

What if nature is bandlimited by a Planck-scale cutoff?

Achim Kempf

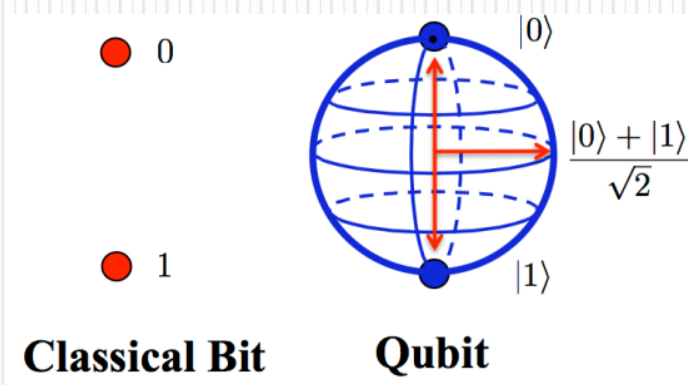
**Departments of Applied Mathematics and Physics
and Institute for Quantum Computing
University of Waterloo
Perimeter Institute for Theoretical Physics**

Karl Schwarzschild Meeting, Frankfurt, 20-24 July 2015

Some philosophy

- Concepts can lose operational meaning:
e.g., temperature, pressure, force, ...
- In quantum gravity: space, time, matter, etc ?
- Most robust: (quantum) information ?

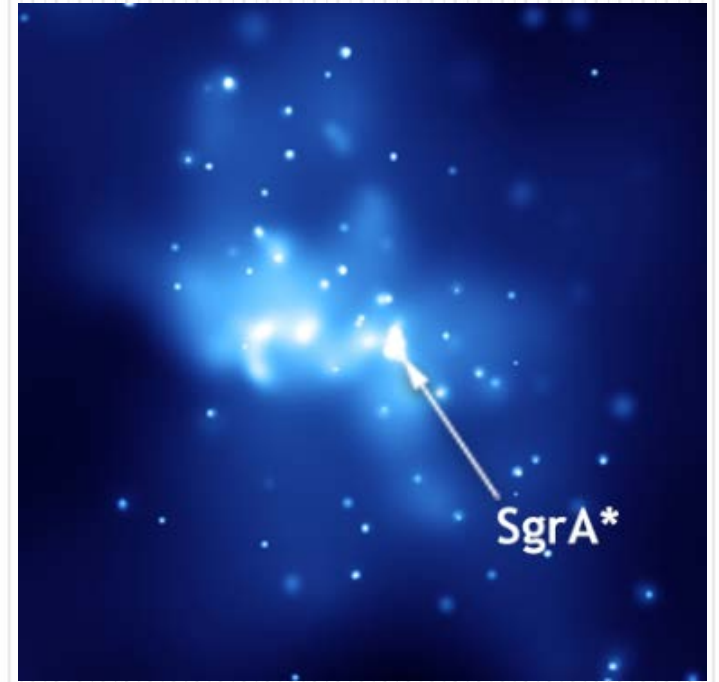
Is information a stronger concept?



Even when the meaning of the units of meters, seconds and kilograms fail, the meaning of bits and qubits may persist.

Information-theoretic foundation for physics ?

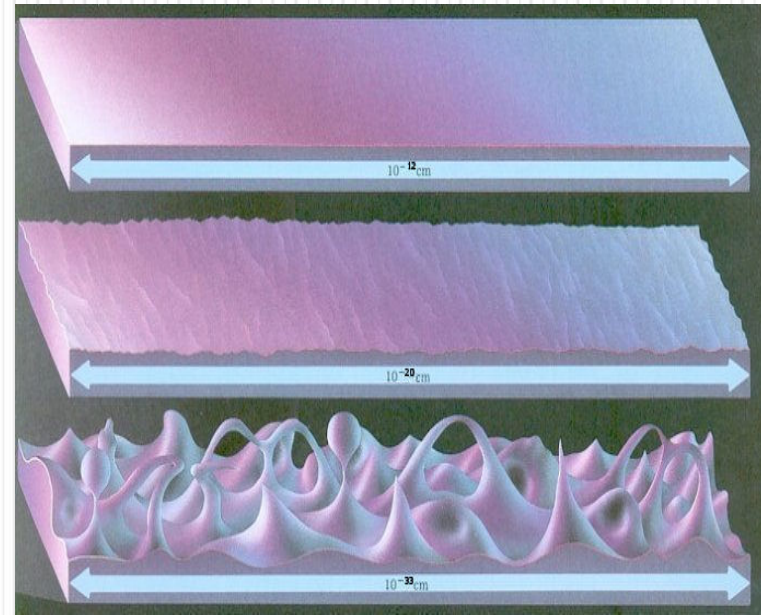
Concrete example:



- **BH entropy may be entanglement entropy.**
(they scale the same way)
- **But for that there must be a natural UV cutoff.**
- **How does spacetime look at the Planck scale ?**

Concretely:

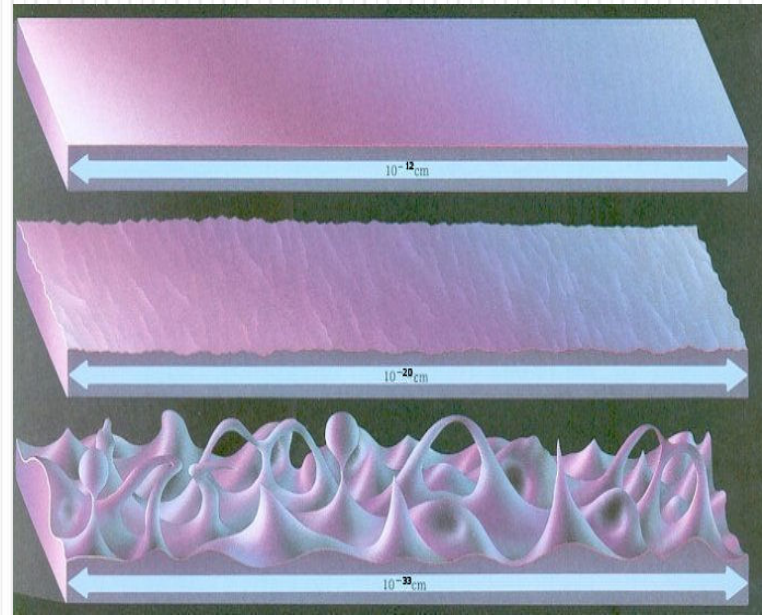
When we zoom in,
does space look like this ?



Concretely:

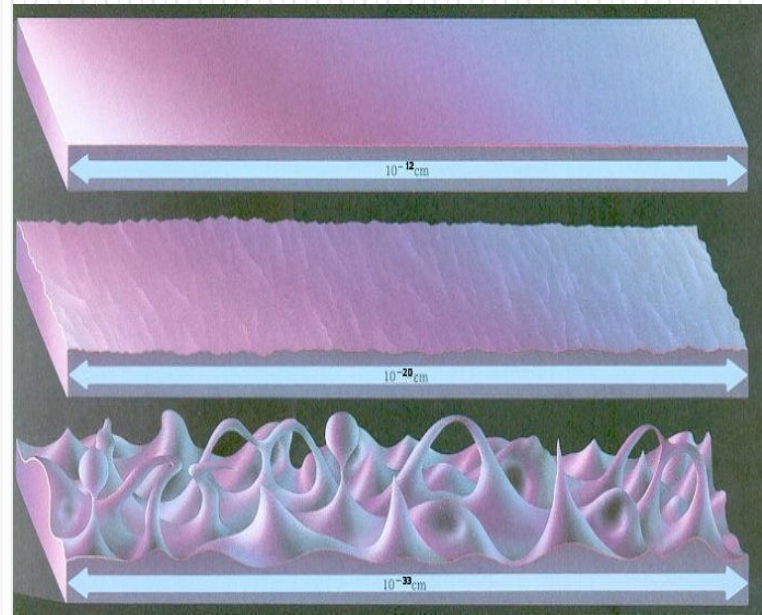
When we zoom in,
does space look like this ?

And does time look like
that ?



Concretely:

When we zoom in,
does space look like this ?



And does time look like
that ?

Check operational meaning!



What happens, operationally, as one approaches the Planck Scale?

Resolve a distance more and more precisely.

=> increasing momentum uncertainty,

=> increasing curvature uncertainty,

=> increasing distance uncertainty.

→ **Cannot resolve distances below $10^{(-35)}\text{m}$.**



What is the structure of spacetime ?

(and does information theory come up naturally?)

Paradox:

General relativity:

- Fields live on a differentiable spacetime manifold.

Quantum field theory:

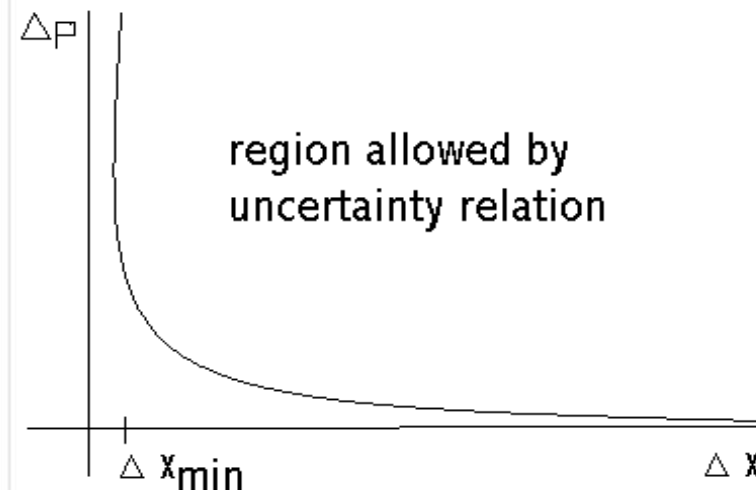
- QFT generally only well defined if spacetime is discrete.



Possible resolution

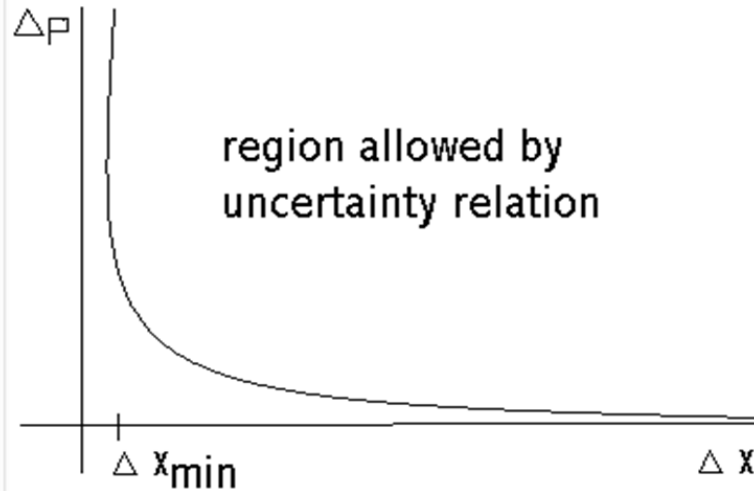
Studies in quantum gravity and string theory

\Rightarrow



Canonical commutation relations and Hilbert space representation:

- A. Kempf, J. Math. Phys. **35**, 4483 (1994), hep-th/9311147
- A. Kempf, G. Mangano, and R.B. Mann, Phys. Rev. D **52**, 1108 (1995), hep-th/9412167



- *If so, fields must possess a finite bandwidth !*
- **Spacetime is both discrete *and* continuous, in the same mathematical way that information is.**

See:

- A. Kempf, Phys. Rev. Lett. **85**, 2873 (2000)
- A. Kempf, Phys. Rev. Lett. **92**, 221301 (2004)
- A. Kempf, R. Martin, Phys. Rev. Lett. **100**, 021304 (2008)
- A. Kempf, Phys. Rev. Lett. **103**, 231301 (2009)
- D. Aasen, T. Bhamre, A. Kempf, Phys. Rev. Lett. **110**, 121301 (2013)

Information theory does come up naturally

Information can be:

- continuous (e.g., music):



- discrete (letters, digits, etc):

R 7 2 S B

Unified in 1949 by Shannon, through: Sampling theory

Applications ubiquitous:

- communication engineering & signal processing
- scientific data taking, e.g., in astronomy.

Shannon's sampling theorem

- Assume f is “bandlimited”, i.e:

$$f(x) = \int_{-\omega_{\max}}^{\omega_{\max}} \tilde{f}(\omega) e^{-2\pi i \omega x} d\omega$$

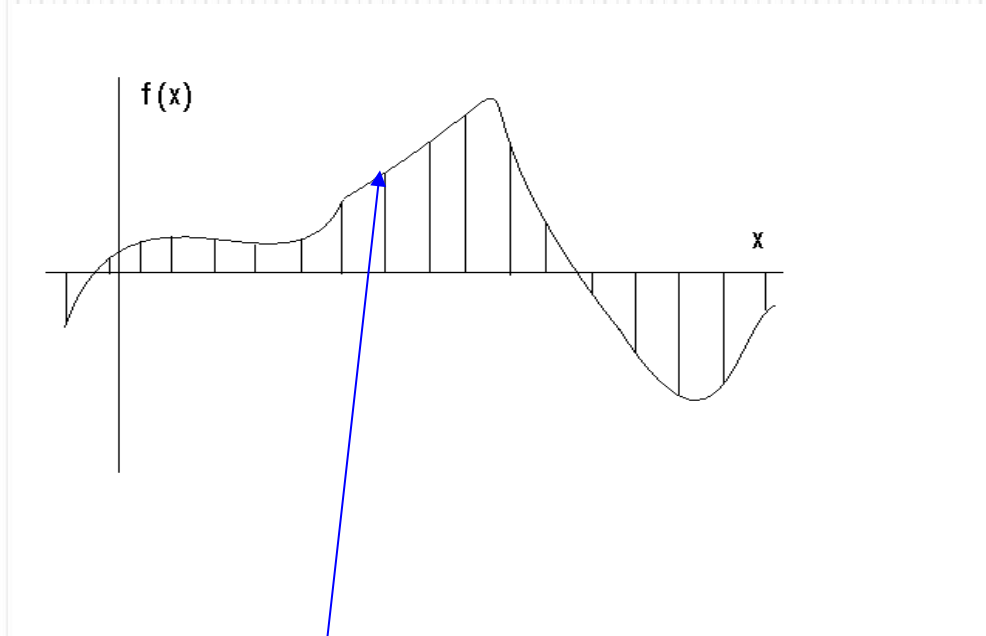
- Take samples of $f(x)$ at Nyquist rate:

$$x_{n+1} - x_n = (2\omega_{\max})^{-1}$$

- Then, exact reconstruction is possible:

$$f(x) = \sum_n f(x_n) \frac{\sin[2\pi(x-x_n)\omega_{\max}]}{\pi(x-x_n)\omega_{\max}}$$

samples



Properties of bandlimited functions

- Differential operators are also finite difference operators.
- Differential equations are also finite difference equations.
- Integrals are also series:

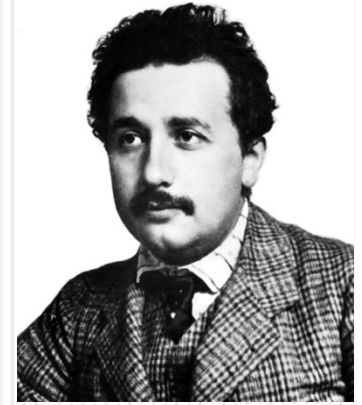
$$\int_{-\infty}^{\infty} f(x)^* g(x) dx = \frac{1}{2\omega_{\max}} \sum_{n=-\infty}^{\infty} f(x_n)^* g(x_n)$$

Remark:

Useful also as a summation tool for series

(traditionally used, e.g., in analytic number theory)

Covariant “bandlimitation” ?



Cut off of the spectrum of
the Laplacian or d'Alembertian.

$$Z[J] = \int_{\mathcal{F}} e^{iS[\phi] + i \int J\phi} d^n x D[\phi]$$

→ The space of fields, \mathcal{F} , in the QFT path integral
is spanned by the eigenfunctions w. eigenvalues:

$$\lambda_i < \Lambda$$

What if physical fields are “bandlimited”?

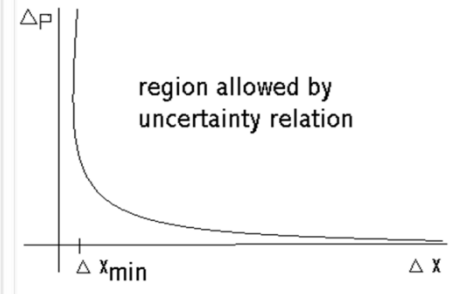
Fields possess equivalent representations

- on a differentiable spacetime manifold

(which shows preservation of external symmetries)

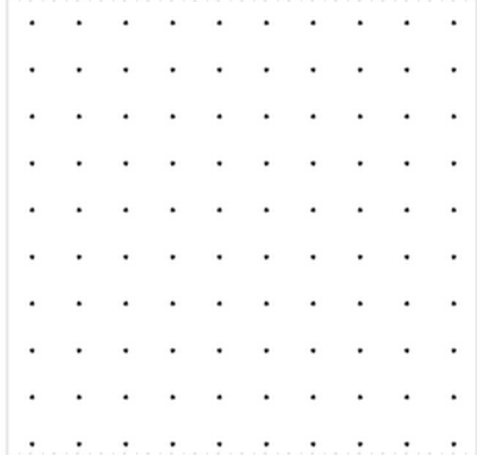
- on any lattice of sufficiently dense spacing

(which shows UV finiteness of QFTs).



Entanglement entropy

$$\nabla^2 \psi - \frac{1}{c^2} \frac{\partial^2 \psi}{\partial t^2} = \frac{m^2 c^2}{\hbar^2} \psi$$

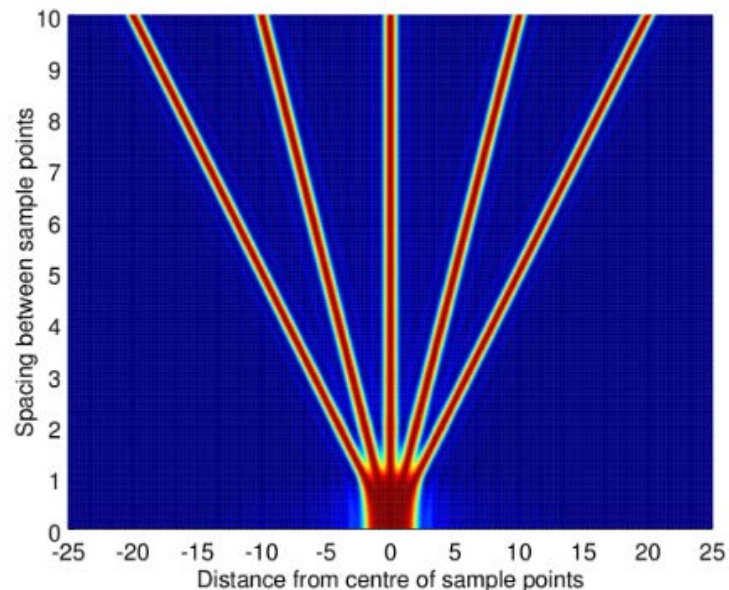


For lattice QFT:

- Coupled harmonic oscillators
- Their ground state is short-range entangled
- Entanglement entropy obeys volume and area laws.

Ent. entropy with bandwidth cutoff

With Jason Pye (UW) and William Donnelly (UCSB):



- Find smooth transition from volume to area law.

The physics of the Nyquist rate?

- Area law kicks in at Nyquist rate.
- On flat space, Nyquist rate is constant.
- How does curvature affect the Nyquist rate?

Density of degrees of freedom ?

We can use [Gilkey 1975]:

Consider any compact 4-dim Riemannian manifold. Then:

$$N = \frac{1}{16\pi^2} \int d^4x \sqrt{|g|} \left\{ \frac{\Lambda^2}{2} + \frac{\Lambda}{6} R + \mathcal{O}(R^2, \Lambda^{-1}) \right\}$$

Can now read off:

- Cosmological constant is density of DOF: $\frac{N}{V} = \frac{\Lambda^2}{32\pi}$
- Curvature is local perturbation of density of DOF.

Re-expressing the Einstein action

=> Einstein action takes the simple form:

$$\begin{aligned} S &= \frac{6\pi}{\Lambda} \frac{1}{16\pi^2} \int d^4x \sqrt{|g|} \left\{ \frac{\Lambda^2}{2} + \frac{\Lambda}{6} R + O(R^2, \Lambda^{-1}) \right\} \\ &= \frac{6\pi}{\Lambda} N \\ &= \frac{6\pi}{\Lambda} \text{Tr}(1) \end{aligned}$$

Notice: The Einstein action is the integral over the density of degrees of freedom, where the cosmological constant sets the baseline, modulated by curvature.

Compare: scalar field action

- In the eigenbasis of the Laplacian is not only the Einstein action diagonal but also the action of a scalar field:

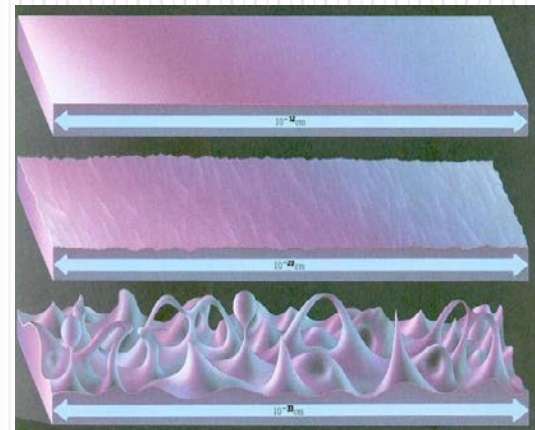
$$\begin{aligned} S_{matter} &= \int d^n x \sqrt{|g|} \frac{1}{2} \phi(x) (\Delta + m^2) \phi(x) \\ &= \sum_{i=1}^N \frac{1}{2} \phi_i (\lambda_i + m^2) \phi_i \\ &= \text{Tr} \left[\frac{1}{2} (\Delta + m^2) | \phi \rangle \langle \phi | \right] \end{aligned}$$

- **Actions are traces, and gravity could be a leading constant.**

But what is bandlimitation for spacetime itself ?

Is there a Shannon-like reconstruction of space from discrete sets of samples?

With bandwidth / min uncertainty cutoff,
what could supercede rulers and clocks ?



Idea:

Noise correlator as proxy for distance

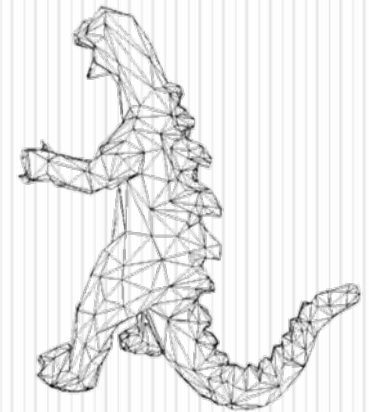
- Quantum field correlators indicate distance.
- Does the entanglement structure of the vacuum encode spacetime's curvature ?

Idea: correlators as proxy for distances

At N points \mathbf{x}_i of a finite piece of the manifold,

sample the propagator's matrix elements:

$$\langle \mathbf{x}_a | 1/\Delta | \mathbf{x}_b \rangle$$



- Work w. Aslanbeigi and Saravani: One can reconstruct the metric.
- Basis independent information \rightarrow eigenvalues of Δ .
- **Does the spectrum tell the shape?**

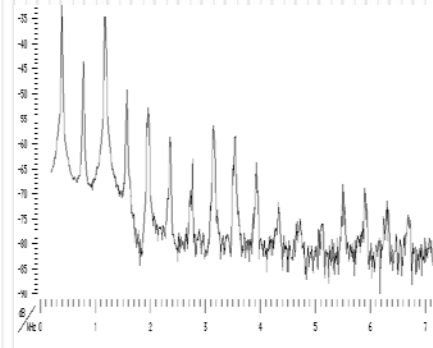
Spectral Geometry:

- “ How far is shape determined by sound ? ”

$$-d^2\phi/dt^2 = \Delta_g \phi$$



(M, g)

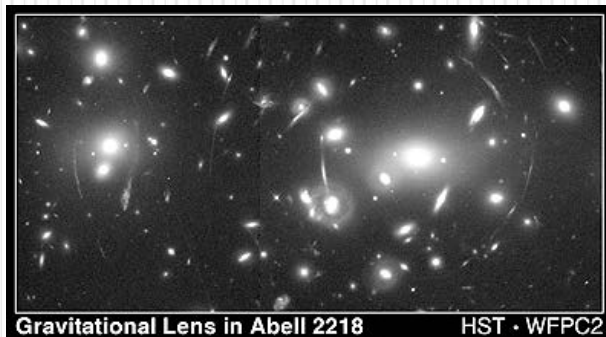


$\text{spec}(\Delta)$

There are some positive results, e.g., on shapes of revolution!

Prospect:

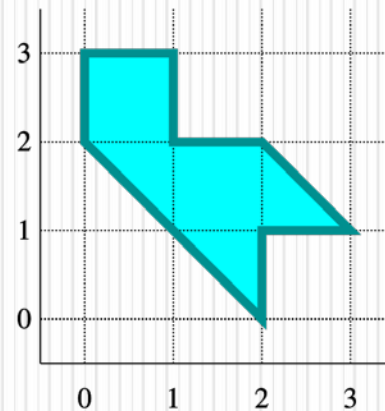
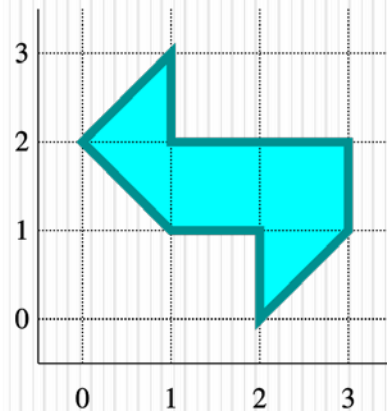
Can one hear a spacetime's curvature in its quantum noise?



Deep link between gravity and quantum theory?

Problem !

- **Spectral geometry has counter examples !**
- Work by Milnor, Sunada, Gordon ...



Solution:

Infinitesimal spectral geometry & tensor spectra

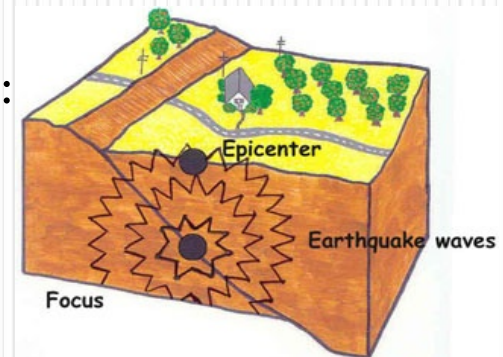
In dimensions $d > 2$, not every perturbation of a Riemannian manifold can be described by a scalar function f .

$$g_{\mu\nu}(x) \rightarrow (1 + f(x)) g_{\mu\nu}(x)$$

Need to use scalar, vector and tensor perturbations:

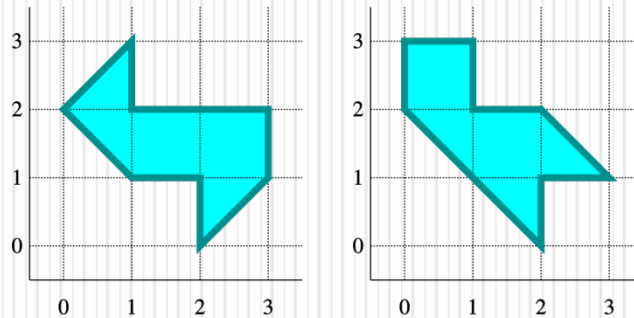
$$g_{\mu\nu}(x) \rightarrow g_{\mu\nu}(x) + \delta s_{\mu\nu}(x) + \delta v_{\mu\nu}(x) + \delta h_{\mu\nu}(x)$$

(Seismic waves of different types carry independent information too)



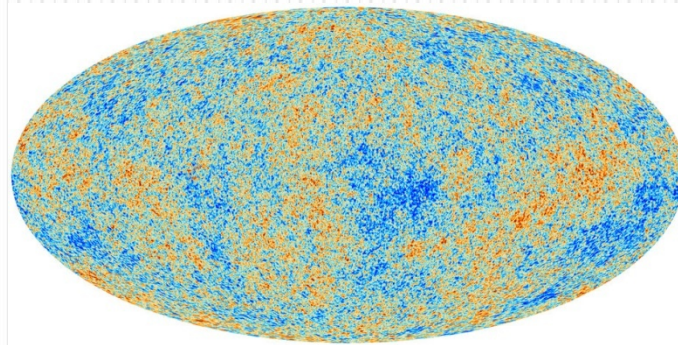
Work with Mikhail Panine (UW)

- We showed that even spectral geometry of planar domains works generically.



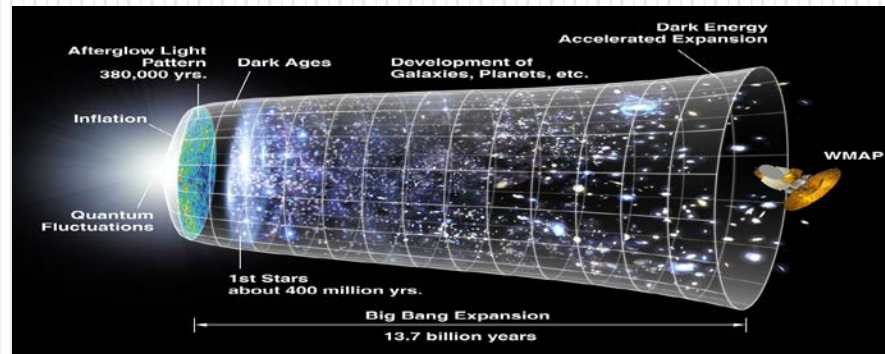
Experimental predictions ?

CMB is closest to Planck scale

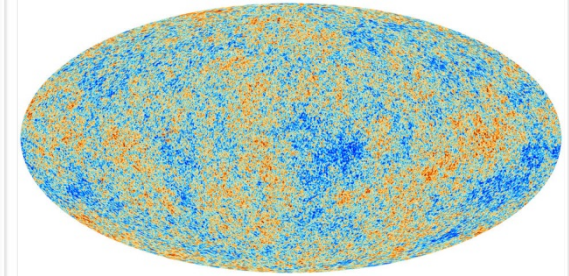


Why?

Hubble scale in inflation only about 5 orders from Planck scale.



Applied to cosmology



Multiple groups have non-covariant predictions for CMB.

- Characteristic, $O(10^{-5})$ or $O(10^{-10})$ modulations
- Characteristic deviation from scalar/tensor consistency relation in B-polarization data.

Big challenge:

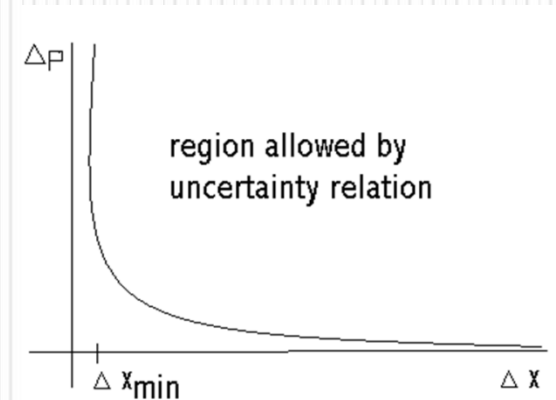
Predictions with local Lorentz covariant bandlimit cutoff!

Upcoming work with:

Aidan Chatwin-Davies (CalTech) and Robert Martin (U. Cape Town)

Summary

- Philosophy: Only information theoretic concepts may survive at Planck scale
- Found: Spacetime may be bandlimited
- Thus discrete = continuous



for spacetime, as is the case for information.

- Notion of distance replaced by info-theoretic notion of correlation.
- Most also already works Lorentz-covariantly (pls ask).

Thank you