



# **QUANTUM CHAOS INSIDE BLACK HOLES**

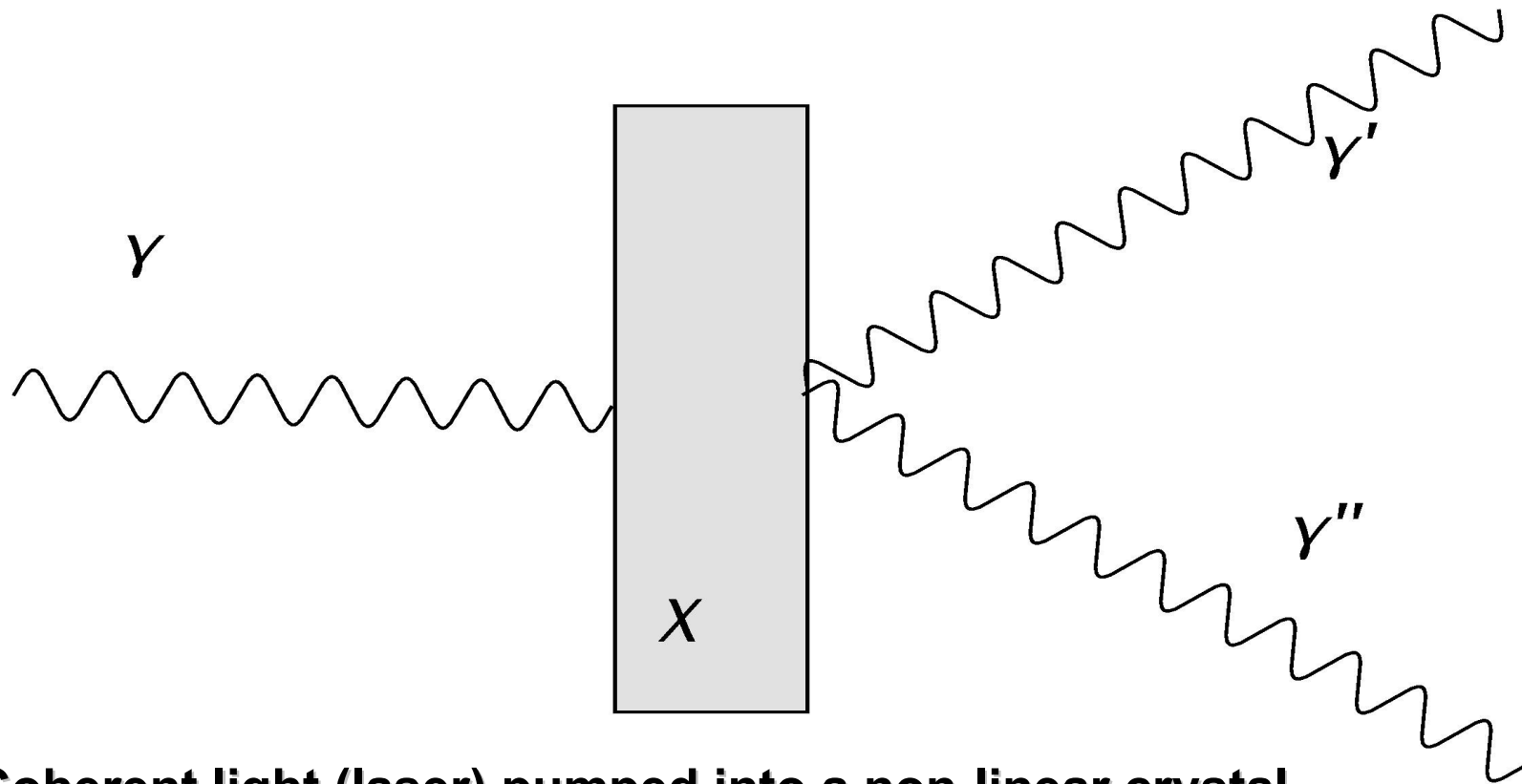
**Andrea Addazi**

**University of L'Aquila, INFN LNGS**

Let's start with a simple consideration:

**how can I create an entangled mixed state  
by an ingoing pure state?**

**Spontaneous parametric down conversion!!!**



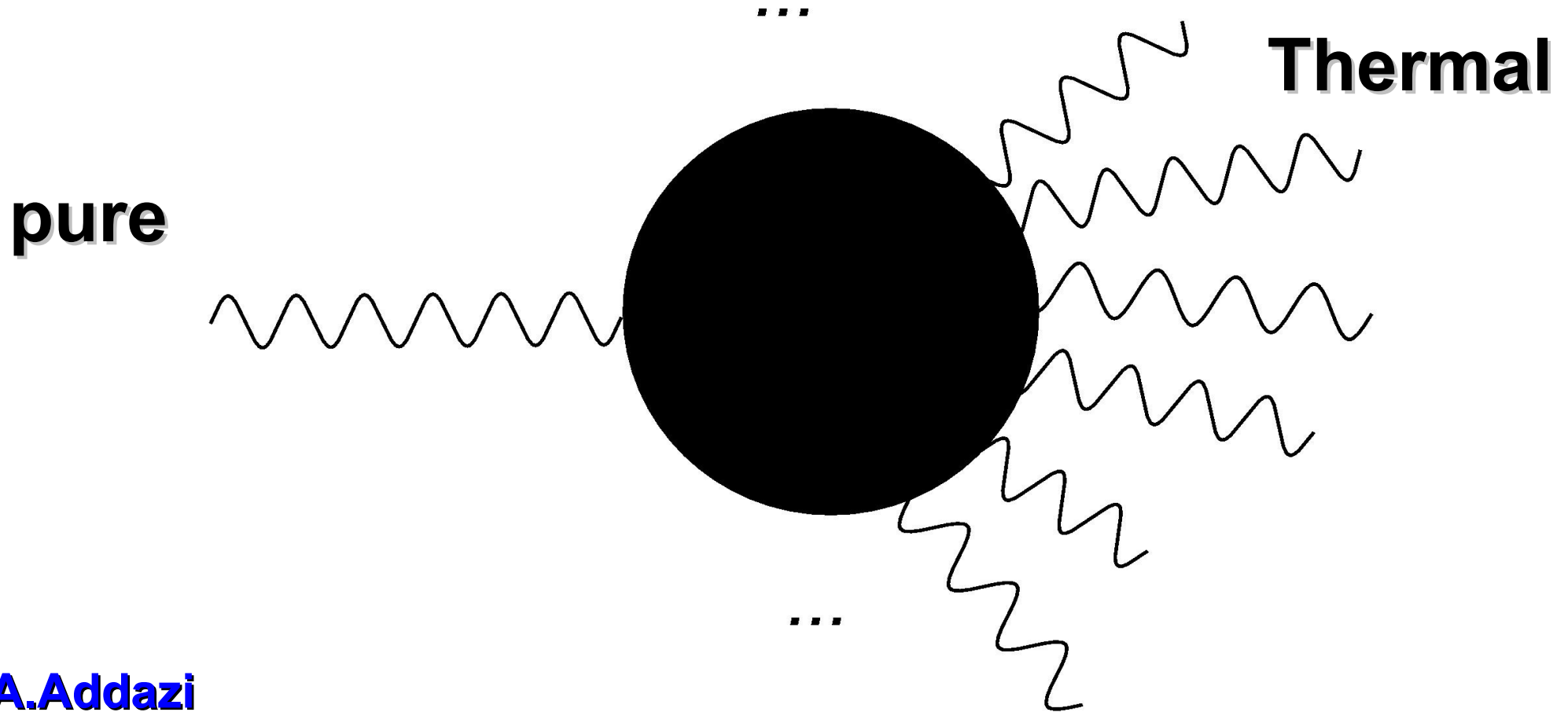
Coherent light (laser) pumped into a non-linear crystal

# Is it a violation of quantum mechanics?

No: the evolution from a pure state to a mixed one is just due to a **quantum decoherence effect**: a part of the information is effectively lost in the complicated system.

A wave function, or an S-matrix approach, for the IN-OUT transition is “practically” unsensed: a **density matrix** approach is practically preferable. The density matrix of the ingoing radiation has a non-unitary evolution, described by Liouville equation

**And now a “Crazy idea”  
(maybe so crazy to be seriously  
considered...):  
a “spontaneous conversion mechanism”  
inside Quantum BH!!!**



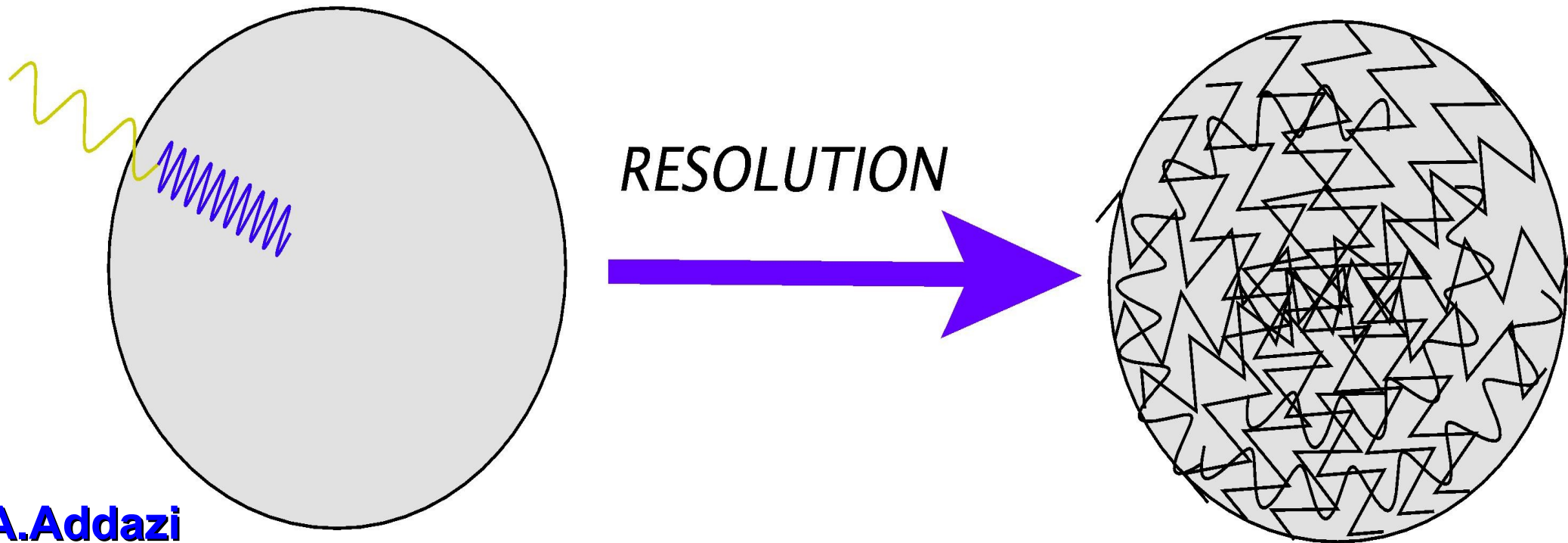
# Let us clarify what we will propose:

- Unitarity is never lost at fundamental level
- Information paradoxes are related to quantum chaotic physics inside Bhs
- We will argue how semiclassical BH can be reconstructed by a superposition of a large number of horizonless singularities
- We will argue how a part of information infalling in this system is “forever” (BH lifetime) trapped in this system.

## THIS IS A SPACE-TIME SINAI BILIARD

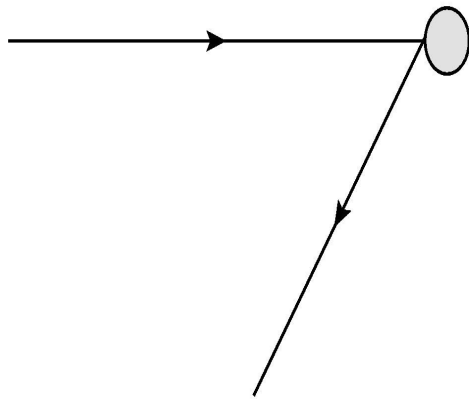
- Let me clarify that this approach cannot have the purpose to solve BH problems... We hope it could be a new sensed and useful point of view

**Basic argument: information is blueshifted and it start to resolve a non-trivial topology. An infalling wave is scattered by “asperities”. A part of this is “forever” trapped in back and forth scatterings...**

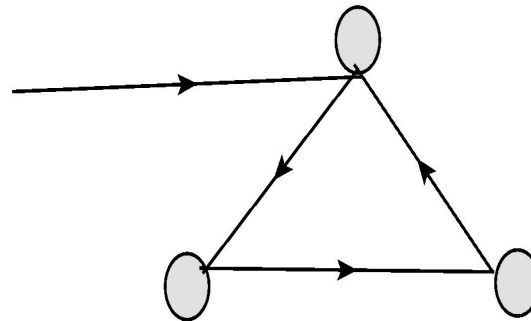


# A simple example: 2d classical chaotic scattering of a particle on N disks

a)

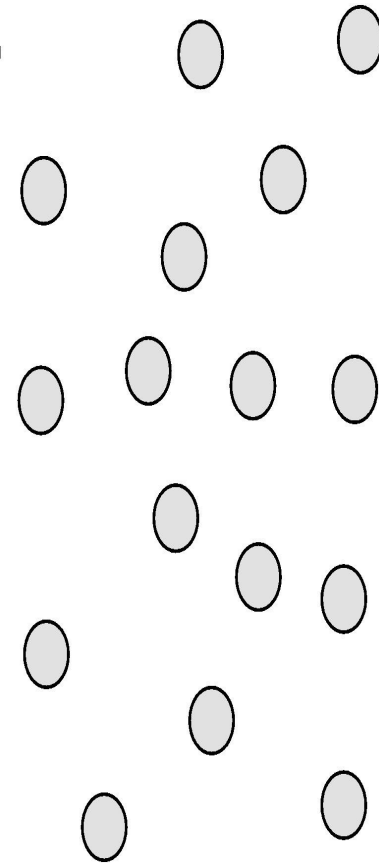


b)



...

c)



**Existence of trapped trajectories in (b). This number diverges as  $2^n$ , where  $n$  is the number of bounces**

# General Properties of Classical Chaotic scatterings (Kolmogorov, Arnold,..., Sinai,..and many others):

- Sensitivity to the initial conditions: parametrized by Lyapunov exponents
- Chaotic saddle: zone of unstable periodic trajectories trapped forever and growing exponentially with time!!! Typical example is the hyperbolic set: a set of hyperbolic unstable trajectories, exponentially growing or decreasing with time, but with a number of directions that is a topological invariant



- Semiclassical approach (n label all trajectories)

$$I = \int_0^t [\mathbf{p} \cdot d\mathbf{r} - H d\tau] \gg \hbar$$

$$K_{WKB}(\mathbf{r}, \mathbf{r}_0, t) \simeq \sum \mathcal{A}_n(\mathbf{r}, \mathbf{r}_0, t) e^{\frac{i}{\hbar} I_n}$$

$$\mathcal{A}_n(\mathbf{r}, \mathbf{r}_0, t) = \frac{1}{(2\pi i \hbar)^{\nu/2}} \sqrt{|\det[\partial_{\mathbf{r}_0} \partial_{\mathbf{r}_0} I_n[\mathbf{r}, \mathbf{r}_0, t]]|} e^{-\frac{i\pi h_n}{2}}$$

On unstable trajectories

$$|\mathcal{A}_n| \sim \exp\left(-\frac{1}{2} \sum_{\lambda_k > 0} \lambda_k t\right)$$

On stable trajectories

$$|\mathcal{A}_n| \sim |t|^{-\nu/2}$$

Time delays and Survival probabilities are determined by a continuous part and the integration over a large number of resonances (Pollicott-Ruelle spectra)

$$\mathcal{T} = \int \frac{d\Gamma_{ph}}{(2\pi\hbar)^{\nu-1}} [\delta(E - H_0 + V) - \delta(E - H_0)] + O(\hbar^{2-\nu}) + 2 \sum_p \sum_p \tau_{a=1}^{\infty} \tau_p \frac{\cos\left(a \frac{S_p}{\hbar} - \frac{\pi a}{2} \mathbf{m}_p\right)}{\sqrt{|\det(\mathcal{M}_p^a)|}} + O(\hbar)$$

$$P(t) \simeq \int \frac{d\Gamma_{ph}}{(2\pi\hbar)^f} \mathcal{I}_D e^{\mathbf{L}_{cl} t} \tilde{\rho}_0 + O(\hbar^{-\nu+1}) + \frac{1}{\pi\hbar} \int dE \sum_p \sum_a \frac{\cos\left(a \frac{S_p}{\hbar} - a \frac{\pi}{2} \mathbf{m}_p\right)}{\sqrt{|\det(\mathbf{m}_p^a - \mathbf{1})|}} \int_p \mathcal{I}_D e^{\mathbf{L}_{cl} t} \tilde{\rho}_0 dt + O(\hbar^0)$$

$$\mathbf{L}_{cl} \phi_n = \{H_{cl}, \phi_n\}_{Poisson} = \lambda_n \phi_n$$

# Semiclassical Path Integral: 1th warm up

$$Z_E = \int \mathcal{D}g \mathcal{D}\phi e^{-I[g, \phi]}$$

$$I_E = - \int_{\mathcal{M}} \sqrt{g} d^4x \left( \mathcal{L}_m + \frac{1}{16\pi} R \right) + \frac{1}{8\pi} \int_{\partial\mathcal{M}} \sqrt{h} d^3x (K - K^0)$$

$$I[\phi, g] = I[\phi_0, g_0] + I_2[\tilde{\phi}, \tilde{g}] + \text{higher orders}$$

$$I_2[\tilde{\phi}, \tilde{g}] = I_2[\tilde{\phi}] + I_2[\tilde{g}]$$

$$\log Z = -I[\phi_0, g_0] + \log \int \mathcal{D}\tilde{\phi} \mathcal{D}\tilde{g} e^{-I_2[\tilde{g}, \tilde{\phi}]}$$

# Semiclassical path Integral: 2th warm up

$$x = 4M \sqrt{1 - \frac{2M}{r}}$$

$$ds_E^2 = \left(\frac{x}{4M}\right)^2 + \left(\frac{r^2}{4M^2}\right)^2 dx^2 + r^2 d\Omega^2$$

$$Z_{ES} \simeq e^{-\frac{\beta^2}{16\pi}} \quad S = \beta(\log Z - \frac{d}{d\beta}(\log Z)) = \frac{\beta^2}{16\pi} = \frac{1}{4}A$$

# Reconstruction of a semiclassical BH by an ensemble of metric tensors

$$\log Z_{TOT} \simeq \sum_{I=1}^N \log Z_I$$

$$-I[g_0, \phi_0] + \log \int \mathcal{D}\tilde{\phi} e^{-I_2[g_0, \phi]} + \log \int \mathcal{D}\tilde{g} e^{-I[\tilde{g}]} \simeq -\sum_J I[g_0^J, \phi_0] + \sum_J \left[ \log \int \mathcal{D}\tilde{\phi} e^{-I_2[g_0^J, \phi]} + \log \int \mathcal{D}\tilde{g}^J e^{-I[\tilde{g}^J]} \right]$$

$$g_0 \simeq \left( \sum_J \sqrt{g_0^J} \right)^2$$

$$\sqrt{g_0} R(g_0) \simeq \sum_J \sqrt{g_0^J} R(g_0^J)$$

$$I_2[\tilde{g}] = \sum_J I_2[\tilde{g}^J]$$

## Carlip-Teitelboim approach:

the semiclassical wave function can be reconstructed  
by a superposition of  $N \gg 1$  wave functions  
corresponding to horizonless conic singularities

$$\frac{1}{i} \frac{\partial \Psi(S_w)}{\partial S_w} = \Theta_{WKB} \Psi(S_w)$$

$$\Theta = -\frac{\tau}{2} \frac{\mathcal{L}_n A_{D-2}}{A_{D-2}} \quad S_w = -2\pi \int_H \frac{\partial L}{\partial R_{\gamma a \delta b}} a \epsilon_{\gamma a} \epsilon_{\delta b} \quad \left\{ \Theta, \frac{1}{2\pi} S_w \right\} = 1$$

$$\Theta_{WKB} = \left[ 2\pi - \frac{1}{iC_1} (S_w - \langle S_w \rangle) + \dots \right]$$

$$\Psi(S_w) = N_1 e^{-2\pi i S_w} e^{-\frac{1}{4\Delta S_w^2} (S_w - \langle S_w \rangle)^2}$$

$$\tilde{\Psi}(\Theta) = N_2 e^{i\langle S_w \rangle \Theta} e^{-\frac{1}{2C_2} (\Theta - \langle \Theta \rangle)^2}$$

These are complicated ways just to say: these systems can emit Hawking's radiation. For  $t \ll t_{\text{BH}}$ , unitarity is apparently broken: the infalling information is converted in a mixed state during this stage. However, **<collapse|S|final evaporation>** is unitary: during the final part of the BH life, asperities are not more sustained by the matter density, they cannot more trap information and this is re-emitted as a “final information burst”, restoring unitarity in the environment. As a consequence, unitarity does not require a pure emitted radiation by this system during  $0 < t < t_{\text{BH}}$ .



# THANK YOU!!!

Andrea Addazi, KSM15 Frankfurt