

Gravitational Waves from a Dark Sector

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Outline

- Motivation for (non-abelian) Dark Sectors
- Phase Transition of $SU(N)$ Theories
- GW Signal and ELISA

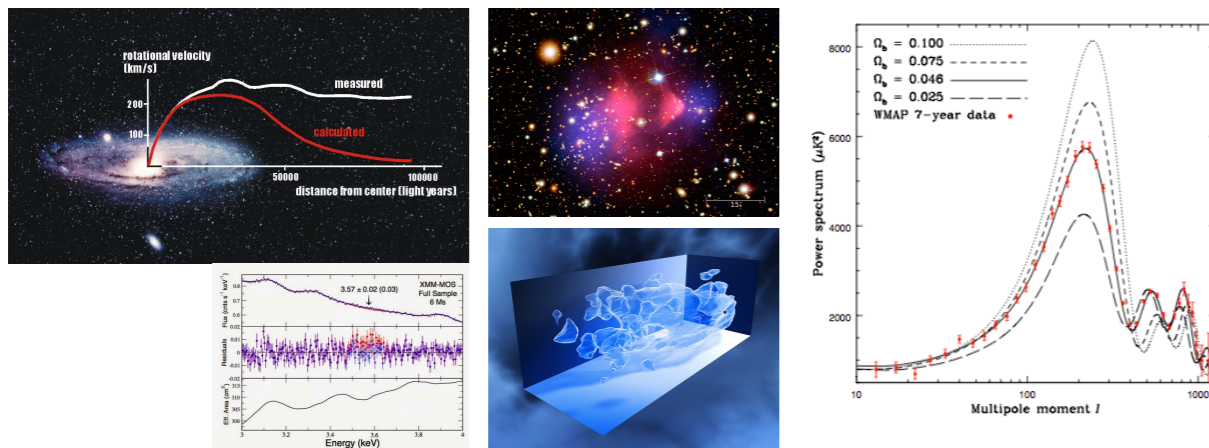
Based on 1504.07263

Motivation

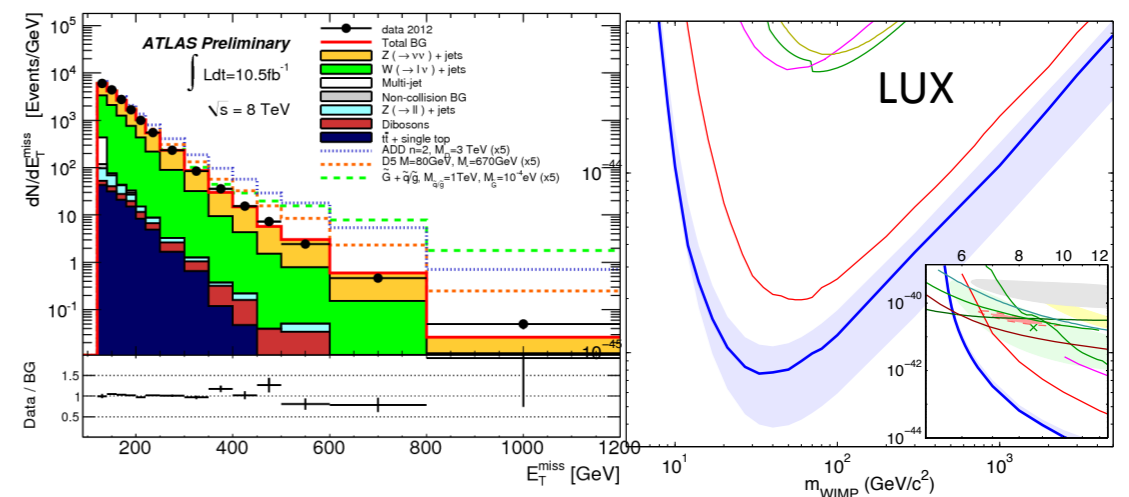
- Dark Sectors:
 - ▶ No or suppressed interaction with SM particles
 - ▶ “Dark” $SU(N_d)$ theory with n_f Dirac fermions
 - ▶ Confinement scale Λ_d
- Motivations:
 - ▶ Non-minimal Dark Matter models $M_{\text{DM}} \sim \Lambda_d$
 - ▶ Naturalness

Dark Matter

We have seen DM in the sky:



But no direct observation



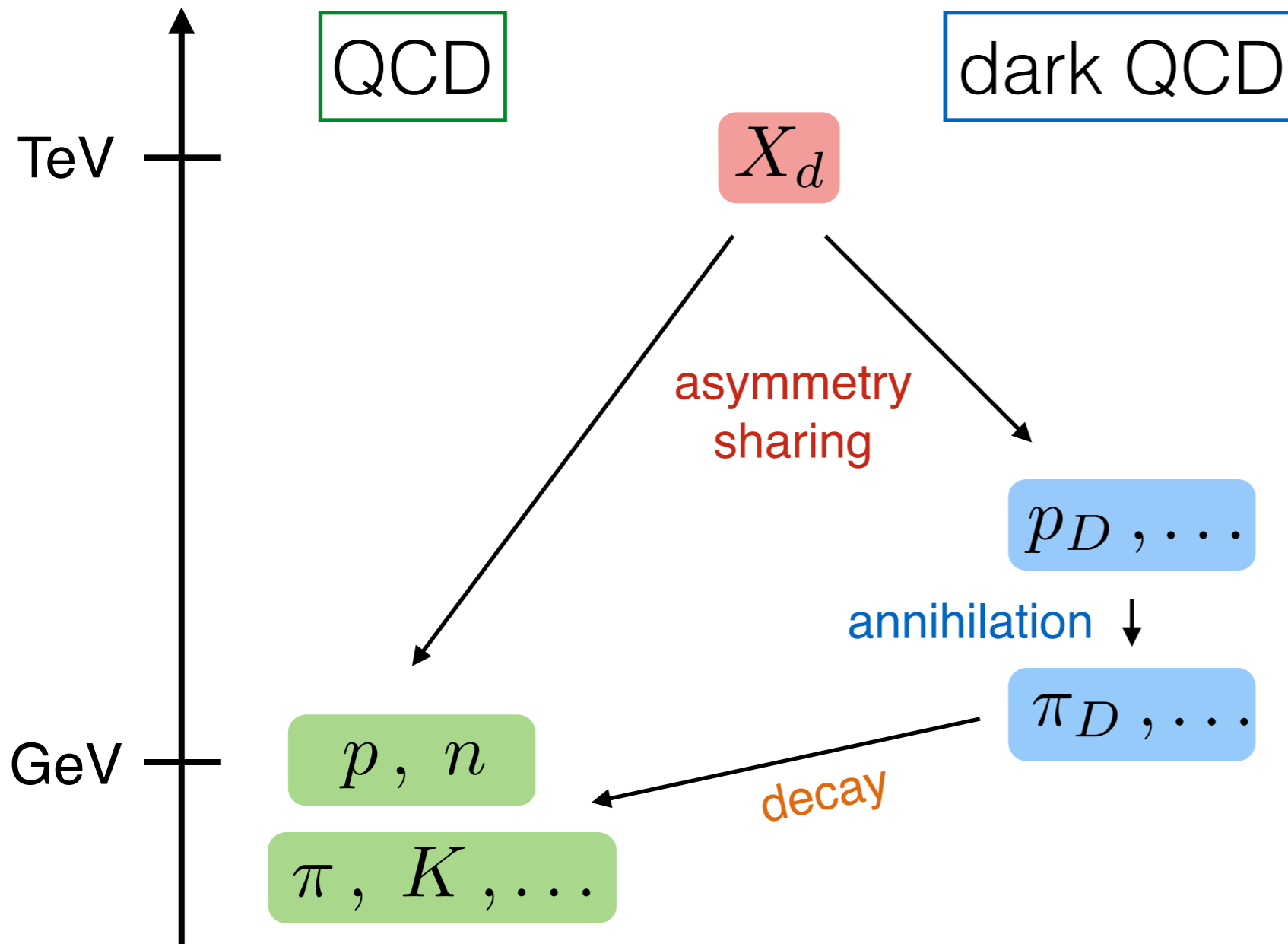
Maybe DM is just part of a larger dark sector

- Example: Proton is massive, stable, composite state
- DM self interactions solve structure formation problems
- New signals, new search strategies!

DM Motivation

- New mechanisms for relic density, extend mass range:
 - Asymmetric DM - GeV-TeV scale
 - Strong Annihilation - 100 TeV scale
 - SIMP - MeV scale
- Advantages of Composite
 - DM mass scale and stability
 - Fast annihilation for ADM
 - Self-interactions for structure formation

Asymmetric, Composite DM



- SU(N) dark sector with neutral **“dark quarks”**
- Confinement scale Λ_{darkQCD}
- DM is composite **“dark proton”**

Bai, PS, PRD 89, 2014
PS, Stolarski, Weiler, JHEP 2015

many other works!

Similar setup e.g.: Blennow et al; Cohen et al; Frandsen et al;
Reviews: Petraki & Volkas, 2013; Zurek, 2013;

Heavy Composite

- Confinement scales up to 100 TeV
 - Saturate unitarity bound for WIMP annihilation!

Griest, Kamionkowski, 1989

- If dark quarks charged under $SU(2) \times U(1)$
 - $M_{\text{DM}} > 10 \text{ TeV}$ for fermionic DM (odd N_d)

Appelquist et al, 1301.1693
 - $M_{\text{DM}} > 100s \text{ GeV}$ for bosonic DM (even N_d)

Appelquist et al, 1503.04205
 - Based on lattice studies of electromagnetic form factors

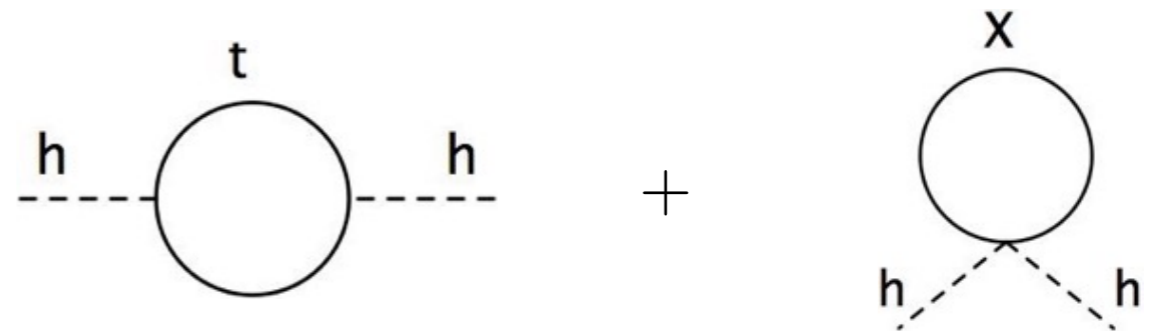
SIMP DM

Hochberg, Kuflik, Volansky, Wacker, 2014
+ Murayama, 2015

- Relic density from $3 \rightarrow 2$ annihilation
 - Suggests $M_{\text{DM}} \sim 100 \text{ MeV}$
 - (WIMP miracle suggests $M_{\text{DM}} \sim 1 \text{ TeV}$)
- SIMPlest model:
 - SU(N) dark sector with $N_d, n_f \geq 3$
 - $3 \rightarrow 2$ annihilation of dark pions through WZW term

Naturalness

- UV sensitivity of Higgs mass from top loop
- Naturalness requires coloured top partners X



Chacko, Goh, Harnik, 2005

- Alternative: Twin Higgs

- ▶ Z_2 symmetric “mirror world”
- ▶ Higgs mass regulated by top partners with dark colour

- ▶ Minimal model:

Craig, Katz, Strassler, Sundrum, 2015

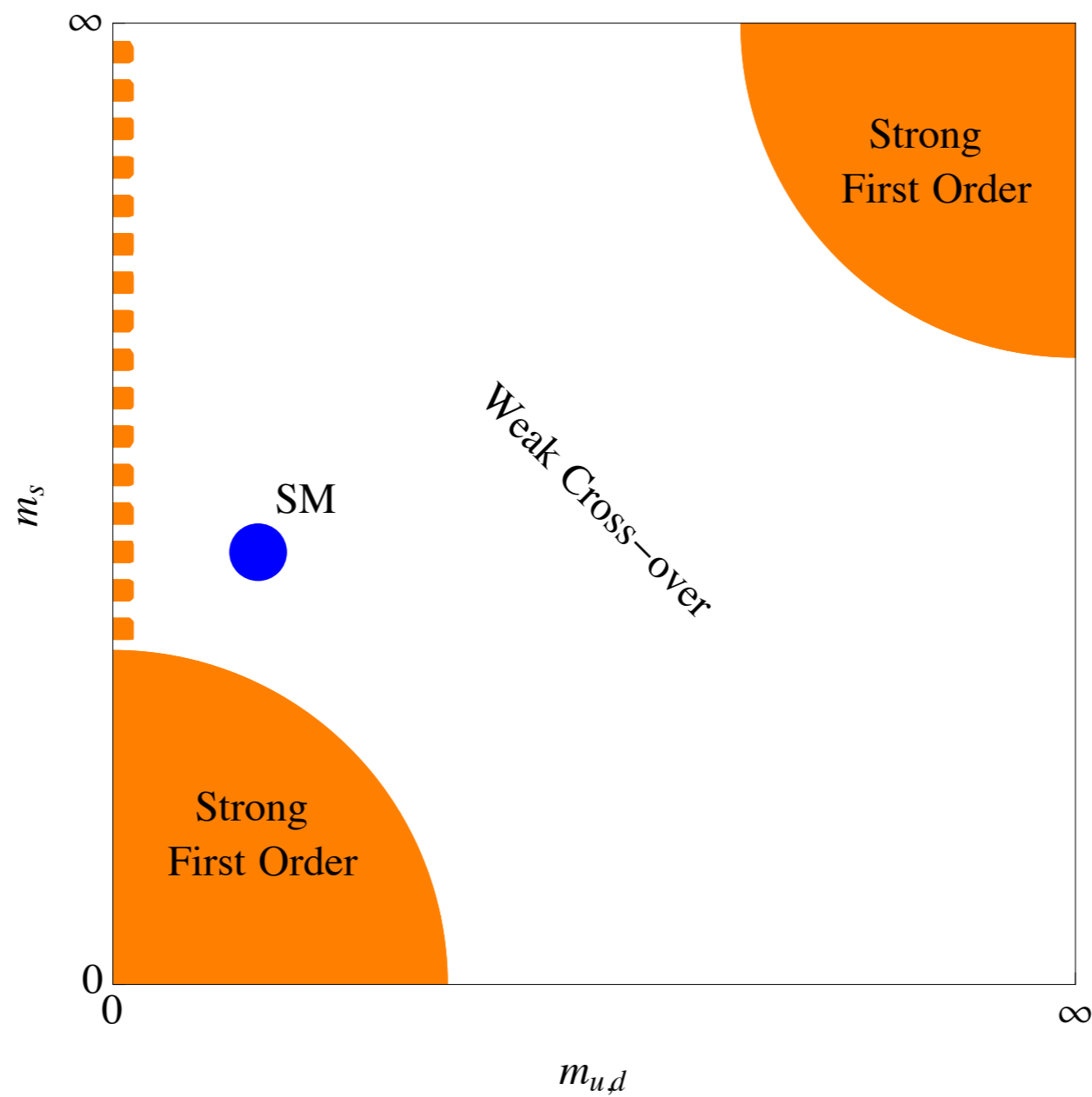
3rd gen. partners only - pure YM dark sector below m'_b

The Dark Phase Transition

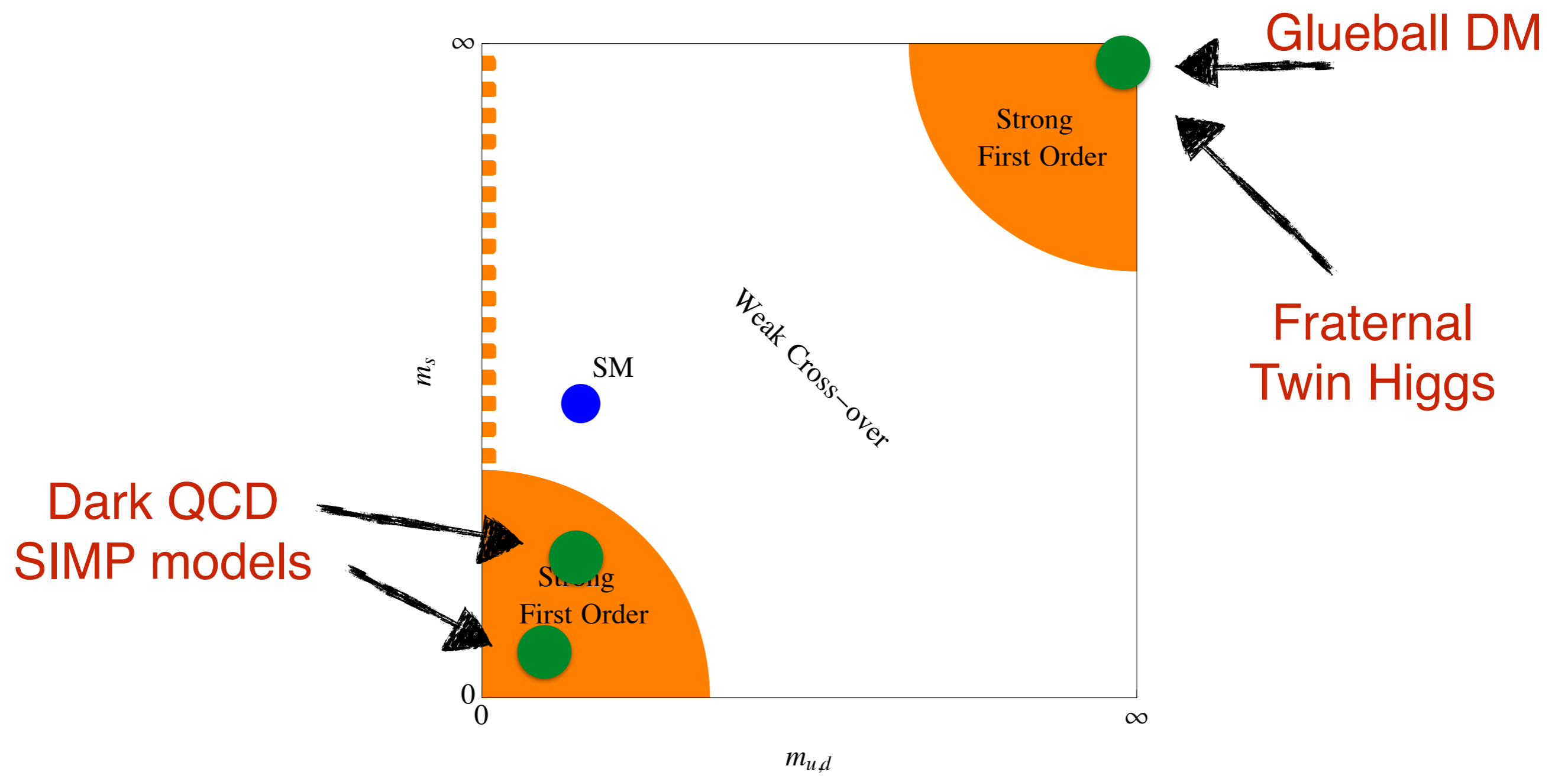
Phase Transition

- SU(N) dark sectors well motivated
- Confinement/chiral symmetry breaking phase transition at scale Λ_d
 - ▶ DM: $\Lambda_d \sim M_{\text{DM}}$ (MeV - 100 TeV)
 - ▶ Naturalness: $\Lambda_d \sim \text{few} \times \Lambda_{\text{QCD}}$
- First order PT in large class of models
- Still possible if LHC finds no new physics

QCD Phase Diagram



Phase Diagram II



SU(N) - PT

- Consider $SU(N_d)$ with n_f massless flavours
- PT is first order for
 - ▶ $N_d \geq 3$, $n_f = 0$
 - ▶ $N_d \geq 3$, $3 \leq n_f < 4N_d$
- Not for:
 - ▶ $n_f = 1$ (no global symmetry, no PT)
 - ▶ $n_f = 2$ (not yet known)

Svetitsky, Yaffe, 1982
M. Panero, 2009

Pisarski, Wilczek, 1983

SU(N) - PT 2

- One more parameter: Θ angle
- Effect on PT not well studied
- N_d, n_f dependence of PT strength?
- Finite density/chemical potentials?

M. Anber, 2013
Garcia-Garcia, Lasenby, March-Russell, 2015

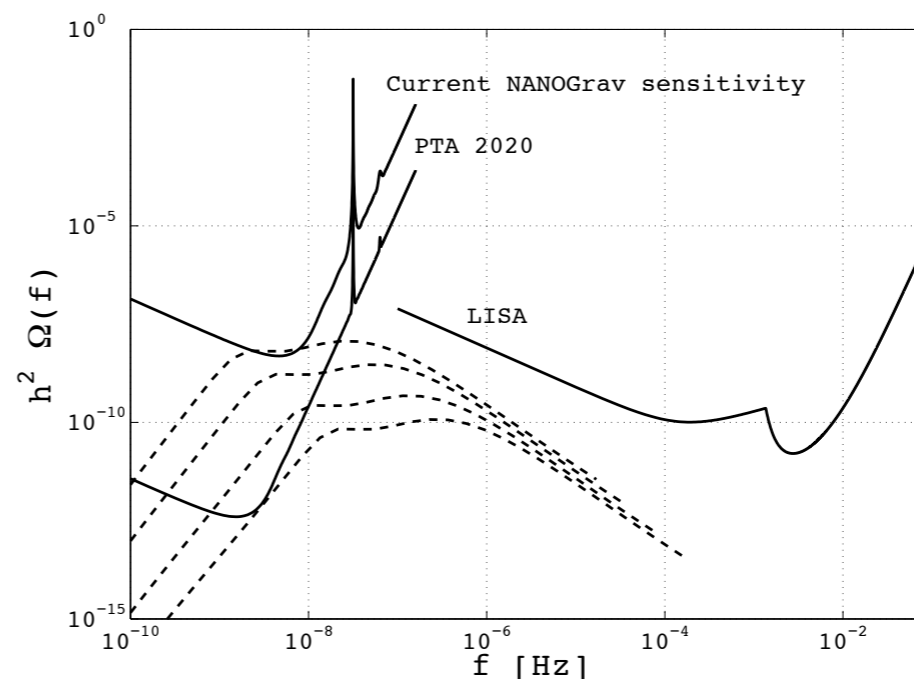
Panero, 2009

- ▶ QCD FOPT?

Schwarz, Stuke, 2009

- ▶ GW signal:

Caprini, Durrer, Siemens, 2009



GW signals

GW spectra

- Lot of work on GW from 1st order PT
 - Still difficult to simulate or model
- Here in addition:
 - Transition is non-perturbative
 - Parameters not known - take an optimistic guess

$$\beta/H_* = 1 - 100$$

$$v = 1$$

$$\frac{\kappa\alpha}{1 + \alpha} = 0.1$$

Shape

- From Bubble collisions and turbulence

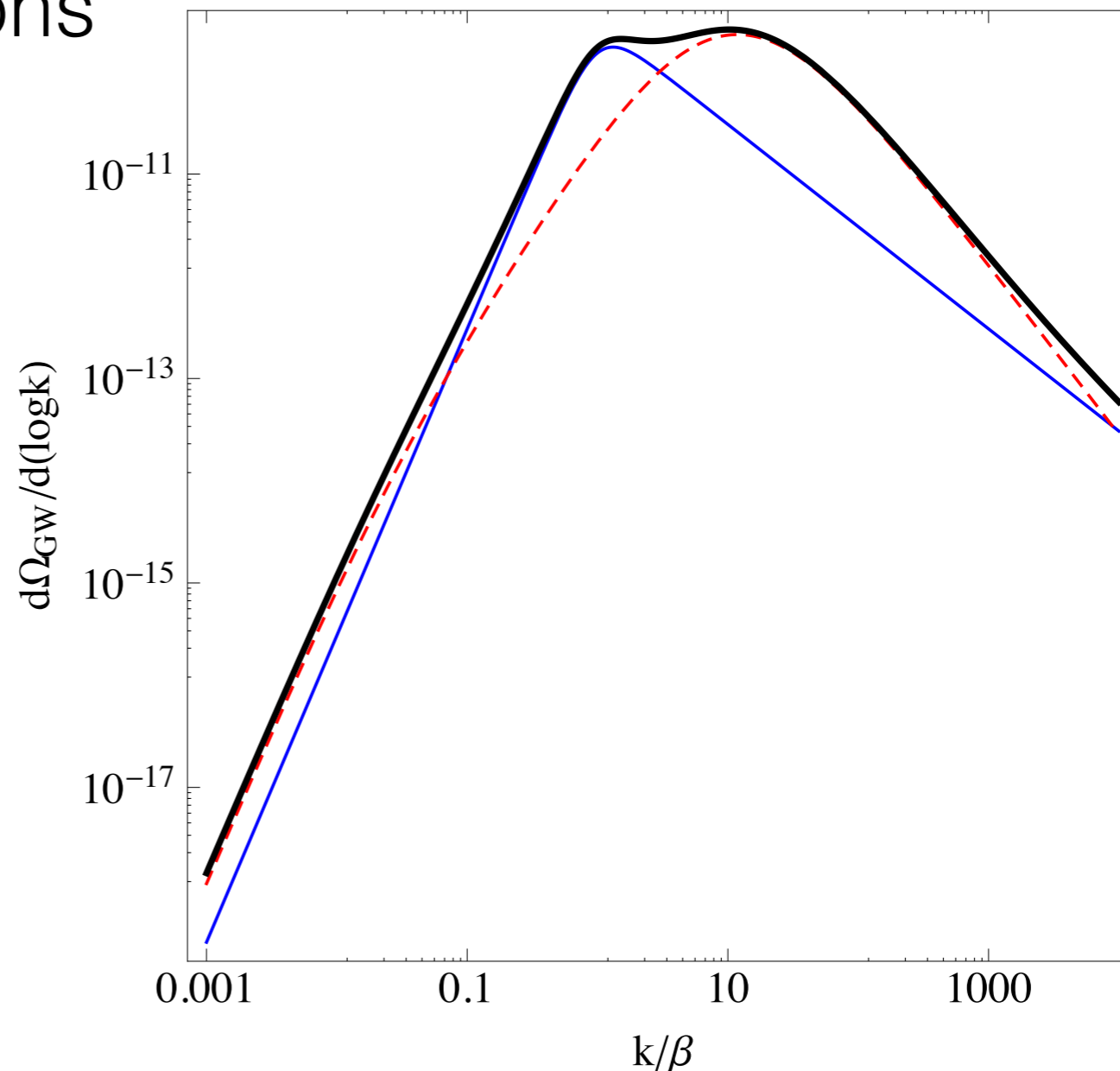
$$\frac{d\Omega_{\text{GW}}^{(B)} h^2}{d \log k} \simeq \frac{2}{3\pi} h^2 \Omega_{r0} \left(\frac{\mathcal{H}_*}{\beta}\right)^2 \Omega_{S_*}^2 v^3 \frac{(k/\beta)^3}{1 + (k/\beta)^4},$$

$$\frac{d\Omega_{\text{GW}}^{(MHD)} h^2}{d \log k} \simeq \frac{8}{\pi^6} h^2 \Omega_{r0} \left(\frac{\mathcal{H}_*}{\beta}\right) \Omega_{S_*}^{3/2} v^4 \frac{(k/\beta)^3}{(1 + 4k/\mathcal{H}_*)(1 + (v/\pi^2)(k/\beta))^{11/3}}$$

Caprini, Durrer, Siemens, 2010
Huber, Konstandin, Servant, ...

- k conformal wave number

$$\beta \sim (1 - 100) \mathcal{H}_*$$



Sound waves not included yet!

Hindmarsh, Huber, Rummukainen, Weir, 2013, 2015

Frequency - DM scale

- Redshift:

$$f = \frac{a_*}{a_0} H_* \frac{f_*}{H_*} = 1.59 \times 10^{-7} \text{ Hz} \times \left(\frac{g_*}{80} \right)^{\frac{1}{6}} \times \left(\frac{T_*}{1 \text{ GeV}} \right) \times \frac{f_*}{H_*}$$

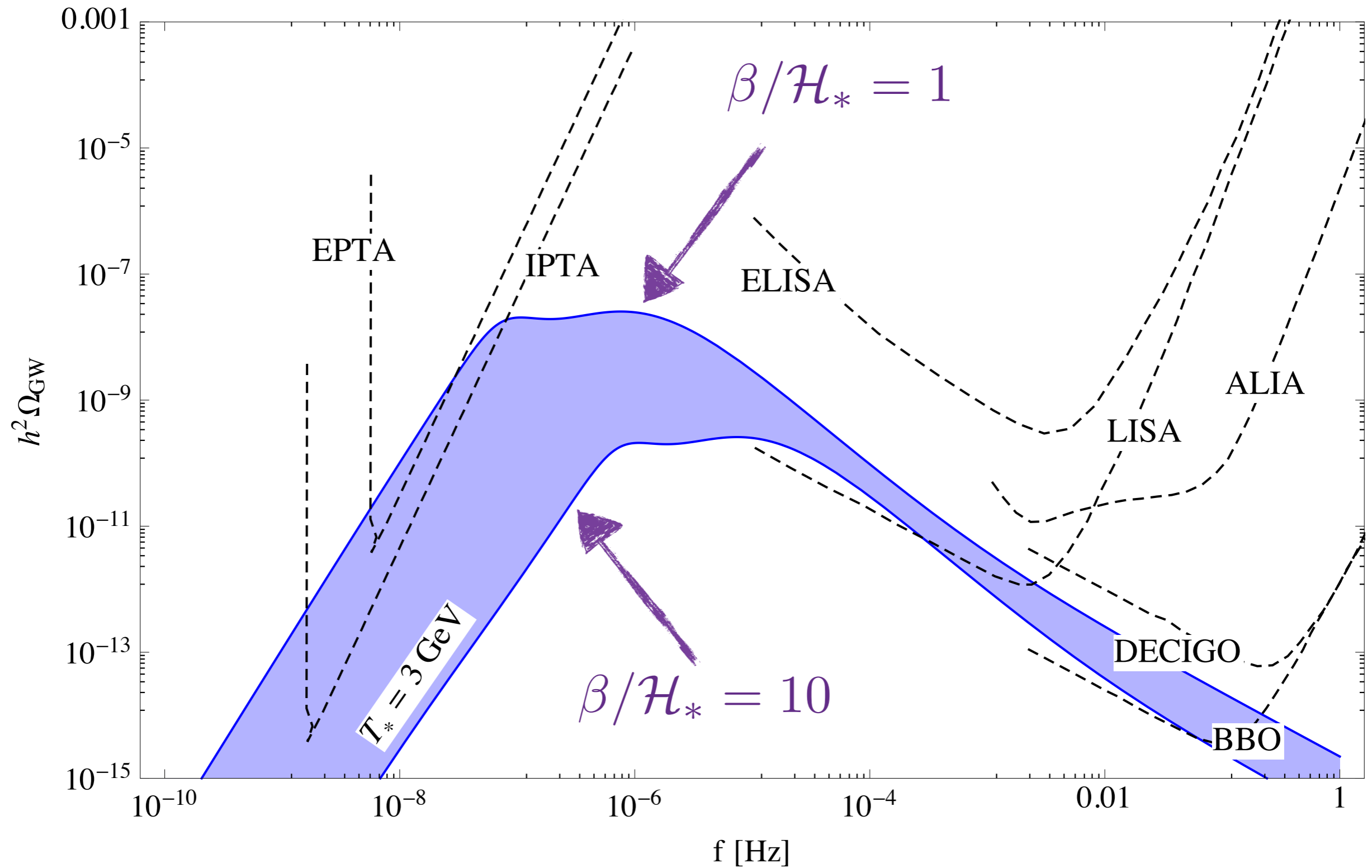
- Peak regions: $k/\beta \approx (1 - 10)$

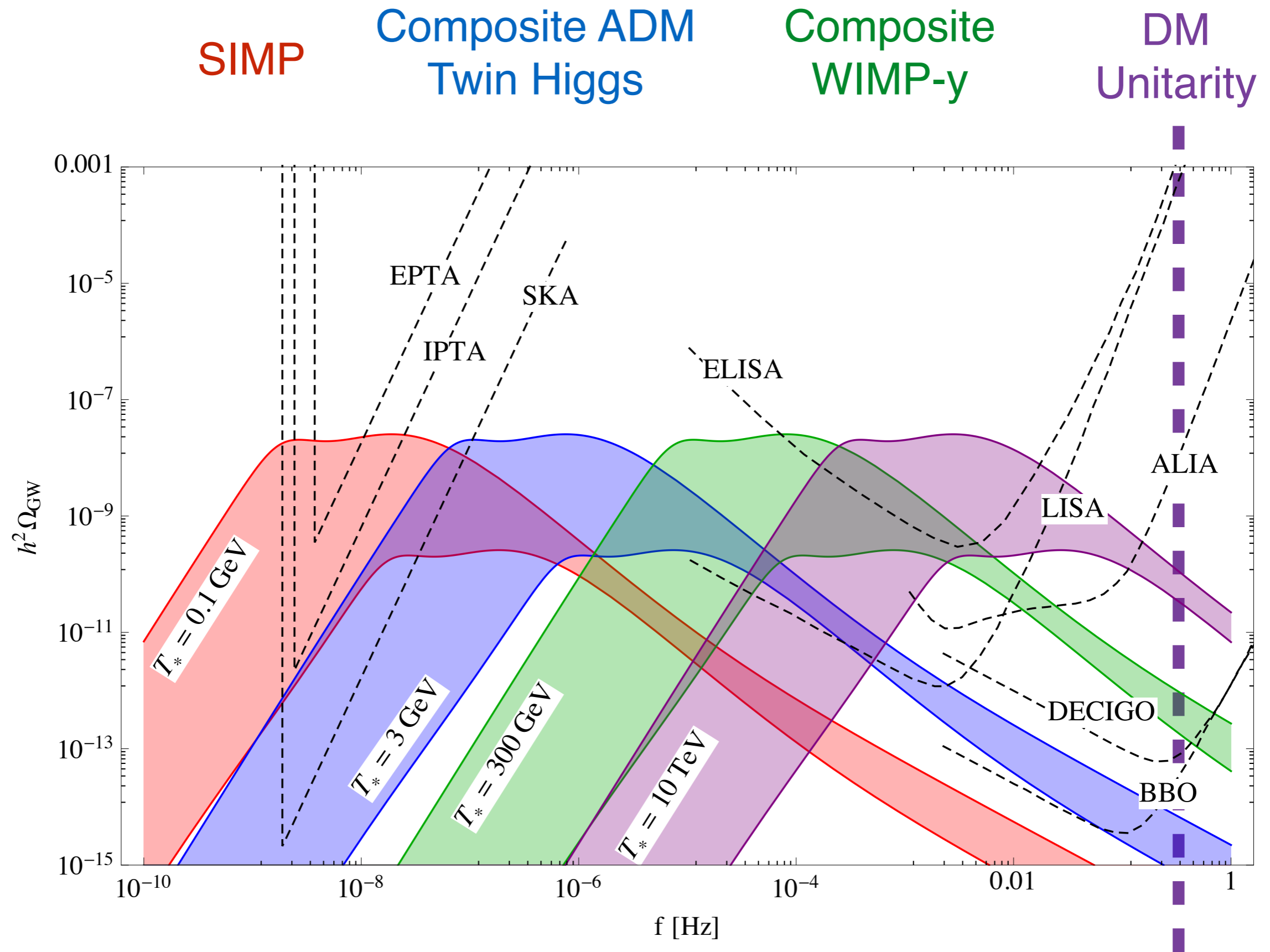
DM mass



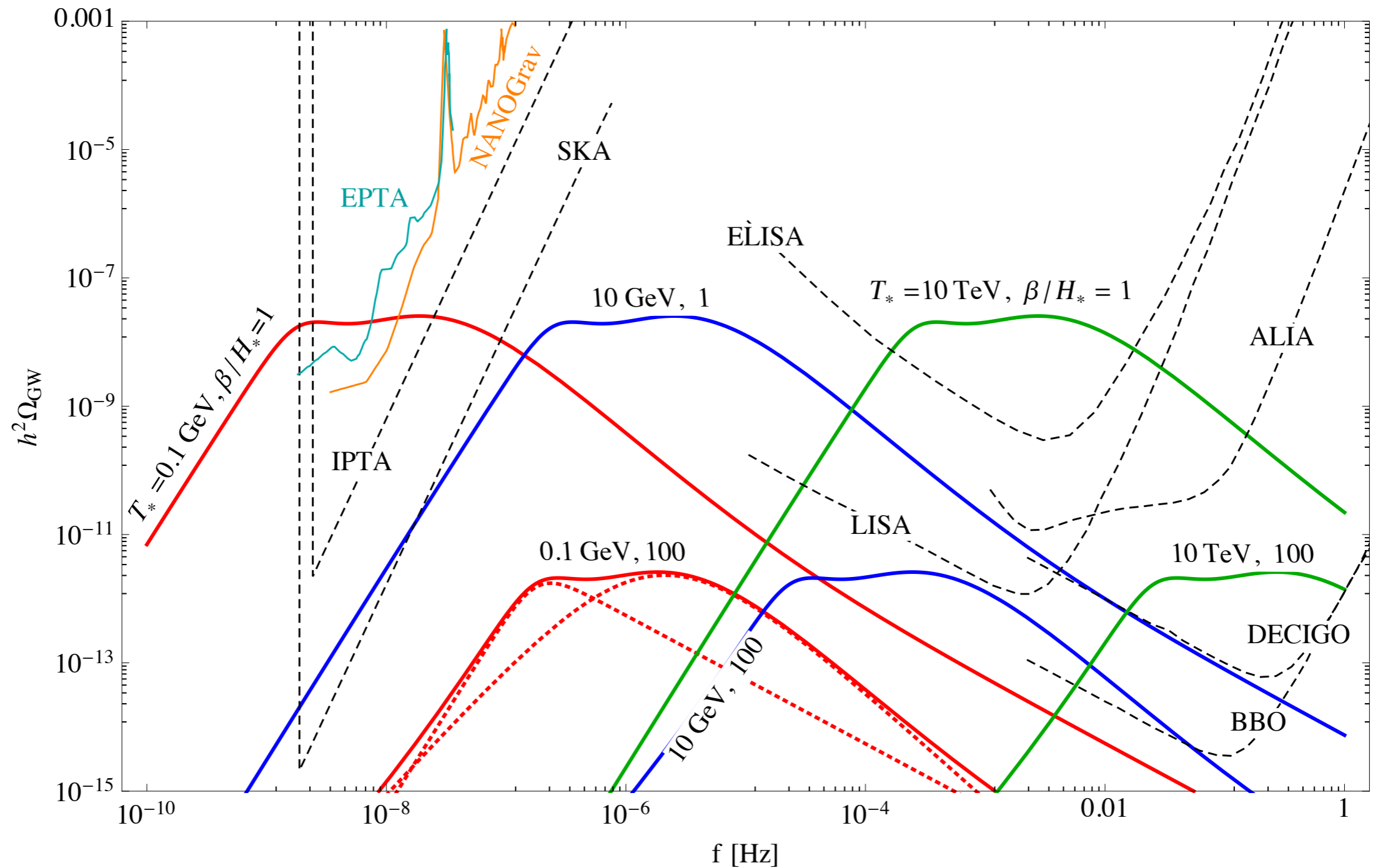
$$f_{\text{peak}}^{(B)} = 3.33 \times 10^{-8} \text{ Hz} \times \left(\frac{g_*}{80} \right)^{\frac{1}{6}} \left(\frac{T_*}{1 \text{ GeV}} \right) \left(\frac{\beta}{\mathcal{H}_*} \right)$$

$T^* \sim \text{Few GeV}$

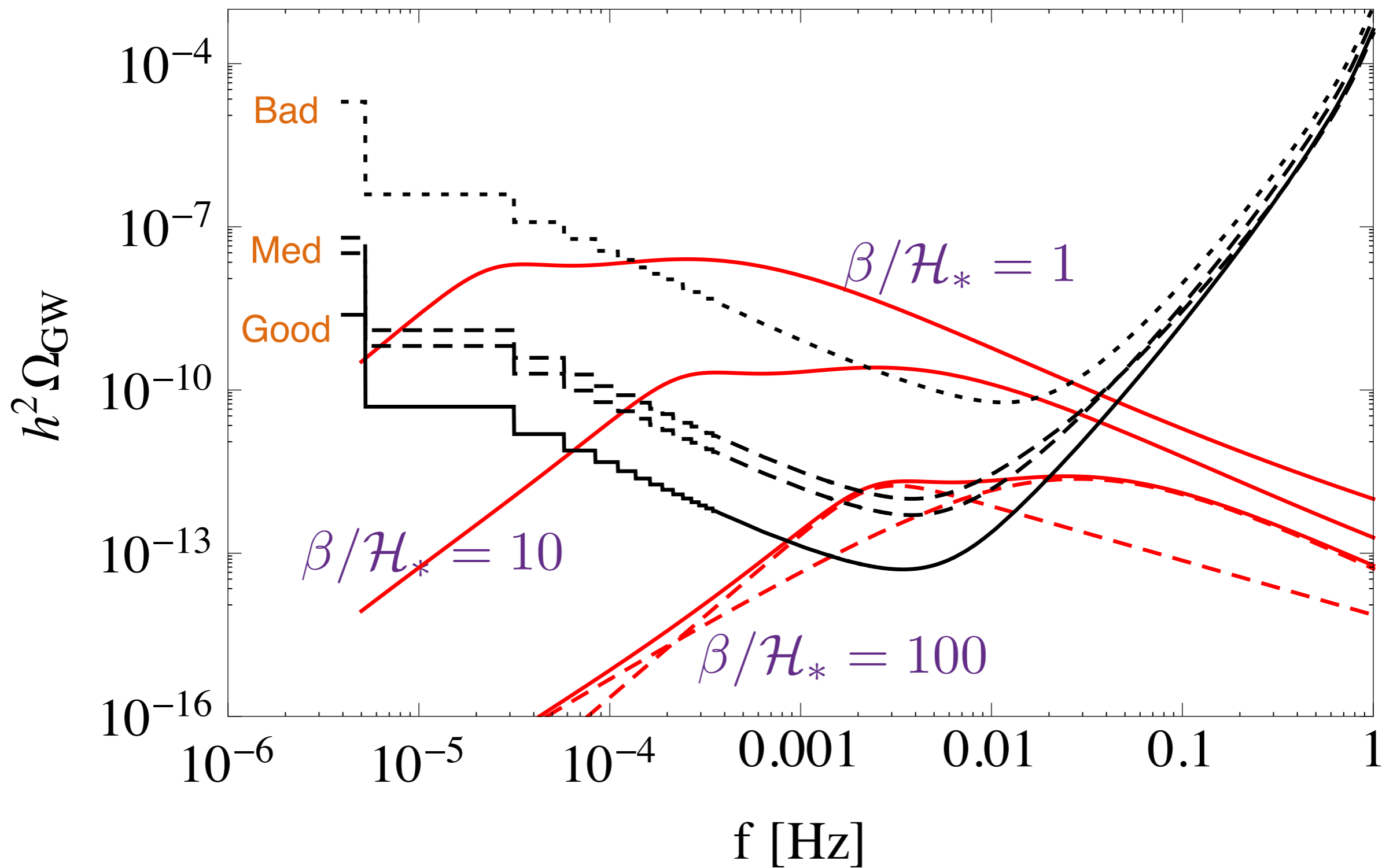




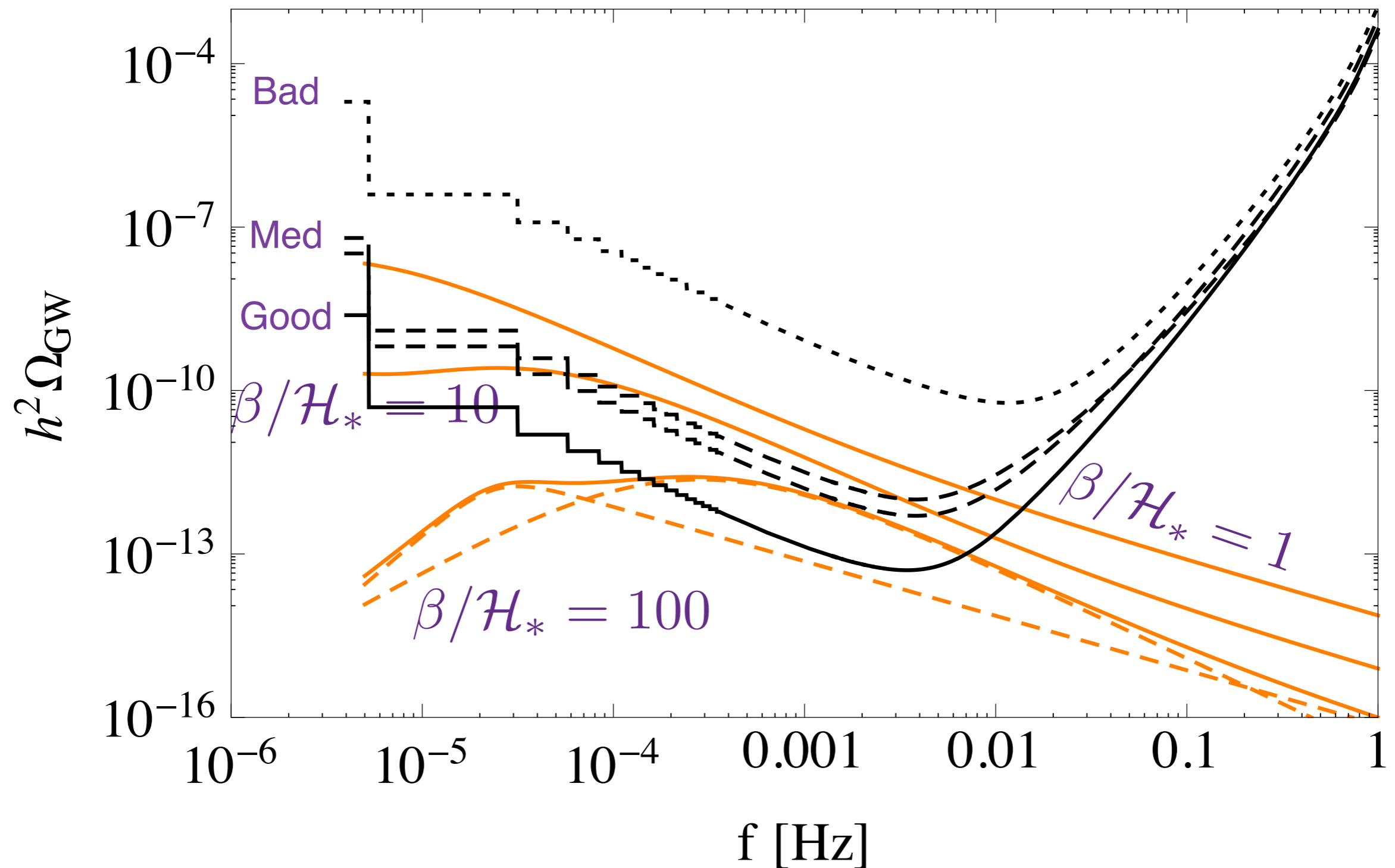
Broader Range of Signals



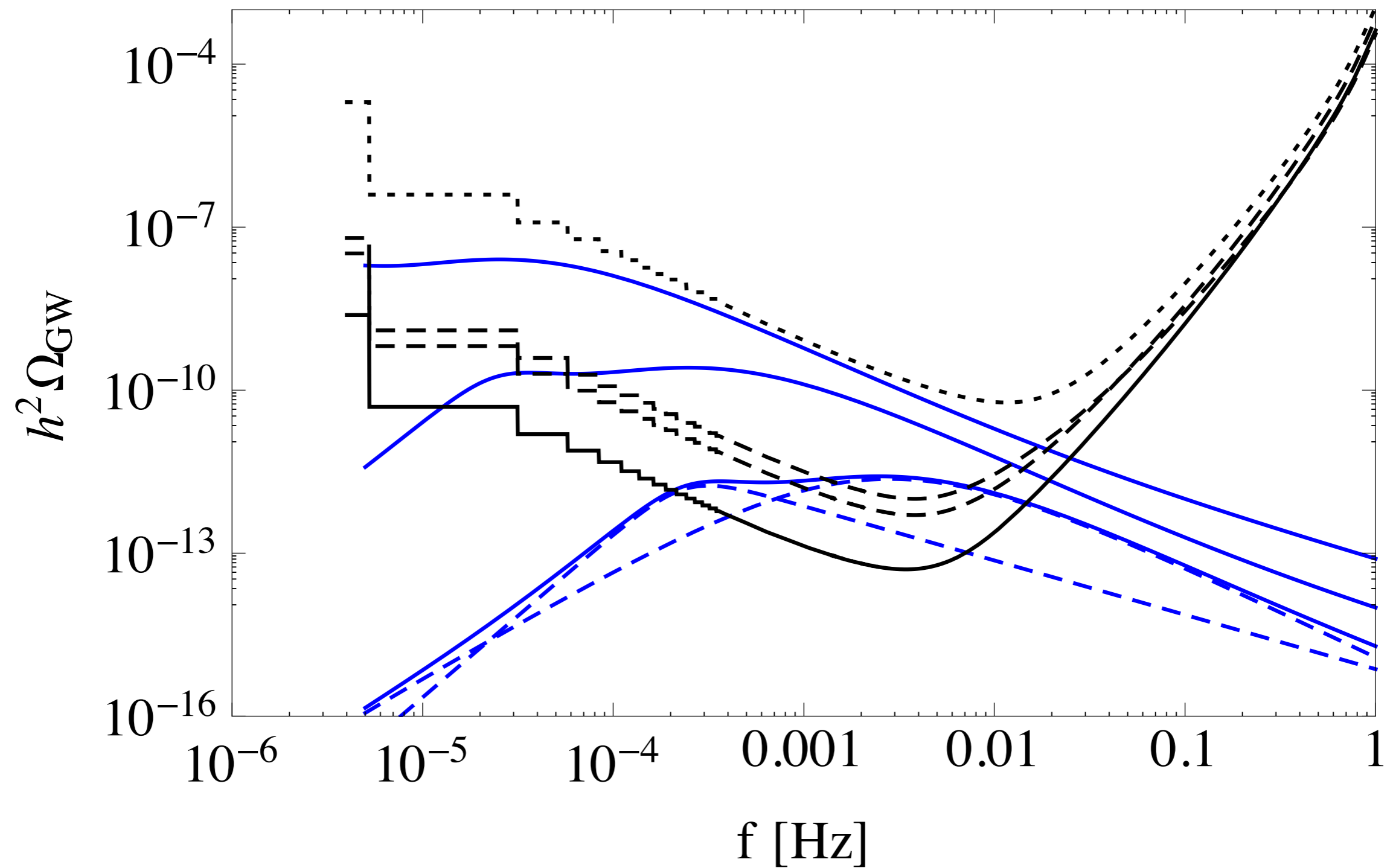
ELISA plots: $T^* = 1$ TeV



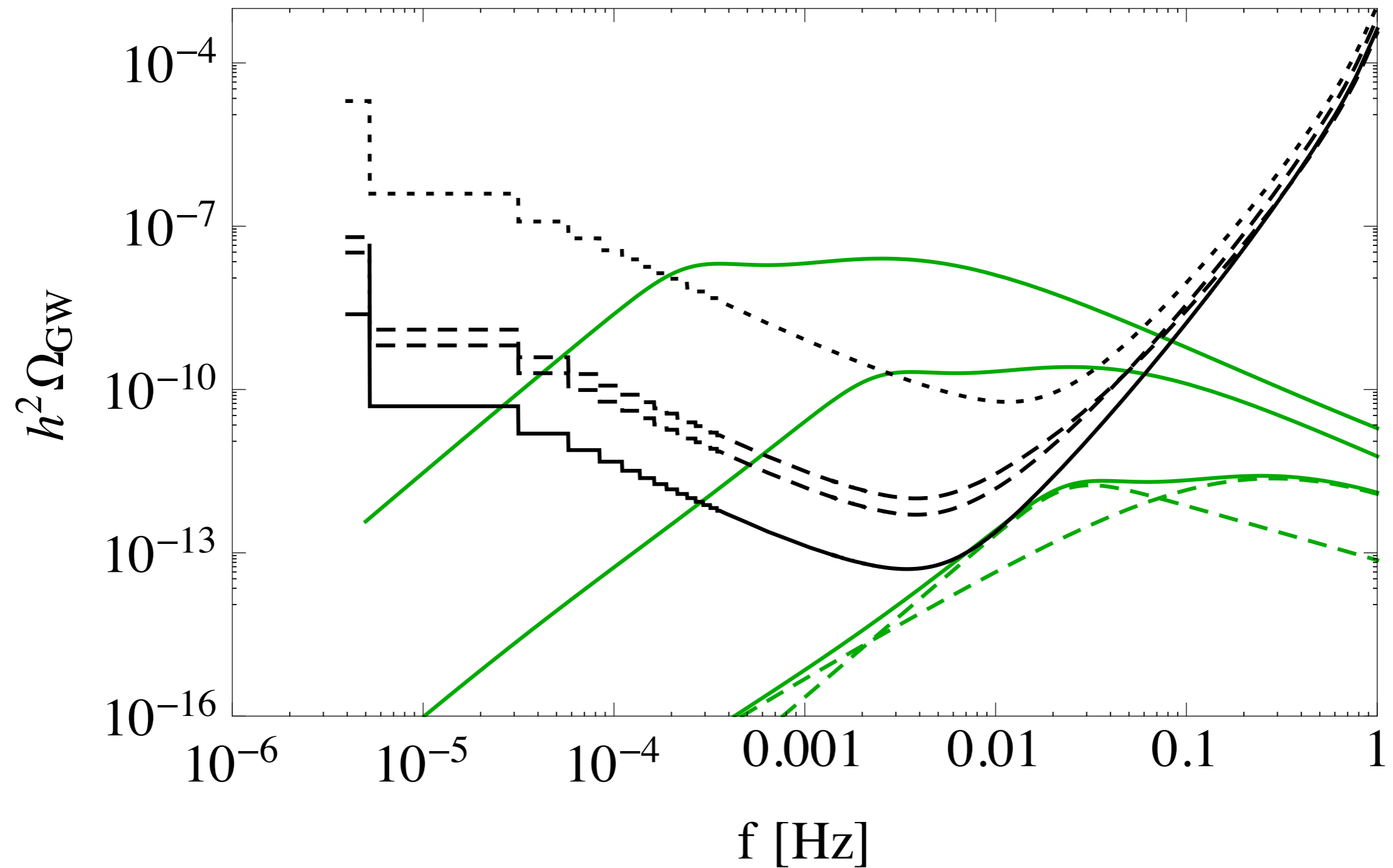
ELISA plots: $T^* = 10$ GeV



ELISA plots: $T^* = 100$ GeV



ELISA plots: $T^* = 10$ TeV



Signal Strength

- Strong signal motivated by radion stabilized RS models (holographic PT)?
- $n_f = 0$ case doable on the lattice
 - ▶ Signal as function of N_d . No other parameters!
 - ▶ Theta dependence?
- $n_f \geq 3$ difficult on the lattice
 - ▶ Holography? Large N?

Summary

- SU(N) dark sectors well motivated, often feature first order PT (also: Dark Baryogenesis!)
- Some Dark Matter models suggests GW signal in range of ELISA
 - ▶ Less dependent on LHC results, frequencies above/below strong EWPT scenario possible
 - ▶ GW - DM connection, possible first trace of very heavy DM
 - ▶ More precise study (lattice) needed for accurate predictions