



# Evolution of Temperature Fluctuation in a Viscous Medium

(arXiv:1510.03154)

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XI Workshop on Particle Correlations and Femtoscopy 2015,  
Warsaw, Poland; 3-8 Nov. 2015

**Fluctuations:** We encounter in everyday examples (temperature fluctuation in a room).

### More Examples:

- ⌘ Critical Opalescence: Density fluctuation near critical point
- ⌘ Signature of temperature fluctuation in particle yield originated from high-energy collisions.
- ⌘ CMB temperature fluctuation

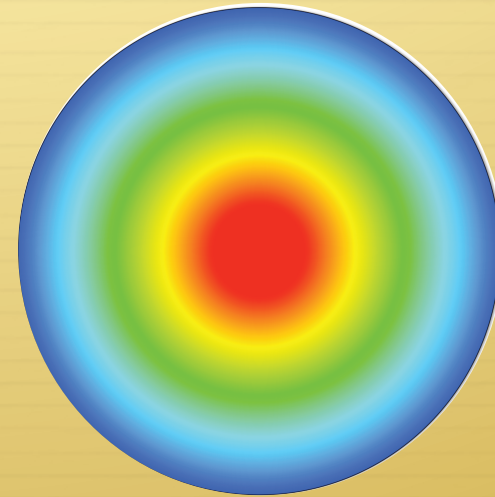
Study of temperature fluctuation may be important for studying QCD phase transition [D K Mishra *et al.* JPG 42, 105105(2015)]

and

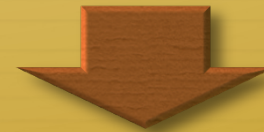
for characterizing QGP [S Basu *et al.* arXiv:1405.3969 [nucl-ex], *ibid.* 1504.04502]



System with same  
temperature everywhere



System with different  
local temperatures



**Hotspots**

**The Model:** Medium consists of weakly interacting hot zones.

Within a certain pre-determined time slice  $\Delta t$ , the temperature of a certain hot zone does not change.

Within  $\Delta t$ , they are represented by a collection of canonical (say, no chemical potential) ensembles and their (inverse) temperature has a particular distribution.

The distribution determines the average inverse temperature  $\beta$  and its fluctuation  $\Delta\beta$  at every hot zone.

With time, temperature distribution changes and so does  $\Delta\beta$ .

We assume particle distribution is affected by both  $\beta$  and  $\Delta\beta$

$$\text{Ansatz: } f = e^{-\beta(1+\Delta\beta)p} = f^{eq} + \Delta f$$

[S. Dodelson, Modern Cosmology]

Feed  $f$  inside the Boltzmann Transport Equation (BTE), which is the evolution equation for inhomogeneous, anisotropic distribution in presence of external force

$$\frac{\partial f}{\partial t} + \vec{v} \cdot \vec{\nabla} f + \vec{F} \cdot \vec{\nabla}_p f = C[f]$$

Momentum space gradient ↓  
External force ↓  
Collision term (takes care of interaction due to which distribution may change) ↓  
Inhomogeneity ↓

We get the evolution equation for  $\Delta f$  and hence for  $\Delta\beta$  assuming **Relaxation Time Approximation (RTA)**

$$C[f] = \frac{f - f^{eq}}{t_R}$$



Relaxation time ( The time within which the distribution changes appreciably )

The Fourier Space (k-space) variation of  $\Delta\beta$  (in a medium of almost massless particles):

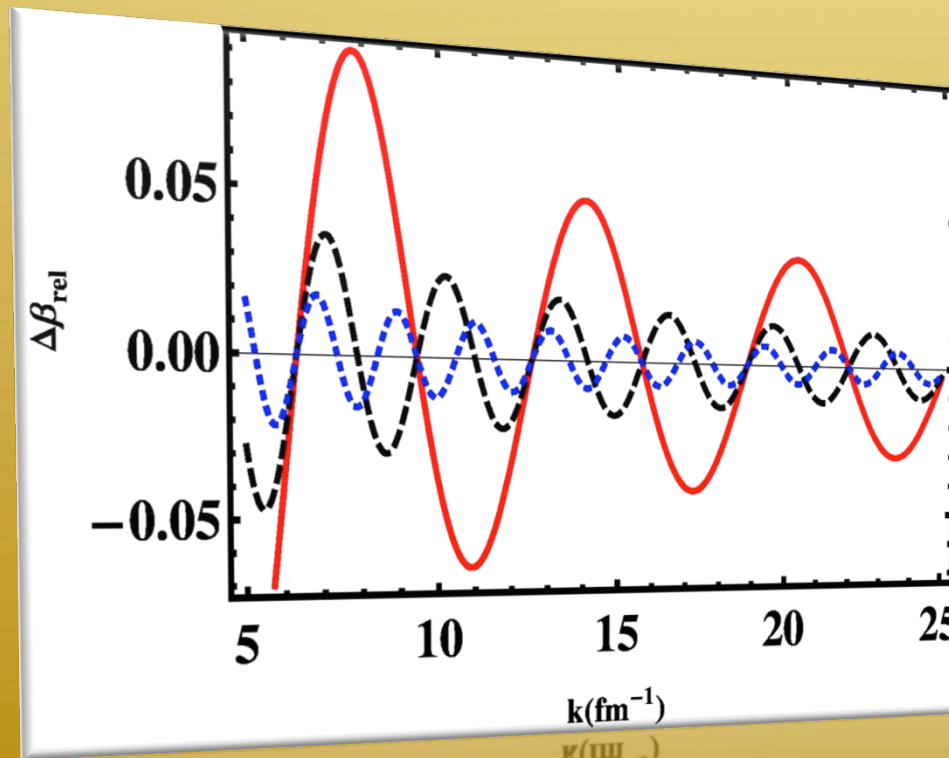
$$\Delta\beta(\vec{k}, \hat{p}; t) = \Delta\beta(\vec{k}, \hat{p}; t^0) e^{-ik\mu(t-t^0)} e^{-\frac{t-t^0}{t_R}}$$

Provided average (inverse) temperature varies slowly with time

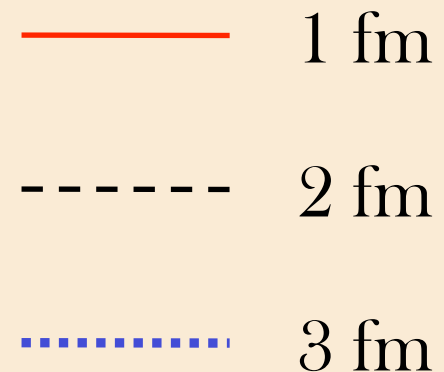


After averaging over  $\mu = \hat{k} \cdot \hat{p}$

$$\Delta\beta_{rel}(\vec{k}, t) = e^{-\frac{t-t^0}{t_R}} \frac{\sin k(t-t^0)}{k(t-t^0)}$$



$(t-t^0)$  values for  
 $t_R = 3$  fm



## Observations:

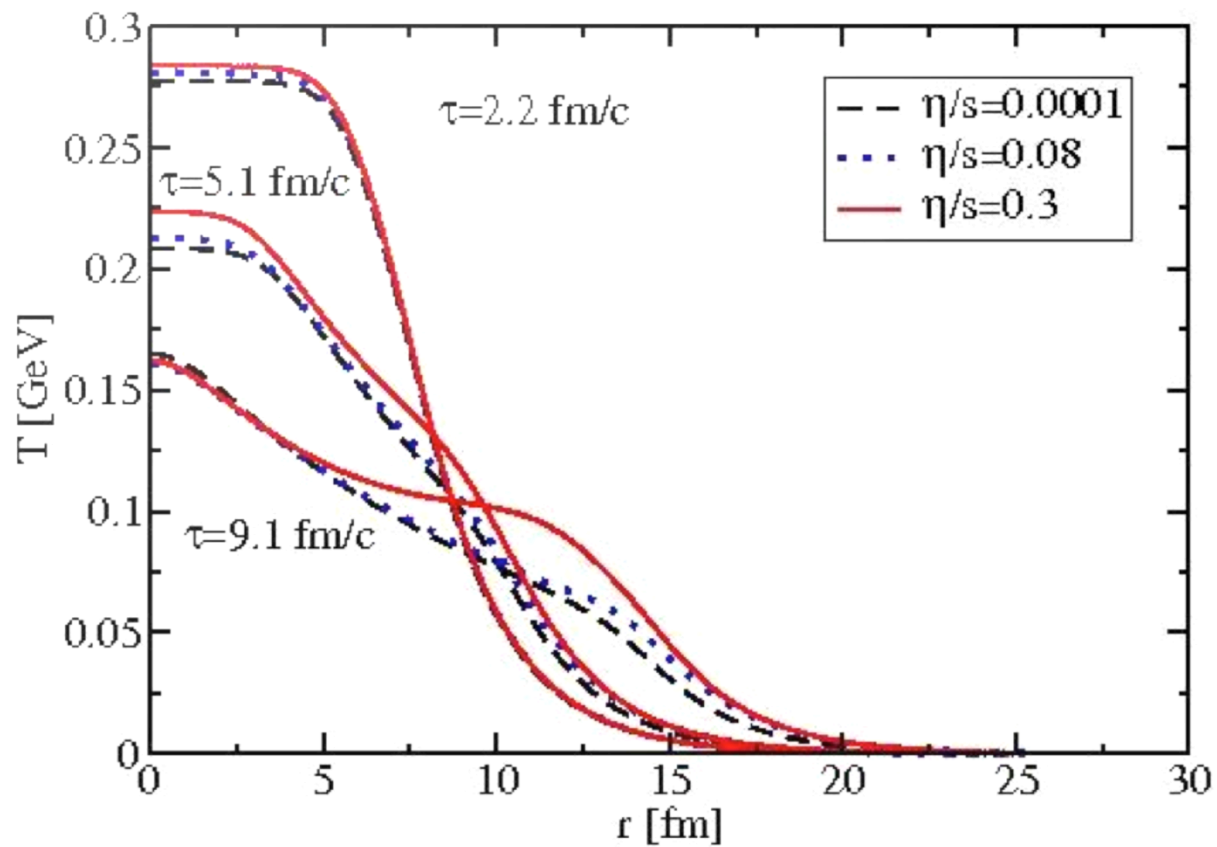
- ⊙ The amplitude of  $\Delta\beta_{rel}$  dies down with time and is more towards the smaller  $k$ , i.e. towards large radius of the system.
- ⊙ Analysis driven by BTE is limited by the constraint over the observation time

$$(t - t^0) \ll t_R$$

- ⊙ Given the constraint,  $\Delta\beta_{rel}$  is independent of  $t_R$

Any generalized analysis possible which will be able to avoid the constraint ?

Yes, to start with, one of the ways is to get the temperature profile of the medium at different stages (i.e. proper time  $\tau$ ) of evolution.



R. Baier P. and Romatschke, EPJC 51, 677(2007)

The temperature profile obtained from the theoretical analysis for a viscous QGP medium (created in a central HIC) evolving hydrodynamically can be described by the following function:

$$\beta_M(r;t) = \beta_0(t) \left[ e^{a(t) \left( \frac{r}{r_0} - 1 \right)} + 1 \right]$$

With the tabulated details:

$\tau(fm/c)$	$\beta_0(GeV^{-1})$	$a$	$r_0(fm)$
2.2	3.45	5.99	7.96
5.1	4.55	3.42	8.41
9.1	5.56	1.91	8.71

From this, we can generate  $\{\beta_{Mi}\}$ , a sequence of radially varying inverse temperature values at different stages.

$$\{\beta_{Mi}\}_\tau \equiv \{\beta_{M0}, \beta_{M1}, \beta_{M2}, \beta_{M3}, \dots, \beta_{Mn}\}_\tau$$

$$\langle \beta_M \rangle = \frac{\beta_{M0} + \beta_{M1} + \beta_{M2} + \beta_{M3} + \dots + \beta_{Mn}}{n}$$

$$\langle \beta_M^2 \rangle = \frac{\beta_{M0}^2 + \beta_{M1}^2 + \beta_{M2}^2 + \beta_{M3}^2 + \dots + \beta_{Mn}^2}{n}$$

So, for a given  $\{\beta_{Mi}\}$ , we can define

❖ An average  $\langle \beta_M \rangle$

❖ Also, a fluctuation on top of it:

$$\begin{aligned}\Delta\beta(r;t) &= \beta_M(r;t) - \langle \beta_M \rangle \\ &= \beta_0(t) e^{a(t)\left(\frac{r}{r_0}-1\right)} + \delta\beta(t)\end{aligned}$$

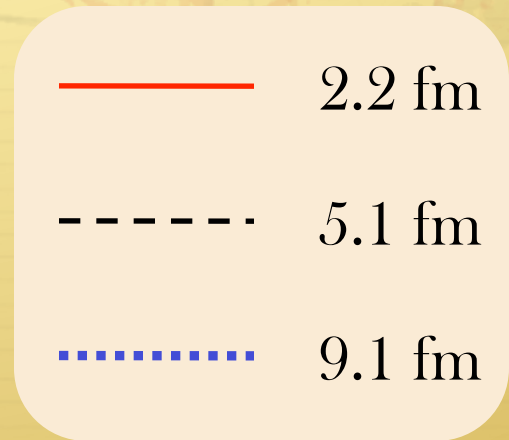
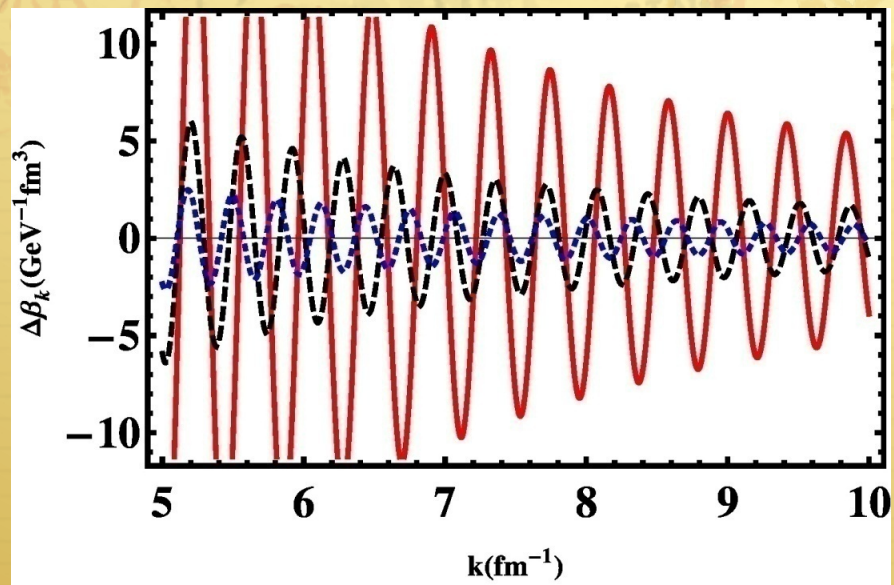
with  $\delta\beta(t) = \beta_0(t) - \langle \beta_M \rangle$



The Fourier Space (k-space) variation of inverse temperature fluctuation:

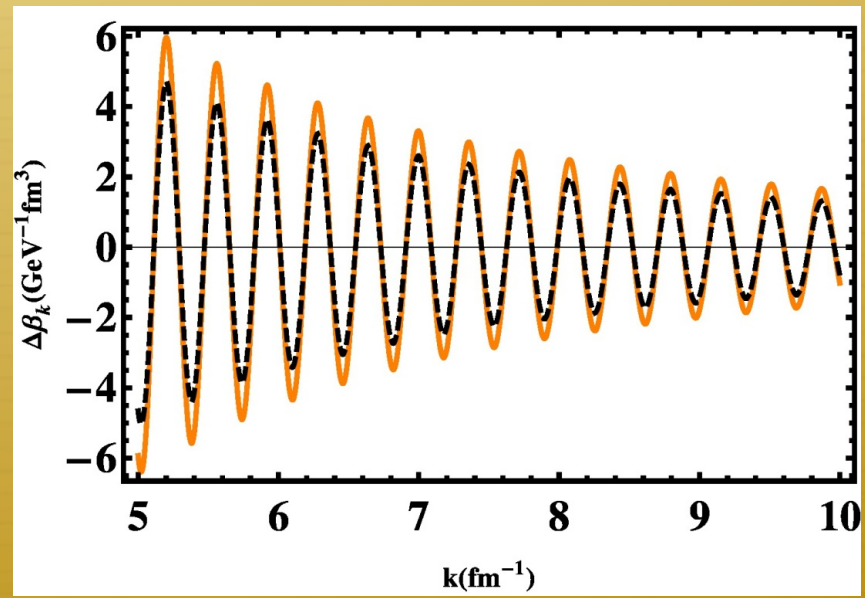
$$\Delta\beta(k, t) = \frac{2\beta_0(t)}{(2\pi)^2 k} \int_0^R e^{a(t)\left(\frac{r}{r_0}-1\right)} r \sin kr \, dr + \delta\beta(t)\delta(\vec{k})$$

where,  $\delta(\vec{k})$  is the Dirac delta function.



—  $\eta / s = 0.08$

- - -  $\eta / s = 0.30$



❖ And, a relative variance for the collection  $\{\beta_{M_i}\}$ :

$$\frac{\langle \beta_M^2 \rangle - \langle \beta_M \rangle^2}{\langle \beta_M \rangle^2} = \frac{\sigma_\beta^2}{\beta^2} = \mathfrak{R}_\beta$$

Recall:

$$\frac{\langle (\frac{1}{T})^2 \rangle - \langle (\frac{1}{T}) \rangle^2}{\langle (\frac{1}{T}) \rangle^2} = q - 1$$

G. Wilk and Z. Włodarczyk, Phys. Rev. Lett. 84 (2000) 2770

$\tau$	$\mathfrak{R}_\beta$
2.2	0.047
5.1	0.011
9.1	0.002

$$\frac{\eta}{s} = 0.08$$

$\frac{\eta}{s}$	$\mathfrak{R}_\beta$
0.08	0.012
0.30	0.011

$$\tau = 5.1 \text{ fm}$$

Let us assume that the system produced by central HIC freezes-out by 9.1 fm and compare theoretically obtained  $\mathcal{R}_\beta$  value at the boundary with the similar experimentally observed (q-1) parameter [G. Wilk and Z. Włodarczyk PRL 84, 2770(2000)] for hadron spectra at  $\sqrt{s_{NN}} = 200$  GeV within 0-10% centrality [Z. Tang *et al.* PRC 79, 051901(R)(2009)].

$\mathcal{R}_\beta^{Theory}$	(q-1)
0.013	$0.018 \pm 0.005$

## Cosmological connections:

- ⌘ Temperature fluctuation of our universe can be explained by the modified Boltzmann-Gibbs formula with  $(q-1)$  value  $0.045 \pm 0.005$  [*A. Bernui et al. PLA 356, 426(2006)*]
- ⌘ Study on the similarity between the surface of the last scattering for CMB radiation and the freeze-out surface in RHICE
- ⌘ Similarity with HIC experimental results: needs review

## Summary and Conclusion:

- ◎ Time evolution of (inverse) temperature fluctuation
- ◎ With time, amplitude of inverse temperature fluctuation decreases
- ◎ With distance amplitude of inverse temperature fluctuation increases
- ◎ Relative fluctuation at the boundary is comparable with the experimental value under similar conditions.

- © More studies required on the evolution of temperature fluctuation with more generalized scenarios.
- © Exploring possible connections with experiment in a more detailed manner.

Thank you