

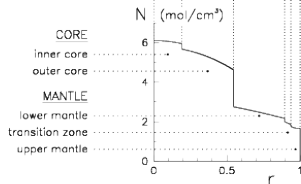
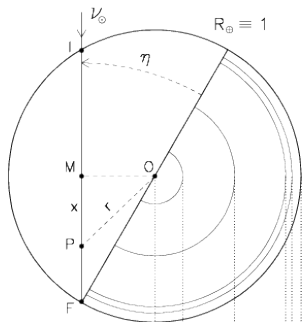
# Terrestrial Matter Effects on Solar Neutrinos

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$I = \nu$  entry point  
 $F = \nu$  endpoint (detector)  
 $M =$  trajectory midpoint  
 $P =$  generic  $\nu$  position  
 $x = MP =$  trajectory coordinate  
 $r = OP =$  radial distance  
 $\eta =$  nadir angle



$$\epsilon \equiv \frac{2VE}{\Delta m_{21}^2} \approx 0.03 \left( \frac{\rho}{3 \frac{\text{g}}{\text{cm}^3}} \right) \left( \frac{7.5 \cdot 10^{-5} \text{eV}^2}{\Delta m_{21}^2} \right) \left( \frac{E}{10 \text{ MeV}} \right) \left( \frac{Y_e}{0.5} \right)$$

$$l_\nu \approx 330 \left( \frac{7.5 \times 10^{-5} \text{eV}^2}{\Delta m_{21}^2} \right) \left( \frac{E}{10 \text{ MeV}} \right) \text{km}$$

$$\phi_{x_k \rightarrow x_n}^m \equiv \int_{x_k}^{x_n} dx \frac{\Delta m_{21}^2}{2E} \sqrt{(\cos 2\theta_{12} - \epsilon(x))^2 + \sin^2 2\theta_{12}}$$

$$\frac{P_N - P_D}{P_D} = -f(\Delta m_{21}^2, \theta_{12}) \frac{1}{2} \int_0^L dx V(x) \sin \phi_{x \rightarrow L}^m$$

where

$$f(\Delta m_{21}^2, \theta_{12}) = \frac{2 \cos 2\theta_{12}^{\odot} \sin^2 2\theta_{12}}{1 + \cos 2\theta_{12}^{\odot} \cos 2\theta_{12}} = \frac{(2P_{ee} - 1) \sin^2 2\theta_{12}}{P_{ee} \cos 2\theta_{12}}$$

$f(\Delta m_{21}^2, \theta_{12}) \simeq -2.3$  for  $\tan^2 \theta = 0.45$  ( $\theta = 34^\circ$ ) and  
 $\Delta m^2 = 7.5 \times 10^{-5} \text{ eV}^2$  ( $P_{ee} \simeq 1/3$ )

$V \rightarrow V \cdot \cos(\theta_{13})^2 \simeq V \cdot 0.98$

## Averaging over neutrino energy

$$A_e = \int dE' g(E', E) \frac{P_N - P_D}{P_D} .$$

$$A_e = -f(\Delta m_{21}^2, \theta_{12}) \frac{1}{2} \int_0^L dx V(x) F(L-x) \sin \phi_{x \rightarrow L}^m,$$

The decrease of  $F$  means that contributions from the large distances to the integral are suppressed.

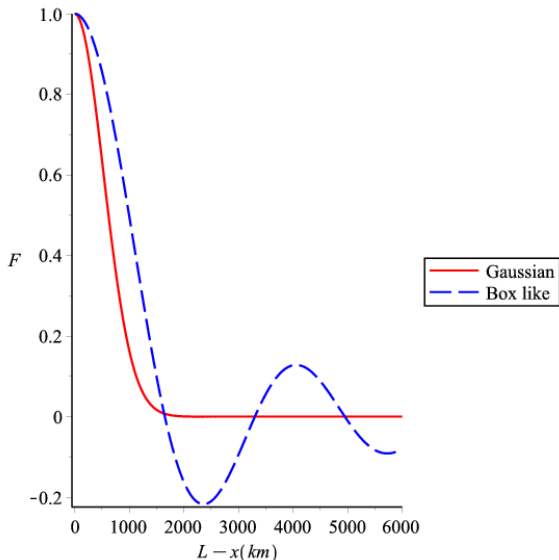
## Gaussian energy resolution function

$$g(E, E') = \frac{1}{\sigma\sqrt{2\pi}} e^{-\frac{(E-E')^2}{2\sigma^2}}, \quad F(L-x) \simeq e^{-2\left(\frac{\pi\sigma(L-x)}{E l_\nu}\right)^2}$$

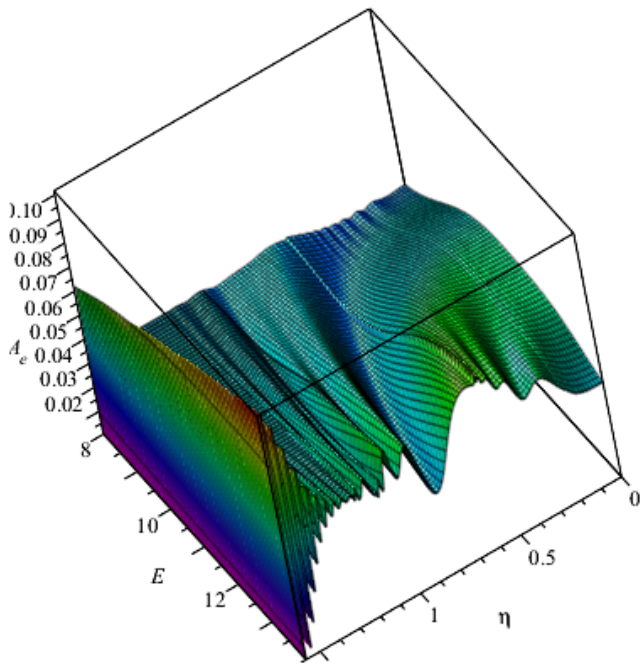
## Box like energy resolution function

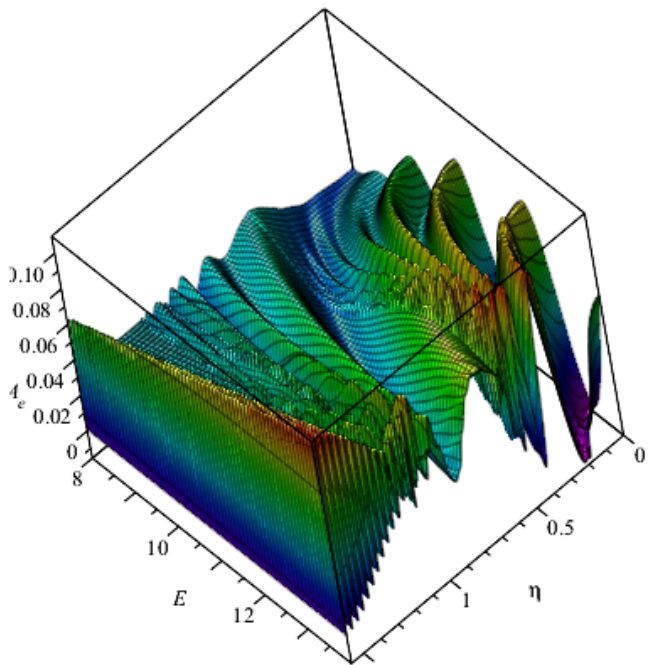
$$A_e = \frac{1}{2\sigma} \int_{E-\sigma}^{E+\sigma} dE' \frac{P_N - P_D}{P_D}$$

$$F(L-x) \simeq \frac{1}{Q(L-x)} \sin Q(L-x), \quad Q(L-x) \equiv \frac{2\pi\sigma(L-x)}{E l_\nu},$$



The attenuation factor  $F$  as function of  $(L-x)$  (distance from detector).  $E = 10$  MeV,  $\sigma = 1$  MeV, and  $\Delta m_{21}^2 = 7.5 \cdot 10^{-5}$  eV<sup>2</sup>



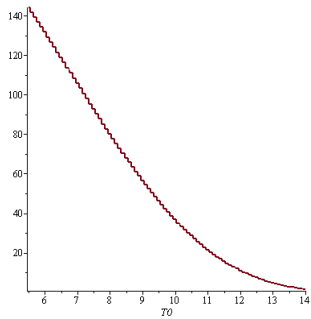
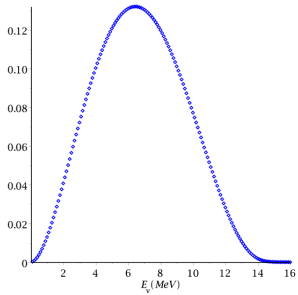


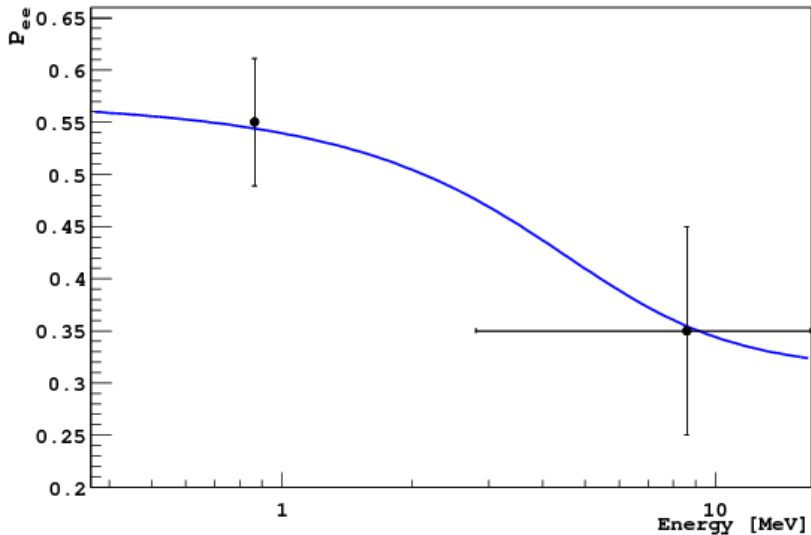
$$A_e(T, \eta) = \frac{\int dT' g(T, T') \int_{T+\frac{m_e}{2}} dE \Delta P_{ee}(E) j_B(E) \left( \frac{d\sigma_{\nu ee}}{dT'} - \frac{d\sigma_{\nu \alpha e}}{dT'} \right)}{\int dT' g(T, T') \int_{T+\frac{m_e}{2}} dE j_B(E) \left( P_{ee} \frac{d\sigma_{\nu ee}}{dT'} + (1 - P_{ee}) \frac{d\sigma_{\nu \alpha e}}{dT'} \right)}$$

$$\Delta P_{ee} = P_N - P_D = \left( \frac{1}{2} - P_{ee} \right) \frac{\sin^2 2\theta_{12}}{\cos 2\theta_{12}} \int_0^L dx V(x) \sin \phi_{x \rightarrow L}^m$$

$$\frac{d\sigma_{\nu ee}}{dT'} / \frac{d\sigma_{\nu \alpha e}}{dT'} \simeq 6$$





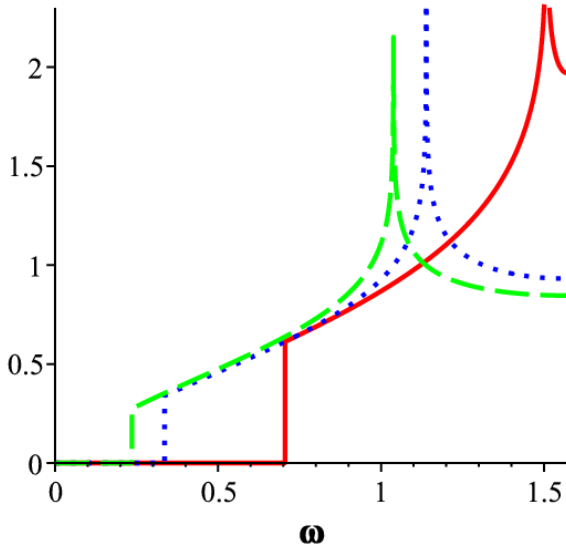


$$P_D \equiv P_{ee} = \frac{1}{2}(1 + \cos 2\bar{\theta}_{12}^{\odot} \cos 2\theta_{12}) + s_{13}^4$$

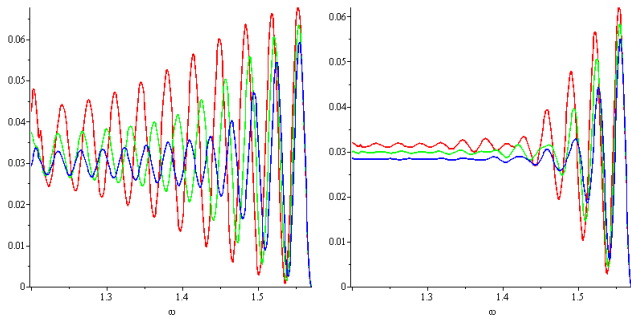
## Averaging over solar neutrino production size

From the Earth  $^8B$  neutrino production region in the Sun is seen as a disk where more neutrinos come from the center of the disk. It turns out that that distribution can be approximated as a normal one with a variance  $\delta_\eta \simeq 1.9 \times 10^{-4}$  :

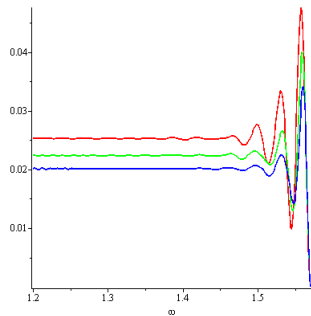
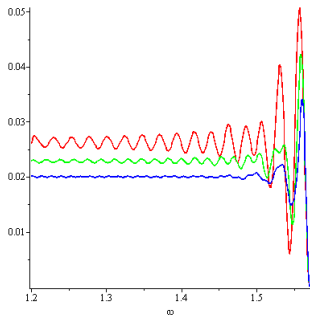
$$f(\eta', \eta) = \frac{1}{\delta_\eta \sqrt{2\pi}} e^{-\frac{(\eta - \eta')^2}{2\delta_\eta^2}} = \frac{1}{\delta_\eta \sqrt{2\pi}} e^{-\frac{(L - 2R_\oplus \cos \eta)^2}{8\delta_\eta^2 R_\oplus^2 \sin^2 \eta}}$$



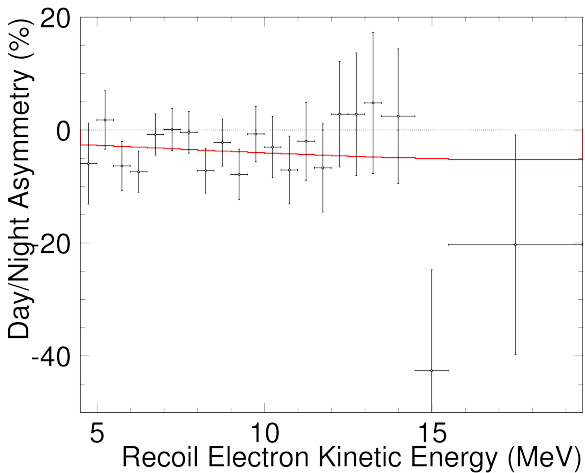
Dependance of average annual weight function on the nadir angle of neutrino trajectory ( $\omega$  is in radians). Outer core is "visible" at Kamioka site about 4.4 hours per day . [Kamioka](#), [SD](#), [Finland](#)



$$T_e = 13, 12, 11 \text{ MeV}$$



$$T_e = 9, 7, 5 \text{ MeV}$$



SK First Indication of Terrestrial Matter Effects on Solar Neutrino Oscillation

Phys. Rev. Lett. 112, 091805, 2014

the change of the solar neutrino flux due to the eccentricity of the Earth orbit ( $\pm 3\%$ ) must be taken into account.

SK recent result  $A = 3.2 \pm 1.1 \pm 0.5$  and they use  $\Delta m^2 \simeq 5 \cdot 10^{-5} \text{ eV}^2$

FULL 3D analyse of the data

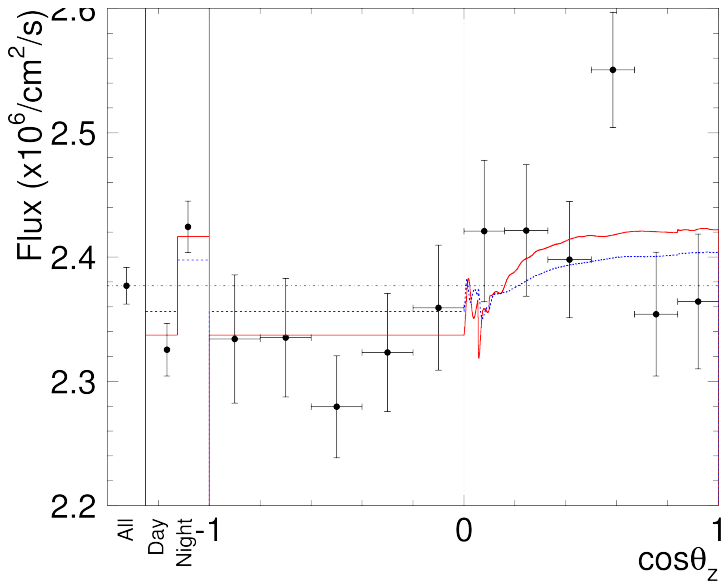
fit with Kamland  $\Delta m^2 = 7.5 \cdot 10^{-5} \text{ eV}^2$

make use "periodograms"

look to the high energy edge of the spectrum



THANK YOU



SK First Indication of Terrestrial Matter Effects on Solar Neutrino Oscillation

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