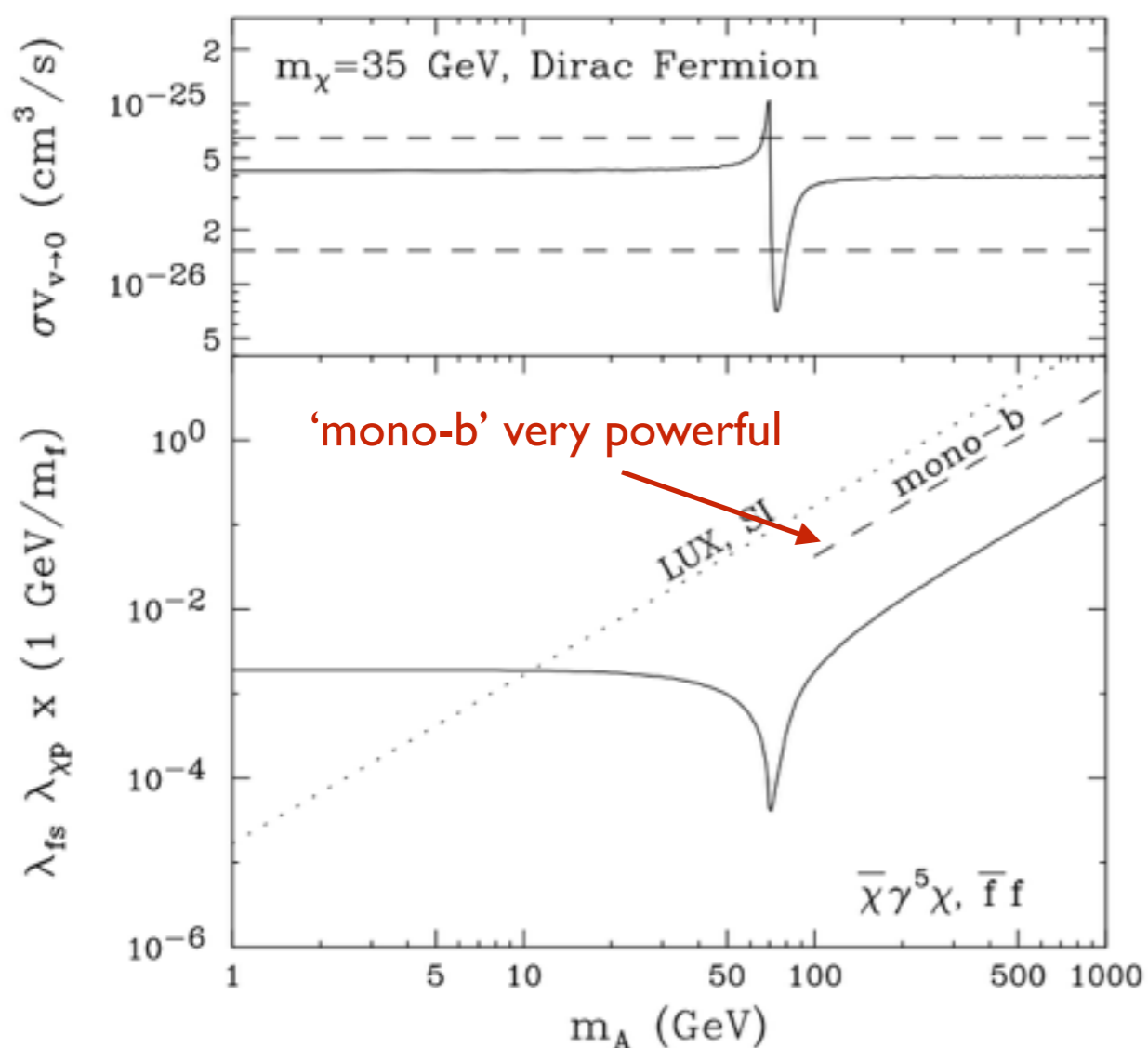
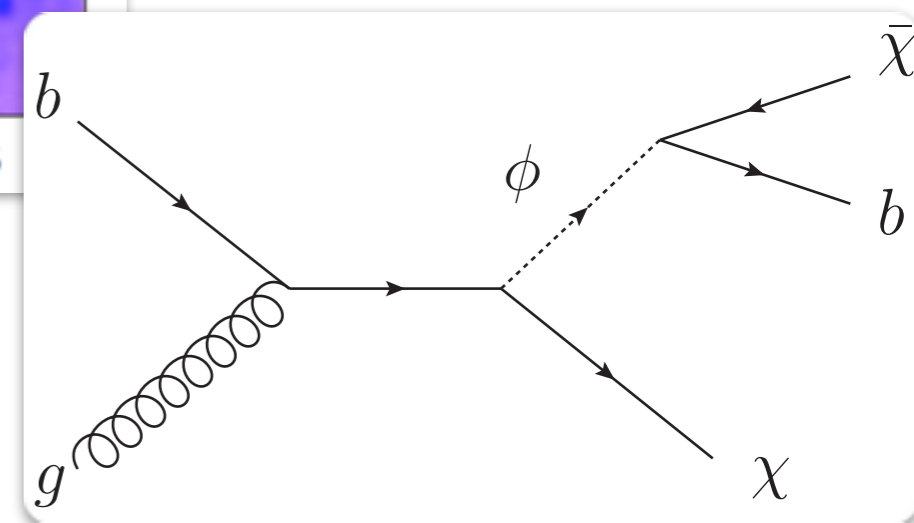
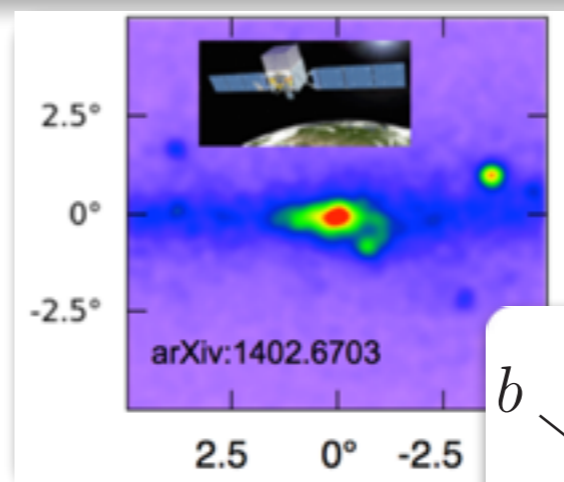


First look into b-FDM samples for 13 TeV

Björn Penning,
Deborah Pinah, Priscilla Pani

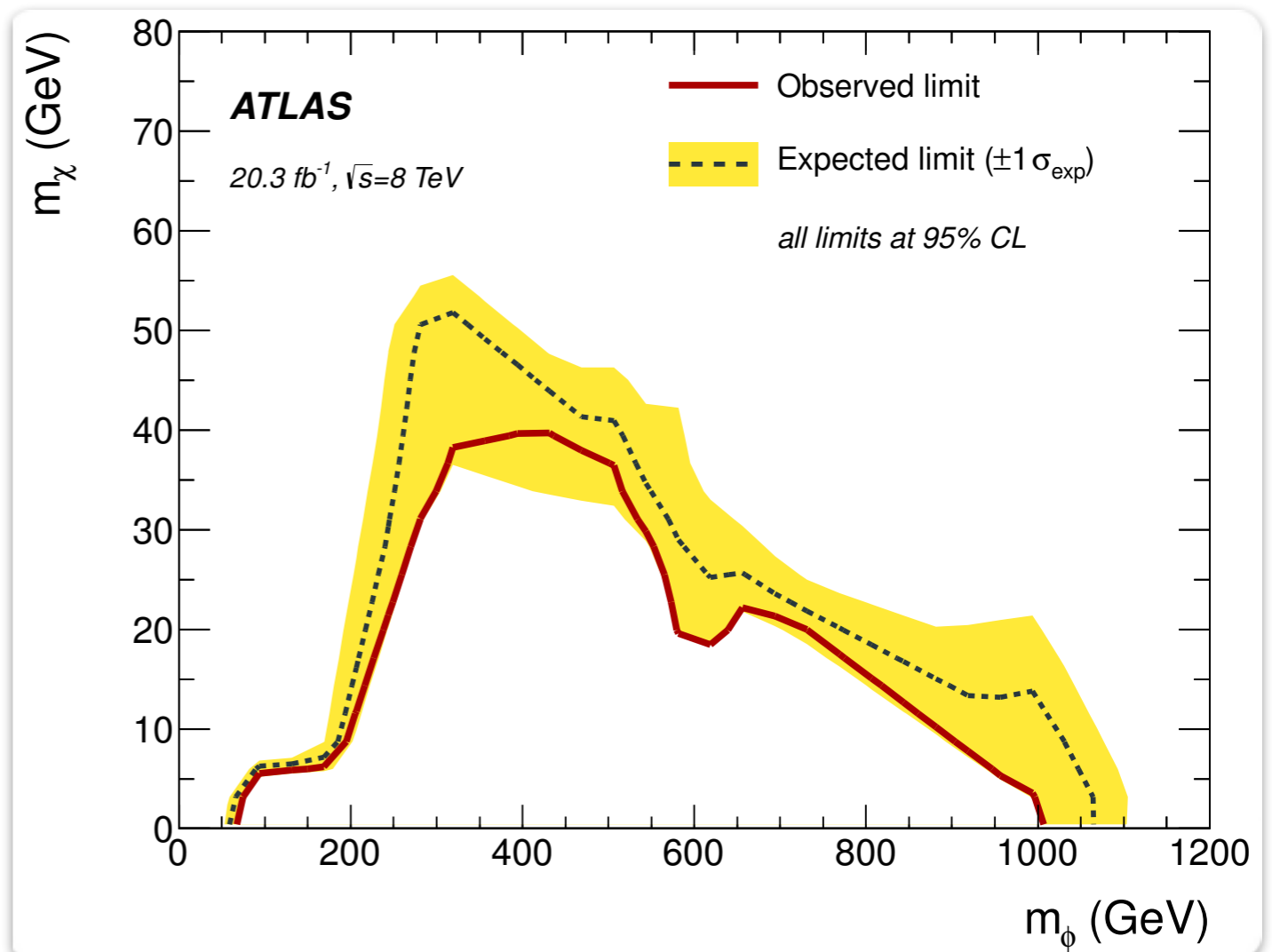
- We consider a model from [Lin et al. \[1404.1373\]](#) where a colored scalar mediator couples to dark matter and a heavy quark (specifically a b-quark).
- Model inspired by the Fermi-LAT excess



- First ‘**mono-b**’ Snowmass projections used by [Hooper et al. \[1404.0022\]](#) that ‘mono-b’ type searches constitute an important tool for low mass DM to probe the Fermi-LAT excess
- First search in ATLAS performed, laying out groundworks for 13 TeV
- Main interest Priscilla, Deborah & I

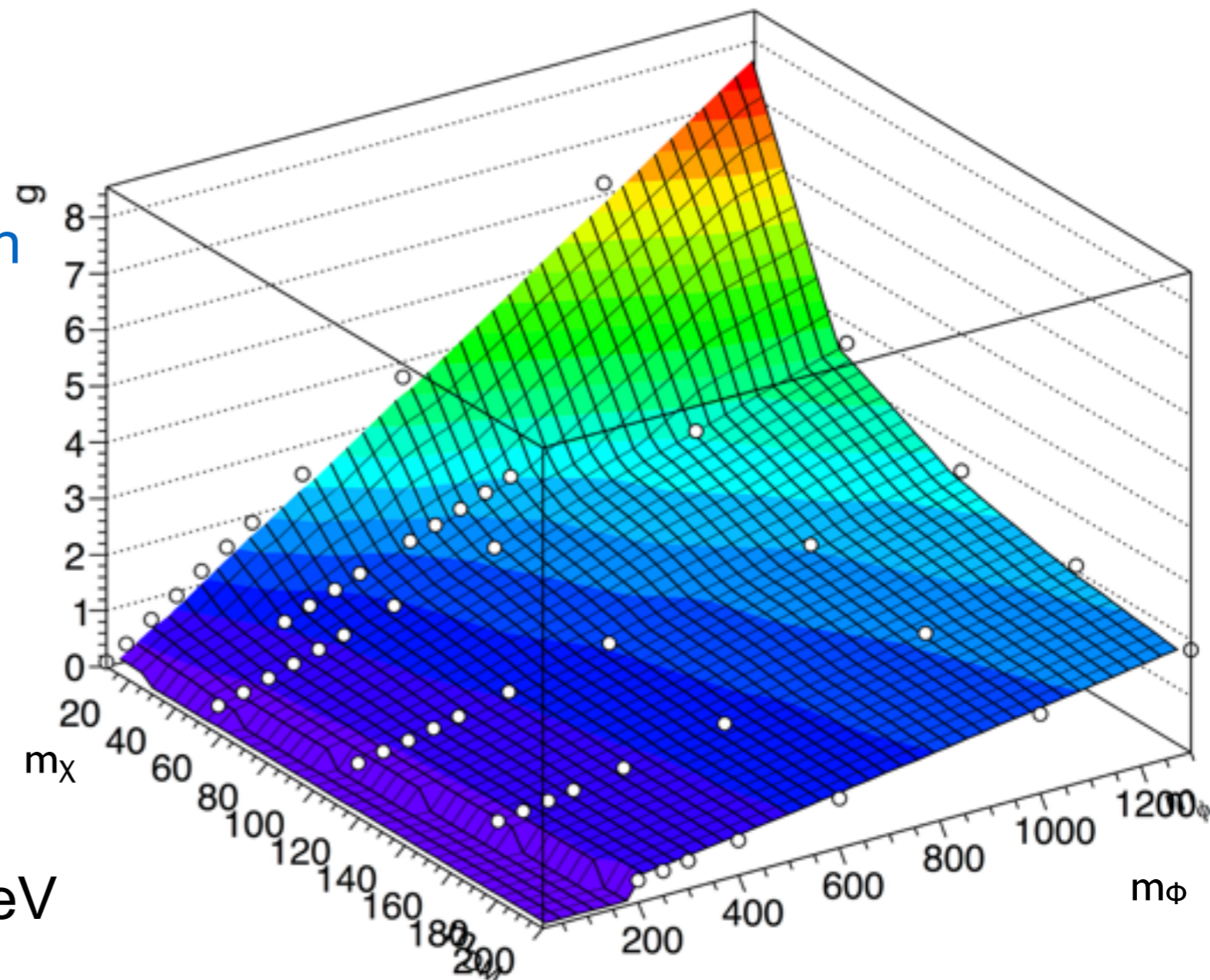
	SR1
Trigger	E_T^{miss}
Jet multiplicity n_j	1-2
b -jet multiplicity n_b	>0 (60% eff.)
Lepton multiplicity n_ℓ	0
E_T^{miss}	>300 GeV
Jet kinematics	$p_T^{b_1} > 100$ GeV
Three-jet invariant mass	
$\Delta\phi(j_i, E_T^{\text{miss}})$	> 1.0, $i = 1, 2$
Angular selections	-
Event shape	-
am_{T2}	-
$m_T^{\ell + E_T^{\text{miss}}}$	-
$E_T^{\text{miss}} / \sqrt{H_T^{4j}}$	-

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- ATLAS published corresponding analysis in the hadronic & l+jets final state (<http://arxiv.org/abs/1410.4031>)
- ATLAS analysis weighted to relic density -> shapes exclusion region to some extent

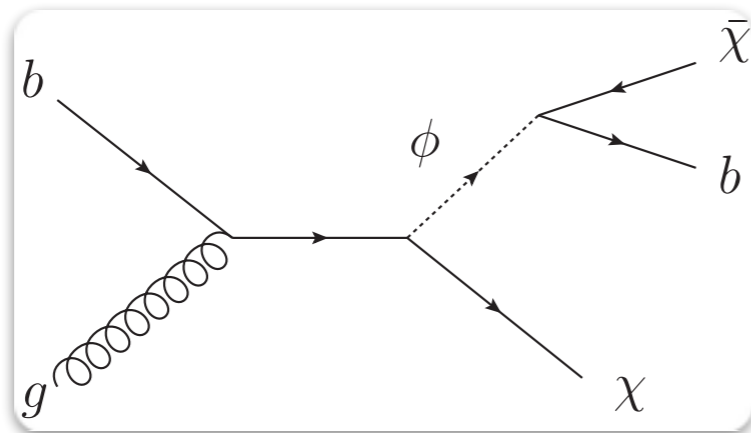
- Relic density weights are between $O(0.1 - 2)$
- Preferably want to simulate still 'normalized grid' and also perform scan in g
- Testing shape changes of various distributions for different couplings
 - If no shape change we may just renormalize
- Simulated
 - $m_{DM}=10-200$ GeV (step size ~ 50 GeV with smaller binning for $m_{DM}<50$)
 - $m_{\Phi} = 10-1300$ GeV (step size ~ 100 , also smaller at low masses)



weights to normalise to $\langle\sigma v\rangle$

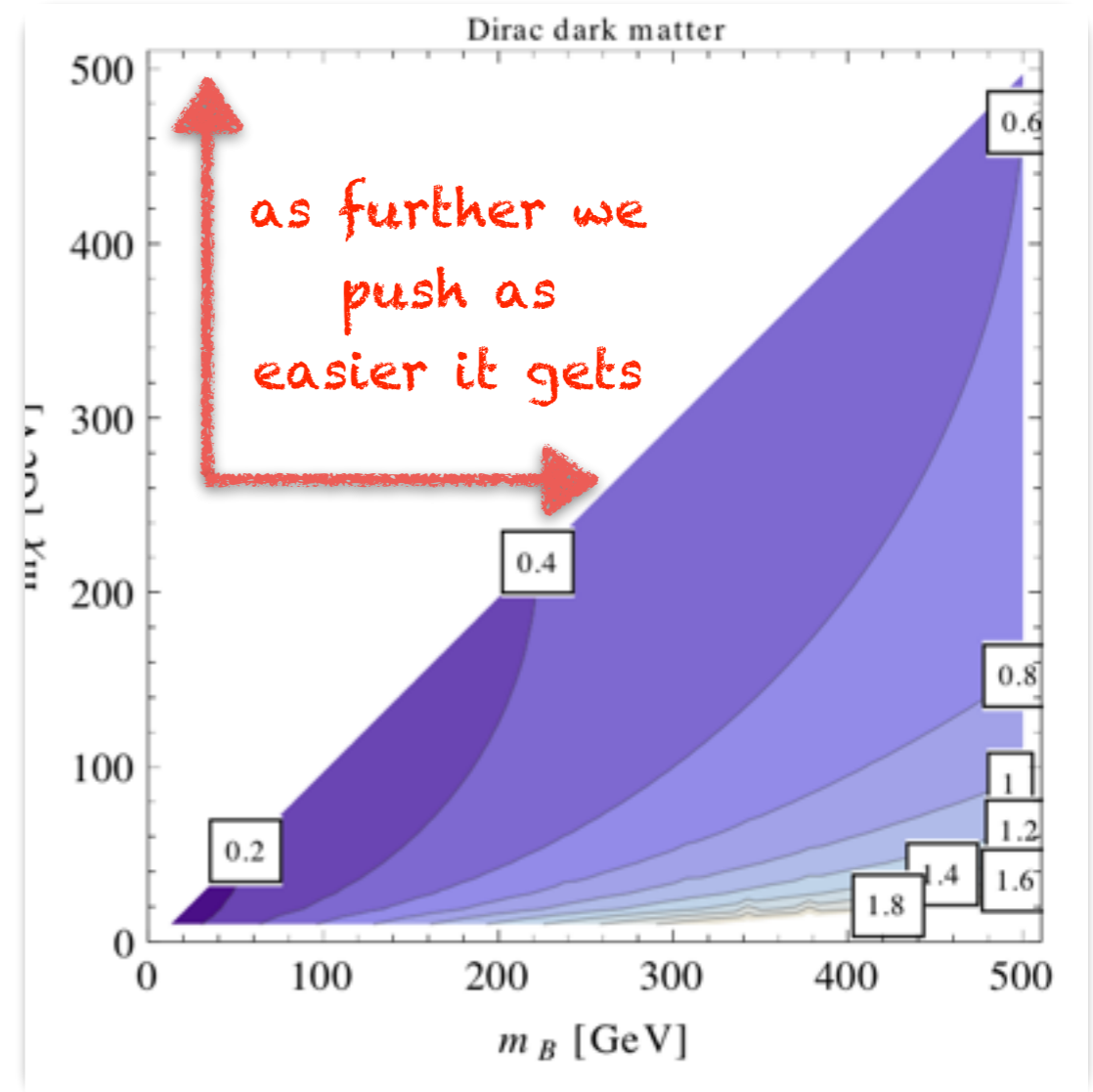
- Rather large xsec increase, reasons:

- b-quark in PDF



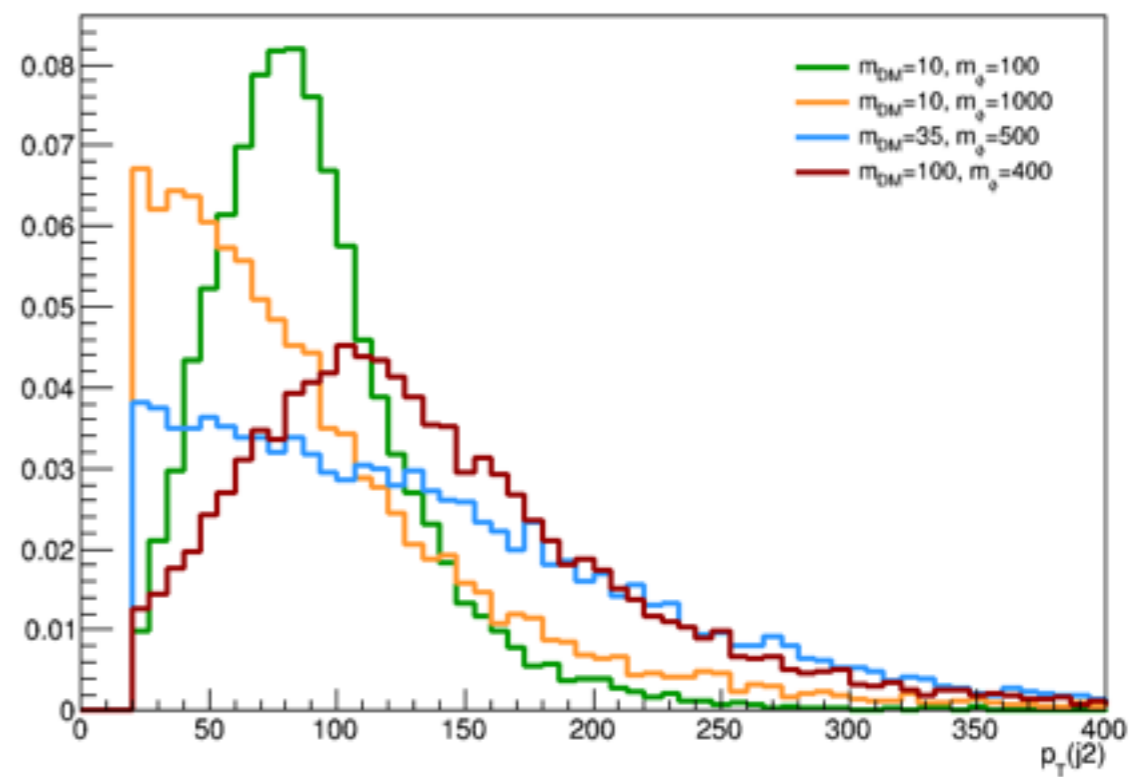
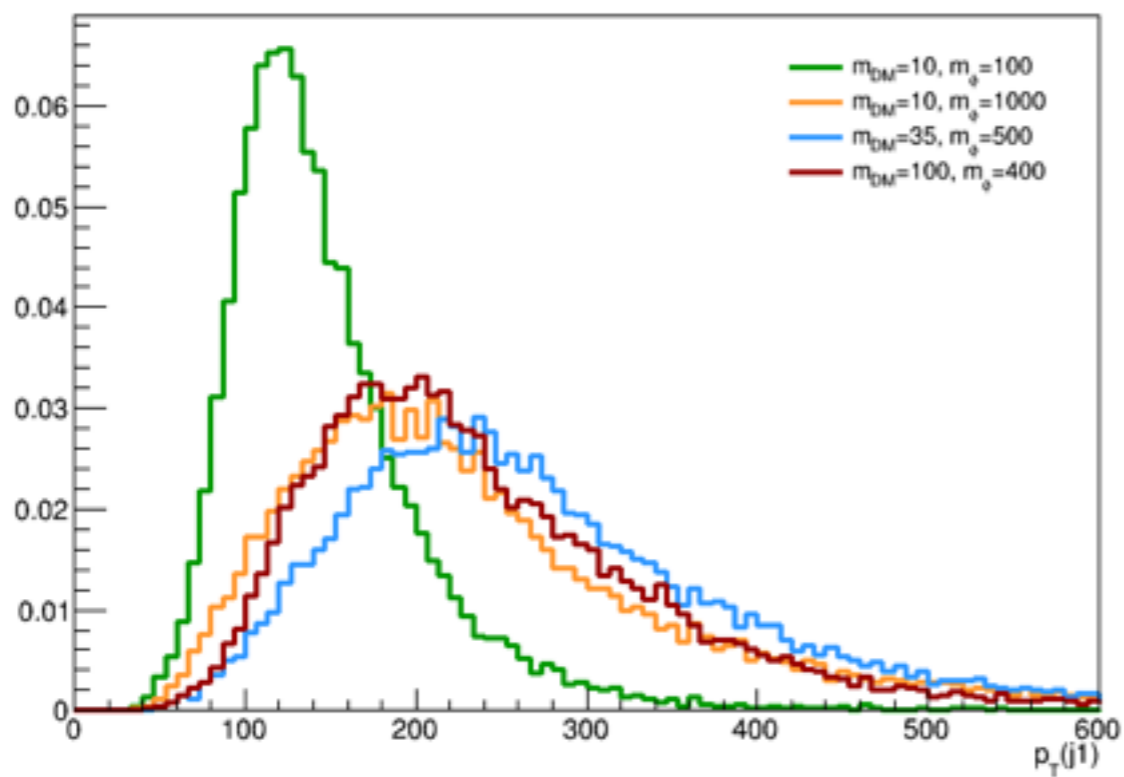
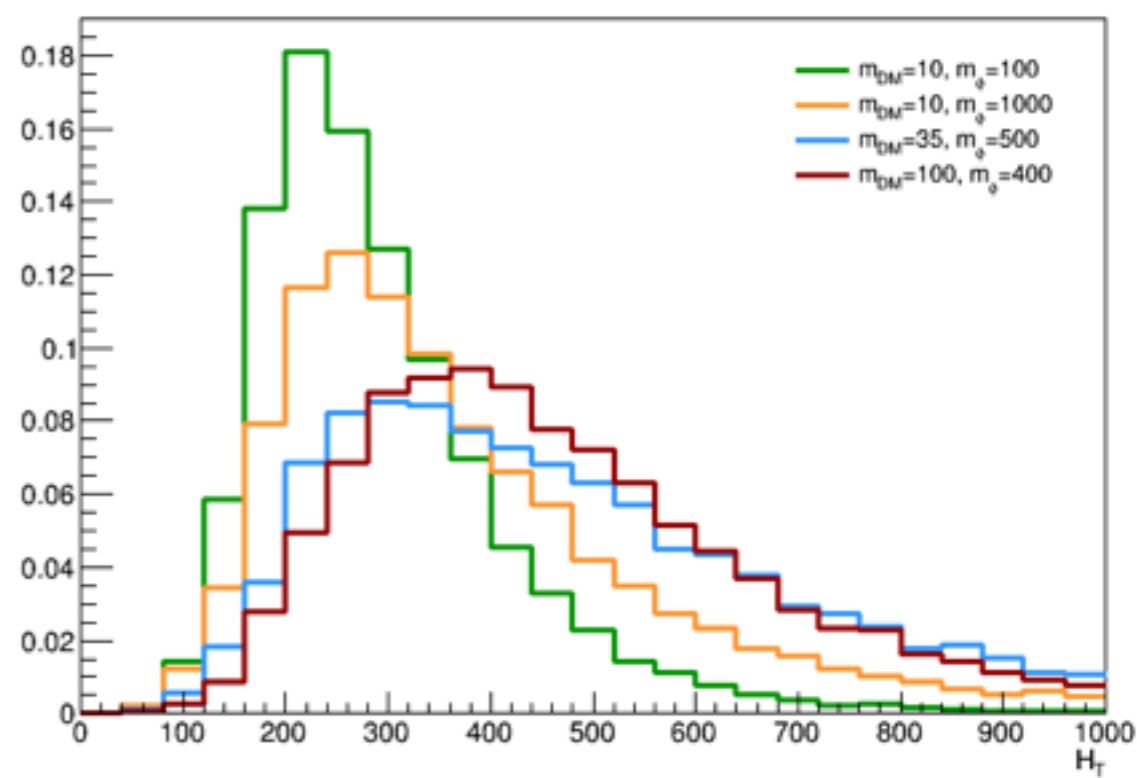
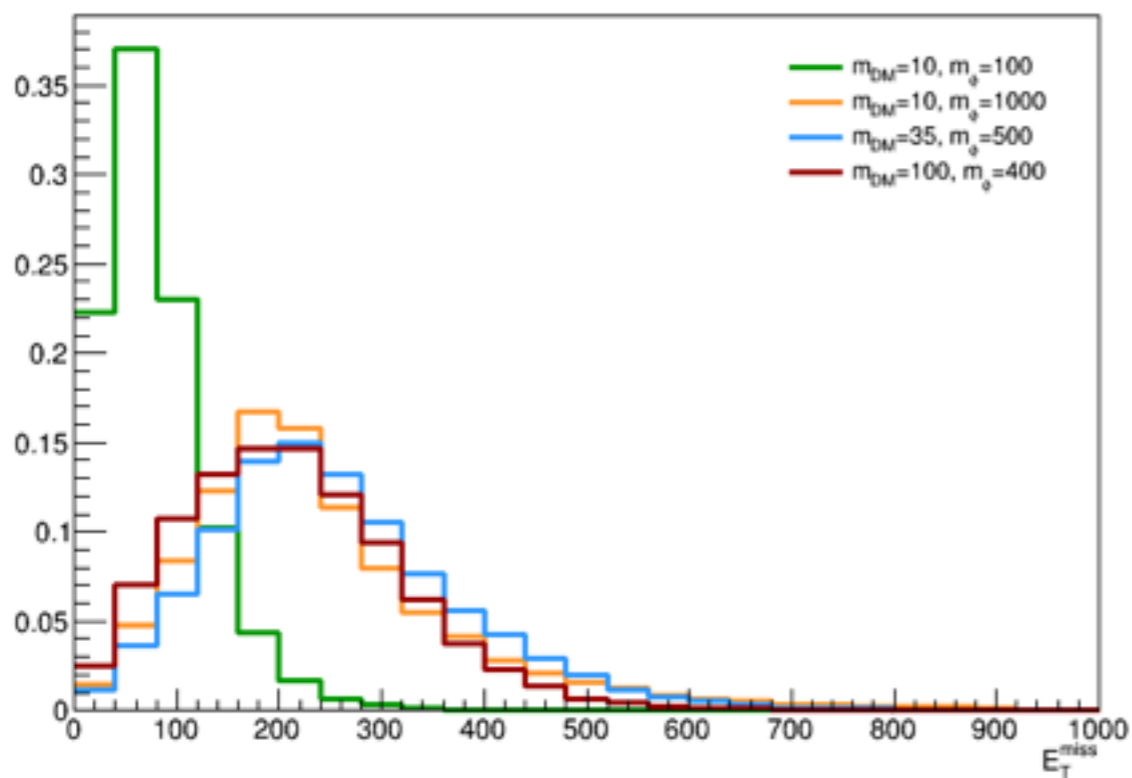
- Relic density weights much less severe

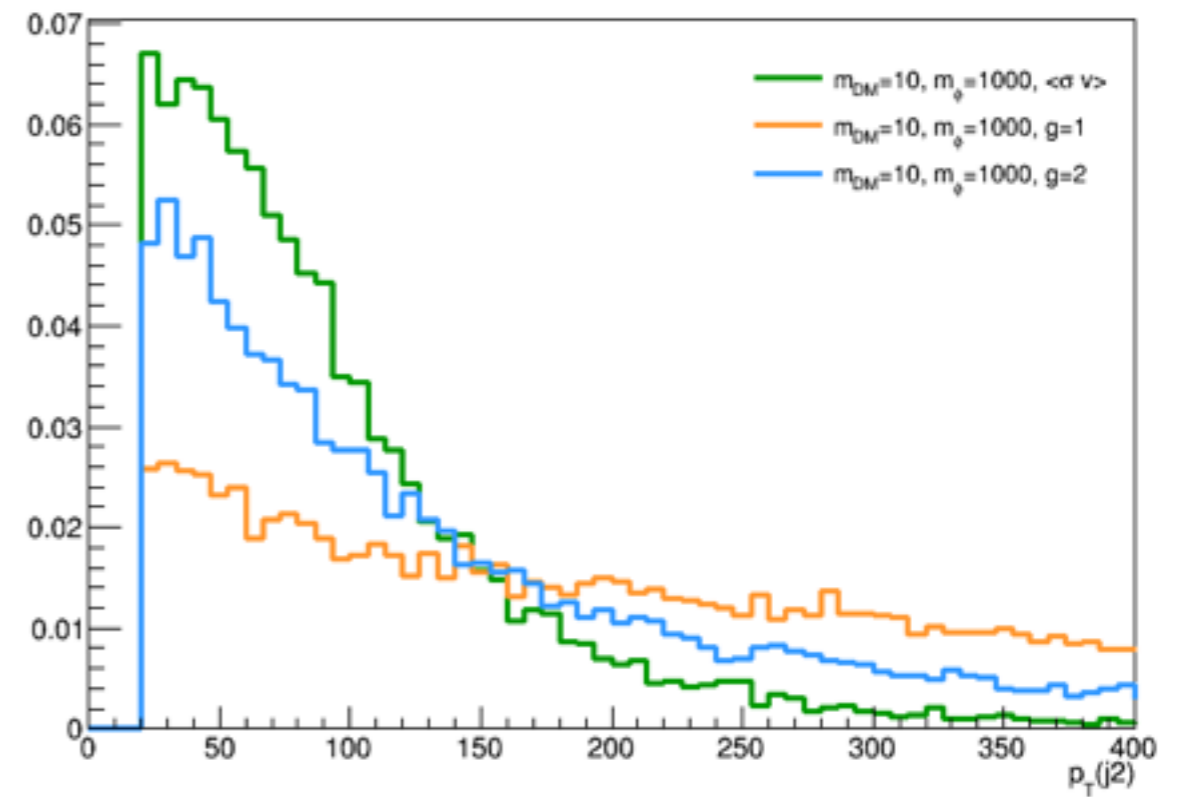
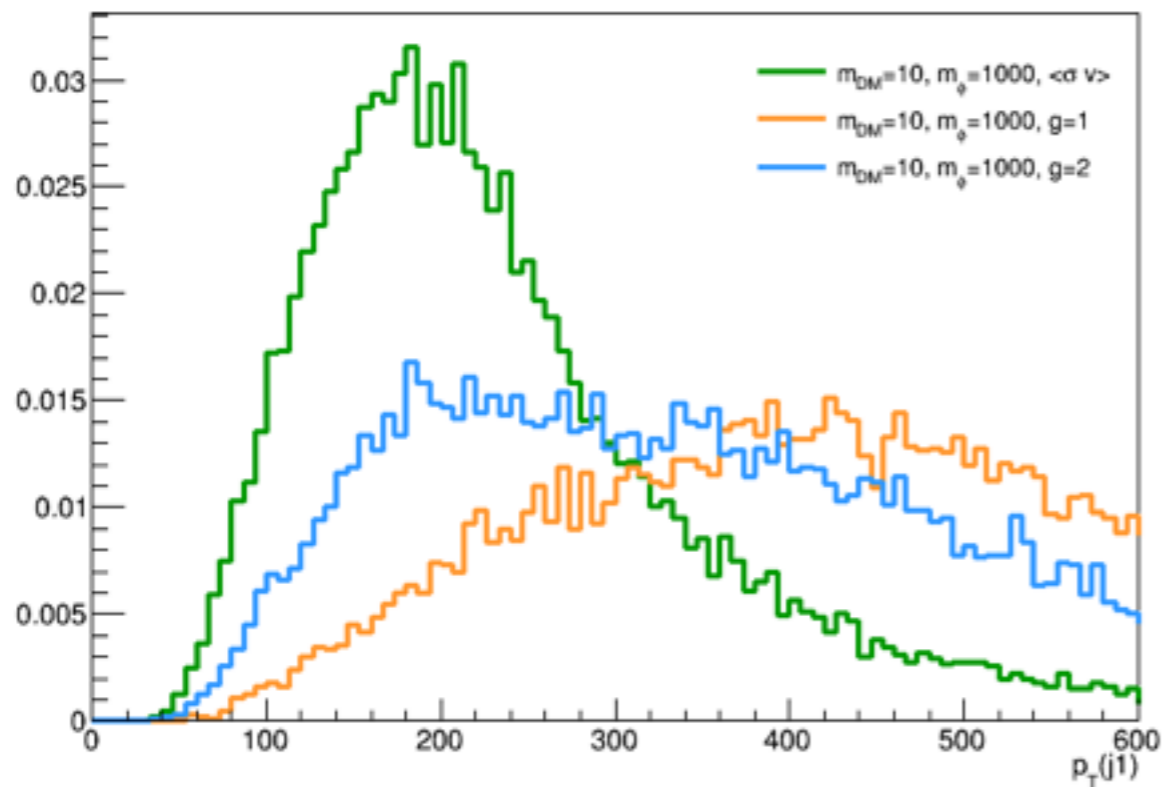
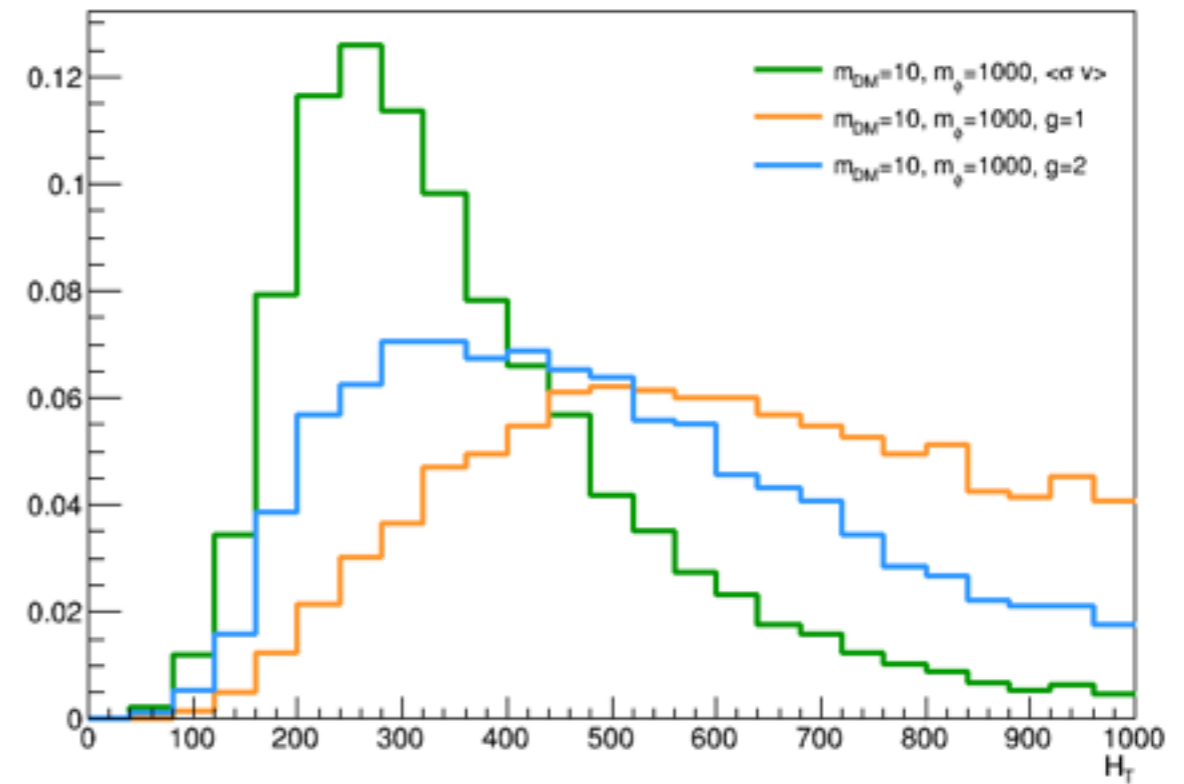
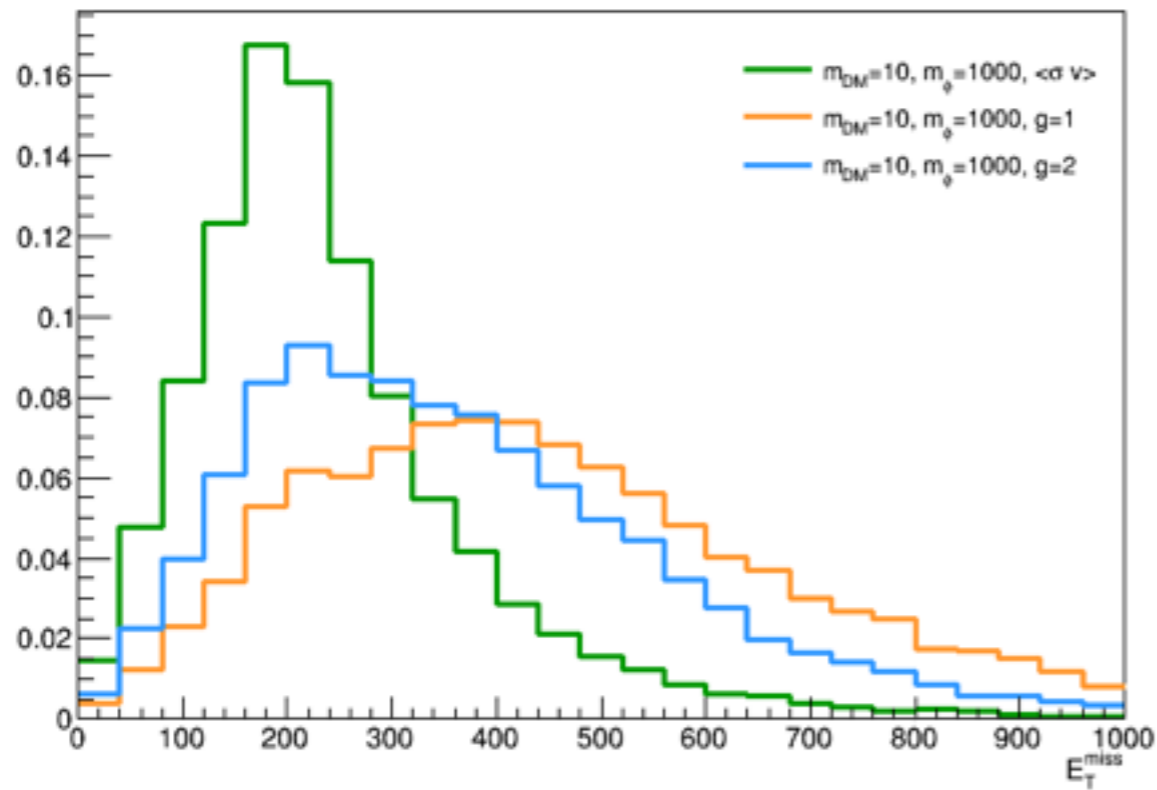
- Reproduce 8 TeV cross sections
- Studies ongoing to verify

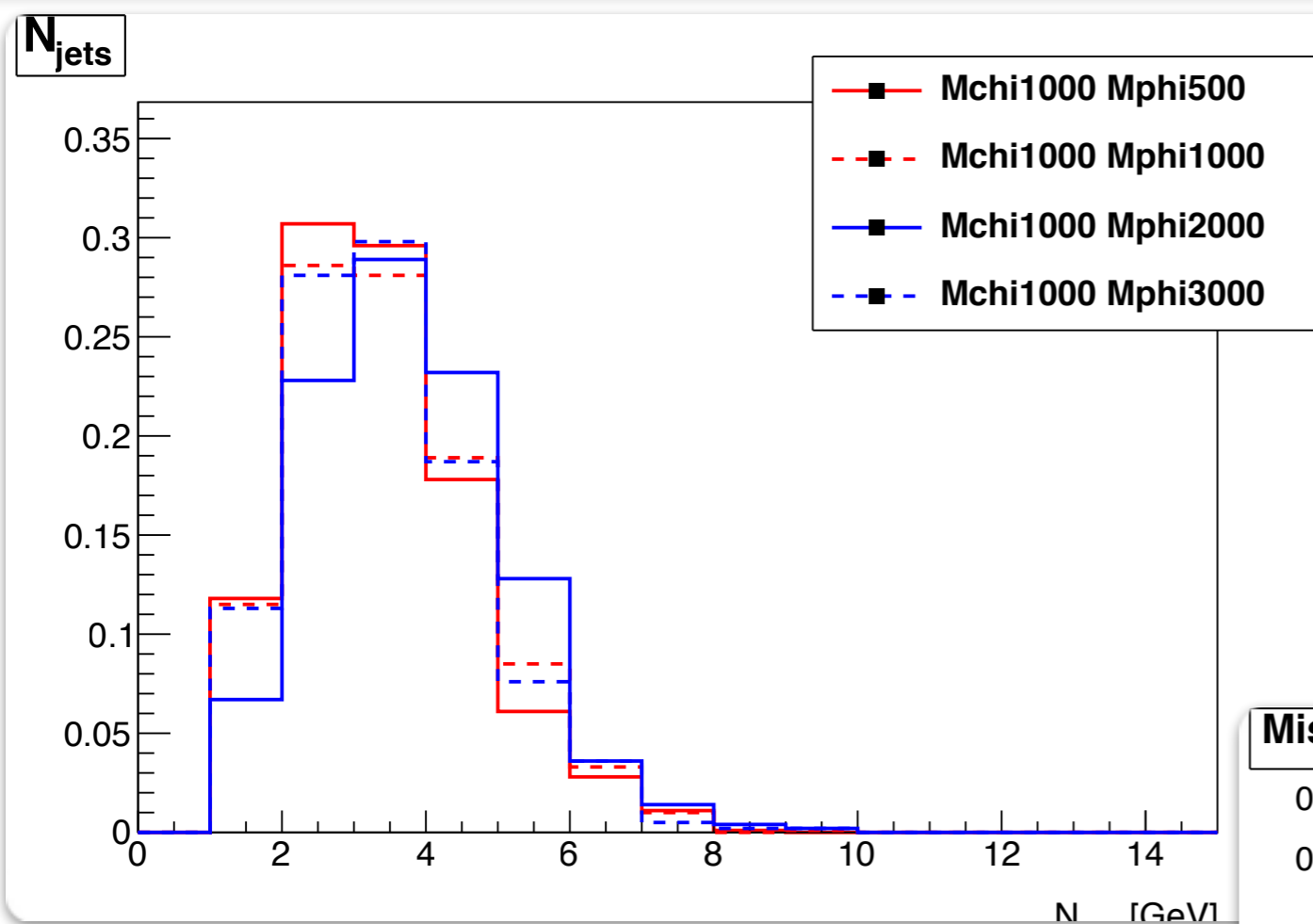


weight for normalizing to $\langle\sigma v\rangle$

$$\langle\sigma v\rangle \approx \frac{3\lambda^4 m_\chi^2 \sqrt{1 - (m_b/m_\chi)^2}}{32\pi (m_B^2 + m_\chi^2 - m_b^2)^2} \approx \frac{3\lambda^4 m_\chi^2}{32\pi (m_B^2 + m_\chi^2)^2}$$

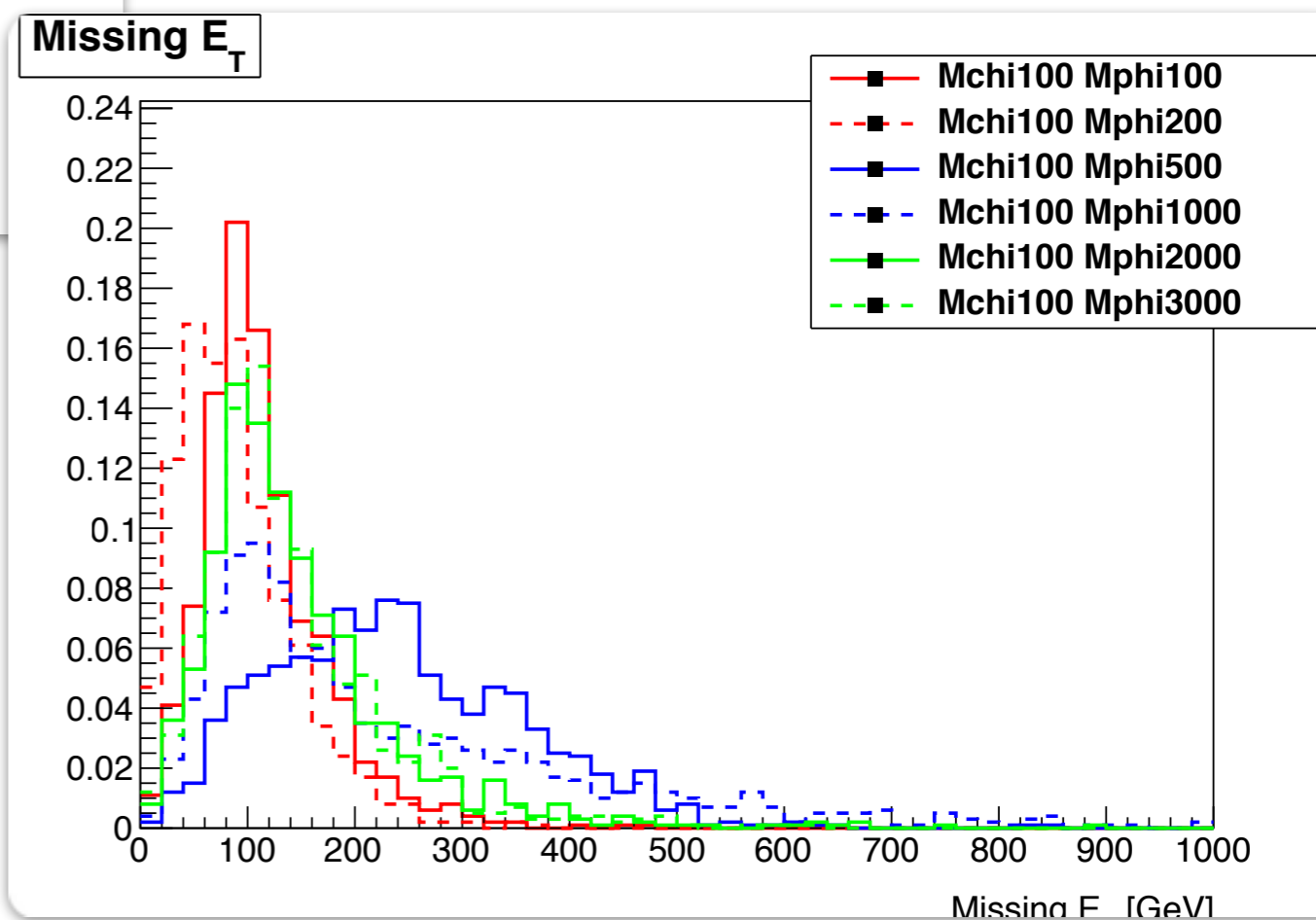






- Also Deborah produced distributions and her plots support prelim. conclusions

- Ongoing studies to compare setups with Priscilla
- We've been busy with DM+tt, DM +b and related responsibilities in individual collaborations





- Super secret (unfortunately an internal tool) indicates that we should generate DM masses up to ~ 300 GeV
- Remaining mass grid looks good and smooth

- Interesting model to probe unique phase space
- Large cross section increase makes such models attractive for early data
- Shapes dependent on g in m_{DM} and m_{Φ} observed
 - Need to generate all couplings to be probed individually
- Sensible to produce samples for coarse grid in $g=1,2..$ and one set normalised to relic density
- In some sense more constrained model, easier to define parameter space

Backup
Just some more distributions

