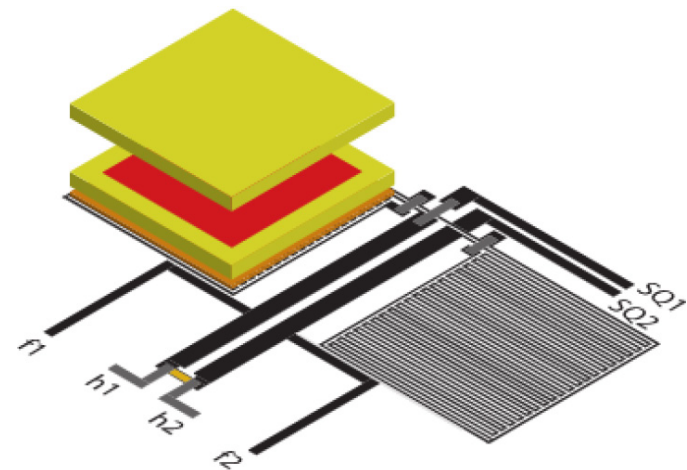




# Electron Capture in $^{163}\text{Ho}$ experiment

Loredana Gastaldo  
for the ECHO Collaboration

Kirchhoff Institute for Physics, Heidelberg University



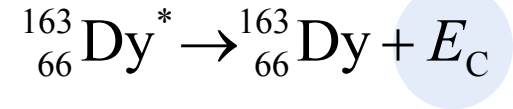
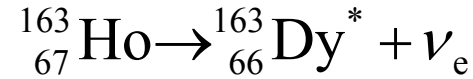
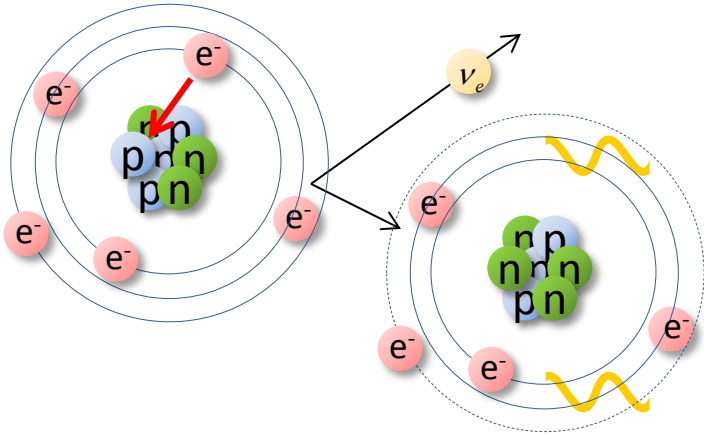
# Contents

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- $^{163}\text{Ho}$  and neutrino mass
- The ECHo neutrino mass experiment
- Conclusions



# $^{163}\text{Ho}$ and neutrino mass



- $\tau_{1/2} \cong 4570 \text{ years}$  ( $2 \cdot 10^{11}$  atoms for 1 Bq)
- $Q_{\text{EC}} = (2.555 \pm 0.016) \text{ keV}^*$   
Experimental range 2.3 keV – 2.8 keV

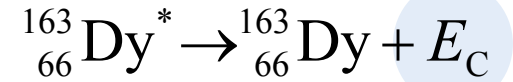
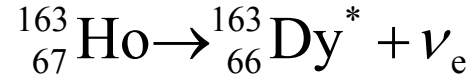
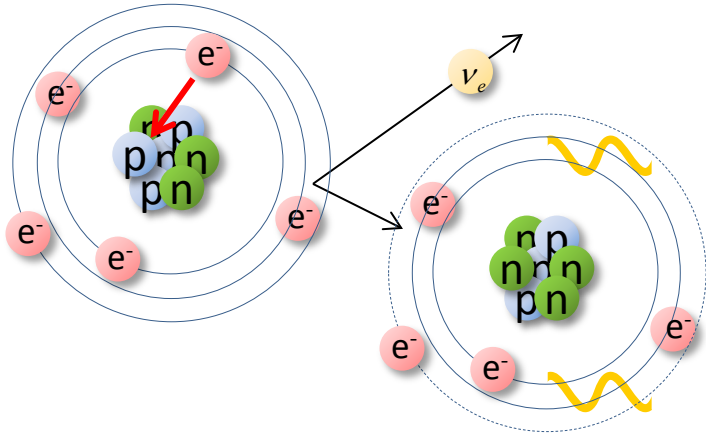
A non-zero neutrino mass affects  
the de-excitation energy spectrum

Atomic de-excitation:

- X-ray emission
- Auger electrons
- Coster-Kronig transitions

} Calorimetric  
measurement

# $^{163}\text{Ho}$ and neutrino mass



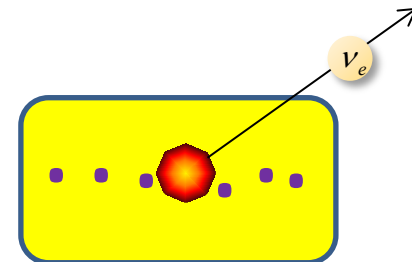
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Atomic de-excitation:

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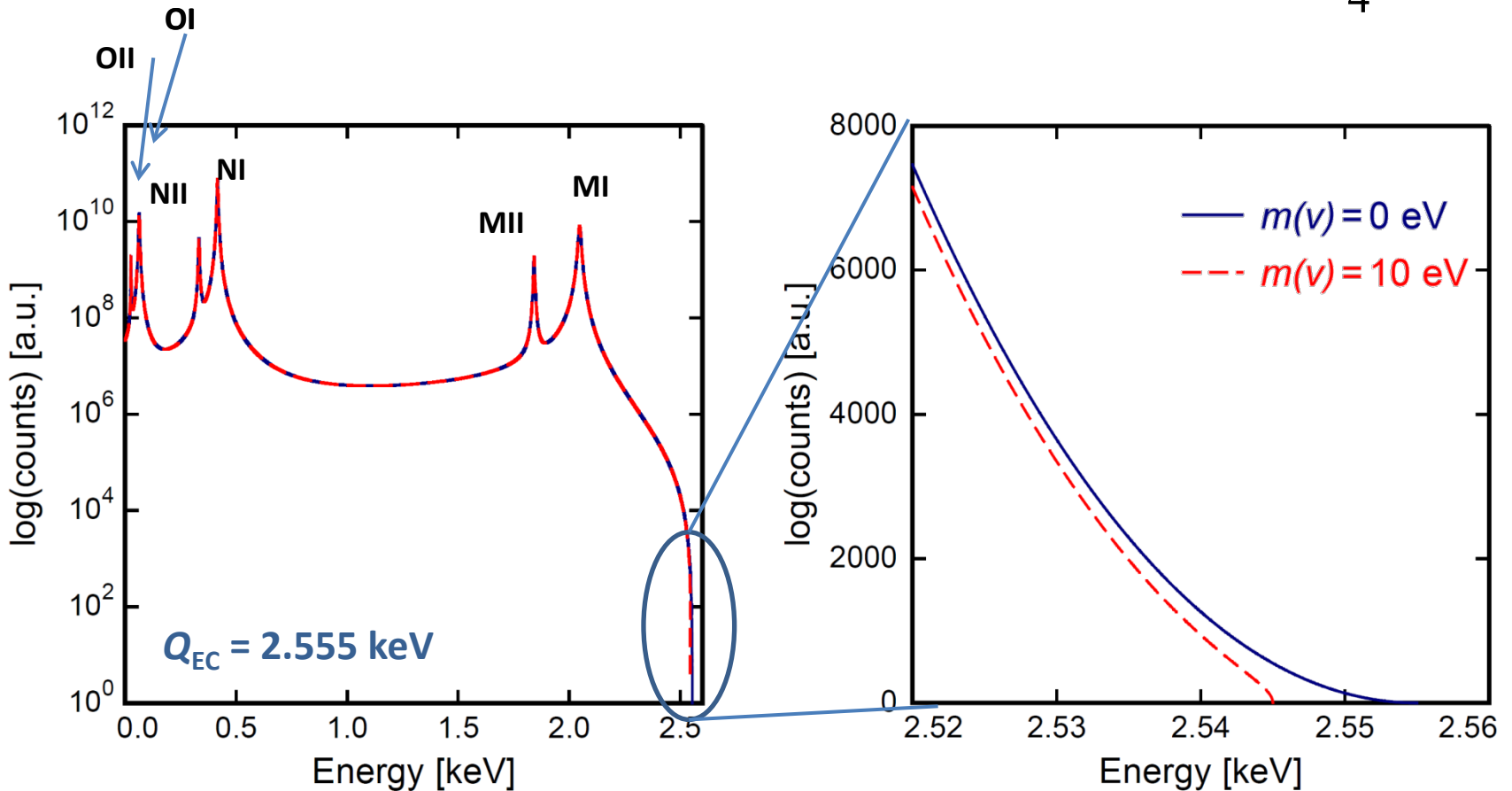
Calorimetric measurement



\*M. Wang, G. Audi et al., *Chinese Phys. C* **36**, 1603, (2012)

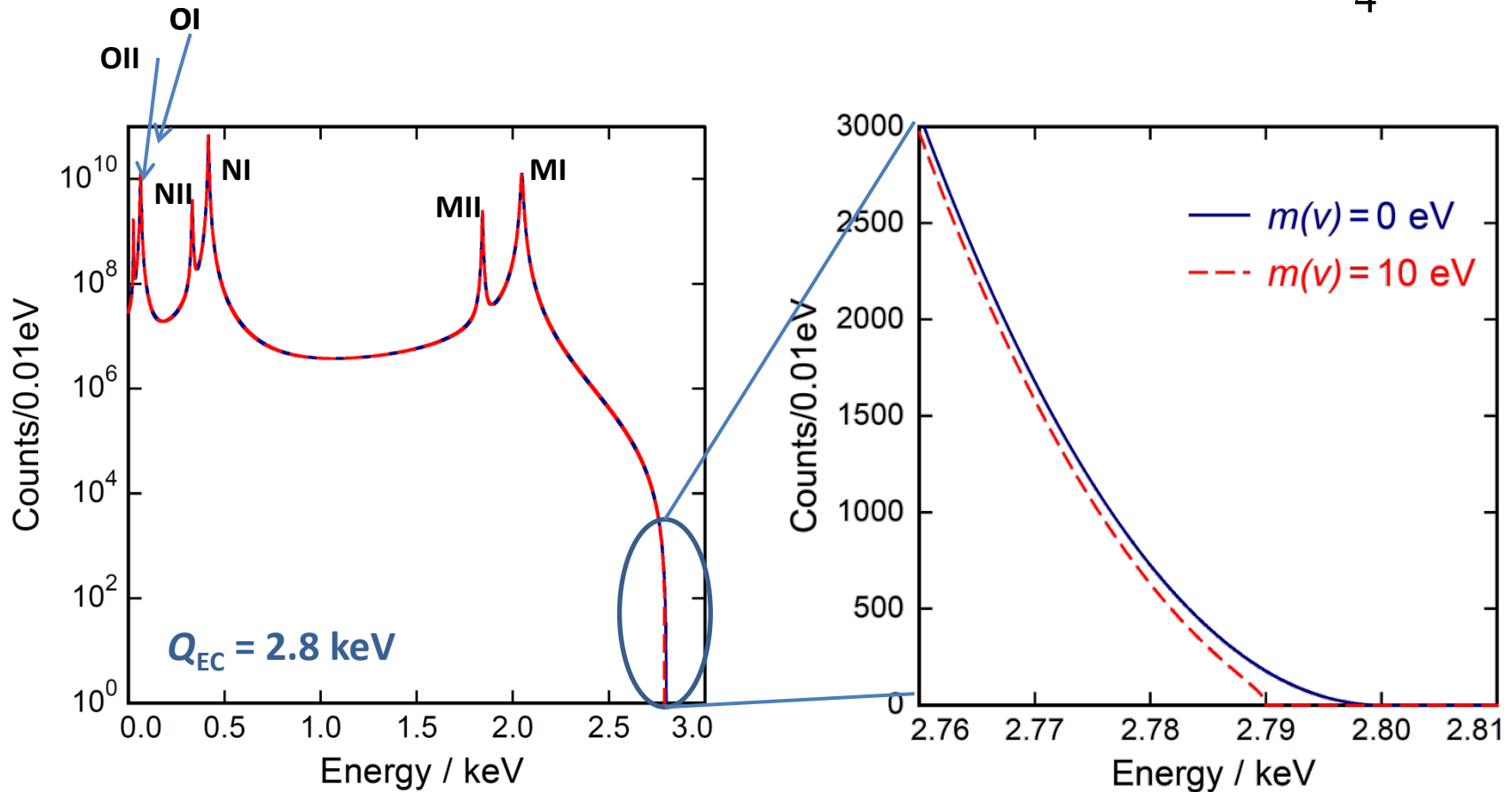
# $^{163}\text{Ho}$ and neutrino mass

$$\frac{dW}{dE_C} = A(Q_{\text{EC}} - E_C)^2 \sqrt{1 - \frac{m_\nu^2}{(Q_{\text{EC}} - E_C)^2}} \sum_H B_H \phi_H^2(0) \frac{\frac{\Gamma_H}{2\pi}}{(E_C - E_H)^2 + \frac{\Gamma_H^2}{4}}$$



# $^{163}\text{Ho}$ and neutrino mass

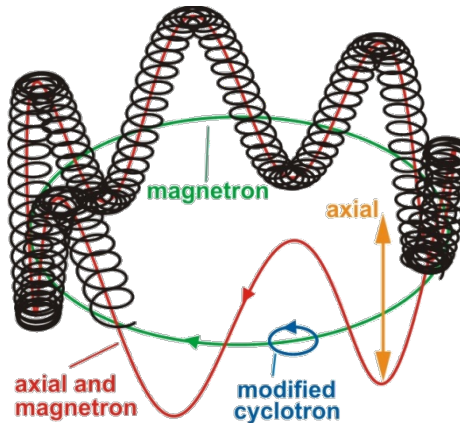
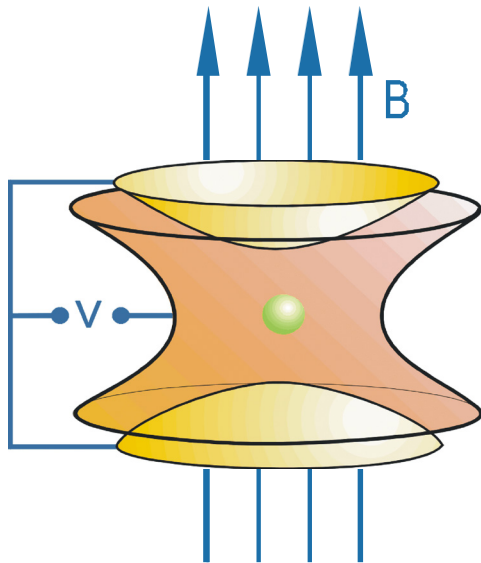
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# $Q_{EC}$ determination of $^{163}\text{Ho}$

## Penning Trap mass spectrometry

- Precise measurement of the  $^{163}\text{Ho}$  and  $^{163}\text{Dy}$  atomic mass



$$v_c^2 = v_+^2 + v_-^2 + v_z^2$$

Talk by Klaus Blaum yesterday

# $Q_{EC}$ determination of $^{163}\text{Ho}$

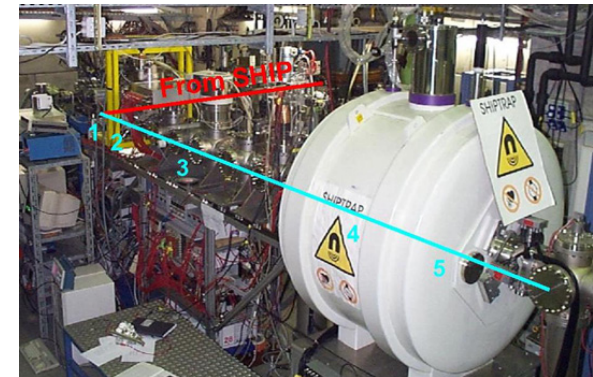
## Penning Trap mass spectrometry

- First experiments at TRIGA-TRAP (Uni-Mainz) in 2014 \*
  - Development of efficient Ho ion source using laser ablation
  - Uncertainties on  $^{163}\text{Dy}$  and  $^{163}\text{Ho}$  mass reduced by a factor of 2
  - Know-how to be applied in SHIPTRAP

\*Preparatory studies for a high-precision Penning trap measurement of the  $^{163}\text{Ho}$  electron capture Q-value  
F. Schneider et al., submitted to EPJ

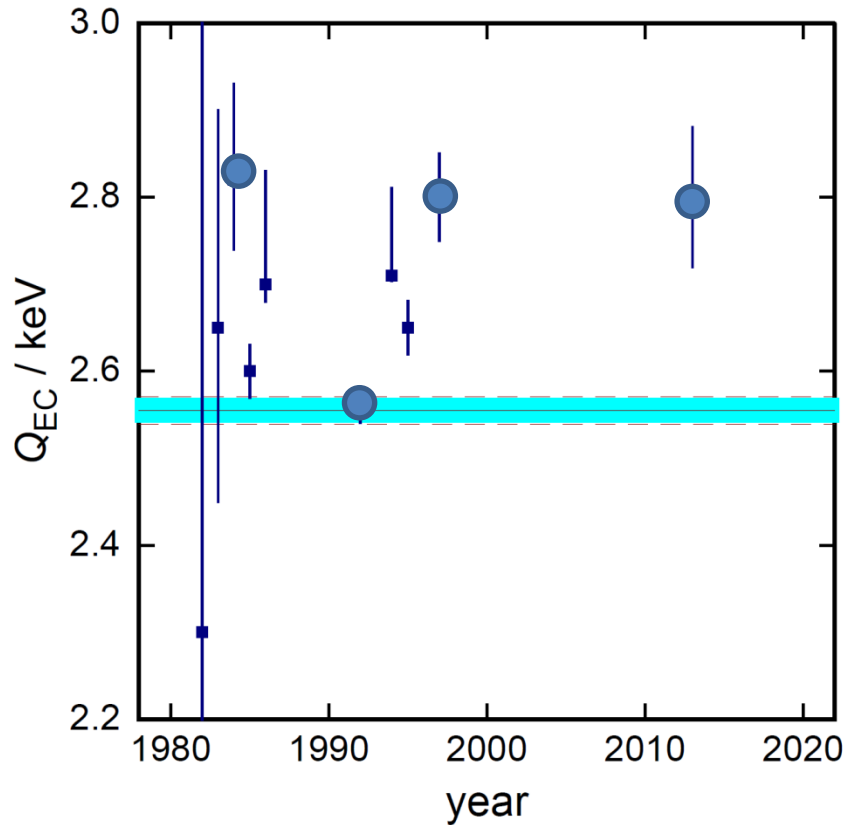
- Presently: SHIPTRAP (GSI) \*\*
  - $Q_{EC}$  determination with smaller uncertainties
  - Define scale of the experiment

\*\* Direct measurement of the mass difference of  $^{163}\text{Ho}$  and  $^{163}\text{Dy}$  as prerequisite to a determination of the electron neutrino mass  
S. Eliseev et al., submitted to PRL





# $Q_{EC}$ determination of $^{163}\text{Ho}$



- Calorimetric measurements
- Other methods

To reduce uncertainties in the analysis:  
 $Q_{EC}$  determination within **1 eV**  
→ **PENTATRAP (MPIK HD)**

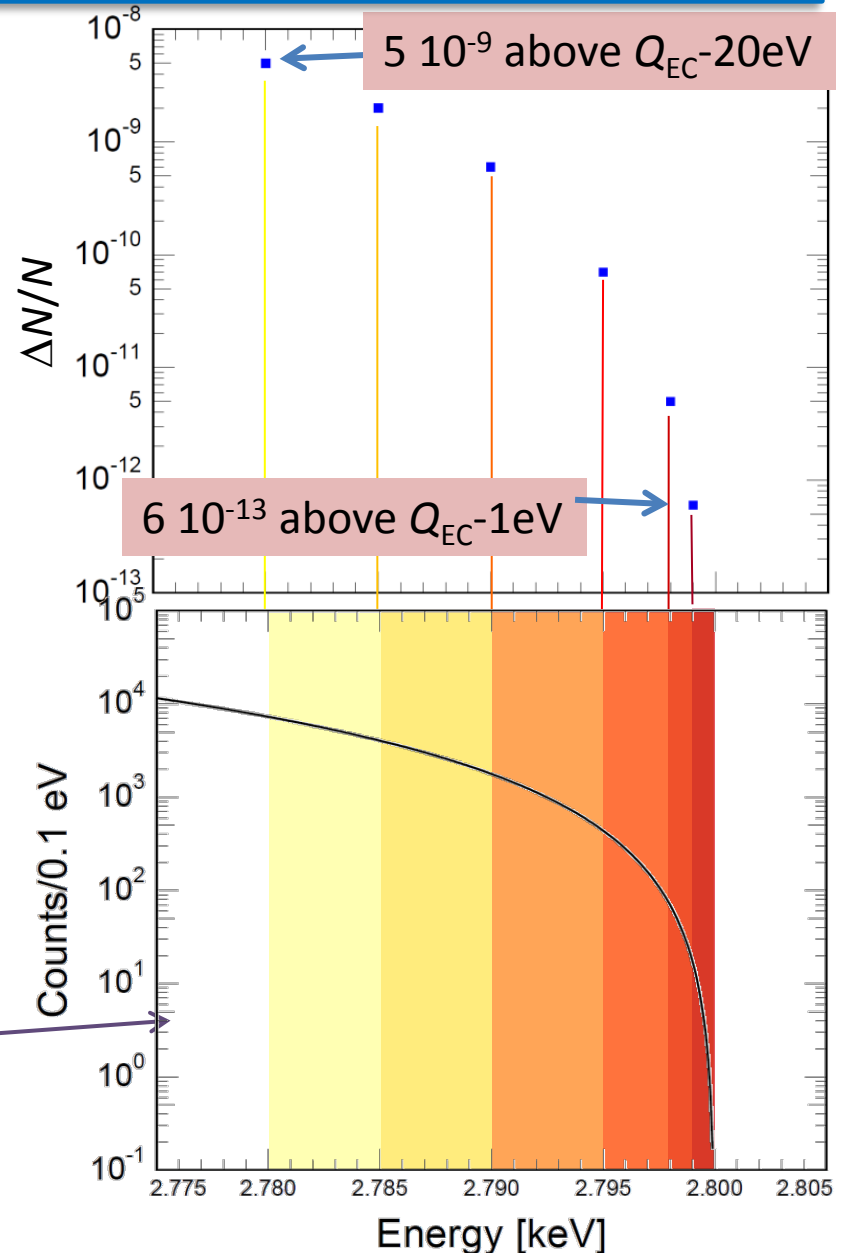
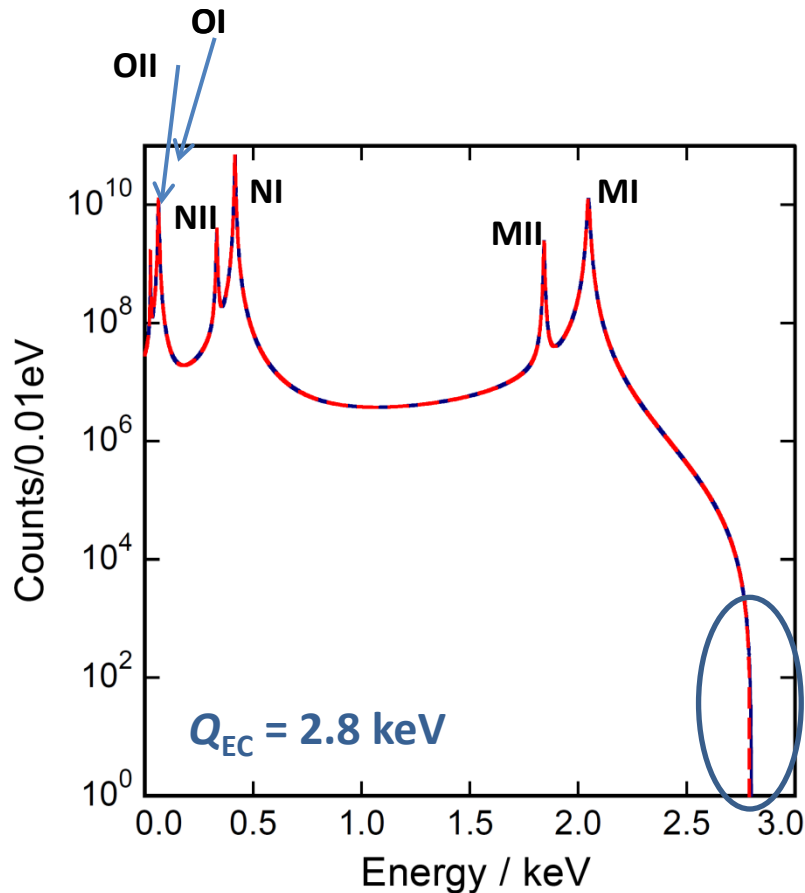
Talk by Klaus Blaum yesterday



# $^{163}\text{Ho}$ and neutrino mass: sub-eV sensitivity

Statistics in the end point region

- $N_{\text{ev}} > 10^{14} \rightarrow A \approx 1 \text{ MBq}$



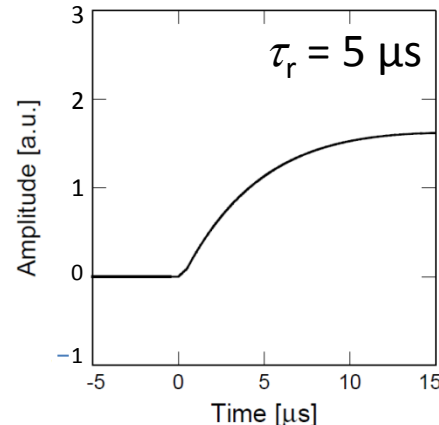
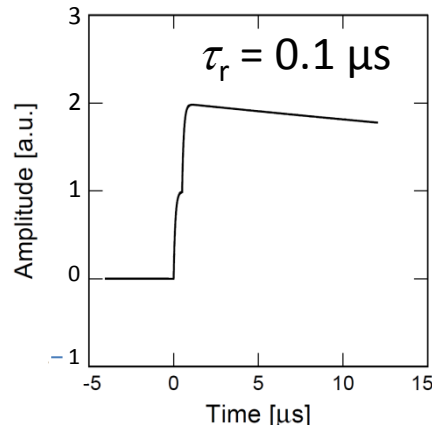
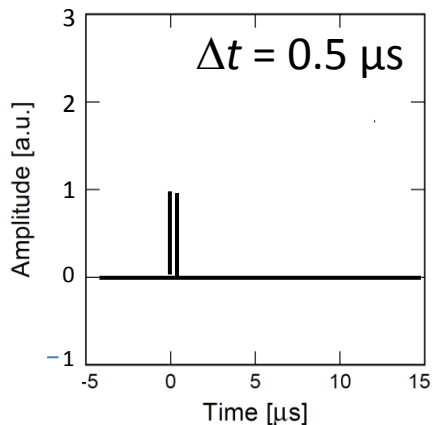
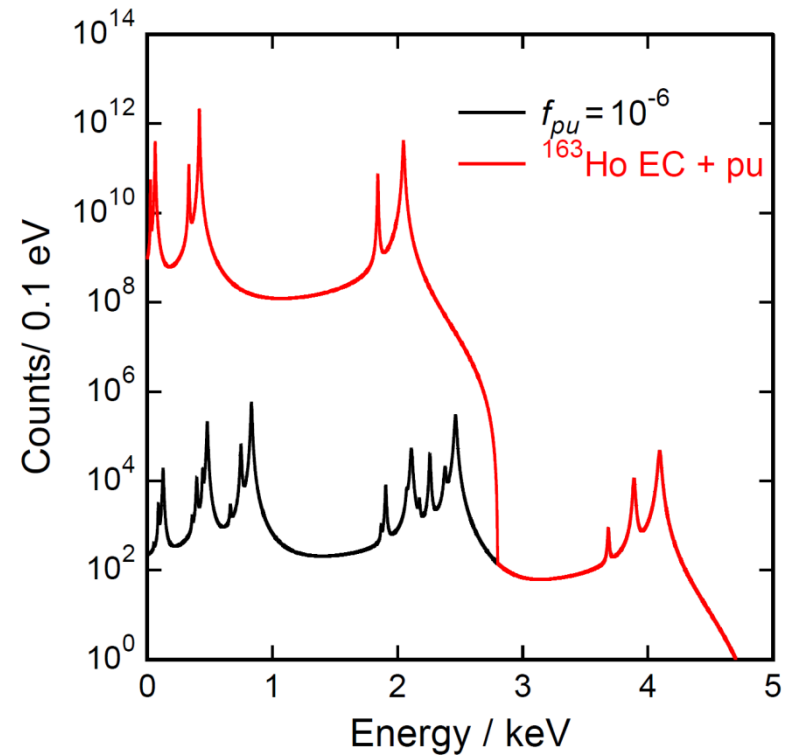
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Unresolved pile-up ( $f_{\text{pu}} \sim a \cdot \tau_r$ )

- $f_{\text{pu}} < 10^{-5}$
- $\tau_r < 1 \mu\text{s} \rightarrow a \sim 10 \text{ Bq}$
- $10^5$  pixels



# $^{163}\text{Ho}$ and neutrino mass: sub-eV sensitivity

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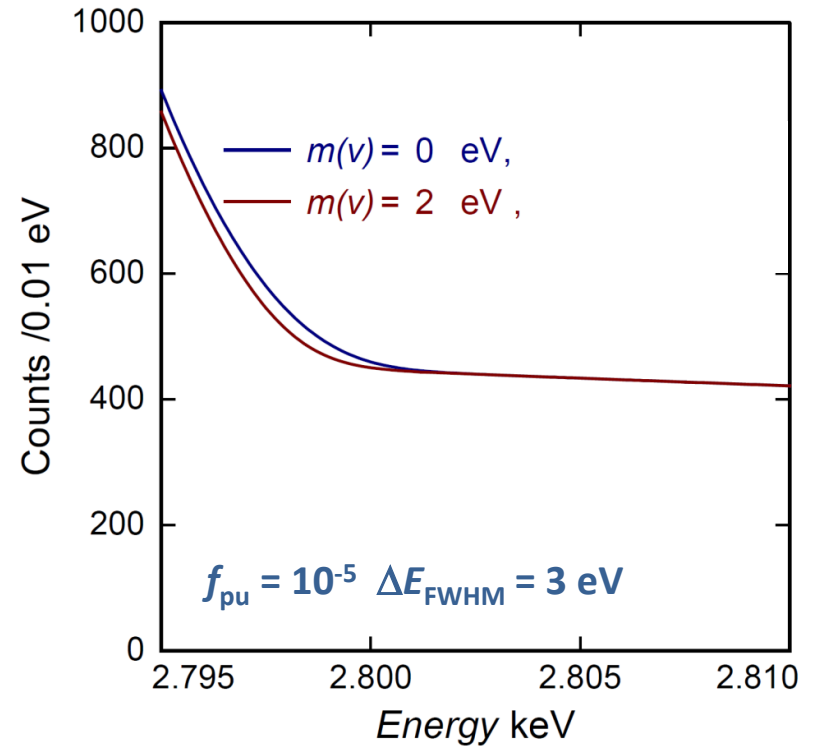
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Precision characterization of the endpoint region

- $\Delta E_{\text{FWHM}} < 2 \text{ eV}$



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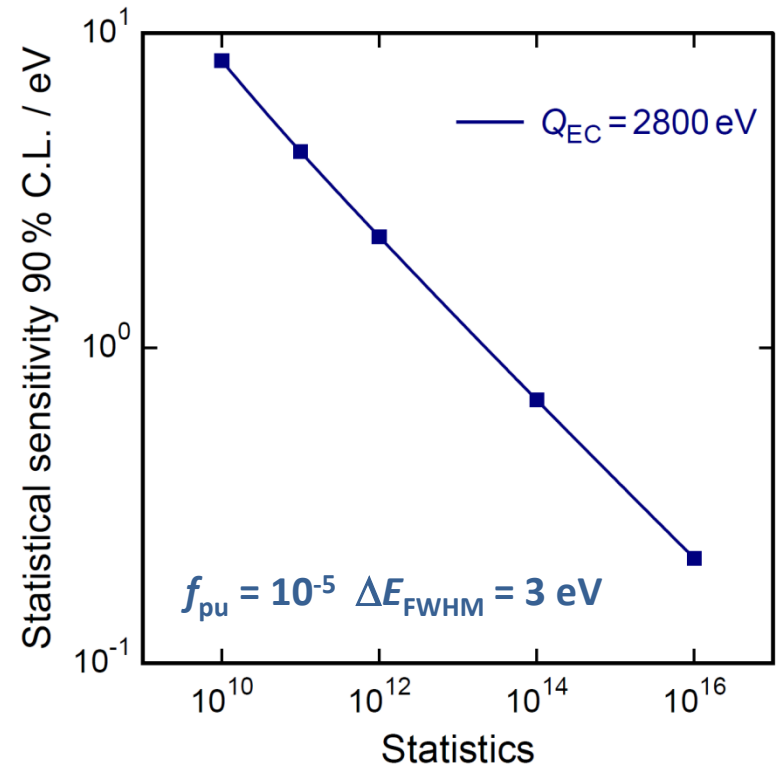
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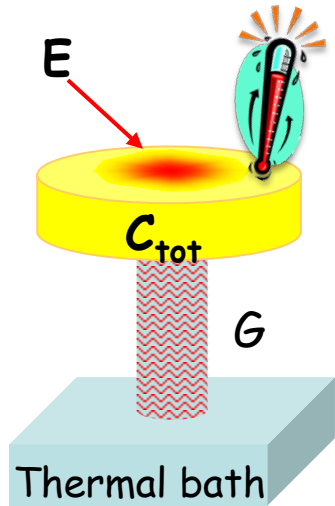
- $\Delta E_{\text{FWHM}} < 2 \text{ eV}$

Low temperature

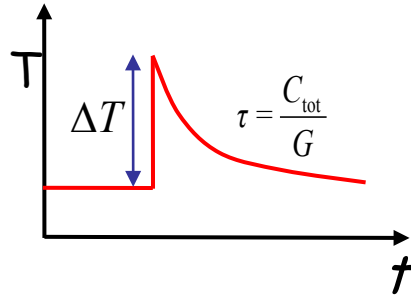
Metallic Magnetic Calorimeter



# Low temperature micro-calorimeters



$$\Delta T \cong \frac{E}{C_{\text{tot}}}$$



$$E = 10 \text{ keV}$$

$$C_{\text{tot}} = 1 \text{ pJ/K}$$

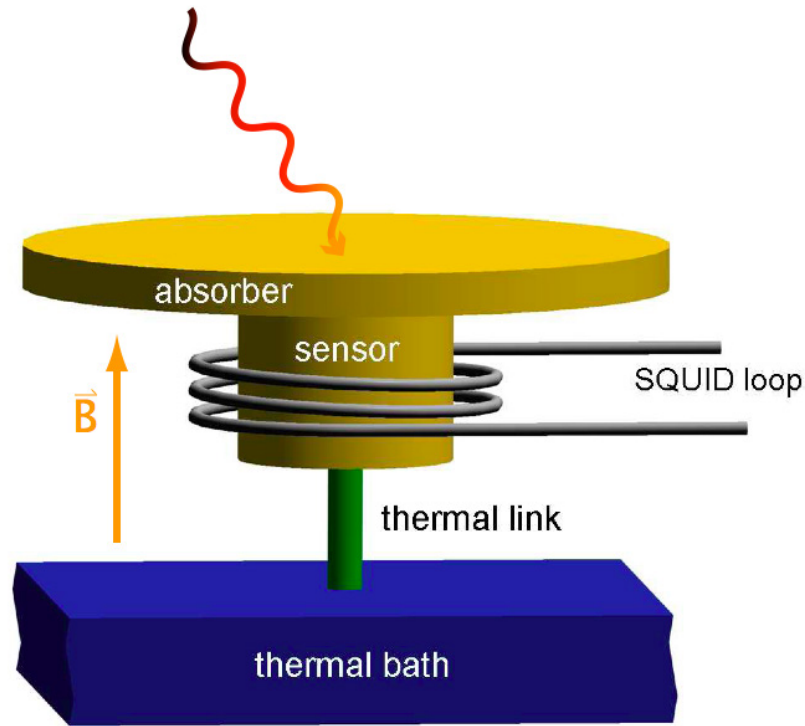
→ ~ 1 mK

- Very small volume
- Working temperature below 100 mK  
small specific heat  
small thermal noise
- Very sensitive temperature sensor

# Metallic magnetic calorimeters (MMCs)

A. Fleischmann et al.,  
*AIP Conf. Proc.* **1185**, 571, (2009)

- Paramagnetic Au:Er sensor



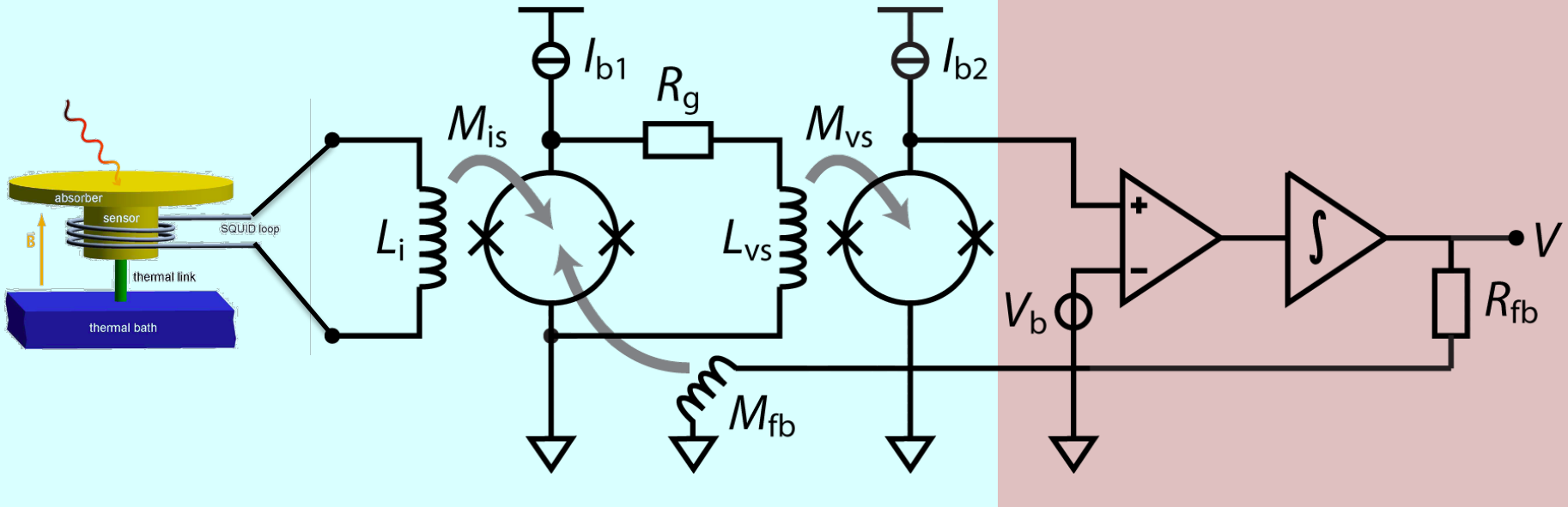
$$\Delta\Phi_s \propto \frac{\partial M}{\partial T} \Delta T \quad \rightarrow \quad \Delta\Phi_s \propto \frac{\partial M}{\partial T} \frac{E}{C_{\text{sens}} + C_{\text{abs}}}$$



# MMCs: Readout

$T \sim 30 \text{ mK}$

$T \sim 300 \text{ mK}$



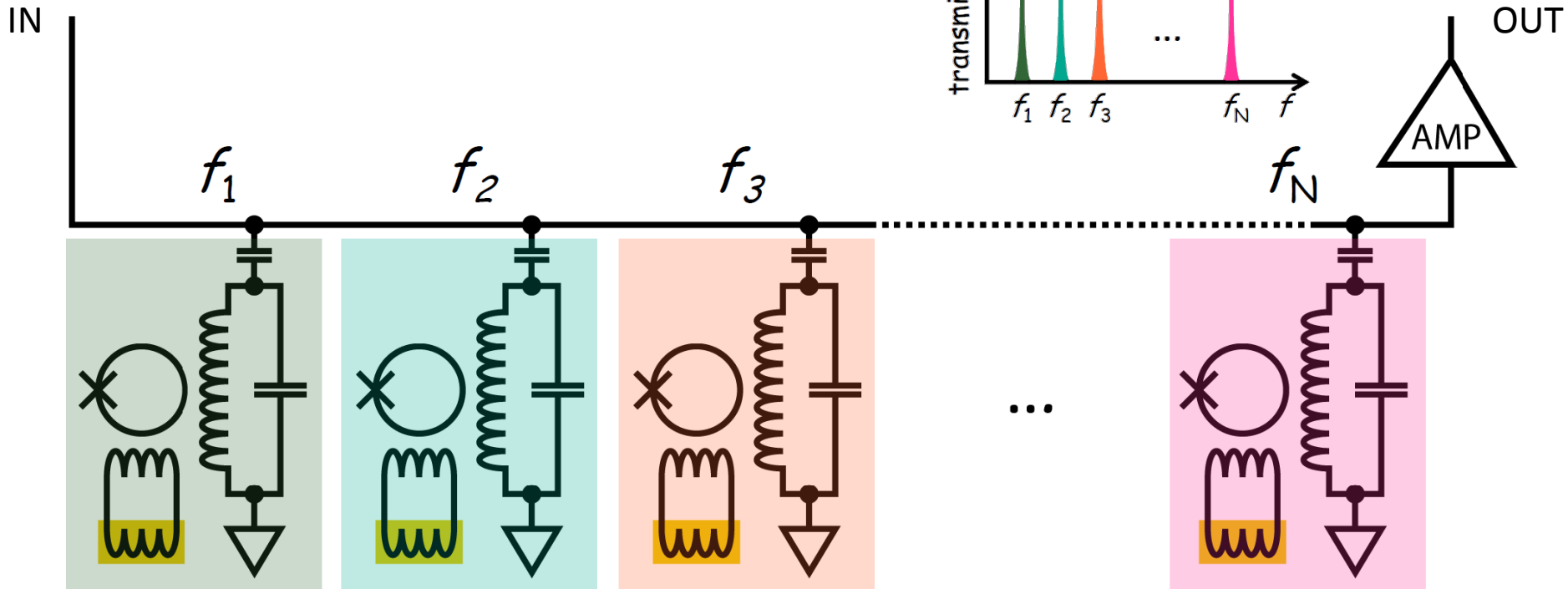
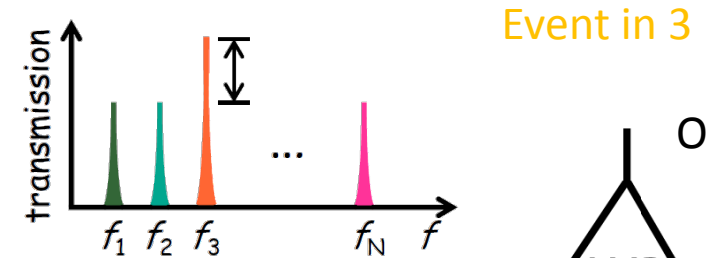
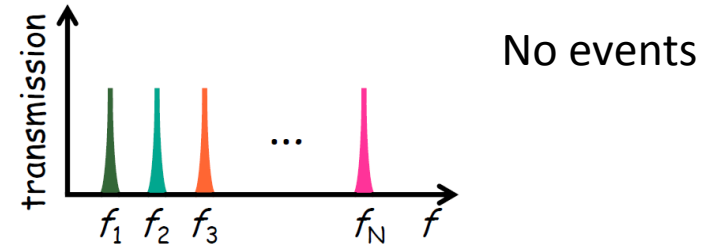
Two-stage SQUID setup with flux locked loop allows for:

- low noise
- large bandwidth / slewrate
- small power dissipation on detector SQUID chip (voltage bias)

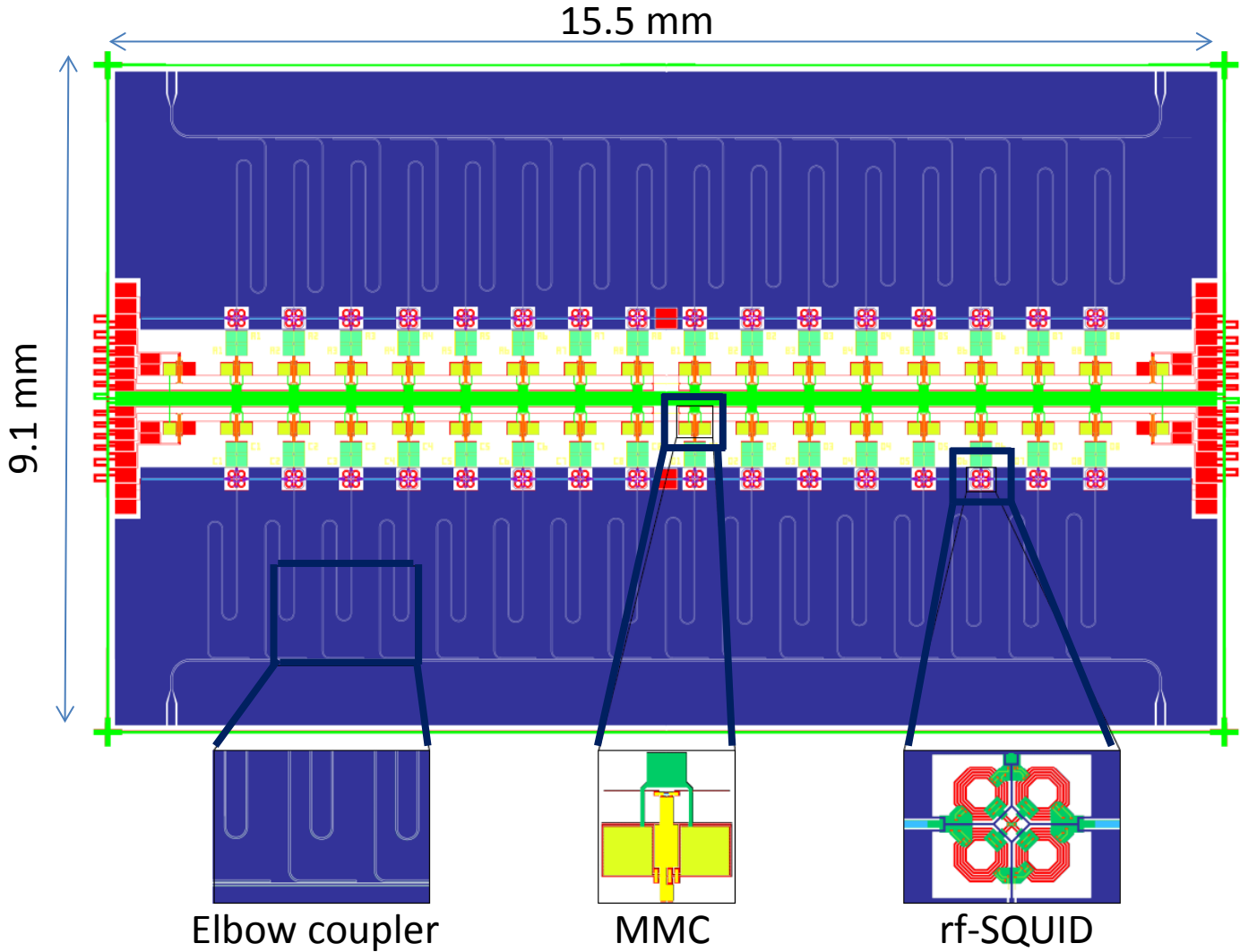


# MMCs: Microwave SQUID multiplexing

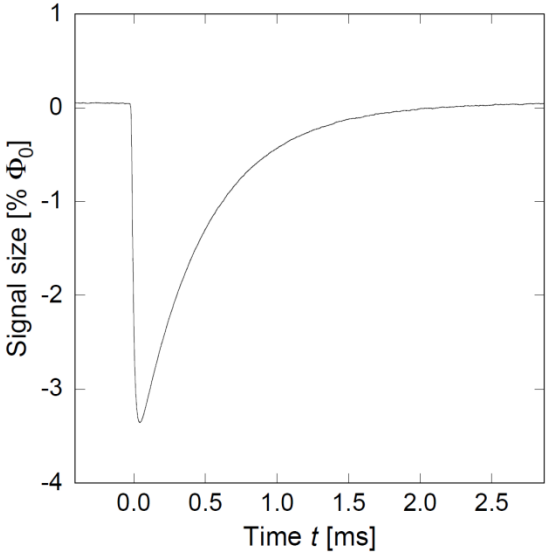
- Single HEMT amplifier and 2 coaxes to read out **100 - 1000** detectors



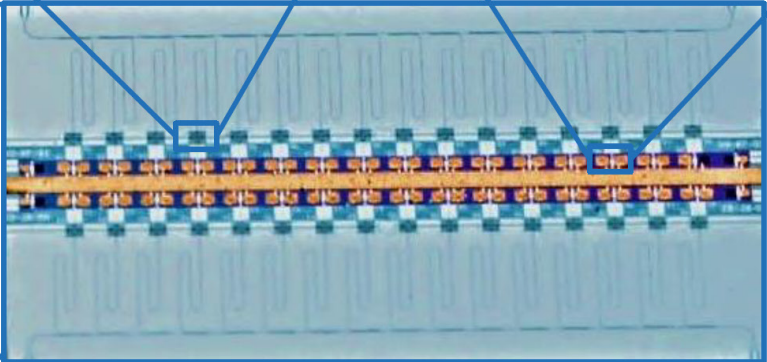
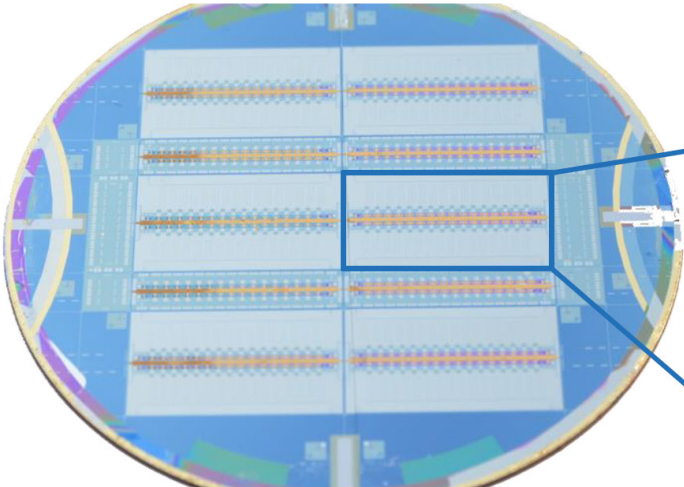
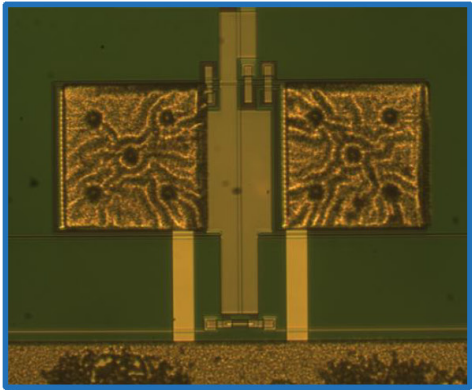
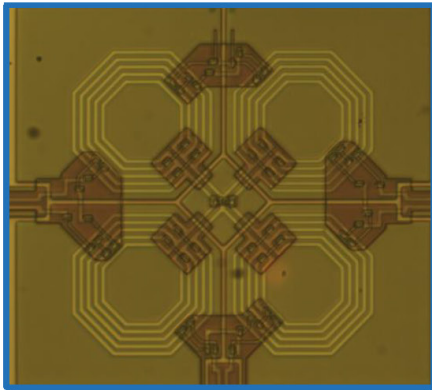
# MMCs: Microwave SQUID multiplexing



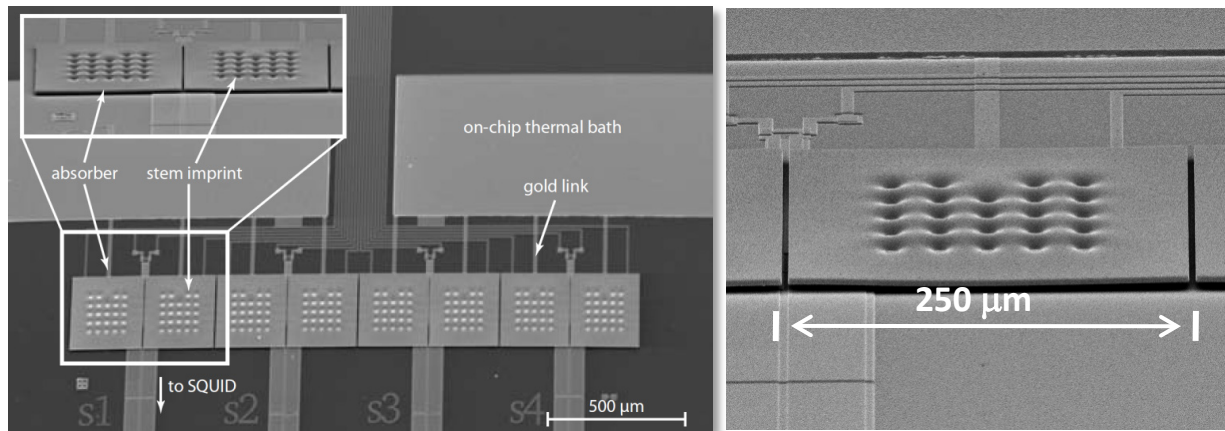
# MMCs: Microwave SQUID multiplexing



Successful production and test of the first prototype

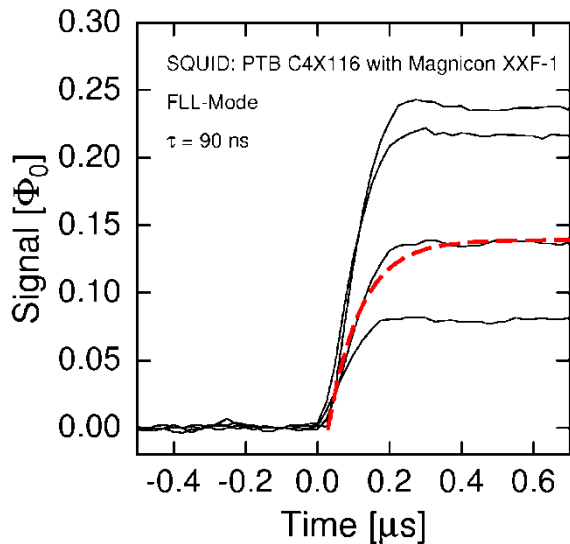


# MMCs: 1d-array for soft x-rays ( $T=20$ mK)

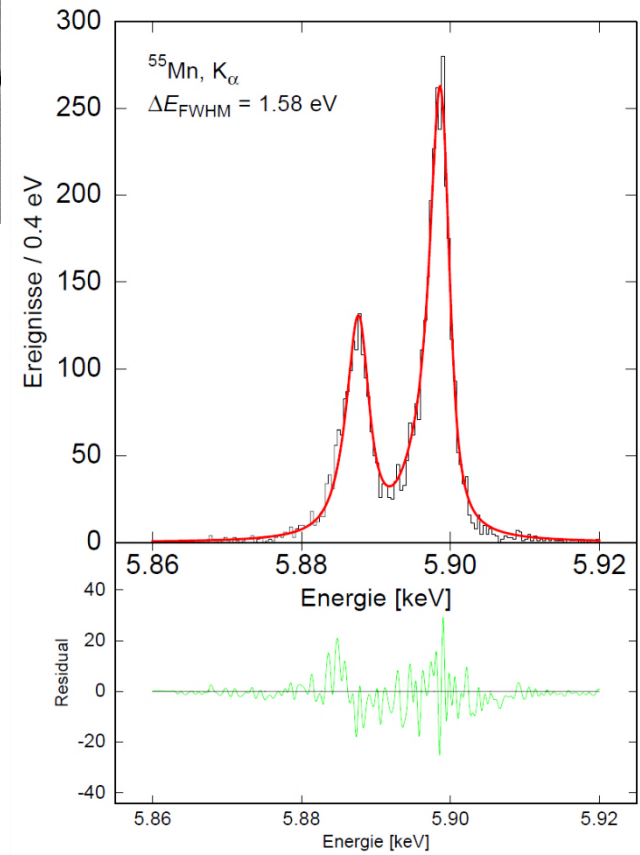
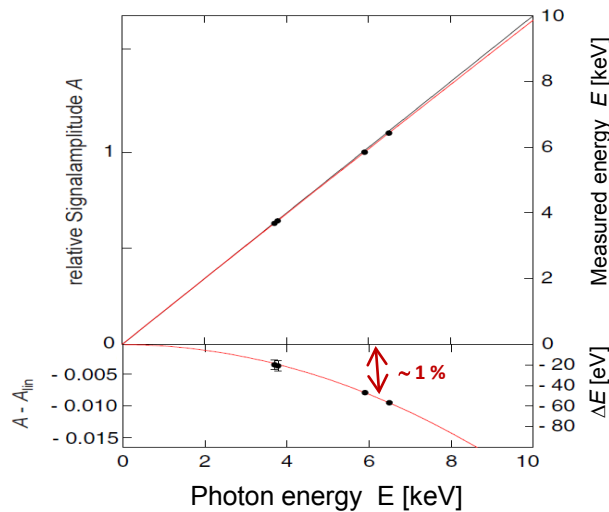


$$\Delta E_{FWHM} = 1.6 \text{ eV @ } 6 \text{ keV}$$

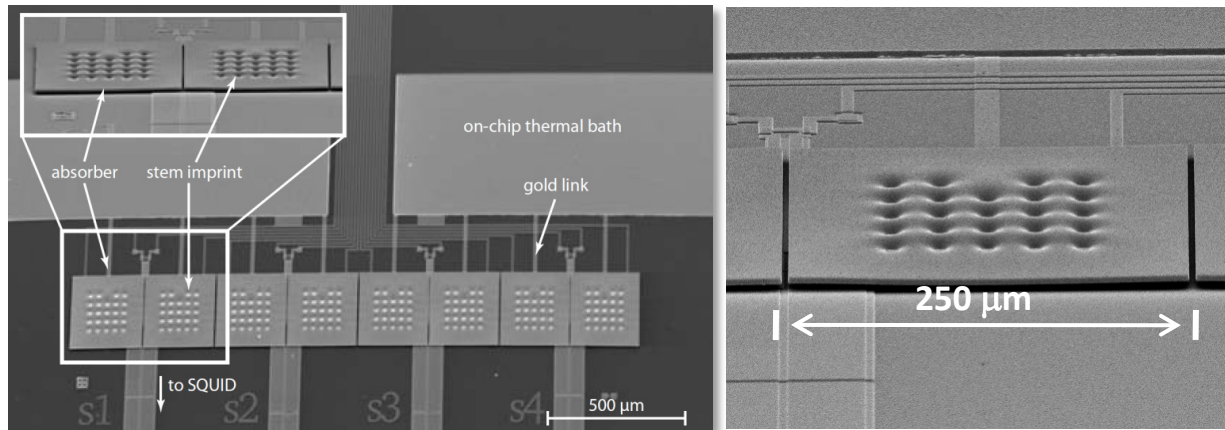
Rise Time: 90 ns



Non-Linearity < 1% @6keV

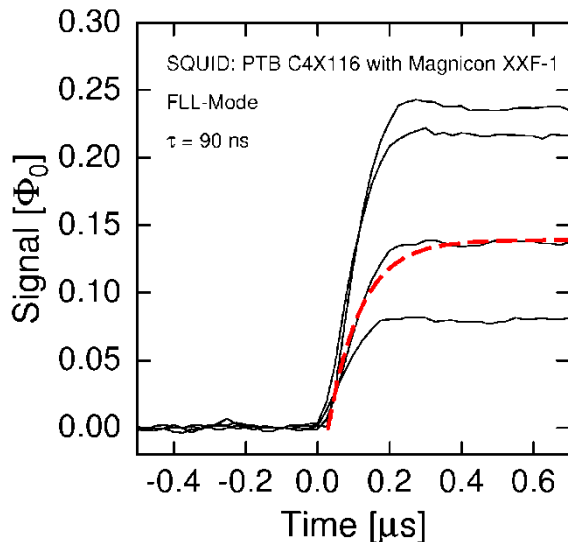


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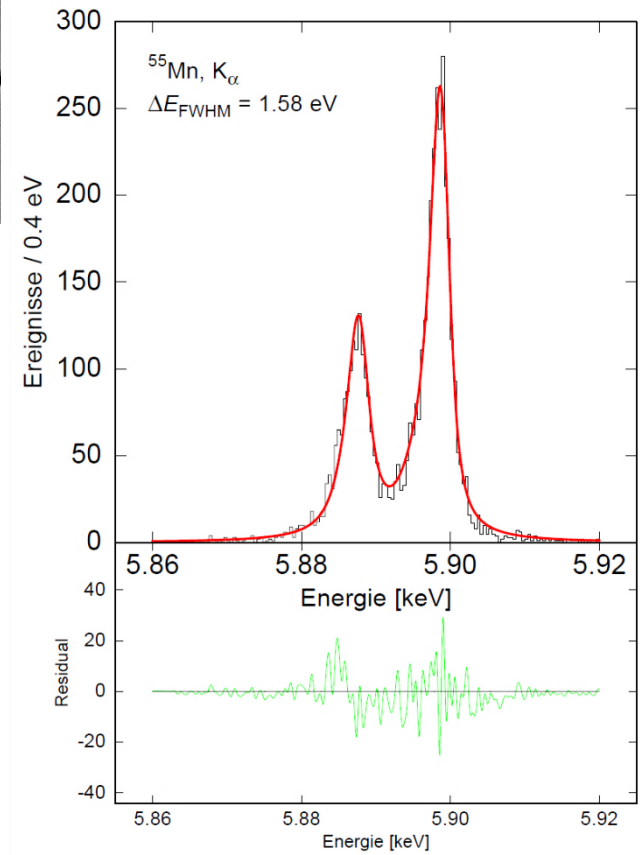
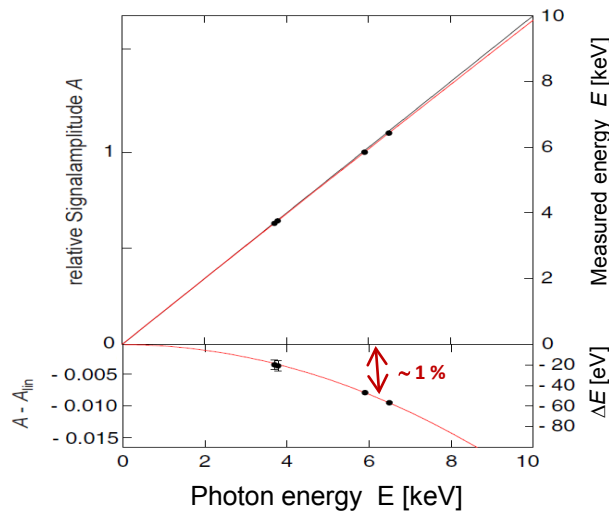


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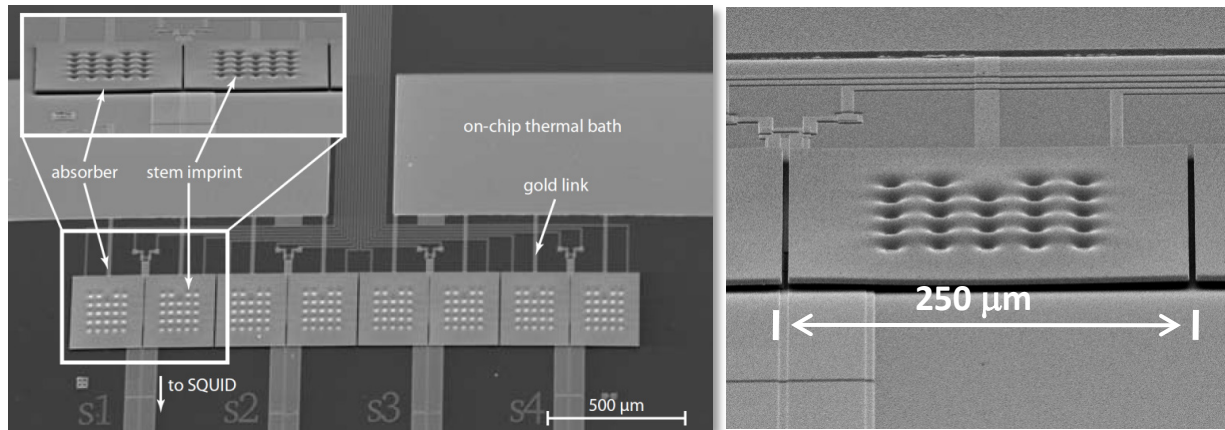


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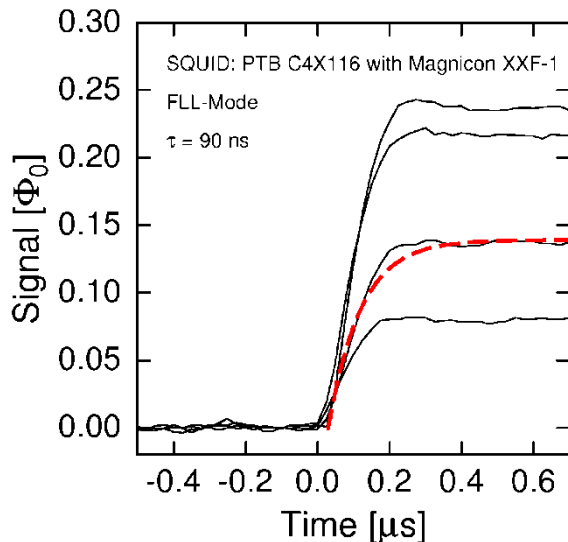
Reduction  
un-resolved pile-up

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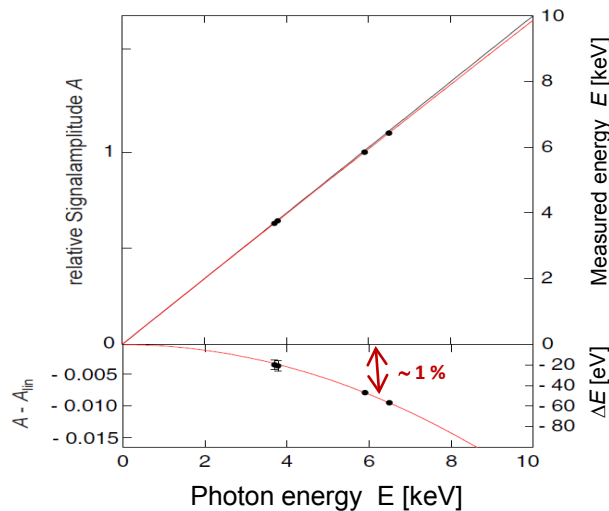
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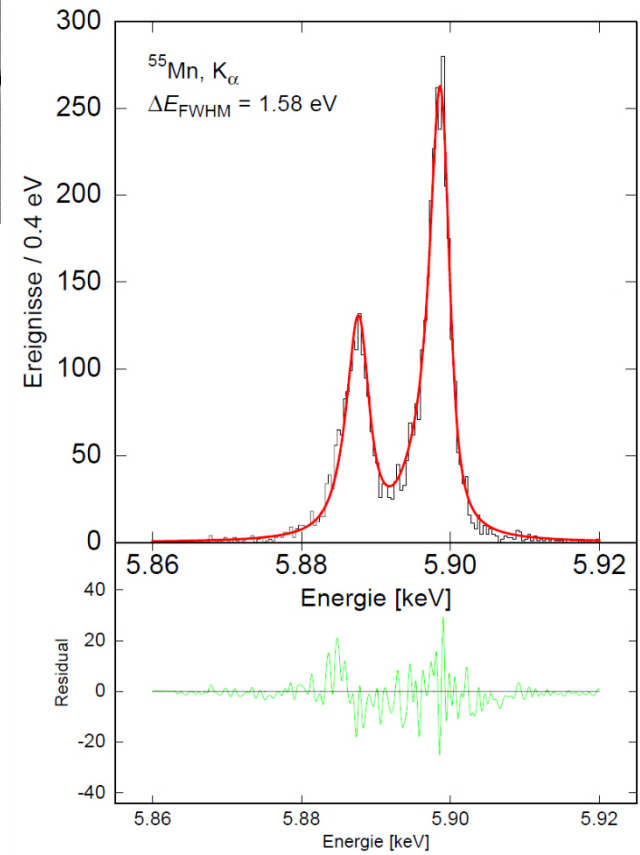


Reduction  
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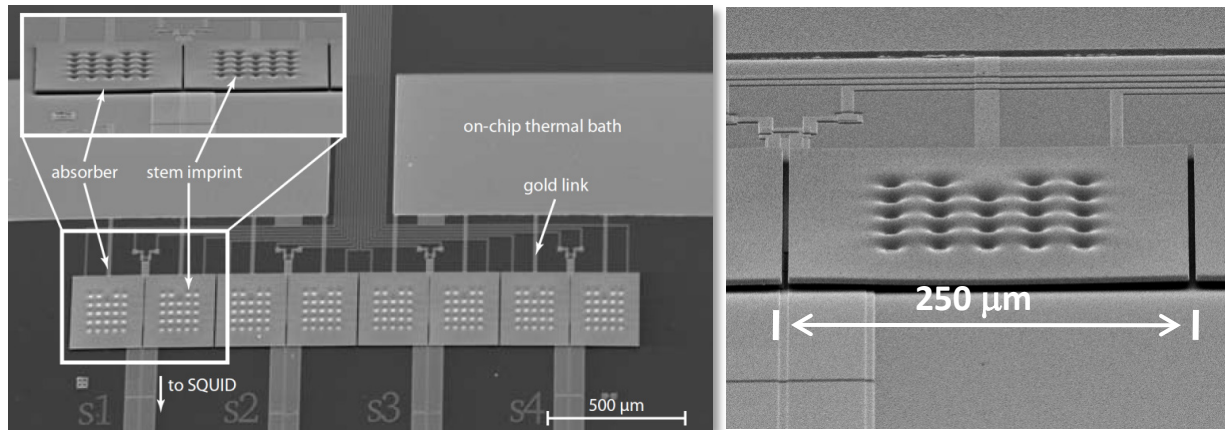


Definition  
of the energy scale



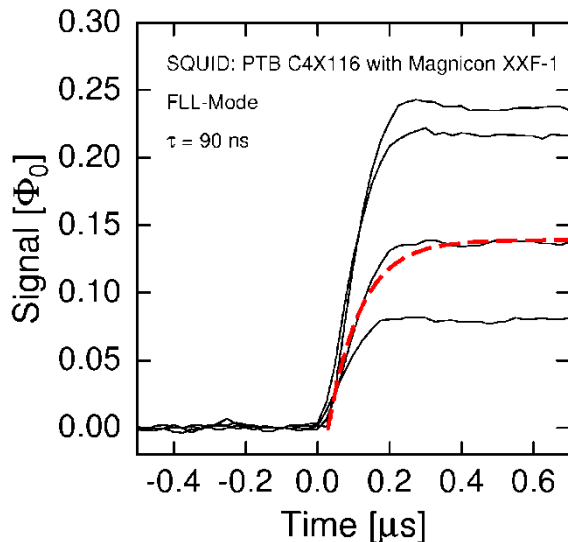


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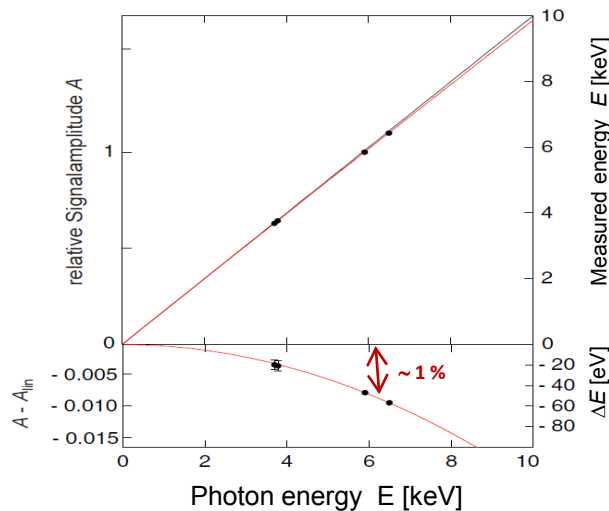
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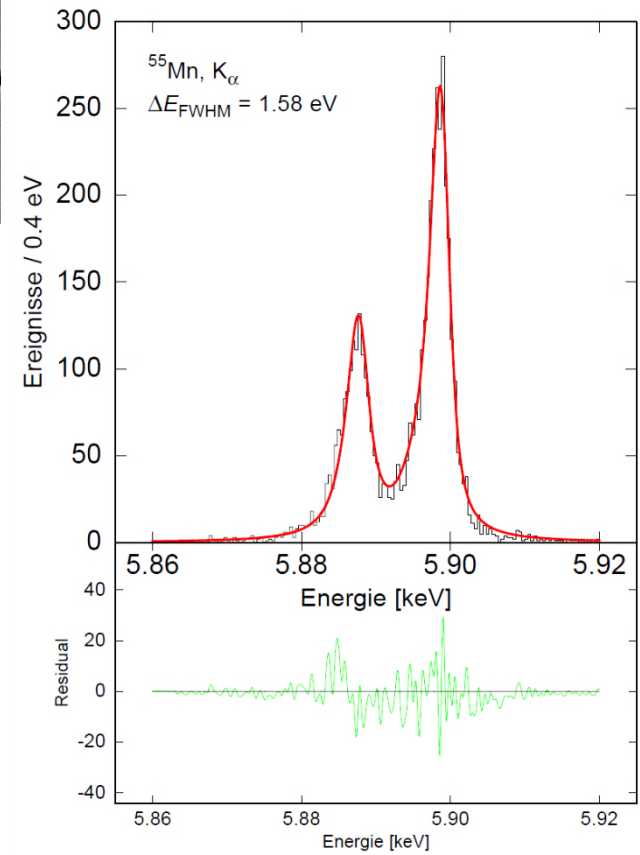


Reduction  
un-resolved pile-up

Non-Linearity < 1% @6keV



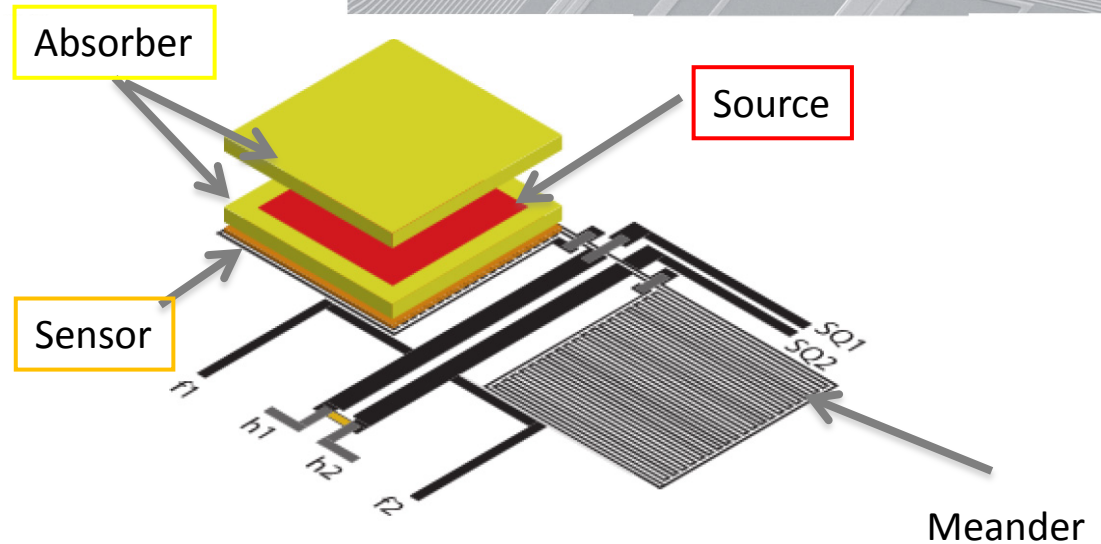
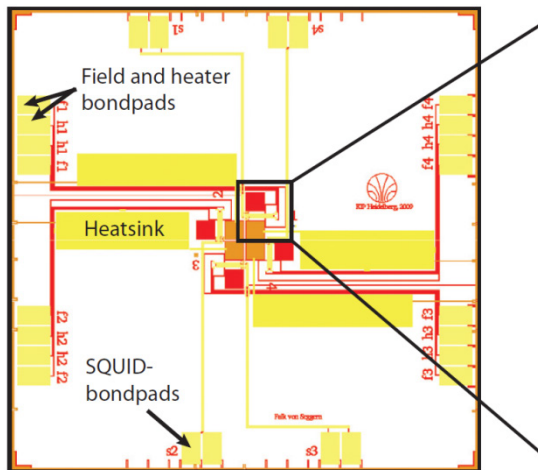
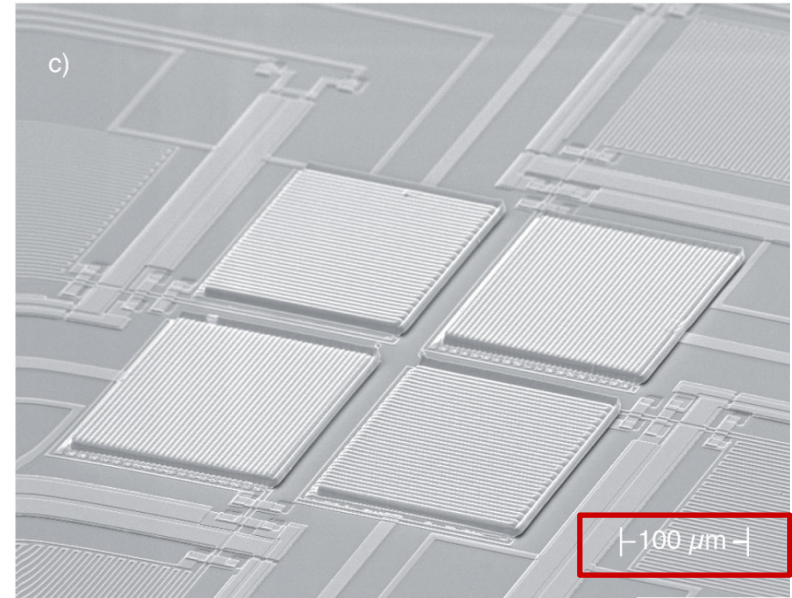
Definition  
of the energy scale



Reduced smearing  
in the end point region

# First detector prototype for $^{163}\text{Ho}$

- Absorber for calorimetric measurement  
→ ion implantation @ ISOLDE-CERN in 2009
- About 0.01 Bq per pixel
- Operated over more than 4 years

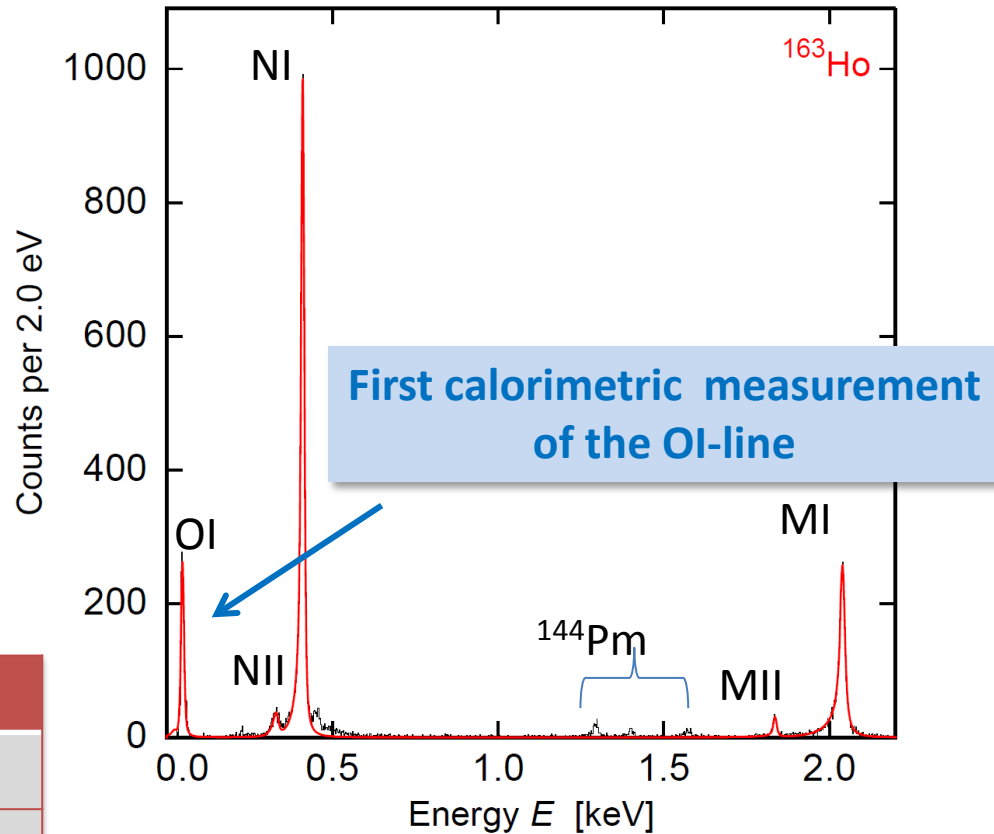




# Calorimetric spectrum

- Rise Time  $\sim 130$  ns
- $\Delta E_{\text{FWHM}} = 7.6$  eV @ 6 keV (2013)  
 $\Delta E_{\text{FWHM}} = 2.4$  eV @ 0 keV (2014)
- Non-Linearity  $< 1\%$  @ 6keV
- Synchronized measurement of 2 pixels
- Presently most precise  $^{163}\text{Ho}$  spectrum

	$E_{\text{H}}$ bind.	$E_{\text{H}}$ exp.	$\Gamma_{\text{H}}$ lit.	$\Gamma_{\text{H}}$ exp
<b>MI</b>	2.047	2.040	13.2	13.7
<b>MII</b>	1.845	1.836	6.0	7.2
<b>NI</b>	0.420	0.411	5.4	5.3
<b>NII</b>	0.340	0.333	5.3	8.0
<b>OI</b>	0.050	0.048	5.0	4.3



$$Q_{\text{EC}} = (2.843 \pm 0.009^{\text{stat}} - 0.06^{\text{syst}}) \text{ keV}$$

# Where to improve

## High purity $^{163}\text{Ho}$ source:

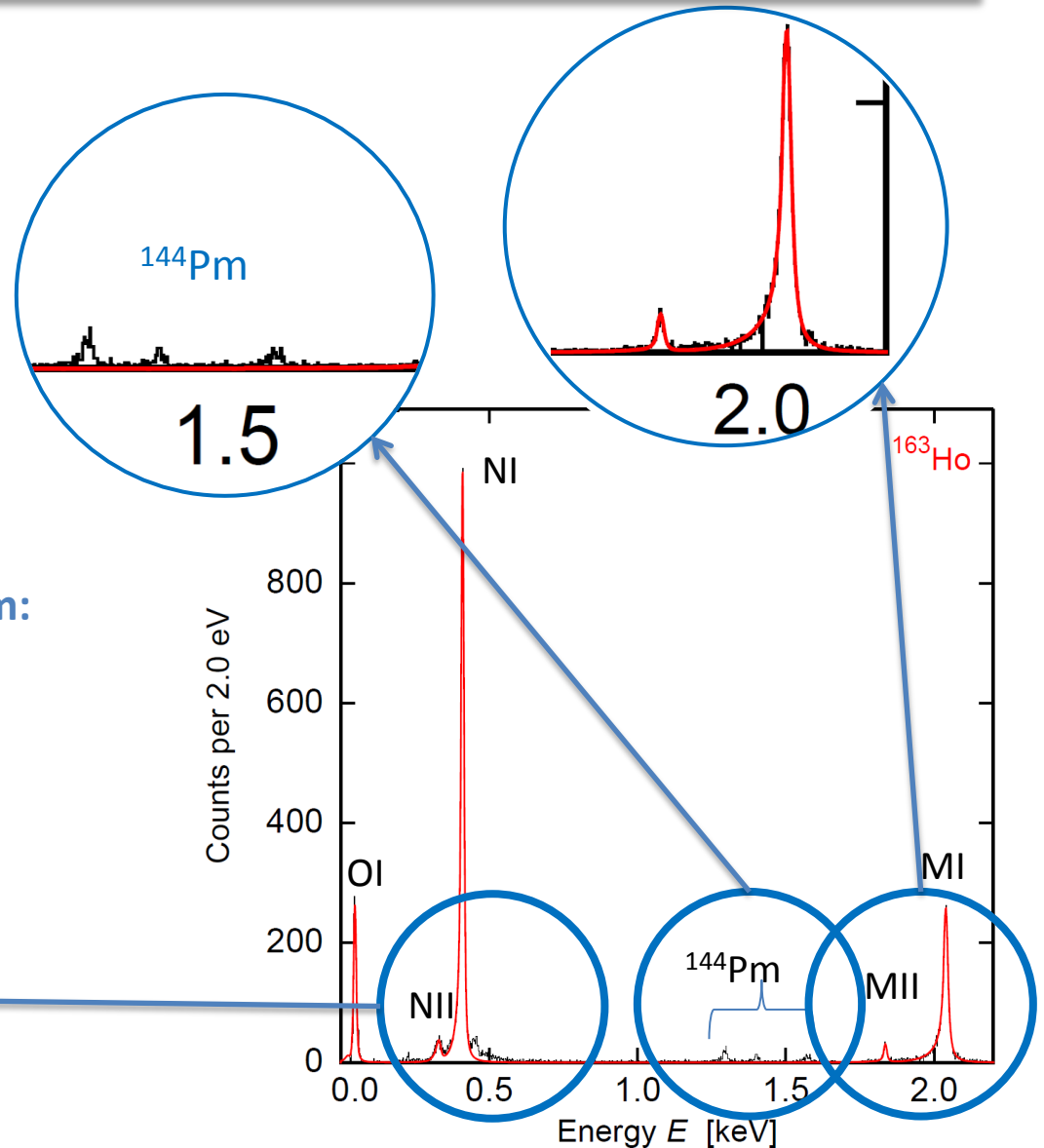
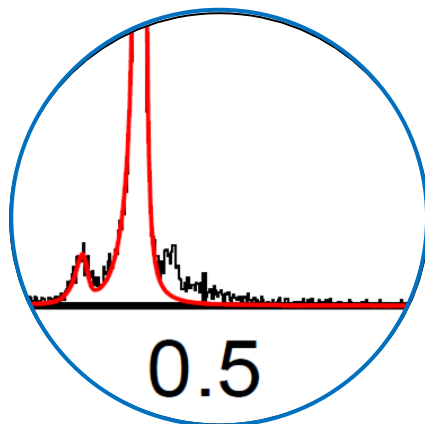
- Background reduction

## Detector design and fabrication:

- Increase activity per pixel
- Remove low energy tail

## Understanding of the $^{163}\text{Ho}$ spectrum:

- Investigate undefined structures



# High purity $^{163}\text{Ho}$ source: (n, $\gamma$ )-reaction on $^{162}\text{Er}$

## June 2012 : one irradiation at BER II Research Rector Berlin :

-Irradiate 5 mg Er for 11 days  $\Rightarrow 1.5 \cdot 10^{16}$  atoms  $^{163}\text{Ho}$

## Summer 2013: two irradiations at ILL Grenoble

- Prior to irradiation:

all elements lighter than Er separated  
(including holmium)

- After irradiation:

all elements heavier than Ho are separated  
 $\rightarrow$  pure holmium fraction

- 30 mg for 55 days  $\Rightarrow 1.6 \cdot 10^{18}$  atoms  $^{163}\text{Ho}$

- 7 mg for 7 days  $\Rightarrow 1.4 \cdot 10^{16}$  atoms  $^{163}\text{Ho}$



Thermal neutron flux  
( $\Phi$ ):  $1.3 \times 10^{15} \text{ cm}^{-2} \text{ s}^{-1}$

  
NEUTRONS  
FOR SCIENCE

PAUL SCHERRER INSTITUT  




# High purity $^{163}\text{Ho}$ source: Chemical separation

$\gamma$  spectrum of the 30 mg sample **after** chemical separation:

$\Rightarrow$  **only  $^{166\text{m}}\text{Ho}$  visible**

**Excellent chemical separation!**

High purity needed for MMC measurements:

$$^{166\text{m}}\text{Ho} / ^{163}\text{Ho} < 10^{-9}$$

# High purity $^{163}\text{Ho}$ source: Mass separation

## RISIKO off-line mass separator

- Optimized resonant laser ionization for Ho
- Focalization of the beam for implantation onto sub-mm detector absorber

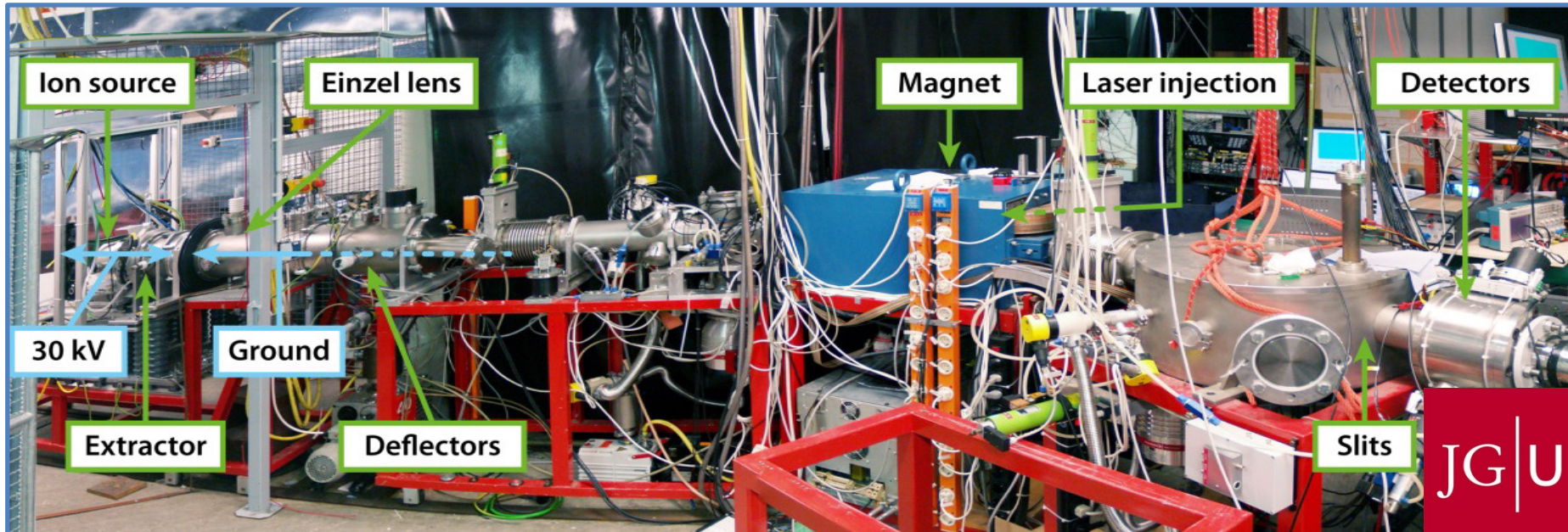
Magnetic sector-field Mass Spect.

30 kV two stage acceleration

60° double focussing separator magnet

Mass resolution:  $\frac{m}{\Delta m} = 500 - 1000$

Suppression of neighboring masses  $> 10^3$





# High purity $^{163}\text{Ho}$ source: Mass separation

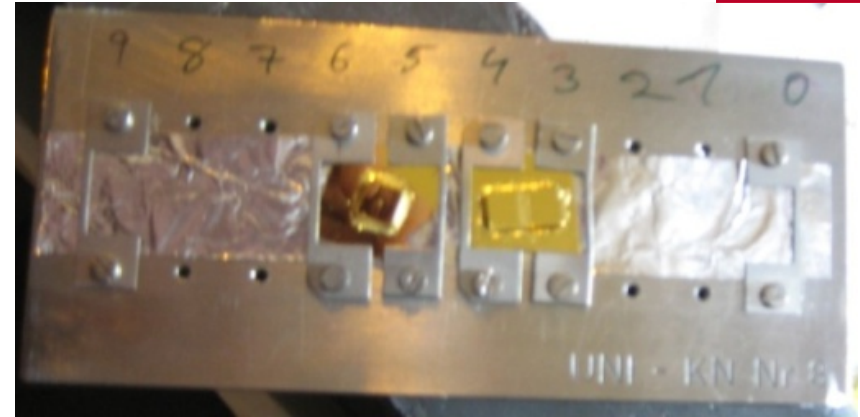
Chemically purified  $^{163}\text{Ho}$  source  
as offline target at ISOLDE/CERN

ISOLDE

JG|U

- Separation performed in December 2014
- 2 new chips to test!  
... each with 16 pixel detector arrays
- First low level gamma measurement in Tübingen

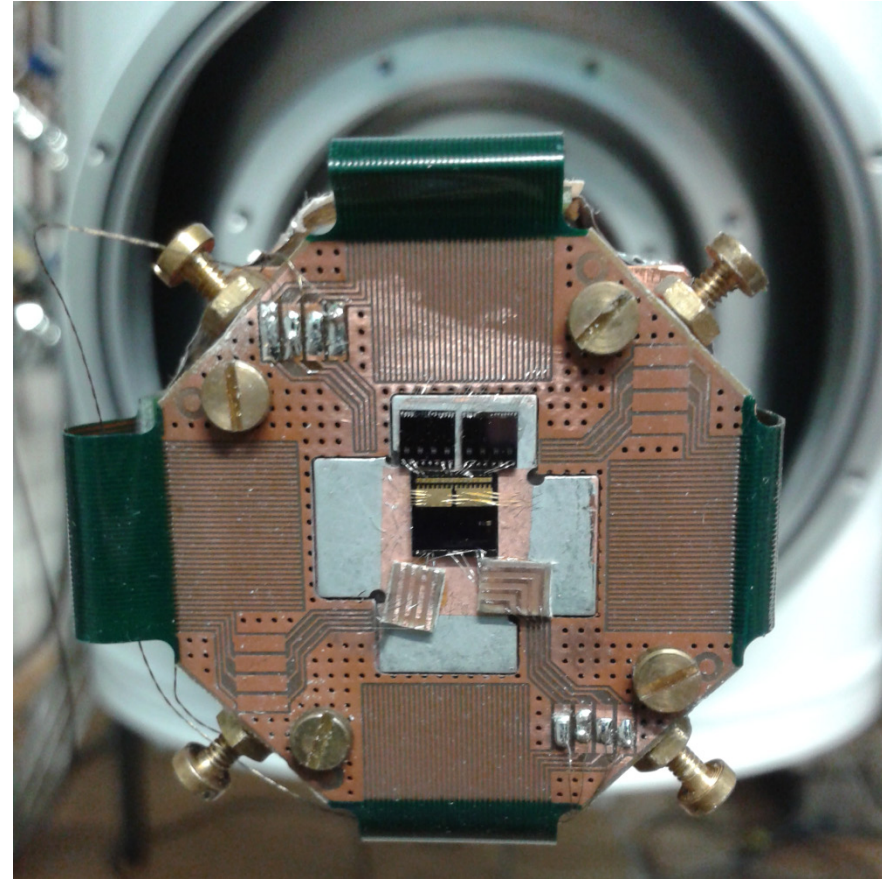
**No evidence of  $^{166\text{m}}\text{Ho}$**



# New detector ready for first tests....

Mounted on a cold arm of a dry cryostat

At 18 mK since last Wednesday.....



- Activity per pixel  $A \sim 0.2 \text{ Bq}$
- Baseline resolution  $\Delta E_{\text{FWHM}} = 5 \text{ eV}$
- No evidence of radioactive contamination in the source

DON'T MISS  
NEXT ECHO TALKS



# Where to improve

## High purity $^{163}\text{Ho}$ source:

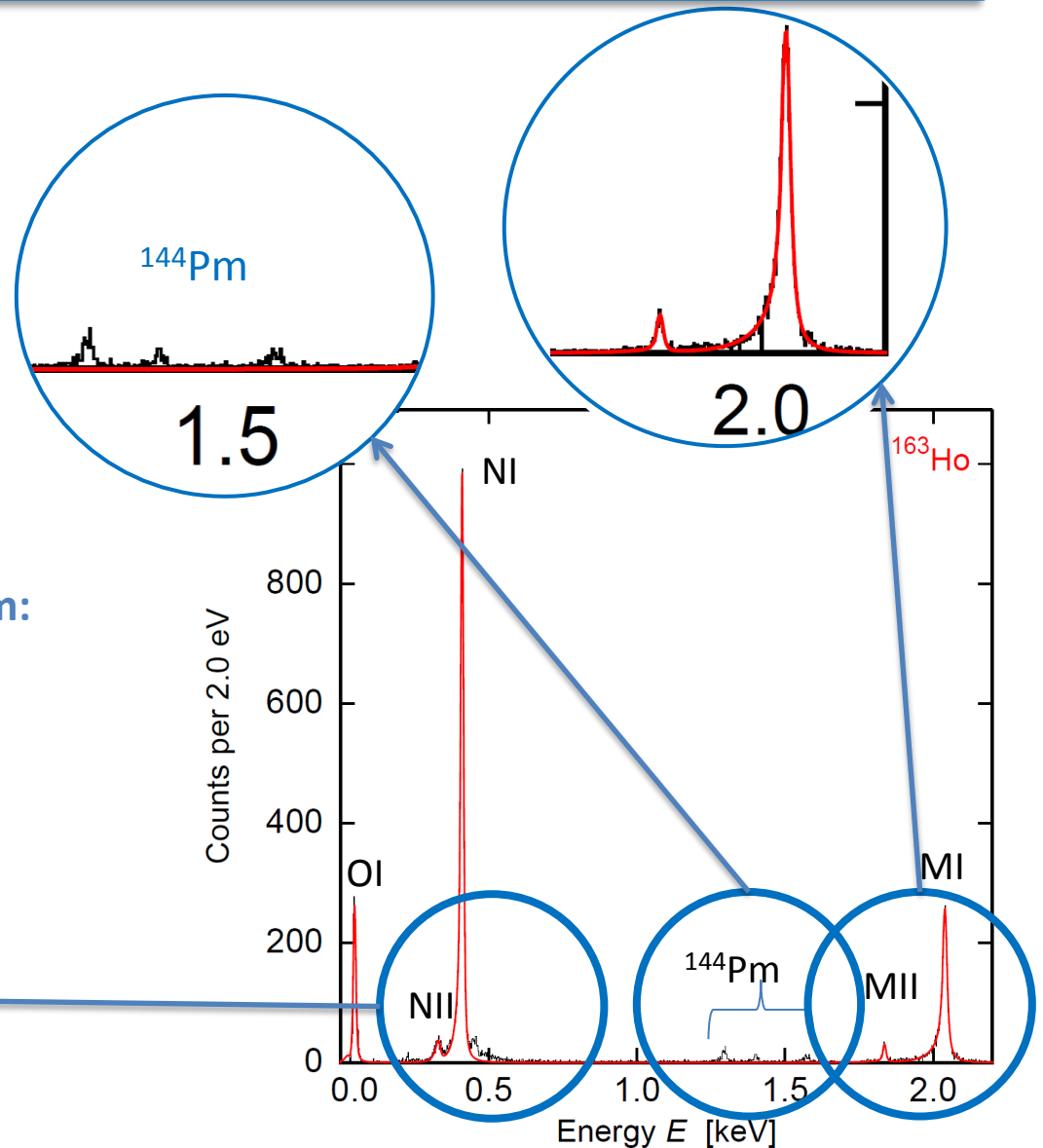
- ✓ Background reduction

## Detector design and fabrication:

- ✓ Increase activity per pixel
- ✓ Remove low energy tail

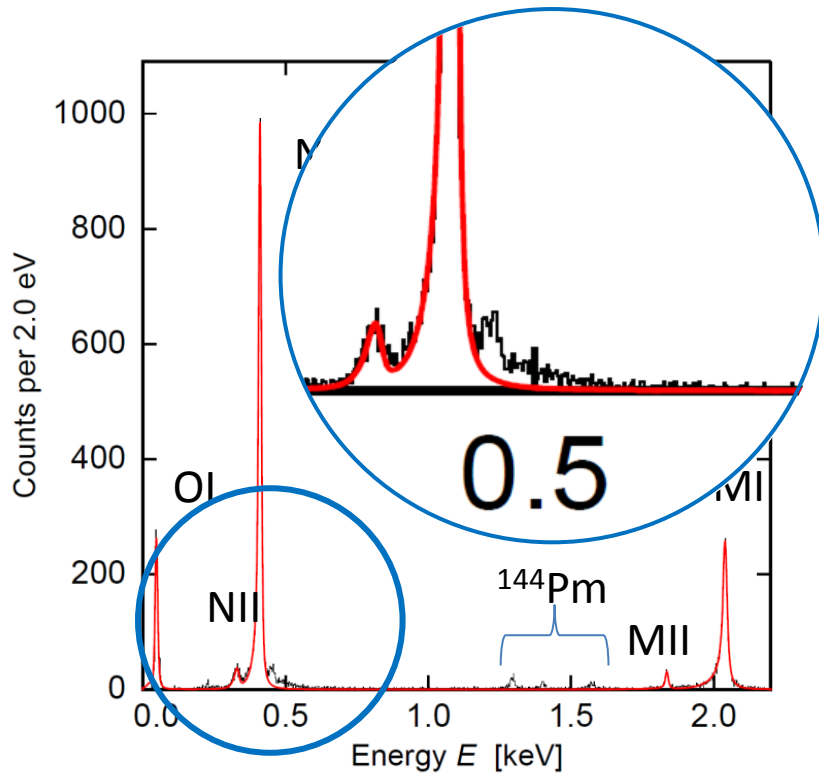
## Understanding of the $^{163}\text{Ho}$ spectrum:

- Investigate undefined structures





# Characterisation of spectral shape



Estimate the effect of

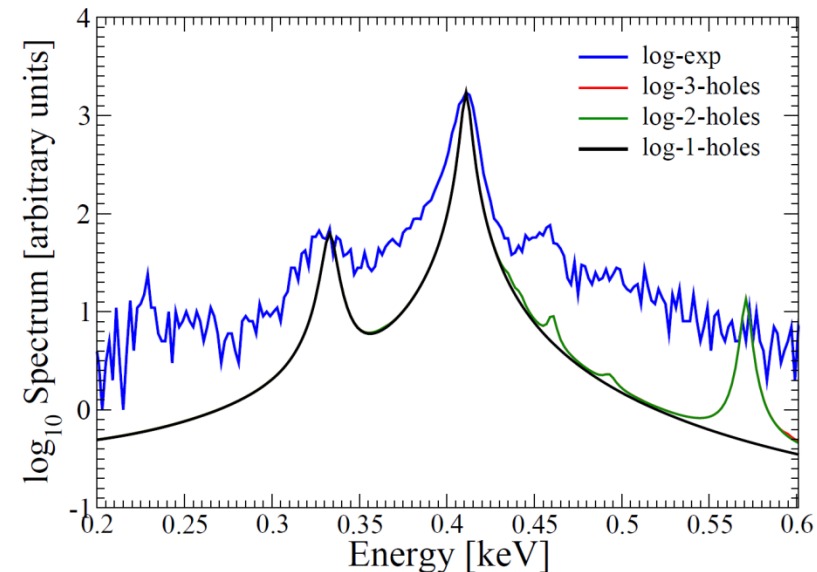
- Higher order excitation in  $^{163}\text{Dy}$
- $^{163}\text{Ho}$  ion embedded in Au

A. Faessler et al.  
*J. Phys. G* **42** (2015) 015108

R. G. H. Robertson  
*Phys. Rev. C* **91**, 035504 (2015)

A. Faessler et al.  
*Phys. Rev. C* **91**, 045505 (2015)

A. Faessler et al.  
*Phys. Rev. C* **91**, 064302 (2015)



# Background

## Background sources:

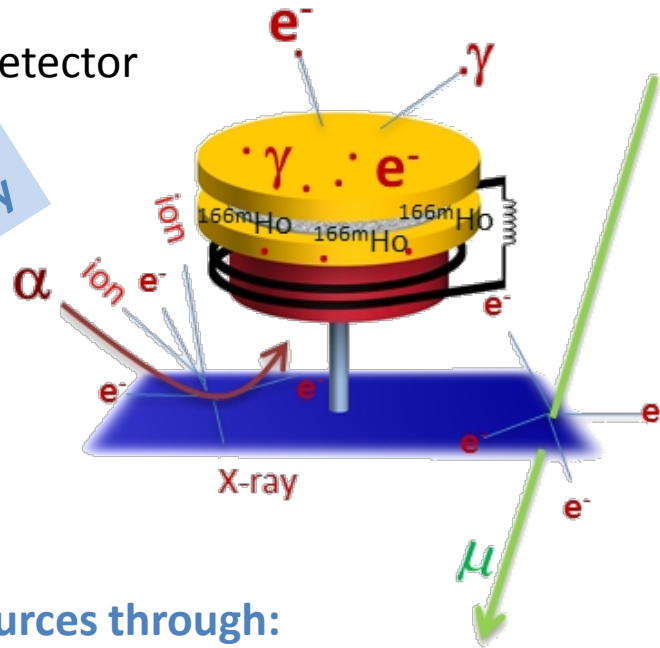
- Environmental radioactivity
- Cosmic rays
- Induced secondary radiation by cosmic rays
- Radioactivity in the detector

—————> Material screening

—————> { Underground labs  
μ-Veto



Underground measurements in **Modane**



Background level  
 $5 \times 10^{-5}$  counts/eV/det/day

## Study of background sources through:

- Monte Carlo simulations
- Dedicated experiments

Screening facilities

- **Uni-Tübingen**
- **Felsenkeller**



# ECHo overview

➤ Prove **scalability** with medium large experiment **ECHo-1K**

- $A \sim 1000$  Bq High purity  $^{163}\text{Ho}$  source (produced at reactor)
- $\Delta E_{\text{FWHM}} < 5$  eV
- $\tau_r < 1$   $\mu\text{s}$
- multiplexed arrays  $\rightarrow$  microwave SQUID multiplexing
  
- 1 year measuring time  $\rightarrow 10^{10}$  counts = Neutrino mass sensitivity  $m_\nu < 10$  eV

Just approved

Research Unit FOR 2202/1

„Neutrino Mass Determination by Electron Capture in Holmium-163 – ECHo“

**DFG** Deutsche  
Forschungsgemeinschaft

➤ **ECHo-1M** towards sub-eV sensitivity

# Thank you!

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