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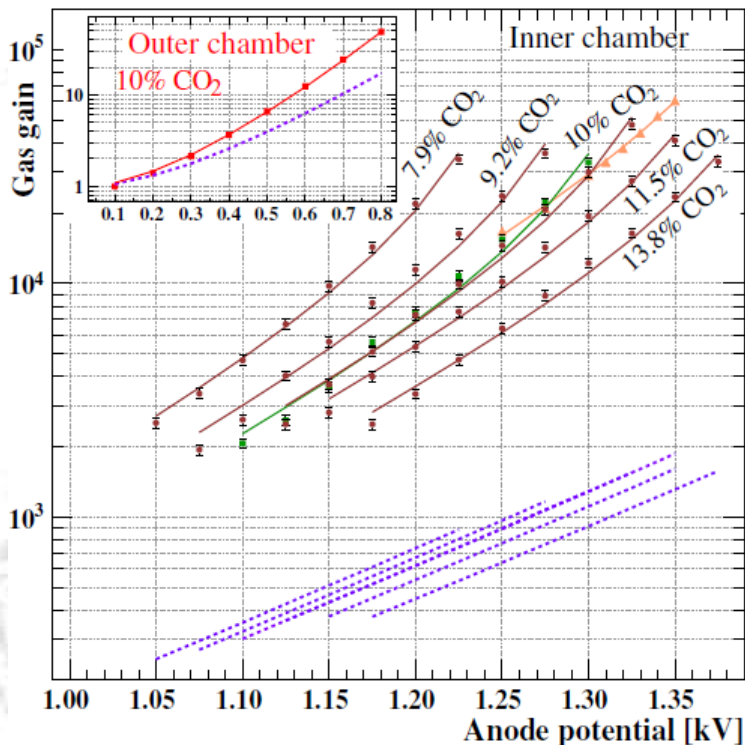
Status of gas gain measurements and calculations in Ne-CO₂ mixtures

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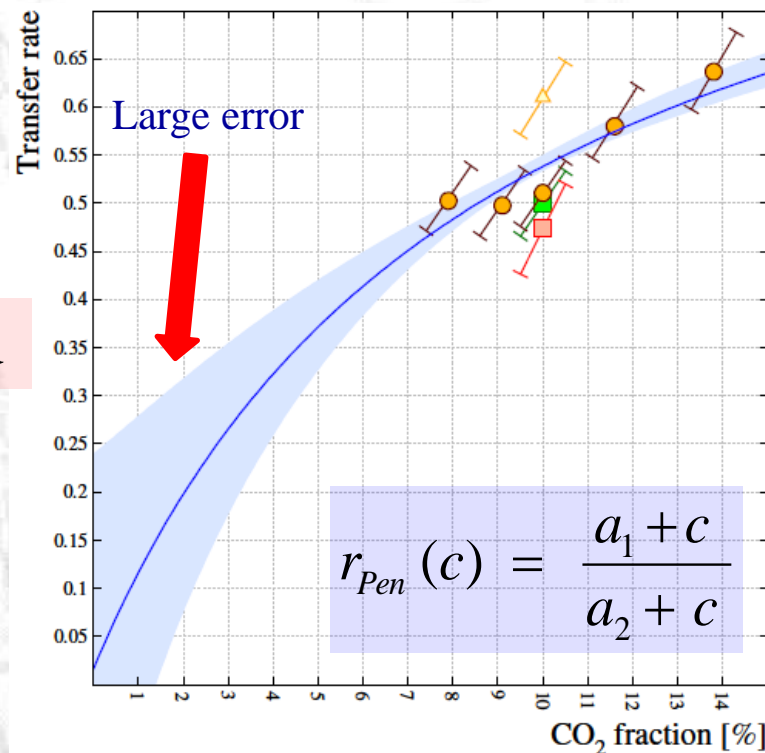
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ALICE TPC (earlier measurements and calculations)



Unpublished



- ❖ Measurements: C. Garabatos and D. Vranic
- ❖ Pressure: 990-1010 mbar
- ❖ Gain: limited range $2 \times 10^3 - 5 \times 10^4$
- ❖ CO₂ fraction: 7.9 % - 13.8 %
- ❖ Uncertainty: CO₂ ± 0.5 % (absolute)

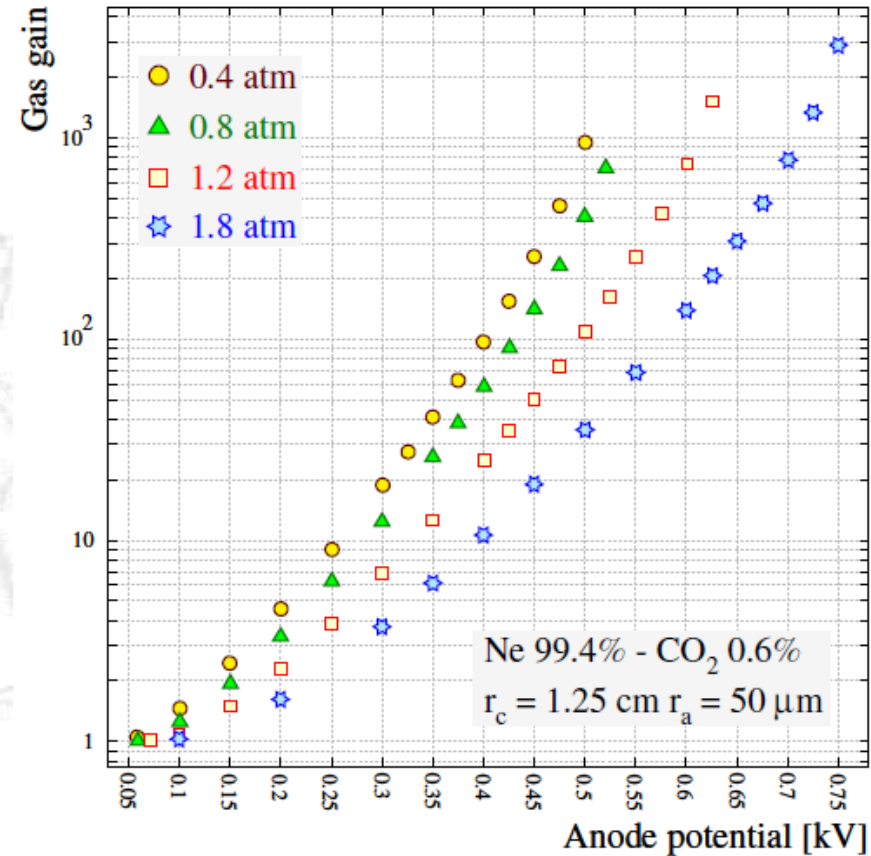


- ❖ Difficult to use a comprehensive transfer fit
- ❖ Radiative transfer rate: $a_1 / a_2 = 0.01 \pm 0.22$
 - ❖ no measurements at low CO₂ fractions
- ❖ **Experimental deficiencies prevent to achieve accurate information about the processes involving Penning transfer.**

Ne – CO₂ gas gain measurements at Krakow

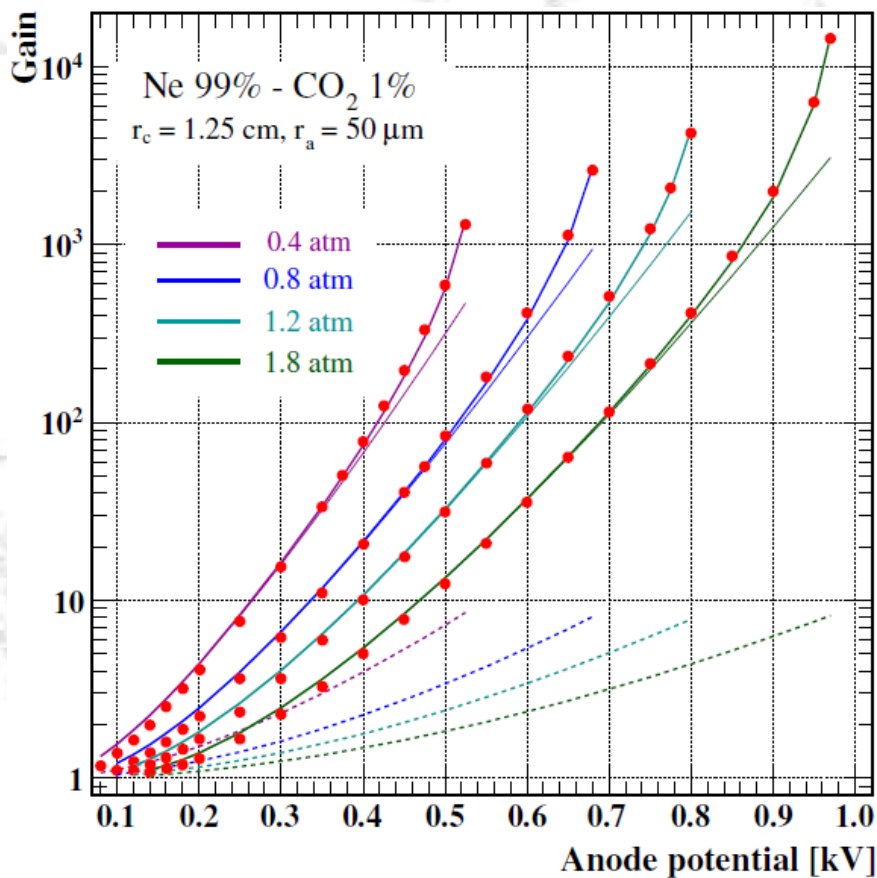
CO₂ percentage

- | | |
|----------------|--------------------------------|
| 1) 0.6% | 8) 15% |
| 2) 1% | 9) 20% |
| 3) 2% | 10) 30% |
| 4) 4% | 11) 40% |
| 5) 5% | 12) 50% |
| 6) 7% | 13) 60% |
| 7) 10% | 14) Pure CO₂ |



- ❖ Gas gains: measured by Tadeusz KOWALSKI,
- ❖ Single wire proportional counters: $r_c = 1.25$ cm, $r_a = 24$ μ m or $r_a = 50$ μ m,
- ❖ Wide gain regime: ionisation to higher than 10^5 ; less than 5% error on gas gain,
- ❖ Pressure range: 0.4 – 1.8 atm; in addition 0.25 atm for a few mixtures.

Ne – CO₂ gas gain calculations



Penning correction

- ❖ $\text{Ne}^* + \text{CO}_2 \rightarrow \text{Ne} + \text{CO}_2^+ + e^-$
- ❖ All of the excited Ne atoms can ionise CO₂

$$\alpha_{\text{Penning}} = \alpha \frac{\sum v_i^{\text{ion}} + \sum r_i v_i^{\text{exc}}}{\sum v_i^{\text{ion}}}$$

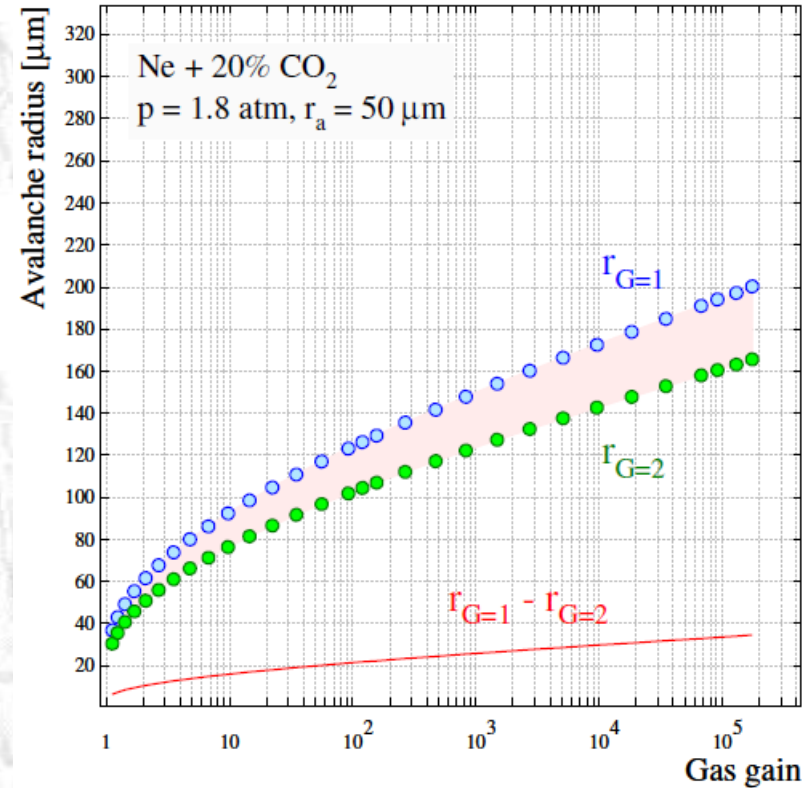
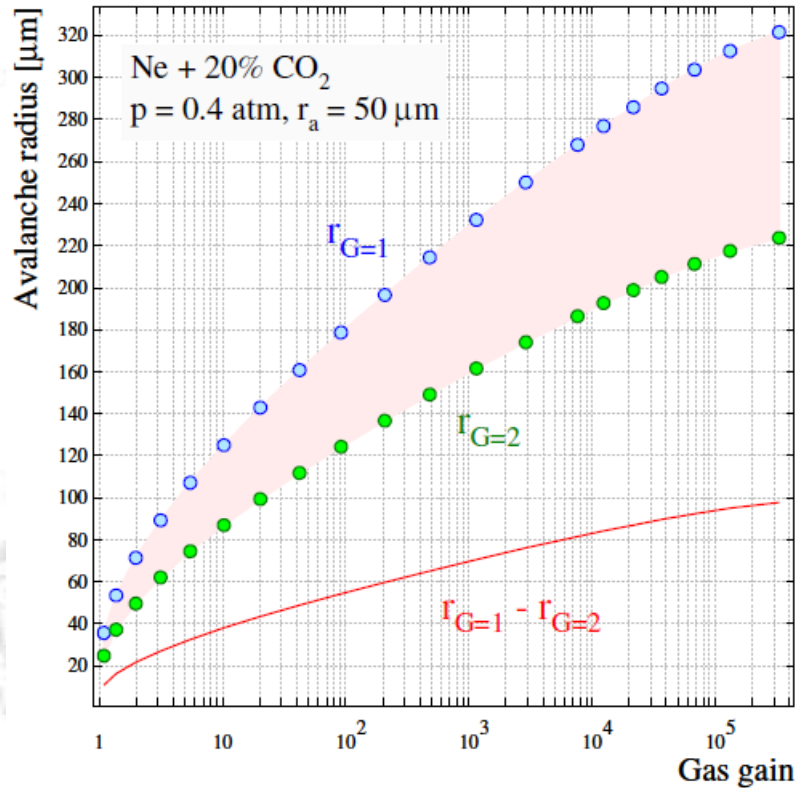
Photon feedback

$$G' = G / (1 - \beta G)$$

Production frequencies of the ionisations and excitations with **Magboltz 10.10**

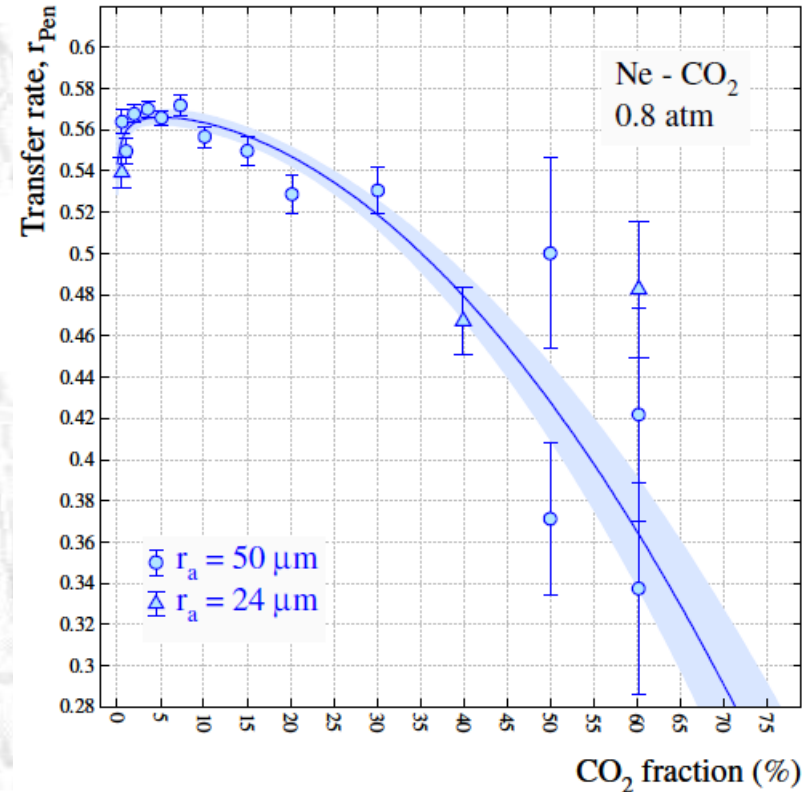
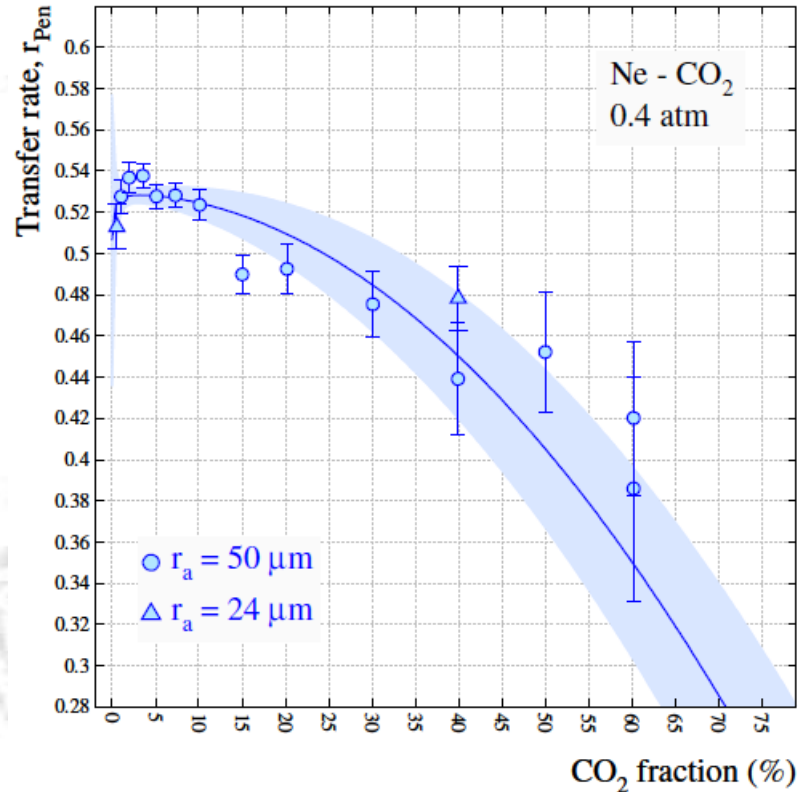
- ❖ Dashed lines: **without** corrections (Penning, feedback),
- ❖ Thin lines: with Penning, **without** feedback corrections,
- ❖ Thick lines: final fits **with** Penning and feedback corrections.

Avalanche sizes



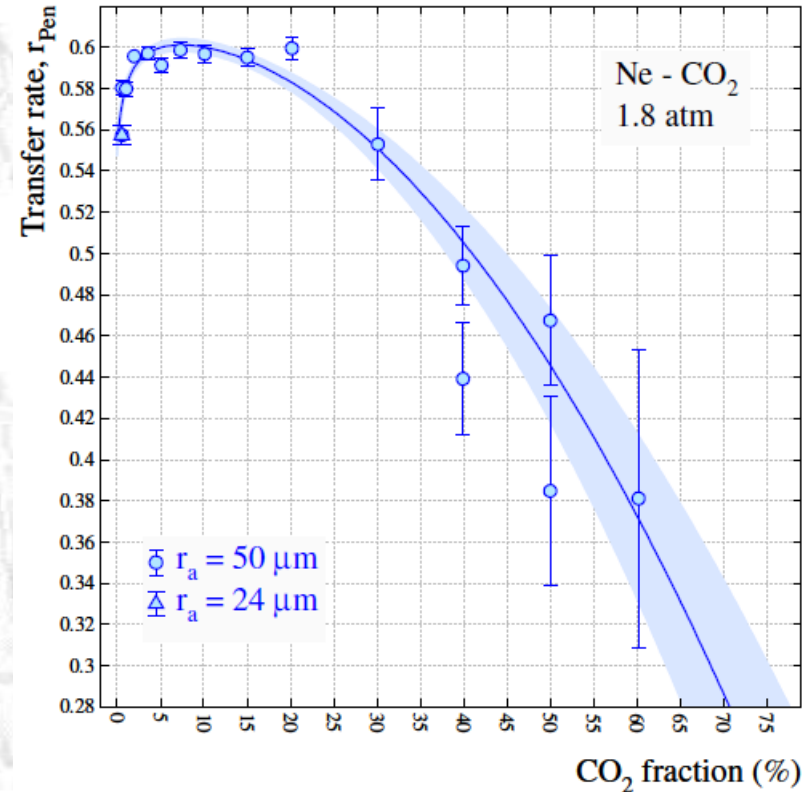
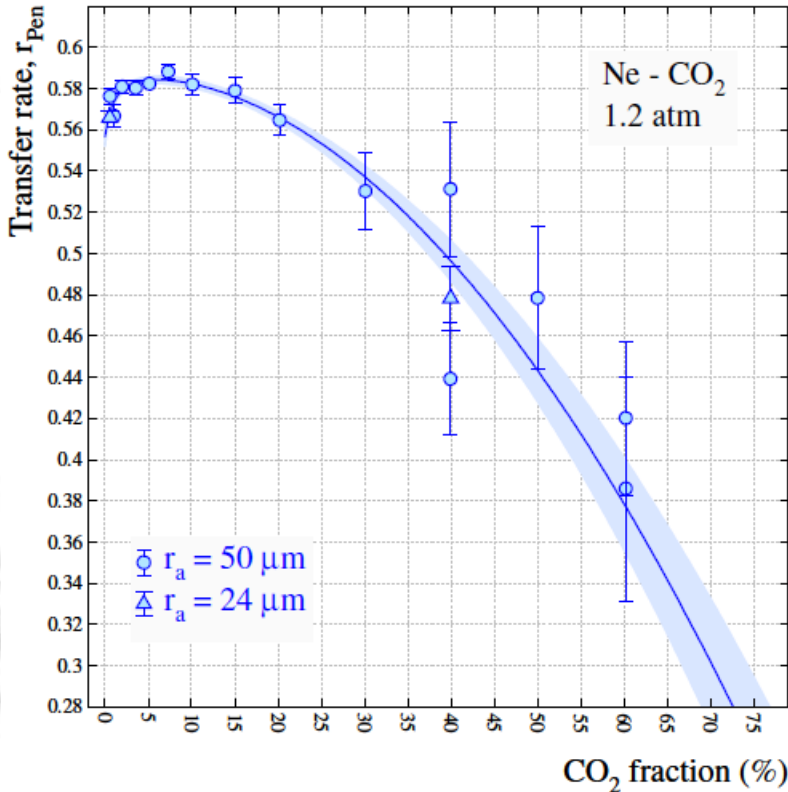
- ❖ The circles $r_{G=1}$, $r_{G=2}$: radius from G=1 and G=2 to the highest measured gas gains
- ❖ Pink bands: intermediate zone between the multiplication and drift region,
 - ❖ electric field strength become radially smaller with the depth of $r_{G=1} - r_{G=2}$
 - ❖ production of the excitations and ionisations compete each other
- ❖ **Less Penning transfer at lower pressure** in the same mixture (will see next slides)

Transfer rates



- ❖ Transfer rates first increase at low CO₂ concentrations and reach a peak
 - ❖ more visible at high pressures
- ❖ After peak, continues decrease of the transfer rates with increasing CO₂ fractions
- ❖ Large uncertainties at high CO₂ fractions due to lack of data
 - ❖ measurements beyond 60% CO₂ on the way

Transfer rates (fit function)



$$r_{Pen}(c) = \frac{a_1 c + a_3}{c + a_2} - a_4 c^2$$

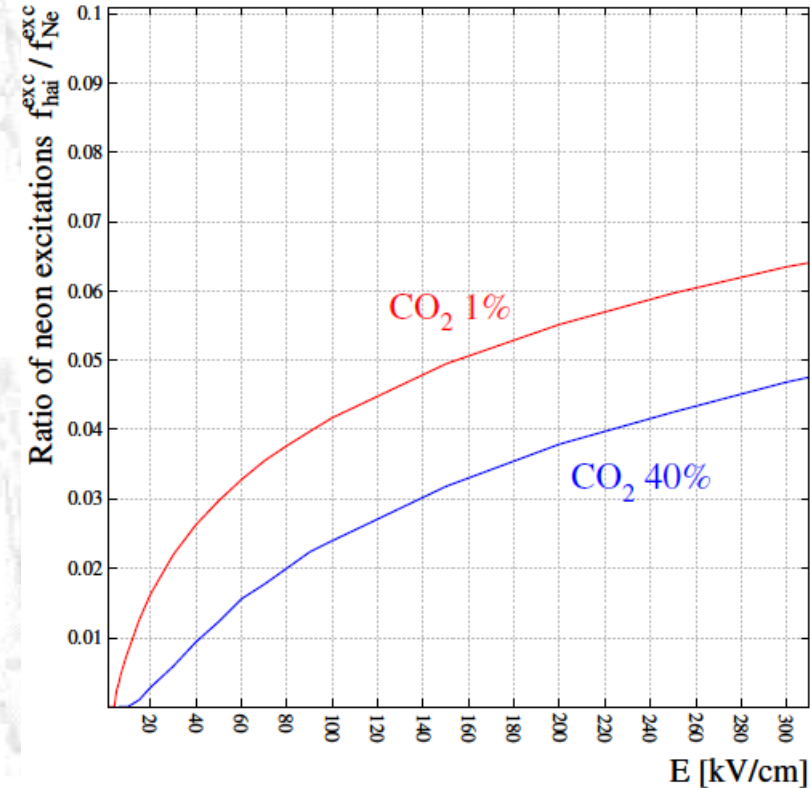
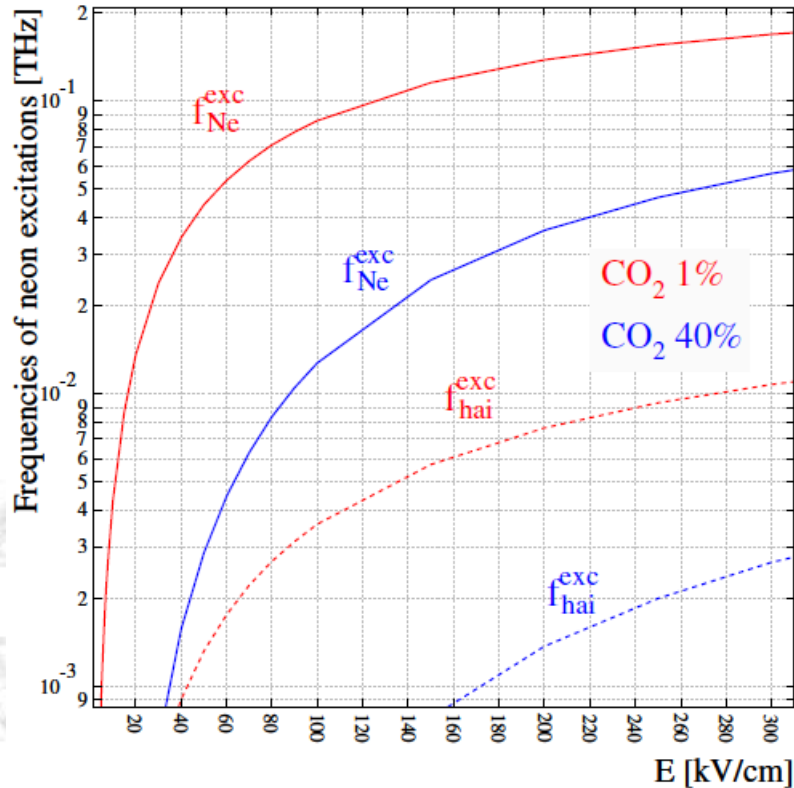
- ❖ $\mathbf{a_4}$: reduction parameter of the rates
- ❖ $\mathbf{c^2}$ dependence of the fit function points
- three – body interactions**

Fitting parameters

Pressure [atm]	a_1	a_2	a_3/a_2	a_4
0.4	0.5298 ± 0.0057	0.0016 ± 0.0002	0.5064 ± 0.0707	0.4981 ± 0.0806
0.8	0.5710 ± 0.0050	0.0074 ± 0.0002	0.5428 ± 0.0147	0.5710 ± 0.0767
1.2	0.5916 ± 0.0026	0.0108 ± 0.0001	0.5564 ± 0.0044	0.5916 ± 0.0646
1.8	0.6133 ± 0.0047	0.0126 ± 0.0002	0.5543 ± 0.0070	0.6133 ± 0.1169

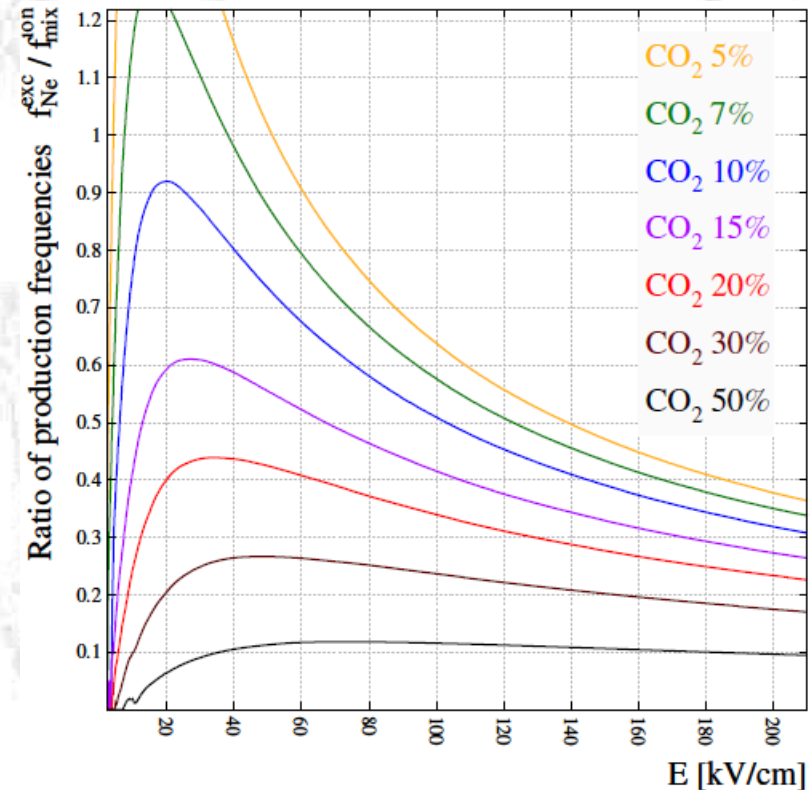
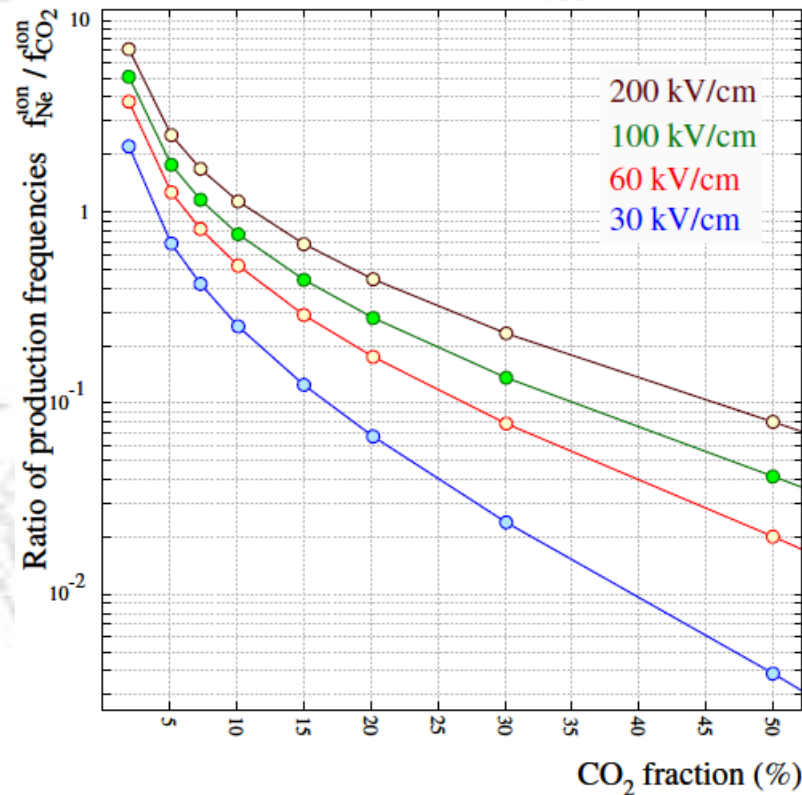
- ❖ a_1 : asymptotic values of the transfer rates,
 - ❖ smaller than 1: energy losses with inelastic collisions; the number of rotation, vibration and polyad modes of CO₂ molecules,
 - ❖ such losses become unlikely to occur at high pressures since a_1 increase
- ❖ a_2 : collisional energy transfer probability, $\text{Ne}^* + \text{CO}_2 \rightarrow \text{Ne} + \text{CO}_2^+ + e^-$
 - ❖ the time between collisions shorter at higher pressure
- ❖ a_3/a_2 : radiative transfer and homonuclear associative ionisation of neon (h.a.i.)
 - ❖ $\gamma + \text{CO}_2 \rightarrow \text{CO}_2^+ + e^-$
 - ❖ $\text{Ne}^* + \text{Ne} \rightarrow \text{Ne}_2^+ + e^-$

Contribution of the h.a.i processes to the a_3/a_2 parameter



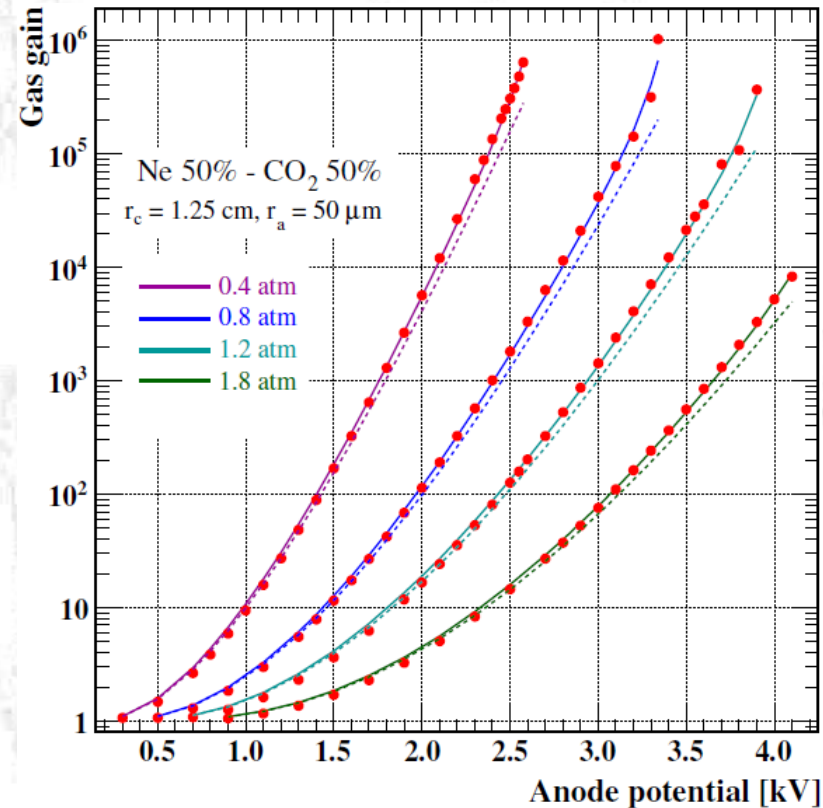
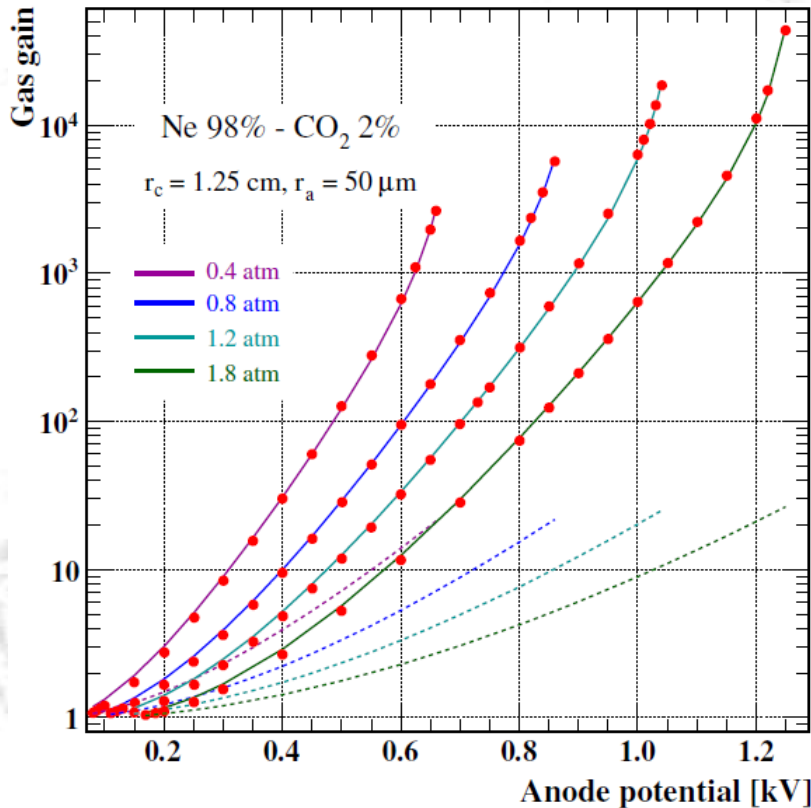
- ❖ h.a.i. states are less frequently produced (threshold is high 20.86 eV)
- ❖ even if all h.a.i states take place in multiplications, their contributions can not excess several percentages,
- ❖ the contribution will be less at high CO2 concentrations,
- ❖ **radiative transfer** remains as the **dominant** process in a_3/a_2 paraneter

Understanding the transfer drops, parameter a_4 (production frequencies)



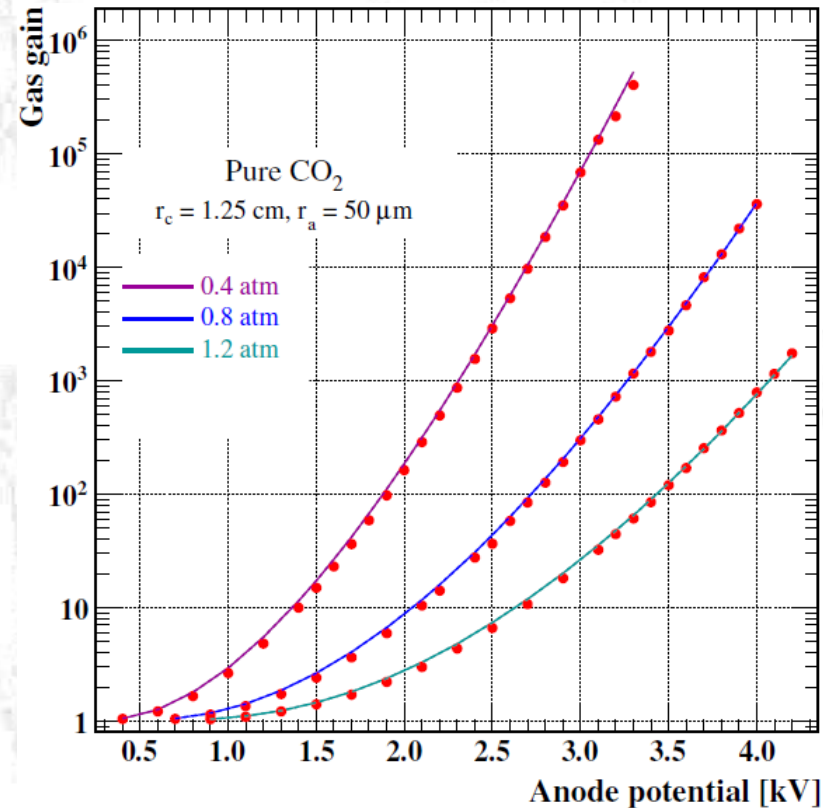
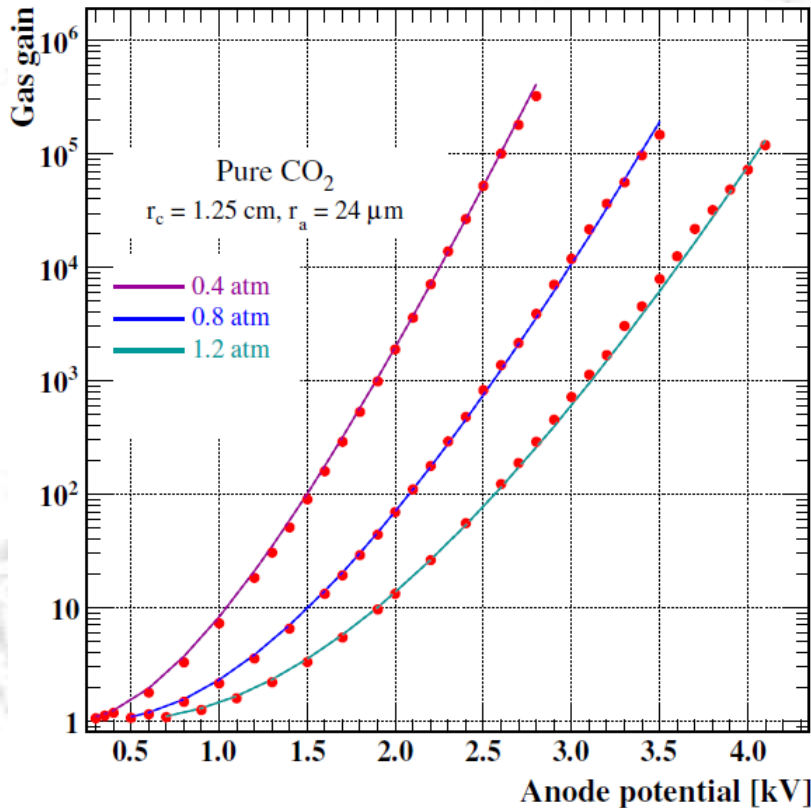
- ❖ the largest part of the gain comes from CO_2 ionisations,
- ❖ ionisation potentials: CO_2^+ 13.78 eV, Ne^+ 21.56 eV
- ❖ fraction of neon excitations become smaller than unity beyond 10% CO_2 ,
- ❖ Penning impact on gain enhancements will be smaller at high CO_2 percentages.

Effect of the production rates



- ❖ decrease on Penning impact does not mean smaller Penning transfer probability,
- ❖ even if the population of the excited neon atoms diminish at high CO₂ fractions, they can still efficiently transfer to ionize CO₂ molecules,
- ❖ transfer drops can not be explained in terms of production rates.

Gain calculations in pure CO₂



- ❖ Are the CO₂ ionisation cross sections are wrong in Magboltz ???
- ❖ Calculations for the pure CO₂ mixtures confirms that Magboltz use perfectly correct cross sections
- ❖ There should really be other physical processes leading the transfer rate drops

Proposed mechanisms for the transfer drops

❖ Neon excimers (three-body reaction):



CO_2^* can dissociate to form neutral fragments like CO, O and C (3)

❖ at the same pressure efficiency of (1) and thereby (2) will decrease with increasing fraction of CO_2 so, neon excimer formation is working inverse way of transfer reductions; not correct answer.

❖ Other three-body reactions:



❖ (4) and (5) can dissociate through; $\text{NeCO}_2^* \rightarrow \text{Ne} + \text{C} + \text{O}_2$ (6)



The mechanisms (4) and (5) are dominated by partial pressure of CO_2 ; **formation of the NeCO_2^* complex is the approach to explain energy transfer drops !!!**



Thanks and ????