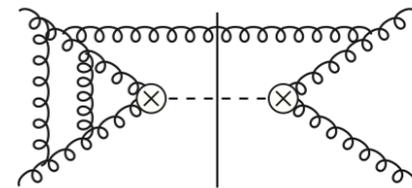
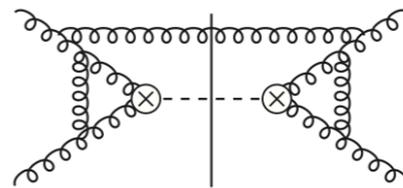
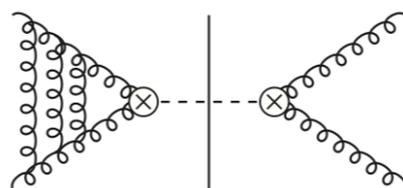
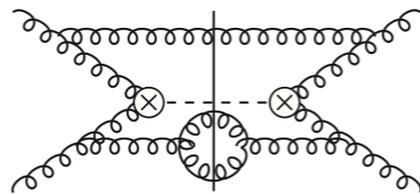
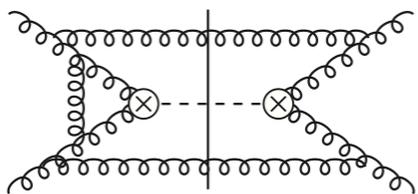


# Physics at the LHC after the Higgs discovery: what's next?

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PH Department  
CERN

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$$|b(\tau, \varepsilon, a, b)| \leq 2$$

$$\varphi(\sigma_1 t) \varphi(\sigma_2 t) = \varphi(\sqrt{\sigma_1^2 + \sigma_2^2} t)$$

$$P(\omega) = \frac{\sum_{k=1}^r P_k^* \log_2 \frac{1}{P_k}}{\sum_{k=1}^r P_k^*}$$

$$i_k \sigma_k^2 = \lambda_i \quad c_i k$$

$$y = \phi(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} e^{-\frac{t^2}{2}} dt$$

$$W_k = \binom{n}{k} p^k (1-p)^{n-k}$$

$$P(\eta < y | \xi = x) = \sup_{y' < y, y' \in R} P(\eta < y' | \xi = x)$$

$$S(\alpha, \tau) = \frac{2}{\pi} \int_0^{\pi} \frac{\sin \alpha t}{t} dt$$

$$P(\eta_{\infty} < x) = F(x)$$

$$\sum_{k=1}^r \int_{b_k}^{x+b_k} \left( \int_0^b \Psi_k^*(\tau) d\tau \right) dt - x \int_0^{b_k} \Psi_k^*(\tau) d\tau = \frac{x^2}{2} B(\omega) + \int_0^x (x-u) \sum_{k=1}^r \Psi_k^*(u) du \quad A(\omega) = \sum_{k=1}^r b_k \Psi_k^*(b_k \omega)$$

$$\log \varphi(u) = -\frac{\sigma^2 u^2}{2}$$

$$i^2 = -1; j^2 = -1; k^2 = -1 \quad \lim_{n \rightarrow \infty} \frac{\binom{2n}{n+c}}{\binom{2n}{n}} = e^{-2c}$$

$$S_n = A_n U \Pi A_n$$

$$|A_n| = \frac{n!}{2} \left| \int_{|x|>A} f(x) \log_2 \frac{1}{f(x)} dx \right| < \varepsilon \quad g^{-1} \cdot g = e$$

$$\int_{+\infty}^{-\infty} dG_k(x) \geq \frac{1}{2} \sum_{k \rightarrow \infty} e^{-\frac{k^2 \pi^2}{x^2}} = H(k)$$

$$\prod_{k \leq b}; \bigcup_{i=1}^{n-1} M_i; \bigcap_{n=0}^{\infty} X_n$$

$$y = \sqrt{\frac{\lambda u}{\nu u}} \left( \frac{\eta_{2u}}{\sqrt{\lambda u}} + \frac{\eta_{2u} - \eta_{2u}}{\sqrt{\lambda u}} \right)$$

$$f(t|y) = \frac{2e^{\frac{y^2}{2}}}{\sqrt{2\pi}} \int_{\frac{y}{t}}^{+\infty} \frac{e^{-\frac{u^2}{2}} du}{\left(1 - \frac{y^2}{u^2}\right)^{\frac{3}{2}}}$$

$$DN = \sum_{k=1}^N \frac{E_k}{u}$$

$$f_{n-1}(t) = \int_0^1 f_n(u) f_1(t-u) du = \frac{\lambda^{n+1} t^n e^{-\lambda t}}{n!} \quad \lim_{t \rightarrow 0} (f(t)) = 0$$

$$C_{iv} = \sum_{j=1}^n a_{ij} b_{jv}$$

$$\lim_{u \rightarrow +\infty} \frac{f(u)}{u} = P_e$$

$$R = \int_{-\infty}^{+\infty} \varphi(t) dt$$

$$U_n^+ = \binom{2n}{n} - \binom{2n}{n-c}$$

$$\log \varphi(t) = i\gamma t - c|t|^\alpha \left[ 1 + i\beta \frac{t}{|t|} \omega(t, \alpha) \right] \quad B(\omega) = \sum_{k=1}^r \Psi^*(b_k \omega)$$

$$\int_{-\infty}^{+\infty} e^{-\frac{u^2}{2}} du = F(x) \left( \frac{1}{\sqrt{2\pi}} \right)^{-1}$$

$$|\Psi_{\xi}(t)| = \left| \int_{-\infty}^{+\infty} e^{itx} dF(x) \right| \leq \int_{-\infty}^{+\infty} e^{-\nu x} dF(x) = \varphi_{\xi}(i\nu)$$

$$g^{-1}Ng = \{g^{-1}ng | n \in N\}$$

$$Q = F^{-1}(q)$$

$$q_n(\alpha) = \frac{P_k^*}{\sum_{j=1}^r P_j^*} \quad P(\pi_2 =$$

$$\prod_m = \prod_r | \prod_{m-r}$$

$$|X \cup Y| = |X| + |Y| - |X \cap Y|$$

$$\lim_{n \rightarrow \infty} \frac{1}{n} h_n \left( \frac{x}{\sqrt{n}} \right) = \frac{1}{\sqrt{2\pi}} e^{-\frac{x^2}{2}}$$

$$P_n(k) = \frac{P_k^*}{P_k^*} \quad P \left( \limsup_{n \rightarrow \infty} \frac{|h_n|}{\sqrt{2n \log \log n}} \leq 1 \right) = 1$$

$$P(t) = 1 - \sqrt{1 - e^{-2t}}$$

$$Q(A) = \int_A \chi(\omega) dP \quad l'(\alpha) = -\log_2 \left( \frac{\sum_{k=1}^r P_k^* \log_2 \frac{1}{P_k}}{\sum_{k=1}^r P_k^*} - \left( \frac{\sum_{k=1}^r P_k^* \log_2 \frac{1}{P_k}}{\sum_{k=1}^r P_k^*} \right)^2 \right)$$

$$f_g(u_i) = f \left( \sum_{j=1}^{\dim V_2} a_{ji} v_j^- \right) = \sum_{j=1}^{\dim V_2} a_{ji} \left( \sum_{k=1}^{\dim V_3} b_{kj} w_k \right) \frac{\binom{2k}{k}}{2^k} \approx \frac{1}{\sqrt{\pi k}}$$

$$q \left( e^{-x} \sqrt{\frac{1-q}{uq}} - 1 \right) = -x \sqrt{\frac{q(1-q)}{u}} + o\left(\frac{1}{u}\right)$$

$$\prod_{k=1}^r \left[ g_k \left( \frac{t}{\sqrt{n}} \right) \right]^{N_0 \alpha_k} = e^{-\frac{t^2}{2}}$$

$$P_{j_k}^{(m)} = \sum_{c=0}^{\infty} P_{j_k}^{(r)} P_{k_k}^{(m-r)} \quad \frac{1}{2\pi} \int_{-\infty}^{+\infty} \operatorname{Re} \left\{ \varphi(t) \frac{e^{-ita} - e^{-itb}}{it} \right\} dt$$

$$\liminf_{N \rightarrow \infty} \int_{-\infty}^{+\infty} f_N(x)^\alpha dx \geq \int_{-\infty}^{+\infty} f(x)^\alpha dx$$

$$M(\sigma_j - 1) = \int_{-\infty}^{+\infty} |x-1|^\alpha e^{-x} dx$$

$$\lim_{N \rightarrow \infty} \int_{-A}^{+A} f_N(x) \log_2 \frac{1}{f_N(x)} dx = \int_{-A}^{+A} f(x) \log_2 \frac{1}{f(x)} dx \quad P(|\omega| > \varepsilon) \leq \frac{C_\varphi}{\log N}$$

$$D^2(J_n) \leq \frac{k}{n} + 2k \left( \frac{1}{2} \sum_{k=1}^n R(k) \right)$$

$$\det(M') = \det(M) + \det(M^*) = \det(M)$$

$$h(x, y) = \frac{1}{2\pi} \left[ \sqrt{2} e^{-\frac{x^2}{2}} - e^{-x^2} \right] \quad |M(\varepsilon_n, \varepsilon_m)| \leq C_2 \sqrt{\frac{n}{m-n}}$$

$$D^2(J_n) \leq \frac{k}{n} + 2k \left( \frac{1}{2} \sum_{k=1}^n R(k) \right)$$

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$$N_{\varepsilon_n - \varepsilon_k} = \binom{2n}{n+k} = \binom{2n}{n-k}$$

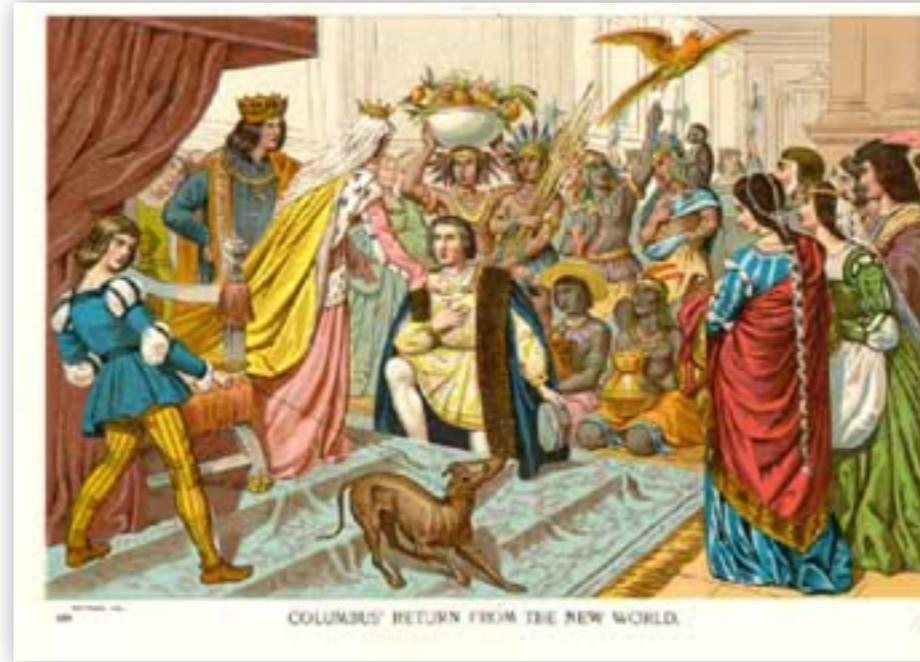
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# A “real” story from the past ...

Barcelona, 15 March 1493 .....



Columbus:

Your Majesty, the fleet needs an **upgrade**, we need to go back to the Indies with **10 times** more ships

King Ferdinand and Queen Isbellagnieszka:

You discovered the Indies, your theory is right, why do you need more?

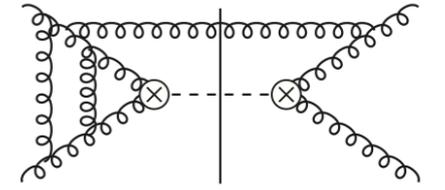
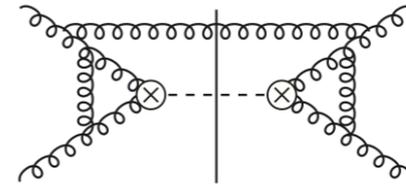
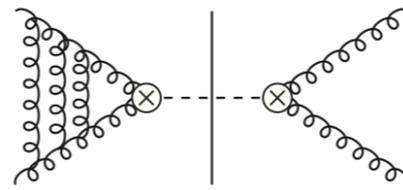
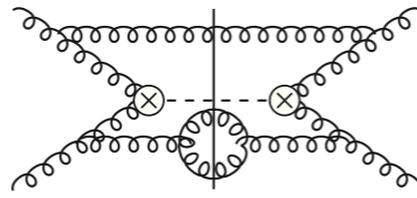
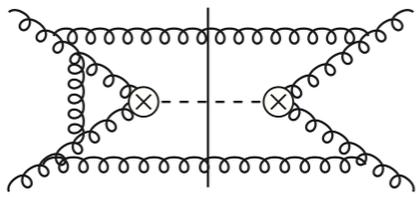
Columbus:

**Theorists\*** say these may not be the **standard Indies**. They calculated the Earth radius, and the standard Indies cannot be so close: these are likely to be **beyond the standard Indies** (*moving eastward ...*)

*\* If the King had listened to theorists to start with, he would have never authorized the mission: everyone one would have died of starvation well before reaching the “standard” Indies ...*

# The explorer's checklist

- Confirm you've landed in the right place
- Explore surroundings
  - look for what you expect
  - search for new “commodities”
- Push further the boundary of uncharted territories



$$|b(\tau, \varepsilon, a, b)| \leq 2$$

$$\varphi(\sigma_1 t) \varphi(\sigma_2 t) = \varphi(\sqrt{\sigma_1^2 + \sigma_2^2} t)$$

$$P(\omega) = \frac{\sum_{k=1}^r P_k^* \log_2 \frac{1}{P_k}}{\sum_{k=1}^r P_k^*}$$

$$i_k \sigma_k^2 = \lambda_i \quad c_i k$$

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$$W_k = \binom{n}{k} p^k (1-p)^{n-k}$$

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$$P(\eta_{\infty} < x) = F(x)$$

$$S_n = A_n U \Pi A_n$$

$$|A_n| = \frac{n!}{2} \left| \int_{|x|>A} f(x) \log_2 \frac{1}{f(x)} dx \right| < \varepsilon$$

$$\int_{-\infty}^{+\infty} dG_k(x) \geq \frac{1}{2} \sum_{k \rightarrow \infty} e^{-\frac{k^2 \pi^2}{x^2}} = H(k)$$

$$f_{n-1}(t) = \int_0^1 f_n(u) f_1(t-u) du = \frac{\lambda^{n+1} t^n e^{-\lambda t}}{n!}$$

$$\log \varphi(t) = i \gamma t - c |t|^\alpha [1 + i \beta \frac{t}{|t|} \omega(t, \alpha)] \quad B(\omega) = \sum_{k=1}^r \Psi^*(b_k \omega)$$

$$\int_{-\infty}^{+\infty} e^{-\frac{u^2}{2}} du = F(x) \left( \frac{1}{\sqrt{2\pi}} \right)^{-1}$$

$$\prod_{m=1}^n = \prod_{r=1}^n \prod_{m=r}^n$$

$$|X \cup Y| = |X| + |Y| - |X \cap Y|$$

$$f: X \rightarrow X \cap W$$

$$Q(A) = \int_A \chi(\omega) dP \quad l'(x) = -\log_2 \left( \frac{\sum_{k=1}^r P_k^* \log_2 \frac{1}{P_k}}{\sum_{k=1}^r P_k^*} - \left( \frac{\sum_{k=1}^r P_k^* \log_2 \frac{1}{P_k}}{\sum_{k=1}^r P_k^*} \right)^2 \right)$$

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$$D^2(J_n) \leq \frac{k}{n} + 2k \left( \frac{1}{2} \sum_{k=1}^n R(k) \right)$$

$$\sum_{k=1}^r \int_{b_k \omega}^{x+b_k \omega} \left( \int_0^b \Psi_k^*(\tau) d\tau \right) dt - x \int_0^{b_k \omega} \Psi_k^*(\tau) d\tau = \frac{x^2}{2} B(\omega) + \int_0^x (x-u) \sum_{k=1}^r \Psi_k^*(u) du \quad A(\omega) = \sum_{k=1}^r b_k \Psi^*(b_k \omega)$$

$$\eta_1 = \sum_{k=1}^n a_k \xi_k$$

$$\log \varphi(u) = -\frac{\sigma^2 u^2}{2}$$

$$i^2 = -1; j^2 = -1; k^2 = -1 \quad \lim_{n \rightarrow \infty} \frac{\binom{2n}{n+c}}{\binom{2n}{n}} = e^{-2c}$$

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$$f(t|y) = \frac{2e^{-\frac{y^2}{2}}}{\sqrt{2\pi}} \left( \frac{e^{-\frac{t^2}{2}}}{\left(1 - \frac{y^2}{u^2}\right)^{\frac{3}{2}}} \right)$$

$$\Delta N = \sum_{k=1}^N \frac{\varepsilon_k}{n}$$

$$\prod_{k \leq b}; \bigcup_{i=1}^{n-1} M_i; \bigcap_{n=0}^{\infty} X_n$$

$$f = \sqrt{\frac{\lambda u}{2n}} \left( \frac{\eta_{2n}}{\sqrt{2n}} + \frac{\eta_{2n} - \eta_{2n}}{\sqrt{2n}} \right)$$

$$U_n^+ = \binom{2n}{n} - \binom{2n}{n-c}$$

$$f_n(t) = \frac{\lambda^n t^{n-1} e^{-\lambda t}}{(n-1)!}$$

$$\lim_{n \rightarrow \infty} \frac{f_n(t)}{n} = P_e$$

$$R = \int_{-\infty}^{+\infty} \varphi(t) dt$$

$$C_{iv} = \sum_{j=1}^r a_{ij} b_{jv}$$

$$\lim_{n \rightarrow \infty} P \left( \frac{\sum_{j=1}^r (j_{n+1} - k_{jn}) - \log \frac{1}{q}}{\sqrt{\frac{1-q}{q}}} \right) C_n(\alpha) \geq \frac{n!}{\prod_{k=1}^n n_k(\alpha)!}$$

$$\frac{u}{m} \varphi(t) = \varphi\left(c \left(\frac{u}{m}\right) t\right)$$

$$\lim_{t \rightarrow 0} \frac{f(t)}{t} = 0$$

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$$\varphi(t) = 1 - \sqrt{1 - e^{2t}}$$

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$$\frac{\binom{2k}{k}}{2^k} \approx \frac{1}{\sqrt{\pi k}}$$

$$\prod_{k=1}^r \left[ g_k \left( \frac{t}{\sqrt{k}} \right) \right]^{N_0 \alpha_k} = e^{-\frac{t^2}{2}}$$

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$$P(|\omega_n| > \varepsilon) \leq \frac{C_\varphi}{\log N}$$

$$M(|\sigma_j - 1|^\alpha) = \int_0^\infty (x-1)^\alpha e^{-x} dx$$

$$N_{\varepsilon_1 - \varepsilon_2} = \binom{2n}{n+k_2} = \binom{2n}{n-k_1}$$

$$\det(M') = \det(M) + \det(M^*) = \det(M) \quad h(x, y) = \frac{1}{2\pi} \left[ \sqrt{2} e^{-\frac{x^2}{2}} - e^{-x^2} \right]$$

$$|M(\varepsilon_n, \varepsilon_m)| \leq C_2 \sqrt{\frac{n}{m-n}}$$

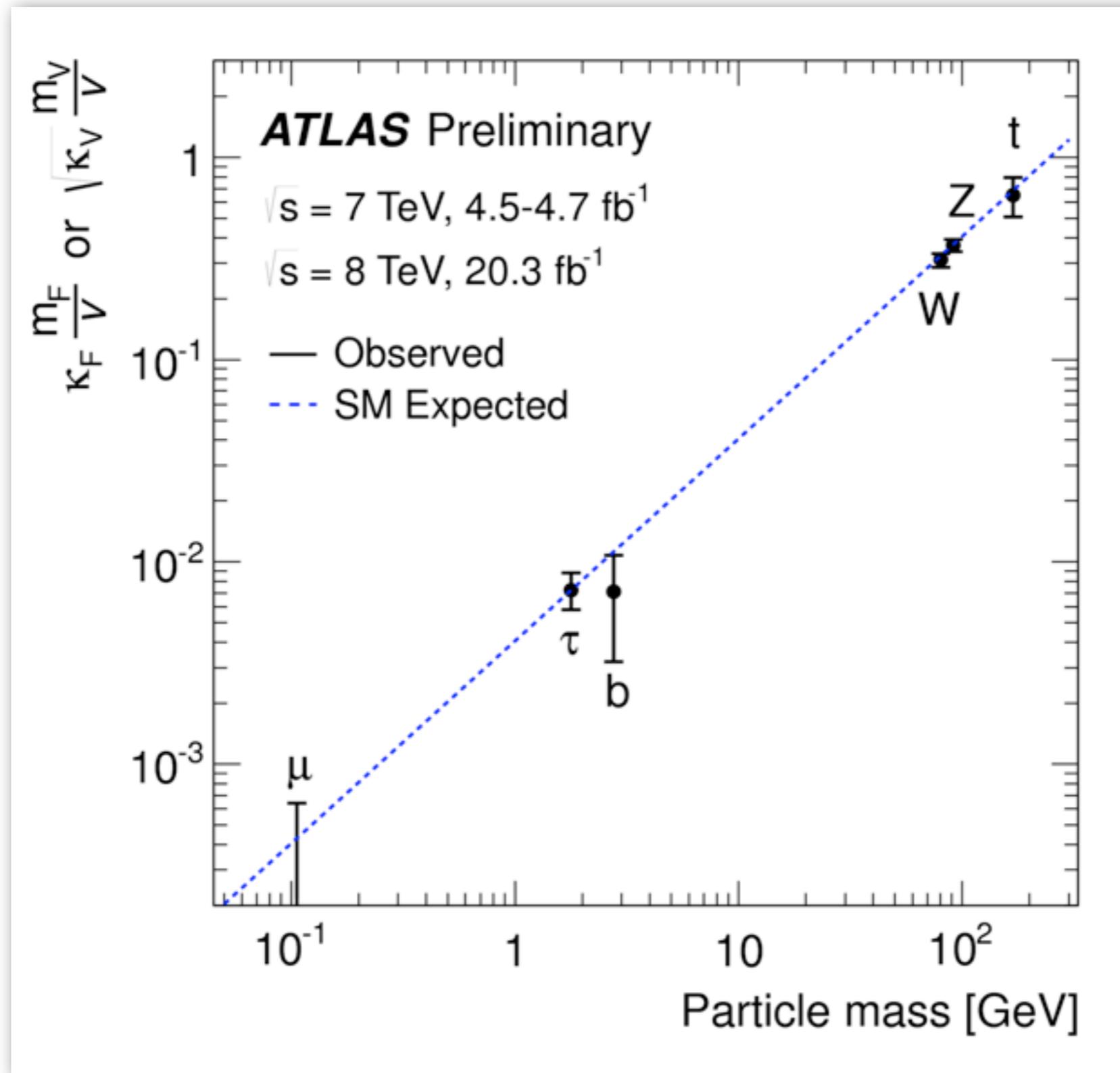
# The LHC checklist

- **Confirm you've landed in the right place**
  - ▶ *is the new particle at 125 GeV the Higgs of the Standard Model?*
- **Explore surroundings**
  - look for what you expect: the *guaranteed deliverables*
    - ▶ *systematically study and measure with ever-growing precision the interactions of the Higgs boson, and of the other SM particles*
  - search for new “commodities” - the *known unknowns*, follow a path ...
    - ▶ *dark matter, origin of neutrino masses, baryon asymmetry of the universe*
- **Push further the boundary of uncharted territories - search the Eldorado**
  - ▶ *supersymmetry, new forces, extra dimensions, ...*

*The Higgs discovery goal provided in the 90's the reference target for the design of the LHC accelerator and experiments.*

*But the physics goals and ambitions of the LHC always went beyond the Higgs discovery .....*

Run I of the LHC determined, with a precision of  $\pm 20\%$ , that the Higgs boson gives a mass to SM particles



# Open Higgs issues for run 2 and beyond

1. This limited precision, due to low statistics, is not sufficient to probe most possible scenarios alternative to the SM: **will the SM withstand more accurate tests?**

Example:  $BR[H \rightarrow \mu\tau] = (0.89 \pm 0.40)\%$  reported by CMS, needs more statistics to confirm (In the SM should be 0)

2. The Higgs mechanism has only been tested on a fraction of the SM particles, due to low statistics: **do the other particles (e.g. muon, charm, etc) interact with the Higgs as predicted by the SM?**

Example: more than  $300 \text{ fb}^{-1}$  required to establish  $H \rightarrow \mu\mu$  at  $5\sigma$

3. Neutrino masses are not a SM ingredient: **how do neutrinos acquire their mass?**

The LHC plays a role in exploring possible answers

4. **Are there more Higgs bosons?**

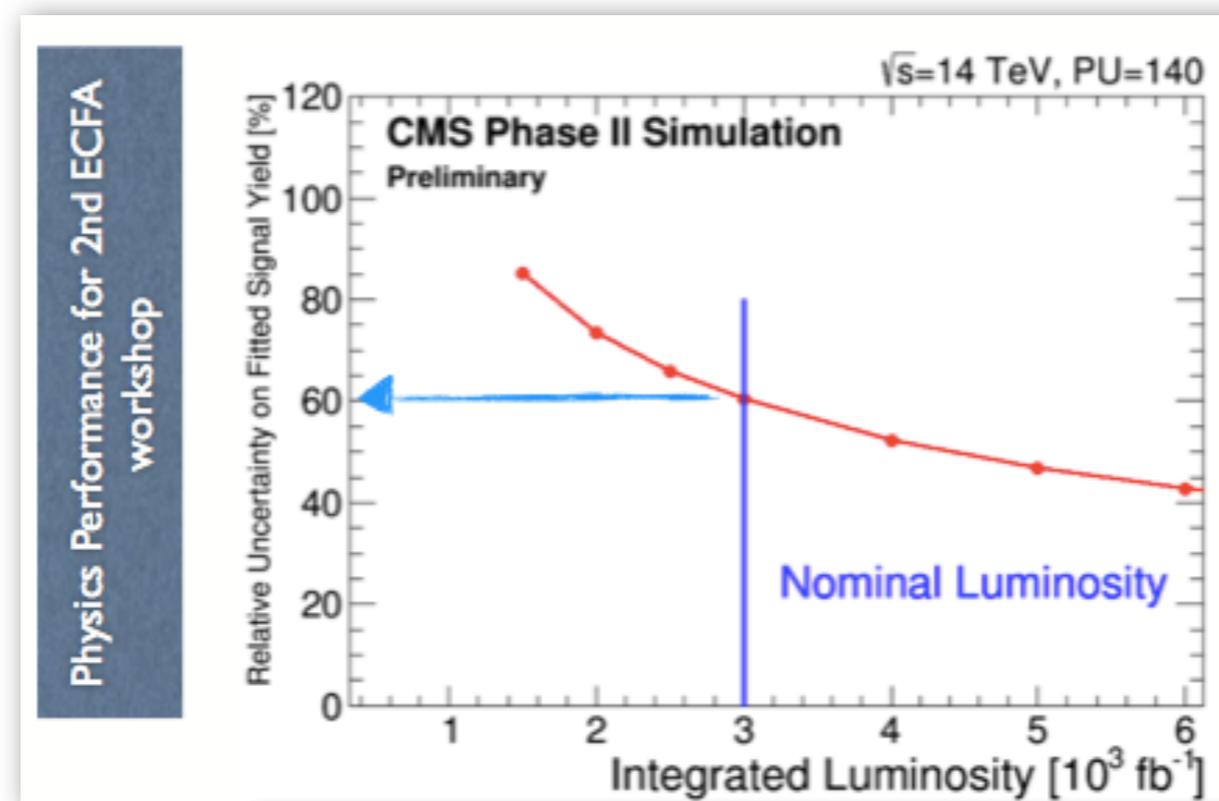
Most theories beyond the SM have more Higgs bosons

# 5. What gives mass to the Higgs ??

Obvious question, with a trivial answer in the SM: the Higgs gives mass to itself!

But less trivial answers can arise in beyond-the-SM scenarios

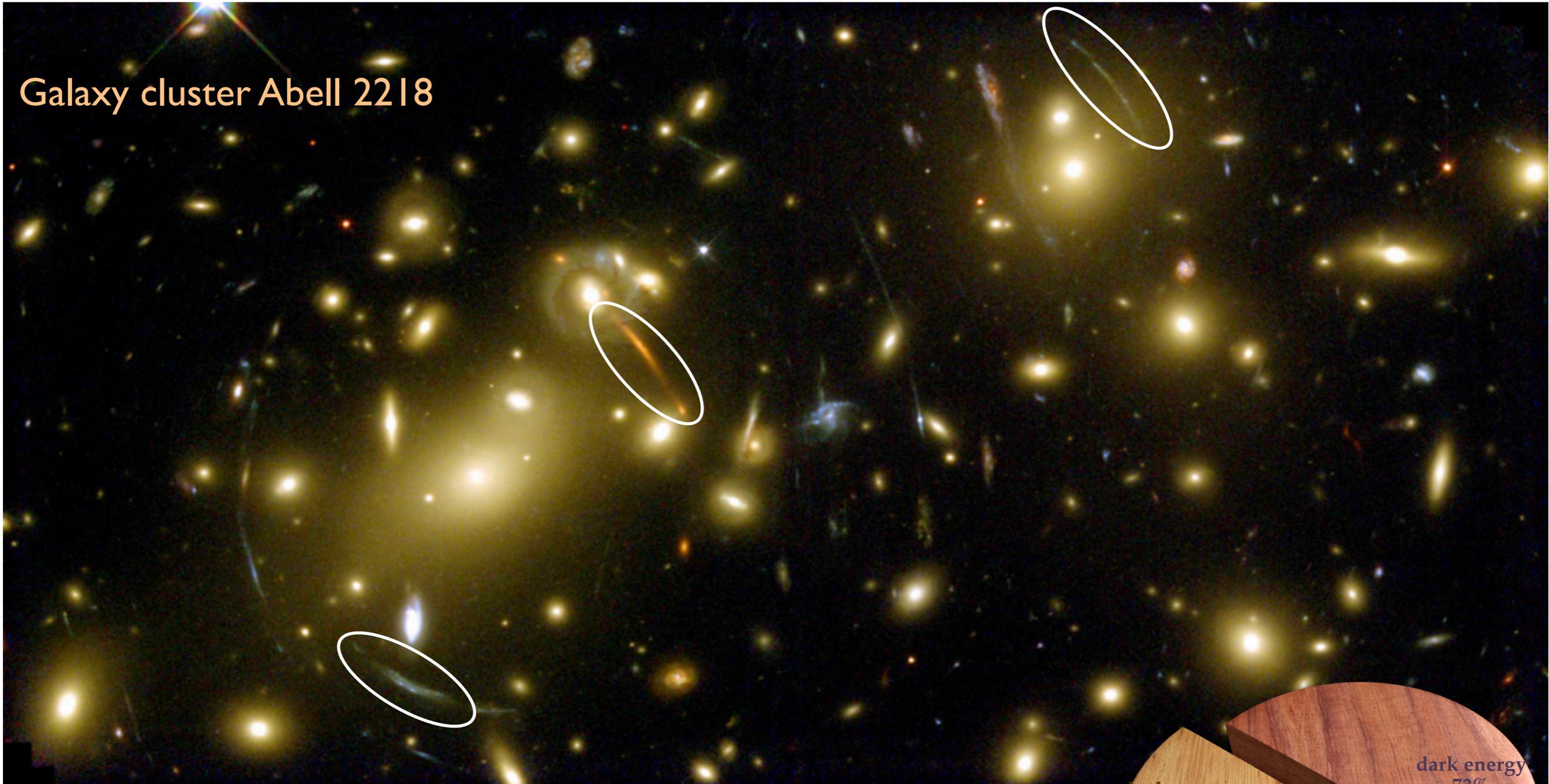
Testing how the Higgs interacts with itself (*this is how we probe the origin of the Higgs mass*) will require the full High Luminosity programme, and possibly more



The measurement of Higgs self-interactions has broad implications on issues such as the nature of the EW phase transition during the Big Bang

# What is Dark Matter?

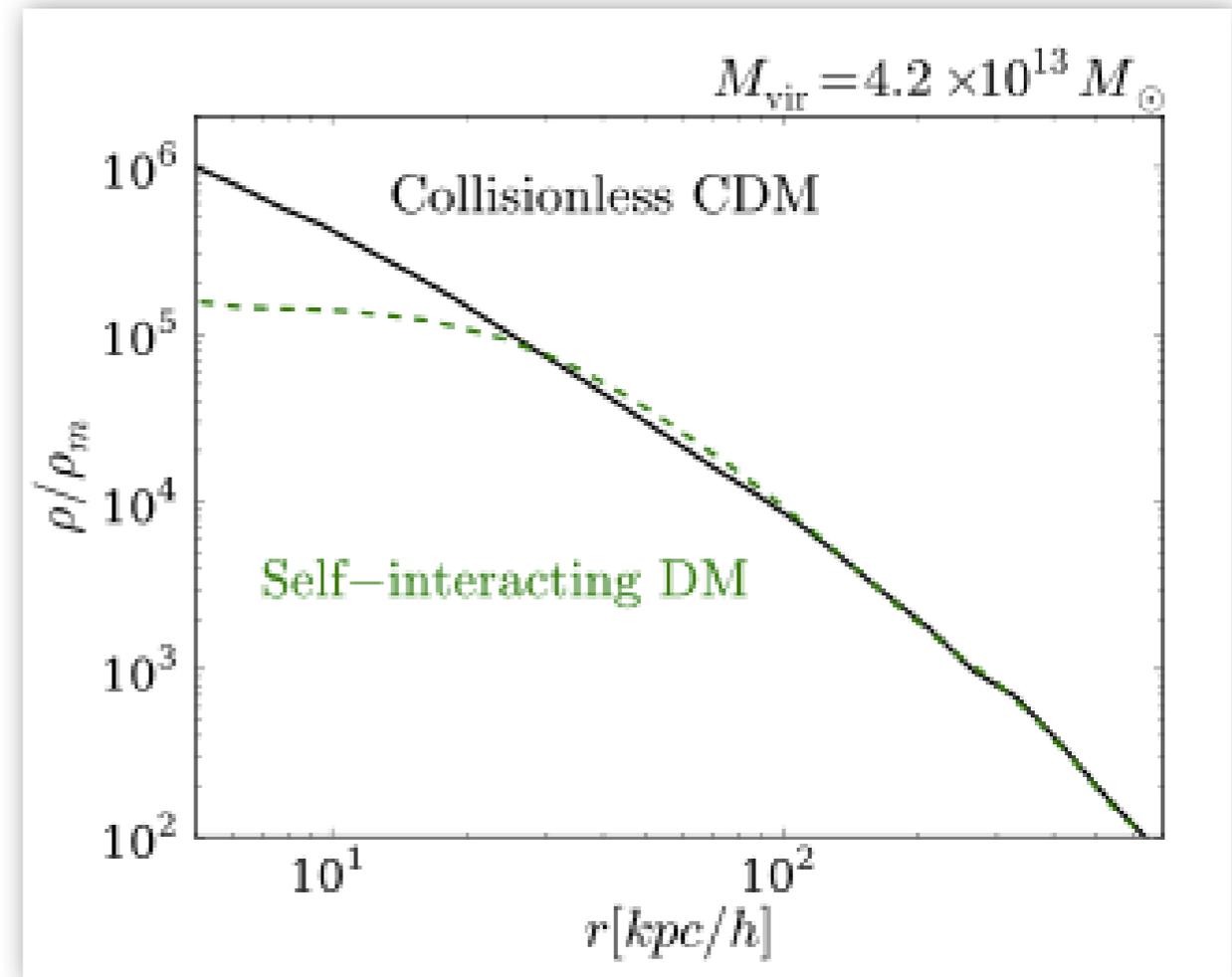
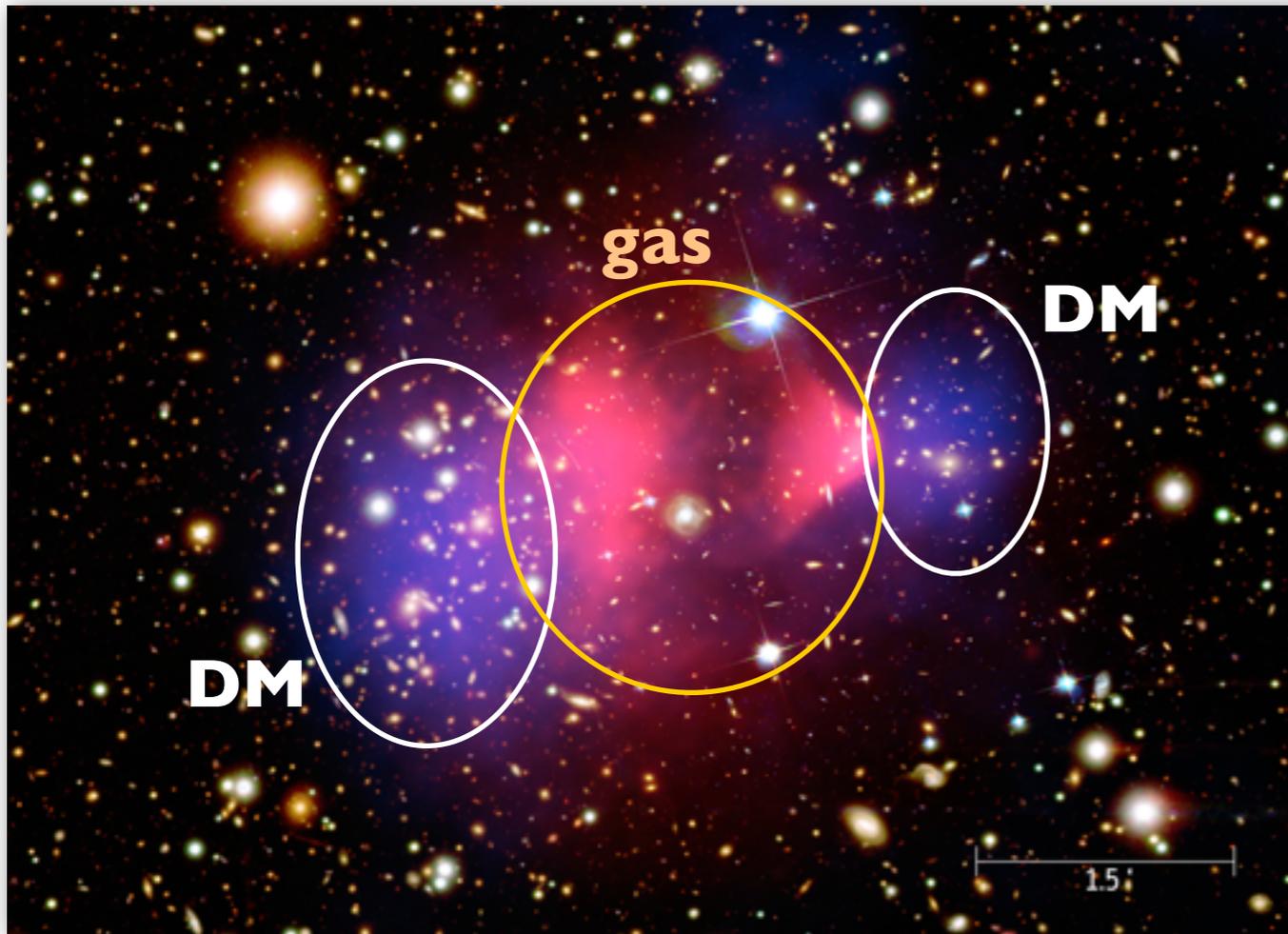
Galaxy cluster Abell 2218



The modeling of Dark Matter has become more and more articulate. From a single source (WIMP, axion, neutrino, ...) to the possibility of dark hidden worlds



# Evidence building up for self-interacting DM



$$\sigma \sim 1 \text{ cm}^2 (m_{\chi}/g) \sim 2 \times 10^{-24} \text{ cm}^2 (m_{\chi}/\text{GeV})$$

$$\text{For a WIMP: } \sigma \sim 10^{-38} \text{ cm}^2 (m_{\chi}/100 \text{ GeV})$$

Growing interest in models with rich sectors of “dark” particles, coupled to the SM ones via weakly interacting “portals”

## **6. Can the Higgs be the portal between the visible and the hidden world?**

Plausible BSM theories of this type exist. They may also

- solve the hierarchy problem in a *natural* way
- connect the mechanisms that create the matter-over-antimatter asymmetry in the Universe, with those generating Dark Matter
- explain why there are similar amounts of visible and dark matter in the Universe

***The opportunities for testing and discovering such scenarios at the LHC are being studied***

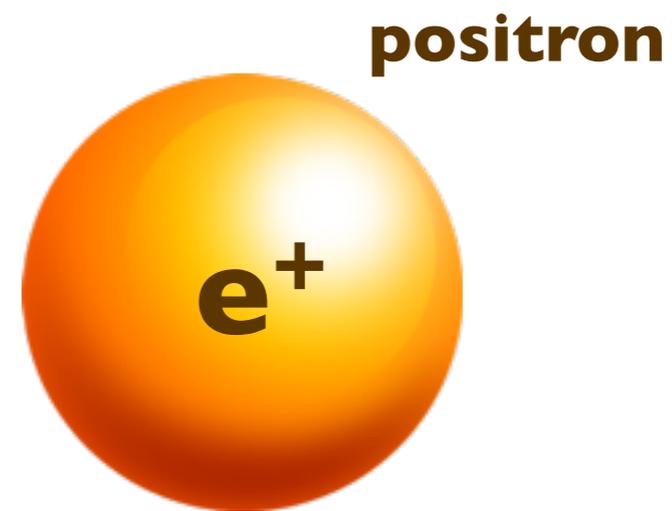
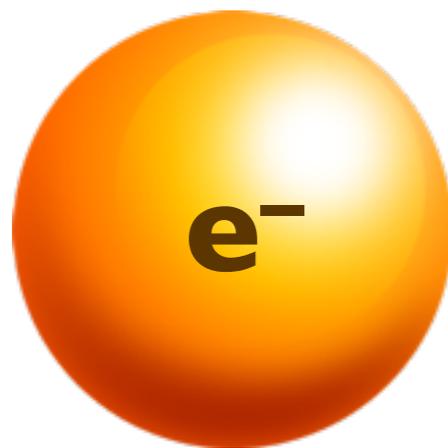
*The search for Dark Matter particles at the LHC continues, independently of these scenarios, and remains one of the key goals of future runs ...*

# Milestones of the XX century

**Special relativity from space-time symmetry**  
**+**  
**quantum mechanics**

**foremost consequence of the two:**

**the electron is not alone !**

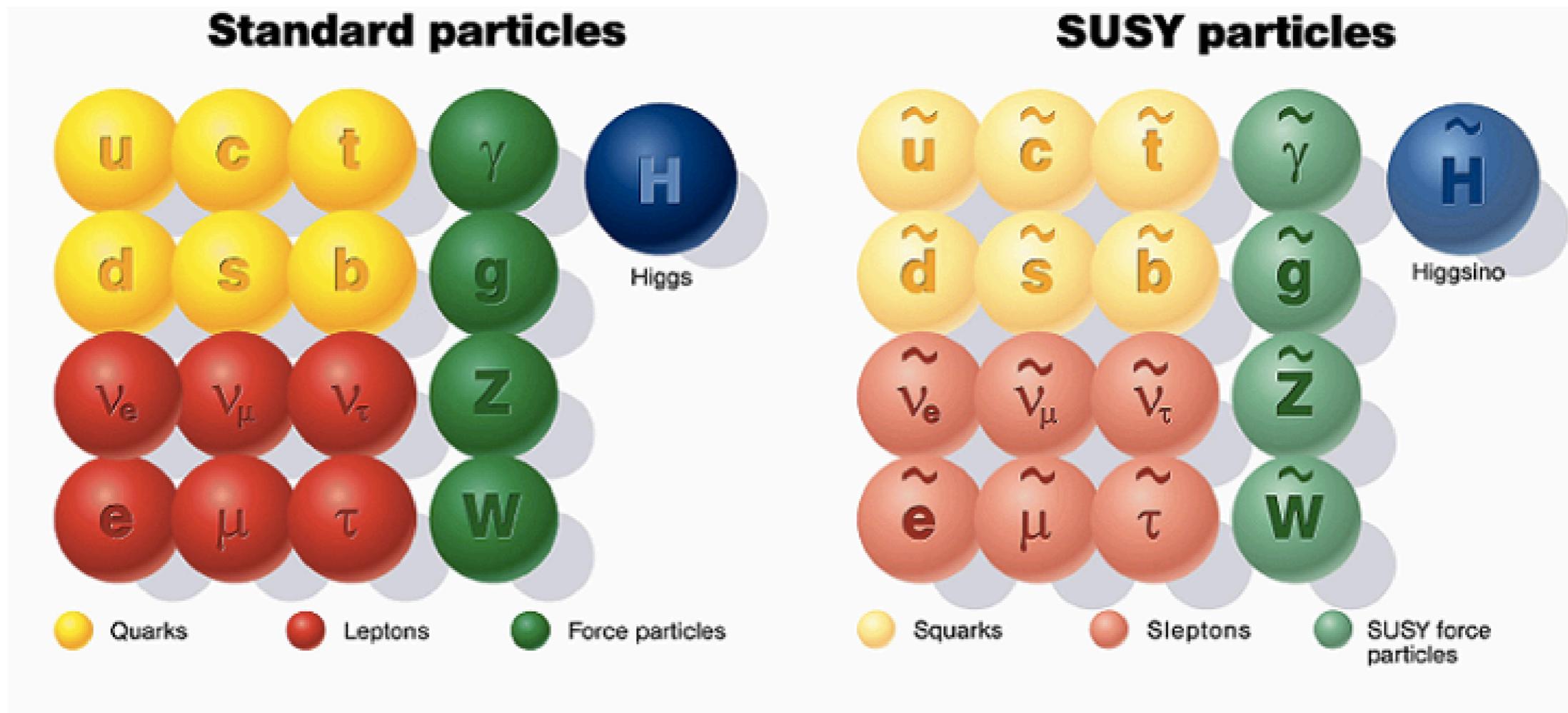


**this is generalized to all other fundamental particles: (except the photon and the Z) they all have an antiparticle !!**

# Supersymmetry: an additional possible symmetry of the space-time

*... in fact, the largest possible symmetry of the space-time as we know it ...*

**foremost consequence: each SM particle has a “supersymmetric partner”**



# Is Supersymmetry (*SUSY*) really an underlying symmetry of our world ?

## Fundamental implications:

- the discovery of a single Susy particle implies the existence of all others!!
- several SUSY particles and interactions directly address, and could solve, open issues of particle physics:
  - breaking of EW symmetry, triggering the Higgs mechanism
  - origin of DM
  - origin of matter/antimatter asymmetry
  - “hierarchy problem”
  - relation between gravity and the “other” forces (strong and electroweak)
  - unification of all forces (GUT), and a framework to understand neutrino masses

***Most other BSM theories address one or more of these issues.***

***None has the same intrinsic “simplicity” (builds on the basic concept of space-time symmetries), with the most extensive range of conceptual and practical implications***