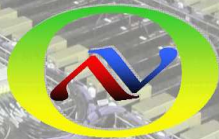


First oscillation results from $\text{NO}\nu\text{A}$

Physics In Collision – University of Warwick
September 18, 2015

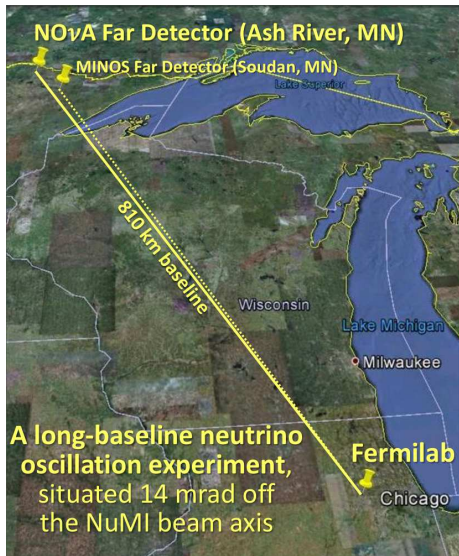
Christopher Backhouse
Caltech
for the $\text{NO}\nu\text{A}$ Collaboration



NO ν A physics goals

$$\nu_{\mu} \rightarrow \nu_e$$

- ▶ Measure θ_{13} via ν_e appearance



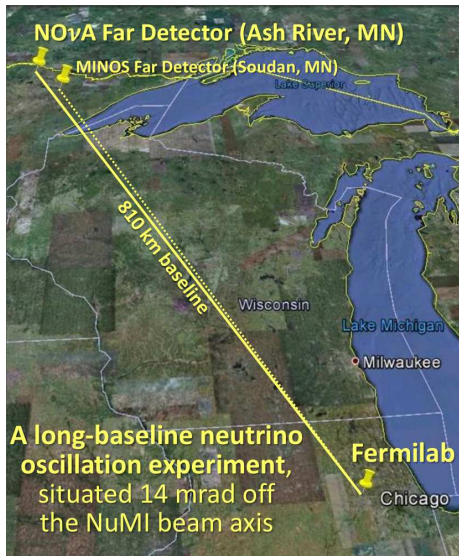
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- ▶ Precision measurements of $|\Delta m_{\text{atm}}^2|$ and θ_{23}
- ▶ Could exclude maximal mixing



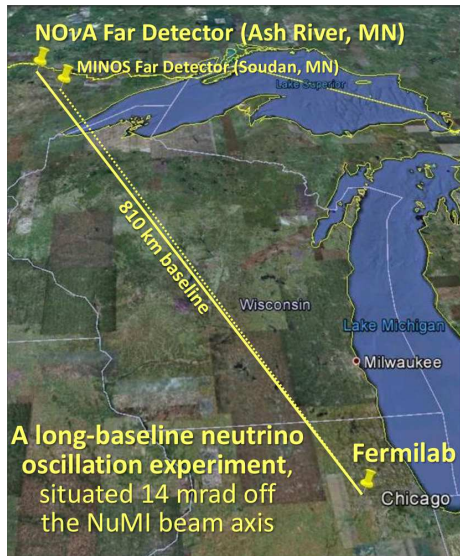
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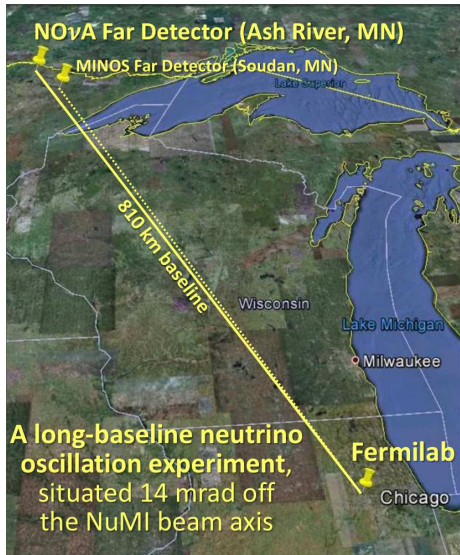
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$$\nu_{\mu} \rightarrow \nu_e \text{ and } \bar{\nu}_{\mu} \rightarrow \bar{\nu}_e$$

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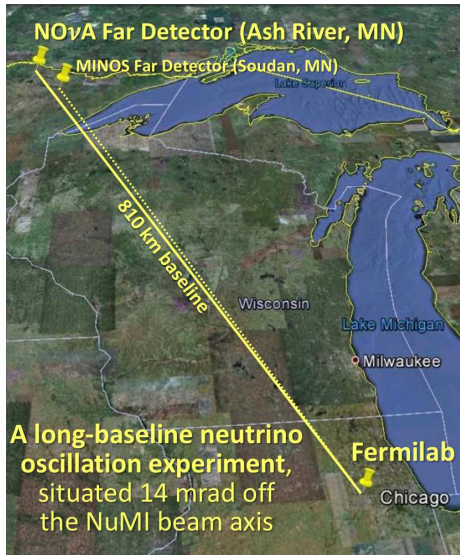
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- ▶ Measure θ_{13} via ν_e appearance
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- ▶ Determine the mass hierarchy
- ▶ Search for $\delta_{CP} \neq 0$

$$\nu_\mu \rightarrow \nu_\mu \text{ and } \bar{\nu}_\mu \rightarrow \bar{\nu}_\mu$$

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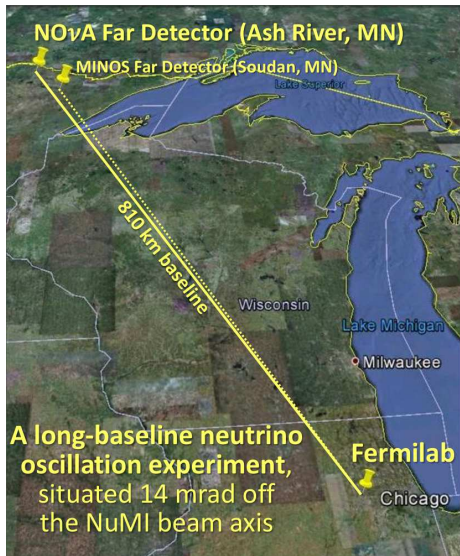
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- ▶ Precision measurements of $|\Delta m_{\text{atm}}^2|$ and θ_{23}
- ▶ Could exclude maximal mixing

And...

- ▶ Cross-sections from the ND
- ▶ Steriles, supernovae, exotica



The NO ν A collaboration

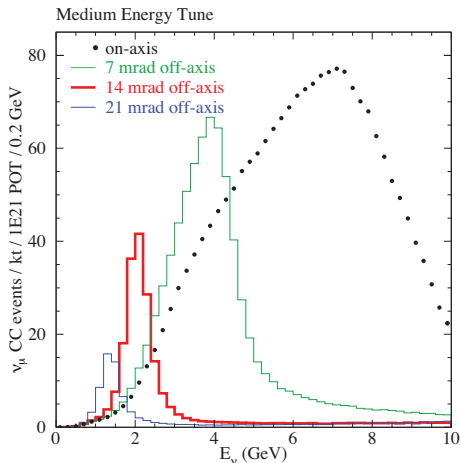


40 institutions, 7 countries, over 200 collaborators

Argonne, Athens, Banaras Hindu, Caltech, CUSAT, Czech Academy of Sciences, Charles, Cincinnati, Colorado State, Czech Technical University, Delhi, Dubna, Fermilab, Goias, IIT-Guwahati, Harvard, IIT-Hyderabad, Hyderabad, Indiana, Iowa State, Jammu, Lebedev Physical Institute, UCL, Michigan State, Minnesota-Twin Cities, Minnesota-Duluth, INR Moscow, Panjab, SDMT, South Carolina, SMU, Stanford, Sussex, Tennessee, Texas-Austin, Tufts, Virginia, Wichita State, William and Mary, Winona State.

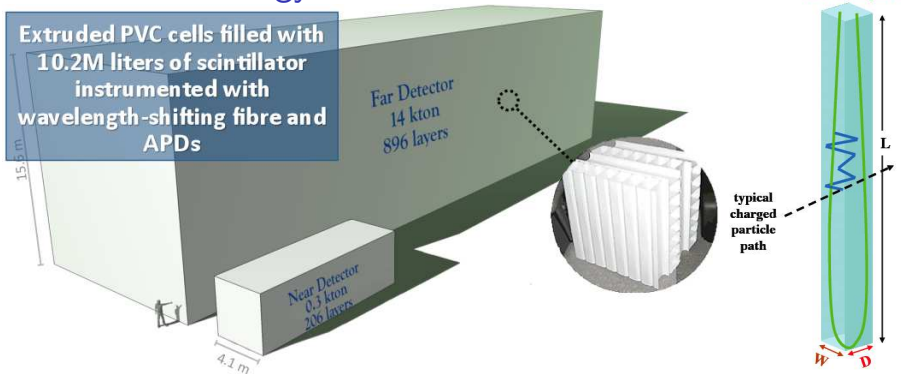
Accelerator and NuMI upgrades

- ▶ 14mrad off-axis location gives sharply-peaked 2GeV ν_μ beam with $\sim 1\%$ ν_e contamination
- ▶ NuMI beam routinely operated at 400kW with 85% uptime.
- ▶ Peak intensity 520kW
- ▶ Using data from Feb 16 2014 to May 15 2015 with detector still under construction
- ▶ Total of 3.45×10^{20} POT
- ▶ Equivalent of 2.74×10^{20} POT with full 14kt detector



14mrad off-axis beam peaks sharply at 2GeV

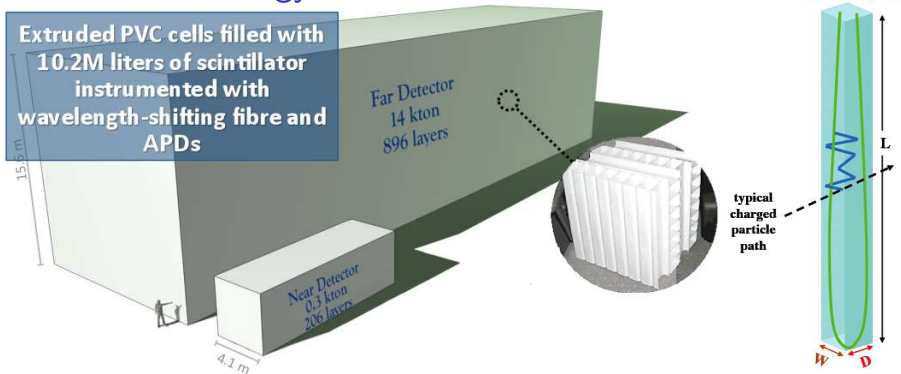
Detector technology



- ▶ Fine-grained low-Z, highly active, tracking calorimeter
- ▶ 64% liquid scintillator by mass
- ▶ WLS fibres looped in 4x6cm cells of PVC extrusion
- ▶ Each to one of 32 pixels of Hamamatsu APD
- ▶ ~85% quantum eff. Gain ~100×. Cooled to -15°C



Detector technology



Far Detector

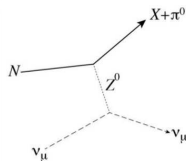
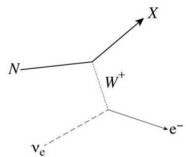
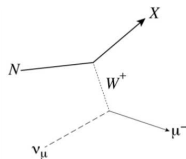
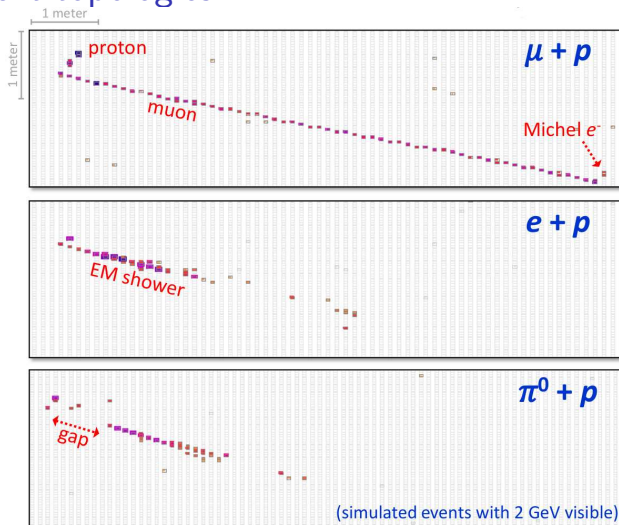
- ▶ 14 kton
- ▶ 344,000 channels
- ▶ On the surface

Near Detector

- ▶ 0.3 kton
- ▶ 18,000 channels

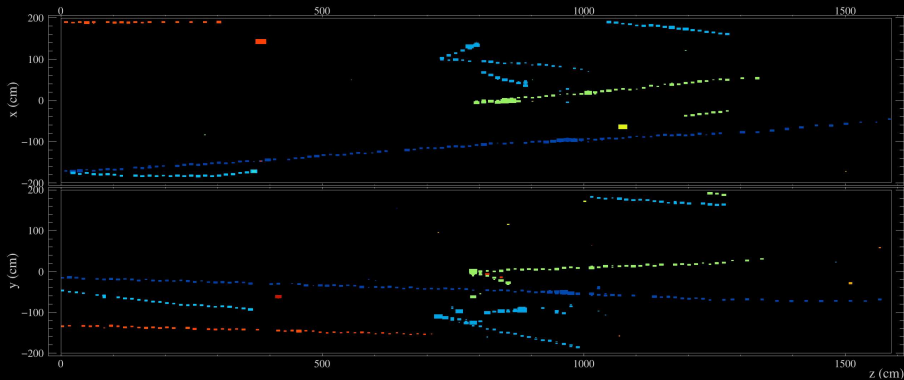


Event topologies



- ▶ Very good granularity, especially considering scale
- ▶ $X_0 = 38\text{cm}$ (6 cell depths, 10 cell widths)

ND neutrinos



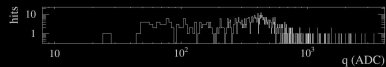
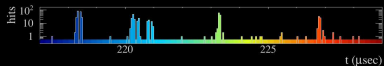
NOvA - FNAL E929

Run: 10407 / 1

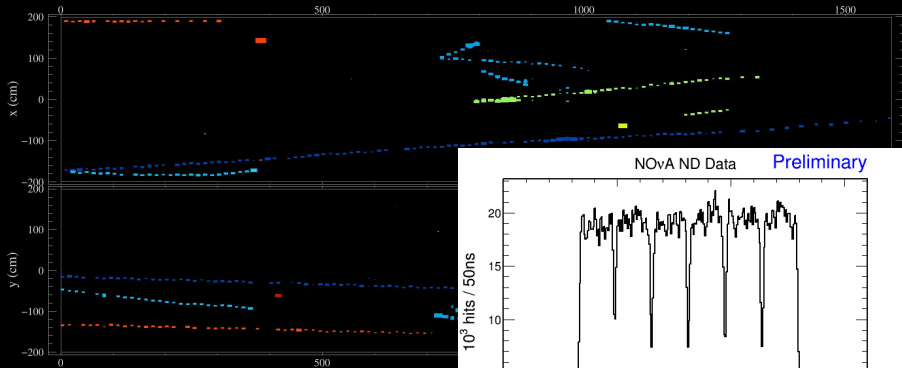
Event: 27950 / --

UTC Thu Sep 4, 2014

05:28:44.034495968



ND neutrinos



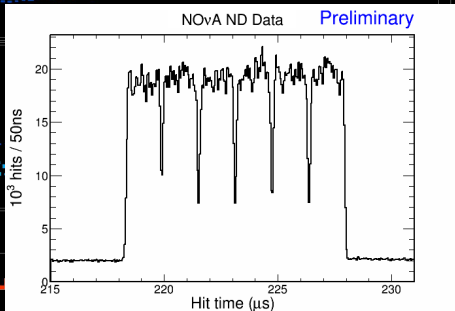
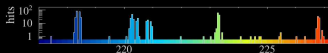
NOvA - FNAL E929

Run: 10407 / 1

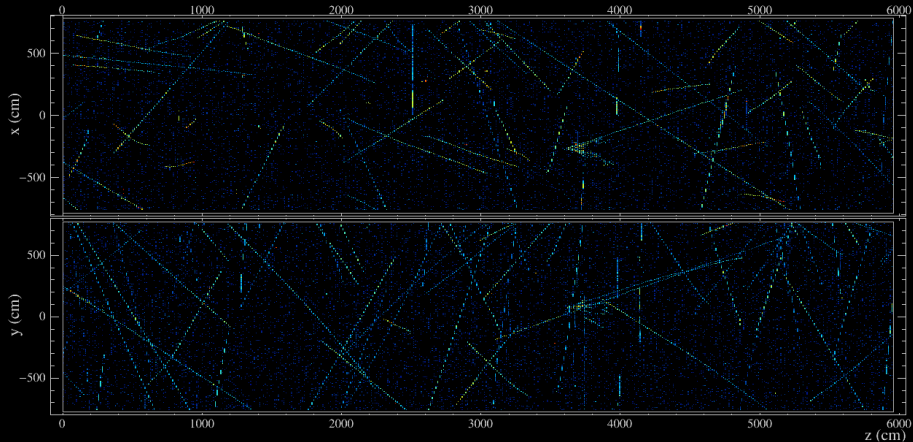
Event: 27950 / --

UTC Thu Sep 4, 2014

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FD neutrino search



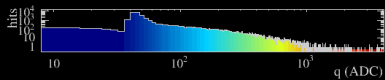
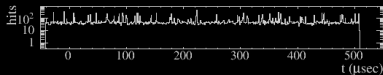
NO ν A - FNAL E929

Run: 18620 / 13

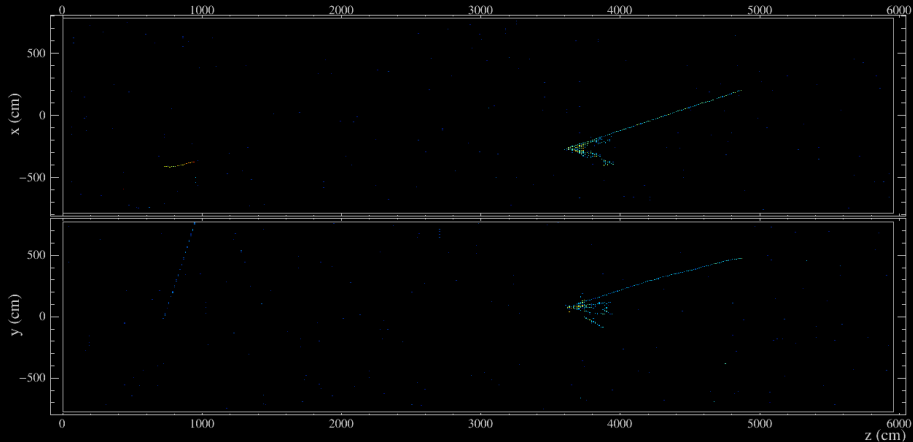
Event: 178402 / --

UTC Fri Jan 9, 2015

00:13:53.087341608



FD neutrino search



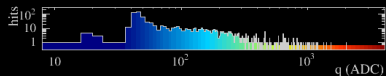
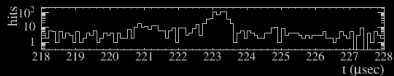
NOvA - FNAL E929

Run: 18620 / 13

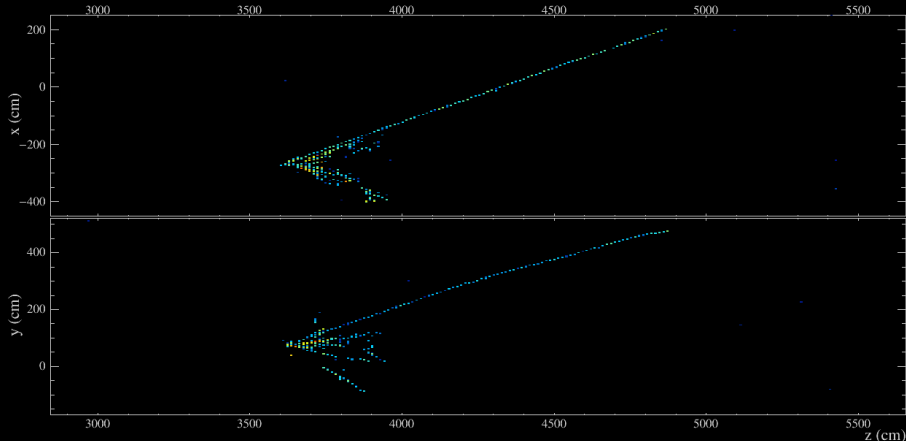
Event: 178402 / --

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FD neutrino search



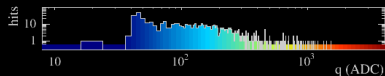
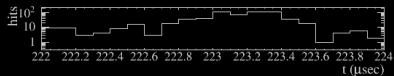
NOνA - FNAL E929

Run: 18620 / 13

Event: 178402 / --

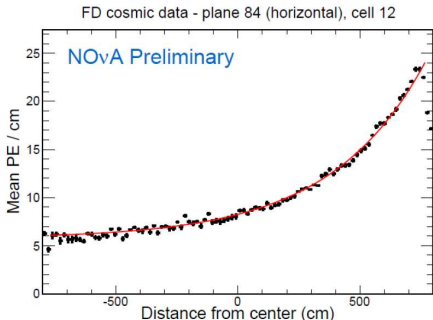
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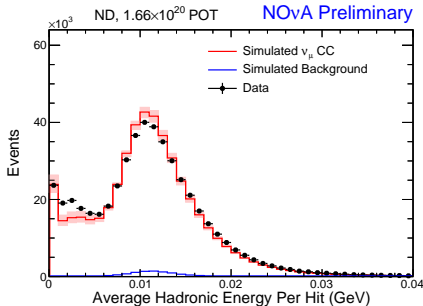
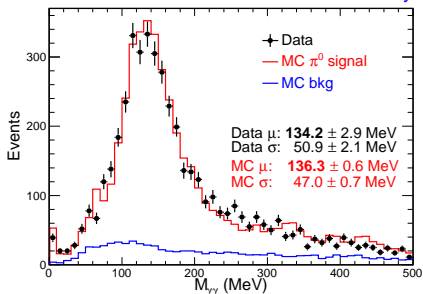
Calibration and energy scale

- ▶ Channel-to-channel and attenuation calibration with cosmic muons
- ▶ Absolute energy scale uses stopping muons as a standard candle
- ▶ Multiple calibration cross-checks
 - ▶ Cosmic muon dE/dx
 - ▶ Beam muon dE/dx
 - ▶ Michel energy spectrum
 - ▶ π^0 mass peak
 - ▶ Hadronic energy/hit



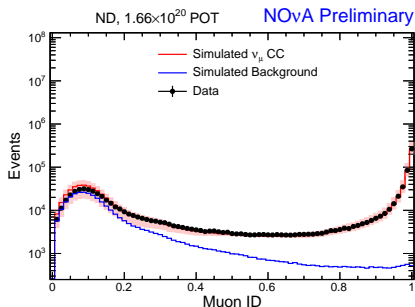
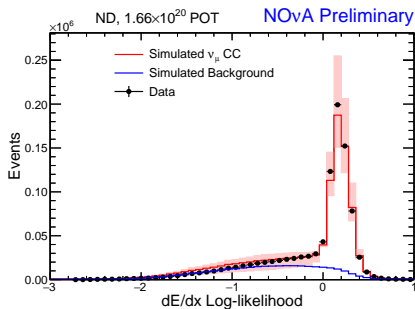
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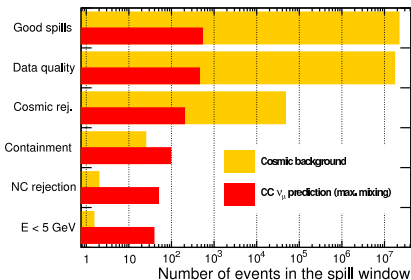


Selecting muon neutrinos

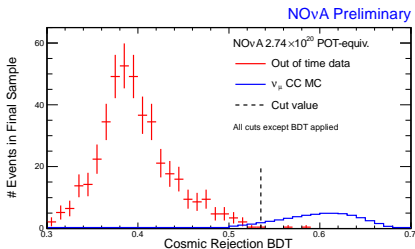
- ▶ Basic containment cuts requiring no activity close to detector walls
- ▶ kNN-based ν_μ classifier using 4 inputs
 - ▶ Track length
 - ▶ dE/dx
 - ▶ Scattering
 - ▶ Fraction of planes that have track-only
- ▶ Selection 70% efficient for ν_μ signal, 2% for NC background



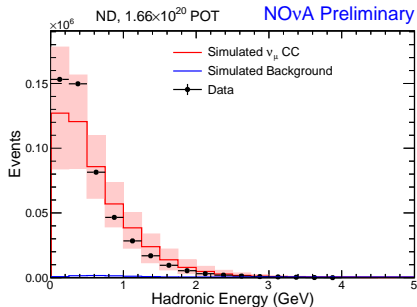
Cosmic rejection for ν_μ analysis



- ▶ Cosmic background rate measured from data adjacent in time to the beam spill window
- ▶ $10\mu\text{s}$ spill window at $\sim 1\text{Hz}$ gives 10^5 rejection
- ▶ Additional factor 10^7 from event topology plus boosted decision tree based on
 - ▶ Track direction
 - ▶ Track start and end points
 - ▶ Track length
 - ▶ Energy
 - ▶ Number of hits

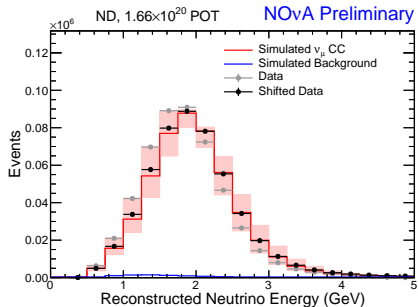
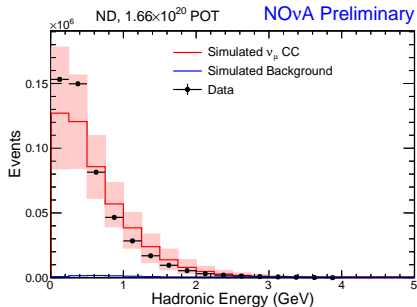


Muon neutrino energy reconstruction



- ▶ Good data/MC agreement for muon neutrino selected events
- ▶ But, 21% more energy in MC hadronic system compared to data
- ▶ Recalibrate to make neutrino energy peak match
- ▶ Take the entire shift as a systematic \rightarrow 6% neutrino energy scale uncertainty

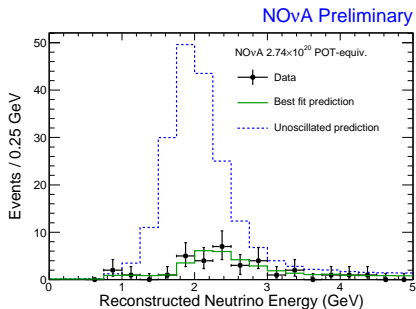
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- ▶ Use ND data to predict FD neutrino spectrum

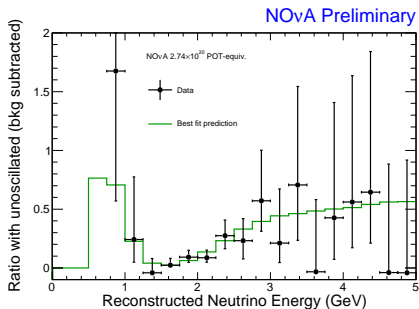
ν_μ disappearance results

- ▶ Expect 201 events w/o oscillations
- ▶ We observe 33 events



ν_μ disappearance results

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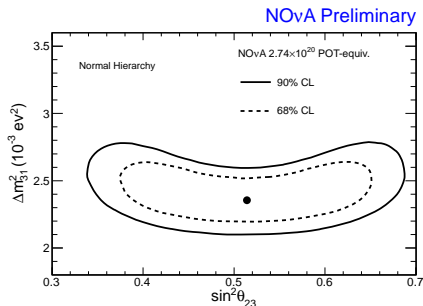
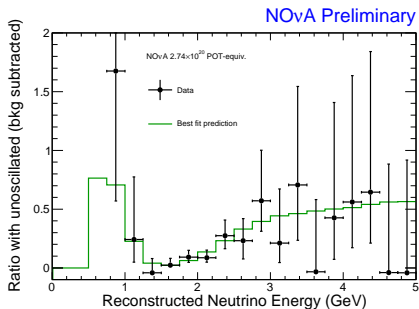


ν_μ disappearance results

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- ▶ $\Delta m_{32}^2 = +2.37_{-0.15}^{+0.16}$ or
 $-2.40_{-0.17}^{+0.14}$

- ▶ $\sin^2 \theta_{23} = 0.51 \pm 0.10$



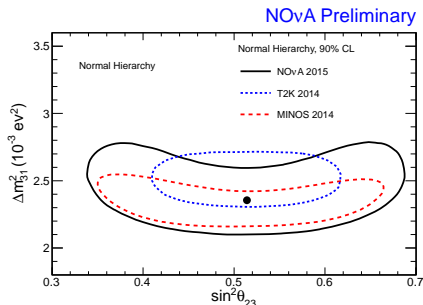
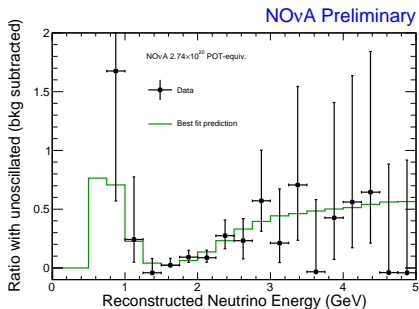
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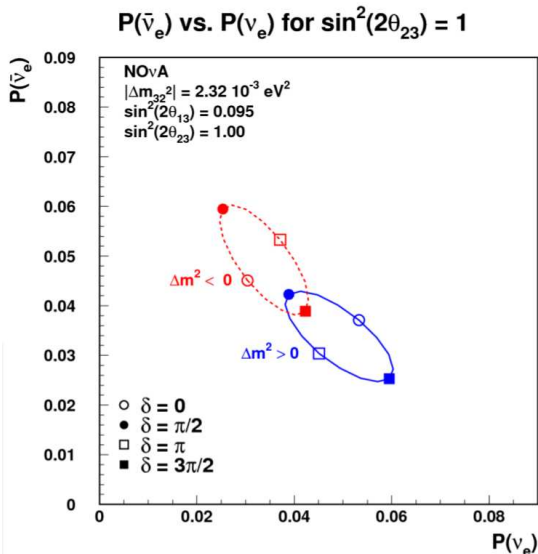
- ▶ $\sin^2 \theta_{23} = 0.51 \pm 0.10$

- ▶ Very competitive with $\sim 8\%$ of nominal exposure



Principle of the ν_e measurement

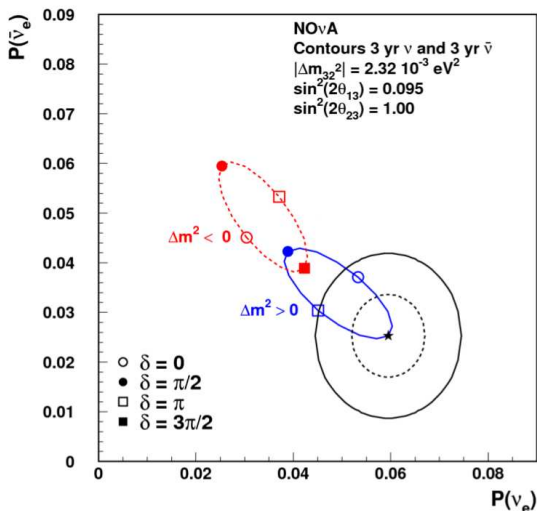
- ▶ To first order, NO ν A measures $P(\nu_\mu \rightarrow \nu_e)$ and $P(\bar{\nu}_\mu \rightarrow \bar{\nu}_e)$ evaluated at 2GeV
- ▶ These depend differently on $\text{sign}(\Delta m^2)$ and δ_{CP}



Principle of the ν_e measurement

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- ▶ These depend differently on $\text{sign}(\Delta m^2)$ and δ_{CP}
- ▶ Ultimately constrain to some region of this space

1 and 2 σ Contours for Starred Point



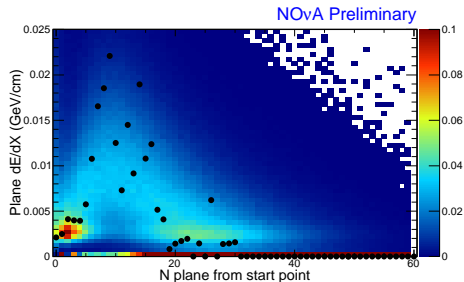
Selecting electron neutrinos

- ▶ Two ν_e classifiers developed based on very different techniques

- ▶ **LID** Uses PDFs of leading shower longitudinal and transverse profiles
- ▶ **LEM** Finds best matches in a Monte Carlo library based on entire event topology

Selecting electron neutrinos

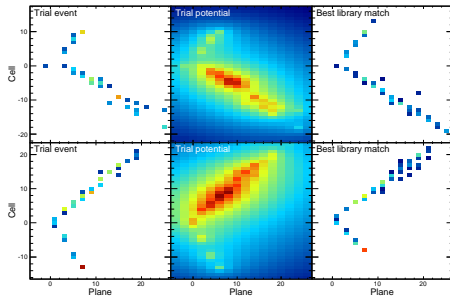
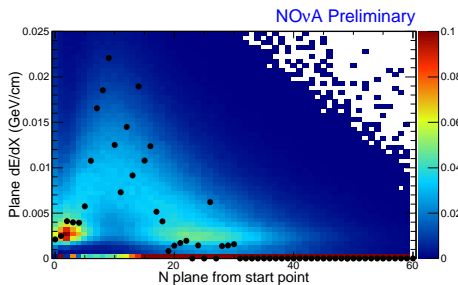
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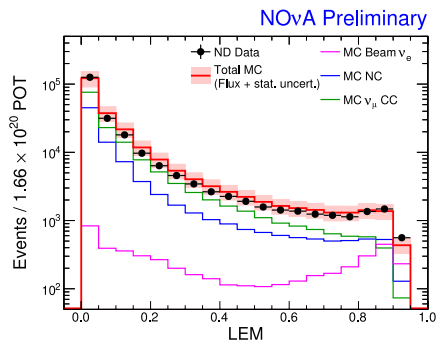
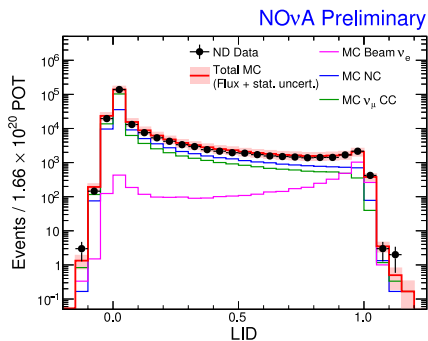
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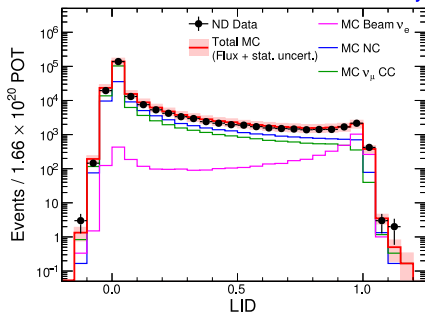


- ▶ Good separation of signal from background, including cosmic backgrounds
- ▶ Identical performance. 35% signal efficiency, 0.7% NC events remain. 62% expected overlap between samples

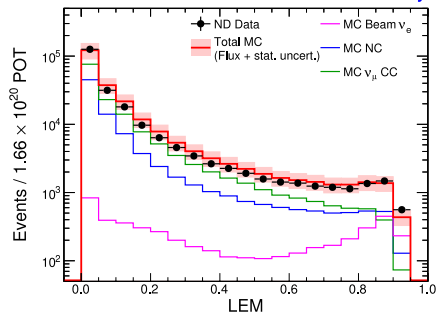
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NOvA Preliminary



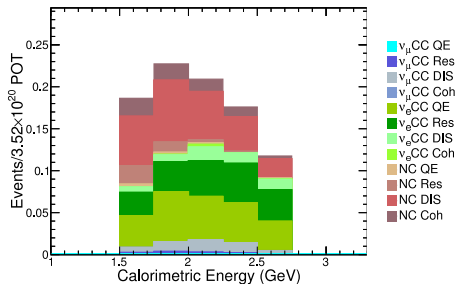
NOvA Preliminary



- ▶ Good separation of signal from background, including cosmic backgrounds
- ▶ Identical performance. 35% signal efficiency, 0.7% NC events remain. 62% expected overlap between samples
- ▶ **Before unblinding FD data, decided to show both results and use LID as the primary selector**

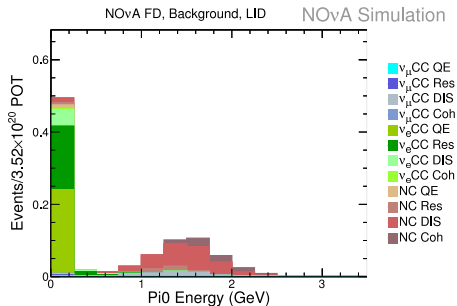
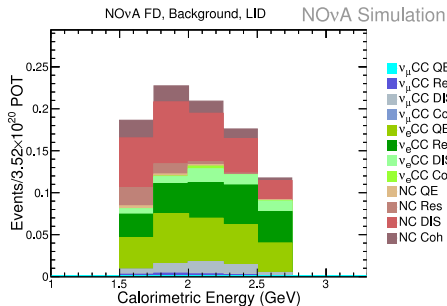
Selecting electron neutrinos

NOvA FD, Background, LID NOvA Simulation



- ▶ Selected FD background dominated by NC DIS and intrinsic beam ν_e contamination

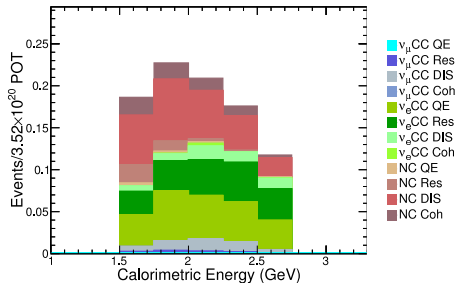
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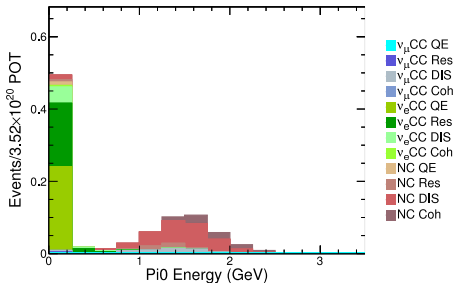
- ▶ Selected FD background dominated by NC DIS and intrinsic beam ν_e contamination
- ▶ $\sim 1.5\text{GeV}$ π^0 in most NCs

Selecting electron neutrinos

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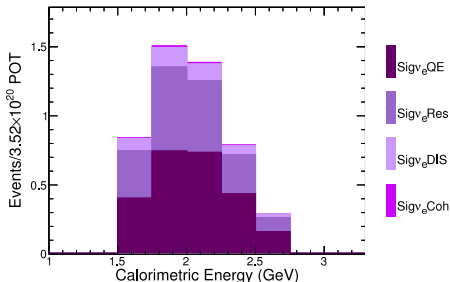


NOvA FD, Background, LID NOvA Simulation



- ▶ Selected FD background dominated by NC DIS and intrinsic beam ν_e contamination
- ▶ $\sim 1.5\text{GeV } \pi^0$ in most NCs
- ▶ Signal dominated by QE and RES, minimal impact dependence on hadronic system

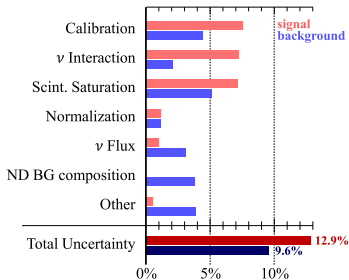
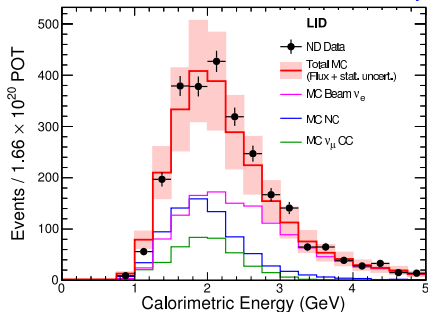
NOvA FD, Signal, LID NOvA Simulation



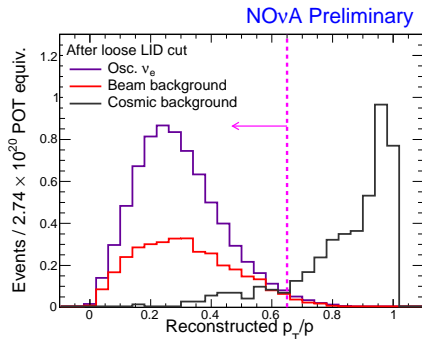
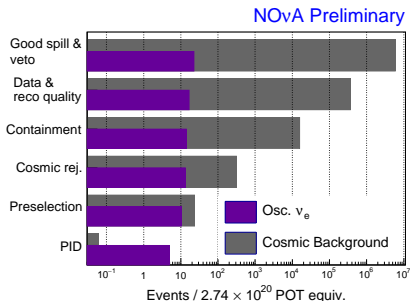
FD prediction

- ▶ Good data/MC agreement for ND energy spectrum
- ▶ Extrapolate 5% ND data background excess bin-by-bin to FD using Far/Near ratio
- ▶ FD signal expectation uses the same procedure and ND ν_μ spectrum as the disappearance analysis
- ▶ Most systematics assessed by modifying the Near and Far MC and calculating the variation in the prediction

NOvA Preliminary



Cosmic rejection for ν_e analysis



- ▶ Containment and topological cuts (e.g. p_T/p)
- ▶ Classifiers themselves provide remaining rejection
- ▶ Achieve 10^8 removal of cosmic rays

- ▶ Measure rate using out-of-time spill data as for ν_μ analysis
- ▶ Cosmic background expectation **0.06 events**

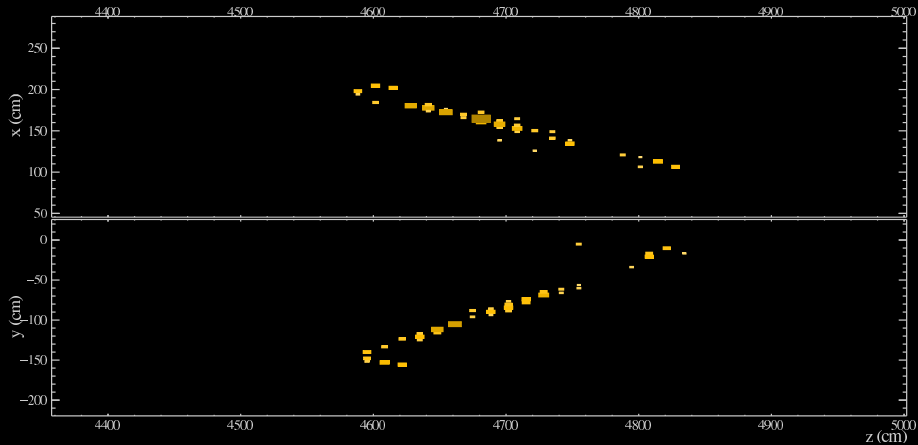
Signal and background predictions

- ▶ Background prediction for both selectors is ~ 1 event $\pm 10\%$ (syst)
- ▶ Few percent dependence on oscillation parameters
- ▶ Dominated by beam ν_e and NC
- ▶ Cosmic background comparable to smallest beam backgrounds

PID	Total bkg	ν_e CC	NC	ν_μ CC	ν_τ CC	cosmic
LID	0.94 ± 0.09	0.46	0.35	0.05	0.02	0.06
LEM	1.00 ± 0.11	0.46	0.40	0.06	0.02	0.06

- ▶ Range of signal predictions
- ▶ NH, $\delta_{CP} = 3\pi/2 \rightarrow 6 \pm 0.7$
- ▶ IH, $\delta_{CP} = \pi/2 \rightarrow 2 \pm 0.3$
(for LID, LEM similar)

Example selected event



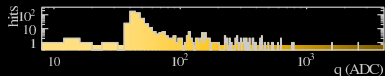
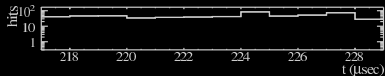
NOνA - FNAL E929

Run: 19165/62

Event: 920415 / -

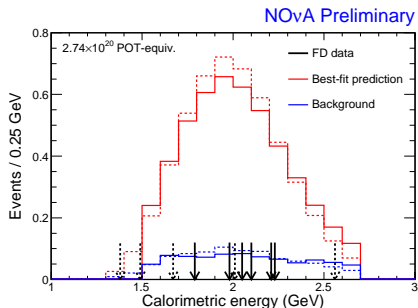
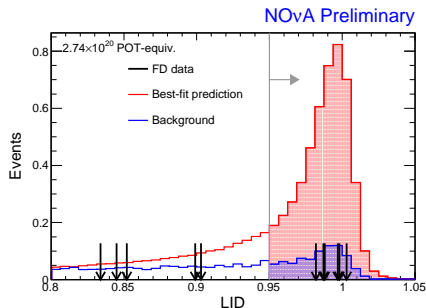
UTC Mon Mar 23, 2015

11:43:54.311669120



Selected events

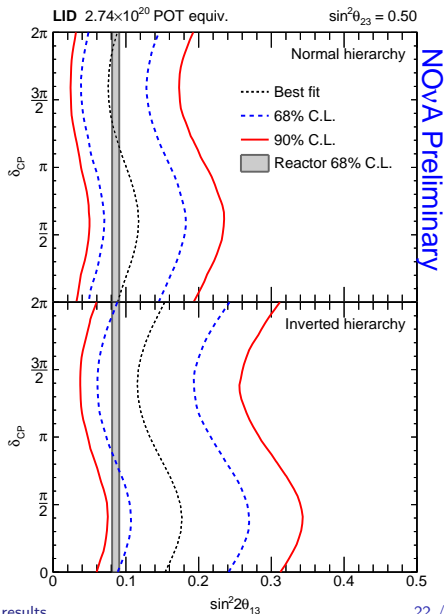
- ▶ LID selects 6 events and LEM selects 11
- ▶ Significance 3.3σ (LID) or 5.5σ (LEM)
- ▶ All 6 LID events are also selected by LEM
- ▶ p-value for $11=6/5/0$ is 9.2%
- ▶ NB, low-end energy cut trained differently for LID and LEM



ν_e appearance results

LID (N=6)

- ▶ Compatible with reactor constraints
- ▶ Slightly better for NH



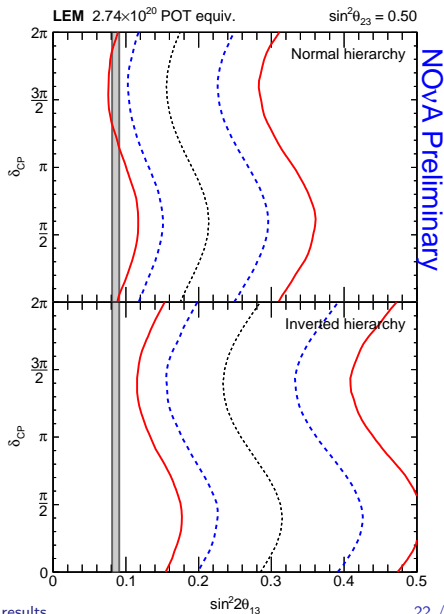
ν_e appearance results

LID (N=6)

- ▶ Compatible with reactor constraints
- ▶ Slightly better for NH

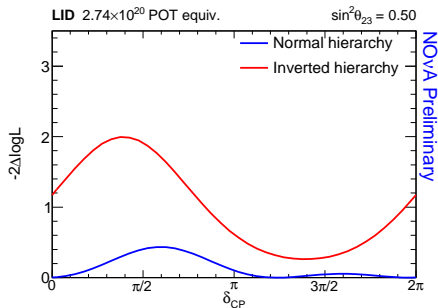
LEM (N=11)

- ▶ Curves shifted right $\sim 2\times$
- ▶ Increases tension



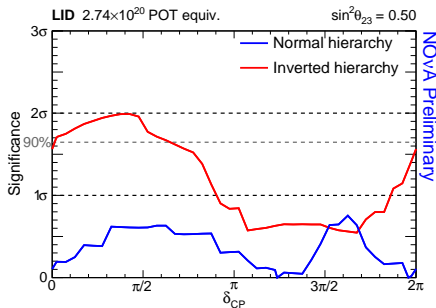
ν_e appearance results

- ▶ Include reactor θ_{13} measurement
- ▶ Plot compatibility as a function of hierarchy and δ_{CP}



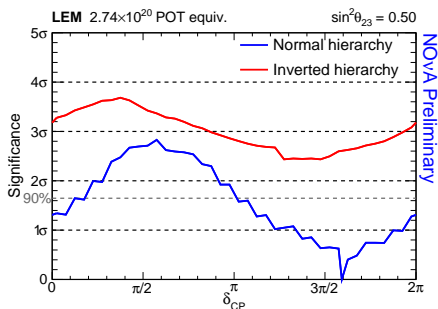
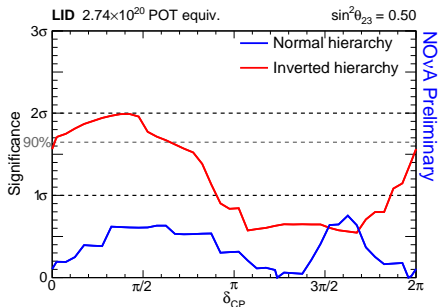
ν_e appearance results

- ▶ Include reactor θ_{13} measurement
- ▶ Plot compatibility as a function of hierarchy and δ_{CP}
- ▶ Deviations from gaussian statistics, discontinuous shape due to discrete set of possible outcomes
- ▶ Disfavour IH for $0 < \delta_{CP} < 0.6\pi$ at 90% with primary selector



ν_e appearance results

- ▶ Include reactor θ_{13} measurement
- ▶ Plot compatibility as a function of hierarchy and δ_{CP}
- ▶ Deviations from gaussian statistics, discontinuous shape due to discrete set of possible outcomes
- ▶ Disfavour IH for $0 < \delta_{CP} < 0.6\pi$ at 90% with primary selector
- ▶ Both selectors prefer NH, both prefer δ_{CP} near $3\pi/2$
- ▶ Disfavour IH, $\delta_{CP} = \pi/2$ at 1.6σ (3.2σ) using LID (LEM) for all $0.4 < \sin^2 \theta_{23} < 0.6$



Conclusion

- ▶ NO ν A has observed muon neutrino disappearance and electron neutrino appearance
- ▶ 6.5% measurement of atmospheric mass splitting, θ_{23} consistent with maximal mixing
- ▶ ν_e appearance signal at 3.3σ
- ▶ Consistent with hints for $\pi < \delta_{CP} < 2\pi$ and NH
- ▶ Only $\sim 8\%$ of nominal exposure, much more to come
- ▶ Stay tuned for next summer!

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www-nova.fnal.gov

Thank you!

Questions?