LUMINOSITY GOALS FOR A 100-TEV PP COLLIDER

Ian Hinchliffe^a, Ashutosh Kotwal^b, Michelangelo L. Mangano^c, Chris Quigg^d, Lian-Tao Wang^e

^a Phyiscs Division, Lawrence Berkeley National Laboratory, Berkeley CA 94720, USA

^b Fermi National Accelerator Laboratory, Batavia, Illinois 60510, USA Duke University, Durham, North Carolina 27708, USA

^c PH Department, TH Unit, CERN, CH-1211 Geneva 23, Switzerland

^d Theoretical Physics Department, Fermi National Accelerator Laboratory P.O. Box 500, Batavia, Illinois 60510 USA Institut de Physique Théorique Philippe Meyer, École Normale Supérieure 24 rue Lhomond, 75231 Paris Cedex 05, France

^e Department of Physics and Enrico Fermi Institute, University of Chicago, Chicago, IL 60637 USA

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Abstract

We consider diverse examples of science goals that provide a framework to assess luminosity goals for a future 100-TeV proton-proton collider.

- what are the physics drivers of the luminosity goals?
- how ambitious should the luminosity goals be?
- is there a minimum acceptable luminosity?

Useful tool to explore luminosity/energy dependence of discovery reach:

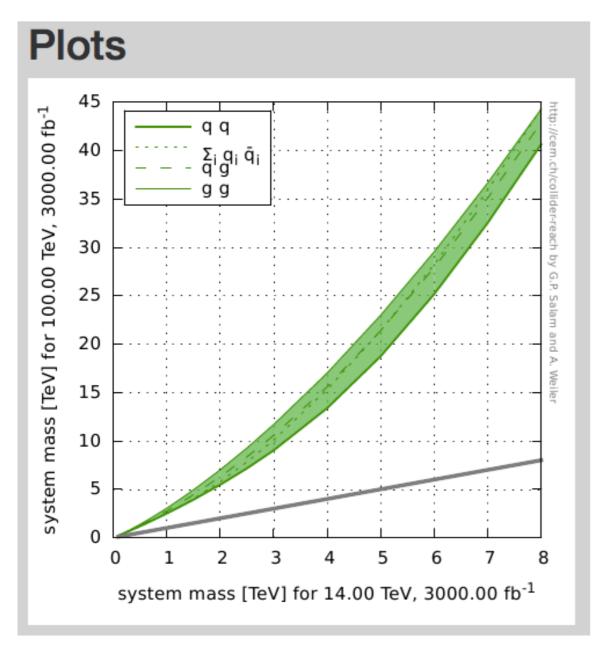
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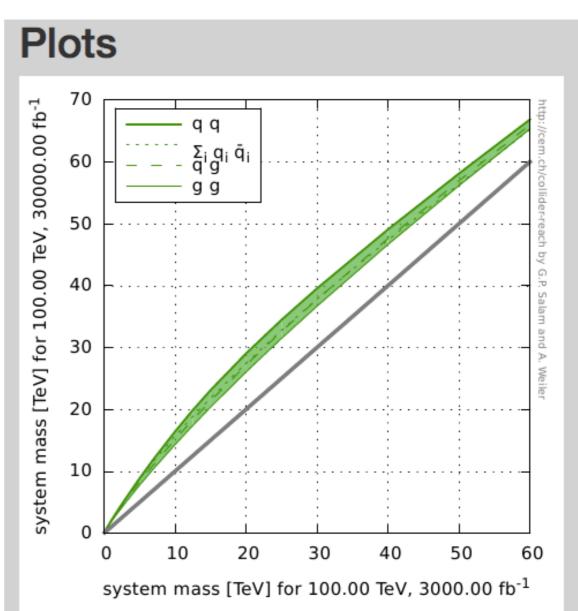
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Collider 1: CoM energy	14	TeV, integrated luminosity	3000	fb⁻1	
Collider 2: CoM energy	100	TeV, integrated luminosity	3000	fb ⁻¹	
PDF: MSTW2008nnlo68cl	\$				

G.Salam and A.Weiler, http://cern.ch/collider-reach

Collider 1: CoM energy	100	TeV, integrated luminosity	3000	fb ⁻¹
Collider 2: CoM energy	100	TeV, integrated luminosity	30000	fb ⁻¹
PDF: MSTW2008nnlo68cl	\$			

Submit





Recent papers addressing the luminosity issue

Mass Reach Scaling for Future Hadron Colliders, T.Rizzo, http://arxiv.org/abs/1501.05583 High Energy Colliding Beams; What Is Their Future? B. Richter, http://arxiv.org/abs/1409.1196 " restricting the luminosity to what will be achieved at HL-LHC gives the new machine a limited vision, and will (and **should**) **seriously lower the likelihood that it will be funded**."

... question is: what does it mean to "restrict the luminosity that will be achieved"? Should L necessarily scale like E_{beam}^2

Ultimate Luminosity must guarantee:

• Extension of the discovery reach at the high mass end

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Initial Luminosity should allow to rapidly (~1st year) surpass the exploration potential of the LHC

$$\sigma(M,g) \propto \frac{g^2}{M^2} L(x = M/\sqrt{S})$$

At fixed mass, cross sections grow when S grows, since

$$L(x) \sim \frac{1}{x^{\alpha}} \log(\frac{1}{x}) , \quad \alpha < 1 \qquad \begin{array}{l} \text{assuming} \\ \mathbf{f}(\mathbf{x}) \sim \mathbf{I}/\mathbf{x}^{\mathbf{I} + \alpha} \end{array}$$

To scale the discovery reach in mass as the growth in energy, means however to keep $x=M/\sqrt{S}$ constant. Then

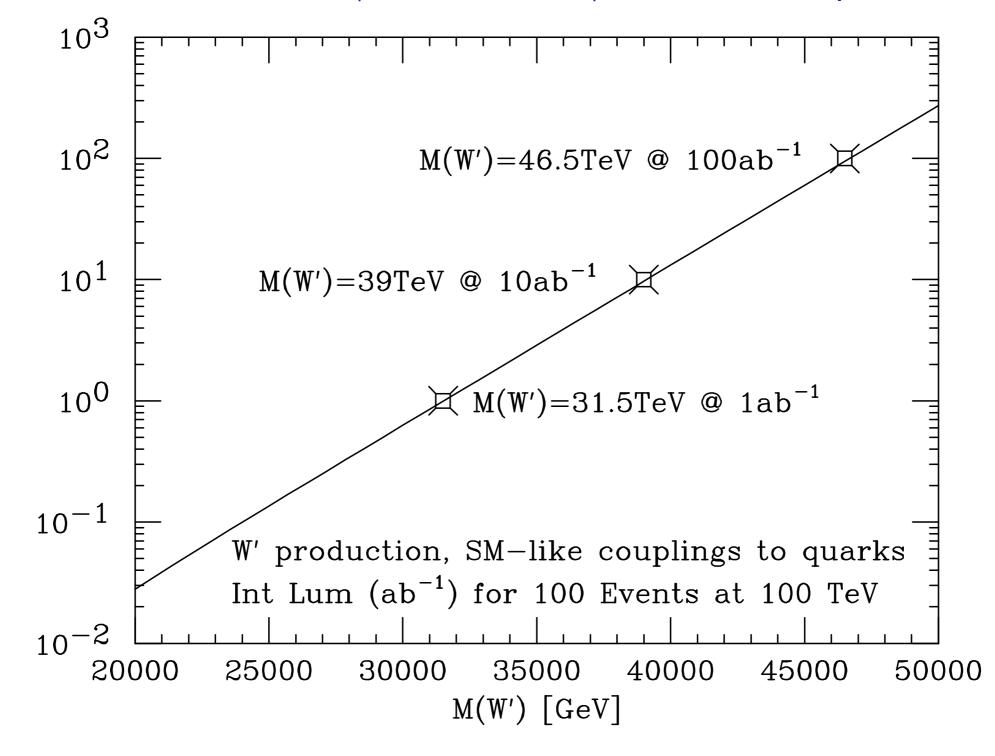
$$\sigma(M,g) \propto \frac{g^2}{S} \frac{L(x)}{x}$$

Thus the cross-sections for searches go like I/S, and the machine luminosity may need to grow accordingly.

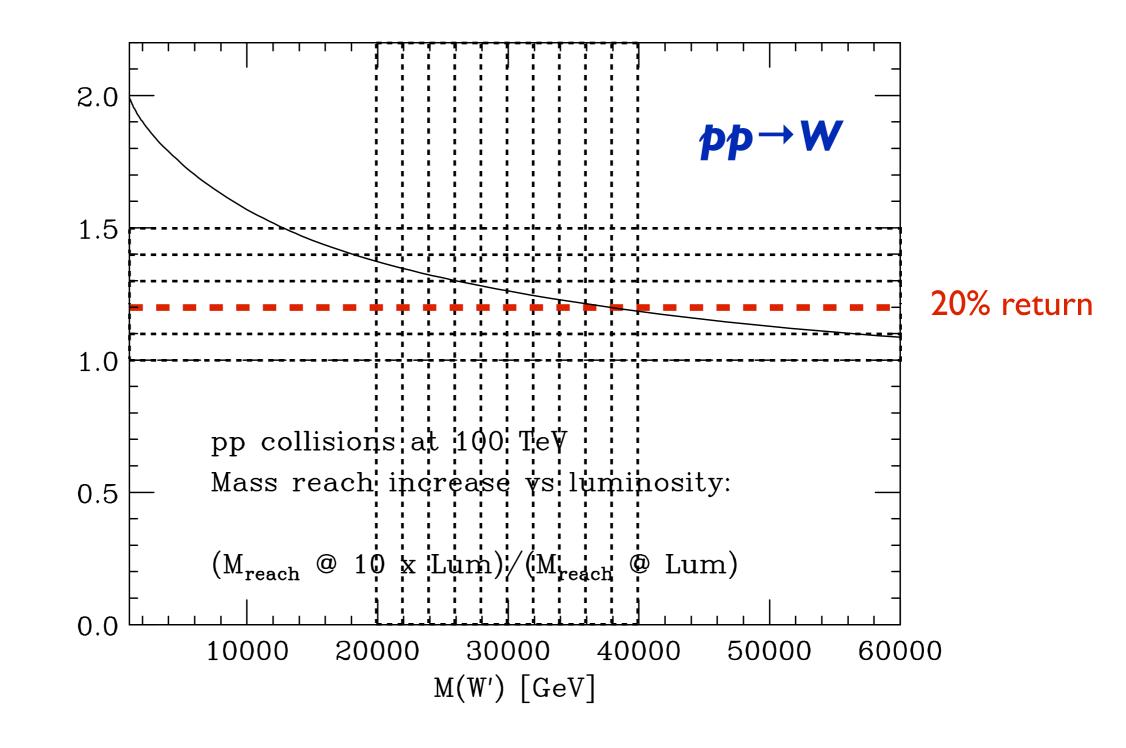
Extension of the discovery reach at high mass

Example: discovery reach of W' with SM-like couplings

NB For SM-like Z', $\sigma_{Z'} BR_{lept} \sim 0.1 \times \sigma_{W'} BR_{lept}$, \Rightarrow rescale lum by ~ 10



At L=O(ab⁻¹), Lum x 10 $\Rightarrow \sim M + 7 \text{ TeV}$



Lum x $10 \Rightarrow$ relative gain much larger at low mass than at high mass

 One could argue that the 10 x increase in lum is not justified if the increase in sensitivity is below a level of O(20%) (unless there is a concrete physics case, e.g. testing a possible recurrent spectrum of resonances)

See e.g. the history of Tevatron achievements: after 1 fb⁻¹, limited progress at the highmass end, but plenty of results at "low" mass (W, top and b physics, Higgs sensitivity,)

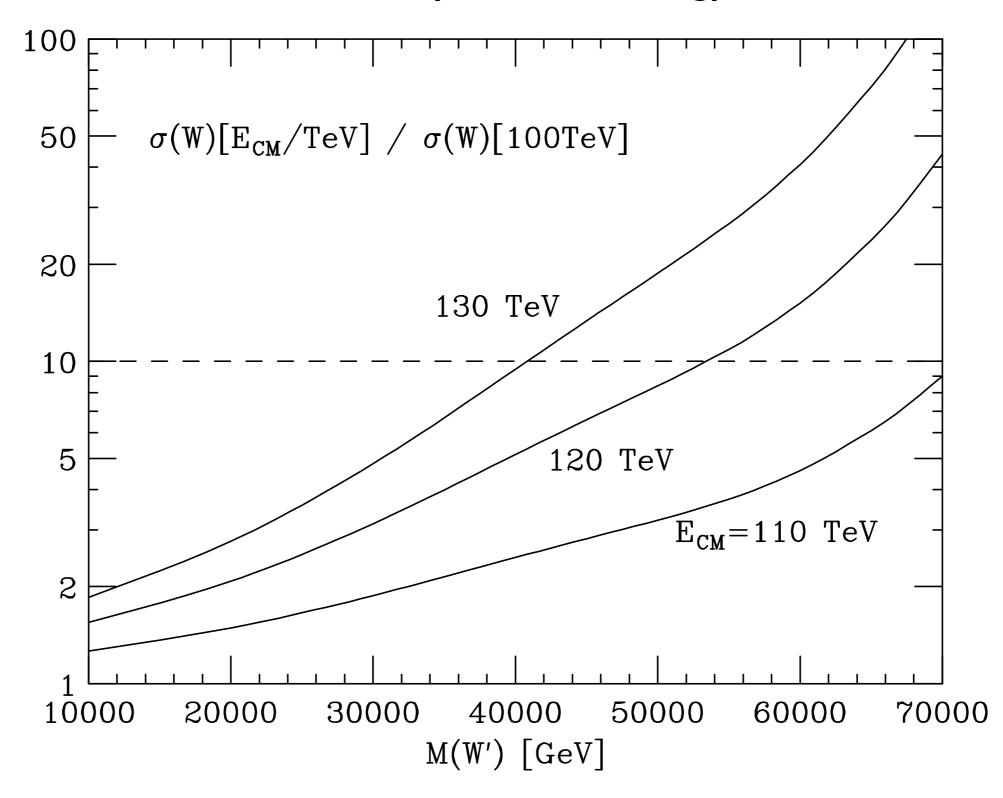
Example from HL-LHC studies: Z' \rightarrow e⁺e⁻

ATLAS/CMS HL docs	300/fb	3000/fb
95% excl (ATLAS)	6.5 TeV	7.8 TeV
5σ (CMS)	5.1 TeV	6.2 TeV

• $\Delta M/M \sim 20\% \Rightarrow$ the LHC reaches the threshold of saturation of the mass reach already at

 300fb^{-1} . Notice that 95% exclusion at 300 makes unlikely the 5 σ discovery at 3000. In fact the main justification for the HL-LHC is the higher-statistics study of the Higgs, not the extension of the mass reach

Luminosity vs CM Energy



- At around 40 TeV, a 20% increase in energy buys a factor of 5 in rate. 30% in energy buys a factor 10 in rate.
- What will be less challenging ? To upgrade the magnets, or to increase Lx10?

Extension of the discovery reach at low mass

- The extension power of higher lum can be important at lower masses, e.g. for processes with very suppressed rates, or difficult to separate from the bg.
- In this case, though, one might benefit more from improved detection efficiency than from pure luminosity.
- The luminosity discussion is extremely process dependent (bg's, detector performance, pileup issues, etc)

Example: direct stop production

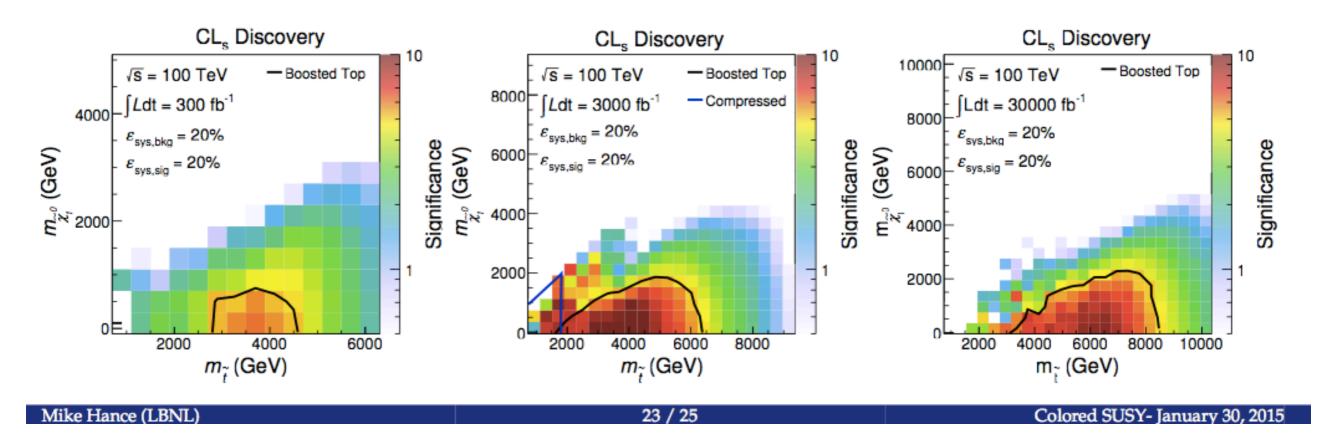
from Mike Hance's talk

$pp \rightarrow \tilde{t}\tilde{t}^* \rightarrow t\tilde{\chi}_1^0 \bar{t}\tilde{\chi}_1^0 - \text{Beyond 3 ab}^{-1}$

arXiv:1406.4512

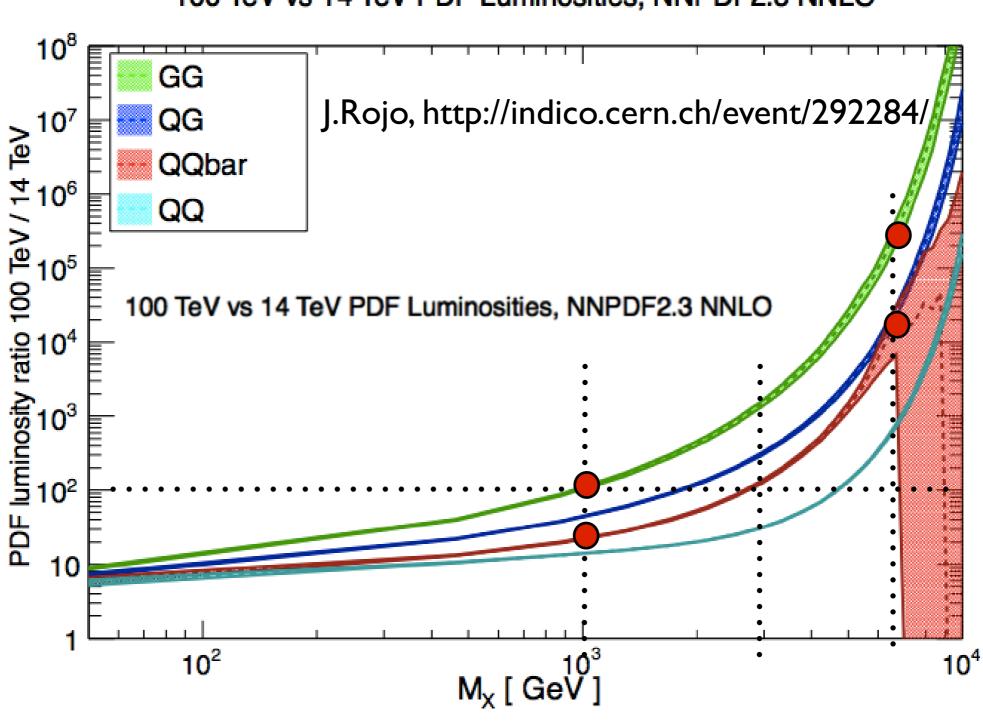
Will 3 ab^{-1} be enough at 100 TeV?

• Scale $E_{\rm T}^{\rm miss}$ cuts for higher masses, going from 0.3 ab⁻¹ to 30 ab⁻¹



Recognizing that higher luminosity is mostly needed to better explore "low" masses, rather than the highest masses, may lead to different perspective on the design of detectors

Higher statistics for studies of particles discovered at the LHC



100 TeV vs 14 TeV PDF Luminosities, NNPDF2.3 NNLO

At the edge of the HL-LHC discovery reach, namely $m_X \sim 6.5 \text{ TeV}$: $\sigma(100 \text{ TeV}) / \sigma(14 \text{ TeV}) \sim \begin{cases} 10^4 \text{ for } q - q \text{ bar} \rightarrow X \\ 10^5 \text{ for } gg \rightarrow X \end{cases}$

 \Rightarrow improve by orders of magnitude the precision of the measurements of particle **X** discovered at the mass-end of the LHC reach

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At lower masses the increase is less pronounced. $m_X \sim I \text{ TeV}:$ $\sigma(100 \text{ TeV}) / \sigma(14 \text{ TeV}) \sim \begin{cases} \sim 25 \text{ for } q - q \text{ bar} \rightarrow X \\ \sim 10^2 \text{ for } gg \rightarrow X \end{cases}$ At the edge of the HL-LHC discovery reach, namely $m_X \sim 6.5 \text{ TeV}$: $\sigma(100 \text{ TeV}) / \sigma(14 \text{ TeV}) \sim \begin{cases} 10^4 \text{ for } q - qbar \rightarrow X \\ 10^5 \text{ for } gg \rightarrow X \end{cases}$

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Once again, it's the "low"-mass physics that benefits the most from luminosity

Higher statistics for Higgs studies

	NLO rates					
	σ(14 TeV)	R(33)	R(40)	R(60)	R(80)	R(100)
ggH	50.4 pb	3.5	4. 6	7.8	11.2	14.7
VBF	4.40 pb	3.8	5.2	<mark>9</mark> .3	13.6	18.6
wн	1.63 pb	2.9	3.6	5.7	7.7	9.7
ZН	0.90 pb	3.3	4.2	<mark>6.</mark> 8	9.6	12.5
ttH	0.62 pb	7.3	11	24	41	61
нн	33.8 fb	6.1	8.8	18	29	42

 $R(E) = \sigma(E \text{ TeV})/\sigma(14 \text{ TeV})$

• Gains in the range 10-50, however

NIO rates

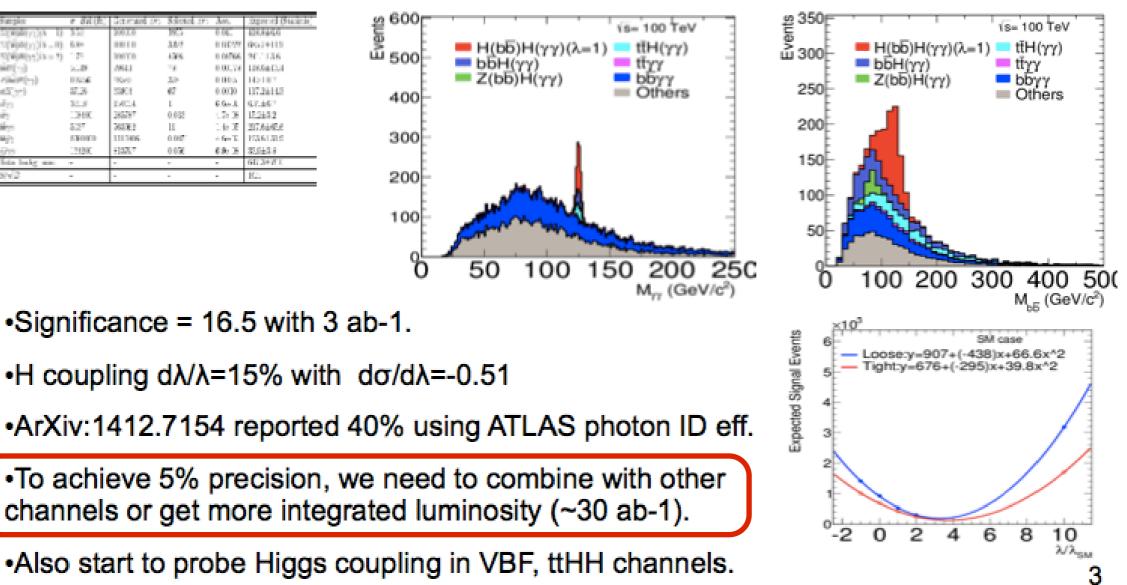
 => needs detailed studies, considering also the prospects to study rare decays, selfcouplings, etc.etc.

Example: H selfcoupling at 100 TeV

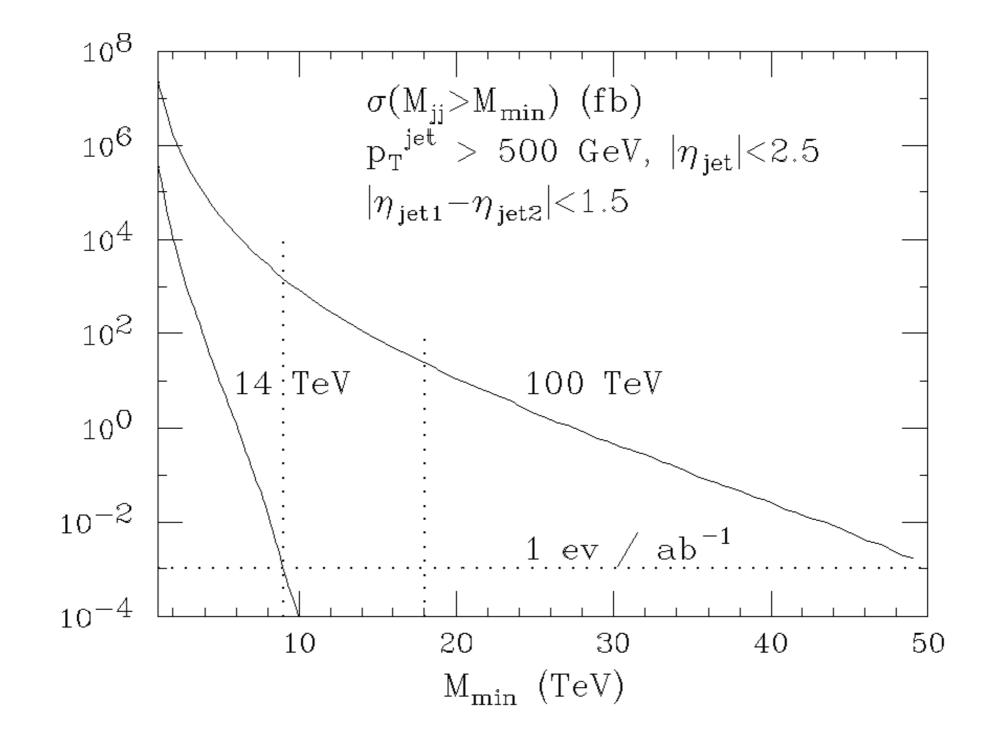
W.Yao, update of http://arxiv.org/abs/1308.6302, shown at "IAS programme on The Future of High Energy Physics", Hong Kong, January 2015 http://ias.ust.hk/program/shared_doc/201501fhep/Weiming%20Yao_Jan%2021.pdf

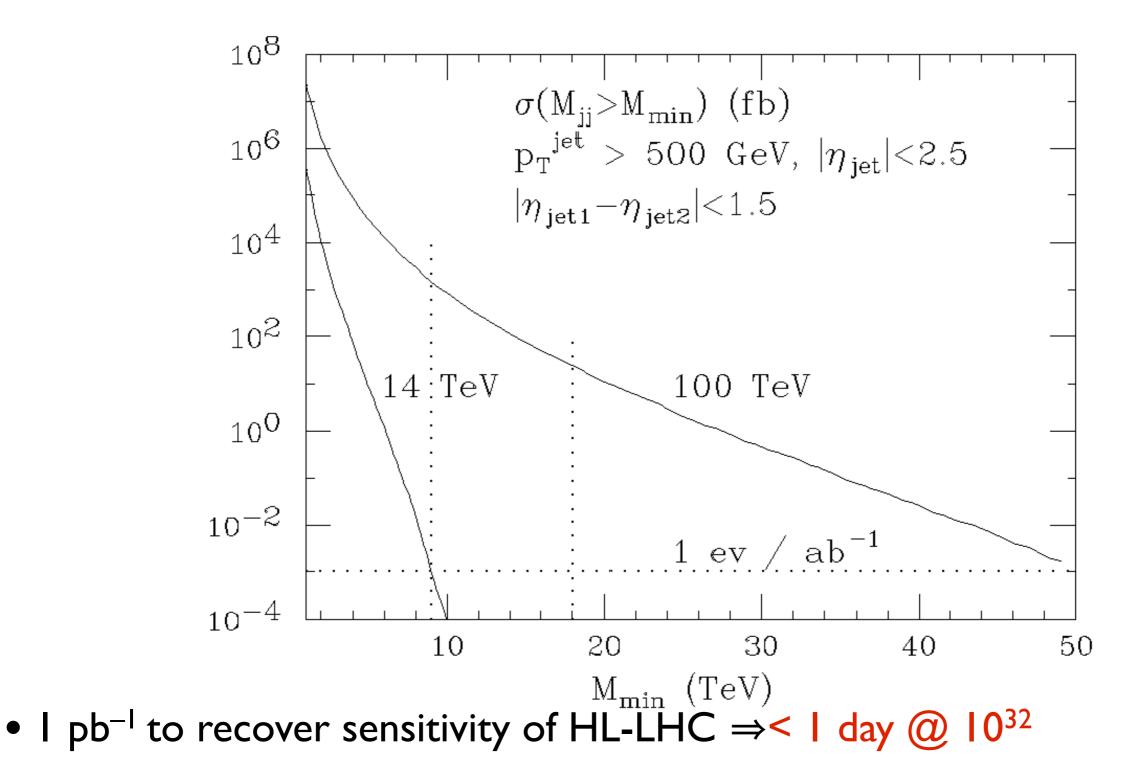
Updating HH \rightarrow bbyy at Tev100

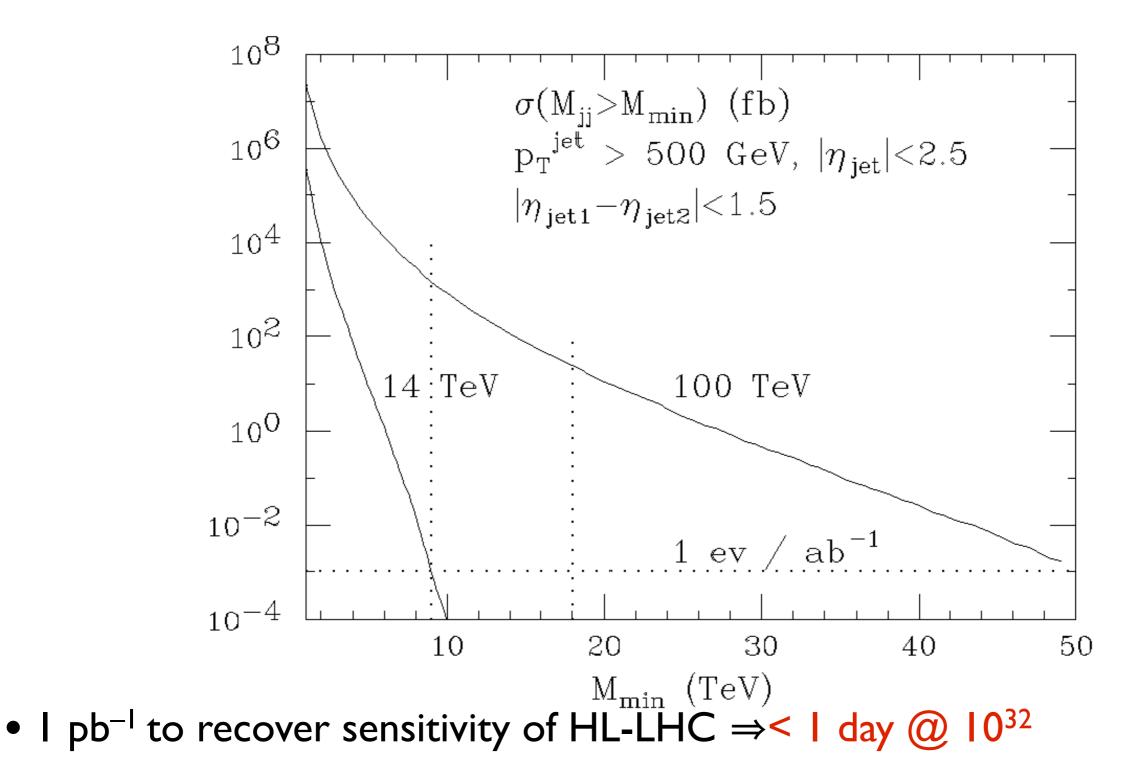
- Using Delphes 3.1.14 and the results depends on detector performace assumed.
- Including jjyy, bbjy, tty, ttyy with ATLAS fy=0.0093e(-Et/27.5) for HL-LHC
- Tighten myy window from 10 GeV used for snowmass to 6 GeV.



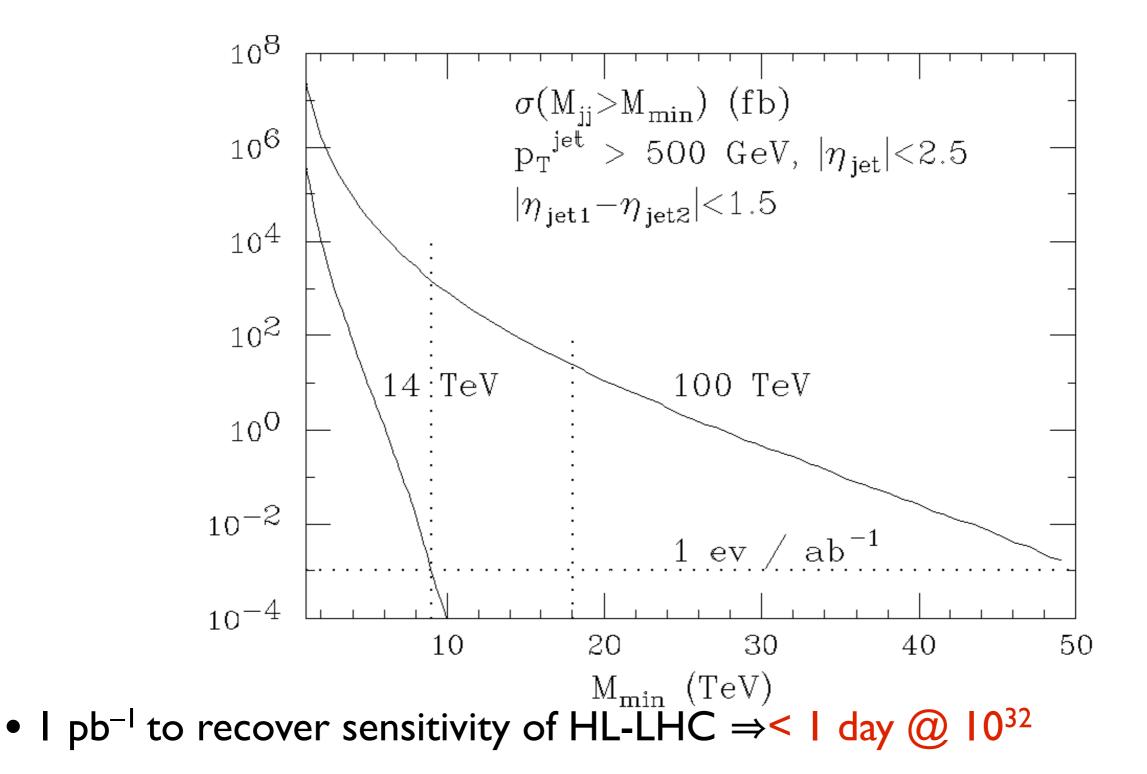
Initial luminosity, or: what's the minimum lum to take us beyond the HL-LHC ?







• 50pb⁻¹ to 2x the sensitivity of HL-LHC \Rightarrow < 1 month @ 10³²



- 50pb⁻¹ to 2x the sensitivity of HL-LHC \Rightarrow < 1 month @ 10³²
- Ifb⁻¹ to 3x the sensitivity of HL-LHC \Rightarrow < 1 year @ 2x10³²

For resonances: at the edge of the HL-LHC discovery reach, namely m_x ~ 6.5 TeV :

$$\sigma(100 \text{ TeV}) / \sigma(14 \text{ TeV}) \sim \begin{cases} 10^4 \text{ for } q - q bar \rightarrow X \\ 10^5 \text{ for } gg \rightarrow X \end{cases}$$

This means:

- If X is discovered at the HL-LHC, it can be confirmed at 100 TeV with $10^{-(4+5)}$ of the HL-LHC luminosity, i.e. O(30-300 pb⁻¹)
 - => $L < 5 \times 10^{31}$ in the 1st year
- A luminosity of O(0.1 1 fb⁻¹) allows the discovery of particles just beyond the HL-LHC reach
 - => L < 2×10^{32} in the 1st year

• The goal of an integrated luminosity in the range of 10-20 ab⁻¹ per experiment, corresponding to an ultimate instantaneous luminosity approaching 2×10³⁵, seems well-matched to our current perspective on extending the discovery reach for new phenomena at high mass scales, high- statistics studies of possible new physics to be discovered at (HL)-LHC, and incisive studies of the Higgs boson's properties.

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- Specific measurements may set more aggressive luminosity goals, but we have not found generic arguments to justify them. The needs of precision physics arising from new physics scenarios to be discovered at the HL-LHC, to be suggested by anomalies observed during the e+e- phase of a future circular collider, or to be discovered at 100 TeV, may well drive the need for even higher statistics. Such requirements will need to be established on a case-by-case basis, and no general scaling law gives a robust extrapolation from 14 TeV. Further work on ad hoc scenarios, particularly for low-mass phenomena and elusive signatures, is therefore desirable.

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- For a large class of new-physics scenarios that may arise from the LHC, less aggressive luminosity goals are acceptable as a compromise between physics return and technical or experimental challenges. In particular, even luminosities in the range of 10³² are enough to greatly extend the discovery reach of the 100 TeV collider over that of the HL-LHC, or to enhance the precision in the measurement of discoveries made at the HL-LHC