Chapter 6 : cosmic ray physics, multiplicities, correlations and spectra

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LHC working group on Forward Physics and Diffraction April the 24th 2015

Chapter 6

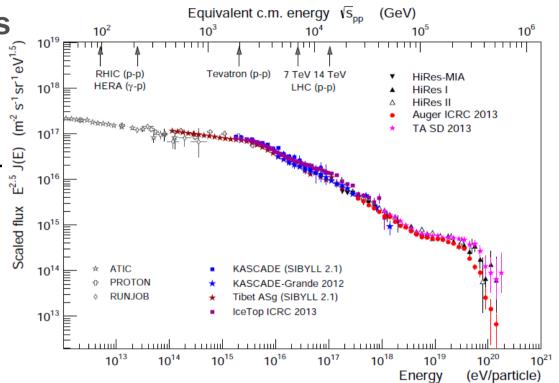
- Measurements of particular interest to improve hadronic models used for air shower simulations
 no real direct test of cosmic ray property
 - but fundamental to reduce uncertainty in air shower measurements (mass composition analysis)
- Min bias type of analysis
 - high cross section processes (~mb)
 - need low luminosity and low pile-up (each event is relevant)
- Contributions from all experiments
 - ➡ ATLAS, LHCf, CMS, TOTEM, LHCb and ALICE

Sections

- 1.Introduction
- 2.LHC and Air Showers
- 3.Energy Flow
- 4.Particle Multiplicities
- 5.Spectra
- 6.Beam

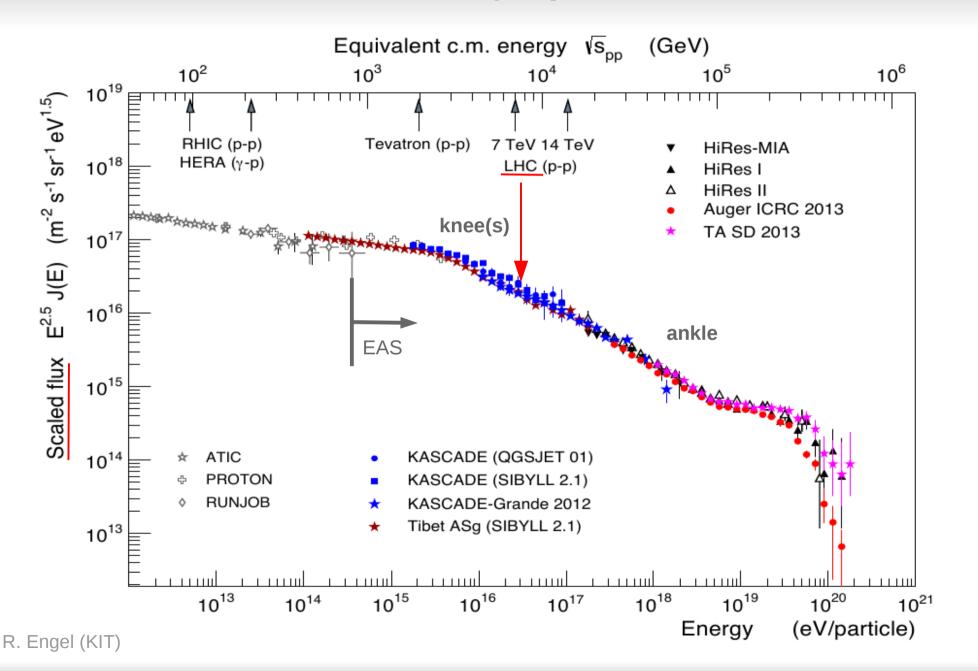
Section 6.1: Introduction

motivations from CR physics
 spectral feature
 mass composition
 test of particle physics at ultrahigh energy
 modified hadronic inter.
 lorentz invariance or extradimensions



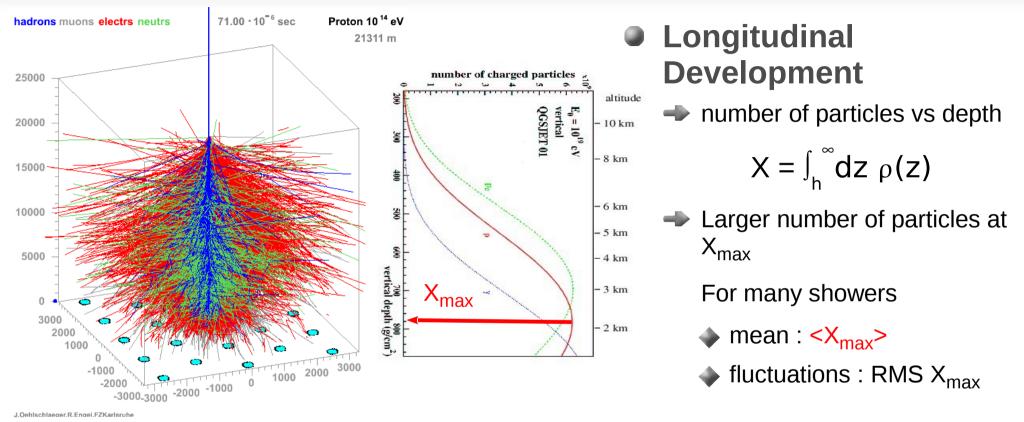
- Hadronic interaction (models)
 - main source of uncertainty in CR (air shower (EAS)) analysis
 - LHC data: highest energy to test and tune models before extrapolation
 - forward measurements most important for EAS development

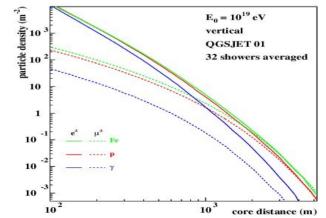
Cosmic Ray Spectrum



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Extensive Air Shower Observables





Lateral distribution function (LDF)

- particle density at ground vs distance to the impact point (core)
- can be muons or electrons/gammas or a mixture of all.

Introduction

Energy Flow

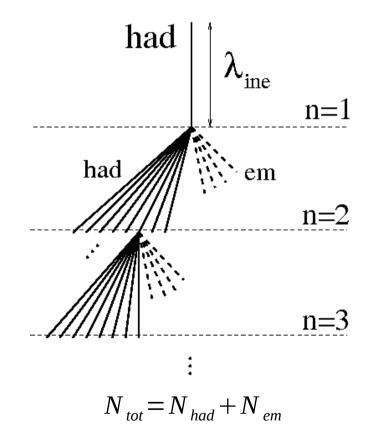
Particle Multiplicitie

Spectra Bea

Simplified Shower Development

Using generalized Heitler model and superposition model :

 $X_{max} \sim \lambda_e \ln \left((1-k) \cdot E_0 / (2 \cdot N_{tot} \cdot A) \right) + \lambda_{ine}$



J. Matthews, Astropart.Phys. 22 (2005) 387-397 Model independent parameters :

- \blacksquare E₀ = primary energy
- A = primary mass
- λ_{a} = electromagnetic mean free path
- Model dependent parameters :
 - k = elasticity
 - N_{tot} = total multiplicity
 - λ_{ine} = hadronic mean free path (cross section)

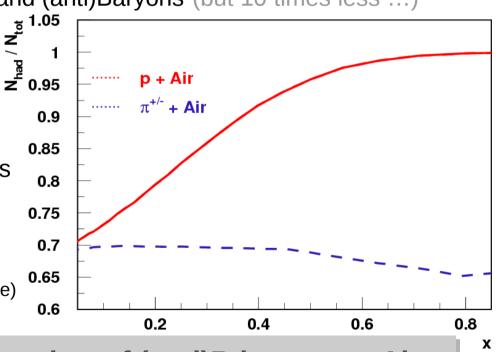
Muon Number

From Heitler

$$N_{\mu} = \left(\frac{E_0}{E_{dec}}\right)^{\alpha}, \quad \alpha = \frac{\ln N_{\pi^{ch}}}{\ln \left(N_{\pi^{ch}} + N_{\pi^0}\right)}$$

 \rightarrow after n generations

- ➡ In real shower, not only pions : Kaons and (anti)Baryons (but 10 times less ...)
- \bullet Baryons do not produce leading π^0
- With leading baryon, energy kept in hadronic channel = muon production
- Cumulative effect for low energy muons
- High energy muons
 - important effect of first interactions and baryon spectrum (LHC energy range)



Muon number depends on the number of (anti)B in p- or π -Air interactions at all energies

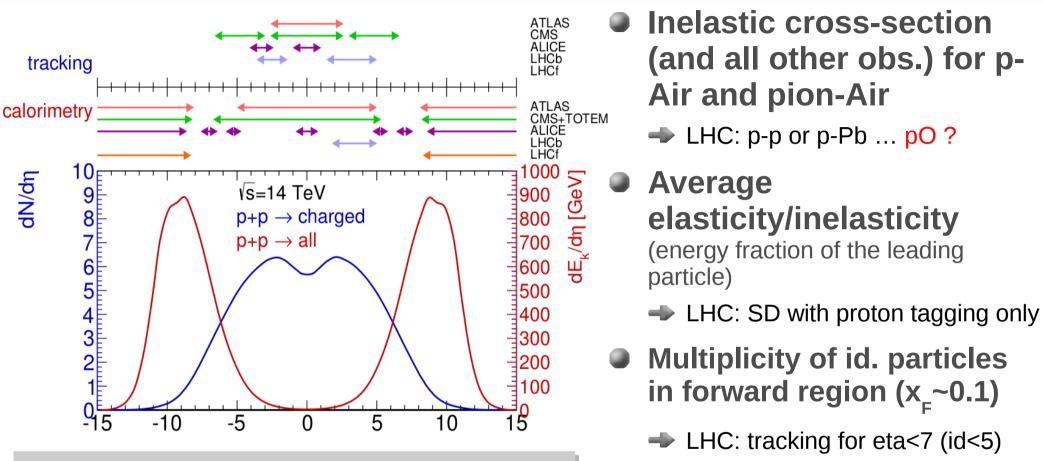
More fast (anti)baryons = more muons

T. Pierog et al., Phys. Rev. Lett. 101 (2008) 171101

T. Pierog, KIT - 8/32

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Ideal Measurements for CR

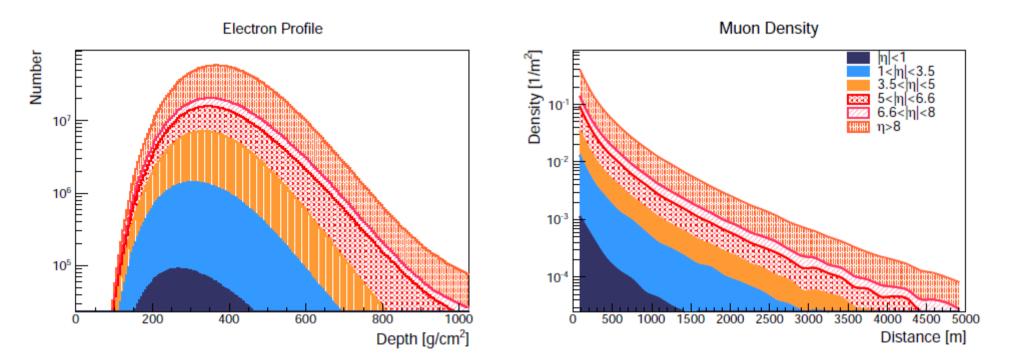


More direct measurement of particles important for air shower development not really possible at LHC ! (excluded by kin. and techn. limits)

- EM/Had Forward Energy flow (x_c>0.1)
 - ➡ LHC: ZDCs for neutral particles

add tracking in ZDC ?

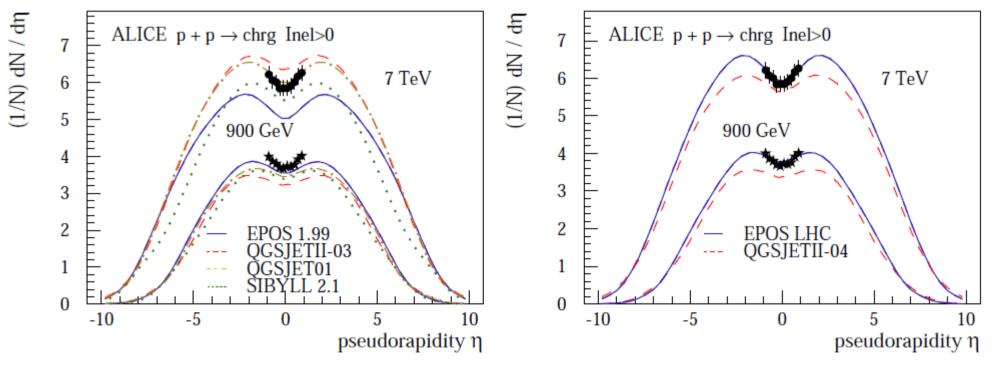
- 2.1.LHC data and hadronic interaction models
 - 2 type of measurements
 - direct connection with air shower development
 necessary to fix physics of the models



comparison old-new models for pseudorapidity

➡ first discussion on experiment (proton tagging, direct measurement, etc ...)

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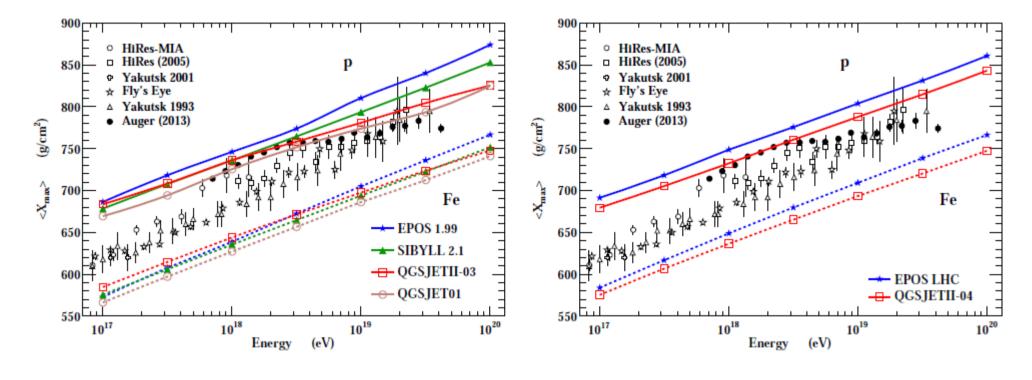
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Chapter 6.2: LHC and Air Showers

2.2.Hadronic interaction models and air showers

X_{max} : difference between models reduced by a factor of 2

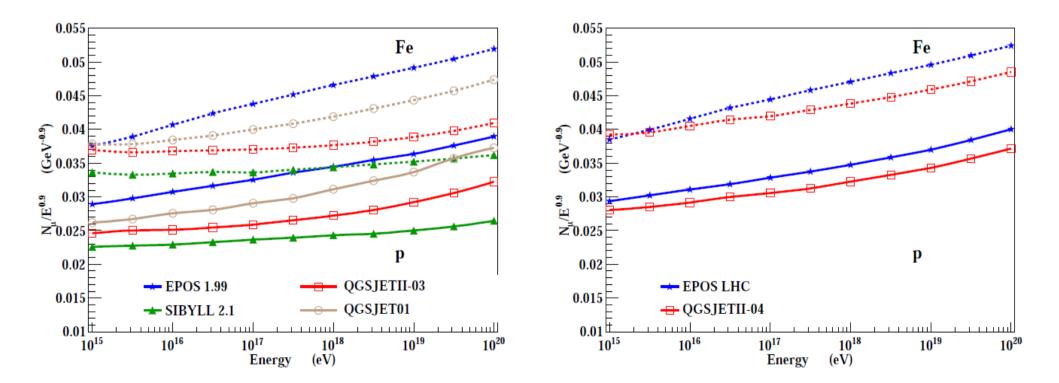


→ N_{mu}

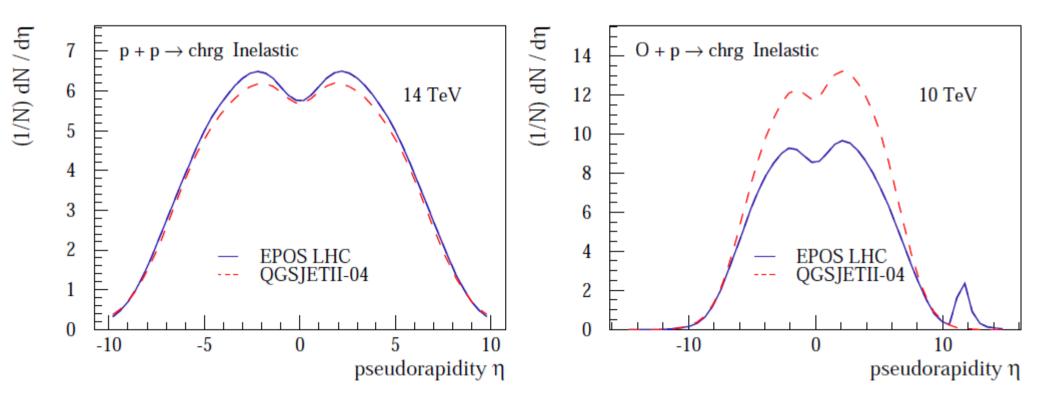
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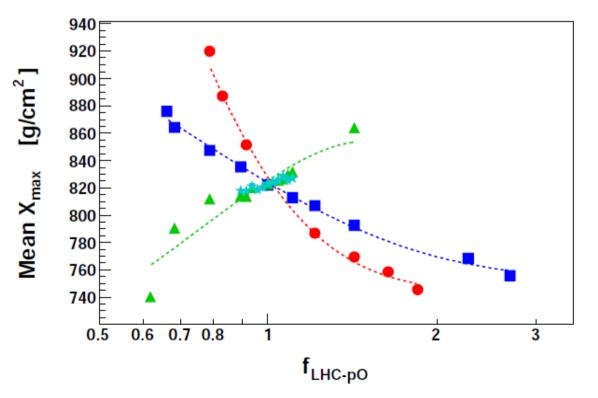
 \rightarrow N_{mu} : difference between models reduced by a factor of 3



- 2.3.Need for measuring p-O interactions
 - comparison p-p and p-O
 - p-Pb not good enough because of difficulty of selecting peripheral collisions and having proper fluctuations (and no cross-section measurement)



- 2.3.Need for measuring p-O interactions
 - comparison p-p and p-O
 - p-Pb not good enough because of difficulty of selecting peripheral collisions and having proper fluctuations (and no cross-section measurement)
 - effect of extrapolation
 - 10% difference in cross-section ~ proton/helium difference in Xmax

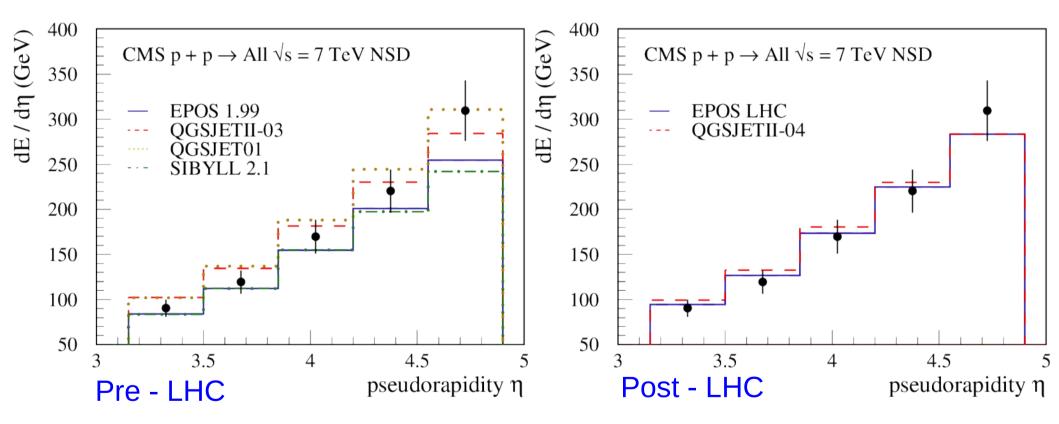


- cross section
- multiplicity
- elasticity
- charge ratio

Chapter 6.3: Energy Flow

3.1.Past measurements of energy flow

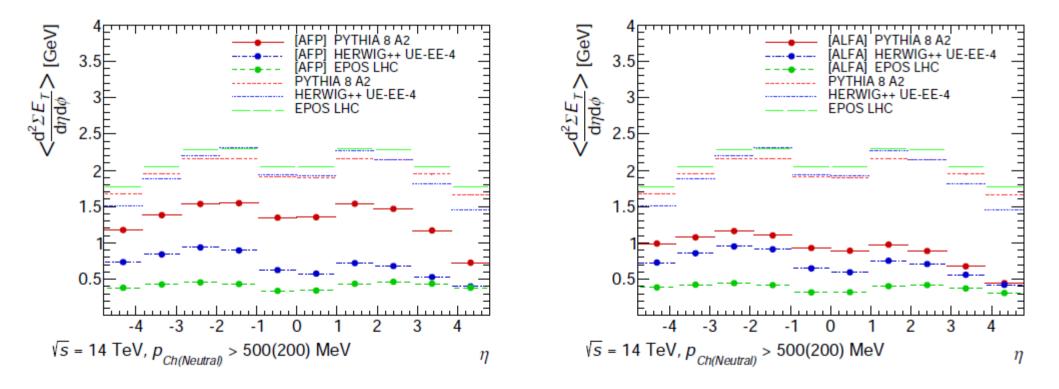
ATLAS, CMS, LHCb



- 3.2.Future measurements of energy flow
 - better model discrimination using ALPHA or AFP
 - ATLAS, CMS, LHCb common fiducial definition for energy flow

Chapter 6.3: Energy Flow

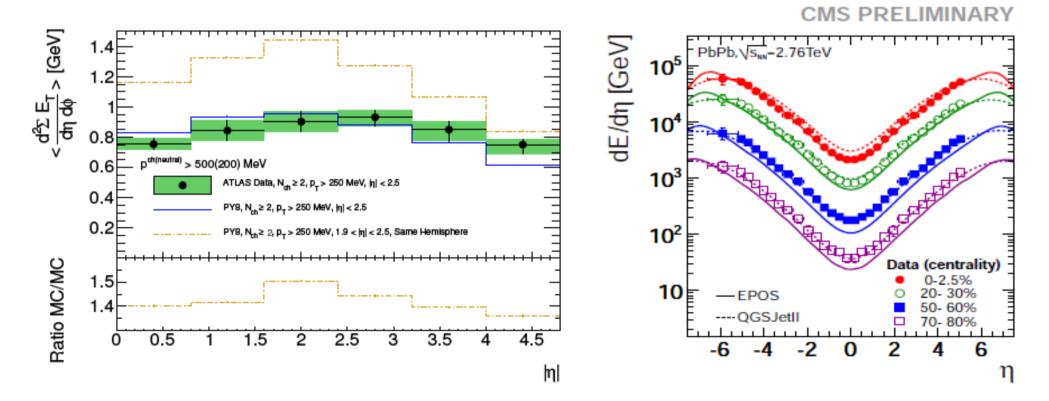
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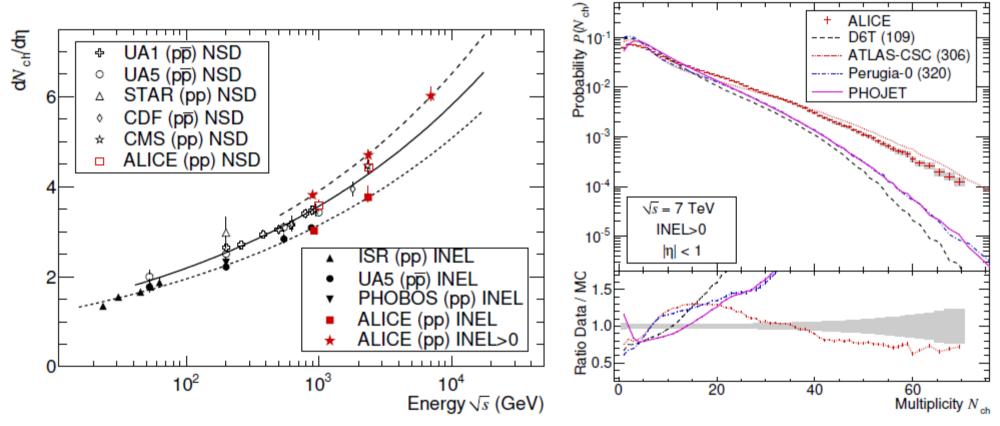
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Chapter 6.4: Particle Multiplicities

4.1.Past measurements of particle multiplicities
 ALICE, CMS/TOTEM, LHCb



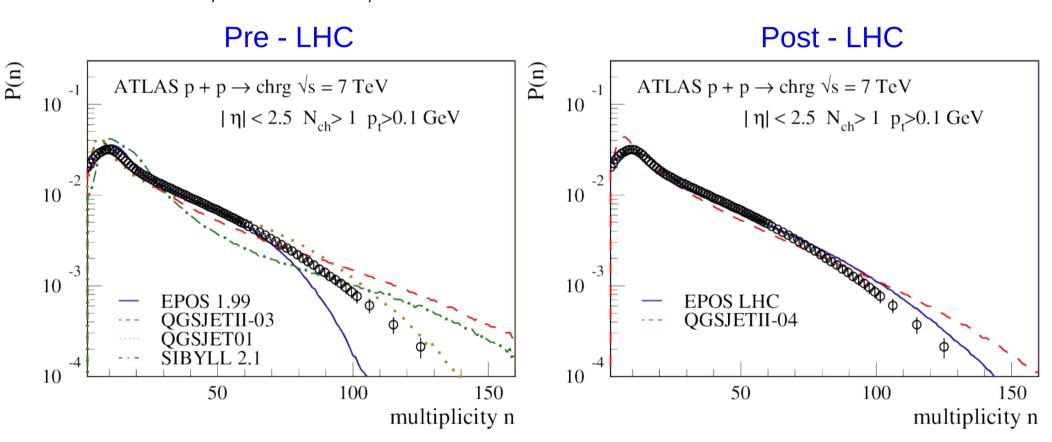
4.2.Future measurements of particle multiplicities

ATLAS with AFP/ALPHA, CMS/TOTEM

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 ALICE, CMS/TOTEM, LHCb



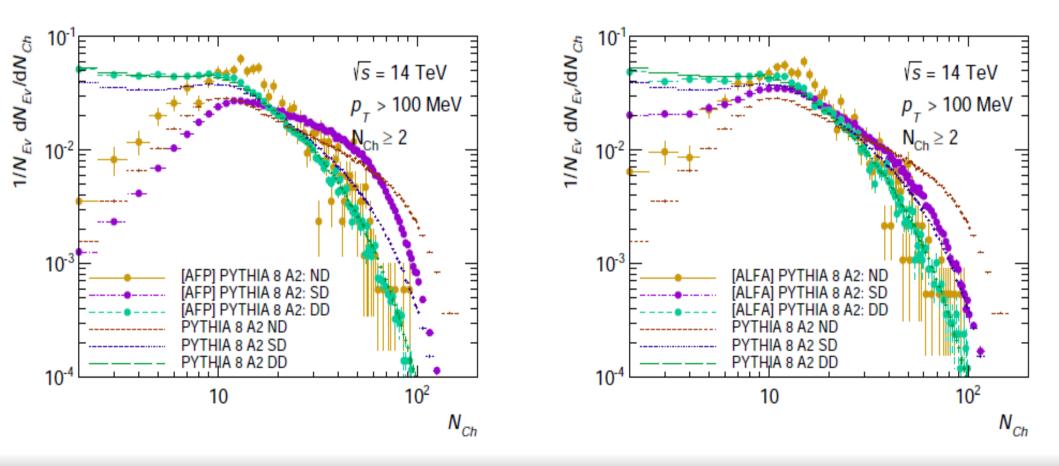
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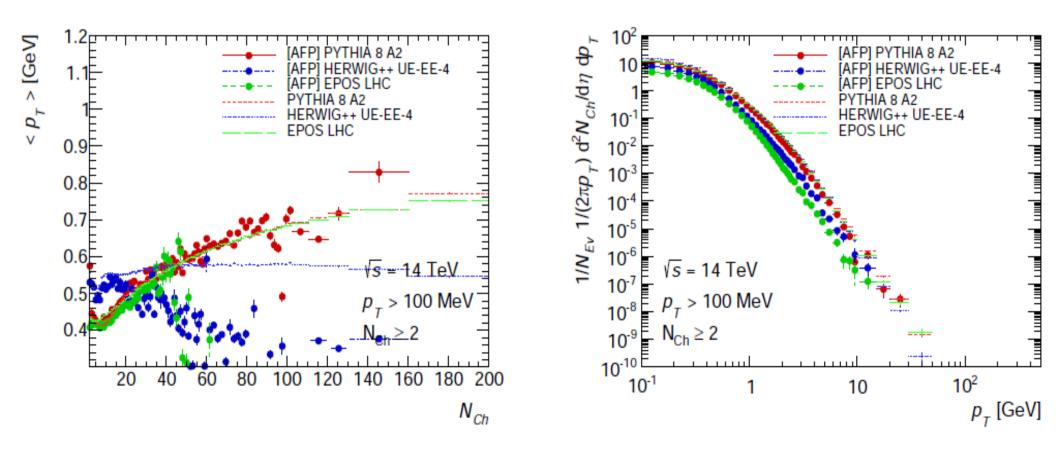


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Beam

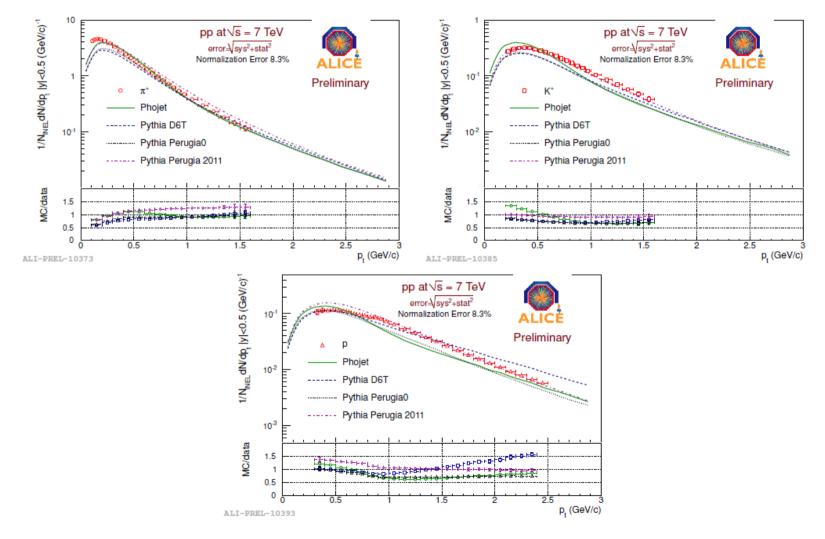
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- 4.1.Past measurements of particle multiplicities
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Chapter 6.5: Spectra

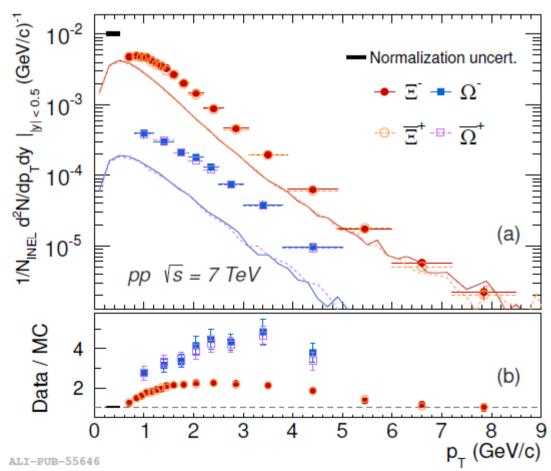
5.1.Identified charged particle spectra in p-p and p-Pb
 ALICE



Chapter 6.5: Spectra

5.1.Identified charged particle spectra in p-p and p-Pb
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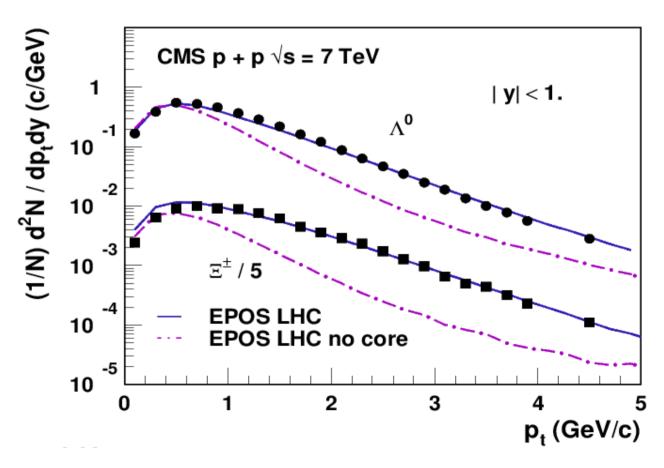
difficulties to reproduce all hadron production: important to understand for air shower development



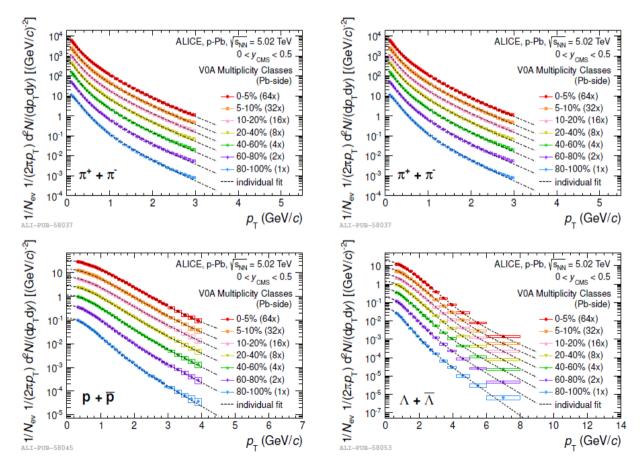
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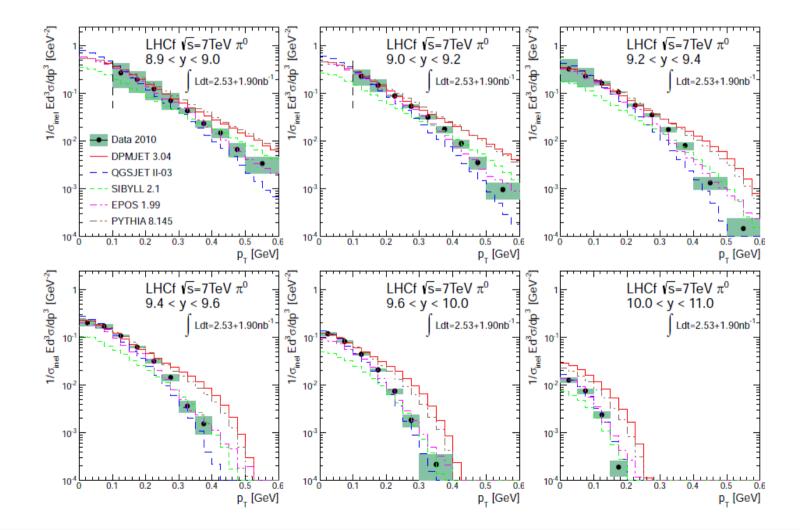
difficulties to reproduce all hadron production: important to understand for air shower development



- 5.1.Identified charged particle spectra in p-p and p-Pb
 ALICE
 - evolution as a function of event multiplicity: important to understand for air shower development



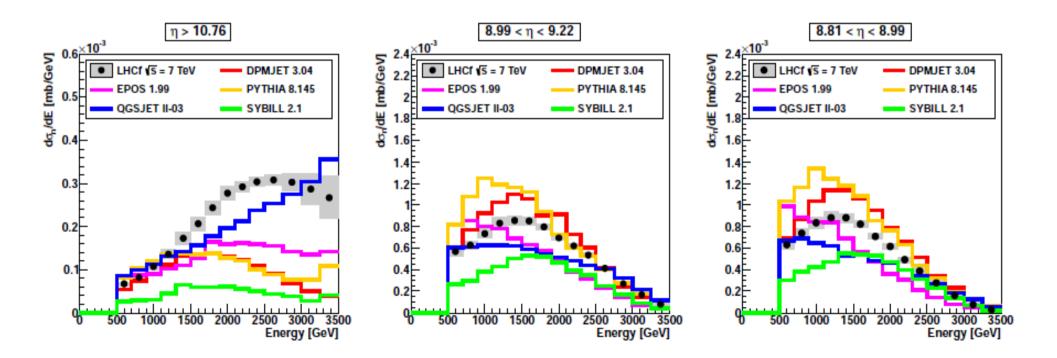
- 5.2.Neutral particle spectra
 - ALICE and LHCf



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Beam

- 5.2.Neutral particle spectra
 - ALICE and LHCf
 - direct measurements relevant for EAS development
 - both mesons and baryons
 - large model discrepancies

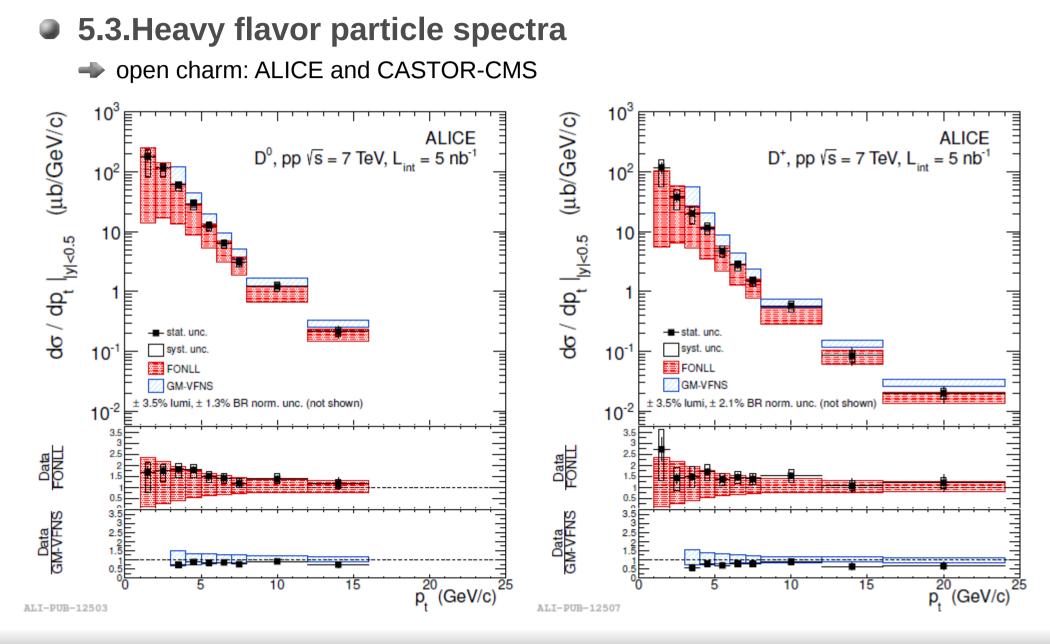


Energy Flow

Particle Multiplicities

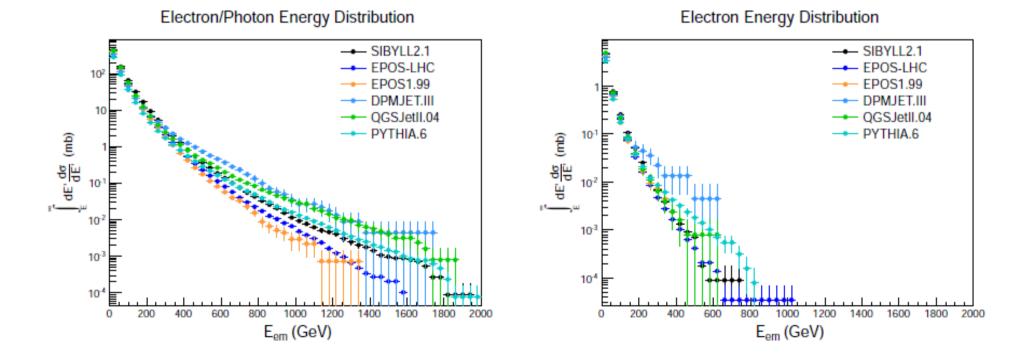
Spectra

Chapter 6.5: Spectra



IFT – April 2015

- 5.3.Heavy flavor particle spectra
 - open charm: ALICE and CASTOR-CMS
 - important for high energy neutrino production
 - background for astrophysical PeV neutrinos measured by IceCube.



Spectra

Beam

Chapter 6.6: Beam

6.1.Proton-proton collisions

Exp	σ ⁻¹ (nb ⁻¹)	Pile-up	\mathscr{L} (cm ⁻² s ⁻¹)	β* (m)	N _b	N _p /b	bunch spacing (ns)
LHCf	5-20	<1	6x10 ²⁸	19	40	10 ¹⁰	
ATLAS		<1					200
							>50
CR		<1					

6.2.Light ion Collisions

➡ fixed target (LHCb) and p-O beam

Summary

- Measurements of particular interest to improve hadronic models used for air shower simulations
 - fundamental to reduce uncertainty in air shower measurements (mass composition analysis)
- Min bias type of analysis
 - need low luminosity and low pile-up (each event is relevant)
- Contributions from all experiments
- Final chapter
 - ➡ 15 authors
 - ➔ 30 pages
 - ✤ 26 figures
 - ➡ 81 references
- Thanks to the referees for the comments and to all contributors !

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