

# Reaching for the stars

## New developments in ground-based astronomy

Markus Kissler-Patig



Director



# Today's ground-based astronomical facilities

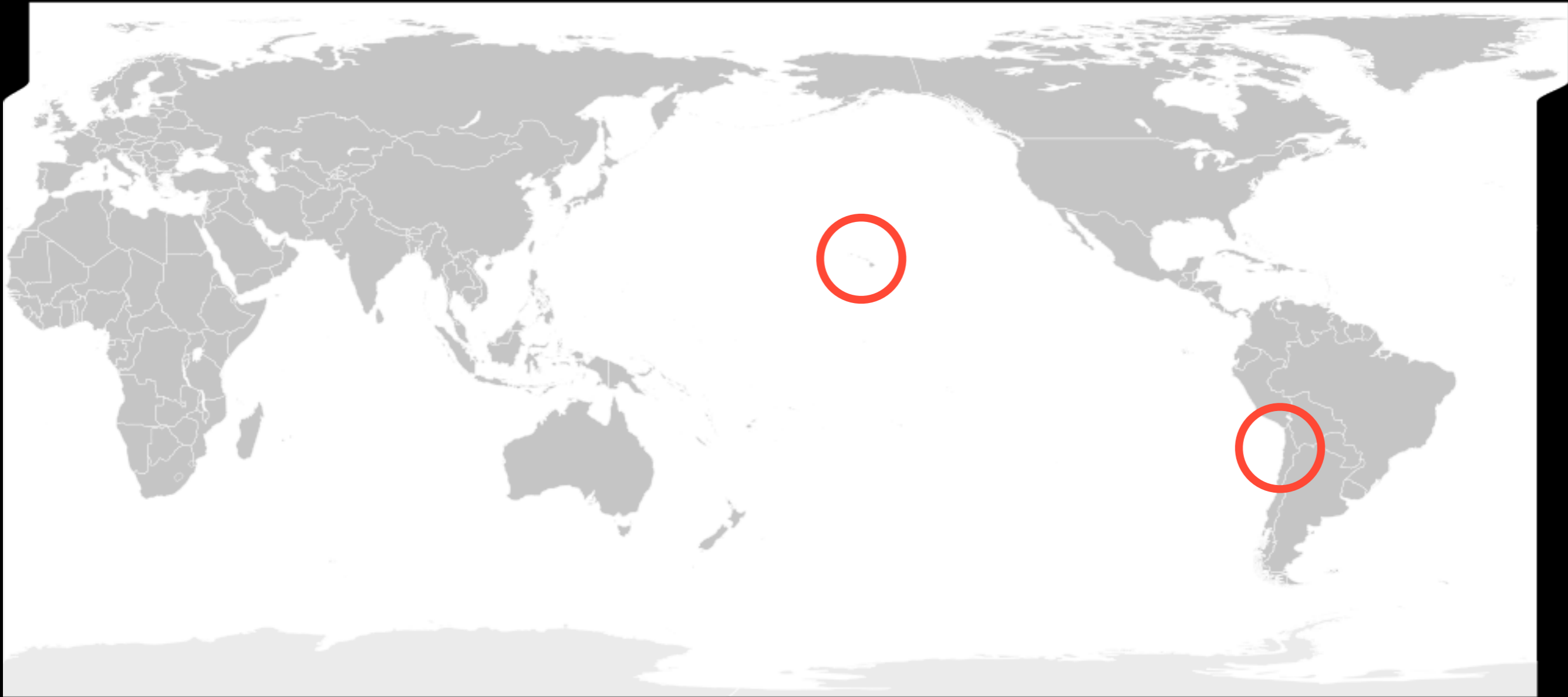
Current research in Exoplanets

Current research in Cosmology

Upcoming facilities (LSST, ELTs)

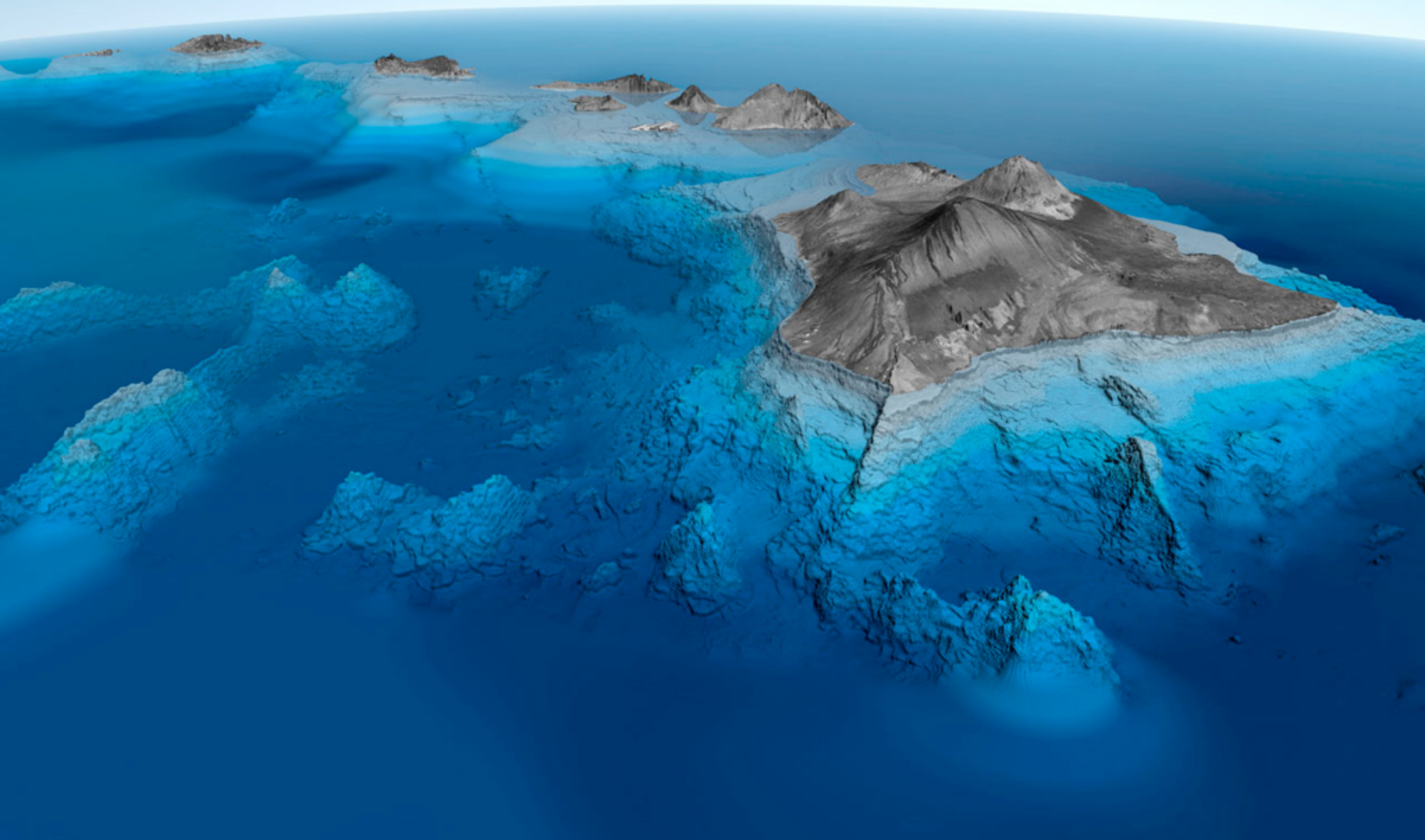
# Today's ground-based astronomical facilities

# The best observing sites on the planet:



Atmospheric conditions (turbulence, water vapor), accessibility, attractive to international staff, ...

# Hawai'i





Hilo, rainiest city in the United States





Maunakea

TMT

Keck

IRTF

Subaru

CFHT

SMA



Gemini

JCMT

UH 88'

CSO

UKIRT

VLBA





# Why so many telescopes?

Astronomers exploit all the light they can get...

Covering the accessible **wavelength range**:

**Radio**: neutral gas (Hydrogen) [VLBA]

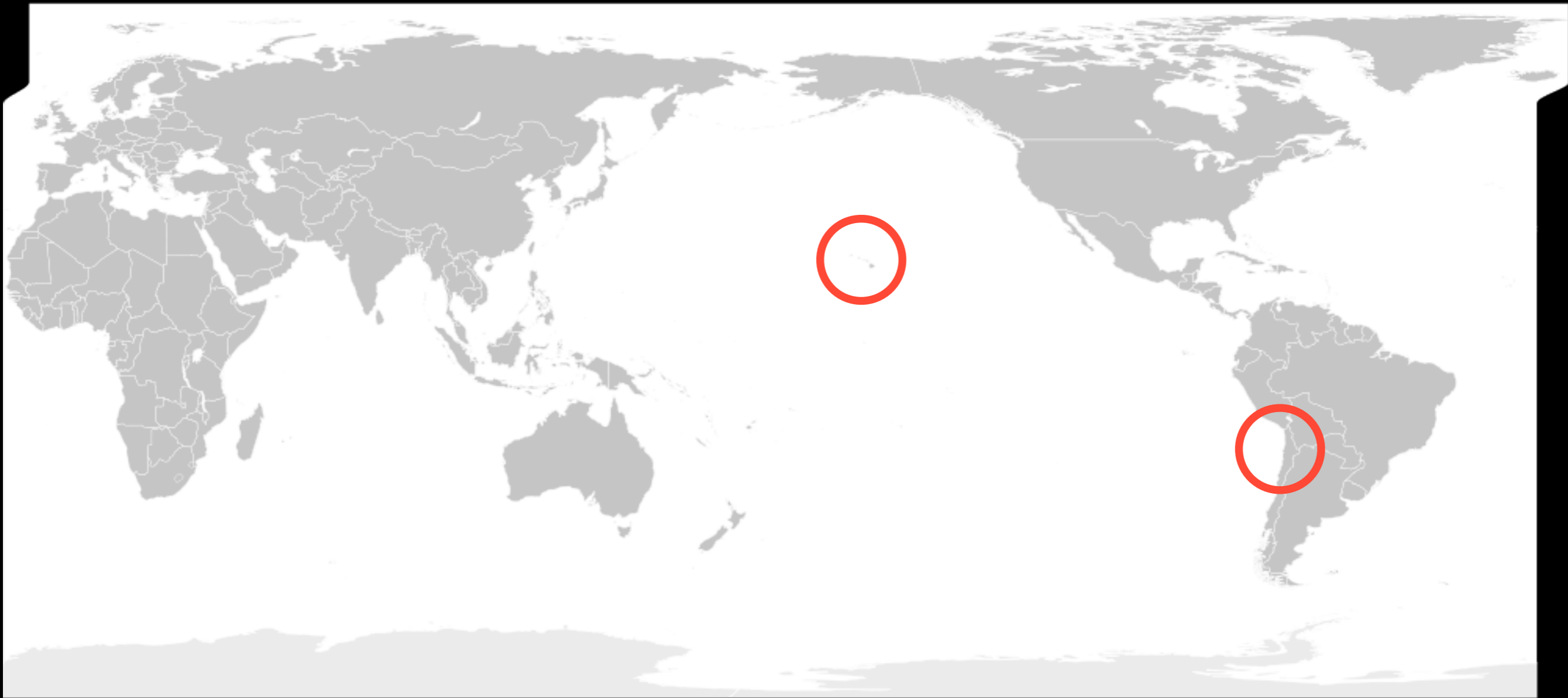
**Sub-mm**: cold (10-100K), molecular gas [CSO, JCMT, SMA]

**Near-infrared**: dust obscured light, star formation, highly redshifted objects [UKRIT, IRTF]

**Optical/UV**: stellar light, galaxies [Keck, Subaru, Gemini, CFHT, UH 88']

Within a wavelength range, **specializing the telescopes** for narrow/wide field imaging and/or spectroscopy

# The best observing sites on the planet:



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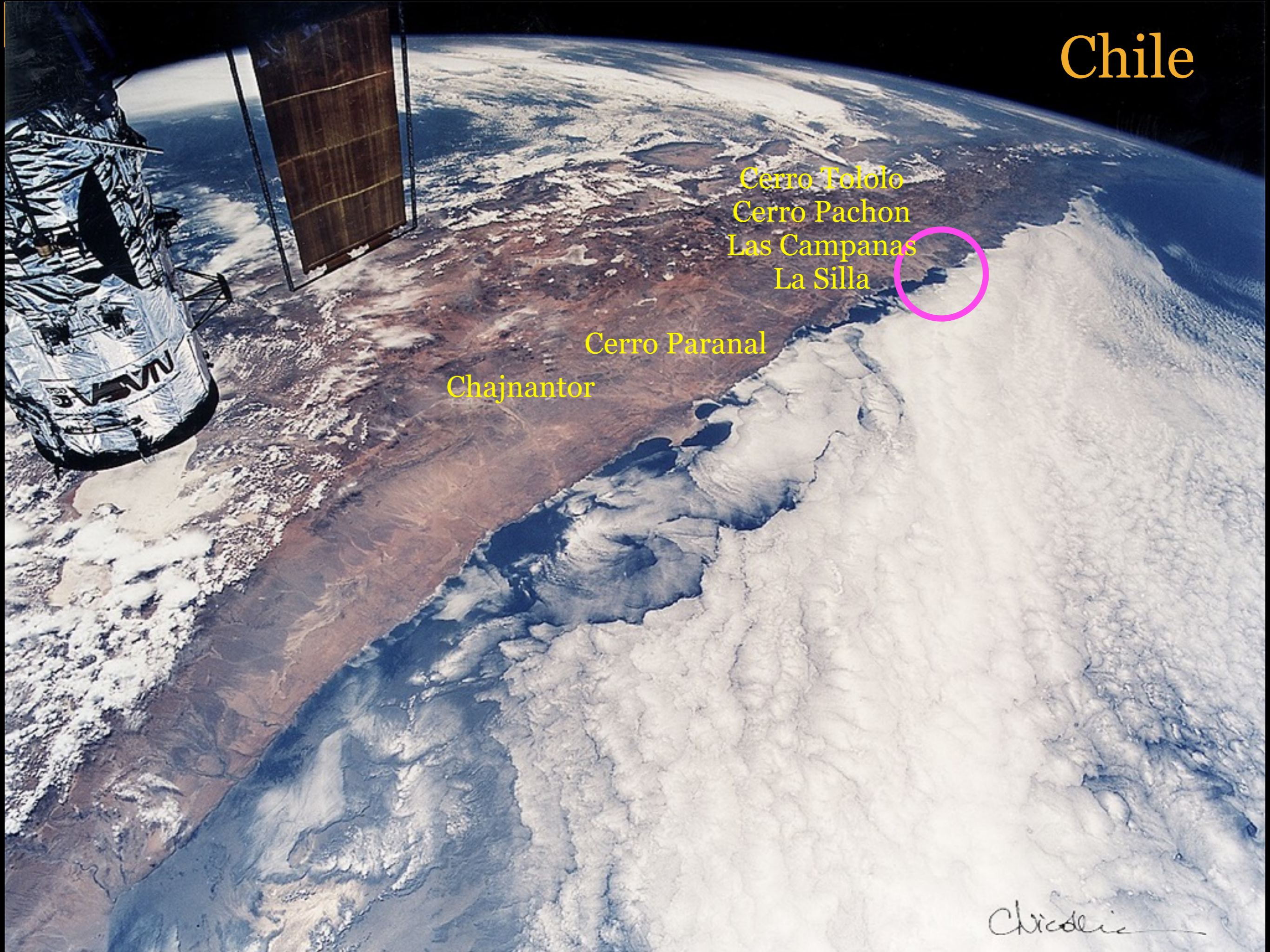
# Chile

Cerro Tololo  
Cerro Pachon  
Las Campanas  
La Silla

Cerro Paranal

Chajnantor

Chiodic



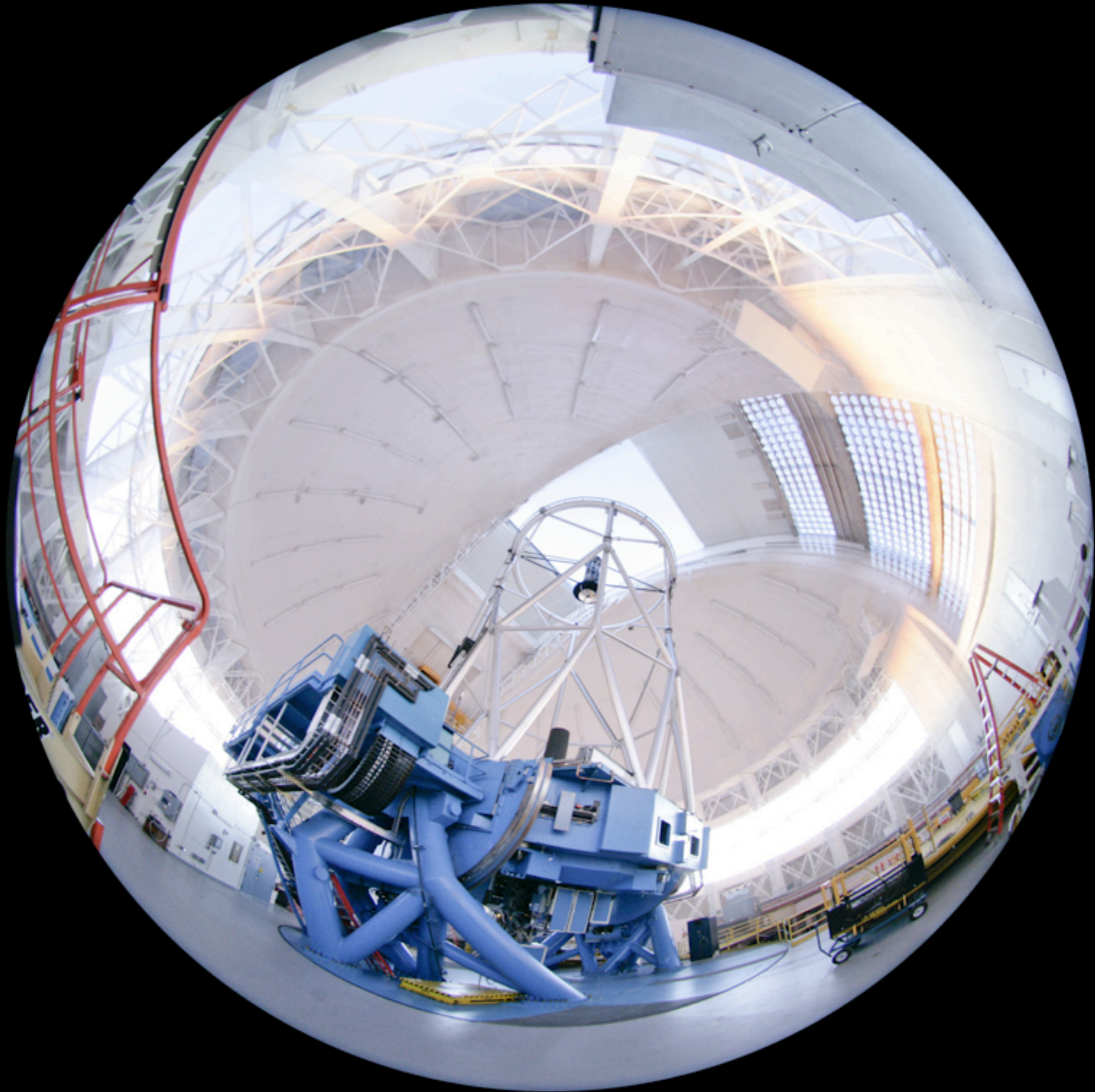




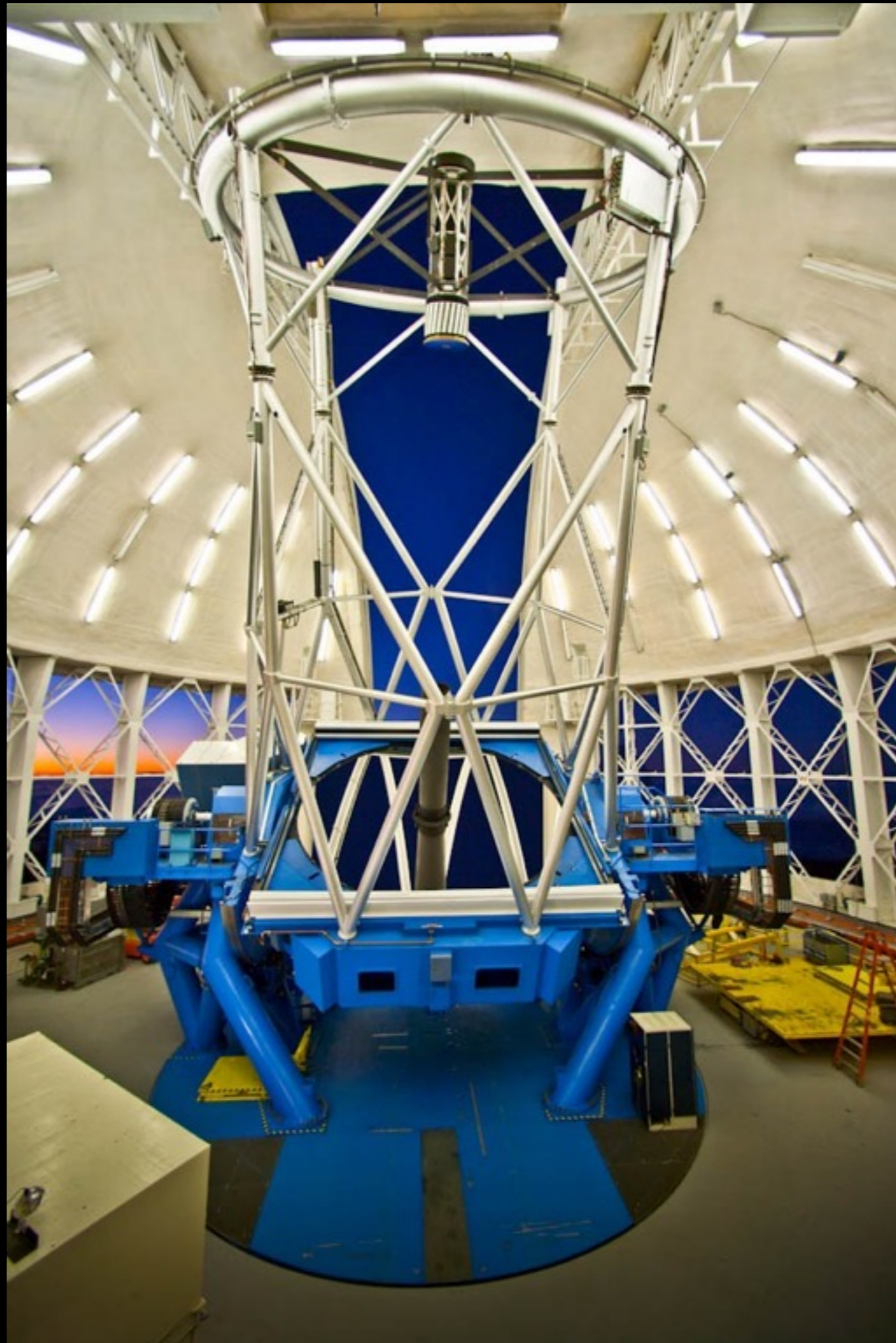
# Telescopes in operations...





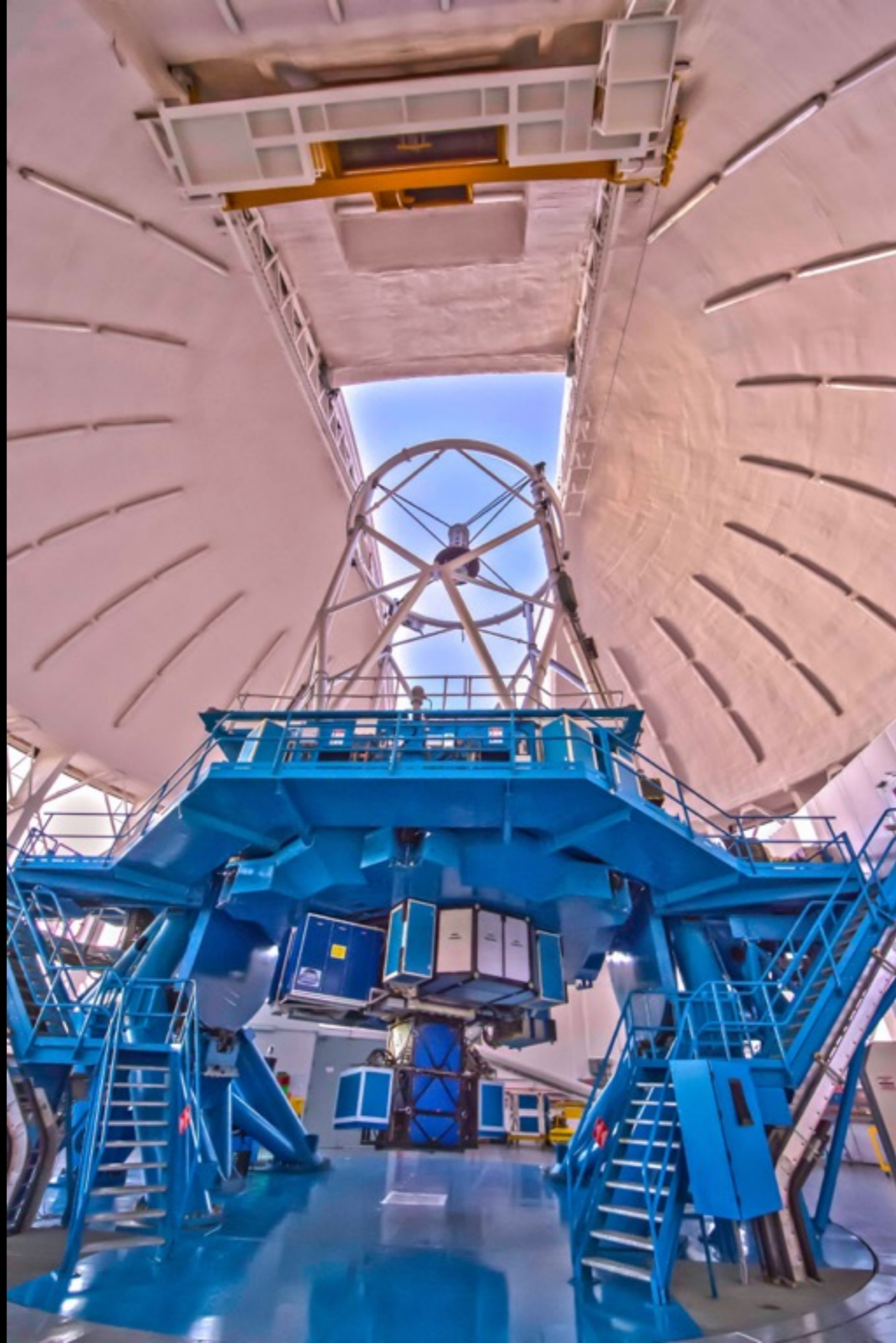


# Telescopes basics



Secondary mirror  
(1.0 m  $\varnothing$ )

Primary mirror  
(8.1 m  $\varnothing$ )





Typical 8-10m telescope cost ~\$100M

Typical instruments: 3x3x3m, 5 tons, ~\$20M

$\lambda$  range (optical, near/mid-infrared)

imaging

spectroscopy

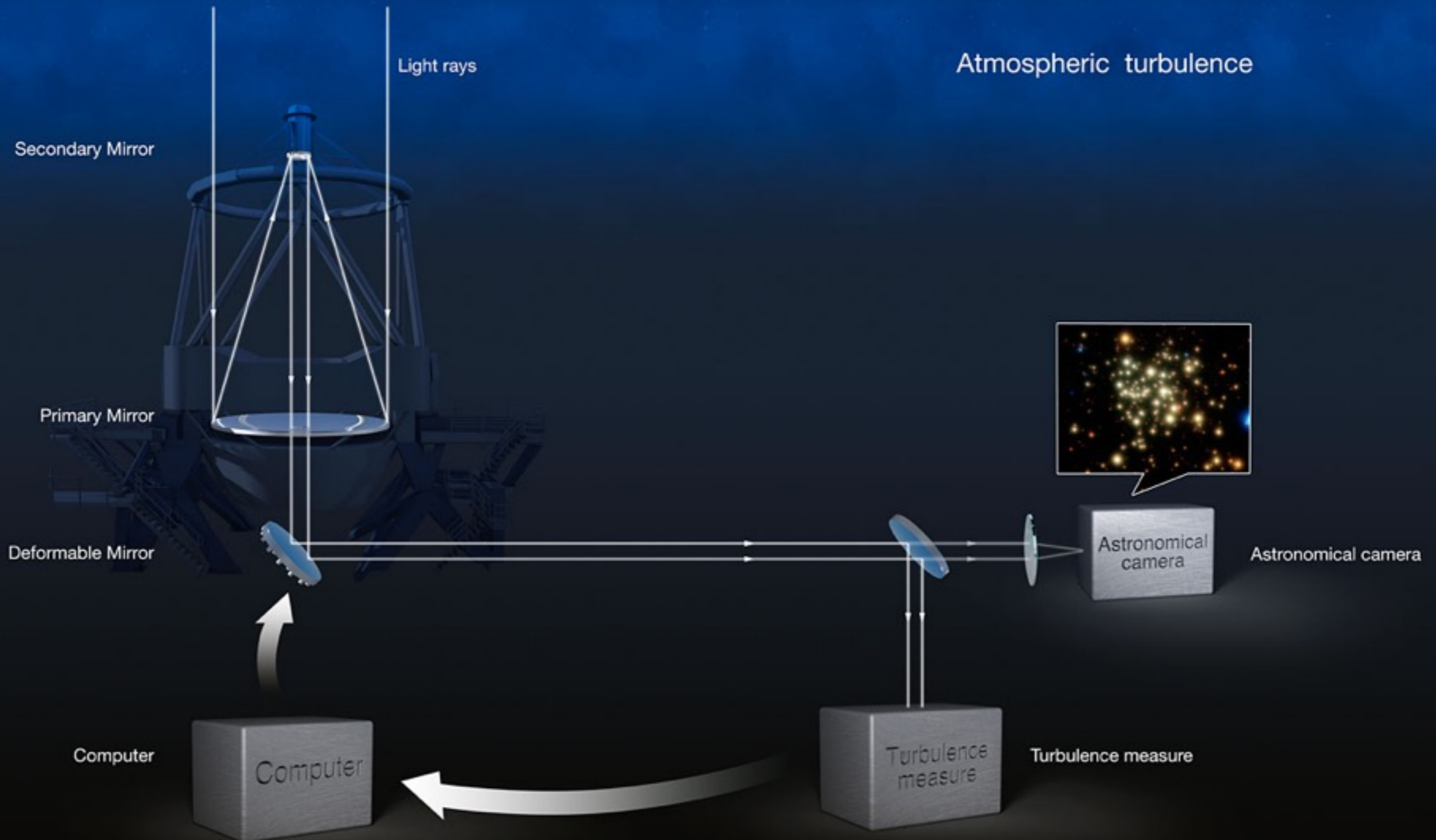
diffraction  
limited

wide-field

high-spectral  
resolution

multi-object /  
integral field

# Lasers and adaptive optics

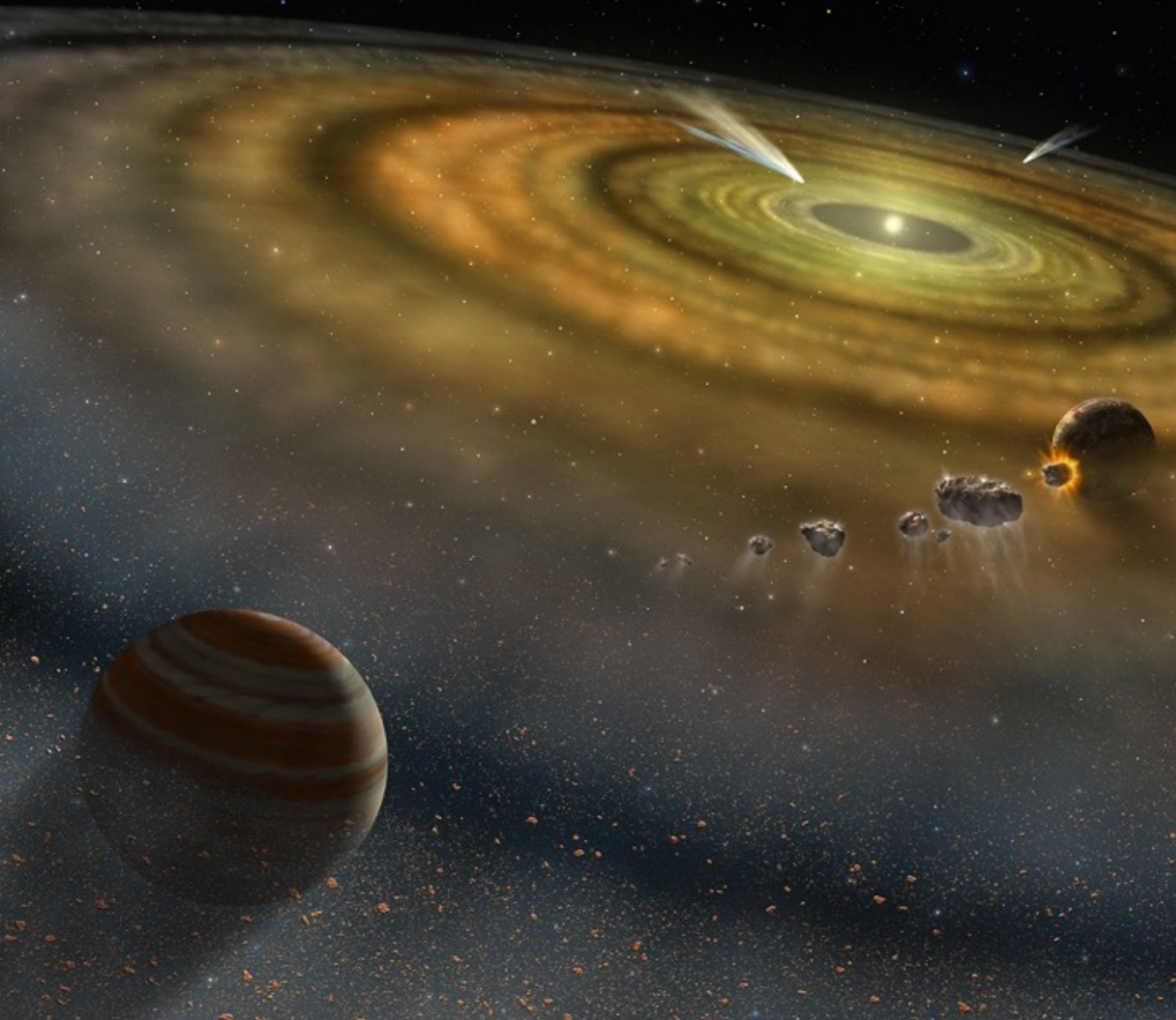






# Current Research in Exoplanets

# Are we alone?



# A Brief history of exoplanet research

**1992:** First exoplanets detected around Neutron Star (**still weird**)

**1995:** First exoplanet detected around a Sun-like star  
(Mayor & Queloz, Université de Genève)

**2003:** Exoplanet #100 detected

**2005:** First image (direct detection) of an exoplanet

**2005:** First detection of water in an exoplanet “atmosphere”

**2007:** First exoplanet discovered in a “habitable zone”

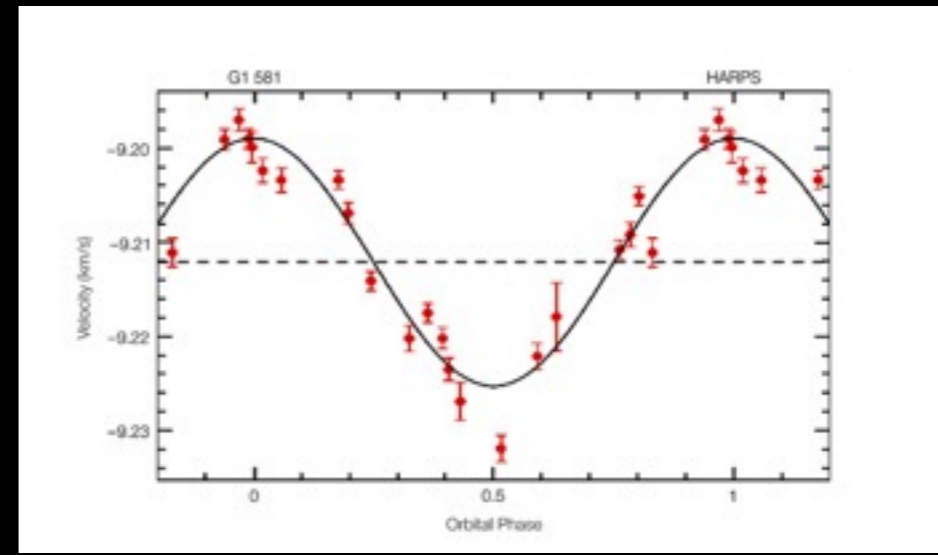
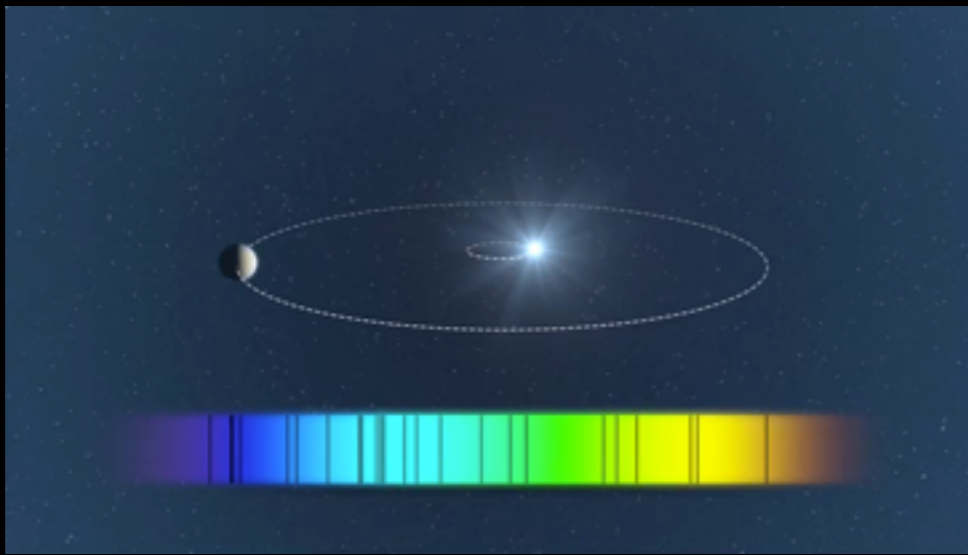
**2009:** First telluric exoplanet in a “habitable zone”

**2010:** Kepler releases its first few hundred planet candidates

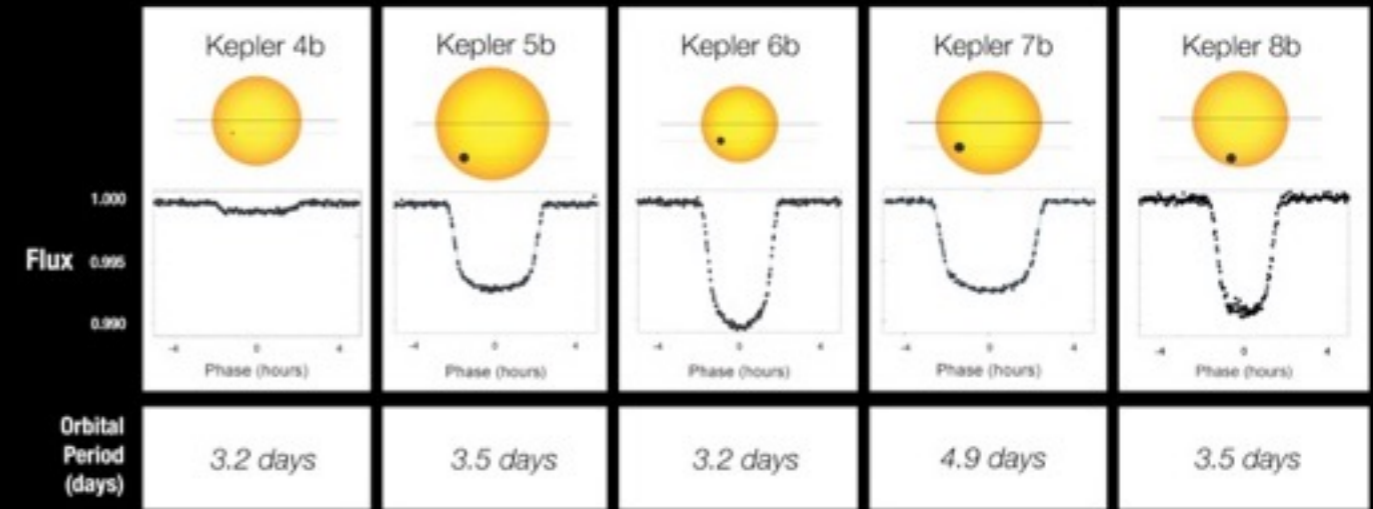
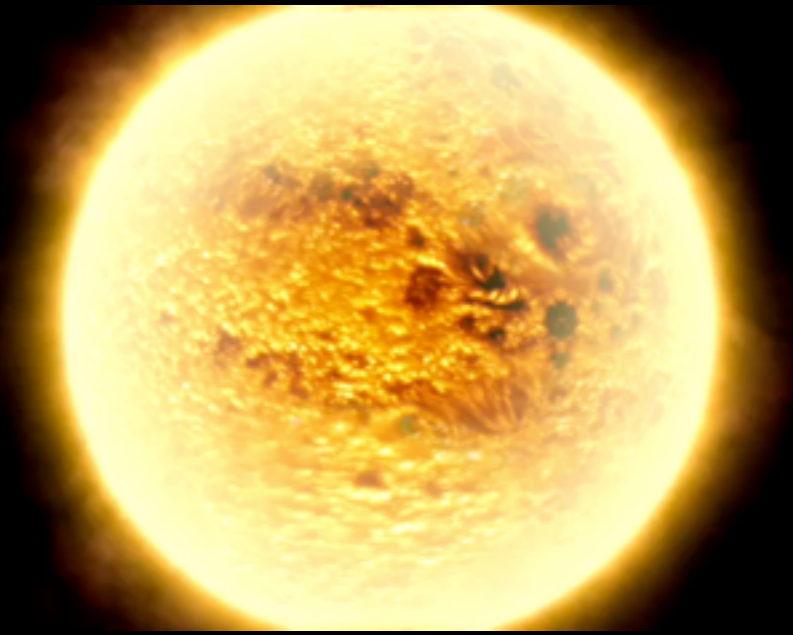


The race for “firsts” continues - the next frontier is to go beyond detection and into characterization of exoplanets (atmospheres)

Radial Velocity



Planet Transit

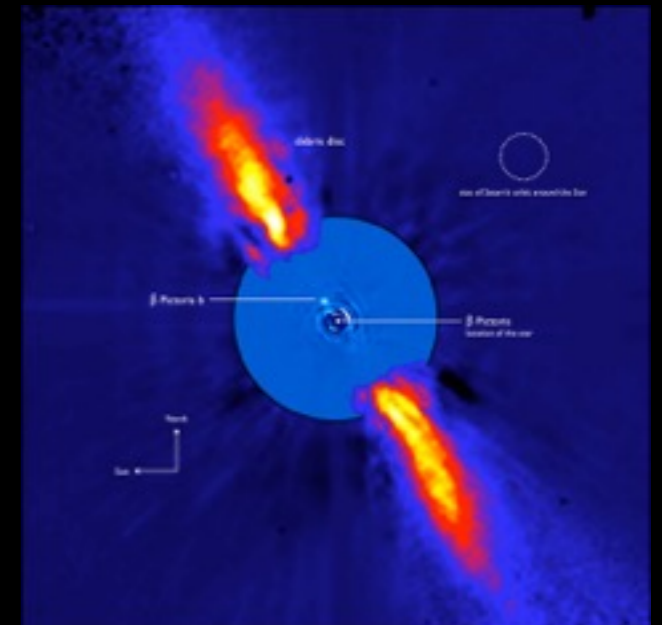
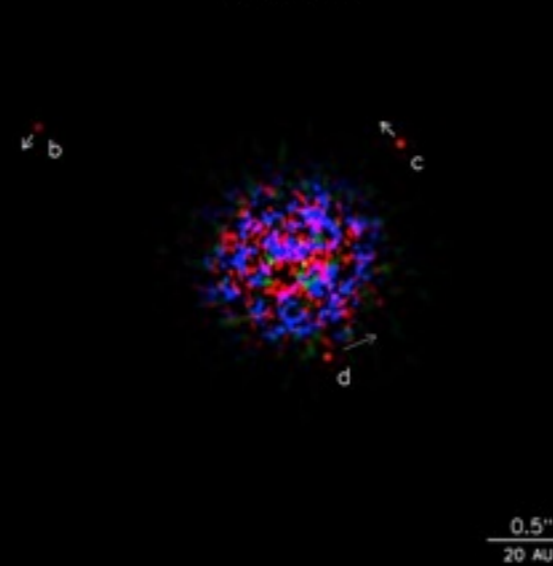


Direct Imaging

2MASSWJ1207334-393254



HR 8799



Detected as of yesterday (July 1, 2015):  
[[www.exoplanet.eu](http://www.exoplanet.eu)]

Radial velocity or astrometry: 604

Transiting exoplanets: 1211

Microlensing: 35

Imaging: 59

Timing: 23

Total: 1932 exoplanets in 1222 systems

# Habitable planets as of yesterday (July 1, 2015):

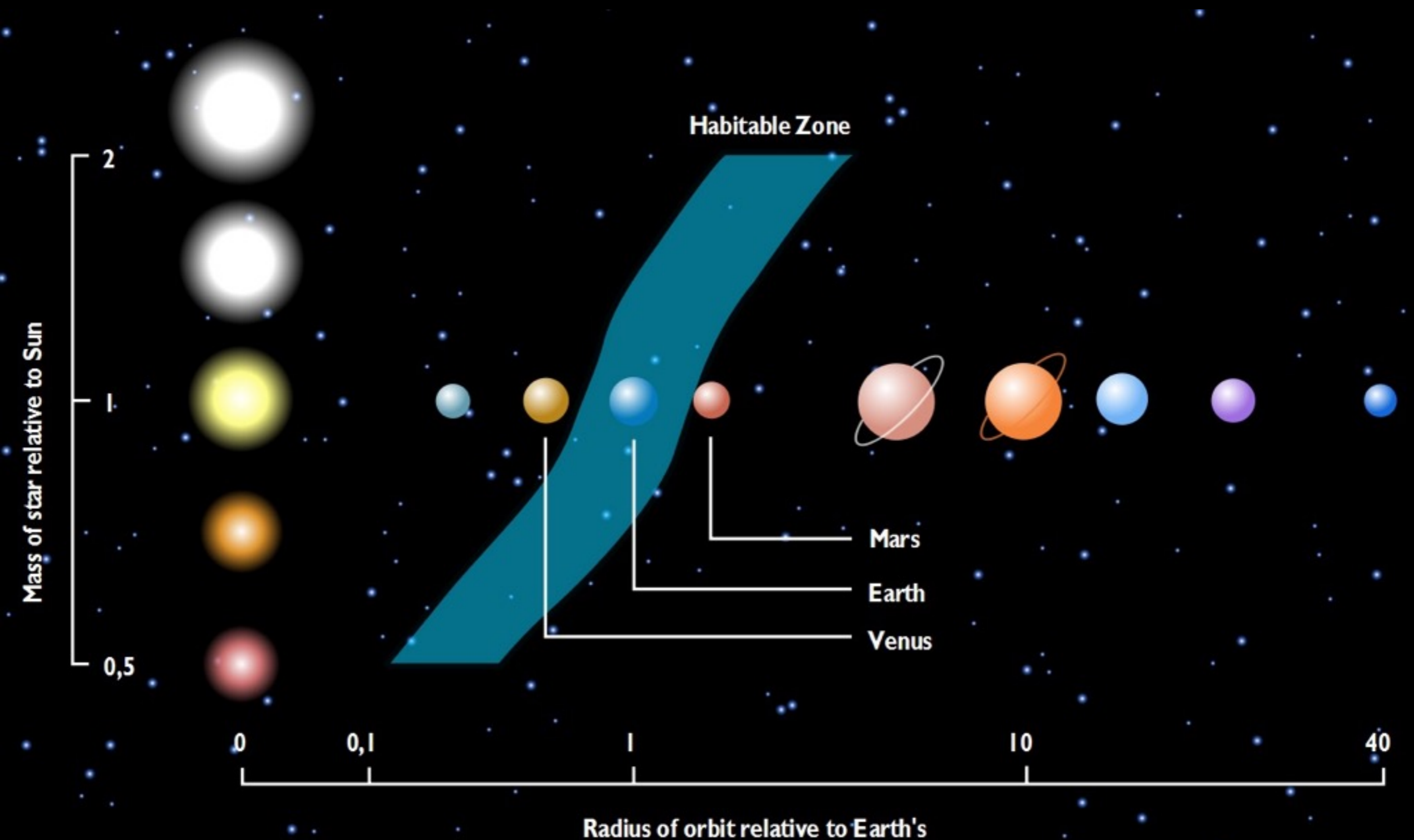
[<http://phl.upr.edu/projects/habitable-exoplanets-catalog/>]

## Potentially Habitable Exoplanets

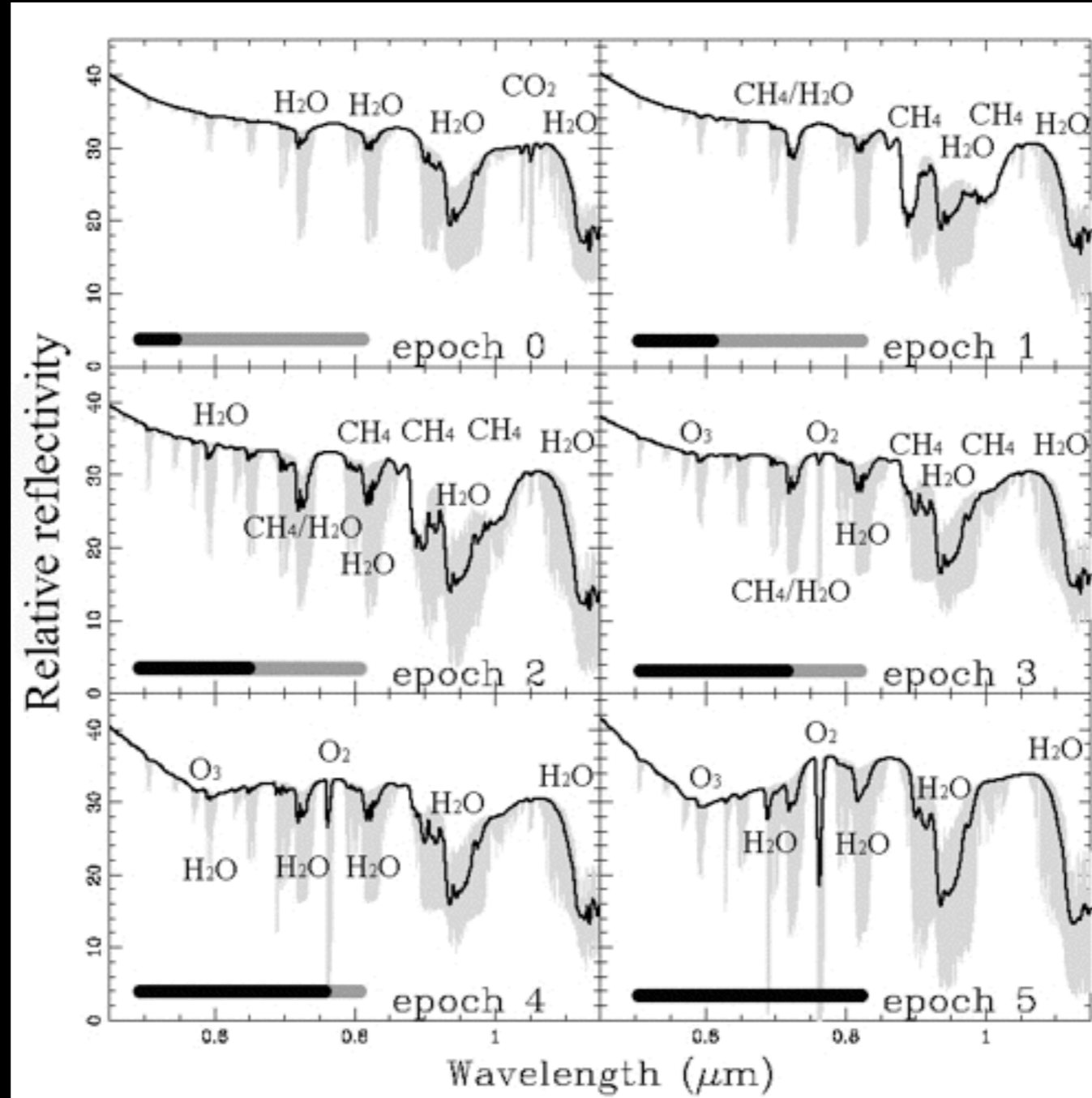
Ranked by Distance from Earth (light years)



# The Habitable Zone (liquid surface water):

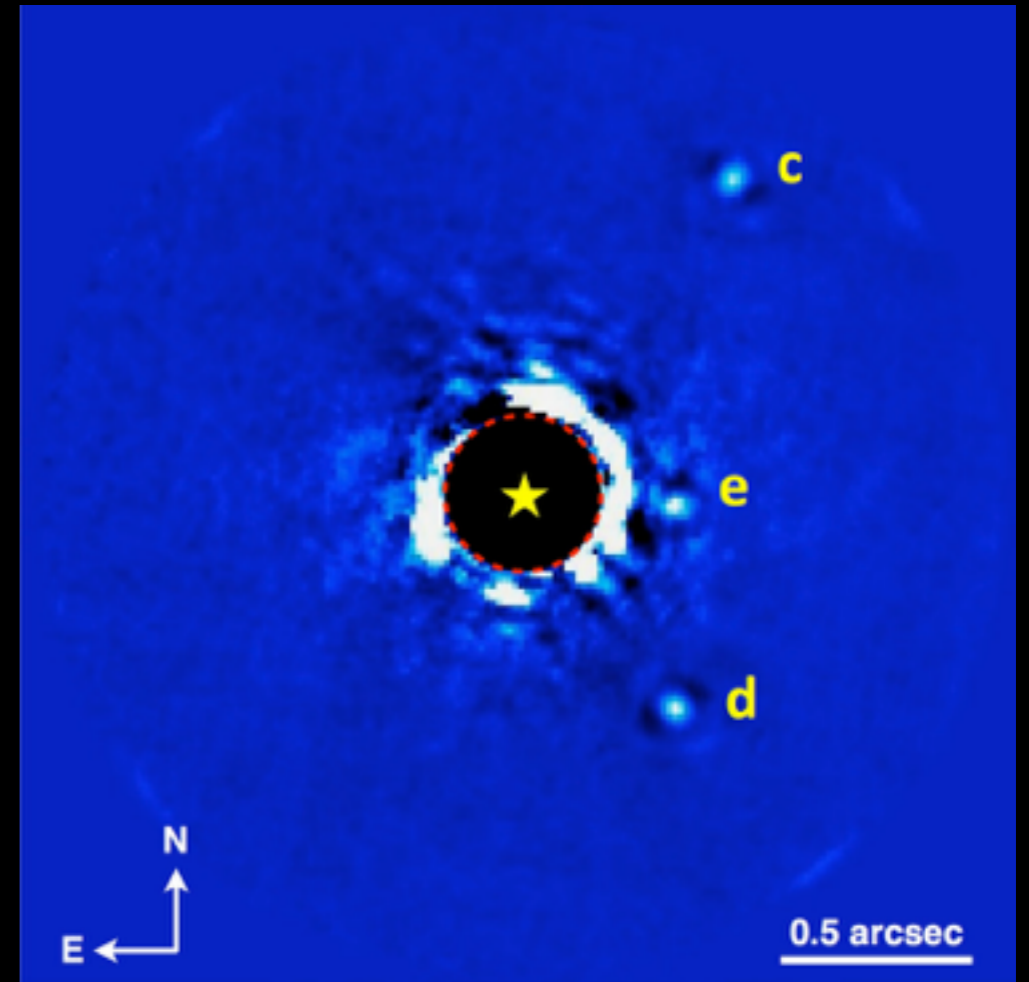


# Bio-signatures: Life and Earth's atmosphere co-evolved





# Current state-of-the-art instruments



The Gemini Planet Imager achieves contrasts of  $1:10^{6-7}$  at  $0.1''-0.2''$  (1-2 AU at 10pc distance) of the parent star

**Extremely large telescopes** will increase the contrast by three to four orders of magnitude

# Current Research in Cosmology



[48 seconds]

# Properties of the Universe

**are determined by five numbers:**

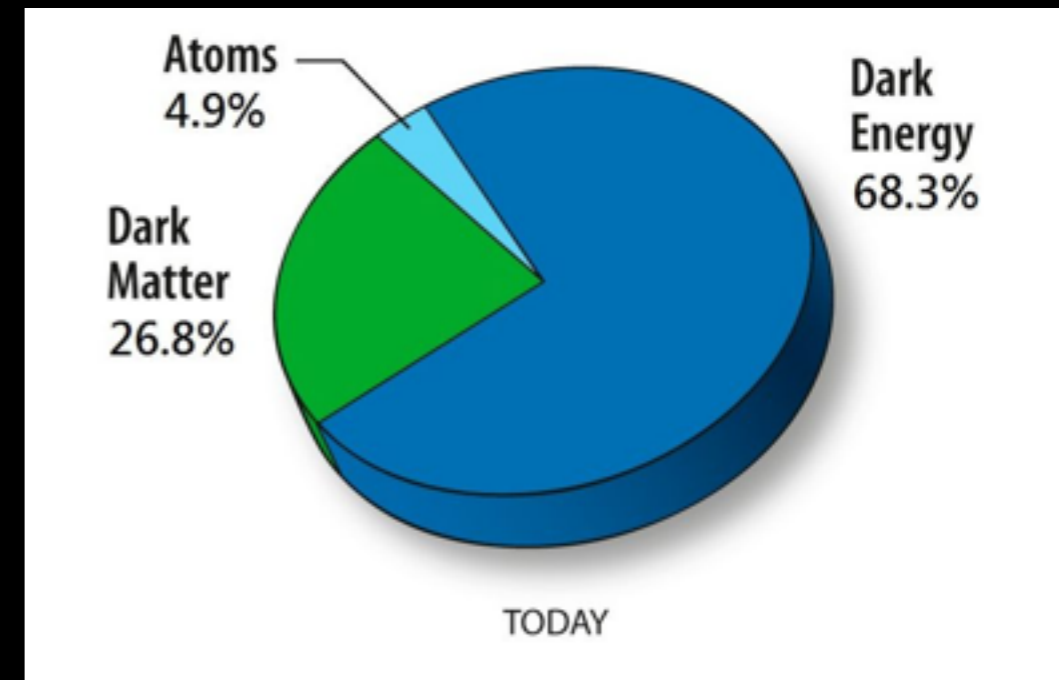
Density of matter

Density of atoms

Age of the Universe

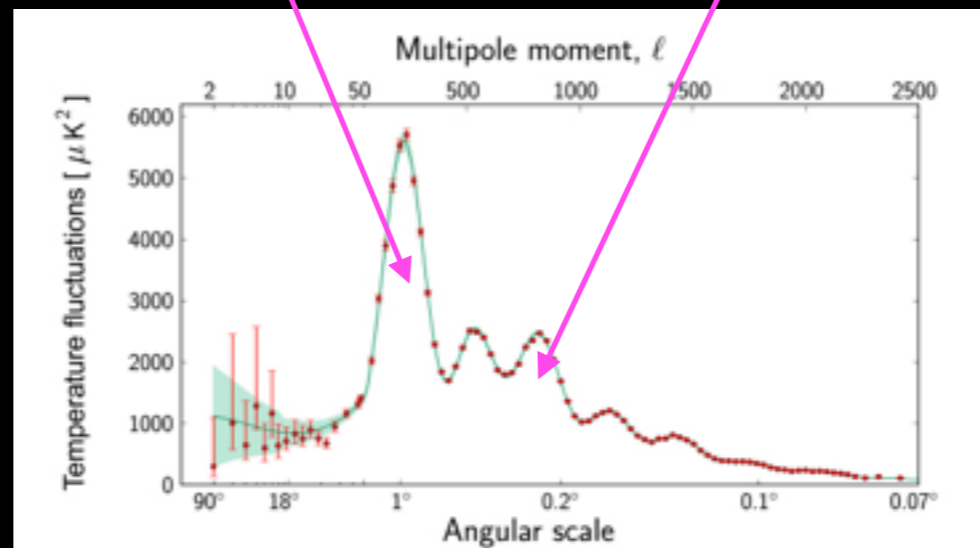
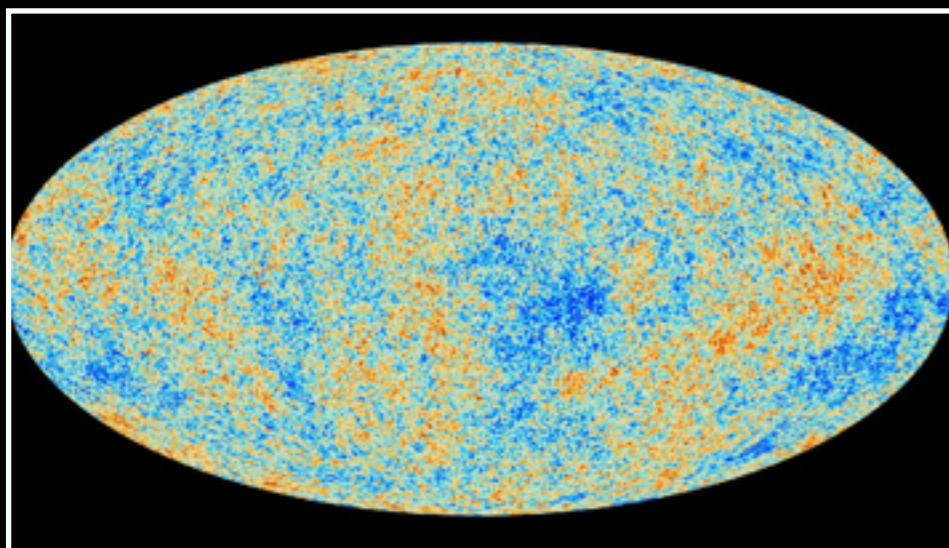
Amplitude of the initial fluctuations

Their scale dependence

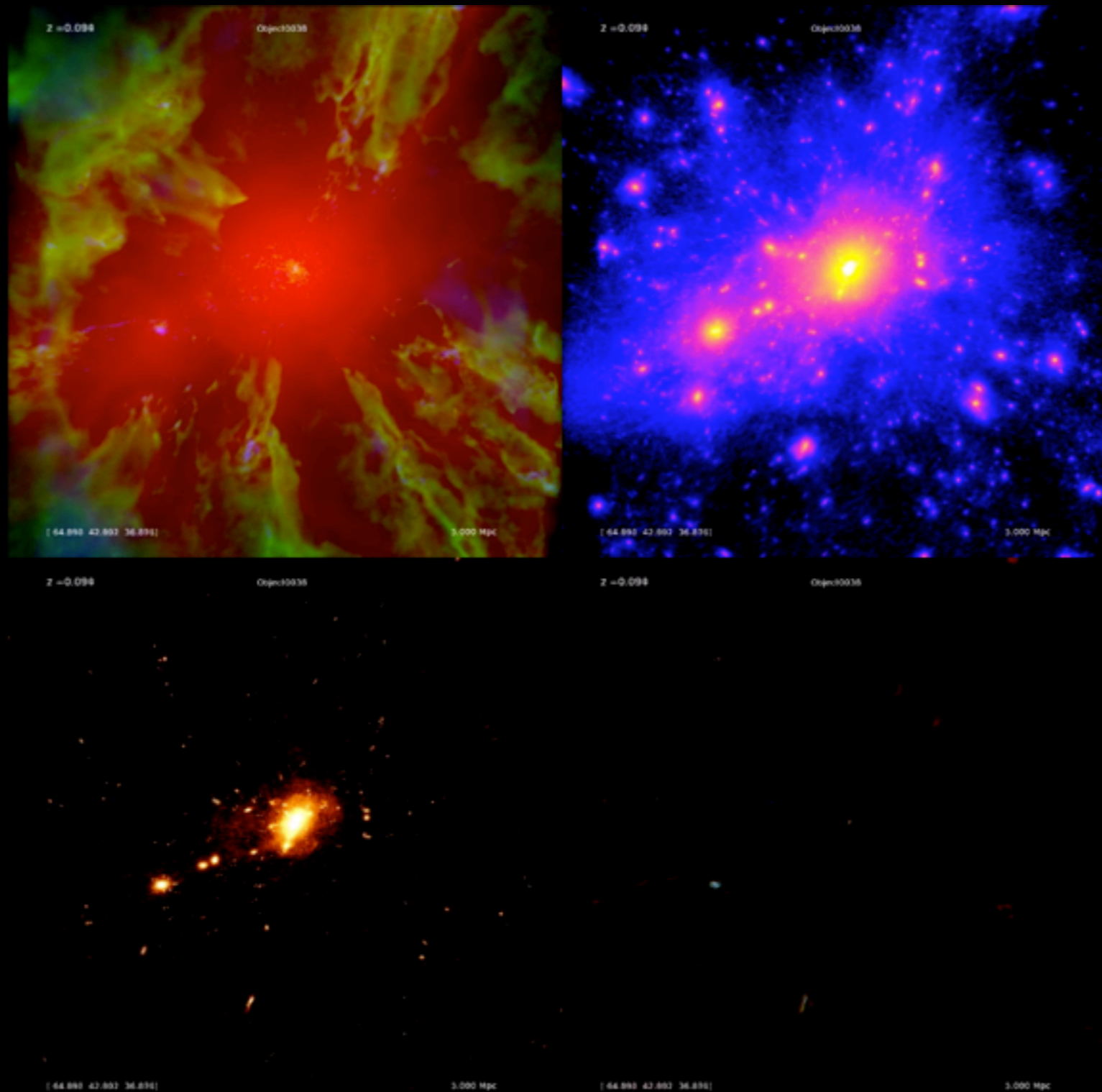


Baryonic matter density

Dark matter density



# Simulations of the Universe



[EAGLE simulations - ICC Durham - 38 seconds]

# Dark Matter

## Evidence:

Velocity dispersion of galaxies in clusters

Galaxy rotation curves

Gravitational lensing

Big Bang Nucleosynthesis

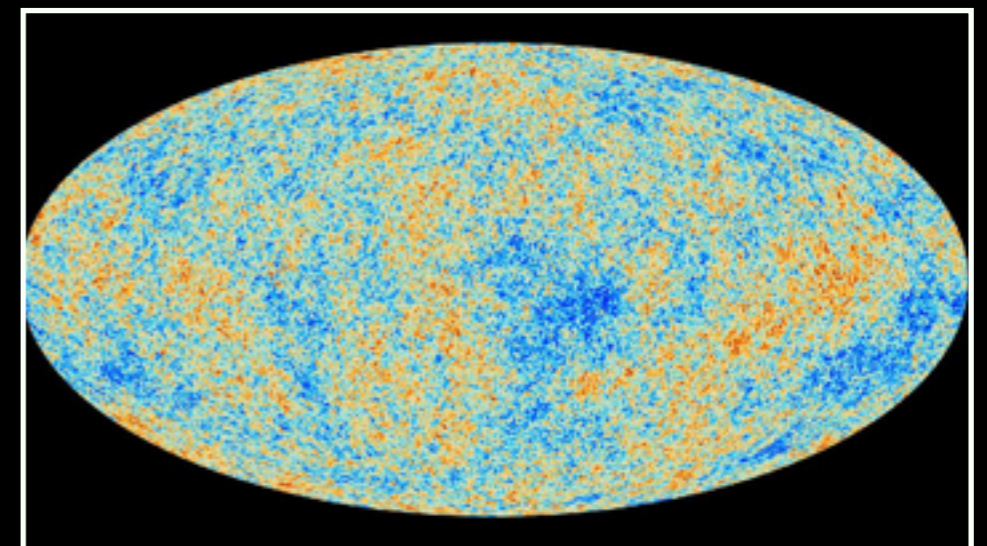
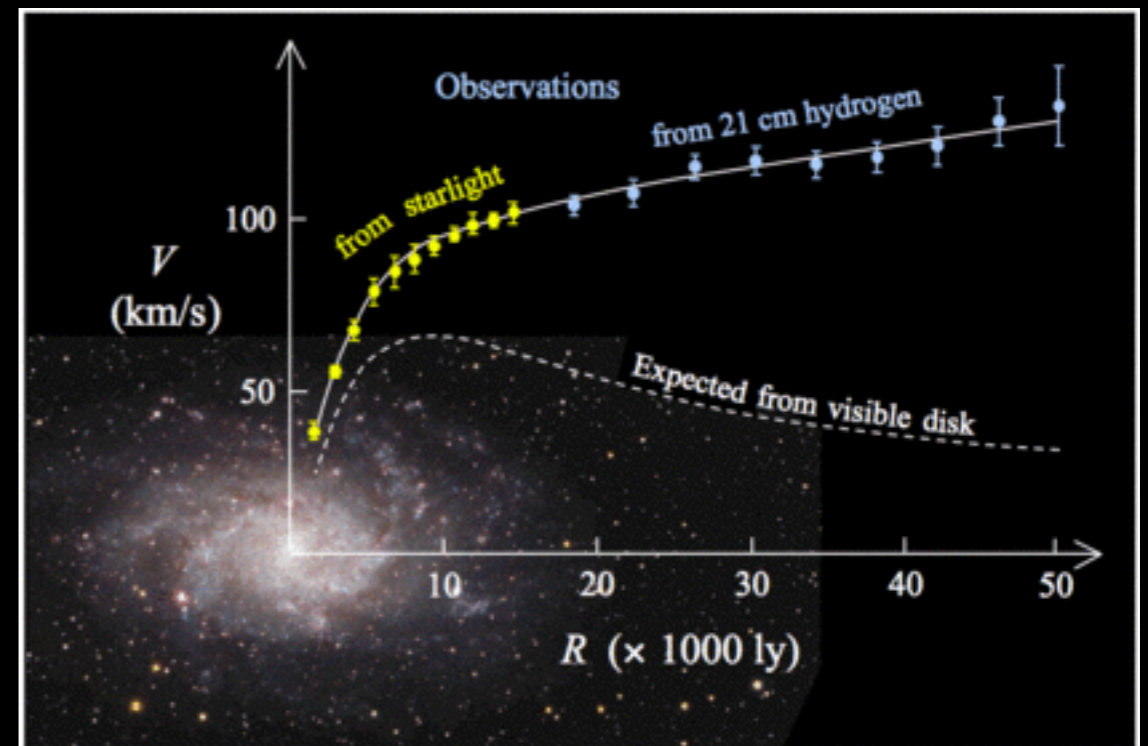
Microlensing

## Candidates:

Cold Dark Matter (e.g. WIMPs)

Warm Dark Matter (e.g. IMPs)

Hot Dark Matter (e.g. WILPs, Neutrinos)





# Dark Energy

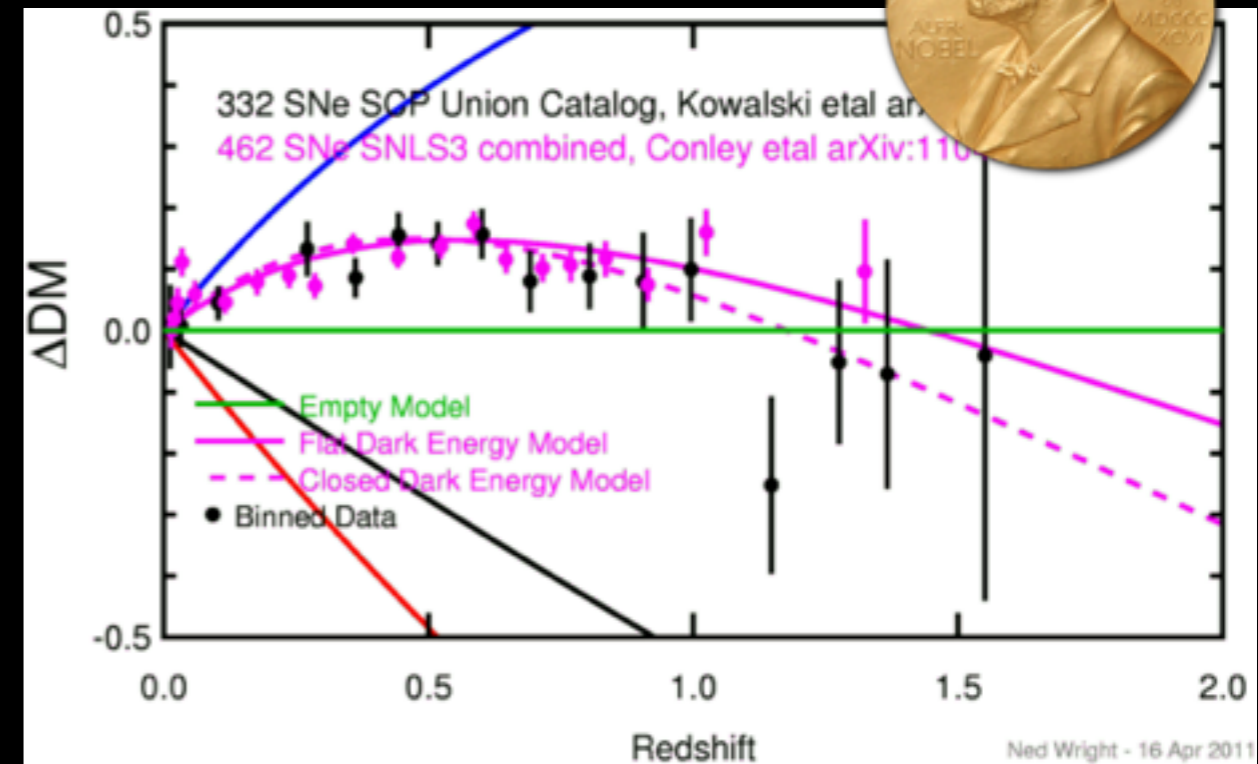
## Evidence:

Supernovae measurement  
(expansion rate)

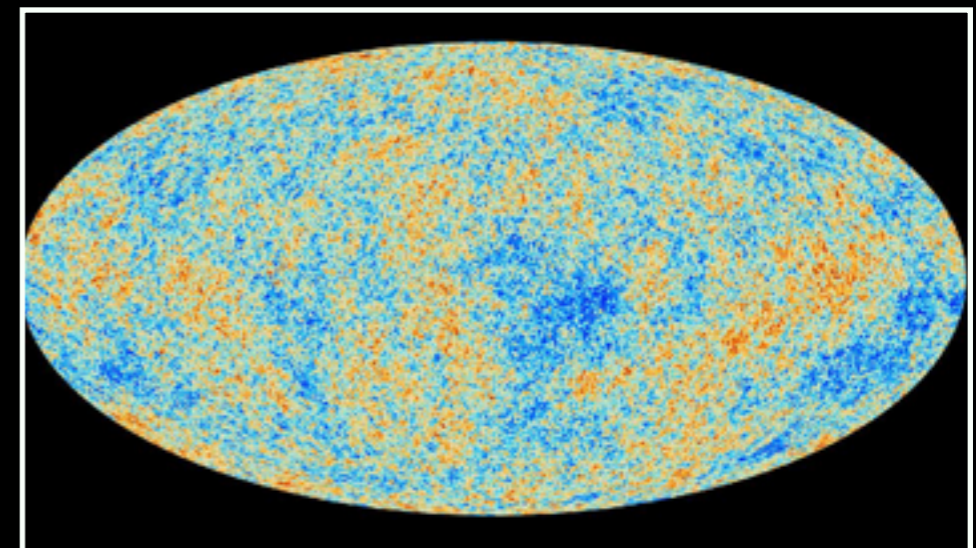
Cosmic Microwave Background  
Large Scale Structure  
(flat universe)

## Candidates:

Cosmological *constant* (vacuum energy)  
Scalar field [f(space,time)] (quintessence)



[Nobel prize in physics 2011]

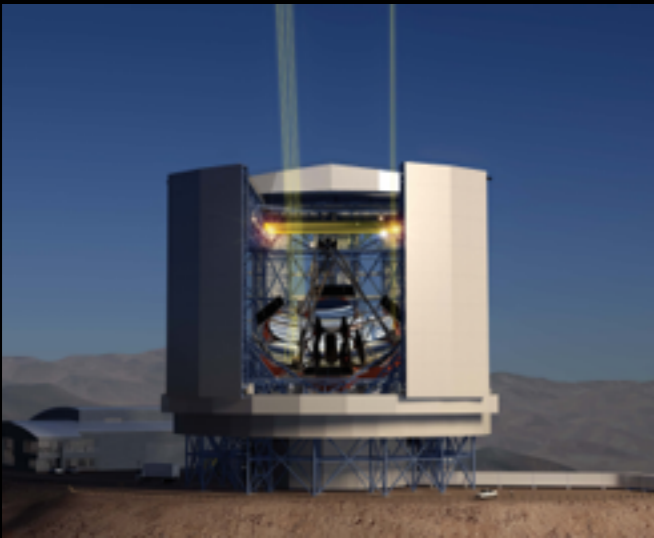




# Upcoming Facilities

# The Extremely Large Telescopes

USA, Korea, Australia, Brazil



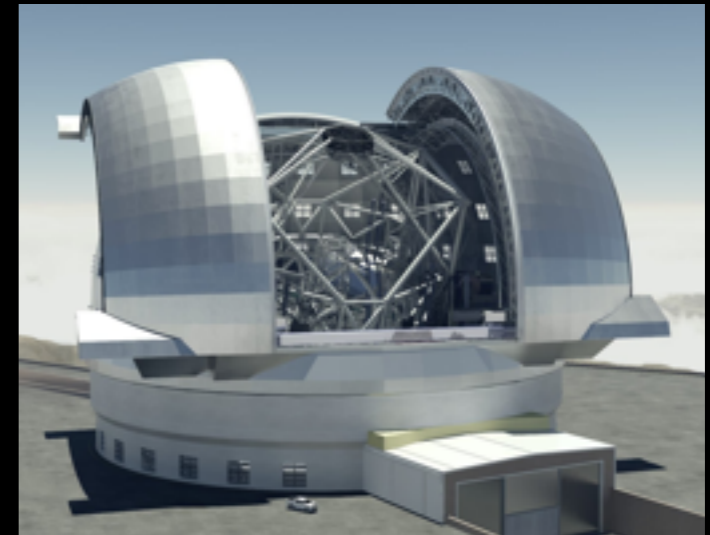
Giant Magellan Telescope  
(GMT) 25m Ø, Chile

USA, Canada, China, India, Japan



Thirty Meter Telescope  
(TMT) 30m Ø, Hawaii

ESO (15 member states)



European Extremely Large Telescope  
(E-ELT) 40m Ø, Chile

**Construction costs: \$1-1.5B**

All three started civil engineering work on site  
First light is expected between 2022-2025

## ELTs key science goals:

### Characterization of Exoplanets

(detection of Earth-like planets, characterization of their atmospheres, ...)

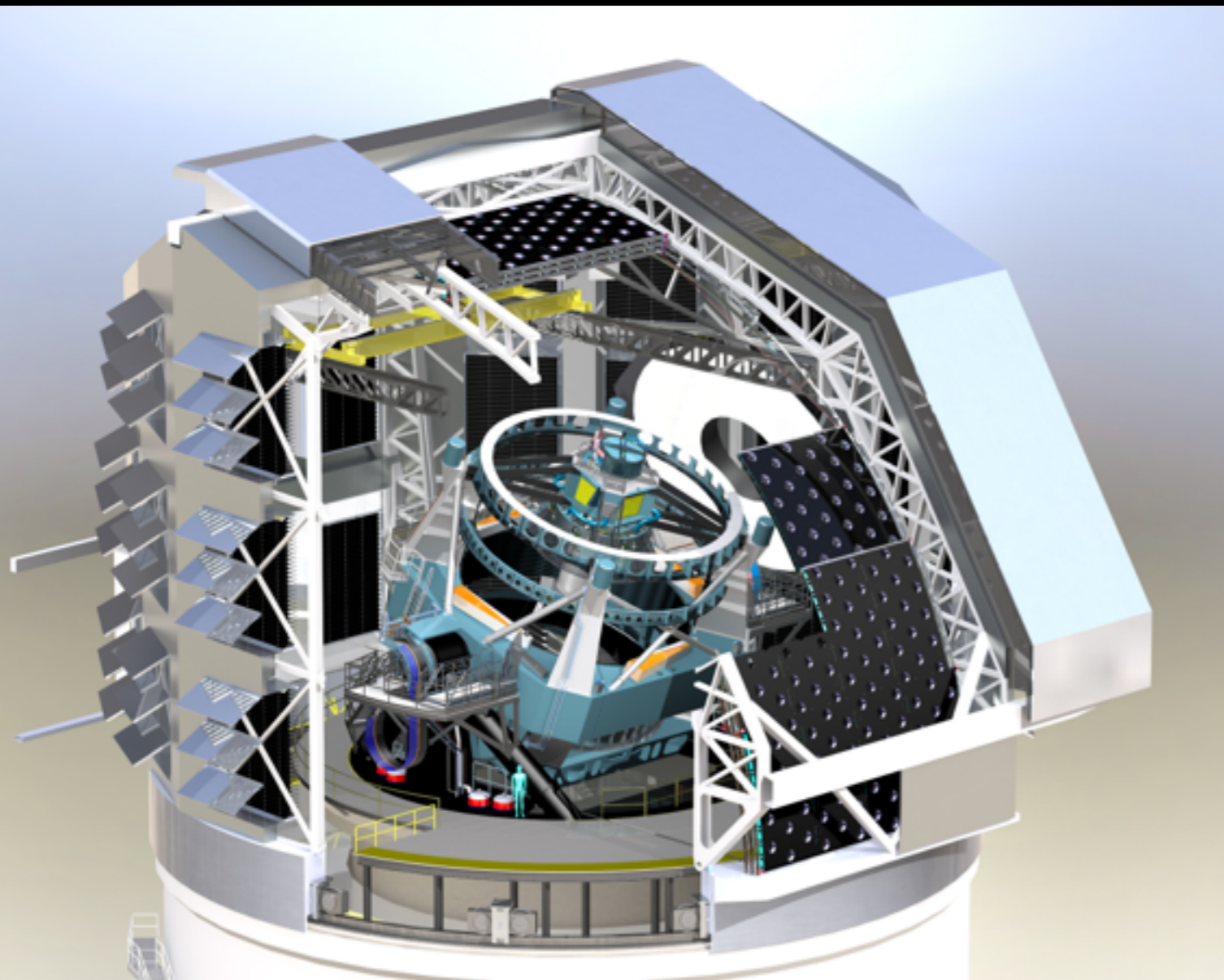
### Physics of high redshift galaxies

( $\gamma$ -ray bursts, supernovae, near-Earth objects, ...)

### Cosmology and Fundamental physics

(variations of fundamental constants, high-redshift supernovae, acceleration of the expansion of the universe, ...)

# The Large Synoptic Survey Telescope



Construction costs: \$400M

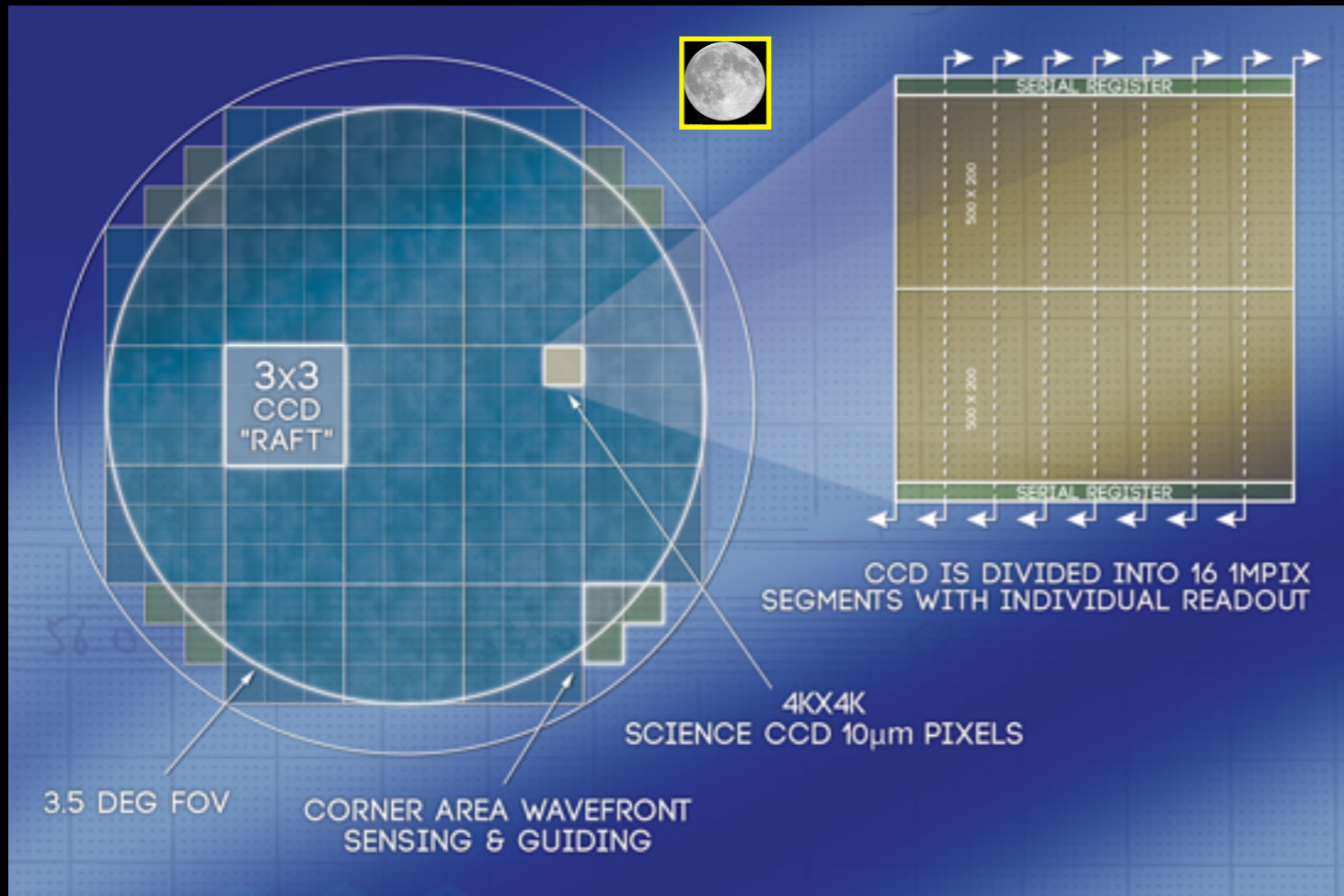
Construction has started  
Site: Cerro Pachon, Chile

Start of survey ~2020  
Database of 30-60 PByte

(LSST) 8m  $\varnothing$

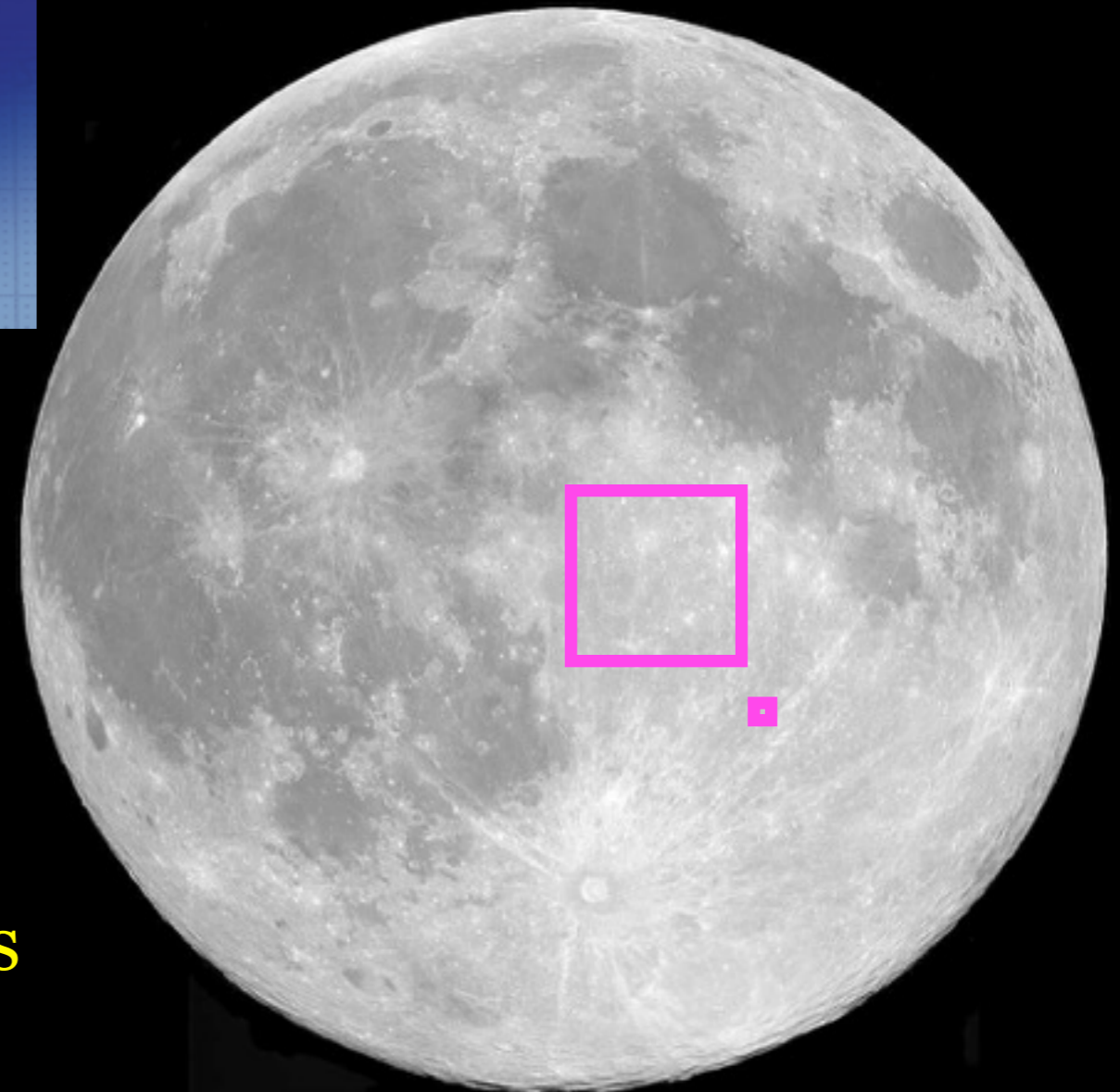
Consortium of US and  
international institutions





ELT field of view  
0.1 (0.001) deg  $\emptyset$

LSST field of view  
3.5 deg  $\emptyset$



The LSST will visit each patch of the southern sky >1000 times in ten years

## LSST key science goals:

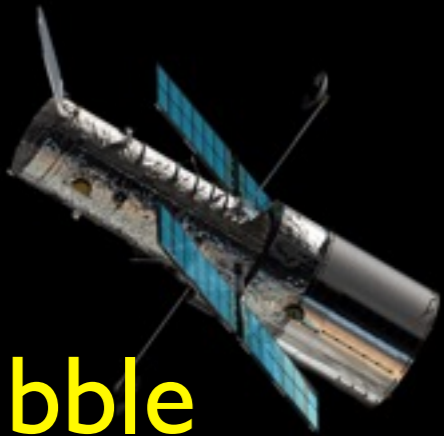
**Nature of Dark Energy (and Dark Matter)**  
(weak lensing cosmic shear, baryonic acoustic oscillations, galaxy cluster counts, ...)

**Explore the Transient Universe**  
( $\gamma$ -ray bursts, supernovae, near-Earth objects, ...)



And of course, there is space-based astronomy...

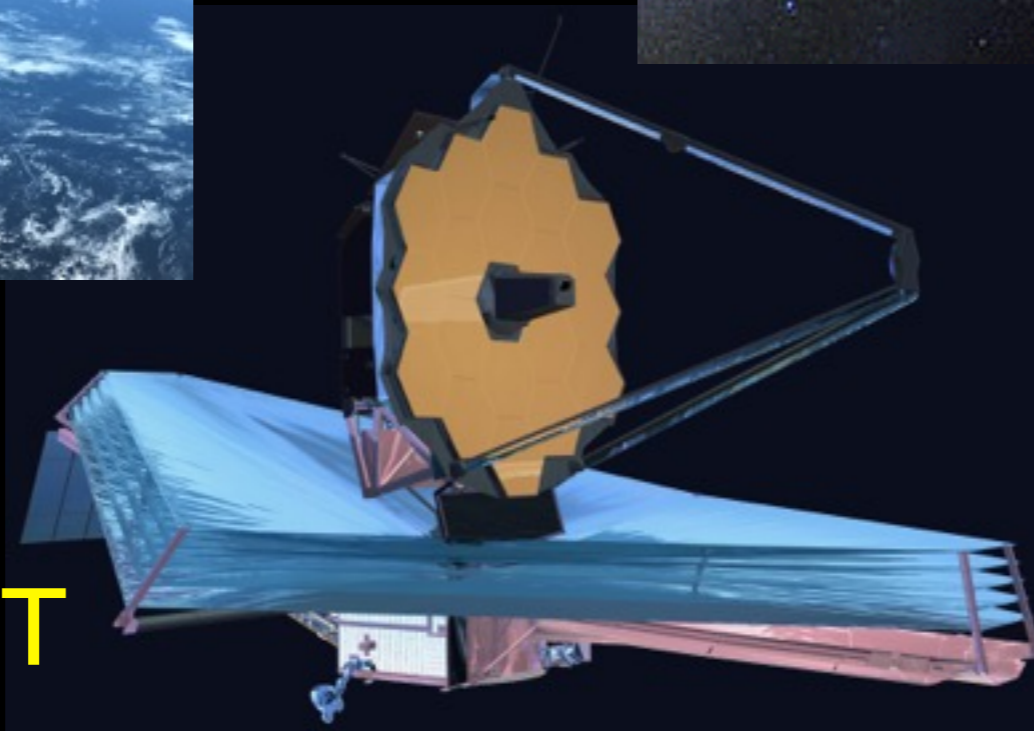
Hubble



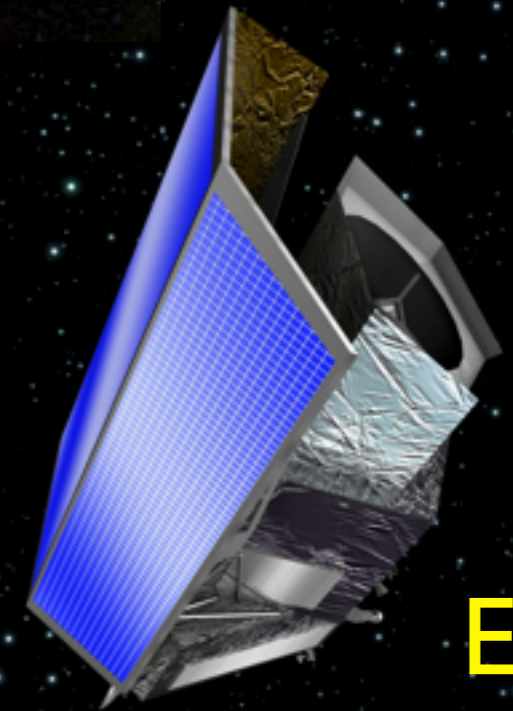
Gaia



JWST



Euclid



... but this would be another talk

Thank you!

