

CERN, April 28 2015

3rd European Hyper-K meeting

Proton decay theory and predictions

Michal Malinský

Institute of Particle and Nuclear Physics Charles University in Prague

Abstract

We propose the *Hyper-Kamiokande* (Hyper-K) detector as a next generation underground water We propose the **Hyper-Kannokando** (11, posses)
Cherenkov detector. It will serve as a far detector of a long baseline neutrino oscillation experiment en-
Cherenkov detector. It will serve as a far detector of a long baseli Cherenkov detector. It will serve as a lar detector capable of observing – far beyond the sensitivity visioned for the upgraded J-PARC, and as a detector capable of observing – far beyond the sensitivity visioned for the upgraded J-PARC, and as a decessory of the Super-Kamiokande (Super-K) detector – proton decays, atmospheric neutrinos, and neutrinos from of the Super-Kamiokande (Super-K) detector – proton decays, atmosph of the Super-Kamiokande (Super-K) detector present is based on the highly successful Super-K, taking full astronomical origins. The baseline design of Hyper-K is based on the highly successful Super-K, taking full advantage of a well-proven technology.

Abstract We propose the Hyper-Kamiokande (Hyper-K) detector as a next generation underground water become neutrino oscillation experiment en-The primary objectives of LBNE, in priority order are the following experiments: tivity from 1. precision measurements of, the parameters that govern $\nu_{\mu} \rightarrow \nu_{e}$ oscillations; this in-
cludes precision measurement of the third mixing angle ng full cludes precision measurement of the third mixing angle, measurement of the CP violating phase δ_{CP} , and determination of the mass and interval in the contraction of the mass and interval interval in the mass and interv lating phase δ_{CP} , and determination of the mass ordering (the sign of Δm_{32}^2). 2. precision measurements of θ_{23} and $|\Delta m_{32}^2|$ in the ν_μ -disappearance channel. 3. search for proton decay, yielding significant improvement in the current limits on the partial lifetime of the proton (τ / BR) in one or more. partial lifetime of the proton (τ/BR) in one or more important candidate decay modes,
e.g. $p \to e^+ \pi^0$ or $p \to K^+ \nu$. e.g. $p \rightarrow e^+ \pi^0$ or $p \rightarrow K^+ \nu$. 4. detection and measurement of the neutrino flux from a core-collapse supernova within
our galaxy, should one occur during the lifetime of LDMR our galaxy, should one occur during the lifetime of LBNE.

Proton decay from the SM perspective

The SM lagrangian conserves B and L

 $\mathcal{L}_{SM}=-\tfrac{1}{2}\partial_{\nu}g_{\mu}^{a}\partial_{\nu}g_{\mu}^{a}-g_{s}f^{abc}\partial_{\mu}g_{\nu}^{a}g_{\mu}^{b}g_{\nu}^{c}-\tfrac{1}{4}g_{s}^{2}f^{abc}f^{ade}g_{\mu}^{b}g_{\nu}^{c}g_{\mu}^{d}g_{\nu}^{e}-\partial_{\nu}W_{\mu}^{+}\partial_{\nu}W_{\mu}^{-}$ $M^2W^+_\mu W^-_\mu - \frac{1}{2}\partial_\nu Z^0_\mu \partial_\nu Z^0_\mu - \frac{1}{2c^2}M^2 Z^0_\mu Z^0_\mu - \frac{1}{2}\partial_\mu A_\nu \partial_\mu A_\nu - igc_w (\partial_\nu Z^0_\mu (W^+_\mu W^-_\nu - W^+_\nu W^-_\mu) \label{eq:Z0} \begin{array}{c} Z^0_\nu(W^+_\mu \partial_\nu \tilde{W}^-_\mu - W^-_\mu \partial_\nu W^+_\mu) + Z^0_\mu (W^+_\nu \partial_\nu W^-_\mu - W^-_\nu \partial_\nu W^+_\mu)) - ig s_w (\partial_\nu A_\mu (W^+_\mu W^-_\nu - W^-_\mu \partial_\nu W^+_\mu)) - i g s_w (\partial_\nu A_\mu (W^+_\mu W^-_\nu - W^-_\mu \partial_\nu W^+_\mu)) - \\ W^+_\nu W^-_\mu) - A_\nu (W^+_\mu \partial_\nu W^-_\mu - W^-_\mu \partial_\nu W^+_\mu) + A_\mu (W^+_\nu \partial_\nu$ $\begin{array}{c}\frac{1}{2}g^2W^+_\mu W^-_\nu W^+_r W^-_\nu+\frac{1}{2}g^2W^+_\mu W^-_\nu W^+_\mu W^-_\nu+g^2c_w^2(Z^0_\mu W^+_\mu Z^0_\nu W^-_\nu-Z^0_\mu Z^0_\mu W^+_\nu W^-_\nu)+\\ g^2s_w^2(A_\mu W^+_\mu A_\nu W^-_\nu-A_\mu A_\mu W^+_\nu W^-_\nu)+g^2s_wc_w(A_\mu Z^0_\nu (W^+_\mu W^-_\nu-W^+_\nu W^-_\mu)-\end{array}$ $2A_{\mu}Z_{\mu}^{0}W_{\nu}^{+}W_{\nu}^{-}\right) - \frac{1}{2}\partial_{\mu}H\partial_{\mu}H - 2M^{2}\alpha_{h}H^{2} - \partial_{\mu}\phi^{+}\partial_{\mu}\phi^{-} - \frac{1}{2}\partial_{\mu}\phi^{0}\partial_{\mu}\phi^{0} \beta_h\left(\frac{2M^2}{g^2}+\frac{2M}{g}H+\frac{1}{2}(H^2+\phi^0\phi^0+2\phi^+\phi^-)\right)+\frac{2M^4}{g^2}\alpha_h-g\alpha_hM\left(H^3+H\phi^0\phi^0+2H\phi^+\phi^-\right) \frac{1}{8}g^2\alpha_h\left(H^4+(\phi^0)^4+4(\phi^+\phi^-)^2+4(\phi^0)^2\phi^+\phi^-+4H^2\phi^+\phi^-+2(\phi^0)^2H^2\right)-gMW_\mu^+W_\mu^-H \frac{1}{2}g_{\mu\nu}^{M}Z_{\mu}^{0}Z_{\mu}^{0}H - \frac{1}{2}ig(W_{\mu}^{+}(\phi^{0}\partial_{\mu}\phi^{-} - \phi^{-}\partial_{\mu}\phi^{0}) - W_{\mu}^{-}(\phi^{0}\partial_{\mu}\phi^{+} - \phi^{+}\partial_{\mu}\phi^{0})) +$ $\frac{1}{2}g\left(W^{\pm}_{\mu}(H\partial_{\mu}\phi^{-}-\phi^{-}\partial_{\mu}H)+W^{-}_{\mu}(H\partial_{\mu}\phi^{+}-\phi^{+}\partial_{\mu}H)\right)+\frac{1}{2}g\frac{1}{c_{\mu}}(Z^{0}_{\mu}(H\partial_{\mu}\phi^{0}-\phi^{0}\partial_{\mu}H)+$ $M\left(\frac{1}{c_w}Z_{\mu}^0\partial_{\mu}\phi^0 + W_{\mu}^+\partial_{\mu}\phi^- + W_{\mu}^-\partial_{\mu}\phi^+\right) - ig\frac{s_w^2}{c_w}MZ_{\mu}^0(W_{\mu}^+\phi^- - W_{\mu}^-\phi^+) + ig s_w MA_{\mu}(W_{\mu}^+\phi^- W^{-}_{\mu}\phi^{+})-ig\frac{1-2c_{w}^{2}}{2c_{w}}Z_{\mu}^{0}(\phi^{+}\partial_{\mu}\phi^{-}-\phi^{-}\partial_{\mu}\phi^{+})+ig s_{w}A_{\mu}(\phi^{+}\partial_{\mu}\phi^{-}-\phi^{-}\partial_{\mu}\phi^{+}) \frac{1}{4}g^2W^+_\mu W^-_\mu (H^2 + (\phi^0)^2 + 2\phi^+\phi^-) - \frac{1}{8}g^2\frac{1}{c^2}Z^0_\mu Z^0_\mu (H^2 + (\phi^0)^2 + 2(2s_w^2 - 1)^2\phi^+\phi^-) \begin{array}{c} \frac{1}{2}g^2\frac{s_w^2}{c_w}Z^0_\mu\phi^0(W^+_\mu\phi^- + W^-_\mu\phi^+) - \frac{1}{2}ig^2\frac{s_w^2}{c_w}Z^0_\mu H(W^+_\mu\phi^- - W^-_\mu\phi^+) + \frac{1}{2}g^2s_wA_\mu\phi^0(W^+_\mu\phi^- + \phi^+) \end{array}$ $\begin{array}{l} \displaystyle W^-_\mu\phi^+)+\frac{1}{2}ig^2s_wA_\mu H(W^+_\mu\phi^--\tilde{W^-_\mu\phi^+)-g^2\frac{s_w}{c_w}(2c_w^2-1)Z^0_\mu A_\mu\phi^+\phi^--g^2s_w^2A_\mu A_\mu\phi^+\phi^-+\frac{1}{2}ig_s\,\lambda^a_{ij}(\bar{q}^\sigma_i\gamma^\mu q^\sigma_j)g^a_\mu-\bar{e}^\lambda(\gamma\partial+m^\lambda_e)e^\lambda-\bar{\nu}^\lambda(\gamma\partial+m^\lambda_\nu)\nu^\lambda-\bar{u}^\lambda_j(\gamma\partial+m^\lambda_u)u^\lambda_j-\bar{d}^\lambda_j(\gamma\partial$ $ig s_w A_\mu \left(-(\bar{e}^{\lambda} \gamma^\mu e^{\lambda}) + \frac{2}{3} (\bar{u}_j^{\lambda} \gamma^\mu u_j^{\lambda}) - \frac{1}{3} (\bar{d}_j^{\lambda} \gamma^\mu d_j^{\lambda}) \right) + \frac{ig}{4c_w} Z^0_\mu \{ (\bar{\nu}^{\lambda} \gamma^\mu (1 + \gamma^5) \nu^{\lambda}) + (\bar{e}^{\lambda} \gamma^\mu (4s_w^2 (1-\gamma^5)e^{\lambda}$ + $(\bar{d}_{3}^{\lambda}\gamma^{\mu}(\frac{4}{3}s_{w}^{2}-1-\gamma^5)d_{3}^{\lambda}) + (\bar{u}_{3}^{\lambda}\gamma^{\mu}(1-\frac{8}{3}s_{w}^{2}+\gamma^5)u_{3}^{\lambda})$ + $\frac{ig}{2\sqrt{2}}W^+_{\mu}\left((\bar{\nu}^{\lambda}\gamma^{\mu}(1+\gamma^5)U^{lep}{}_{\lambda\kappa}e^{\kappa})+(\bar{u}^{\lambda}_{j}\gamma^{\mu}(1+\gamma^5)C_{\lambda\kappa}d^{\kappa}_{j})\right)+$ $\frac{ig}{2\sqrt{2}}W^-_\mu\left((\bar{e}^\kappa U^{lep\dagger}_{\kappa\lambda}\gamma^\mu(1+\gamma^5)\nu^\lambda) + (\bar{d}_j^\kappa C^\dagger_{\kappa\lambda}\gamma^\mu(1+\gamma^5)u_j^\lambda)\right) +$ $\frac{ig}{2M\sqrt{2}}\phi^+\left(-m_e^{\kappa}(\bar{\nu}^{\lambda}U^{lep}\lambda\kappa(1-\gamma^5)e^{\kappa})+m_{\nu}^{\lambda}(\bar{\nu}^{\lambda}U^{lep}\lambda\kappa(1+\gamma^5)e^{\kappa}\right)+$ $\frac{ig}{2M\sqrt{2}}\phi^{-}\left(m_e^{\lambda}(\bar{e}^{\lambda}U^{lep}^{\dagger}_{\lambda\kappa}(1+\gamma^5)\nu^{\kappa})-m_{\nu}^{\kappa}(\bar{e}^{\lambda}U^{lep}^{\dagger}_{\lambda\kappa}(1-\gamma^5)\nu^{\kappa}\right)-\frac{g}{2}\frac{m_{\nu}^{\lambda}}{M}H(\bar{\nu}^{\lambda}\nu^{\lambda}) \frac{g m_e^{\lambda}}{M} H(\bar{e}^{\lambda} e^{\lambda}) + \frac{ig m_e^{\lambda}}{M} \phi^0(\bar{\nu}^{\lambda} \gamma^5 \nu^{\lambda}) - \frac{ig m_e^{\lambda}}{2} \phi^0(\bar{e}^{\lambda} \gamma^5 e^{\lambda}) - \frac{1}{4} \bar{\nu}_{\lambda} M_{\lambda \kappa}^R (1 - \gamma_5) \hat{\nu}_{\kappa} \frac{1}{4}\overline{\nu}\lambda M_{\lambda\kappa}^R(1-\gamma_5)\hat{\nu}_{\kappa}+\frac{ig}{2M\sqrt{2}}\phi^+\left(-m_d^{\kappa}(\bar{u}_j^{\lambda}C_{\lambda\kappa}(1-\gamma^5)d_j^{\kappa})+m_u^{\lambda}(\bar{u}_j^{\lambda}C_{\lambda\kappa}(1+\gamma^5)d_j^{\kappa}\right)+$ $\frac{ig}{2M\sqrt{2}}\phi^{-}\left(m_d^{\lambda}(\bar{d}_j^{\lambda}C^{\dagger}_{\lambda\kappa}(1+\gamma^5)u_j^{\kappa})-m_u^{\kappa}(\bar{d}_j^{\lambda}C^{\dagger}_{\lambda\kappa}(1-\gamma^5)u_j^{\kappa}\right)-\frac{g}{2}\frac{m_u^{\lambda}}{M}H(\bar{u}_j^{\lambda}u_j^{\lambda})-\frac{g}{2}\frac{m_d^{\lambda}}{M}H(\bar{d}_j^{\lambda}d_j^{\lambda})+$ $\frac{ig m_u^{\lambda}}{2} \phi^0(\bar{u}_{\lambda}^{\lambda} \gamma^5 u_{\lambda}^{\lambda}) - \frac{ig m_d^{\lambda}}{2} \phi^0(\bar{d}_{\lambda}^{\lambda} \gamma^5 d_{\lambda}^{\lambda})$

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The SM lagrangian conserves B and L

$$
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$$

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The SM lagrangian conserves B and L

$$
\frac{ig}{2\sqrt{2}}W^+_{\mu}\left((\bar{\nu}^{\lambda}\gamma^{\mu}(1+\gamma^5)U^{lep}\lambda\kappa^{e\kappa})+(\bar{u}^{\lambda}_{j}\gamma^{\mu}(1+\gamma^5)C_{\lambda\kappa}d^{\kappa}_{j})\right)+\frac{ig}{2\sqrt{2}}W^-_{\mu}\left((\bar{e}^{\kappa}U^{lep^{\dagger}_{\kappa\lambda}\gamma^{\mu}(1+\gamma^5)\nu^{\lambda})+(\bar{d}^{\kappa}_{j}C^{\dagger}_{\kappa\lambda}\gamma^{\mu}(1+\gamma^5)u^{\lambda}_{j})\right)+\frac{ig}{2M\sqrt{2}}\phi^+(-m^{\kappa}_{e}(\bar{\nu}^{\lambda}U^{lep}\lambda\kappa(1-\gamma^5)e^{\kappa})+m^{\lambda}_{\nu}(\bar{\nu}^{\lambda}U^{lep}\lambda\kappa(1+\gamma^5)e^{\kappa})+\frac{ig}{2M\sqrt{2}}\phi^-\left(m^{\lambda}_{e}(\bar{e}^{\lambda}U^{lep^{\dagger}_{\lambda\kappa}(1+\gamma^5)\nu^{\kappa})-m^{\kappa}_{\nu}(\bar{e}^{\lambda}U^{lep^{\dagger}_{\lambda\kappa}(1-\gamma^5)\nu^{\kappa}}\right)-\frac{g}{2}\frac{m^{\lambda}_{\nu}}{M}H(\bar{\nu}^{\lambda}\nu^{\lambda})-\frac{g}{2}\frac{m^{\lambda}_{e}}{M}H(\bar{e}^{\lambda}e^{\lambda})+\frac{ig}{2}\frac{m^{\lambda}_{\nu}}{M}\phi^0(\bar{\nu}^{\lambda}\gamma^5\nu^{\lambda})-\frac{ig}{2}\frac{m^{\lambda}_{e}}{M}\phi^0(\bar{e}^{\lambda}\gamma^5e^{\lambda})-\frac{1}{4}\bar{\nu}_{\lambda}M^R_{\lambda\kappa}(1-\gamma_5)\hat{\nu}_{\kappa}-\frac{1}{4}\bar{\nu}_{\lambda}M^R_{\lambda\kappa}(1-\gamma_5)\hat{\nu}_{\kappa}+\frac{ig}{2M\sqrt{2}}\phi^+\left(-m^{\kappa}_{d}(\bar{u}^{\lambda}_{j}C_{\lambda\kappa}(1-\gamma^5)d^{\kappa}_{j})+m^{\lambda}_{u}(\bar{u}^{\lambda}_{j}C_{\lambda\kappa}(1+\gamma^5)d^{\kappa}_{j}\right)+\frac{ig}{2M\sqrt{2}}\phi^-\left(m^{\lambda}_{d
$$

always a $\overline{\Psi}\gamma^{\mu}\Psi$ structure - B perturbatively conserved

B & L violation in the SM

Only by anomalies (at the renormalizable level)

Instantons (at zero T) cause $9q + 3l \leftrightarrow \emptyset$ with immeasurably small rates

$$
{}^3He \rightarrow e^+\mu^+\overline{\nu}_\tau
$$

$$
\mathcal{A} \sim e^{-2\pi/\alpha} \sim 10^{-\mathcal{O}(100)}
$$

Sphalerons (at high T) make the tunneling more efficient \Box leptogenesis Kuzmin, V. Rubakov, M. Shaposhnikov, PLB155, 1985 Fukugita, Yanagida, PLB174, 1986

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Renormalizability is nowadays considered a quantitative feature

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SM as an effective theory at d=5 level

There is only one d=5 effective operator in the SM!

BTW: good to have the "complete Higgs doublet" :-)

 $\Lambda \sim (10^{12}-10^{14}) \text{ GeV}$

Baryon number violation from the SM perspective

B. Grzadkowski et al., JHEP 10 (2010) 085, arXiv: 1008.4884

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Let's do the same trick that Schwinger & co. played with the Fermi theory:

Elementary vertex:

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Elementary vertex:

QED-like seed of a renormalizable theory

Elementary vertices:n $\mathbf n$ W W

p

$\mathbf{Example:} \quad (d_R^T C u_R)(Q_L^T C L_L) =$

Example:
$$
(d_R^T C u_R)(Q_L^T C L_L) =
$$

\nScalar exchange

\n
$$
(3, 1, -\frac{1}{3}) \oplus (\overline{3}, 1, +\frac{1}{3})
$$
\n
$$
\Delta
$$

First

\n
$$
\text{Example:} \quad (d_R^T C u_R)(Q_L^T C L_L) = \left[\overline{(u_R)^c} \gamma_\mu Q \right] \left[\overline{(d_R)^c} \gamma_\mu L \right]
$$
\nScalar exchange

\n
$$
(3, 1, -\frac{1}{3}) \oplus (\overline{3}, 1, +\frac{1}{3})
$$
\n
$$
\Delta
$$

 $\Gamma_p\sim \frac{mp}{M^4}<(10^{34} \text{y})^{-1}~$ Such a new physics should be above 10¹⁵ GeV !?? $\frac{m_p^5}{M^4}$ < $(10^{34} \text{y})^{-1}$

SM running gauge couplings

Running gauge couplings in the SM:

$$
\mu \frac{\mathrm{d}}{\mathrm{d}\mu} g = \beta(g, \ldots)
$$

calculable in perturbation theory

$$
\beta = \frac{g^3}{16\pi^2} \left(-\frac{11}{3} C_2(G) + \frac{2}{3} \sum_{f_W} T_2^G(R_{f_W}) + \frac{1}{3} \sum_{s_C} T_2^G(R_{s_C}) \right) + \dots
$$

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\beta = \frac{g^3}{16\pi^2} \left(-\frac{11}{3} C_2(G) + \frac{2}{3} \sum_{fw} T_2^G(R_{fw}) + \frac{1}{3} \sum_{s_C} T_2^G(R_{s_C}) \right) + \dots
$$

b
better coordinates:
$$
\alpha_i \equiv \frac{g_i^2}{4\pi} \qquad t = \frac{1}{2\pi} \log \frac{\mu}{M_Z}
$$

$$
\frac{\mathrm{d}}{\mathrm{d}t}\alpha_i^{-1} = -b_i
$$

first order linear differential equation with constant coefficients (at the leading order)

Running gauge couplings in the SM

$$
\begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix} = -\frac{11}{3} \begin{pmatrix} 0 \\ 2 \\ 3 \end{pmatrix}_{gauge} + 2 \begin{pmatrix} \frac{10}{3} \\ 2 \\ 2 \end{pmatrix}_{ferm.} + \frac{1}{3} \begin{pmatrix} \frac{1}{2} \\ \frac{1}{2} \\ 0 \end{pmatrix}_{scal.}
$$

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Running gauge couplings in the SM d=6 BNV mediators 10 20 30 40 50 60 $\alpha_i =$ g_i^2 $\overline{4\pi}$ α_2^{-1} α_3^{-1} $t=\frac{1}{2\pi}$ $rac{1}{2\pi} \log \frac{\mu}{M}$ t_G $M_G \sim 10^{16} \text{GeV}$ $\overline{1}$ $\mathbb T$ b_1 b_2 b_3 $\sum_{i=1}^{n}$ $= -\frac{11}{3}$ $\sqrt{2}$ \mathbb{R}^n 0 2 3 $\sum_{i=1}^{n}$ \mathbb{R} *gauge* $+2$ $\sqrt{2}$ $\mathbb R$ 10 3 2 2 $\sum_{i=1}^{n}$ A *f erm.* $+$ 1 3 $\sqrt{2}$ $\mathbb R$ $\overline{1}$ $\overline{2}$ $\underline{\mathbb{1}}$ 2 0 $\sum_{i=1}^{n}$ \mathbb{R} *scal.*

1 2 3 4 5

 M_Z

Running gauge couplings in the SM $+X + \Delta$ d=6 BNV mediators

$$
\begin{pmatrix} b_1 \\ b_2 \\ b_3 \end{pmatrix} = -\frac{11}{3} \begin{pmatrix} 0+ \frac{25}{3} \\ 2+3 \\ 3+2 \end{pmatrix}_{gauge} + 2 \begin{pmatrix} \frac{10}{3} \\ \frac{2}{2} \\ 2 \end{pmatrix}_{ferm.} + \frac{1}{3} \begin{pmatrix} \frac{1}{2}+ \frac{1}{3} \\ \frac{1}{2} \\ 0+ \frac{1}{2} \end{pmatrix}_{scal.}
$$

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Running gauge couplings in the SM $+X + \Delta$ d=6 BNV mediators

$$
\begin{pmatrix}\n\frac{3}{5}b_1 \\
b_2 \\
b_3\n\end{pmatrix} = -\frac{11}{3} \begin{pmatrix}\n5 \\
5 \\
5\n\end{pmatrix}_{gauge} + 2 \begin{pmatrix}\n2 \\
2 \\
2\n\end{pmatrix}_{ferm.} + \frac{1}{3} \begin{pmatrix}\n\frac{1}{2} \\
\frac{1}{2} \\
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$$

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\frac{1}{2} \\
\frac{1}{2}\n\end{pmatrix}_{scal.}
$$

Running gauge couplings in the SM $+X + \Delta$ d=6 BNV mediators

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\begin{pmatrix}\n\frac{3}{5}b_1 \\
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5 \\
5\n\end{pmatrix}_{gauge} + 2 \begin{pmatrix}\n2 \\
2 \\
2\n\end{pmatrix}_{ferm.} + \frac{1}{3} \begin{pmatrix}\n\frac{1}{2} \\
\frac{1}{2} \\
\frac{1}{2}\n\end{pmatrix}_{scal.}
$$

Grand unification of the EW & strong interactions

VOLUME 32, NUMBER 8

PHYSICAL REVIEW LETTERS

25 FEBRUARY 1974

Unity of All Elementary-Particle Forces

Howard Georgi* and S. L. Glashow Lyman Laboratory of Physics, Harvard University, Cambridge, Massachusetts 02138 (Received 10 January 1974)

Strong, electromagnetic, and weak forces are conjectured to arise from a single fundamental interaction based on the gauge group SU(5).

We present a series of hypotheses and speculations leading inescapably to the conclusion that $SU(5)$ is the gauge group of the world-that all elementary particle forces (strong, weak, and electromagnetic) are different manifestations of the same fundamental interaction involving a single coupling strength, the fine-structure constant. Our hypotheses may be wrong and our speculations idle, but the uniqueness and simplicity of our scheme are reasons enough that it be taken seriously.

of the GIM mechanism with the notion of colored quarks⁴ keeps the successes of the quark model and gives an important bonus: Lepton and hadron anomalies cancel so that the theory of weak and electromagnetic interactions is renormalizable.⁵

The next step is to include strong interactions. We assume that strong interactions are mediated by an octet of neutral vector gauge gluons associated with local color SU(3) symmetry, and that there are no fundamental strongly interacting scalar-meson fields.⁶ This insures that

Uniqueness of $SU(5)$ @ rank=4

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H.Georgi, S.Glashow, Phys.Rev.Lett. 30 (1974)

 $SU(3)_c \otimes SU(2)_L \otimes U(1)_Y$

$$
(1, 2, -\frac{1}{2}) \quad \begin{pmatrix} \nu_e \\ e \end{pmatrix} \quad \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix}
$$

$$
(1, 1, +1) \quad e^c \quad \mu^c
$$

$$
(3, 2, +\frac{1}{6})
$$

$$
(\frac{u}{d})
$$

$$
(\frac{v}{3}, 1, -\frac{2}{3})
$$

$$
(u)(\frac{u}{d})
$$

$$
u^{c}
$$

$$
(\frac{v}{3}, 1, +\frac{1}{3})
$$

$$
d^{c}
$$

$$
s^{c}
$$

H.Georgi, S.Glashow, Phys.Rev.Lett. 30 (1974)

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H.Georgi, S.Glashow, Phys.Rev.Lett. 30 (1974)

 $24 = (1, 1, 0) \oplus (8, 1, 0) \oplus (1, 3, 0) \oplus (3, 2, -\frac{5}{6}) \oplus (\overline{3}, 2, +\frac{5}{6})$

GUT proton lifetime estimates

Expected near(?) future sentsitivity improvements

can cover most of the predicted range of the predicted range of the predicted range of the leading GUT models.

Hyper-K p-decay sensitivity projection

can cover most of the predicted range of the predicted range of the predicted range of the leading GUT models.

Hyper-K p-decay sensitivity projection

Accuracy of a **factor of few** in Γ_{p} estimates needed to make a case!

can cover most of the predicted range of the predicted range of the predicted range of the leading GUT models.

Hyper-K p-decay sensitivity projection

Accuracy of a **factor of few** in Γ_p estimates needed to make a case! (At least) **NLO PRECISION REQUIRED**

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Proton lifetime estimates in GUTs

Now I'll focus solely on the BNV **theory accuracy...**

Proton lifetime estimates in GUTs

Proton lifetime estimates in GUTs

- requires a **very good** understanding of the **whole** spectrum

NB. SUSY is "schizophrenic" in this respect...

Example:

$$
\frac{g^2}{M_{1/6}^2} C_{ijk} \overline{u^c} \gamma^\mu d_i \overline{d_j^c} \gamma_\mu \nu_k \qquad C_{ijk} = (V_{d^c}^\dagger V_{d})_{ji} (V_{u^c}^\dagger V_{\nu})_{1k}
$$

- RH rotations enter here
- simple Yukawa sector desirable!

Y. Aoki, E. Shintani, A. Soni, Phys.Rev. D89 (2014) 014505 (lattice)

- finite shifts in the gauge matching, can be as large as $\; \Delta \alpha_i^{-1} \sim 1 \;$

Larsen, Wilczek, NPB 458, 249 (1996) G. Veneziano, JHEP 06 (2002) 051 Calmet, Hsu, Reeb, PRD 77, 125015 (2008) G. Dvali, Fortsch. Phys. 58 (2010) 528-536

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- finite shifts in the gauge matching, can be as large as $\; \Delta \alpha_i^{-1} \sim 1 \;$

NO POINT IN WORKING @ NLO WITHOUT TAMING THESE!

What to do about the Planck-scale effects (in matching)?

$$
\mathcal{L}\ni\frac{\kappa}{\Lambda}F^{\mu\nu}\langle\Phi\rangle F_{\mu\nu}
$$

- absent ω d=5 if, e.g., Φ is not in $(Adj. \otimes Adj.)_{sym}$

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SU(5) GUTs:

$$
(24 \otimes 24)_{sym} = 24 \oplus 75 \oplus 200
$$

not many options - the rank should not get reduced...

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$$

- absent ω d=5 if, e.g., Φ is not in $(Adj. \otimes Adj.)_{sum}$

SU(5) GUTs:

$$
(24 \otimes 24)_{sym} = 24 \oplus 75 \oplus 200
$$

not many options - the rank should not get reduced...

SO(10) GUTs:

$$
(45 \otimes 45)_{sym} = 54 \oplus 210 \oplus 770
$$

these, however, are the "usual" choices (**though not minimal**)...

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Minimal SO(10) GUT

The minimal SO(10) unification

The minimal SO(10) unification

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The minimal SO(10) unification

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Taming the Planck-scale effects in the minimal SO(10)

The leading Planck-scale effects absent in SO(10) GUTs broken by 45!

$$
\mathcal{L} \ni \frac{\kappa}{\Lambda} F^{\mu\nu} \langle 45 \rangle F_{\mu\nu} = 0
$$
The minimal SO(10) unification

SO(10) broken by 45, rank reduced by 126

Scalar potential: $V = V_{45} + V_{126} + V_{mix}$

$$
V_{45} = -\frac{\mu^2}{2} (\phi \phi)_0 + \frac{a_0}{4} (\phi \phi)_0 (\phi \phi)_0 + \frac{a_2}{4} (\phi \phi)_2 (\phi \phi)_2 ,
$$

\n
$$
V_{126} = -\frac{\nu^2}{5!} (\Sigma \Sigma^*)_0
$$

\n
$$
+ \frac{\lambda_0}{(5!)^2} (\Sigma \Sigma^*)_0 (\Sigma \Sigma^*)_0 + \frac{\lambda_2}{(4!)^2} (\Sigma \Sigma^*)_2 (\Sigma \Sigma^*)_2
$$

\n
$$
+ \frac{\lambda_4}{(3!)^2 (2!)^2} (\Sigma \Sigma^*)_4 (\Sigma \Sigma^*)_4 + \frac{\lambda'_4}{(3!)^2} (\Sigma \Sigma^*)_4 \cdot (\Sigma \Sigma^*)_4
$$

\n
$$
+ \frac{\eta_2}{(4!)^2} (\Sigma \Sigma)_2 (\Sigma \Sigma)_2 + \frac{\eta_2^*}{(4!)^2} (\Sigma^* \Sigma^*)_2 (\Sigma^* \Sigma^*)_2 ,
$$

\n
$$
V_{\text{mix}} = \frac{i\tau}{4!} (\phi)_2 (\Sigma \Sigma^*)_2 + \frac{\alpha}{2 \cdot 5!} (\phi \phi)_0 (\Sigma \Sigma^*)_0
$$

\n
$$
+ \frac{\beta_4}{4 \cdot 3!} (\phi \phi)_4 (\Sigma \Sigma^*)_4 + \frac{\beta'_4}{3!} (\phi \phi)_4 (\Sigma \Sigma^*)_4
$$

\n
$$
+ \frac{\gamma_2}{4!} (\phi \phi)_2 (\Sigma \Sigma)_2 + \frac{\gamma_2^*}{4!} (\phi \phi)_2 (\Sigma^* \Sigma^*)_2 .
$$

 $(\phi \phi)_0 (\phi \phi)_0 \equiv \phi_{ij} \phi_{ij} \phi_{kl} \phi_{kl}$ $(\phi \phi)_2 (\phi \phi)_2 \equiv \phi_{ij} \phi_{ik} \phi_{lj} \phi_{lk}$ $(\phi \phi)_0 \equiv \phi_{ij} \phi_{ij}, \ \ (\Sigma \Sigma^*)_0 \equiv \Sigma_{ijklm} \Sigma^*_{ijklm}$ $(\Sigma \Sigma^*)_0 (\Sigma \Sigma^*)_0 \equiv \Sigma_{ijklm} \Sigma_{ijklm}^* \Sigma_{nopar} \Sigma_{nonar}^*$ $(\Sigma\Sigma^*)_2(\Sigma\Sigma^*)_2 \equiv \Sigma_{ijklm}\Sigma_{ijkln}^* \Sigma_{oparm}\Sigma_{oparn}^*$ $(\Sigma \Sigma^*)_4 (\Sigma \Sigma^*)_4 \equiv \Sigma_{ijklm} \Sigma_{ijkno}^* \Sigma_{pqrlm} \Sigma_{parno}^*$ $(\Sigma \Sigma^*)_{4'} (\Sigma \Sigma^*)_{4'} \equiv \Sigma_{ijklm} \Sigma^*_{ijkno} \Sigma_{pgrln} \Sigma^*_{parmo}$ $(\Sigma \Sigma)_2 (\Sigma \Sigma)_2 \equiv \Sigma_{ijklm} \Sigma_{ijkln} \Sigma_{oparm} \Sigma_{oparn}$ $(\phi)_2(\Sigma\Sigma^*)_2 \equiv \phi_{ij}\Sigma_{klmni}\Sigma^*_{klmnj}$ $(\phi \phi)_0 (\Sigma \Sigma^*)_0 \equiv \phi_{ij} \phi_{ij} \Sigma_{klmno} \Sigma^*_{klmno}$ $(\phi \phi)_4 (\Sigma \Sigma^*)_4 \equiv \phi_{ij} \phi_{kl} \Sigma_{mnoij} \Sigma_{mnokl}^*$ $(\phi \phi)_{4'}(\Sigma \Sigma^*)_{4'} \equiv \phi_{ij} \phi_{kl} \Sigma_{mnoik} \Sigma_{mnojl}^*$ $(\phi \phi)_2 (\Sigma \Sigma)_2 \equiv \phi_{ij} \phi_{ik} \Sigma_{lmnoj} \Sigma_{lmnok}$ $(\phi \phi)_2 (\Sigma^* \Sigma^*)_2 \equiv \phi_{ij} \phi_{ik} \Sigma_{lmnoj}^* \Sigma_{lmnok}^*$

SO(10) broken by 45, rank reduced by 126

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\n
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\n
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+ \frac{\lambda_0}{(5!)^2} (\Sigma \Sigma^*)_0 (\Sigma \Sigma^*)_0 + \frac{\lambda_2}{(4!)^2} (\Sigma \Sigma^*)_2 (\Sigma \Sigma^*)_2
$$

\n
$$
+ \frac{\lambda_4}{(3!)^2 (2!)^2} (\Sigma \Sigma^*)_4 (\Sigma \Sigma^*)_4 + \frac{\lambda'_4}{(3!)^2} (\Sigma \Sigma^*)_4 \cdot (\Sigma \Sigma^*)_4
$$

\n
$$
+ \frac{\eta_2}{(4!)^2} (\Sigma \Sigma)_2 (\Sigma \Sigma)_2 + \frac{\eta_2^*}{(4!)^2} (\Sigma^* \Sigma^*)_2 (\Sigma^* \Sigma^*)_2 ,
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V_{\text{mix}} = \frac{i\tau}{4!} (\phi)_2 (\Sigma \Sigma^*)_2 + \frac{\alpha}{2 \cdot 5!} (\phi \phi)_0 (\Sigma \Sigma^*)_0
$$

\n
$$
+ \frac{\beta_4}{4 \cdot 3!} (\phi \phi)_4 (\Sigma \Sigma^*)_4 + \frac{\beta'_4}{3!} (\phi \phi)_4 (\Sigma \Sigma^*)_4
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\n
$$
+ \frac{\gamma_2}{4!} (\phi \phi)_2 (\Sigma \Sigma)_2 + \frac{\gamma_2^*}{4!} (\phi \phi)_2 (\Sigma^* \Sigma^*)_2 .
$$

 $(\phi \phi)_0 (\phi \phi)_0 \equiv \phi_{ij} \phi_{ij} \phi_{kl} \phi_{kl}$ $(\phi \phi)_2 (\phi \phi)_2 \equiv \phi_{ij} \phi_{ik} \phi_{lj} \phi_{lk}$ $(\phi \phi)_0 \equiv \phi_{ij} \phi_{ij}, \ \ (\Sigma \Sigma^*)_0 \equiv \Sigma_{ijklm} \Sigma^*_{ijklm}$ $(\Sigma \Sigma^*)_0 (\Sigma \Sigma^*)_0 \equiv \Sigma_{ijklm} \Sigma_{ijklm}^* \Sigma_{nopar} \Sigma_{nonar}^*$ $(\Sigma\Sigma^*)_2(\Sigma\Sigma^*)_2 \equiv \Sigma_{ijklm}\Sigma_{ijkln}^* \Sigma_{oparm}\Sigma_{oparn}^*$ $(\Sigma \Sigma^*)_4 (\Sigma \Sigma^*)_4 \equiv \Sigma_{ijklm} \Sigma_{ijkno}^* \Sigma_{pqrlm} \Sigma_{parno}^*$ $(\Sigma\Sigma^*)_{4'}(\Sigma\Sigma^*)_{4'} \equiv \Sigma_{ijklm}\Sigma_{ijkm}^* \Sigma_{pgrln} \Sigma_{parmo}^*$ $(\Sigma \Sigma)_2 (\Sigma \Sigma)_2 \equiv \Sigma_{ijklm} \Sigma_{ijkln} \Sigma_{oparm} \Sigma_{oparn}$ $(\phi)_2(\Sigma\Sigma^*)_2 \equiv \phi_{ij}\Sigma_{klmni}\Sigma^*_{klmnj}$ $(\phi \phi)_0 (\Sigma \Sigma^*)_0 \equiv \phi_{ij} \phi_{ij} \Sigma_{klmno} \Sigma^*_{klmno}$ $(\phi \phi)_4 (\Sigma \Sigma^*)_4 \equiv \phi_{ij} \phi_{kl} \Sigma_{mnoij} \Sigma_{mnokl}^*$ $(\phi \phi)_{4'}(\Sigma \Sigma^*)_{4'} \equiv \phi_{ij} \phi_{kl} \Sigma_{mnoik} \Sigma_{mnojl}^*$ $(\phi \phi)_2 (\Sigma \Sigma)_2 \equiv \phi_{ij} \phi_{ik} \Sigma_{lmnoj} \Sigma_{lmnok}$ $(\phi \phi)_2 (\Sigma^* \Sigma^*)_2 \equiv \phi_{ij} \phi_{ik} \Sigma_{lmnoj}^* \Sigma_{lmnok}^*$

"Ruled out" in 1980's

$$
m_{(8,1,0)}^2 = 2a_2(\omega_R - \omega_Y)(\omega_R + 2\omega_Y)
$$

$$
m_{(1,3,0)}^2 = 2a_2(\omega_Y - \omega_R)(\omega_Y + 2\omega_R)
$$

Yasuè 1981, Anastaze, Derendinger, Buccella 1983, Babu, Ma 1985

"Ruled out" in 1980's

$$
m_{(8,1,0)}^2 = 2a_2(\omega_R - \omega_Y)(\omega_R + 2\omega_Y)
$$

$$
m_{(1,3,0)}^2 = 2a_2(\omega_Y - \omega_R)(\omega_Y + 2\omega_R)
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Aaarrrggh... tachyonic spectrum unless $\frac{1}{2} < |\omega_Y/\omega_R| < 2$

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$$
m_{(8,1,0)}^2 = 2a_2(\omega_R - \omega_Y)(\omega_R + 2\omega_Y)
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$$

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Aaarrrggh... tachyonic spectrum unless $\frac{1}{2} < |\omega_Y/\omega_R| < 2$

$$
\langle 45 \rangle = \left(\begin{array}{ccc} \omega_Y & & & \\ & \omega_Y & & \\ & & \omega_R & \\ & & & \omega_R \end{array}\right) \otimes \tau_2
$$

SU(5)-like vacua only, **not far from the sick "SM running"**!

"Ruled out" in 1980's

$$
m_{(8,1,0)}^2 = 2a_2(\omega_R - \omega_Y)(\omega_R + 2\omega_Y)
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SU(5)-like vacua only, **not far from the sick "SM running"**!

Quantum salvation in 2010

$$
\Delta m_{(1,3,0)}^2 = \frac{1}{4\pi^2} \left[\tau^2 + \beta^2 (2\omega_R^2 - \omega_R \omega_Y + 2\omega_Y^2) + g^4 (16\omega_R^2 + \omega_Y \omega_R + 19\omega_Y^2) \right] + \log s,
$$

\n
$$
\Delta m_{(8,1,0)}^2 = \frac{1}{4\pi^2} \left[\tau^2 + \beta^2 (\omega_R^2 - \omega_R \omega_Y + 3\omega_Y^2) + g^4 (13\omega_R^2 + \omega_Y \omega_R + 22\omega_Y^2) \right] + \log s,
$$

Bertolini, Di Luzio, MM, PRD 81, 035015 (2010)

The minimal SO(10) unification

Quantum salvation in 2010

$$
\Delta m_{(1,3,0)}^2 = \frac{1}{4\pi^2} \left[\tau^2 + \beta^2 (2\omega_R^2 - \omega_R \omega_Y + 2\omega_Y^2) + g^4 (16\omega_R^2 + \omega_Y \omega_R + 19\omega_Y^2) \right] + \log s,
$$

\n
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$$

Bertolini, Di Luzio, MM, PRD 81, 035015 (2010)

Conclusions / outlook

It's almost impossible to calculate the proton lifetime accurately enough to make a clear case...

The long-ago cursed (but recently resurrected) SO(10) GUT broken by the adjoint scalar is the best hope.

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Thanks for your kind attention!

Backup slides

"Consistency is the last refuge of people without imagination"

Oscar Wilde

Chang, Mohapatra, Gipson, Marshak, Parida (1985) Deshpande, Keith, Pal (1993) Bertolini, Di Luzio, MM (2009)

"Consistency is the last refuge of people without imagination"

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too heavy LH neutrinos!? Simple estimates: $M_{\rm seesaw} \sim 10^{10} \,\text{GeV}$

multiple Yukawa finetuning?

Chang, Mohapatra, Gipson, Marshak, Parida (1985) Deshpande, Keith, Pal (1993) Bertolini, Di Luzio, MM (2009)

"Consistency is the last refuge of people without imagination"

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multiple Yukawa finetuning? \Rightarrow too heavy LH neutrinos!? Simple estimates: $M_{\text{seesaw}} \sim 10^{10} \text{ GeV}$

Enough to make the fine-tunning (if you like) elsewhere.

Chang, Mohapatra, Gipson, Marshak, Parida (1985) Deshpande, Keith, Pal (1993) Bertolini, Di Luzio, MM (2009)

"Consistency is the last refuge of people without imagination"

Oscar Wilde

39 / many

multiple Yukawa finetuning? too heavy LH neutrinos!? Simple estimates: $M_{\text{seesaw}} \sim 10^{10} \text{ GeV}$

Enough to make the fine-tunning (if you like) elsewhere.

Two other potentially realistic minimally fine-tuned & consistent scenarios with "light" scalars:

$$
(8, 2, +\frac{1}{2}) \qquad (6, 3, +\frac{1}{3})
$$

Bertolini, Di Luzio, MM, PRD85 095014 2012

Case I: light $(8, 2, +\frac{1}{2})$ **@ one loop** Bertolini, Di Luzio, MM, PRD 85, 095014 (2012)

Case I: light $(8, 2, +\frac{1}{2})$ **@ one loop** Bertolini, Di Luzio, MM, PRD 85, 095014 (2012)

Michal Malinsky, IPNP Prague

 2) can vary over many orders of magnitude in the lower part of the desert, and it is pushed in the desert, and it is pushe

Case I: light $(8, 2, +\frac{1}{2})$ **@ one loop** Bertolini, Di Luzio, MM, PRD 85, 095014 (2012)

The octet should be light!!!
 The octet should be light!!!

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The octet should be light!!!
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Case I: light $(8, 2, +\frac{1}{2})$ **@ LO** Bertolini, Di Luzio, MM, PRD 85, 095014 (2012)

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