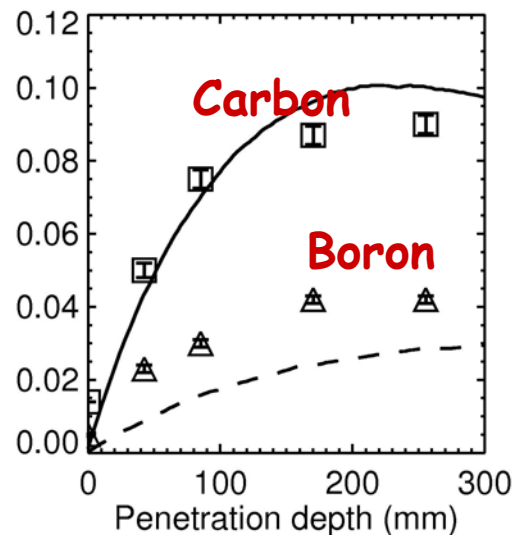
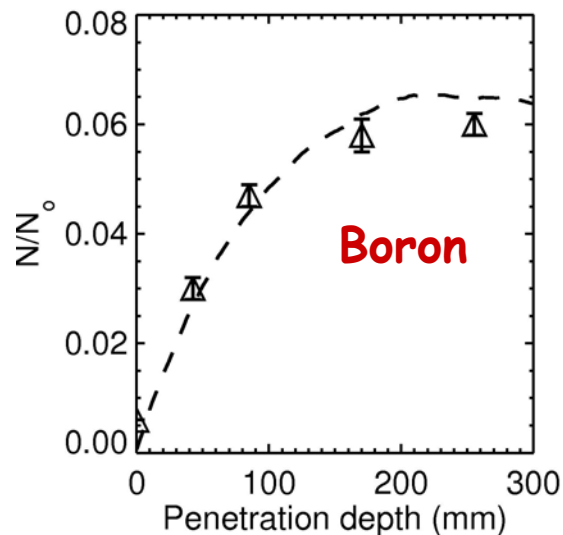
$^{16}\text{O}$

672 MeV/u



NIM B 117,  
221 (1996)

Attenuation  
of primary  
beam

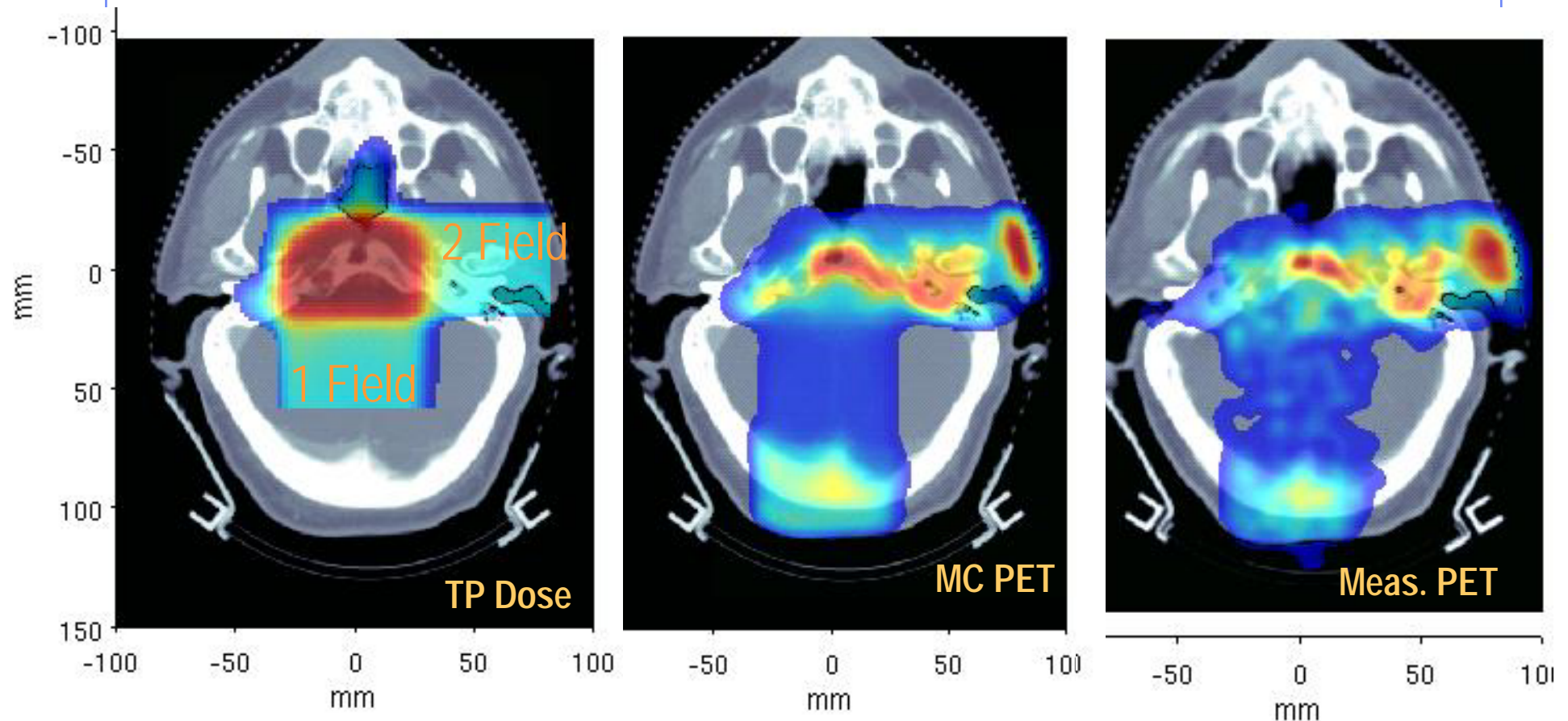


Build-up  
of  
secondary  
fragments  
As a  
function  
of depth

Simulations by F. Sommerer, K. Parodi, W. Enghardt, A. Aiginger  
Forschungszentrum Rossendorf (Dresden), Vienna Univ. of technology and CERN

# Preliminary results: Head K. Parodi et al., to be published

Clival Chordoma, 0.96 GyE / field,  $\Delta T_1 \sim 26$  min,  $\Delta T_2 \sim 16$  min



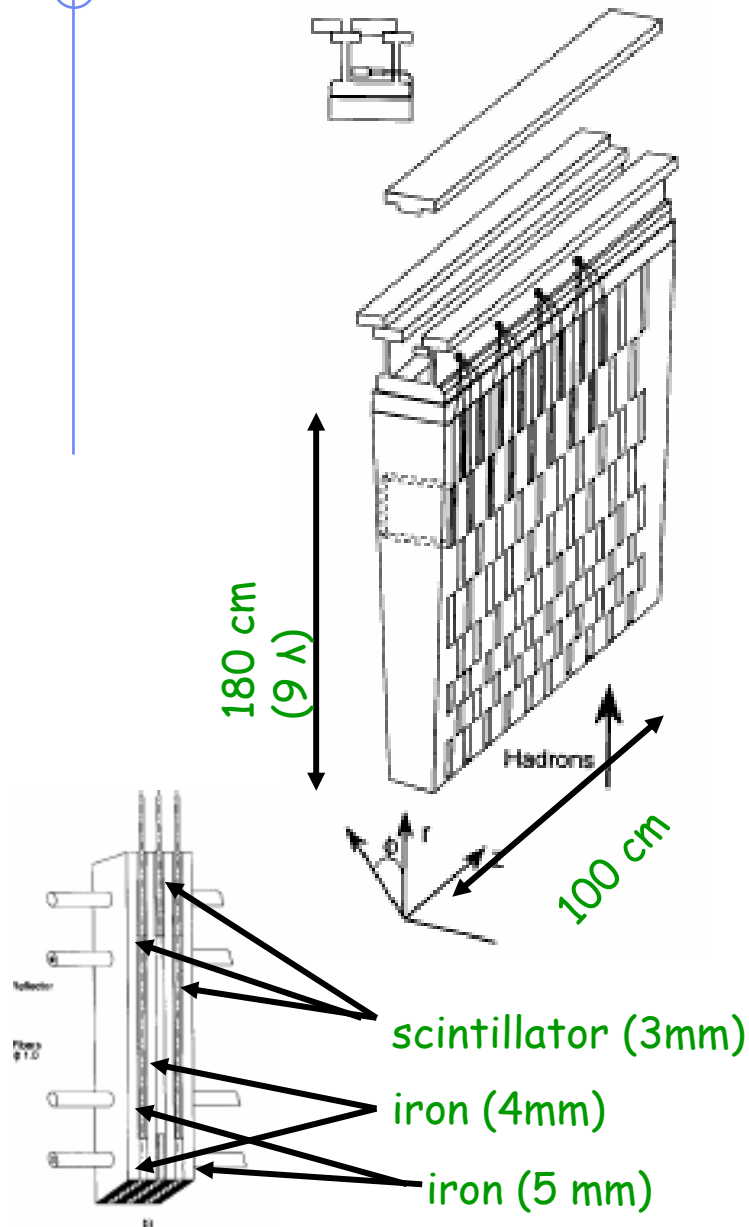
\* K. Parodi, H. Paganetti and T. Bortfeld, **Massachusetts General Hospital**



# ***Calorimeters***

# ATLAS TILE Calorimeter (1994 setup)

1994 Test beam : 5 modules, positrons and positive pions beam, 20-300 GeV/c  
NIM A394,384 (1994)



proton contamination in the beam measured (Cerenkov counters) in a later testbeam:

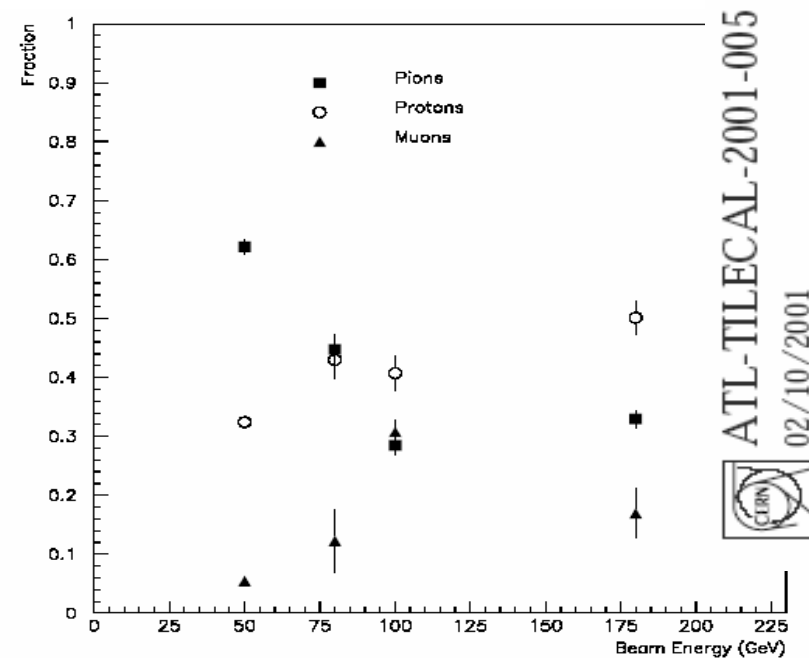
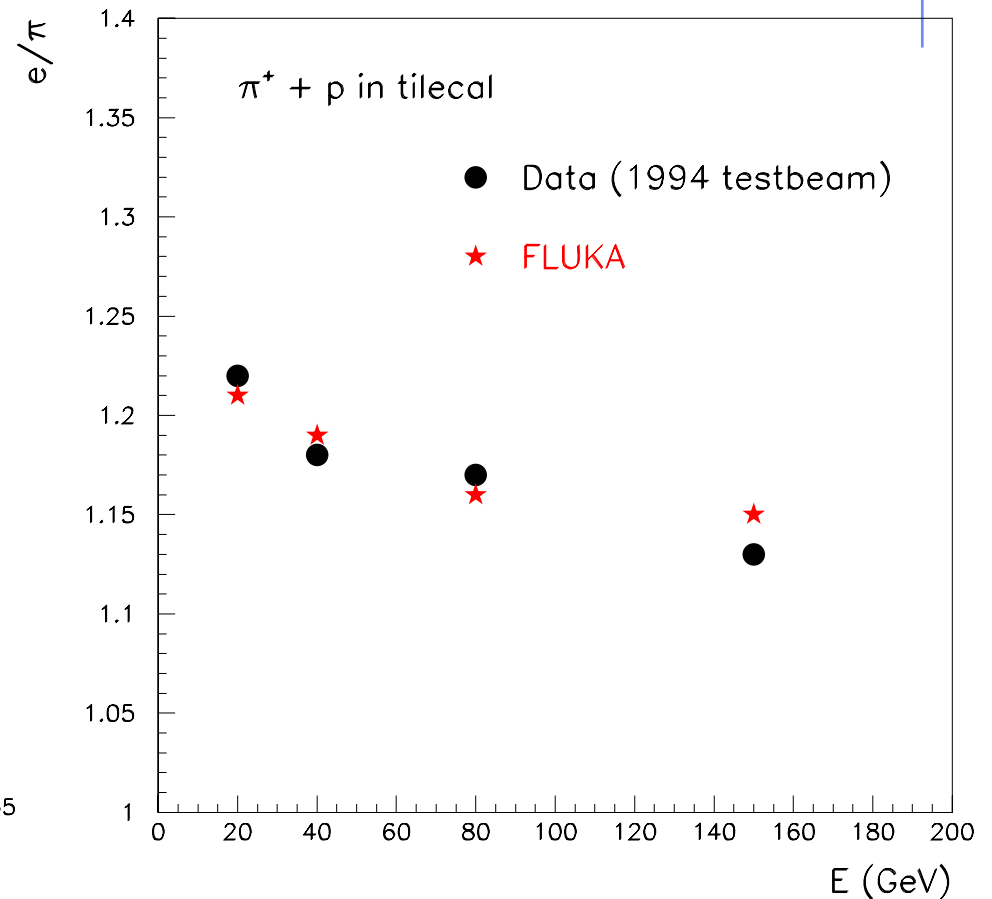
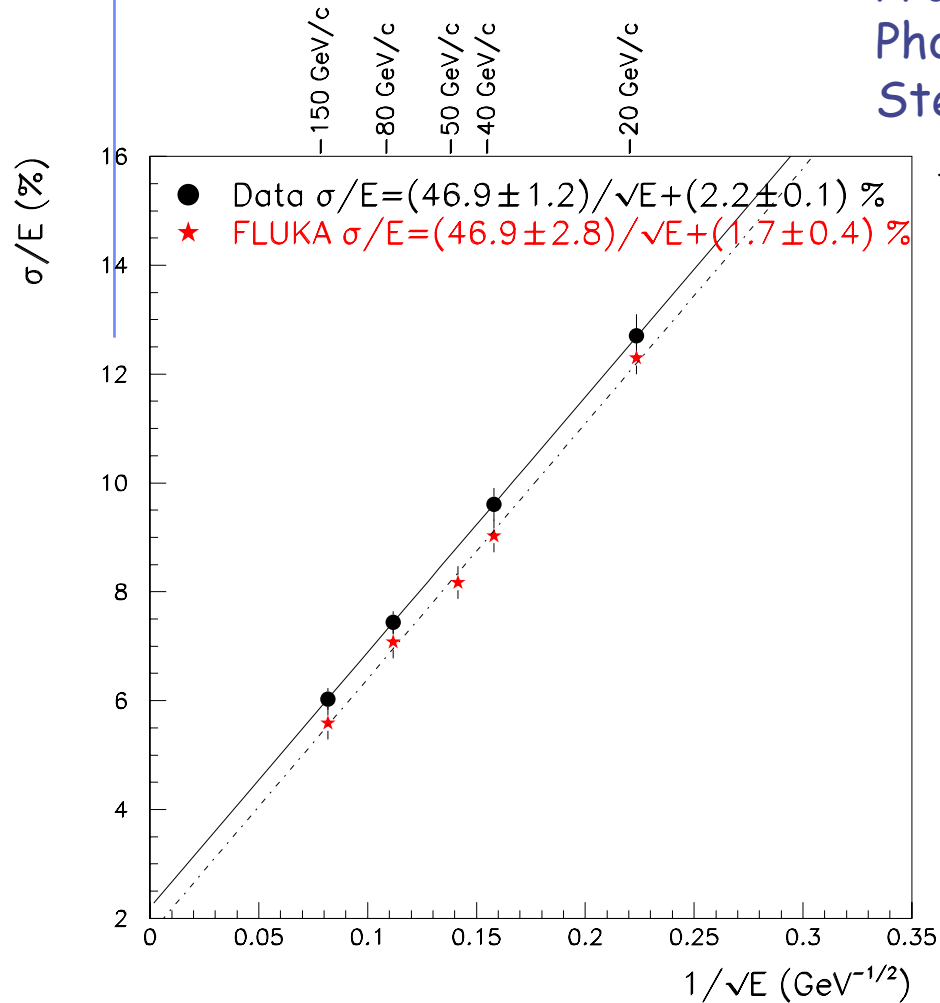


Figure 3: The fractions of pions, protons and muons in the positive pion beam energy.

# TILE Calorimeter: resolution and $e/\pi$

Beam at 20° incidence

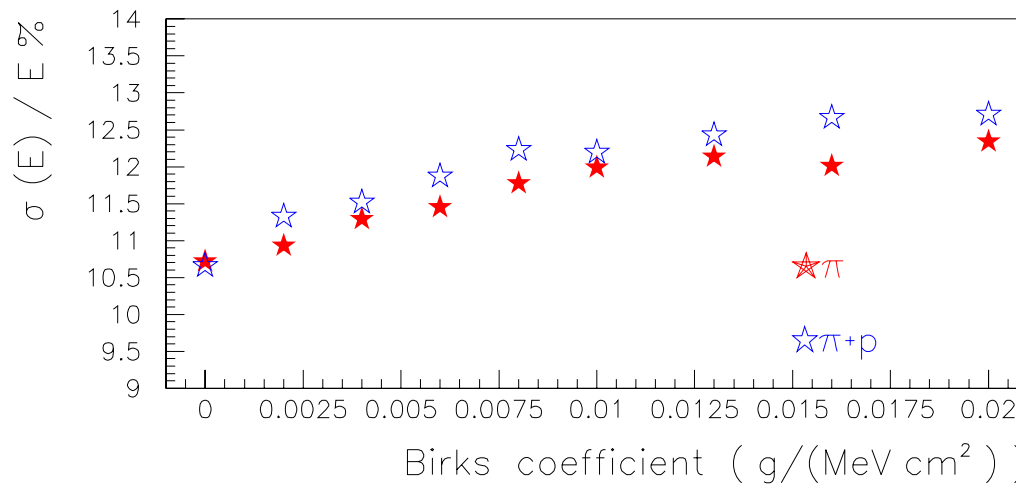
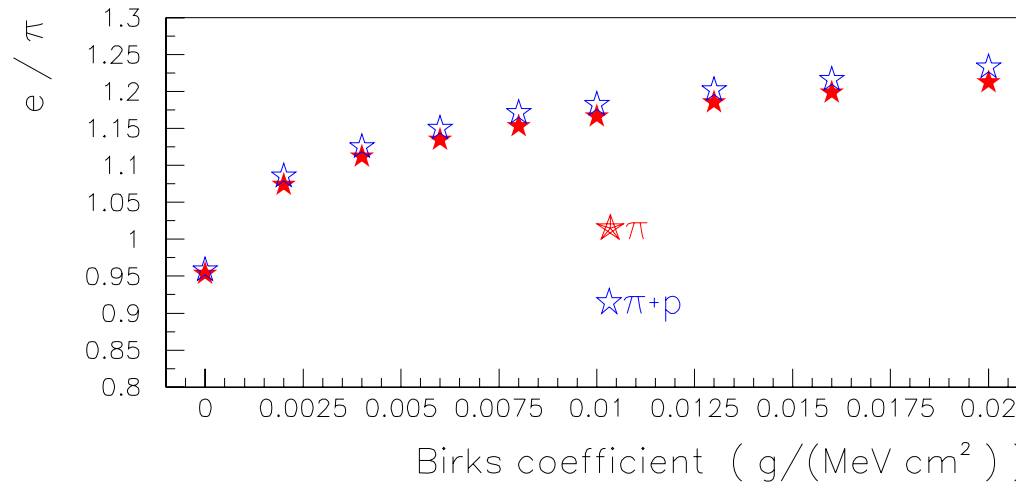
FLUKA simulations with:  
Proton contamination  
Photostatistics convoluted offline  
Step-by-step signal quenching online





# Tile Calorimeter: effect of quenching

Tilecal response to 20 GeV/c  $\pi^+$  vs. quenching



FLUKA simulations,  
20 GeV/c  
20° incidence  
different quenching  
parameters and  
effect of beam  
contamination  
on  $e/\pi$   
and resolution



# Tilecal: pion/proton difference

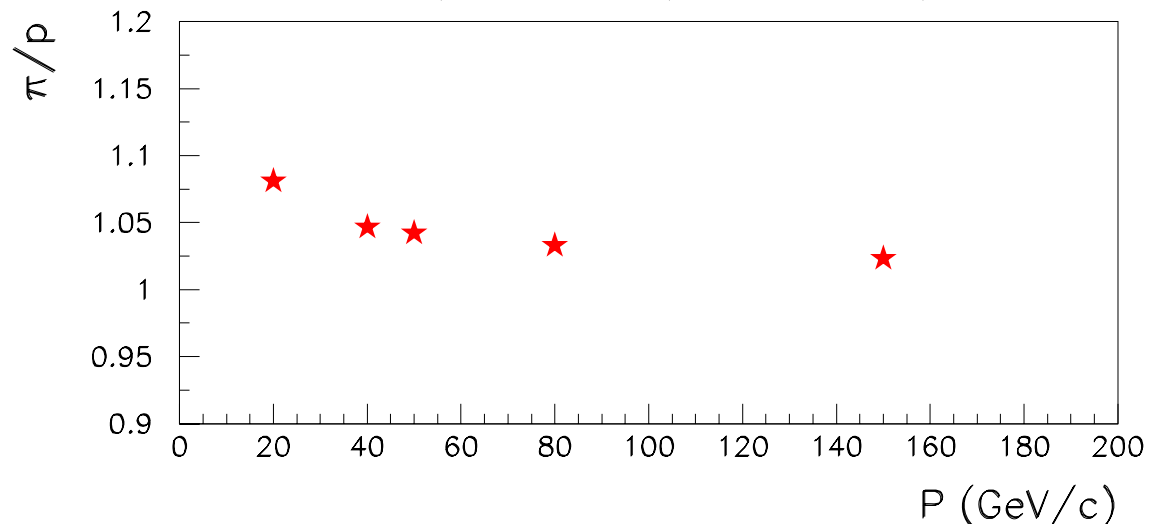
P (GeV/c)	$\pi$ signal/ proton signal		$\pi$ resolution/ proton resolution	
	Data	Fluka	Data	Fluka
50	1.048	1.042	1.11	1.07 $\pm 0.03$
80	1.036	1.034	1.17	1.09 $\pm 0.03$

Data :  
ATL-TILECAL-2001-005

WARNINGS:  
data/fluka have  
different tile  
configurations  
angle of incidence is  
also different

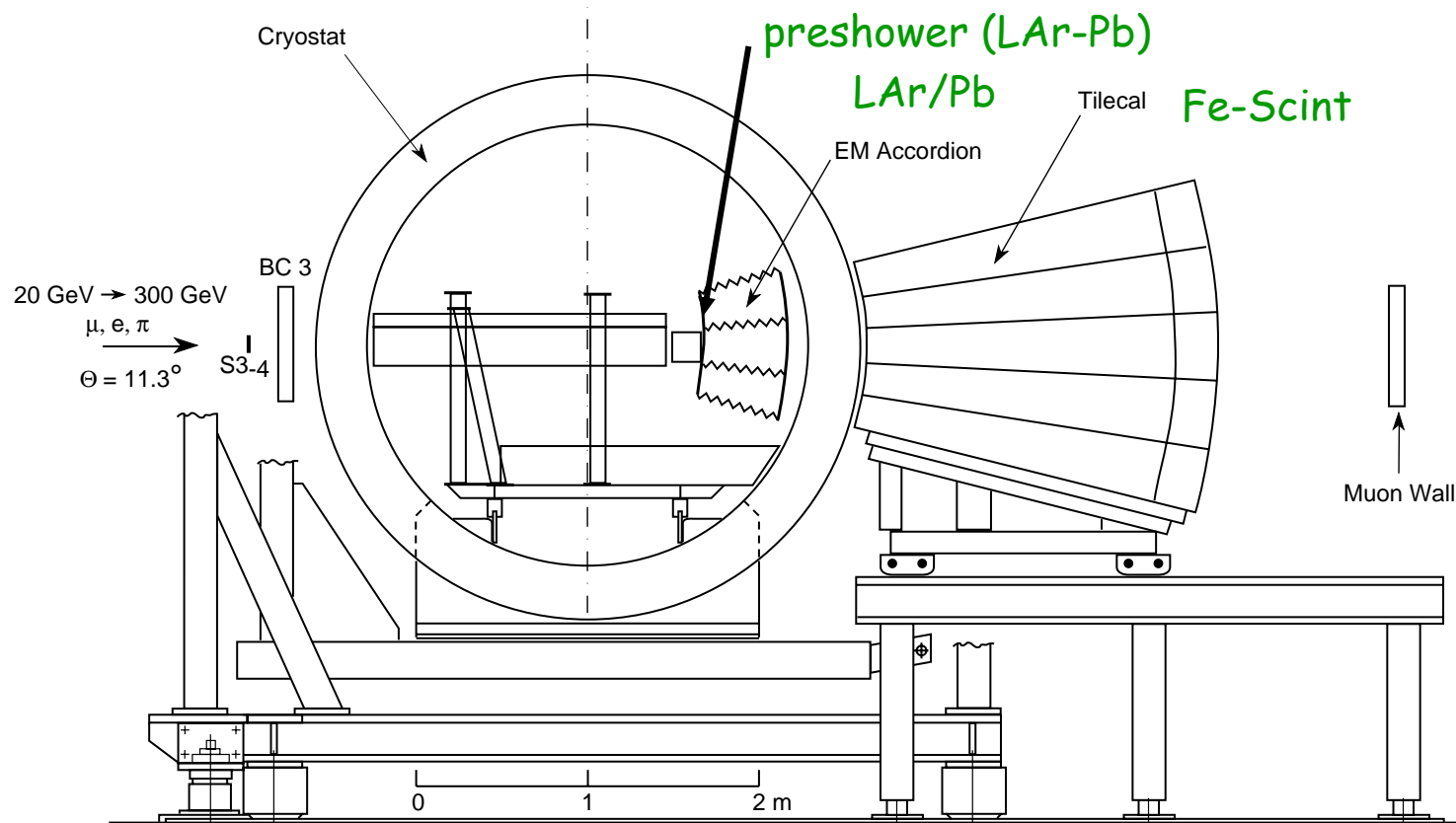
FLUKA simulations:  
energy dependence

Tilecal response to pions and protons



# ATLAS combined calorimeter test beam

Layout of the experimental set-up (NIM A387,333(1997), NIM A449,461 (2000))  
Test beams in 1994 and 1997, electrons and positive pions



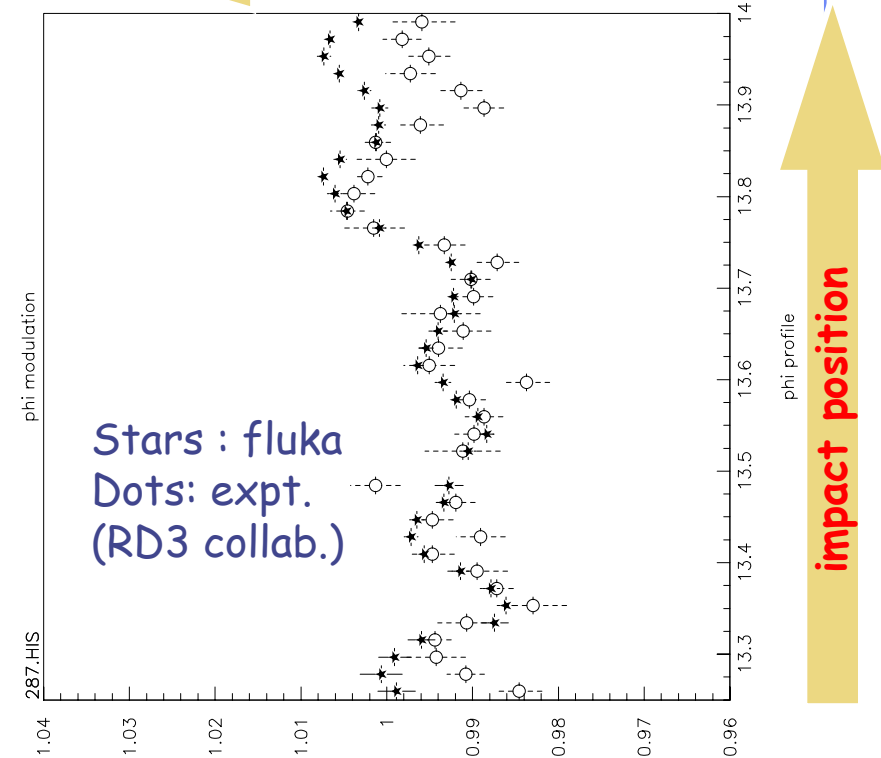
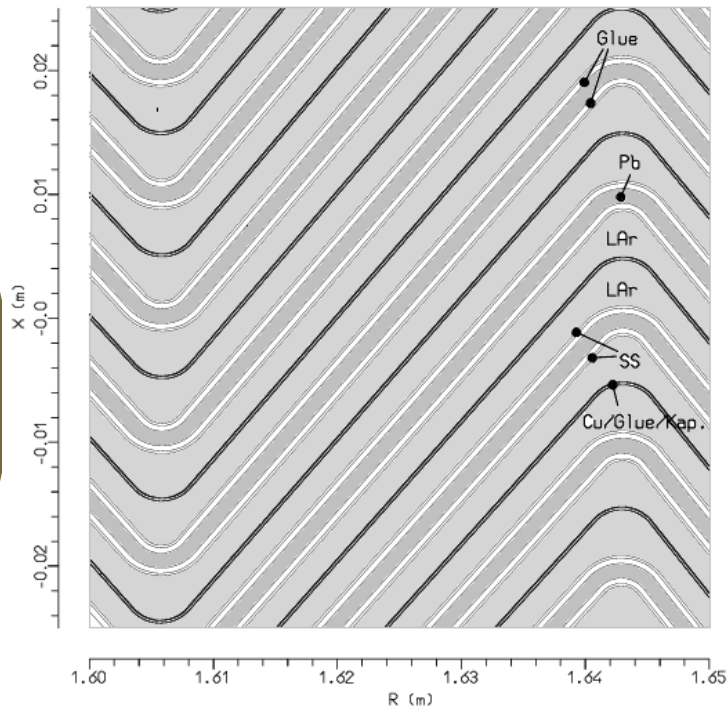
# The EM calo ( from standalone test beams)

Detail of the FLUKA geometry and

response vs. electron impact position

deposited energy

287 GeV  
electron  
beam



Energy resolution 10-100 GeV:

$$Exp : \frac{\sigma}{E} = \frac{9.8 \pm 0.4\%}{\sqrt{E}}$$

$$Fluka : \frac{\sigma}{E} = \frac{9.2 \pm 0.3\%}{\sqrt{E}}$$

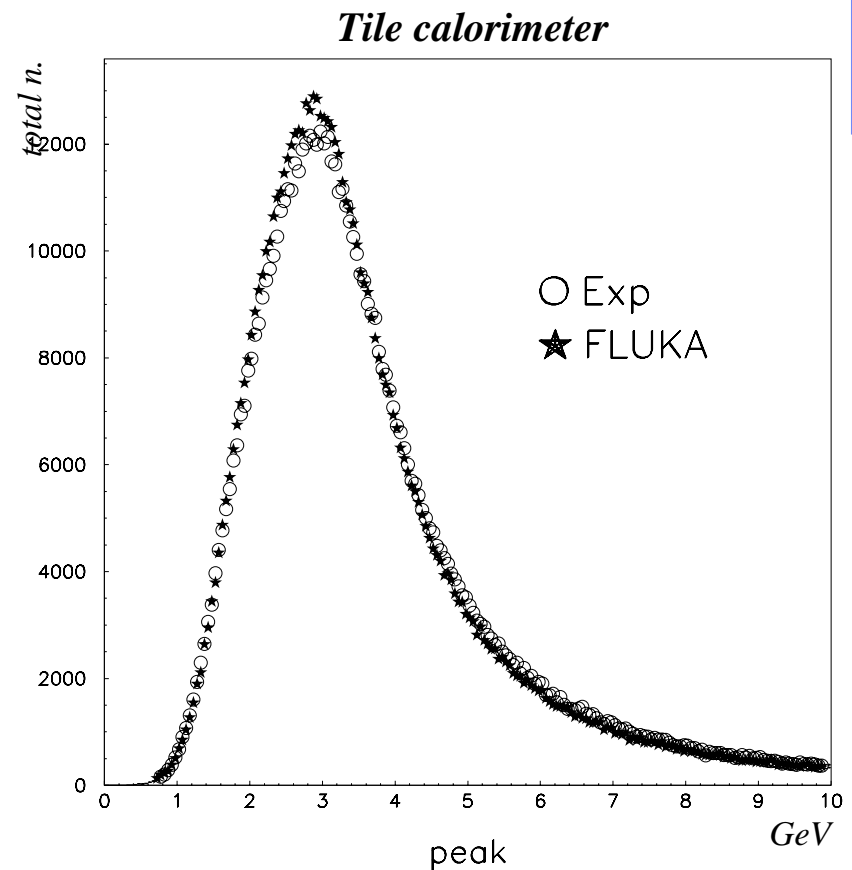
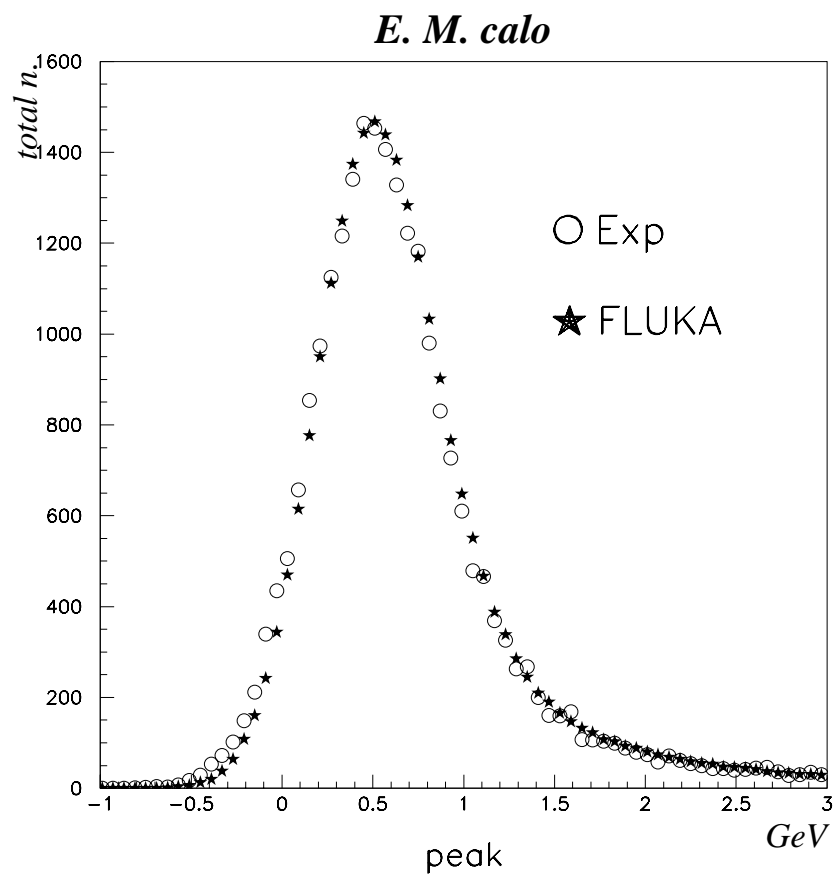
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Longitudinal Development

E GeV	E1/E		r.m.s.	
	Data	Fluka	Data	Fluka
100	0.68	0.69		
287	0.61	0.58	.091	.094

# 300 GeV $\mu$ in ATLAS combined calo

Calibration in electron scale, electronic noise added



Proc. of Calor96

# Comparison with PION data SIMULATIONS

## DATA

Cut on mip in presampler  
Cut on beam position  
(beam chambers)

Calibration in electron scale  
Scintillator quenching included  
Photostatistics convoluted  
Noise added  
Proton contamination taken  
into account  
Cut on preshower

Energy reconstructed using the ``benchmark'' technique

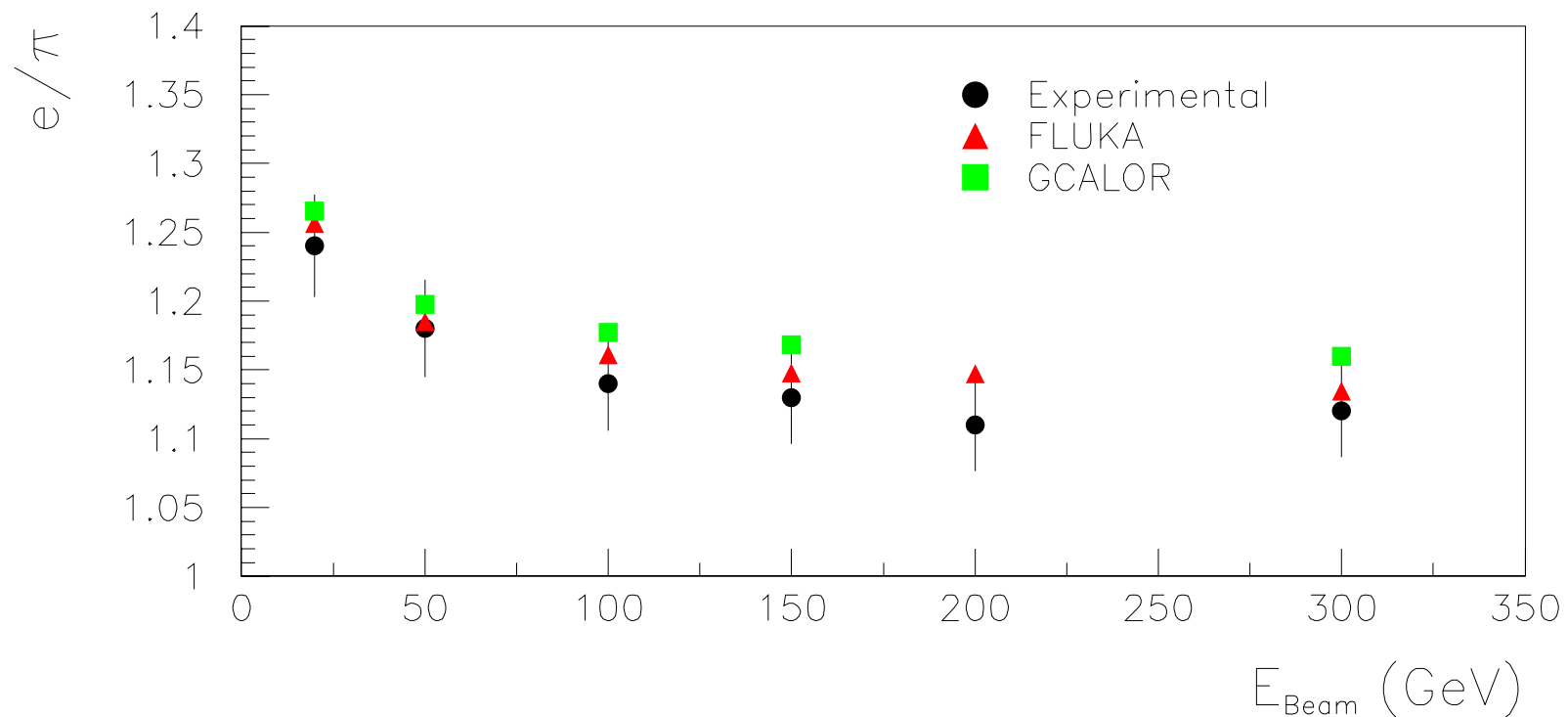
$$E_0 = E_{EM} + a \cdot Q_{had} + b \cdot \sqrt{E_{EM3} \cdot a \cdot Q_{had1}} + c \cdot E_{EM}^2$$

All parameters fixed to minimize  $\sigma/E_0$  at 300~GeV

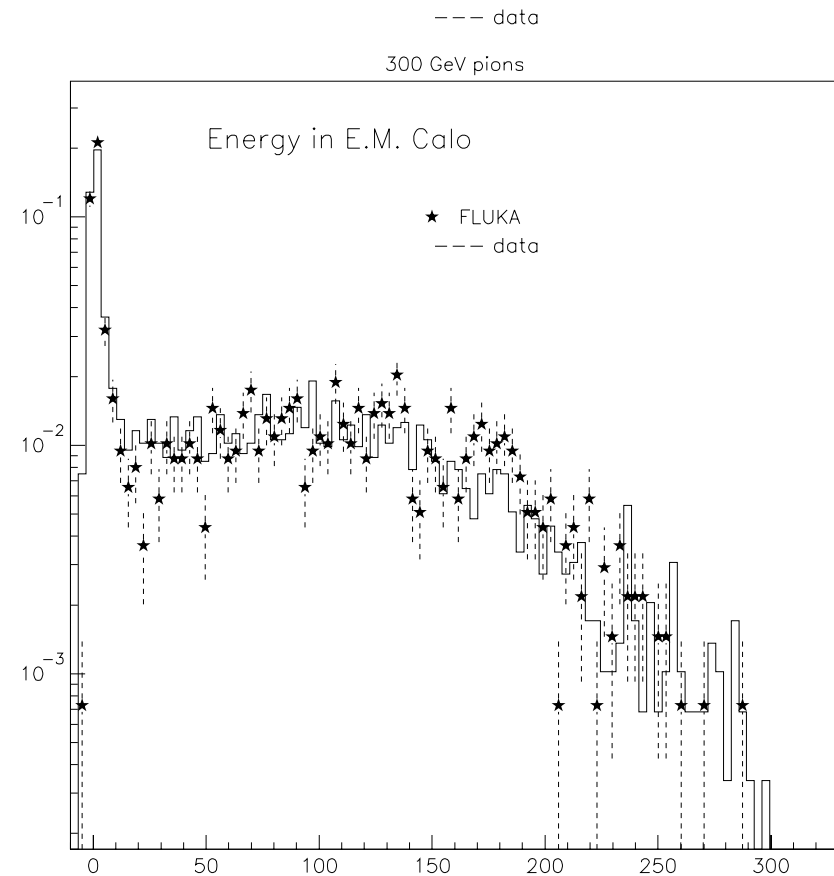
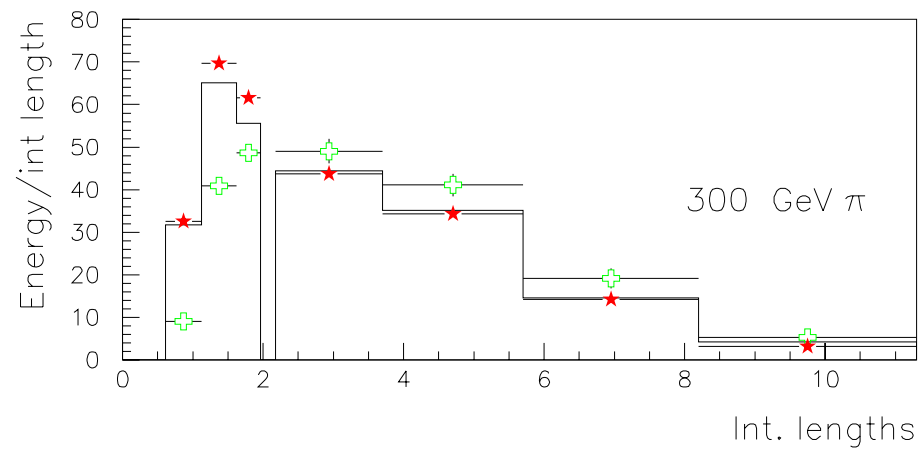
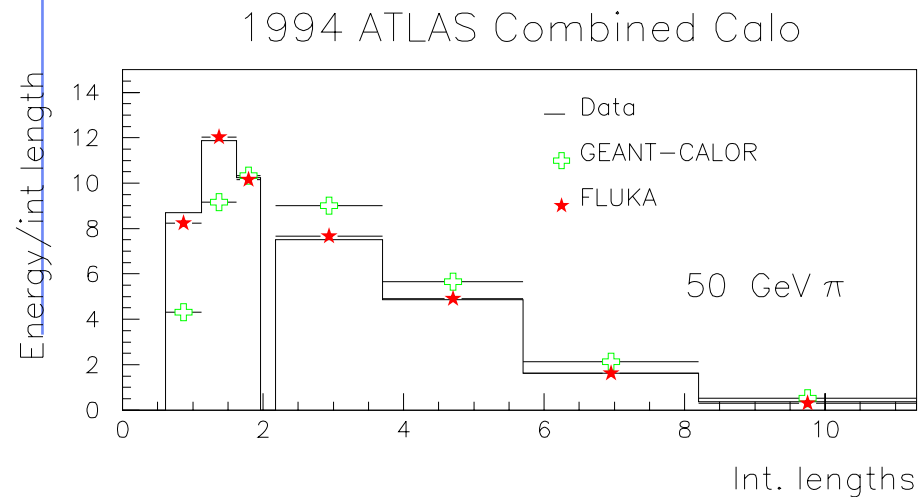
# Atlas combined calo test beam: 1994 data

Experimental electron  
scale calibration  
available

	$a$ (GeV/pC)	$b$	$c$ (GeV <sup>-1</sup> )
FLUKA	0.172(2)	0.38(2)	-0.00038(10)
EXP	0.172	0.44	-0.00038

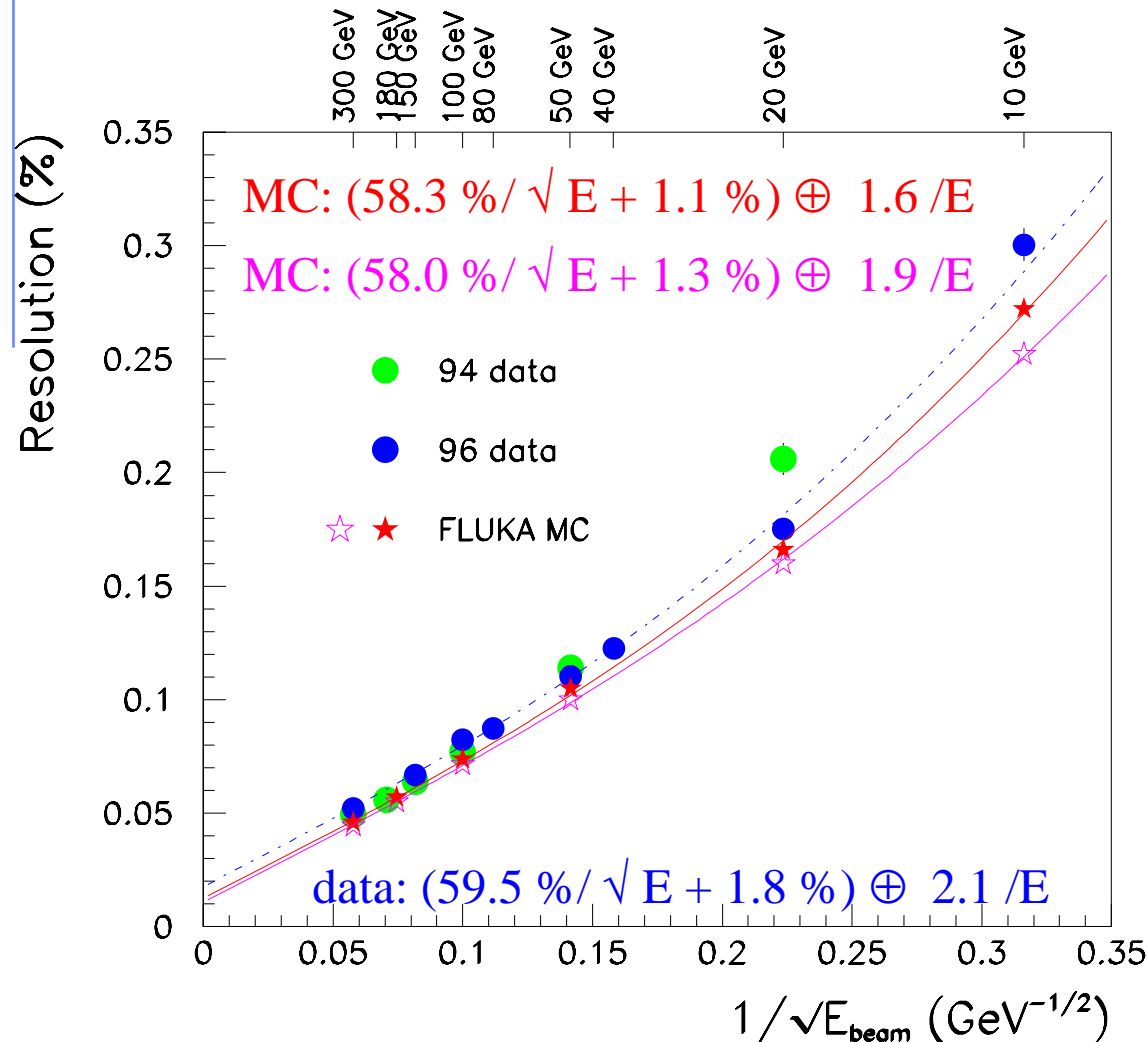


# Atlas combined calo test beam: 1994 data



# Atlas combined calo test beam: resolution

## Energy resolution, 1994+1996 data



1996 Data  
Less electronic noise  
Better presampler  
No electron calibration

Note the  
experimental point  
at 20 GeV

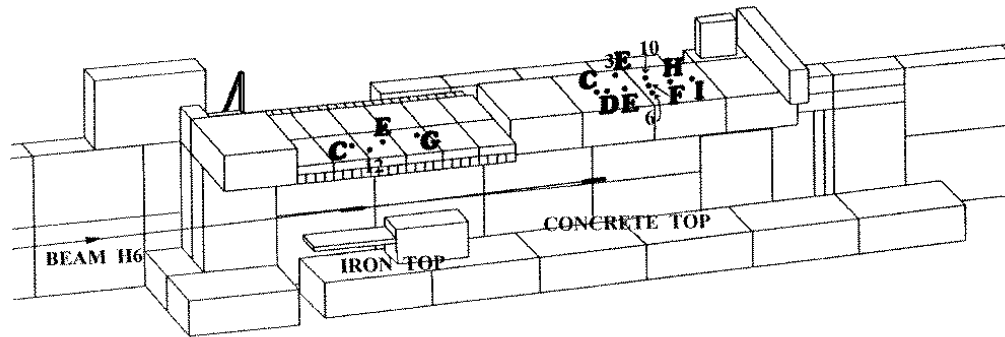
2 simulated curves:  
different algorithms  
for preshower  
reconstruction



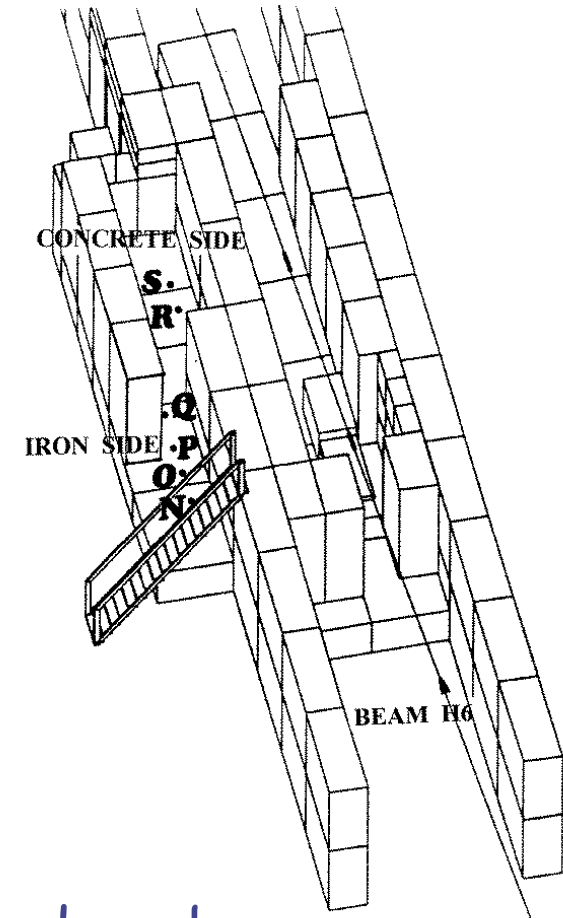


# **CERN-EU High-Energy Reference Field (CERF) facility**

# CERF: neutron measurements



Top (left, one side removed) and side (right, roof removed) views of the CERF facility with the measuring positions



120 GeV secondary SPS mixed hadron beam

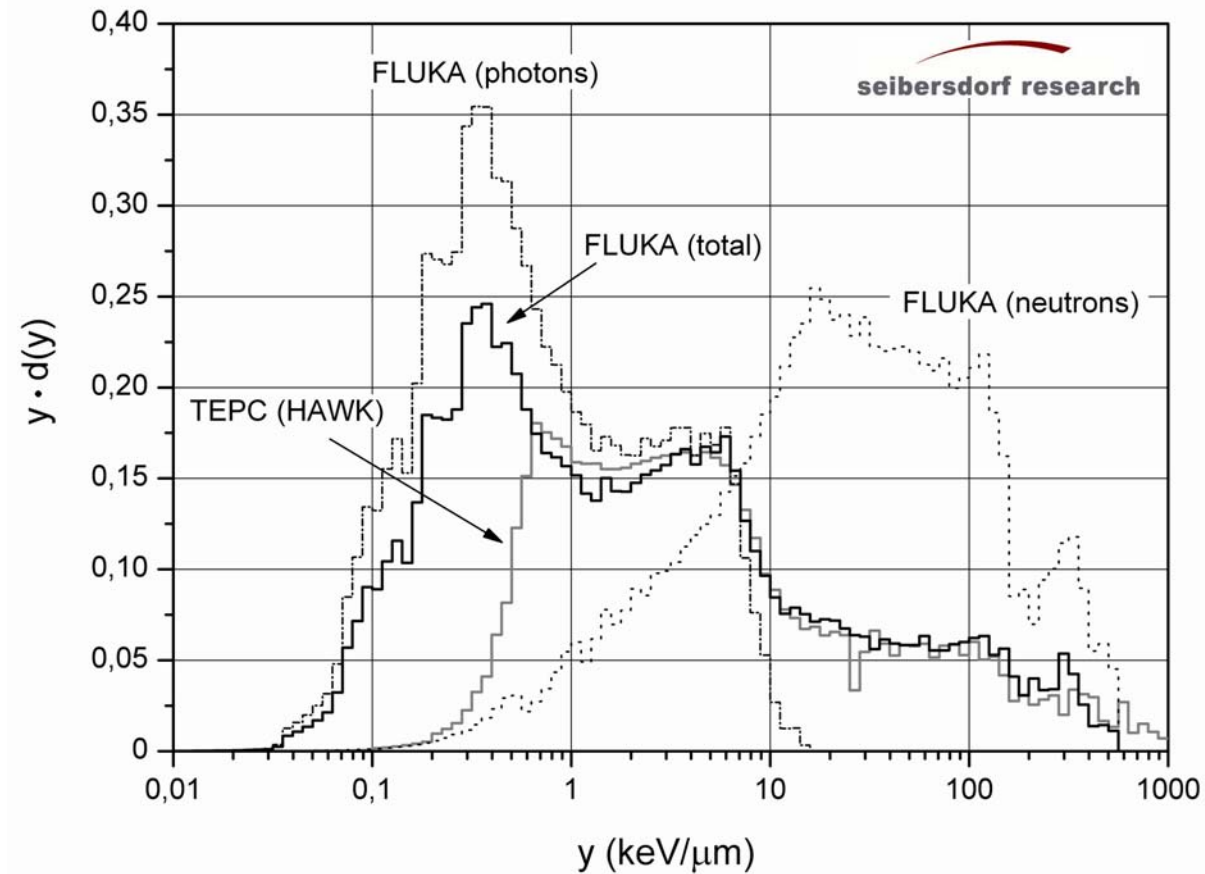
# CERF: results

	experimental		FLUKA		experimental		FLUKA	
	cts/PIC	%	cts/PIC	%	cts/PIC	%	cts/PIC	%
	CONCRETE TOP "E"				IRON TOP "C"			
LINUS rem counter*	0.364	0.36	0.409	2.2	1.78	0.30	1.68	2.1
SNOOPY rem counter*	0.200	0.59	0.207	3.3	1.83	0.75	1.71	2.0
233 sphere	0.788	0.33	0.899	3.7	9.28	0.28	9.23	2.0
178 sphere	0.989	0.36	1.01	3.4	16.1	0.24	16.9	1.9
133 sphere	1.02	0.30	0.981	3.2	19.2	0.19	21.2	1.9
108 sphere	0.942	0.35	0.883	3.1	17.7	0.20	19.2	1.9
83 sphere	0.704	0.30	0.717	3.1	11.2	0.26	12.1	1.9

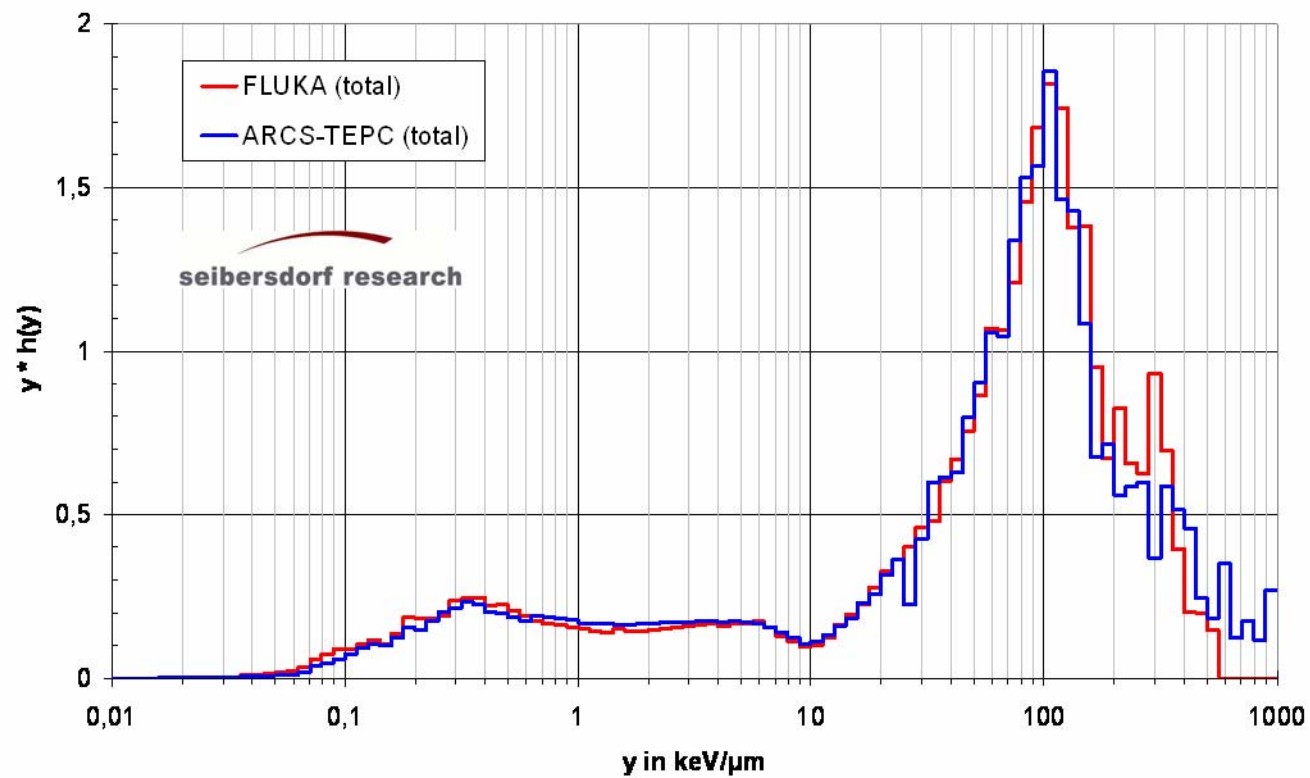
Comparison between the FLUKA predictions and the experimental response of the various detectors in stray radiation fields at CERN\*. The percent statistical (%) uncertainty is indicated

\* C.Birattari et al, Rad.Prot.Dos., 76 (1998), 135

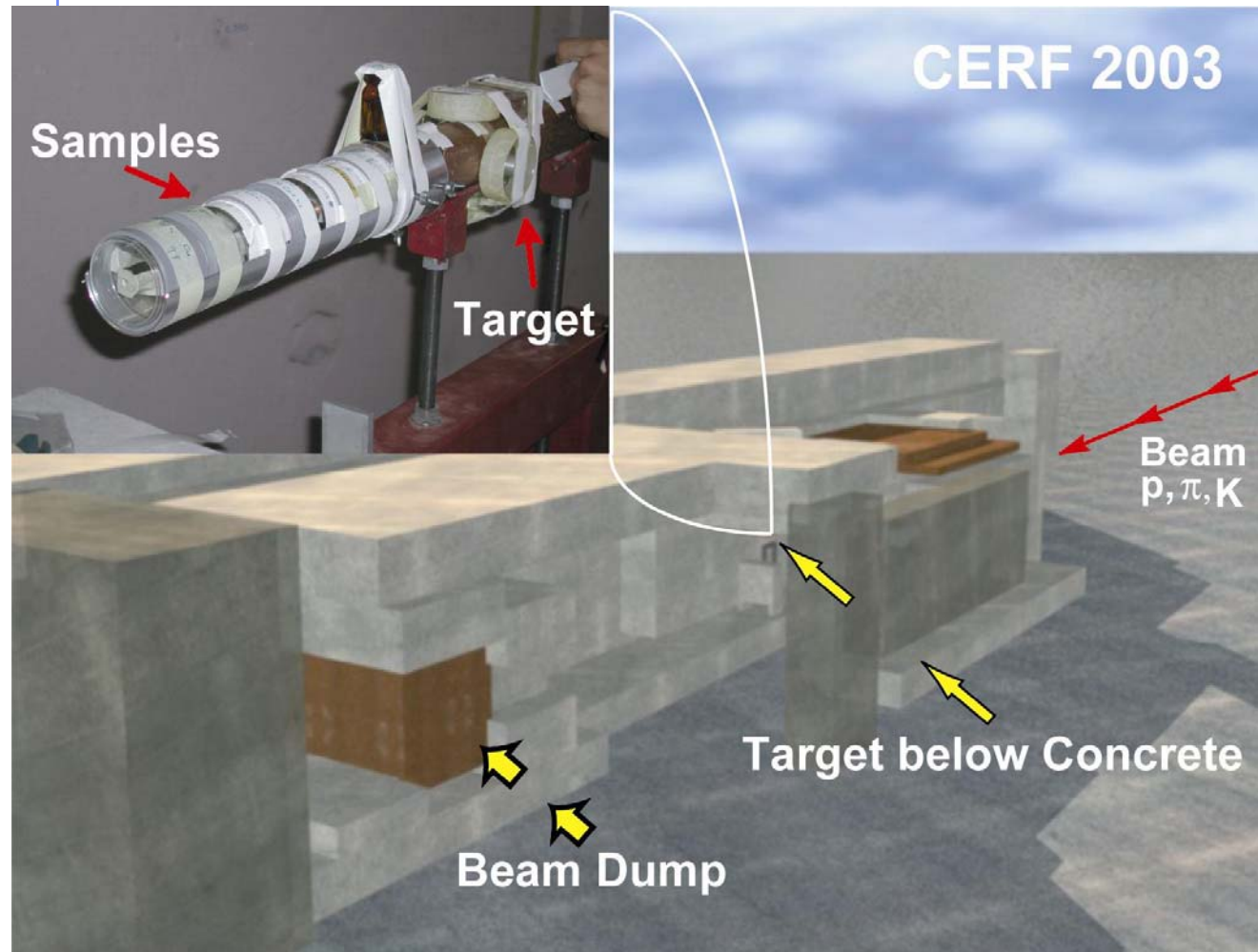
## TEPC Absorbed Dose Measurements and FLUKA Simulation CERF facility for mixed Radiation Field



# TEPC Dose Equivalent Measurements and FLUKA Simulation CERF facility for mixed Radiation Field



# CERN-EU High-Energy Reference Field (CERF) facility



## Location of Samples:

Behind a 50 cm long, 7 cm diameter copper target, centred with the beam axis

# Calculation of Induced Activity with FLUKA

- Simulation of particle interactions and transport in the target, the samples, as well as the tunnel/cavern walls
- Separate simulations for proton and pion beam
- Simulations of isotope production via
  - High-energy processes
  - Low-energy neutron interactions
- Transport thresholds
  - Neutrons: down to thermal energies
  - Other hadrons: until stopped or captured
  - No electromagnetic cascade was simulated
- Calculated quantities
  - Radioactive isotope production per primary particle
  - (Star density and particle energy spectra in the samples)
- Calculation of build-up and decay of radioactive isotopes for specific irradiation and cooling patterns including radioactive daughter products

# Activation: Stainless Steel

Table 1: Stainless Steel, cooling times 1d 6h 28m, 17d 10h 39m

Isotope	t <sub>1/2</sub>	Exp Bq/g ± %		OLD FLUKA/Exp ± %		FLUKA/Exp ± %	
Be 7	53.29d	0.205	24	0.096	34	1.070	30
Na 24	14.96h	0.513	4.3	0.278	8.6	0.406	13
K 43	22.30h	1.08	4.6	0.628	8.7	0.814	11
Ca 47	4.54d	0.098	25	0.424	44	(0.295	62)
Sc 44	3.93h	13.8	4.8	0.692	5.8	0.622	6.2
mSc 44	58.60h	6.51	7.1	1.372	8.1	1.233	8.6
Sc 46	83.79d	0.873	8.3	0.841	9.1	0.859	9.5
Sc 47	80.28h	6.57	8.2	0.970	9.7	1.050	13
Sc 48	43.67h	1.57	5.2	1.266	8.4	1.403	11
V 48	15.97d	8.97	3.1	1.464	3.8	1.354	4.8
Cr 48	21.56h	0.584	6.7	1.084	11	1.032	12
Cr 51	27.70d	15.1	12	1.261	13	1.231	13
Mn 54	312.12d	2.85	10	1.061	10	1.060	11
Co 55	17.53h	1.04	4.6	1.112	7.7	0.980	10
Co 56	77.27d	0.485	7.6	1.422	9.0	1.332	10
Co 57	271.79d	0.463	11	1.180	12	1.140	12
Co 58	70.82d	2.21	5.9	0.930	6.3	0.881	6.9
Ni 57	35.60h	3.52	4.5	1.477	6.5	1.412	8.2

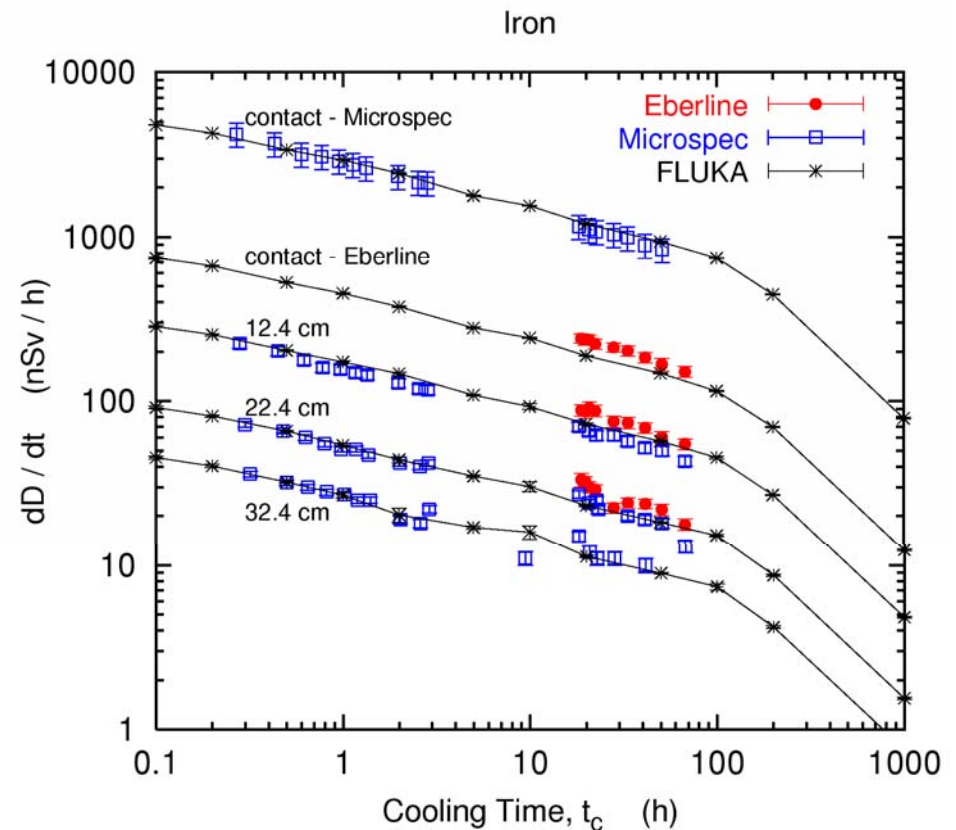
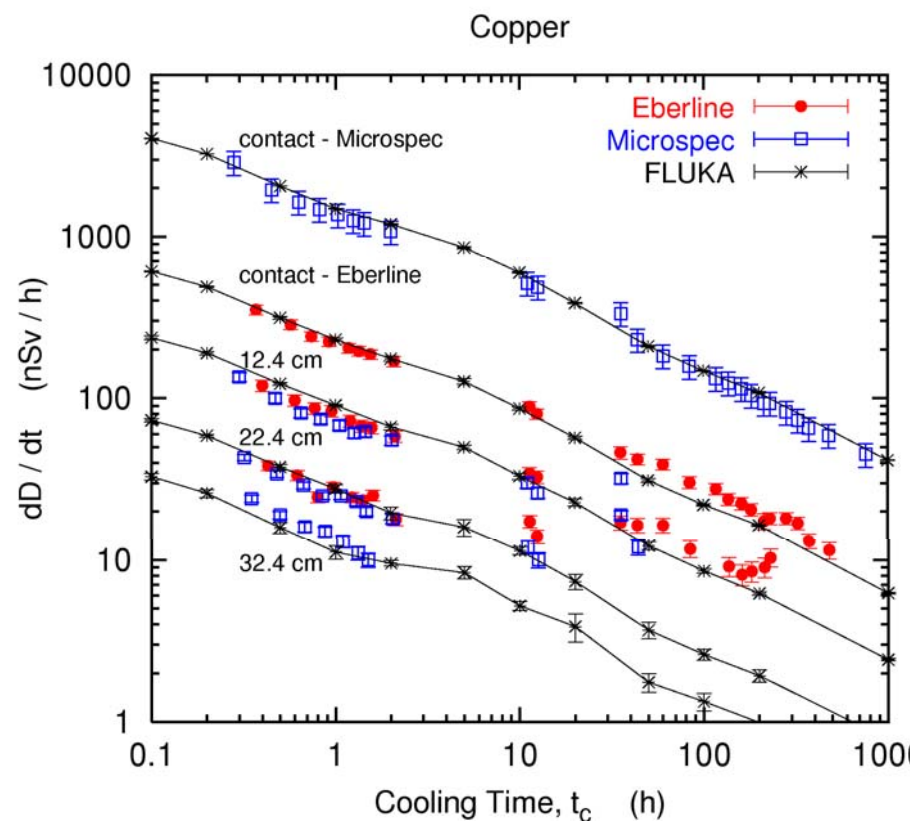
M. Brugger,  
*et al.*,  
Proceedings  
of the Int.  
Conf. on  
Accelerator  
Applications  
(AccApp'05),  
Venice, Italy,  
2005



# Benchmark experiment - *Results 1*

M. Brugger *et al.*, Radiat. Prot. Dosim. 116 (2005) 12-15

Dose rate as function of cooling time  
for different distances between sample and detector





# ***Cosmic Rays***

## FLUKA and Cosmic Ray physics: Atmospheric Showers

### Two different streams:

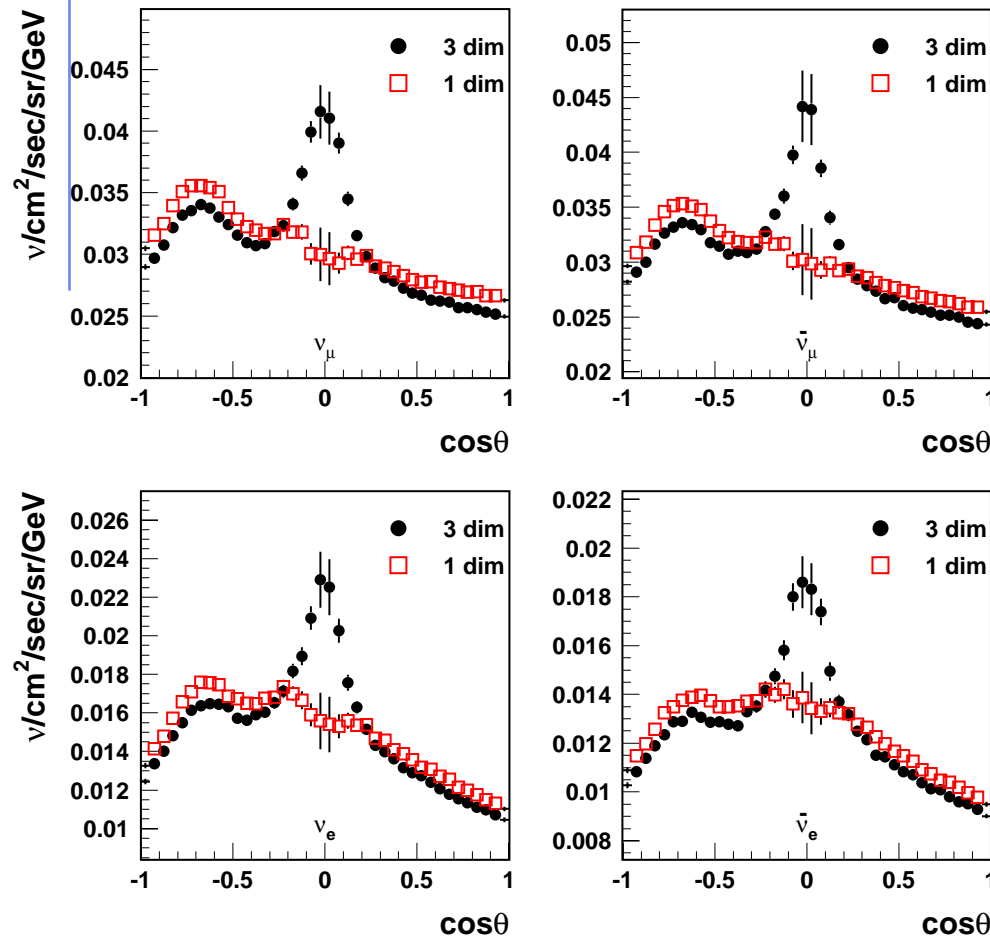
- Basic research on Cosmic Ray physics (muons, neutrinos, EAS, underground physics,...)
- Application to dosimetry in civil aviation (DOSMAX  
Collaboration: Dosimetry of Aircrew Exposure to Radiation  
During Solar Maximum, research project funded by the EU)

### Available dedicated FLUKA library + additional packages including:

- Primary spectra from  $Z = 1$  to  $Z = 28$  (derived from NASA and updated to most recent measurements.)
- Solar Modulation model (correlated to neutron monitors)
- Atmospheric model (MSIS Mass-Spectrometer-Incoherent-Scatter)
- 3D geometry of Earth + atmosphere
- Geomagnetic model

## (3D) Calculation of Atmospheric $\nu$ Flux

Sub-GeV flux at Kamioka



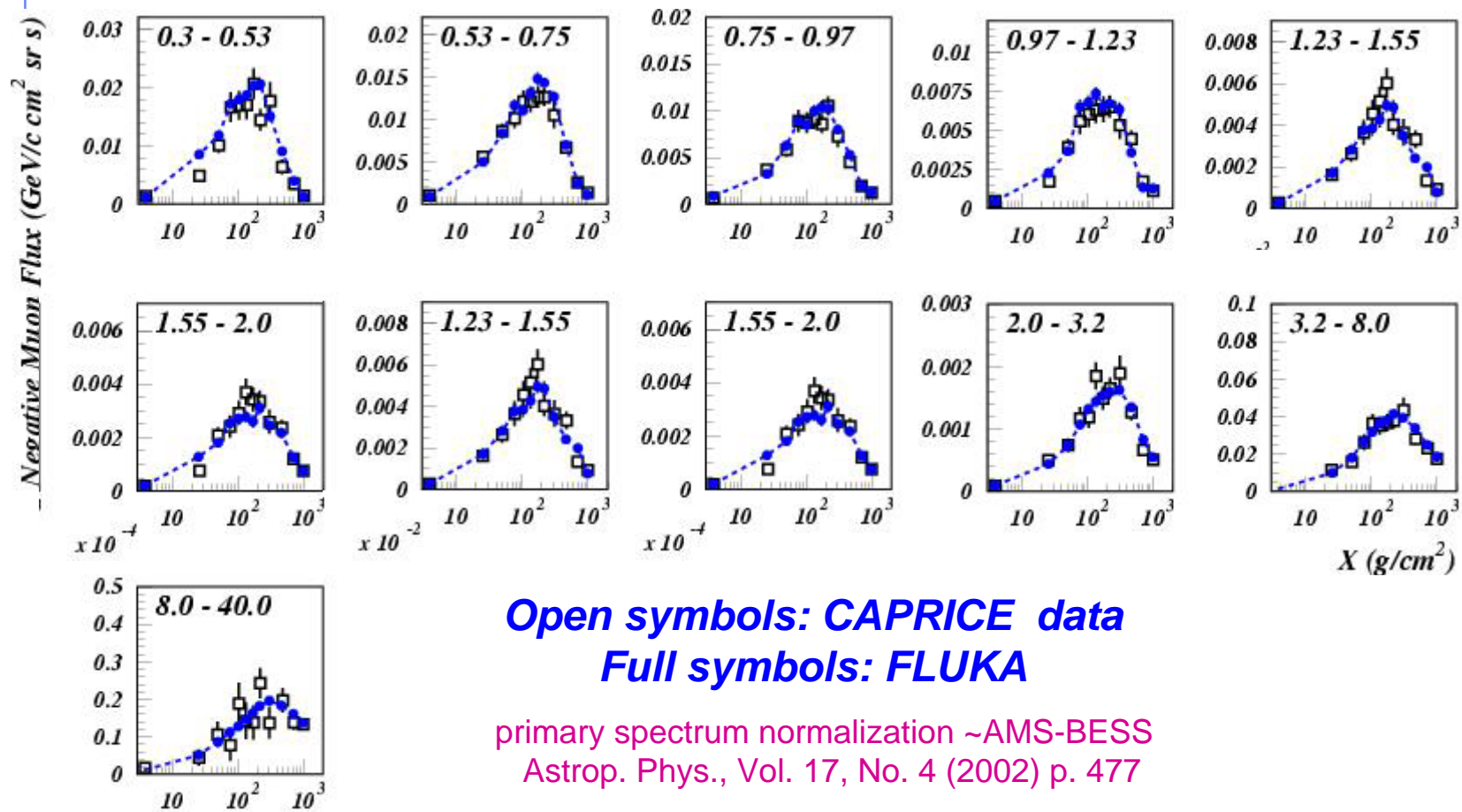
The first 3-D calculation of atmospheric neutrinos was done with FLUKA.

The enhancement in the horizontal direction, which cannot be predicted by a 1-D calculation, was fully unexpected, but is now generally acknowledged.

In the figure: angular distribution of  $\nu_\mu$ ,  $\bar{\nu}_\mu$ ,  $\nu_e$ ,  $\bar{\nu}_e$ .

In red: 1-D calculation

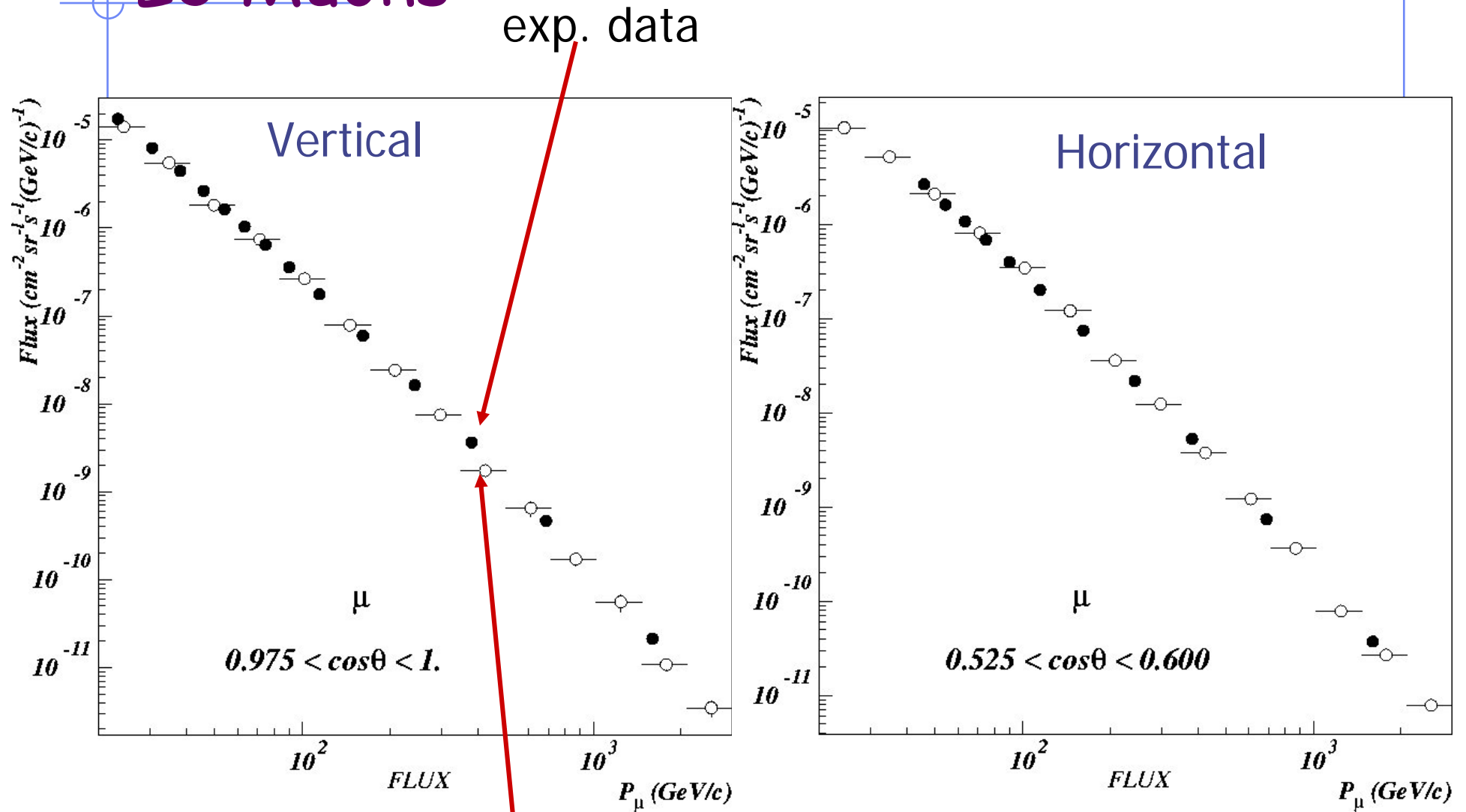
# Negative muons at floating altitudes: CAPRICE94



# Some recent achievements:

(S. Muraro, PhD thesis Milano)

## L3 Muons

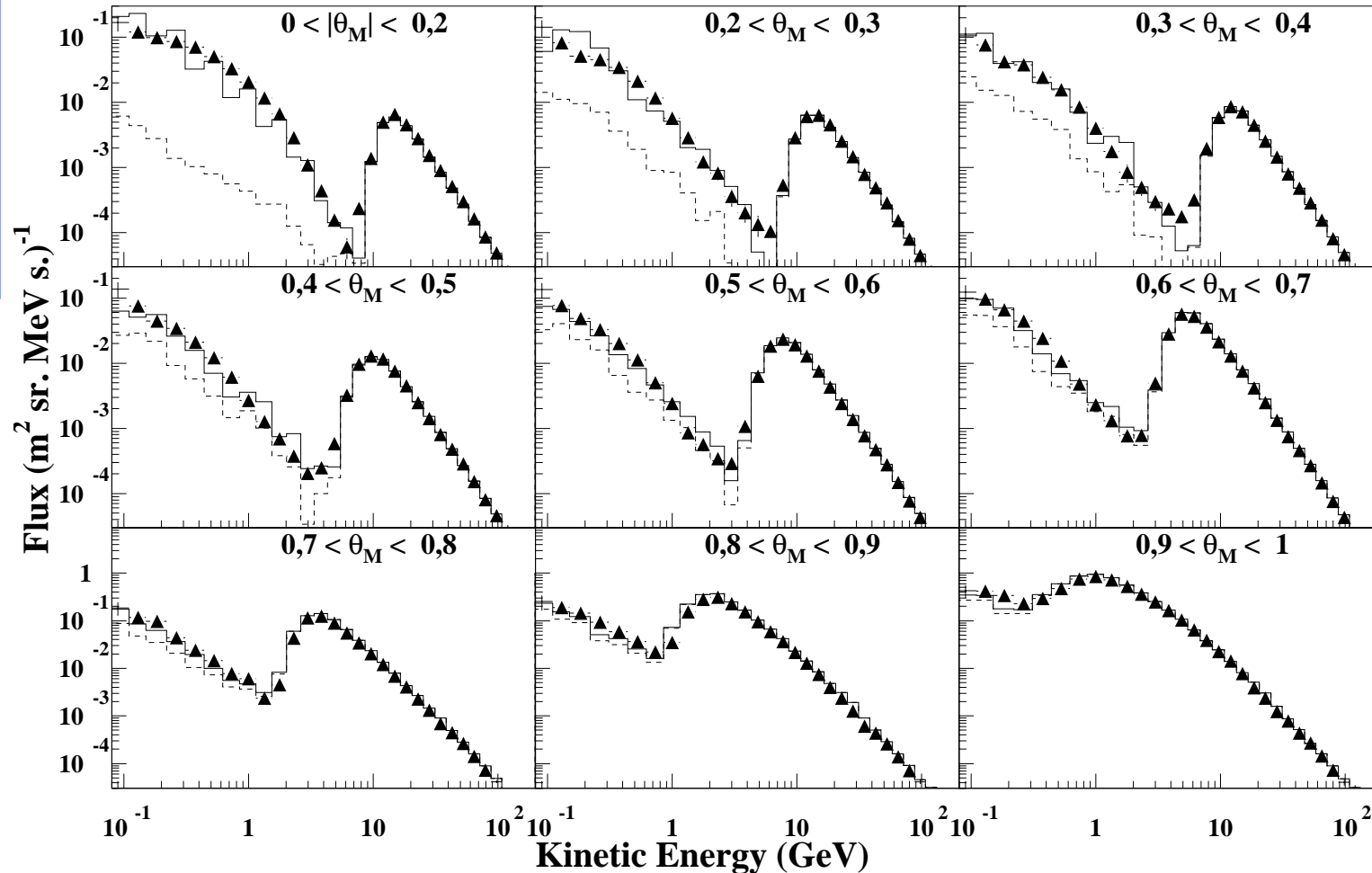


FLUKA simulation

Paola Sala, HSS06

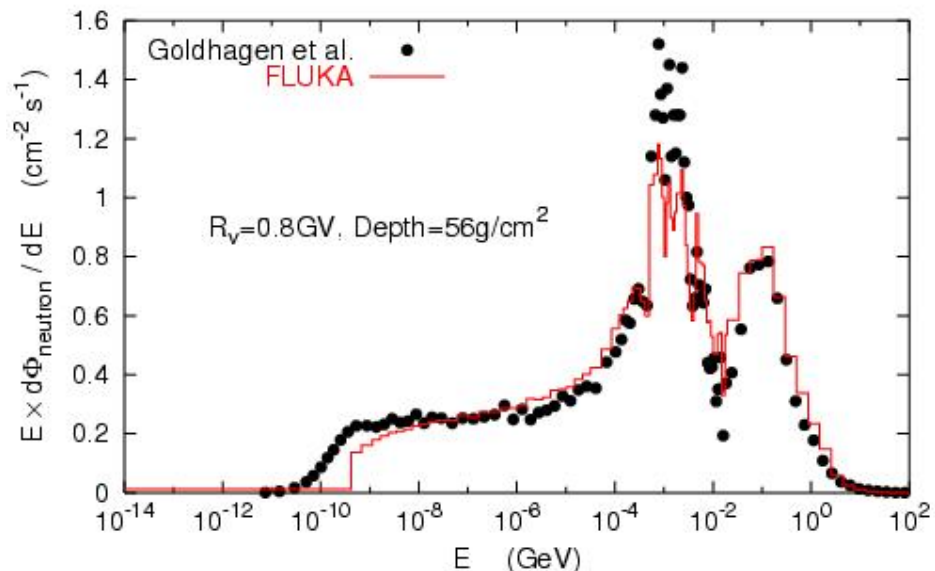
# Comparison with AMS data

Protons and leptons below the geomagnetic cutoff have been measured by the AMS experiment at altitudes 370-390 Km, latitude  $\pm 51.7^\circ$  Astrop. Phys. 20,221 (2003)



Downgoing proton flux, simulation(solid line) AMS data(triangles).  $\theta_M$  is the geomagnetic latitude in radians

# Neutrons on the ER-2 plane at 21 km altitude



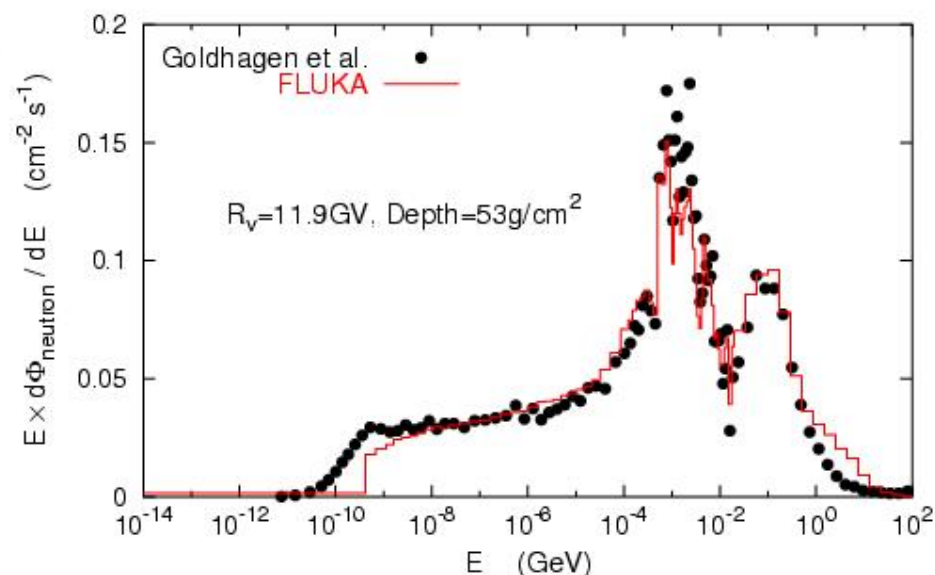
FLUKA calculations:

Roesler et al., Rad. Prot. Dosim. 98, 367 (2002)

Measurements:

Goldhagen et al., NIM A476, 42 (2002)

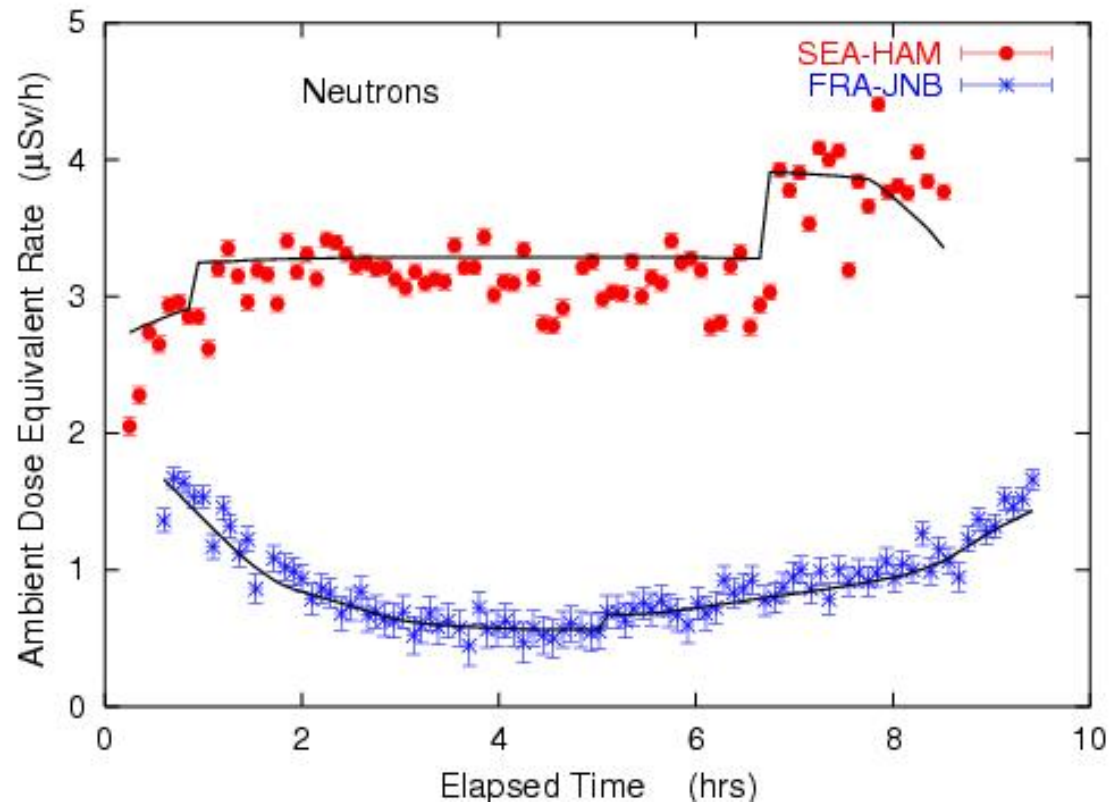
Note one order of magnitude difference depending on latitude





# Dosimetry Applications

Roesler et al.,  
Rad. Prot. Dosim.  
98, 367 (2002)



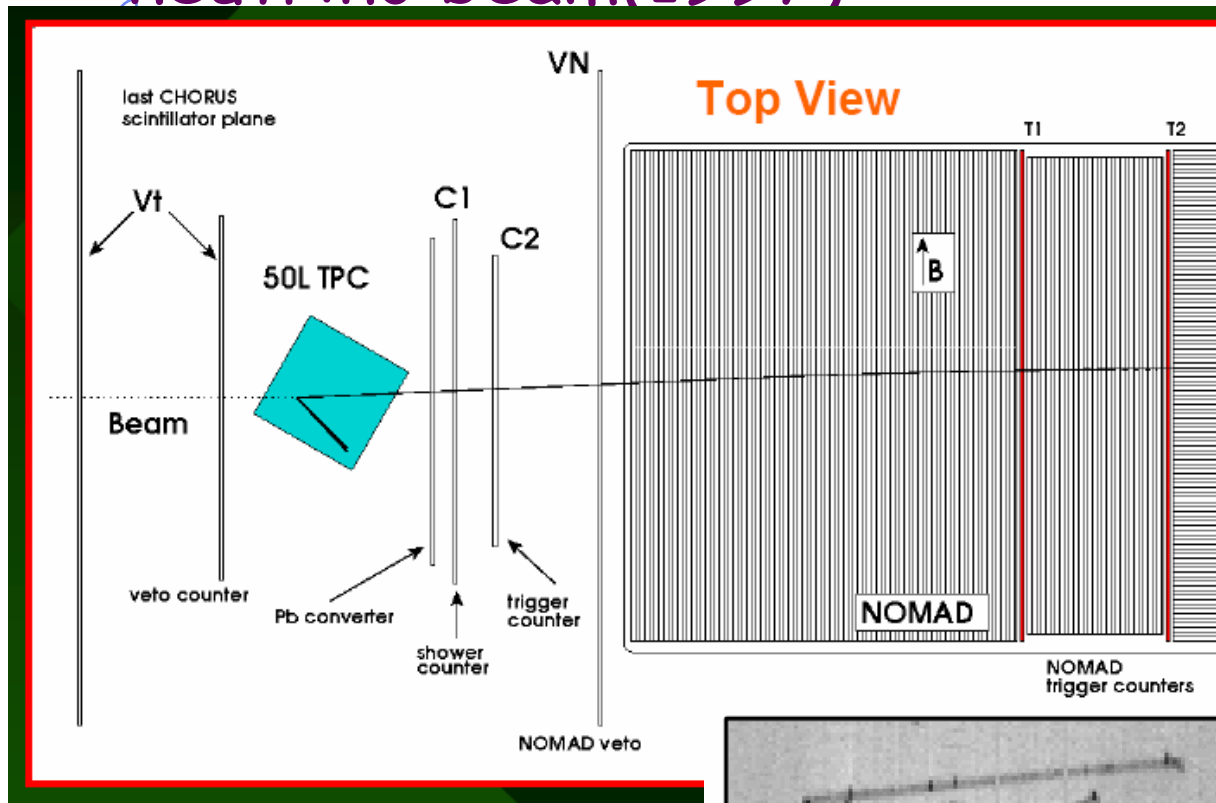
Ambient dose equivalent from neutrons at solar maximum on commercial flights from Seattle to Hamburg and from Frankfurt to Johannesburg.

Solid lines: FLUKA simulation



# **Neutrino Interactions**

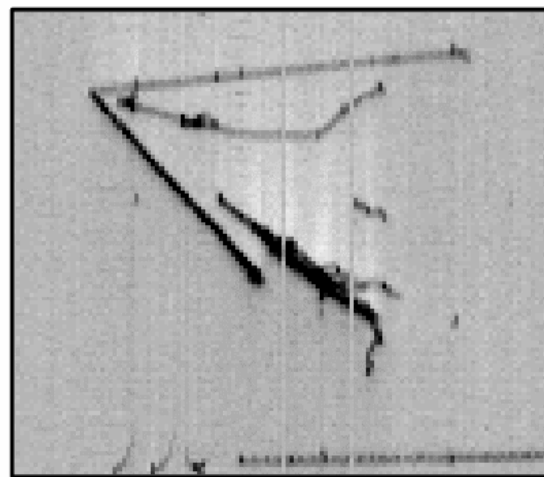
# The 50l LAr TPC in the WANF neutrino beam(1997)



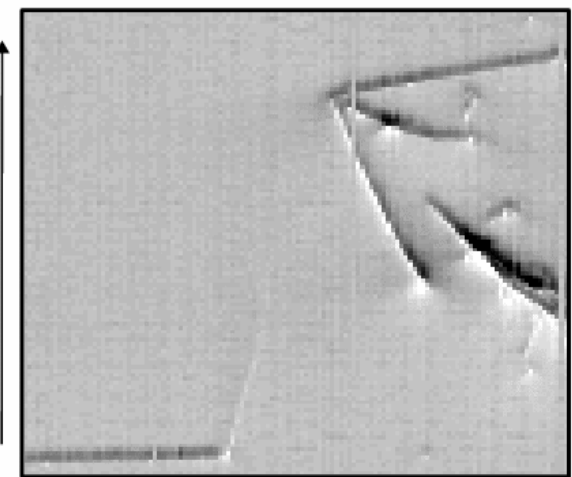
Trigger and  $\mu$   
reconstruction: NOMAD

Event selection:  
"GOLDEN sample"  
== 1  $\mu$  and 1 proton

ICARUS coll. Nufact06

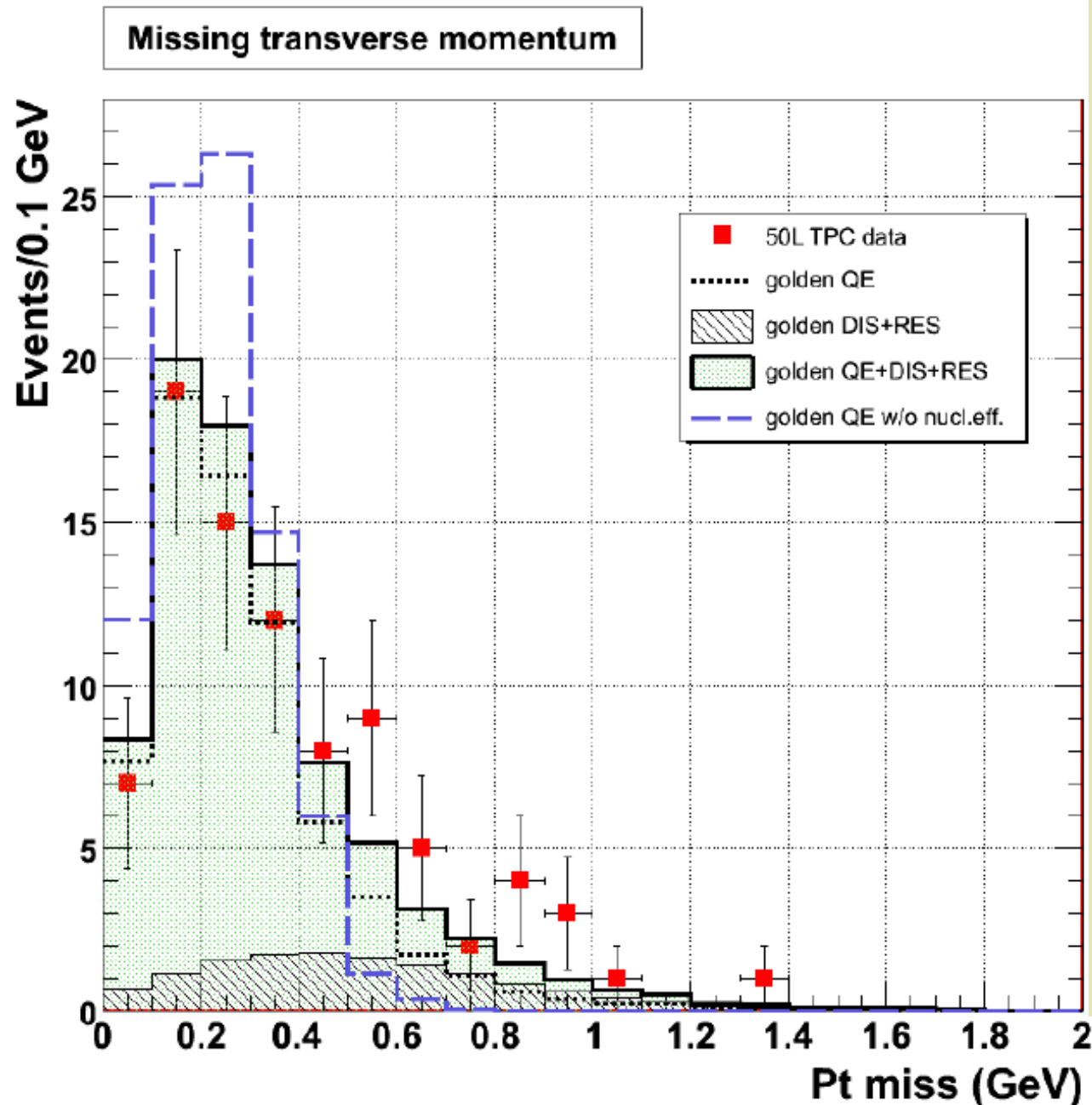


Collection wires. (128 wires: 32 cm.)



Induction wires. (128 wires: 32 cm.)

Time (1300 samples: 47 cm)

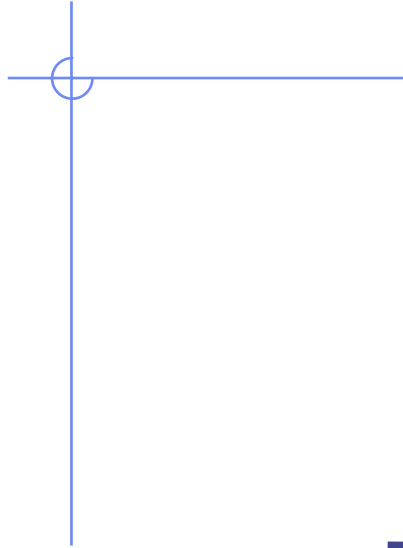


- from 400 QE - golden fraction 16%
- background - additional 20% finally expected

$80 \pm 9(\text{stat.}) \pm 13(\text{syst.})$   
→ mainly QE fraction and beam simul)

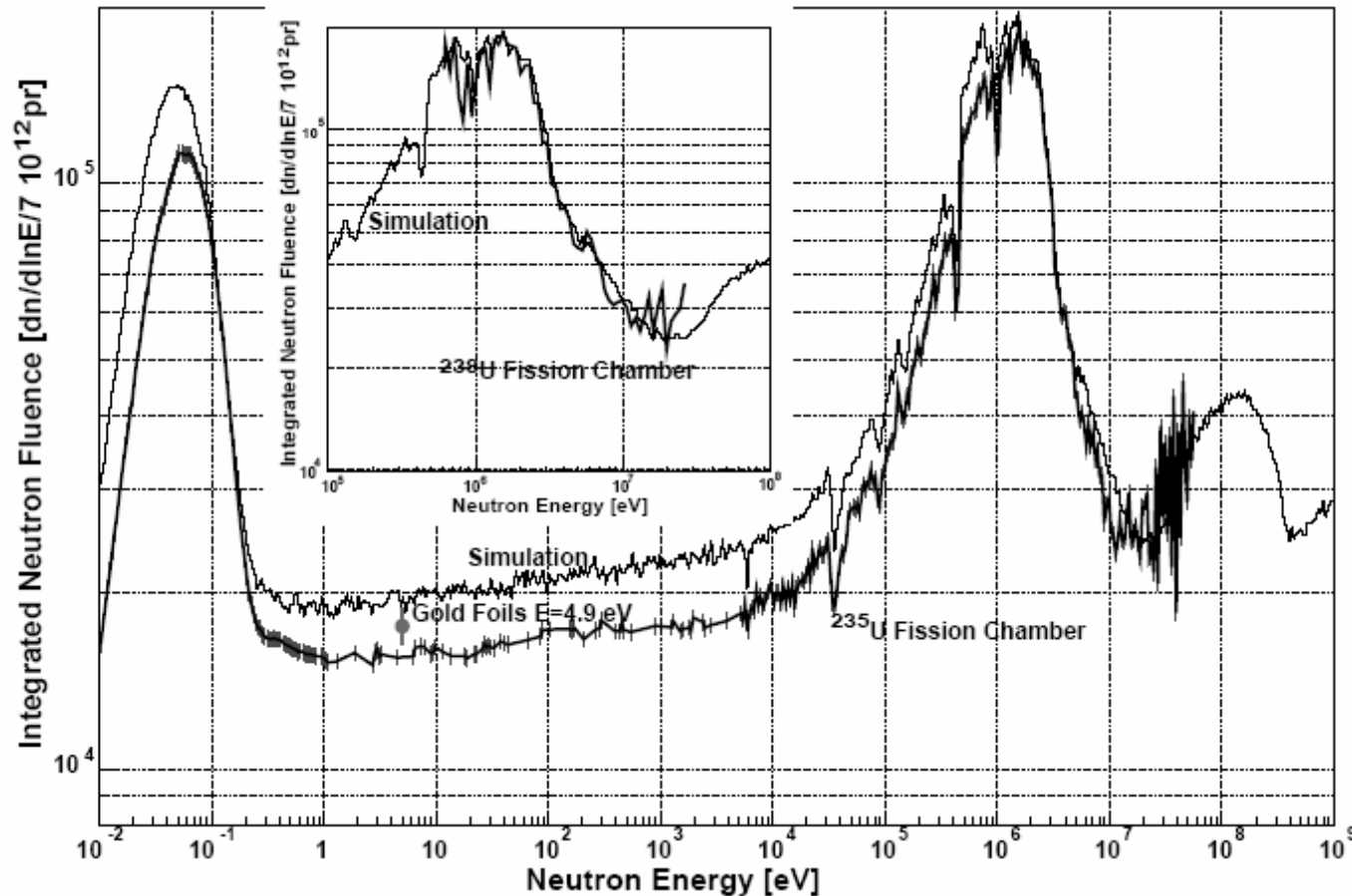
to be compared with **86** events observed

Very good consistency with expectations



**n-tof**

# n-TOF



The n-tof facility  
at CERN:  
neutron beam  
with excellent  
energy resolution  
for cross section  
studies

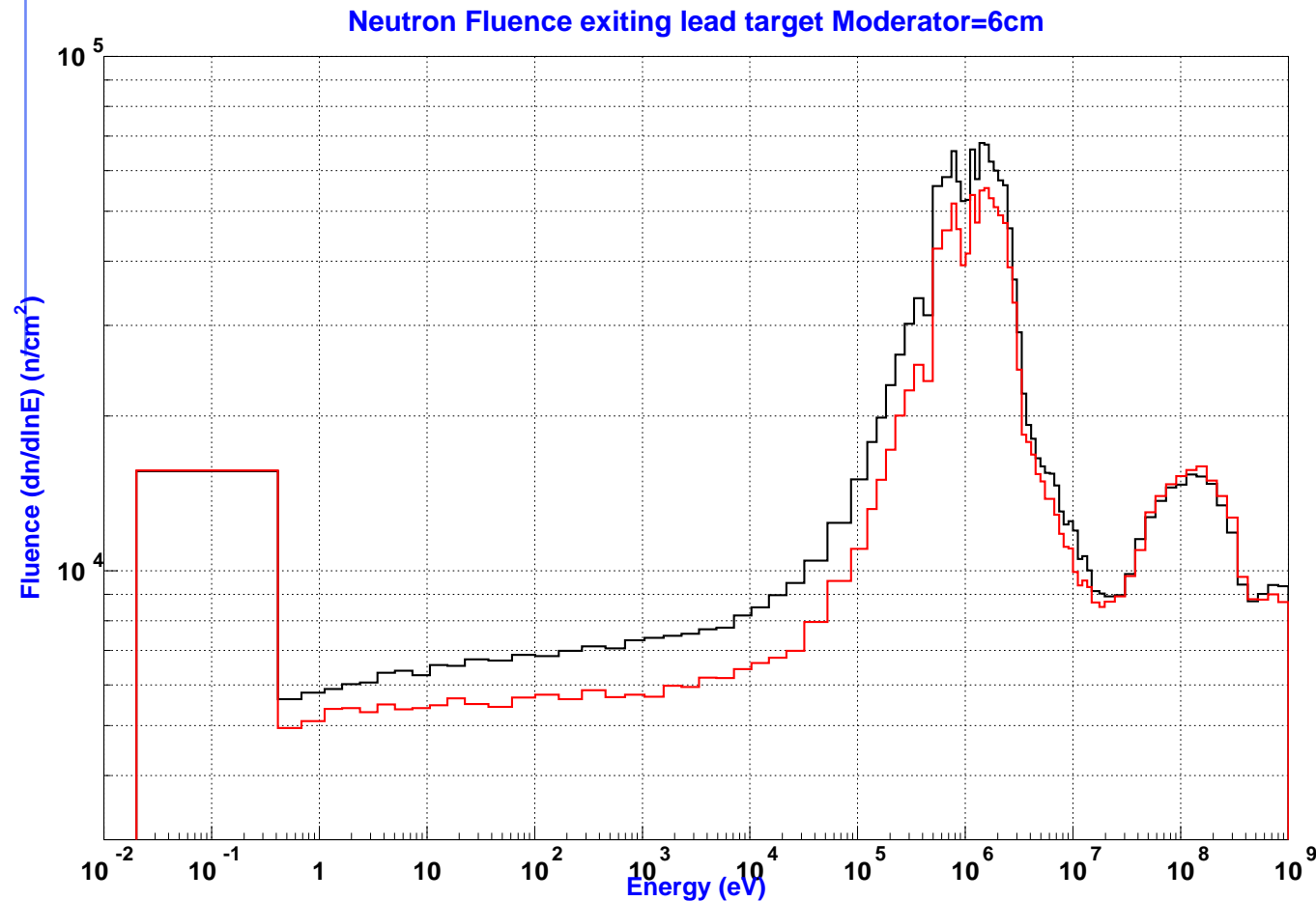
beam from PS :  
20 GeV/c  
protons +  
Huge Lead target  
Water moderator  
neutron beam line

Simulations : FLUKA + C. Rubbia's detailed low energy neutron transport

Assumption : 5 cm water moderator as in the design of the facility

Comparison with measured neutron spectrum shows up to 20% difference in the range  
1-10<sup>5</sup> eV ( published data)

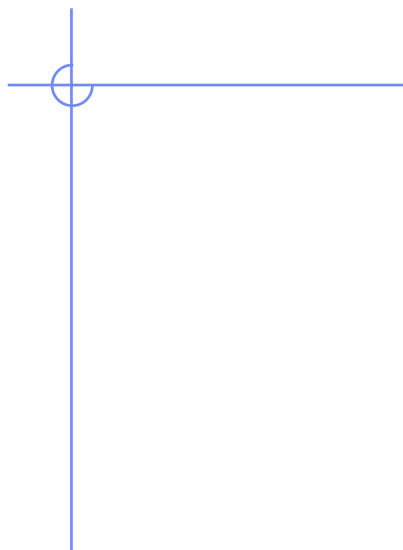
# n-TOF



Preparing for  
Lead target  
dismount-  
Discovery that  
the water layer  
is 6 cm thick  
instead of 5

FLUKA  
simulations with  
6 cm water  
(black)  
compared with 5  
cm (red)

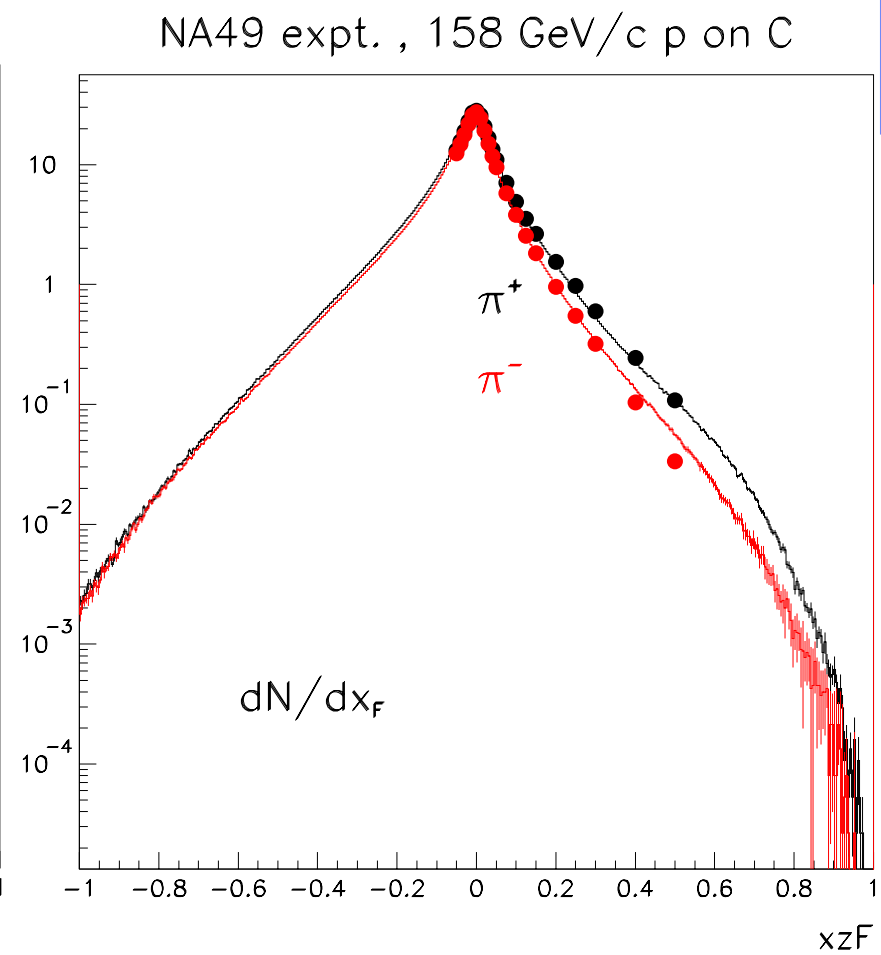
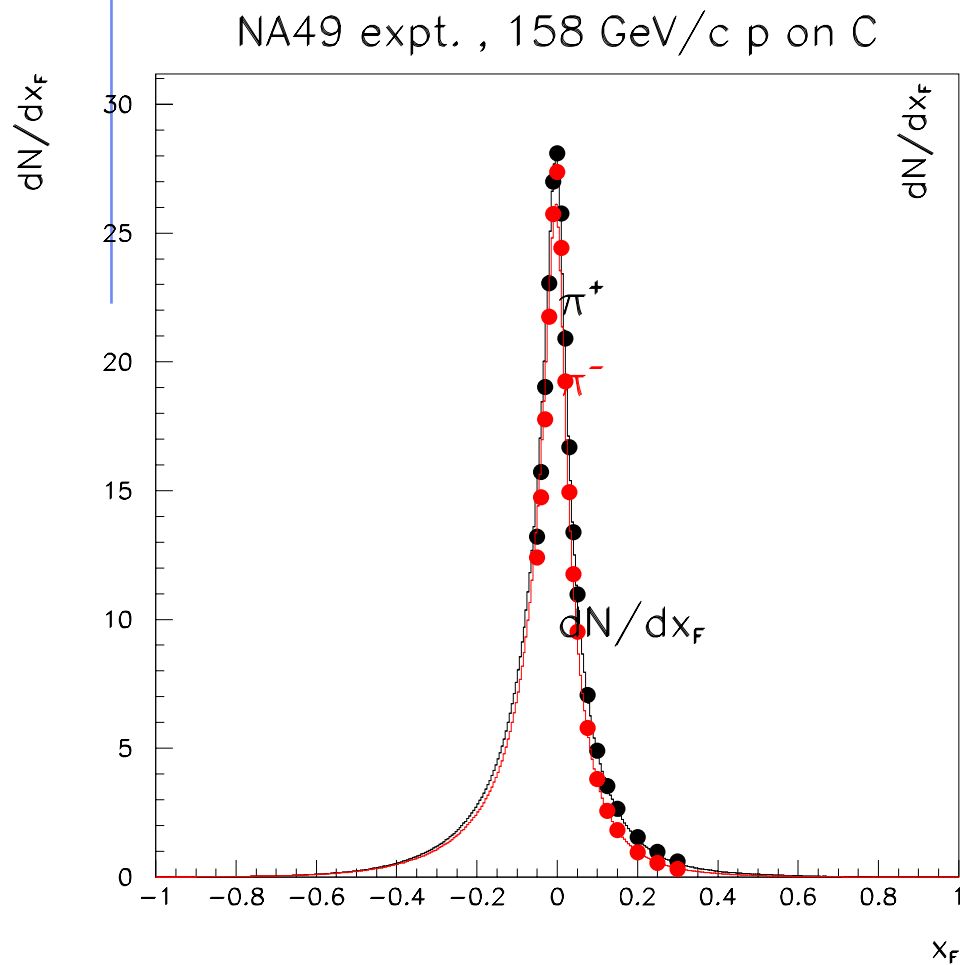
PRELIMINARY, thanks to V. Vlachoudis-CERN



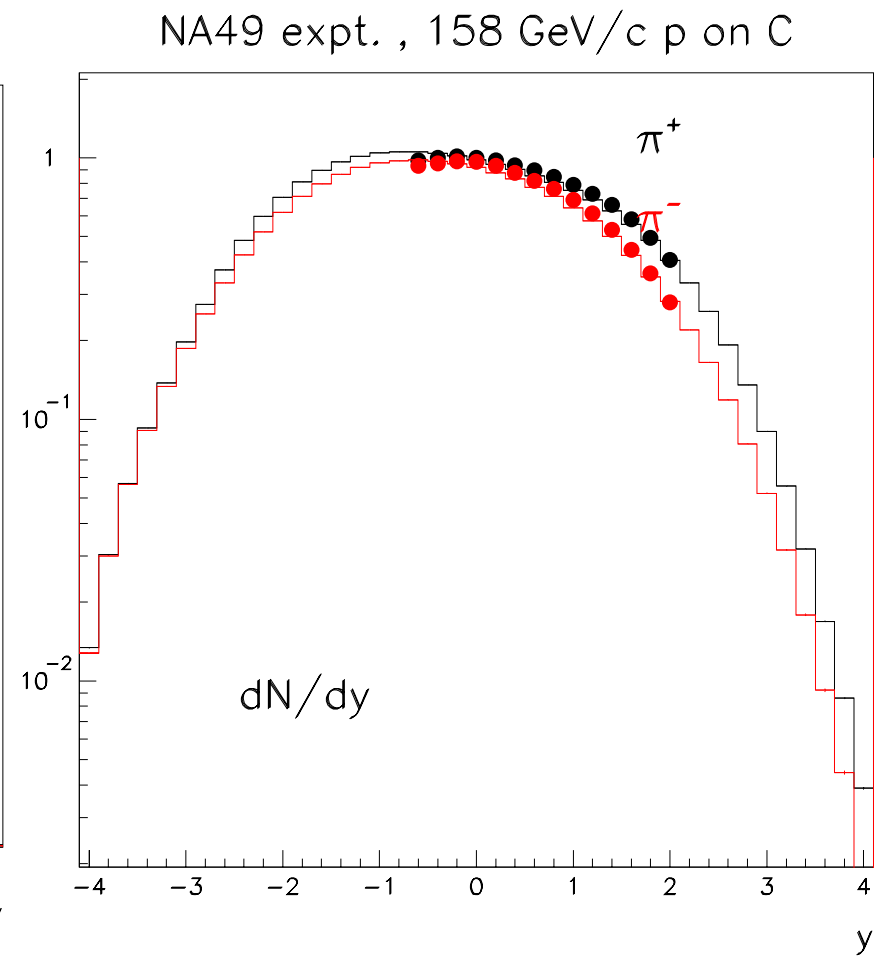
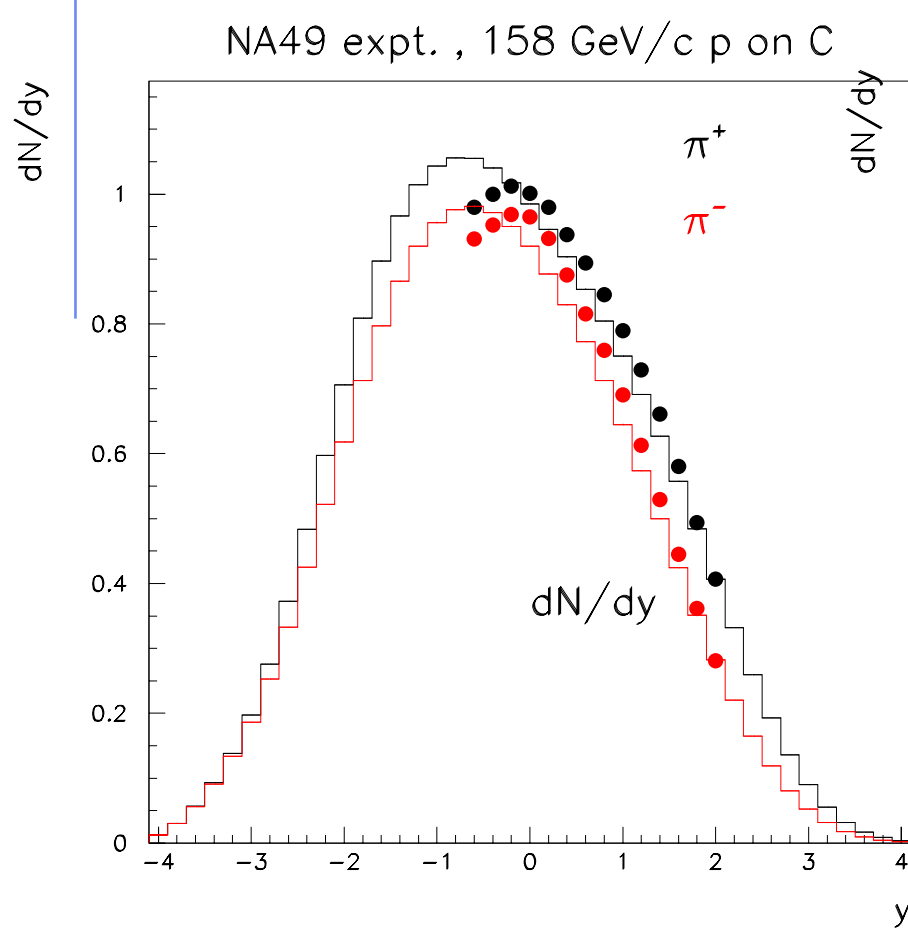
**END**



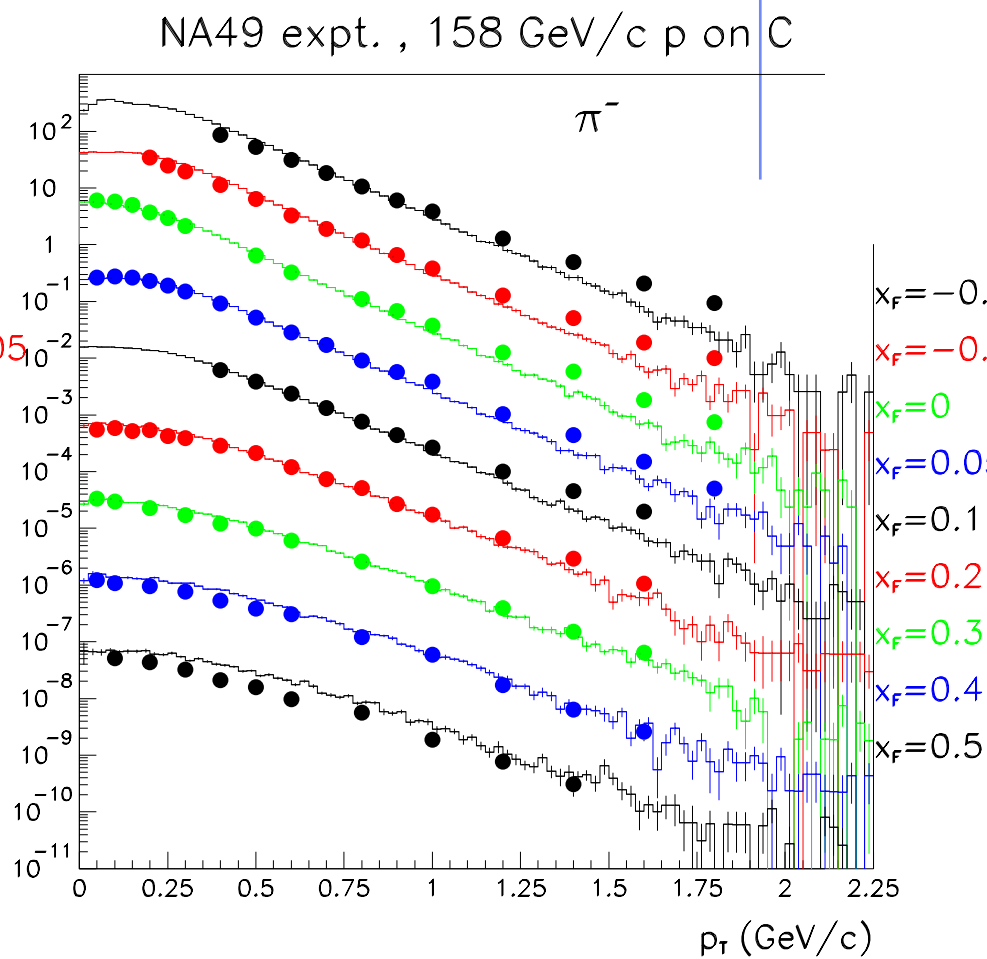
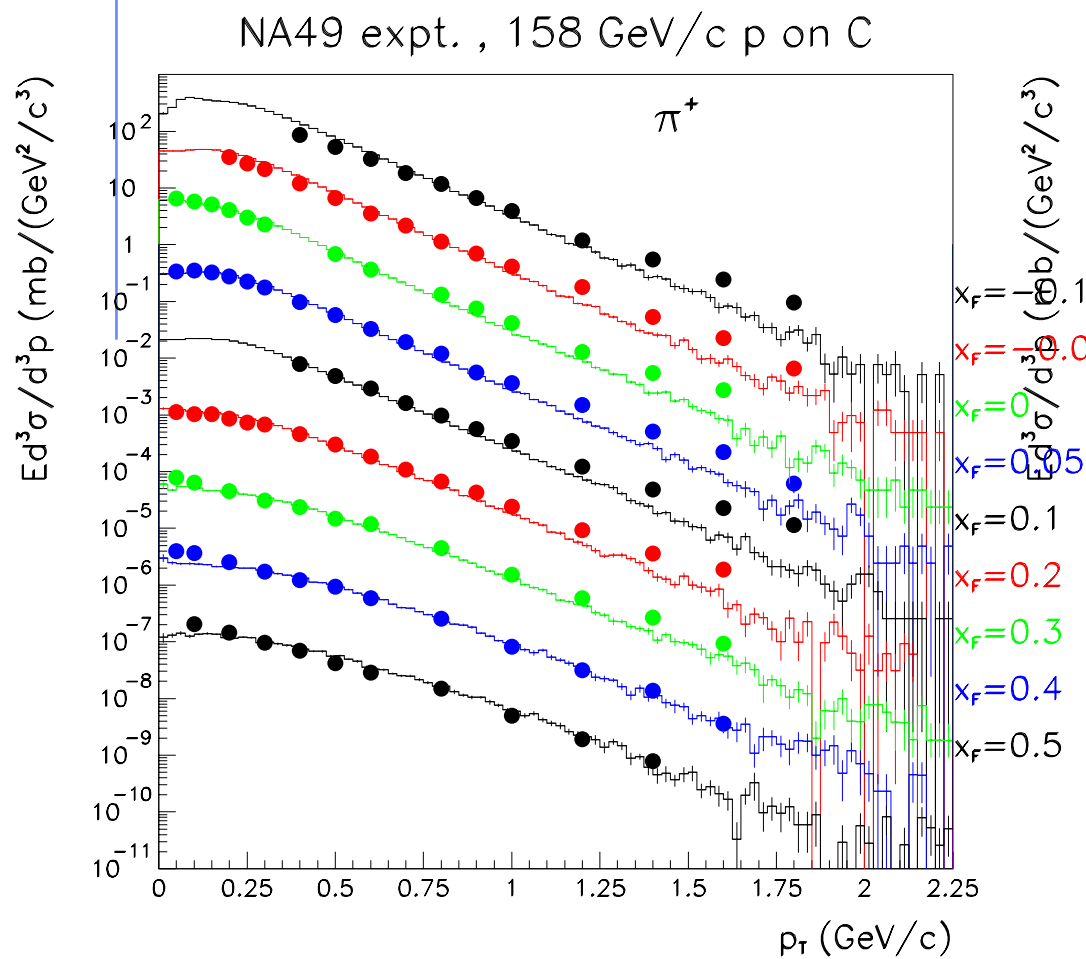
# NA49 p on C



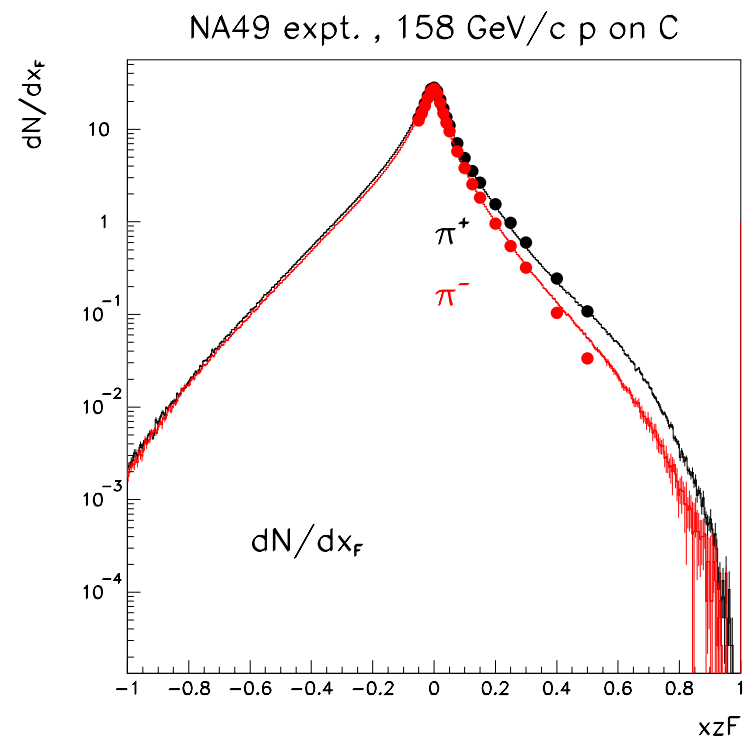
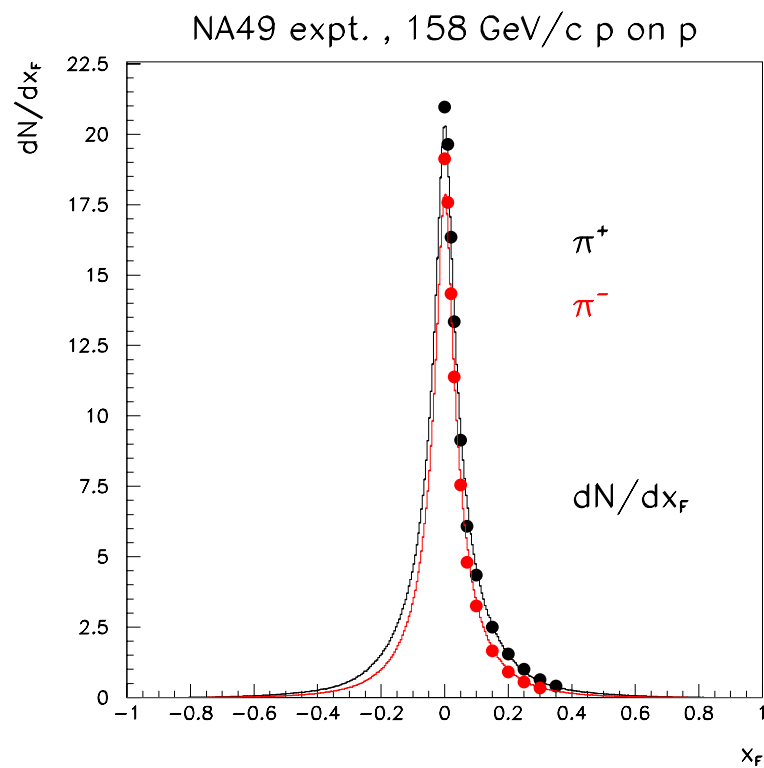
# NA49 p on C



# NA49 p on C



# NA49 p on p



# NA49 p on p

