Beyond the Standard Model

CERN summer student lectures 2015



Lecture 3/5

Christophe Grojean

DESY (Hamburg) ICREA@IFAE (Barcelona) (christophe.grojean@cern.ch)

Outline

Monday

O general introduction, units

• Tuesday

• Higgs physics as a door to BSM

□Wednesday

Higgs and Naturalness: small and large numbers in a quantum world
 Thursday

• grand unification, proton decay

O supersymmetry

0 extra dimensions

Friday

O cosmological interplay

Higgs and Flavor

In SM, the Yukawa interactions are the only source of the fermion masses



Not true anymore if the SM fermions mix with vector-like partners or for non-SM Yukawa

$$y_{ij}\left(1+c_{ij}\frac{|H|^2}{f^2}\right)\bar{f}_{L_i}Hf_{R_j} = \frac{y_{ij}v}{\sqrt{2}}\left(1+c_{ij}\frac{v^2}{2f^2}\right)\bar{f}_{L_i}f_{R_j} + \left(1+3c_{ij}\frac{v^2}{2f^2}\right)\frac{y_{ij}}{\sqrt{2}}h\bar{f}_{L_i}f_{R_j}$$

Look for SM forbidden Flavor Violating decays $h \rightarrow \mu \tau$ and $t \rightarrow hc$

• weak indirect constrained by flavor data (e.g. $\mu \rightarrow e\gamma$): BR<10% • ATLAS and CMS have the sensitivity to set bounds O(1%) • ILC/CLIC/FCC-ee can certainly do much better

(*) e.g. Buras, Grojean, Pokorski, Ziegler '11

Christophe Grojean

Higgs and Flavor

In SM, the Yukawa interactions are the only source of the fermion masses



HEP with a Higgs boson

" If you don't have the ball, you cannot score

Now with the Higgs boson in their feets, particle physicists can... play as well as Barça players



Profound change in paradigm:

missing SM particle is tool to explore SM and venture into physics landscape beyond

Christophe Grojean

Higgs and EW vacuum Stability



$$V(h) = -\frac{1}{2}\mu^2 h^2 + \frac{1}{4}\lambda h^4$$

vev: $v^2 = \mu^2/\lambda$ mass: $m_H^2 = 2\lambda v^2$
the vacuum is not empty even classically $(\hbar \to 0)$

How is Quantum Mechanics changing the picture?



Christophe Grojean

Higgs and EW vacuum Stability



Christophe Grojean

Higgs and EW vacuum Stability



Quantum Instability of the Higgs Mass

so far we looked only at the RG evolution of the Higgs quartic coupling (dimensionless parameter). The Higgs mass has a totally different behavior: it is highly dependent on the UV physics, which leads to the so called hierarchy problem



$$\delta m_H^2 = \left(2m_W^2 + m_Z^2 + m_H^2 - 4m_t^2\right) \frac{3G_F \Lambda^2}{8\sqrt{2}\pi^2}$$
$$\vdots$$
$$m_H^2 \sim m_0^2 - (115 \text{ GeV})^2 \left(\frac{\Lambda}{700 \text{ GeV}}\right)^2$$

Christophe Grojean

BSM

46

CERN, July 2015







Naturalness principle

Following the arguments of Wilson, 't Hooft (and others):

only small numbers associated to the breaking of a symmetry survive quantum corrections



m

courtesy to N. Craig @ Blois '15

 $\delta m \propto \Lambda$

The Higgs mass in the SM doesn't break any (quantum*) symmetry

* it does break classical scale invariance, as the running of the gauge couplings does too!

Christophe Grojean

BSM

Naturalness principle @ work

Following the arguments of Wilson, 't Hooft (and others):

only small numbers associated to the breaking of a symmetry survive quantum corrections

Beautiful examples of naturalness to understand the need of "new" physics see for instance Giudice '13 (and refs. therein) for an account

 \blacktriangleright the need of the positron to screen the electron self-energy: $\Lambda < m_e/\alpha_{\rm em}$

▶ the rho meson to cutoff the EM contribution to the charged pion mass: $\Lambda < \delta m_{\pi}^2 / \alpha_{em}$ ▶ the kaon mass difference regulated by the charm quark: $\Lambda^2 < \frac{\delta m_K}{m_K} \frac{6\pi^2}{G_F^2 f_K^2 \sin^2 \theta_C}$

the light Higgs boson to screen the EW corrections to gauge bosons self-energies
 ...

▶ New physics at the weak scale to cancel the UV sensitivity of the Higgs mass?

Playing with cracks: The way forward

Small numbers are not necessarily theoretically inconsistent but they require some conspiracy at different scales

Better to find an explanation with new degrees of freedom that cancel the sensitivity to the details of the physics at high-energy

Theoretical inconsistencies

* 4 Fermi interactions to describe muon decay $\mathcal{A} \sim G_{\mathcal{F}}E^2 \gg$ W boson

* $W_L W_L$ scattering $\mathcal{A} \sim E^2/v^2 \gg H$ boson

Naturalness arguments

* positron
* rho
* charm quark
* susy?

Small Numbers in a Quantum World

the mass spectrum of the fermions is intriguing



Small Numbers in a Quantum World

the mass spectrum of the fermions is intriguing



but this spectrum is stable under radiative corrections

 $\delta m_e \propto m_e$ if the electron mass is small, it will remain small in a quantum world

the origin of this intriguing spectrum might come from dynamics at much higher scales that will never be explored at colliders

When a small number is "protected" by a symmetry, quantum corrections won't affect it and we can safely postpone the question "why is it small?" to higher energy scales The Higgs mass is a priori not protected and its "smallness" requires an explanation now!

The Higgs, as a fundamental scalar field, is the worm inside the SM

Higgs & Naturalness

The last unknown parameter of the SM has been measured

$$V(H) = -\mu^2 |H|^2 + \lambda |H|^4$$
$$\mu \approx 88.8 \,\text{GeV} \quad \lambda \approx 0.13$$



Why $m_H/M_{PI} \sim 10^{-16}$? Why m_H so close to the critical boundary?

Higgs & Naturalness

The last unknown parameter of the SM has been measured

$$V(H) = -\mu^2 |H|^2 + \lambda |H|^4$$
$$\mu \approx 88.8 \,\text{GeV} \qquad \lambda \approx 0.13$$



Why $m_H/M_{PI}\sim 10^{-16}$? Why m_H so close to the critical boundary? Change the metric? e.g. $d=(E/M_{PI})^{0.018}$

Higgs self-couplings and Naturalness

In the SM, |H|² is the only relevant operator and it is the source of the hierarchy/naturalness/fine-tuning problem It presence has never been tested!

Reconstructing the Higgs potential before EW symmetry breaking from measurements around the vacuum is difficult in general but we can easily test gross features, like the presence of the relevant operator

Solution
Solution
Solution

$$V = -\mu^2 |H|^2 + \lambda |H|^4$$

$$V(h) = \frac{1}{2}m_h^2 h^2 + \frac{1}{6}\frac{3m_h^2}{v}h^3 + \dots$$

$$V(h) = \frac{1}{2}m_h^2 h^2 + \frac{1}{6}\frac{7m_h^2}{v}h^3 + \dots$$
200% correction
to SM prediction
+
allows 1st phase transition

Symmetries to Stabilize a Scalar Potential



These symmetries cannot be exact symmetry of the Nature. They have to be broken. We want to look for a soft breaking in order to preserve the stabilization of the weak scale.

Christophe Grojean

BSM

EWSB might be unnatural

nothing to say but the usual words:

O cosmological constant problem...

O multiverse...

O landscape of vacua...

O laws of physics are environmental.

O anthropic solution...

O end of reductionism...

will be tested to an unprecedented level (10⁻⁴)





CERN, July 2015

Christophe Grojean

BSM