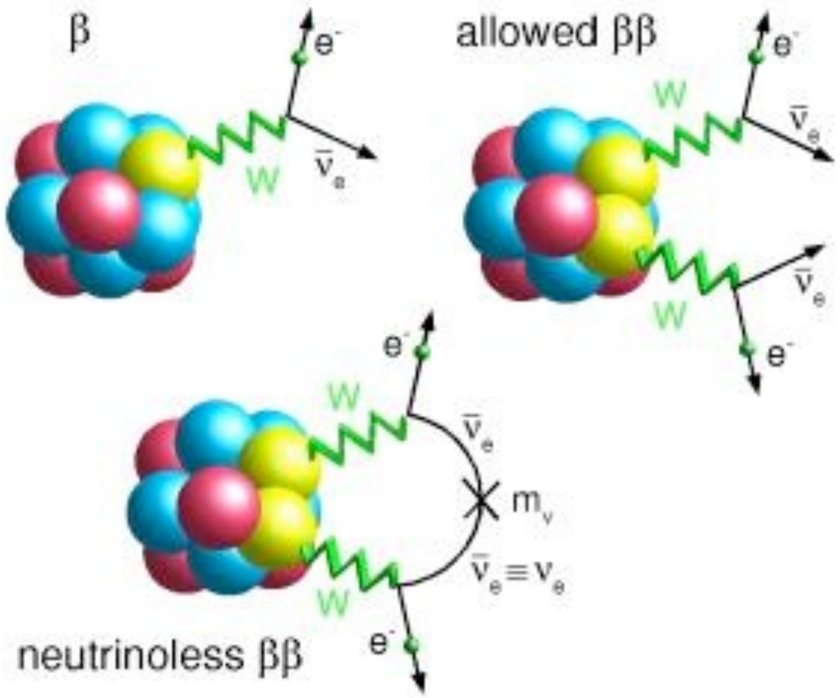
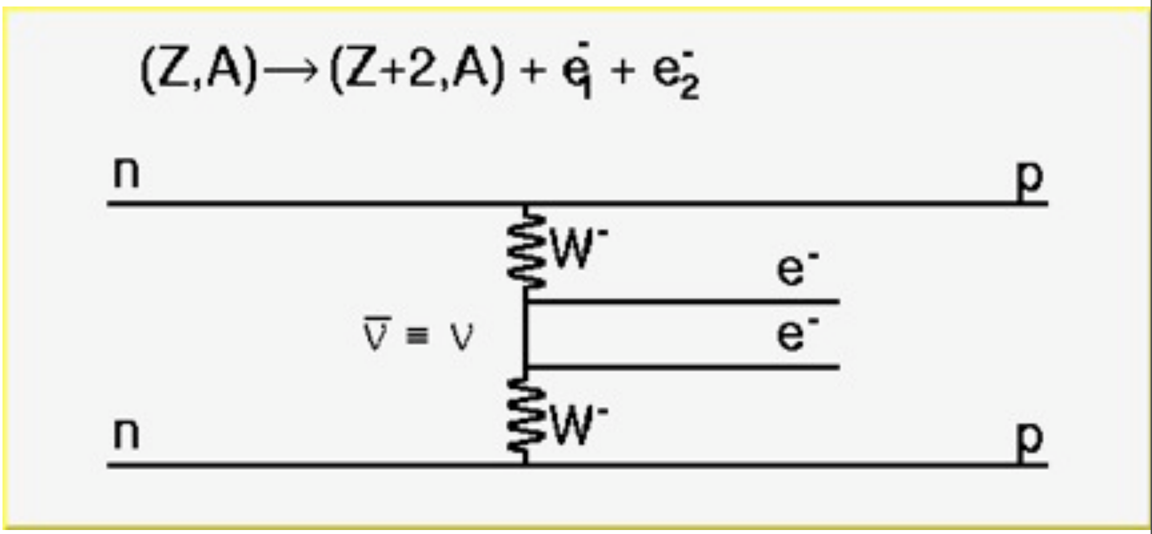




Neutrino mass and neutrinoless double beta decay (0νββ)

Ruben Saakyan, UCL

Annual IOP HEPP Conference

Oxford

8 April 2009

Outline

- Double Beta Decay and neutrino mass
- Experimental techniques
- Current and Future projects
  - Disclaimer: Weighted/biased by UK participation

Neutrinos are massive and they mix

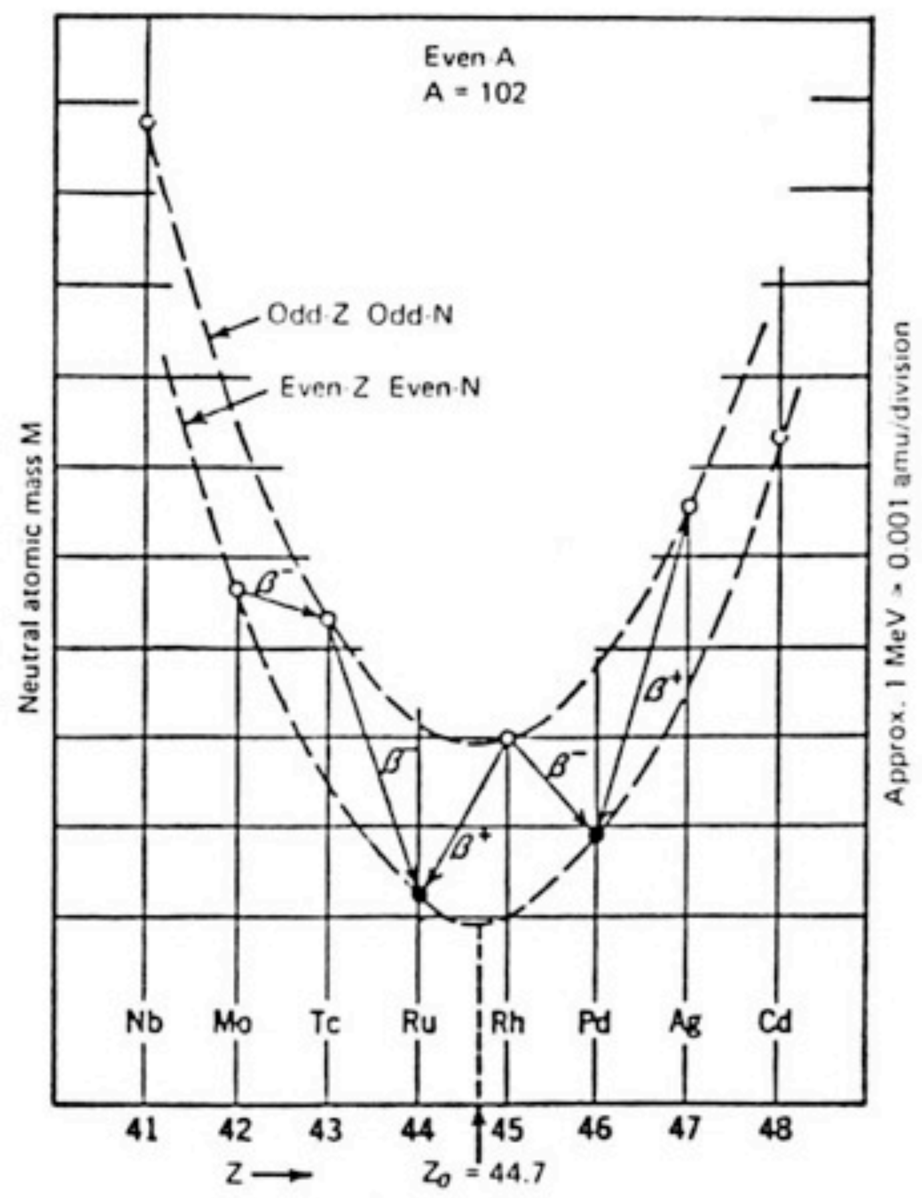
What else do we want to know?

- Number of neutrinos: Are there sterile neutrinos?
- θ_{13}, Precision values of mixing angles and Δm^2's
- Absolute neutrino mass value. Only limits so far.
  Tritium: $m_{\bar{\nu}_e} < 2.3$ eV Cosmology: $\sum m_{\nu_i} < 1$ eV
- Neutrino mass spectrum: Normal ($m_1 < m_2 < m_3$)
  Inverted ($m_3 < m_1 < m_2$) or Quasi-degenerate ($m_1 \approx m_2 \approx m_3$)?
- Origin of matter-antimatter asymmetry.
  CP-violation in lepton sector: $\delta \neq 0, \pi$ and/or $\alpha, \beta \neq 0, \pi$?
- Nature of Neutrinos: Majorana (ν = anti-ν) or Dirac ($\nu \neq$ anti-ν)?
  Full lepton number violation (required in most Grand Unification Theories) .

addressed by $0\nu\beta\beta$ decay

Nuclear Physics and Standard Model ββ decay

For most even-even nuclei only ββ decay is possible (recall pairing term in SEMF!)



^{76}Ge , ^{100}Mo, ^{82}Se, ^{116}Cd, ^{150}Nd, ^{48}Ca, ^{136}Xe ,^{96}Zr, ^{130}Te

$$(A,Z) \rightarrow (A,Z+2) + 2e^- + 2\bar{\nu}_e$$

$$(A,Z) \rightarrow (A,Z-2) + 2e^+ + 2\nu_e$$

$$2e^- + (A,Z) \rightarrow (A,Z-2) + 2\nu_e$$

phase space

$$\frac{1}{T_{1/2}^{2\nu}} = G^{2\nu}(Q_{\beta\beta}, Z)\left|M^{2\nu}\right|^2$$

NME is **measured** in $2\nu\beta\beta$

NME: ~~Nasty~~ Nuclear Matrix Element

Measured for 10 nuclei

Important input for $0\nu\beta\beta$ NME calculation!

Important to understand its background contribution

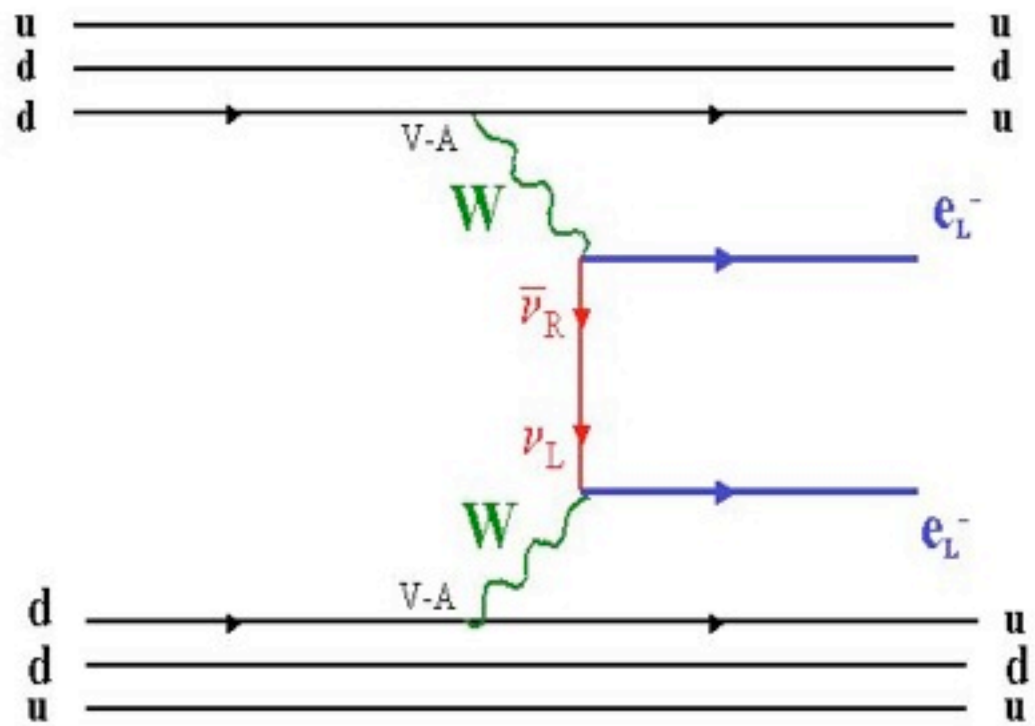
Neutrinoless double beta decay ($0\nu\beta\beta$)

$$(A,Z) \longrightarrow (A,Z+2) + 2\,e^-$$

(also $\beta^+\beta^+$ and 2K-capture possible)

$\Delta L = 2$

Lepton number violation!!!



Light neutrino exchange

Majorana neutrino ($\nu=\bar{\nu}$)

Access to absolute neutrino mass

Phase space factor **Nuclear matrix element**

$$T_{1/2}^{-1} = F(Q_{\beta\beta}^{5},Z)\,|M^{0\nu}|^2\,\langle m_\nu\rangle^2$$

$$\langle m_\nu\rangle = m_1|U_{e1}|^2 + m_2|U_{e2}|^2.e^{i\alpha} + m_3|U_{e3}|^2.e^{i\beta}$$

$|U_{ei}|$: mixing matrix elements

α and β: Majorana phases

Other possible process :

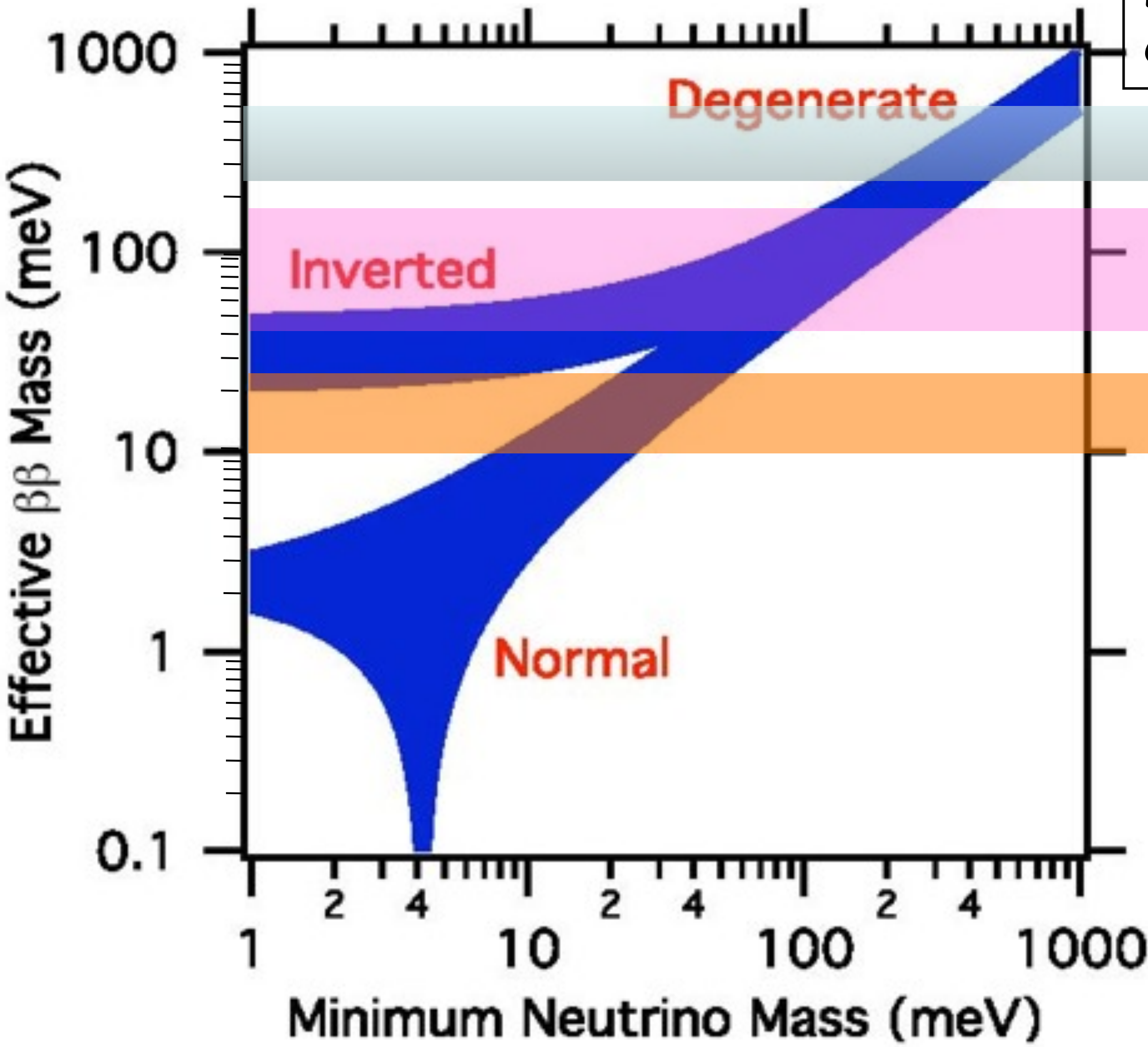
V+A current : $\langle m_\nu\rangle, \langle\lambda\rangle, \langle\eta\rangle$

Majoron emission : $\langle g_M\rangle$

Supersymmetry : $\lambda'_{111},\ \lambda'_{113}$

Schechter-Valle theorem:

$\beta\beta(0\nu) \longrightarrow$ **Majorana neutrinos**



NEMO3, CUORICINO running,
Complete verification by GERDA-I, SuperNEMO-Demonstrator

Klapdor et al. 0.24 – 0.58 eV

Coming up: x100 kg experiments: SuperNEMO, CUORE, GERDA, EXO, SNO+,...

Future 1 ton-scale experiments

Assumptions:

Majorana neutrinos
No cancellations

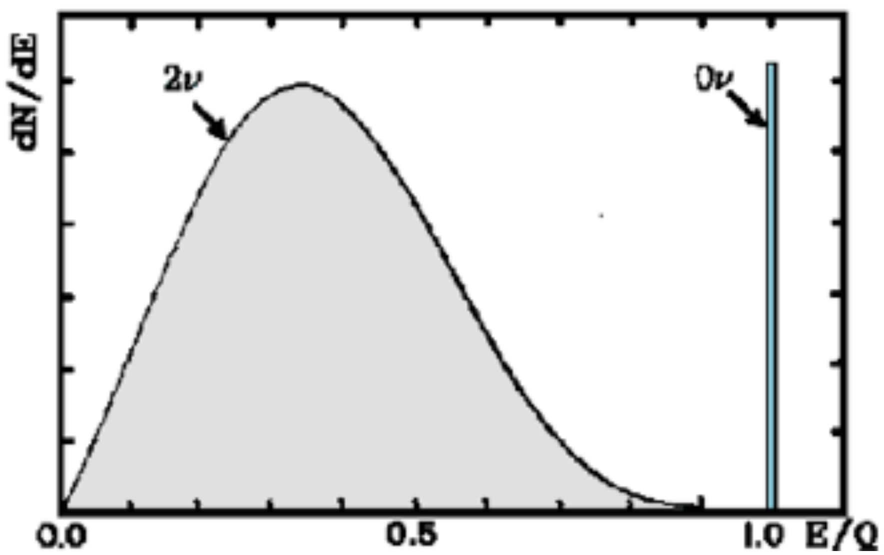
Plot from Avignone, Elliott, Engel arXiv:0708.1033 (2007)

depending on Majorana CPV phases:

$|\langle m_\nu \rangle| \simeq 0.1 - 0.5$ eV   Quasi-degenerate hierarchy ($m_1 \approx m_2 \approx m_3$) – best case

$|\langle m_\nu \rangle| \simeq 0.02 - 0.05$ eV   Inverted hierarchy ($m_3 < m_1 < m_2$) – tough but doable

$|\langle m_\nu \rangle| \simeq 0.0 - 0.006$ eV   Normal hierarchy ($m_1 < m_2 < m_3$) – really tough



Experimental Observables/Signatures

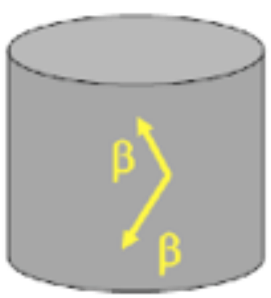
(1) Two coincident electrons from the the same vertex

(2) $E_{e1} + E_{e2} = Q_{\beta\beta}$

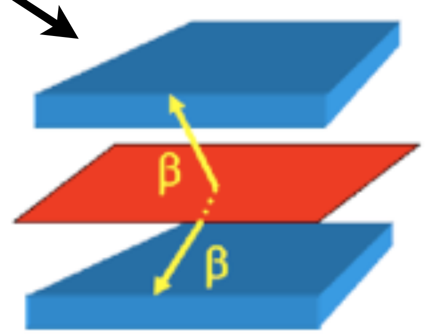
(3) Angular distribution

(4) Daughter nucleus ID (would be great)

Two Approaches



$$\frac{1}{T_{1/2}^{0\nu}} = G^{0\nu}(Q_{\beta\beta}, Z)\left|M^{0\nu}\right|^2 \eta^2$$



Lepton number violating parameter

η can be due to $\langle m_\nu \rangle$

V+A, Majoron, SUSY, H-- or a combination of them

Source = detector (calorimeters)

Great ΔE/E

compact

Source ≠ detector (foil+tracking+calorimetry)

isotope flexibility

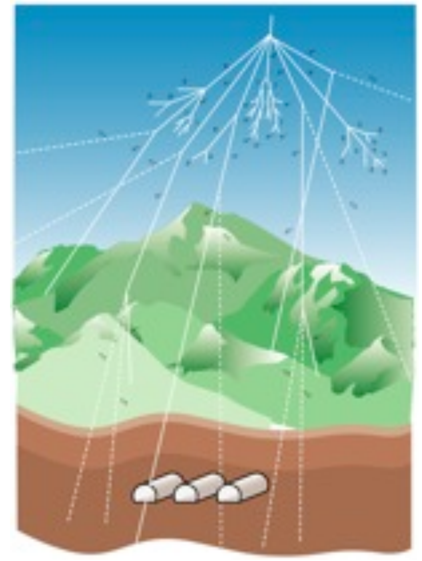
"smoking" gun

Topological bkg suppression

Need detectors which can probe different mechanisms (and different isotopes)

Backgrounds

Cosmic rays



Underground Lab is a must

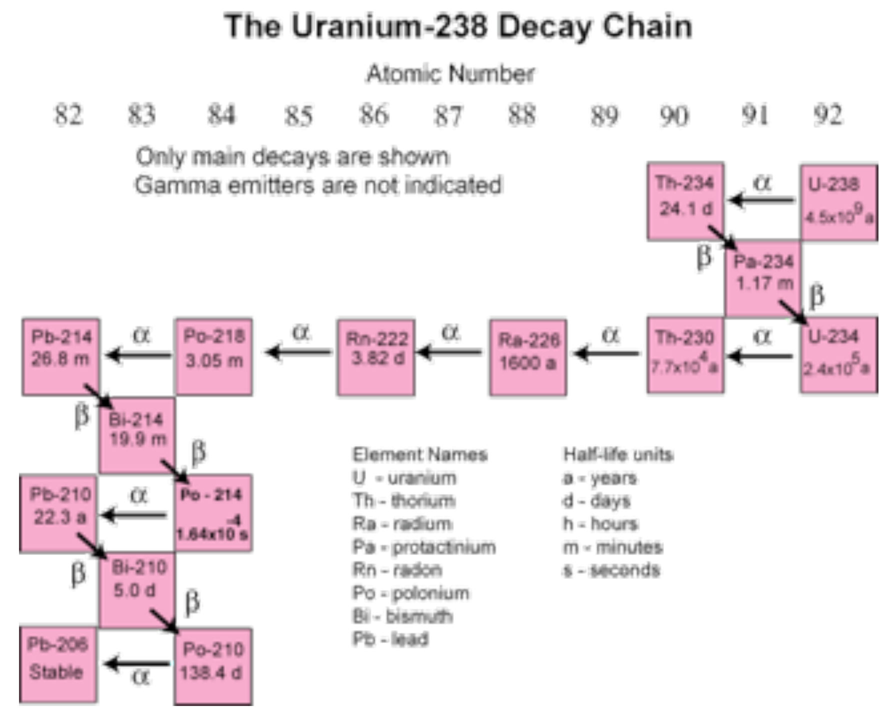
Natural Radioactivity

$T_{1/2}(^{238}U, ^{232}Th) \sim 10^{10}$ yr
$T_{1/2}(0\nu\beta\beta) > 10^{25}$ yr

Main threat from:

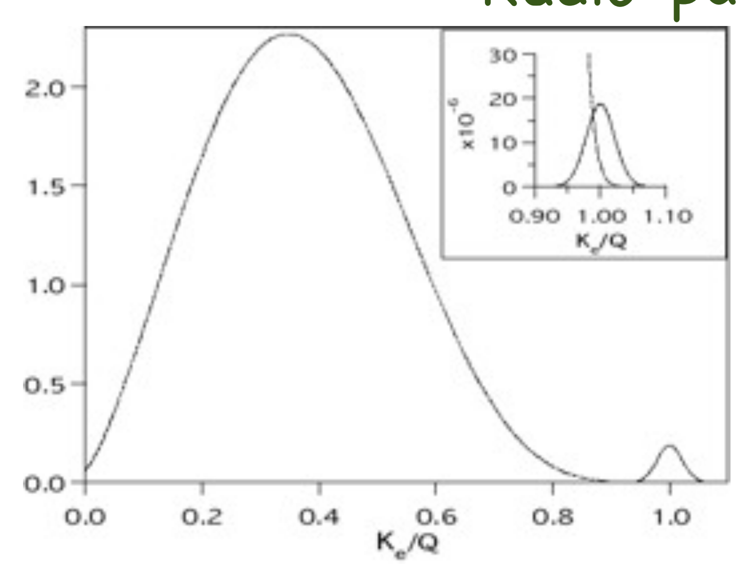
^{214}Bi (Q_β = 3.27 MeV)

^{208}Tl(Q_β = 4.99 MeV)



Radio-purity and bkg identification

Standard Model $2\nu\beta\beta$



Irreducible bkg
Energy resolution
is the only
weapon here

Isotopes

Deuterium	2
Helium gas	3
Lithium	6 7
Carbon	13
Silicon	28 29 30
Sulphur	32 33 34 36
Chromium	50 52 54
Iron	54 56 57 58
Zinc /Depleted Zinc/	64 66 67 68 70
Gallium	69 71
Germanium	70 72 73 74 76
Selenium	74 76 77 78 80 82
Krypton	78 80 82 83 84 86
Molybdenum	96 97 98 100
Cadmium	108 110 111 112 113 114 116
Tellurium	122 123 124 128 130
Xenon	124 126 128-132 134 136
Osmium	184 186 187 188 189 190 192
Lead	204 206 207 20

High $Q_{\beta\beta}$ is important due to phase space and natural radioactivity considerations

Isotope	$Q_{\beta\beta}$ (MeV)	Abund (%)
$^{48}Ca \rightarrow ^{48}Ti$	4.271	0.187
$^{76}Ge \rightarrow ^{76}Se$	2.040	7.8
$^{82}Se \rightarrow ^{82}Kr$	2.995	9.2
$^{96}Zr \rightarrow ^{96}Mo$	3.350	2.8
$^{100}Mo \rightarrow ^{100}Ru$	3.034	9.6
$^{110}Pd \rightarrow ^{110}Cd$	2.013	11.8
$^{116}Cd \rightarrow ^{116}Sn$	2.802	7.5
$^{124}Sn \rightarrow ^{124}Te$	2.228	5.64
$^{130}Te \rightarrow ^{130}Xe$	2.533	34.5
$^{136}Xe \rightarrow ^{136}Ba$	2.479	8.9
$^{150}Nd \rightarrow ^{150}Sm$	3.367	5.6

Centrifuge enrichment well established
x100 kg production possible

Indicative price: 30-60 k€/kg

Unfortunately not possible for
^{150}Nd, ^{48}Ca, ^{96}Zr
Alternative for these isotopes: AVLIS (Atomic Vapour Laser Isotope Separation)
Interesting developments at the MENPHIS facility in France.

General Strategy

$$T_{1/2}^{0\nu}(y) > \frac{\ln 2 \,.\, N}{k_{C.L.}} \cdot \frac{\varepsilon}{A} \sqrt{\frac{M \,.\, t}{N_{Bkg} \,.\, \Delta E}}$$

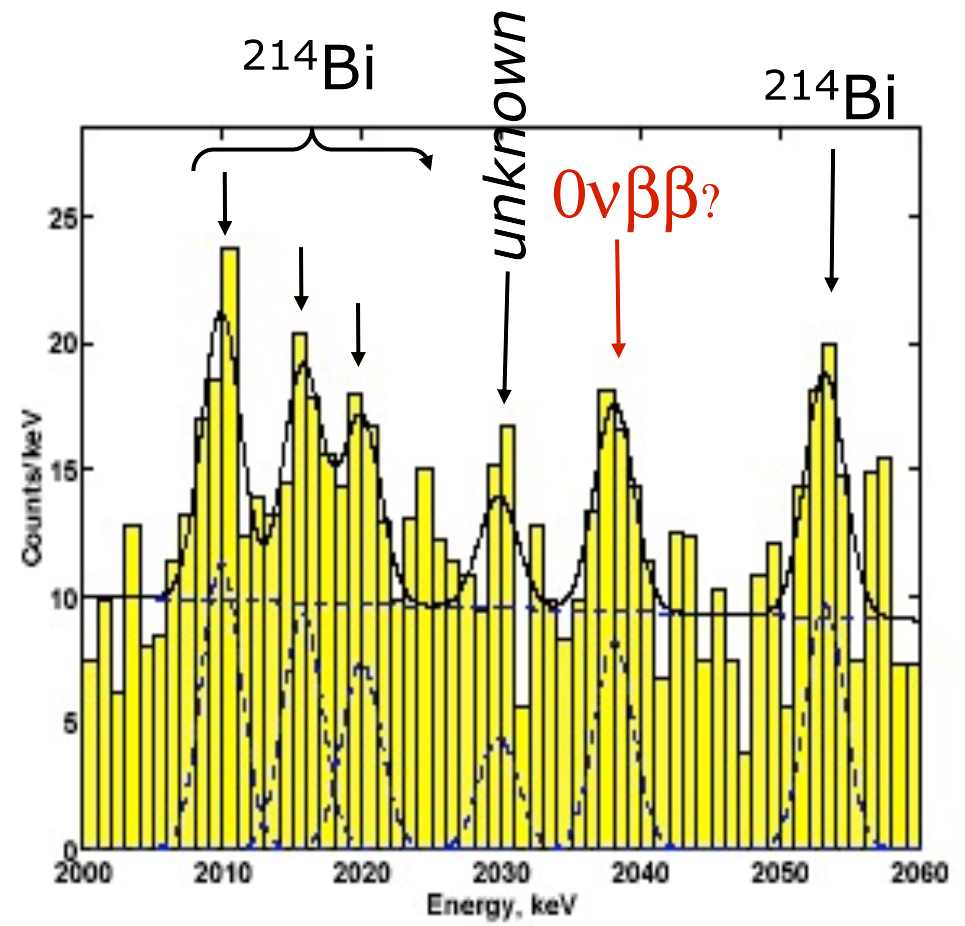
M: mass (kg)
ε : efficiency
K$_{C.L.}$: Confidence level
N: Avogadro number
t: time (y)
N$_{Bckg}$: Background events ($keV^{-1}.kg^{-1}.y^{-1}$)
ΔE: energy resolution (keV)

- Large mass (~100 kg to reach 0.05-0.1 eV level)
- High Efficiency
- Low $N_{Bkg} \,.\, \Delta E$ **Note:** The product is important when considering backgrounds
  - "Squeeze" ΔE - improve resolution
  - Lower N_{Bkg} - improve non-ββ background rejection (topology, particle ID etc)

The Claim

- HPGe detector (86% enriched)
- Full stat (71.71 kg y)
- Outstanding resolution 3.27 keV
- Unknown line at 2038 keV found
- I = 28.75 ± 6.86 events, 4.2σ
- $T_{1/2}$ = (0.69-4.18) 10^{25} y (3σ range, best fit = 1.19)

KKDC claim (subset of Heidelberg-Moscow collaboration)



$< m_\nu > = 0.1 - 0.9\,\mathrm{eV}$ (KKDC, 2004)

Can not be dismissed out of hand

BUT

- ✘ Background under-estimated
- ✘ Relative intensities problem with ^{214}Bi lines
- ✘ Unknown line in the same region

- ^{56}Co by cosmic rays (γ 2034keV+6keV X-ray)
- $^{76}Ge(n\gamma)^{77}Ge$ (2038 keV)
- An unknown line
- A combination of the above

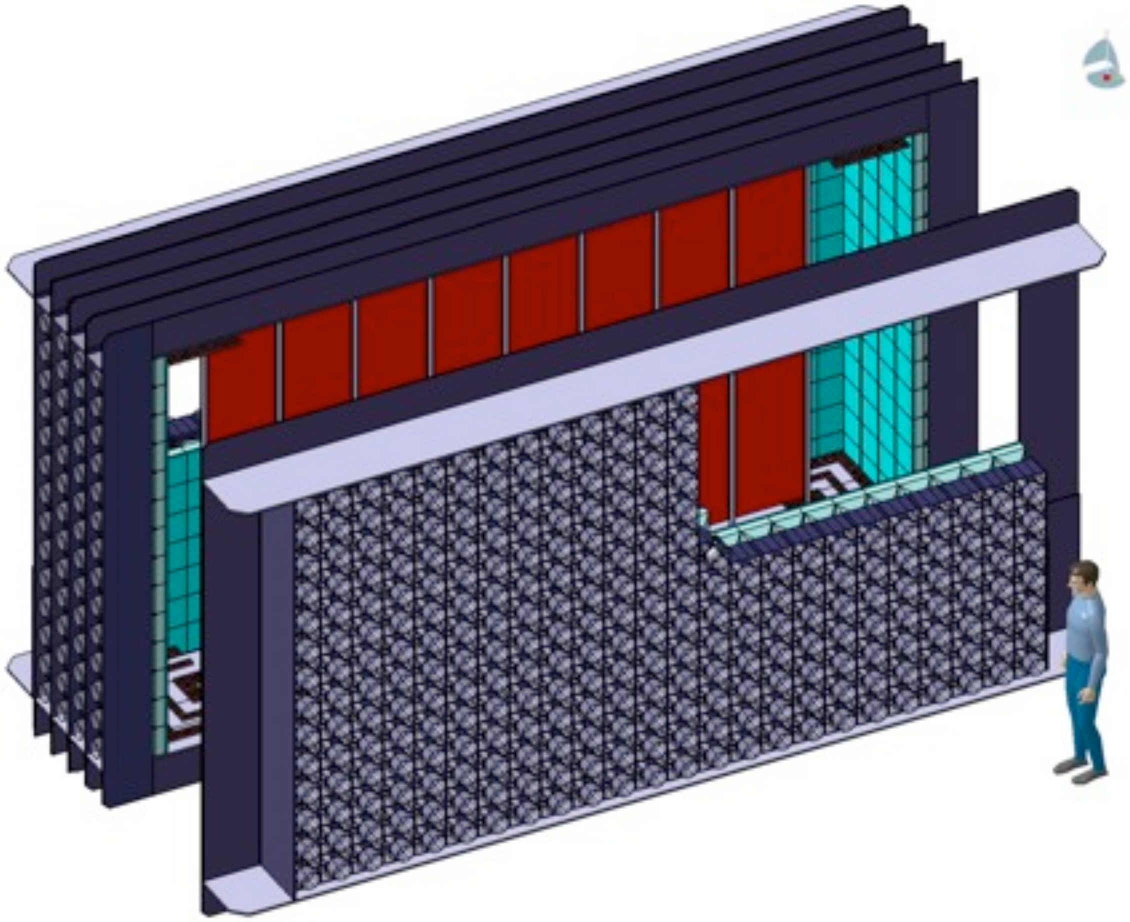
…

Tracking + Calorimetry, source ≠ detector

NEMO-III and SuperNEMO (~90 people)

NEMO-III

SuperNEMO

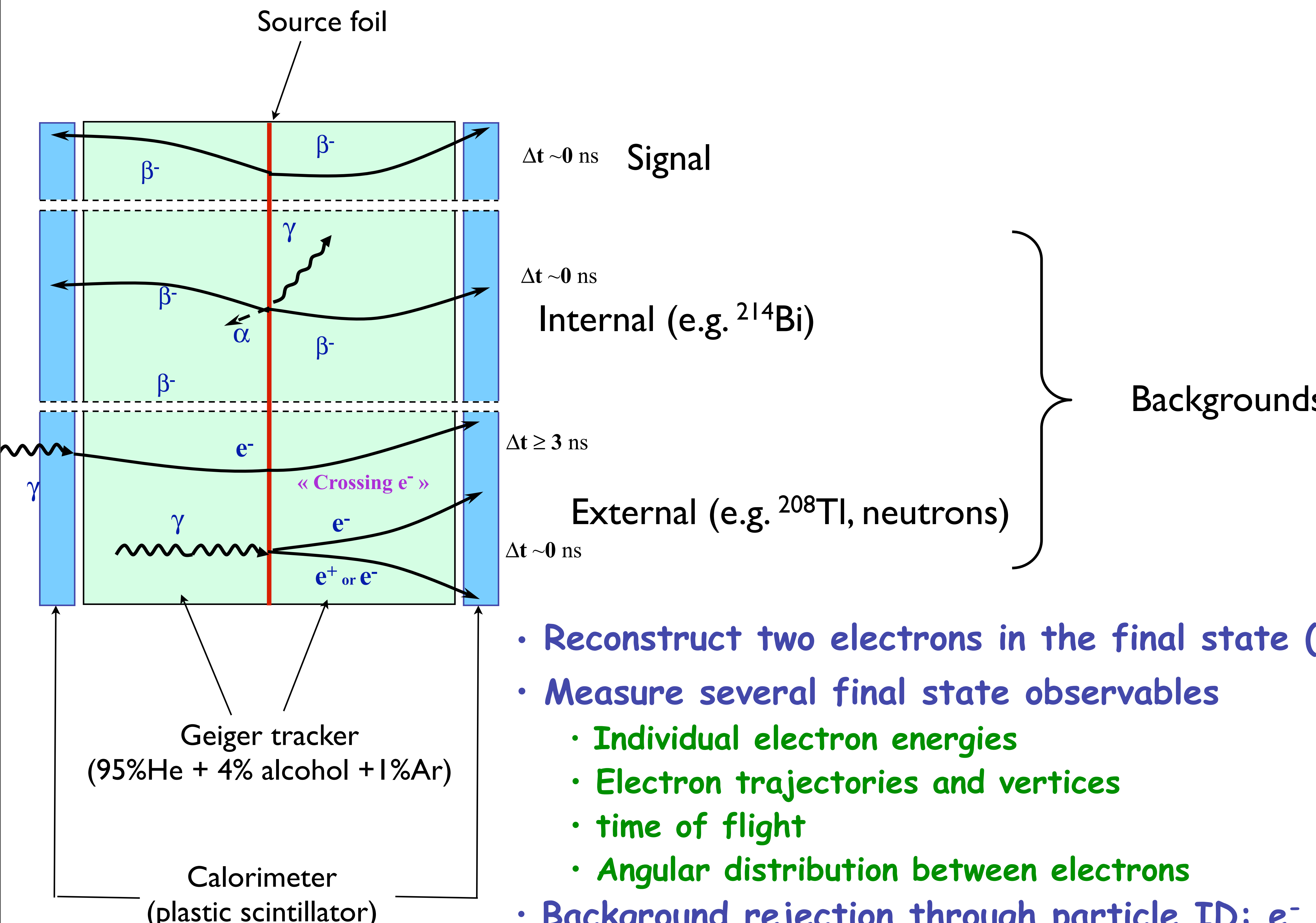


Major contribution from UK and France:
~80% shared equally

UK collaboration: UCL-HEP, UCL-MSSL, University of Manchester, IC.

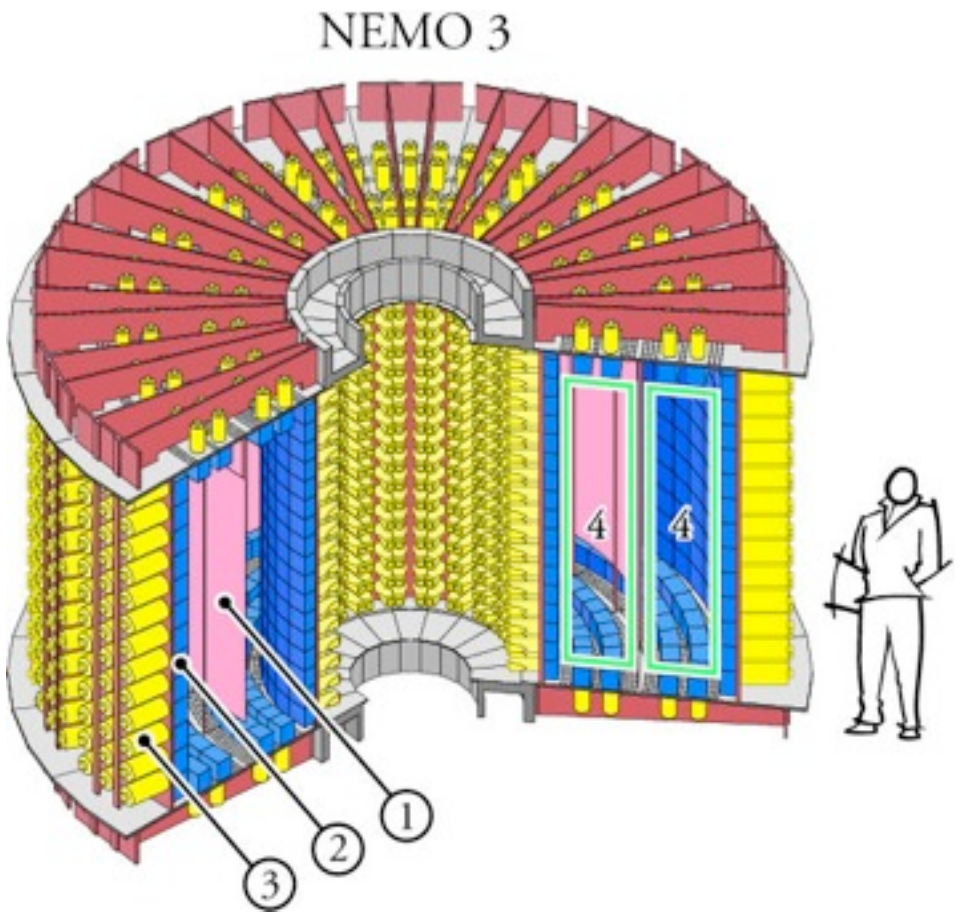
NEMO-III and SuperNEMO

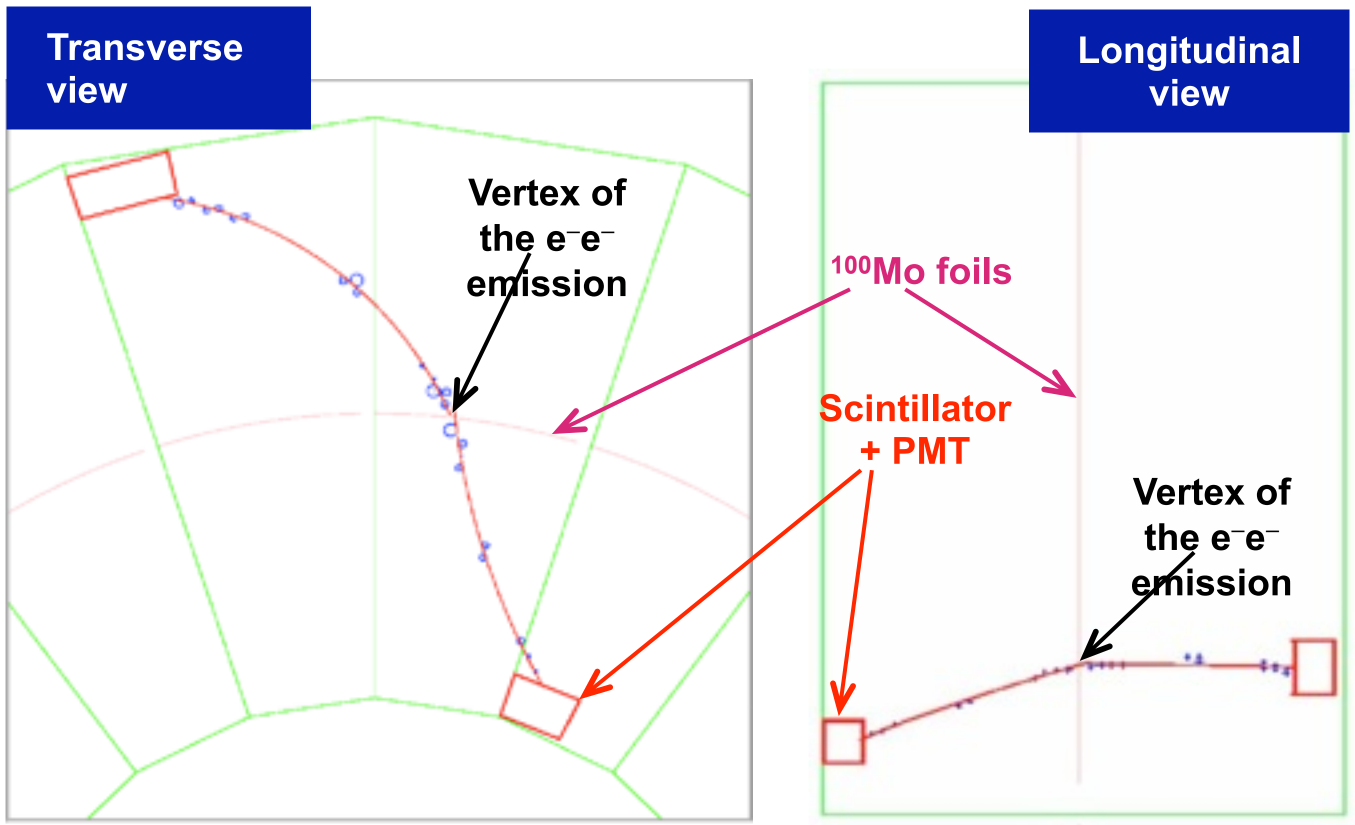
Detection principle: reconstruct topological signature



- Reconstruct two electrons in the final state ($E_1+E_2 = Q_{\beta\beta}$)
- Measure several final state observables
  - Individual electron energies
  - Electron trajectories and vertices
  - time of flight
  - Angular distribution between electrons
- Background rejection through particle ID: e^-, e^+, α, γ

NEMO3/SuperNEMO approach. Calorimetry + Tracking





Sources separated from detector ⇒
⇒ can measure different isotopes
"Smoking gun" signature: ⇒
⇒ two e⁻ with **common** vertex

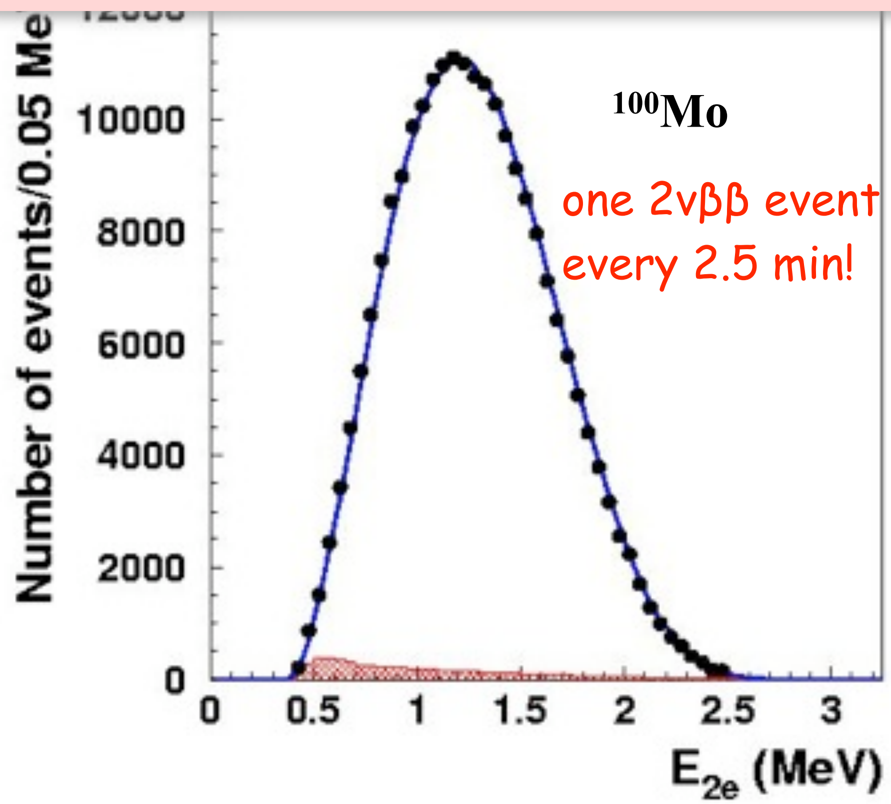
} unique to NEMO

- Search for **any** lepton violating process including with continuum spectrum (e.g. Majoron)
- Attempt to disentangle the underlying physics mechanism through electron's angular distribution and individual energy analysis

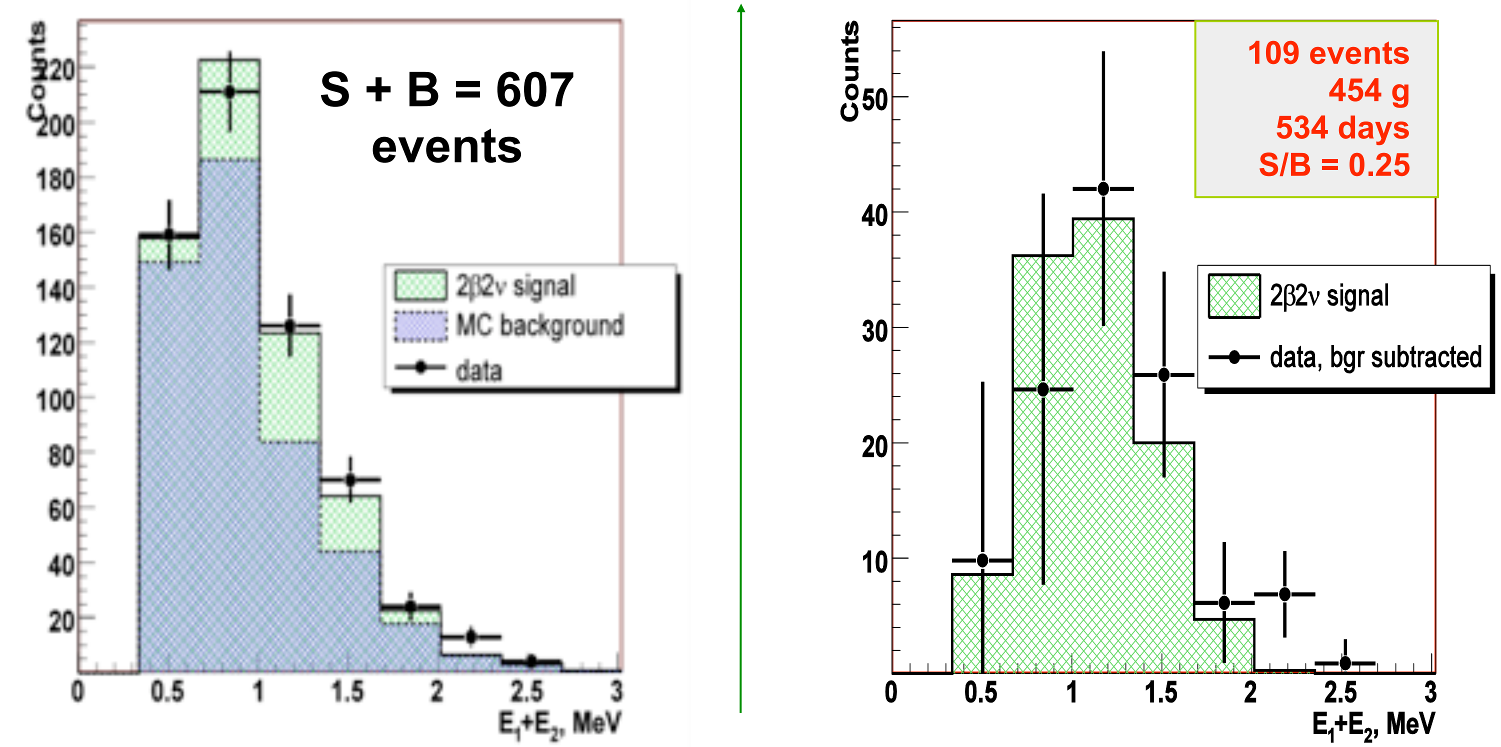
NEMO3 Results

Unprecedented accuracy of 2νββ measurements

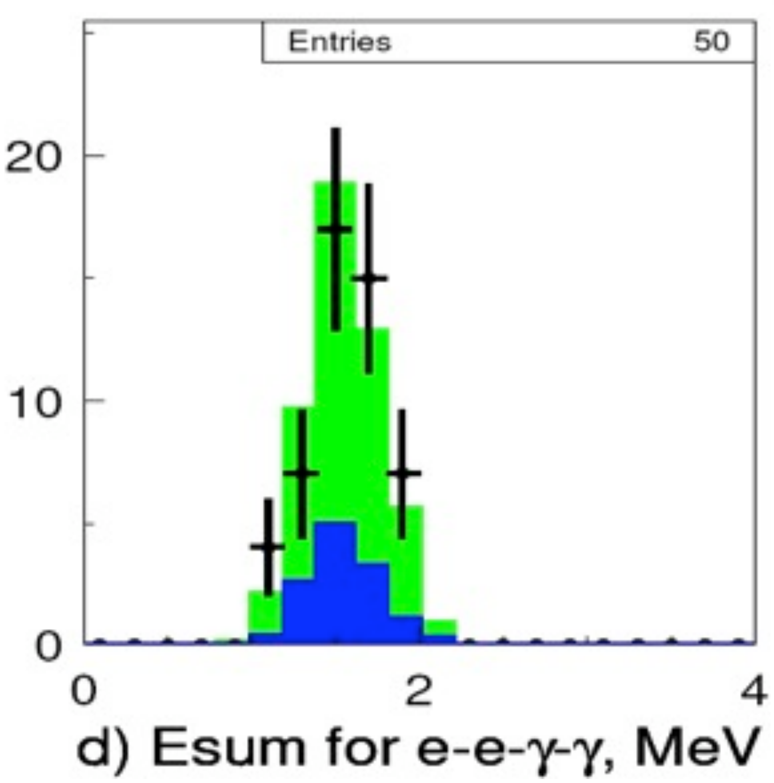
$T_{1/2}(\beta\beta 2\nu) = 7.11 \pm 0.02\ (stat) \pm 0.54\ (syst) \times 10^{18}$ yrs



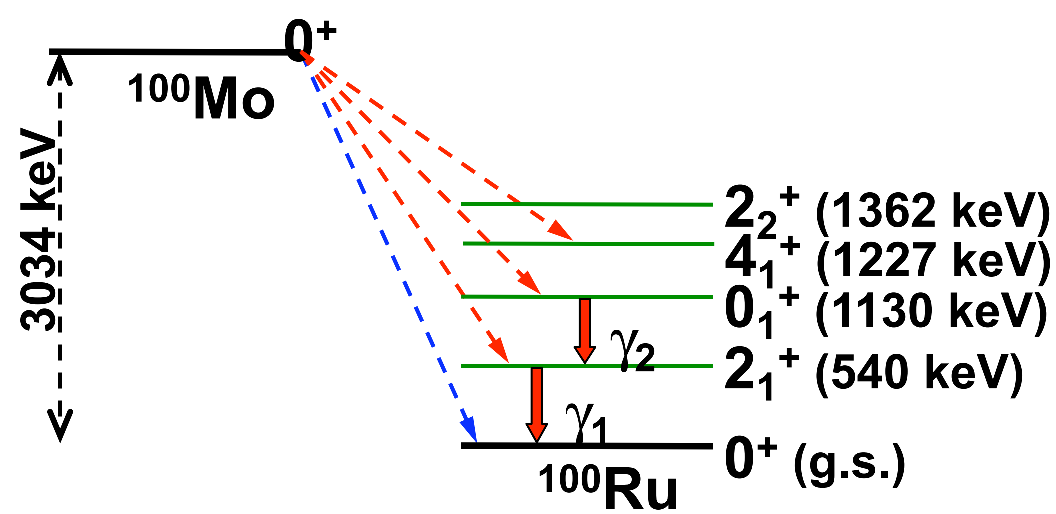
$T_{1/2} =$ [6.9 ± 0.9 (stat) ± 0.8 (syst)] $\times 10^{20}$ y (preliminary)



Long awaited ^{130}Te result



decay to excited states



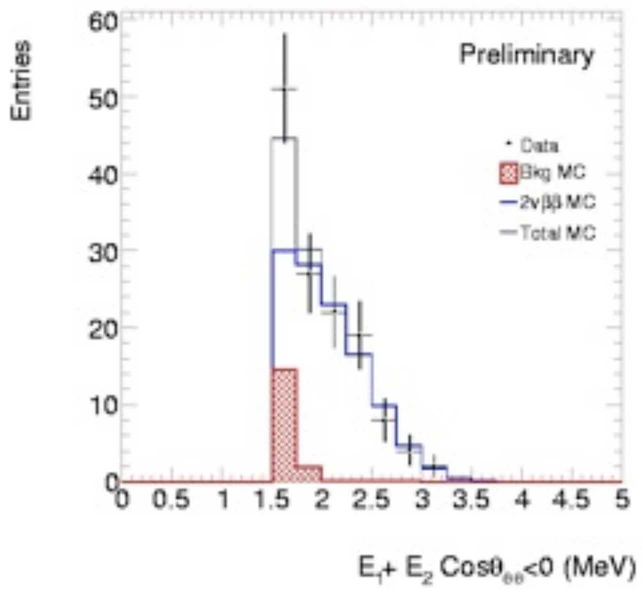
Also 2νββ measurements for ^{82}Se, ^{116}Cd, ^{150}Nd, ^{48}Ca, ^{96}Zr

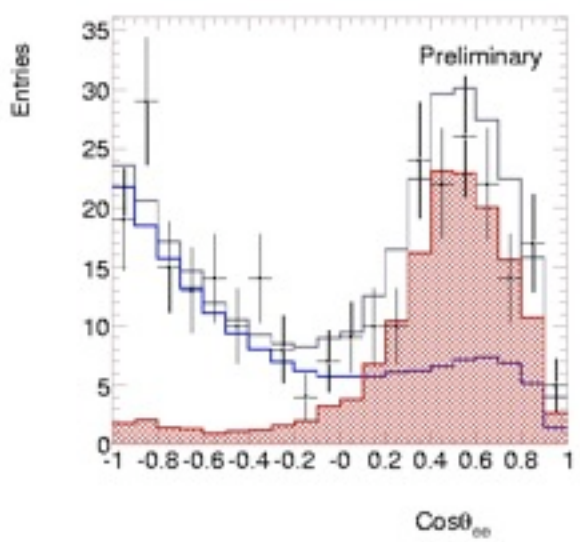
NEMO3 Results

Large UK analysis effort

^{48}Ca

$T_{1/2}(2\nu\beta\beta) = [4.4\ ^{+0.5}_{-0.4}\ (stat) \pm 0.4\ (syst)] \times 10^{19}$ y

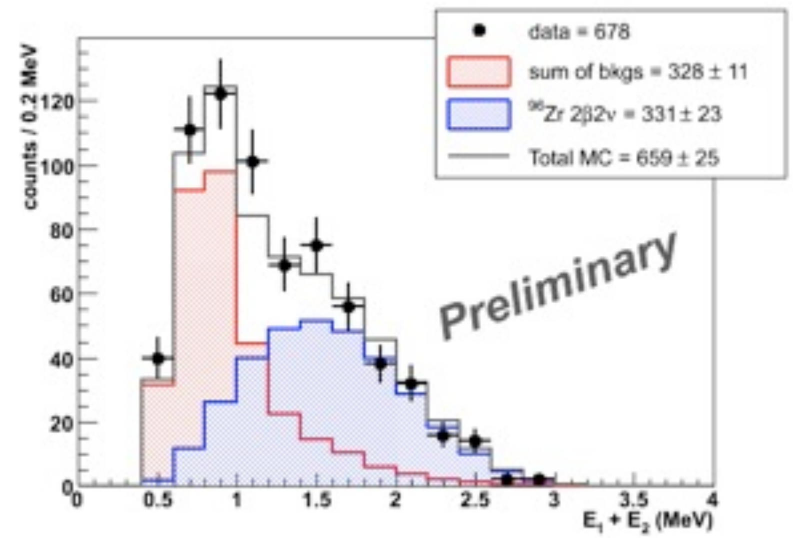


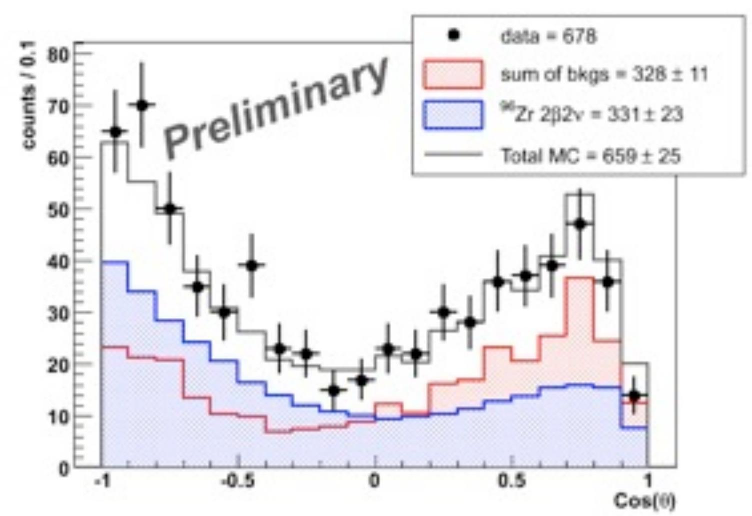


paper in preparation

^{96}Zr

$T_{1/2}(2\nu\beta\beta) = [2.3 \pm 0.2(stat) \pm 0.2(syst)] \times 10^{19}$ y

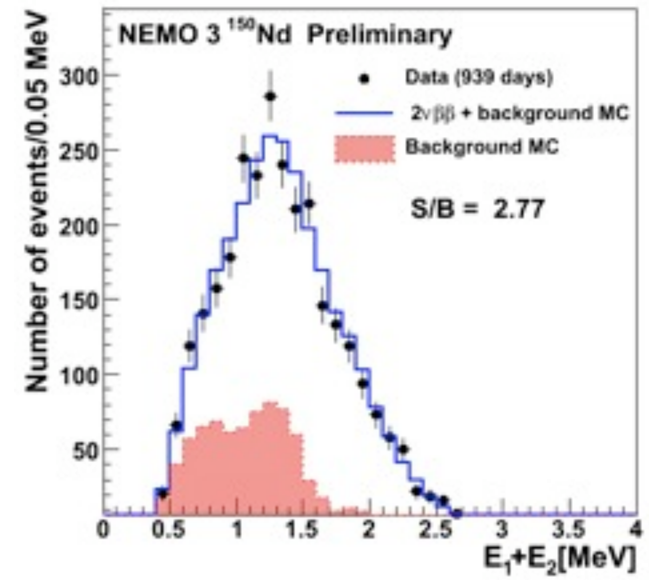


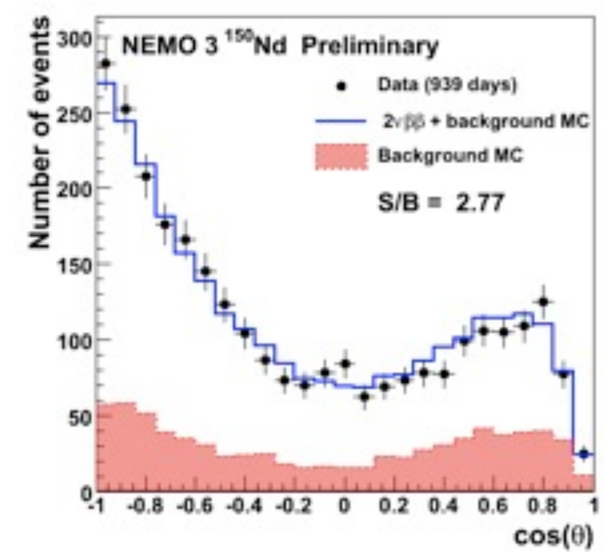


paper about to be submitted to Nucl. Phys. A

^{150}Nd

$T_{1/2}(2\nu\beta\beta) = [9.20\ ^{+0.25}_{-0.22}\ (stat) \pm 0.62\ (syst)] \times 10^{18}$ y





submitted to PRL

NEMO3 Results

Large UK analysis effort

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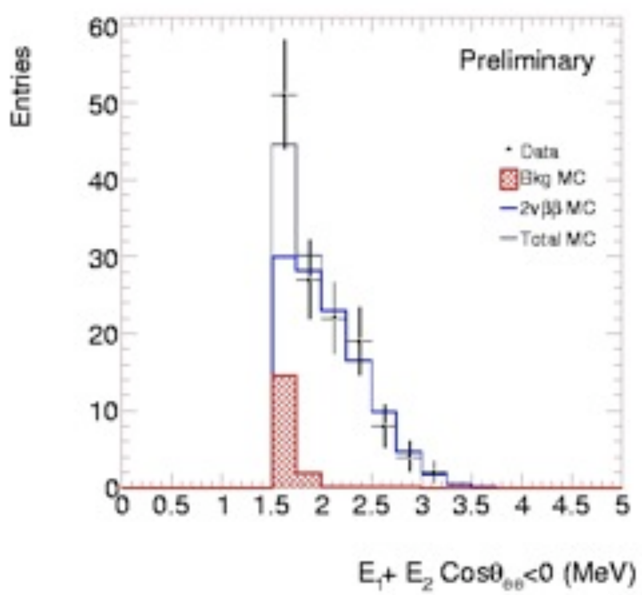

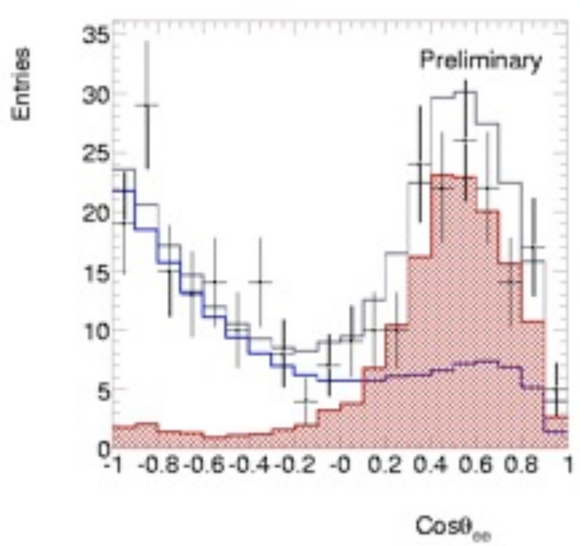

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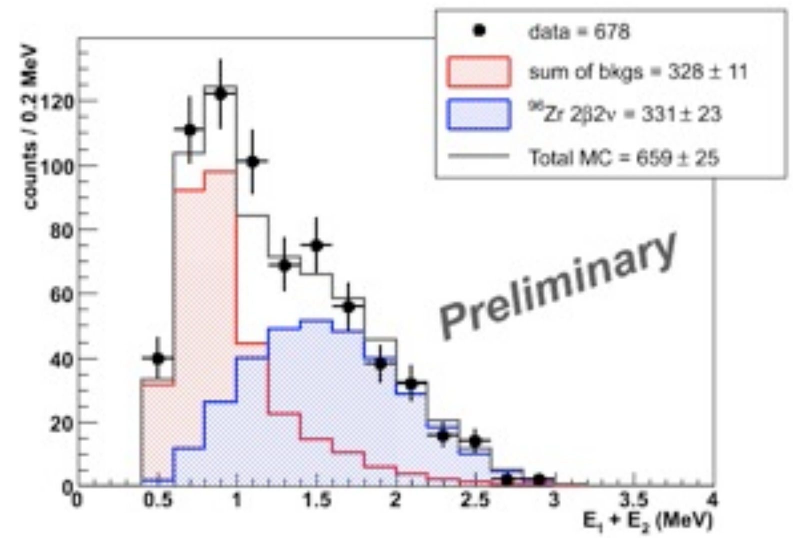

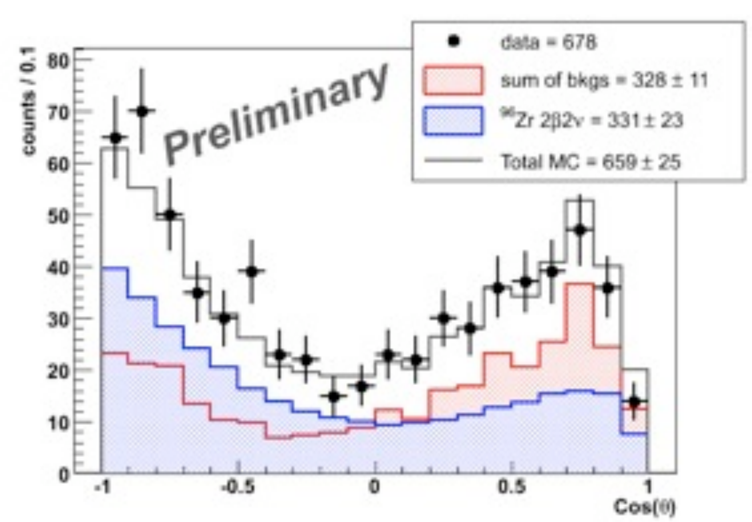

Talk by M. Kauer later today

^{150}Nd

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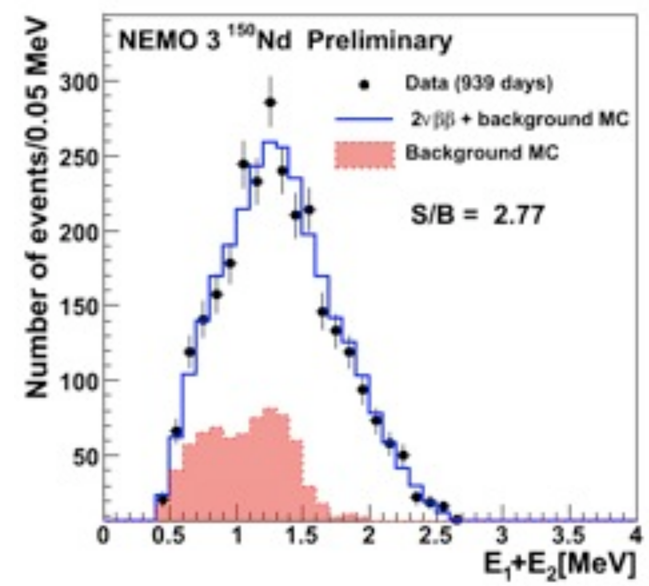
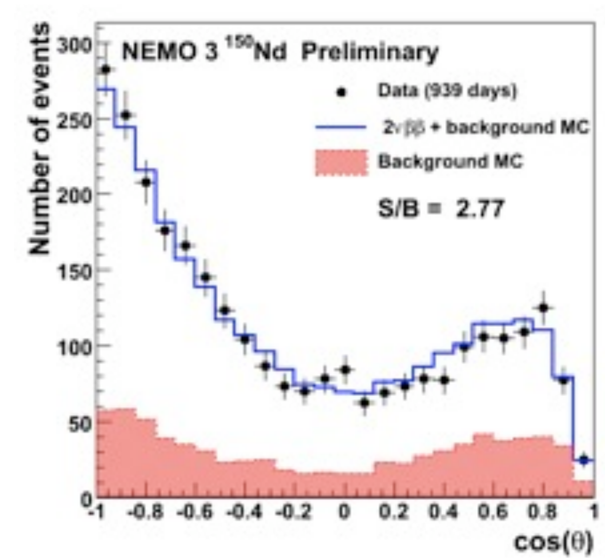
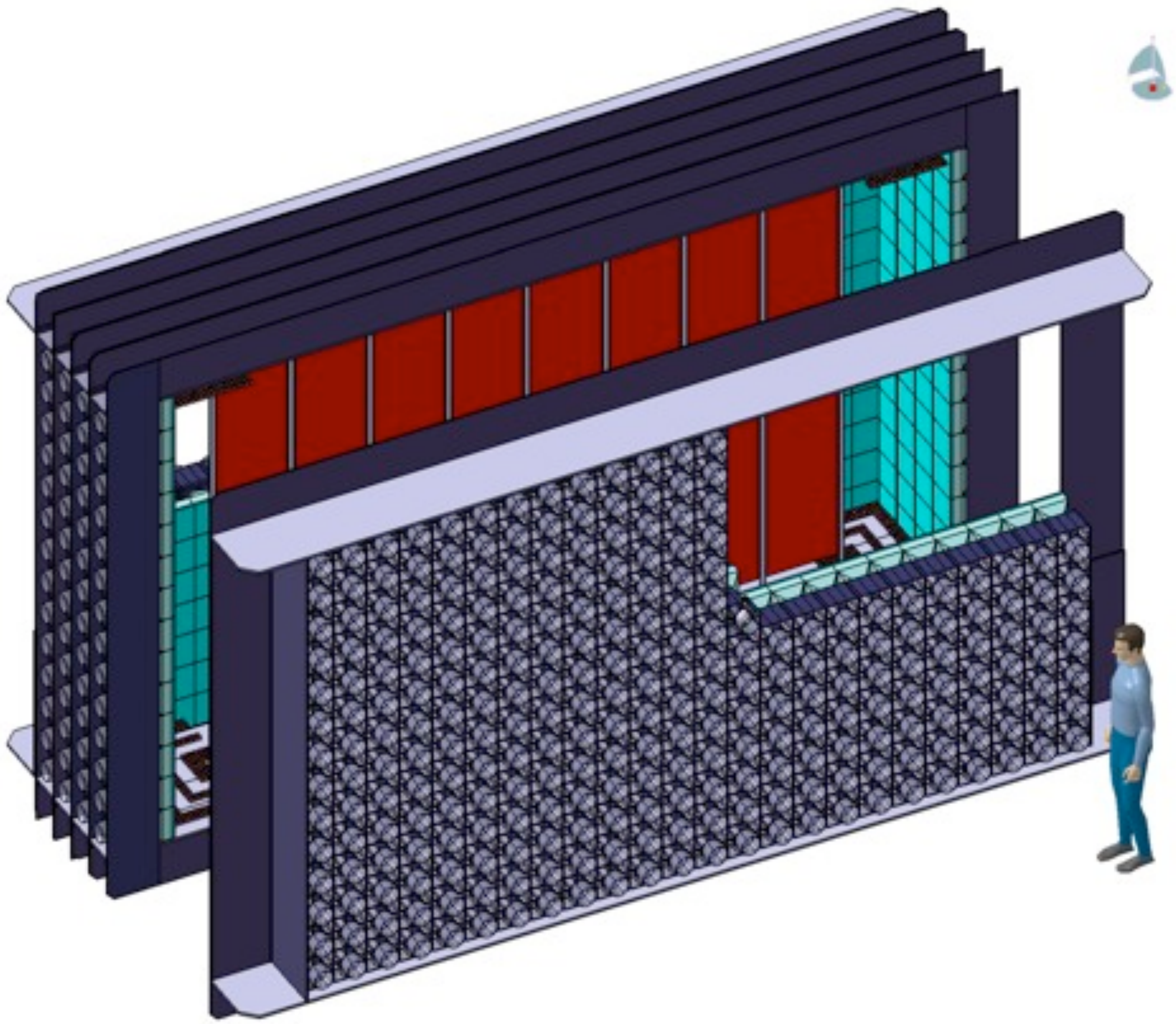




submitted to PRL

SuperNEMO Detector



Planar and modular design:
~ 100 kg of enriched isotopes
(20 modules x 5 kg)

1 super-module:

Source (40 mg/cm^2)   4 x 3 m^2

Tracking : drift chamber

~2000 cells in Geiger mode

Calorimeter: scintillators + PMTs

~ 700 PMTs if scint. blocks

~ 250 PMTs if scint. bars

Modules surrounded by water passive shielding in an underground lab (new-LSM)

Funded Design Study 2006-2009

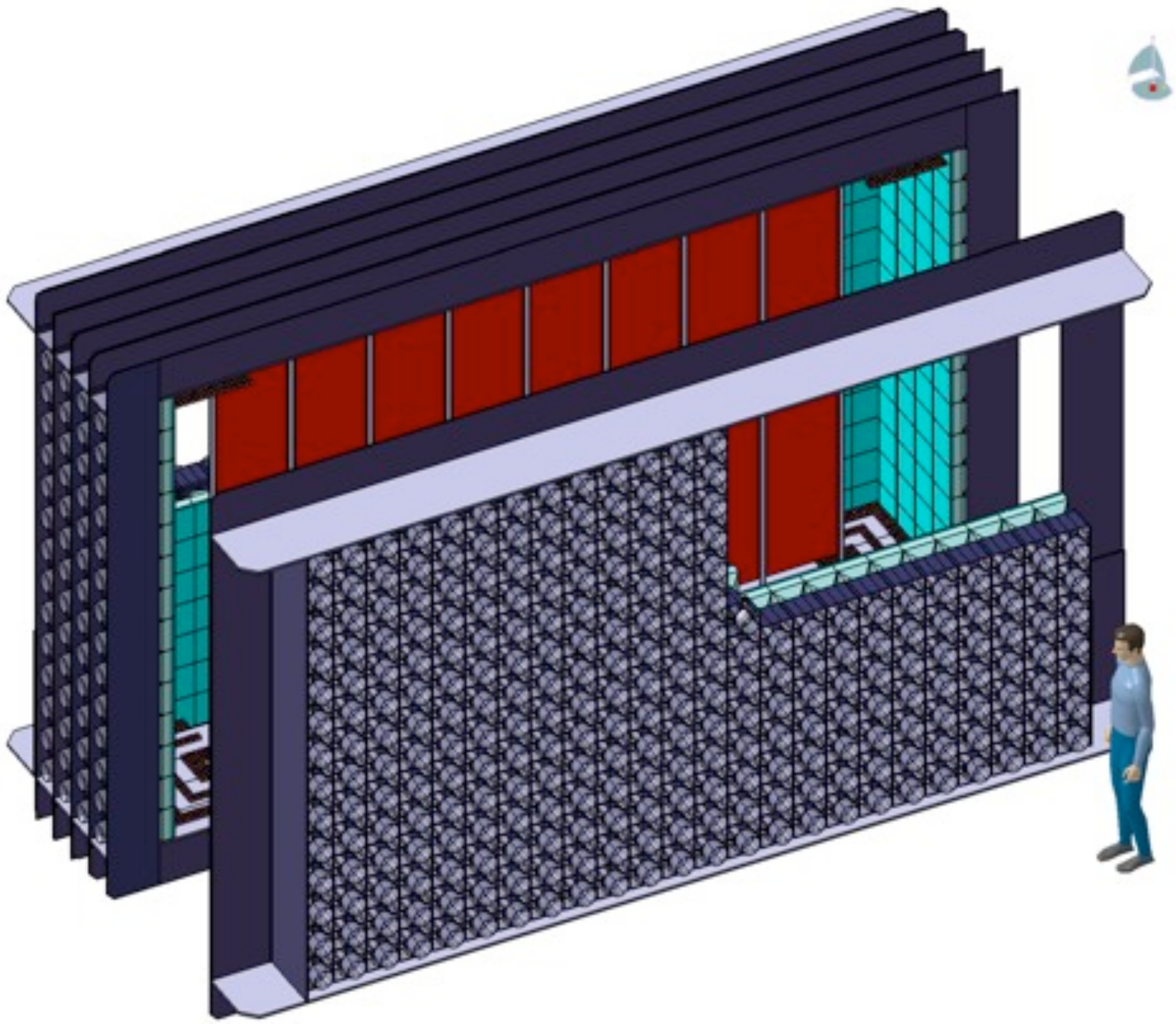
Calorimeter $\Delta E/E = 7\%/\sqrt{E}$ (MeV) $\Rightarrow$ 4% at $Q_{\beta\beta}$

Source radiopurity: ^{208}Tl < 2 μBq/kg,
^{214}Bi < 10 μBq/kg (if ^{82}Se)

Tracker optimization – automated wiring

All deliverables on track to be completed by mid-2009

SuperNEMO Detector



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Tracker optimization – automated wiring

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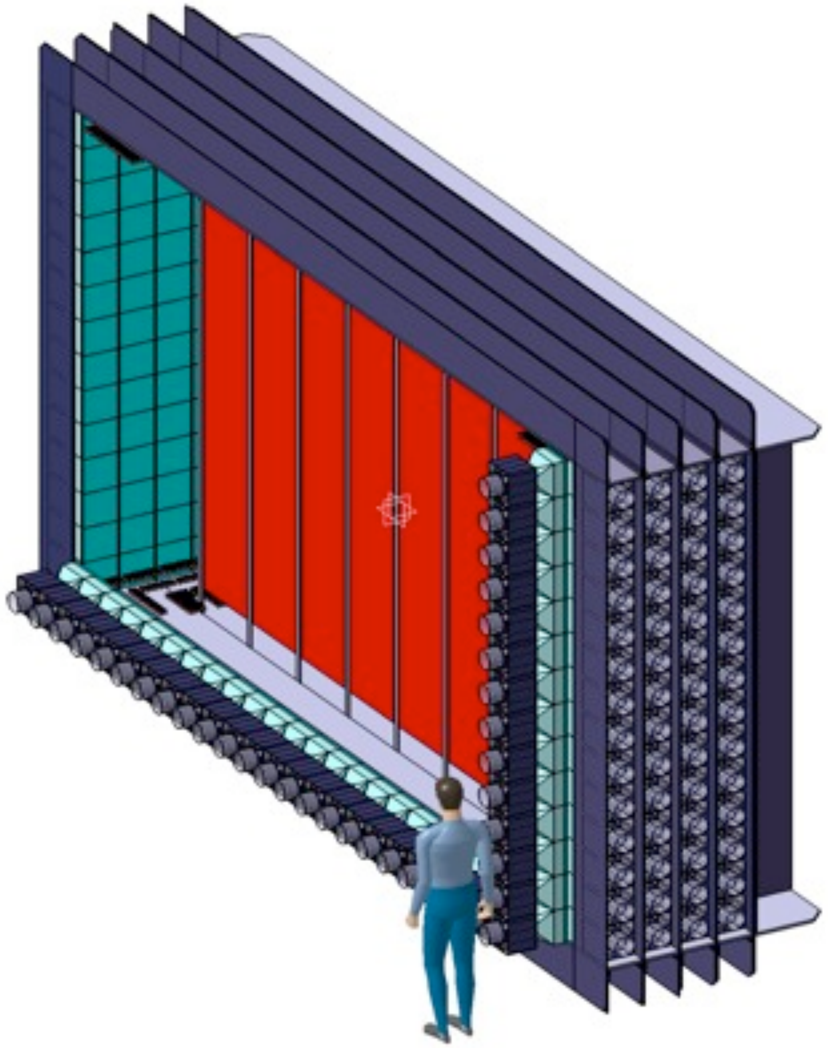
Next step: To build 1st SuperNEMO module - Demonstrator

Goals

- Demonstrate feasibility of large scale mass production
- To measure backgrounds especially from radon emanation
  - Only possible with a realistic super-module
- To finalise detector design
- To produce a competitive physics measurement



Simple counting experiment: 0.3 bkg. events in R.O.I with 7kg of ^{82}Se in 1.5 yr

$$T_{1/2}^{0\nu}(90\%CL) > \frac{M}{A} N_A \frac{\varepsilon \ln 2}{n_{90\%}} t = \frac{7kg\cdot 10^3}{82g} 6\cdot 10^{23} \frac{0.3\cdot 0.69}{2.4} 1.5yr = 6.6\cdot 10^{24}\,yr$$

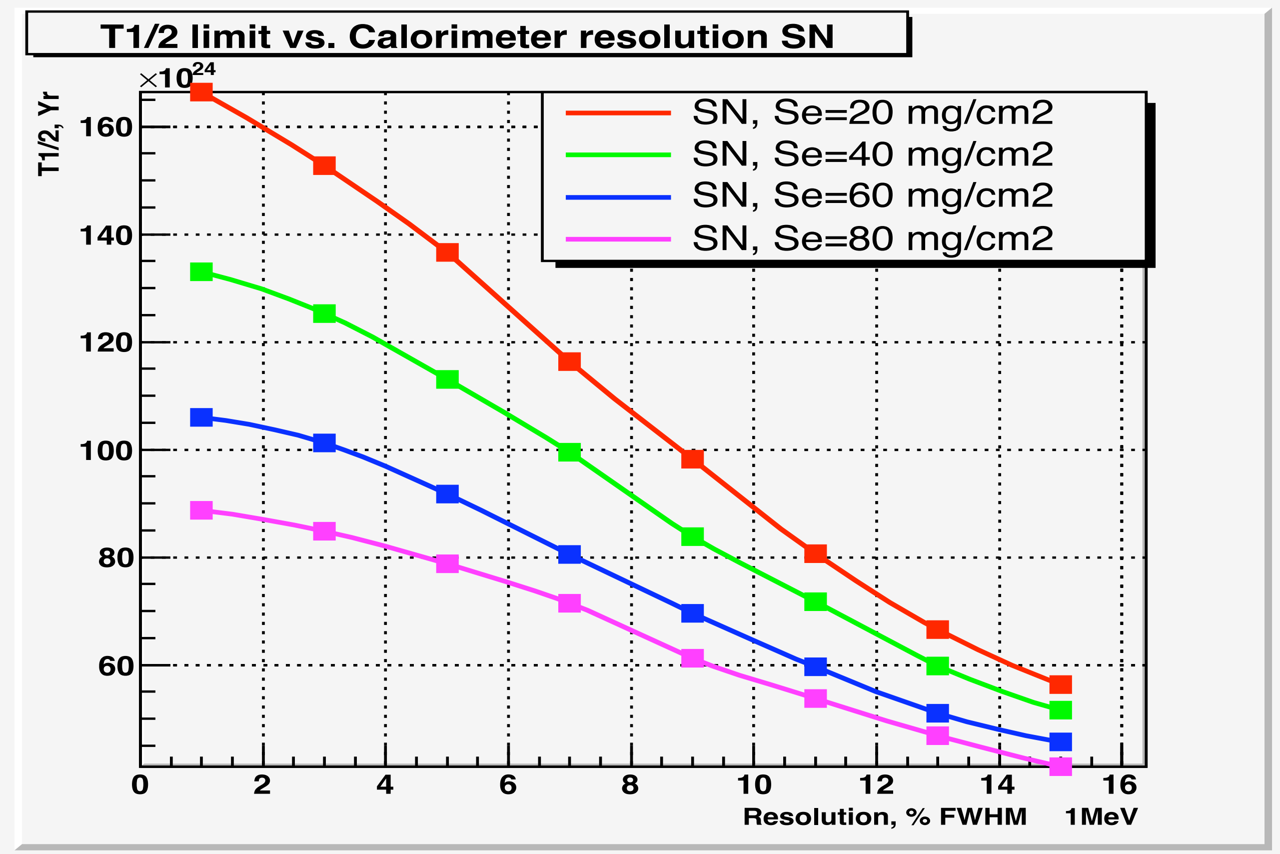
Assuming equal NMEs and known differences in phase space values, it is equivalent to **$3\cdot 10^{25}$ yr** for ^{76}Ge (GERDA-PhaseI)

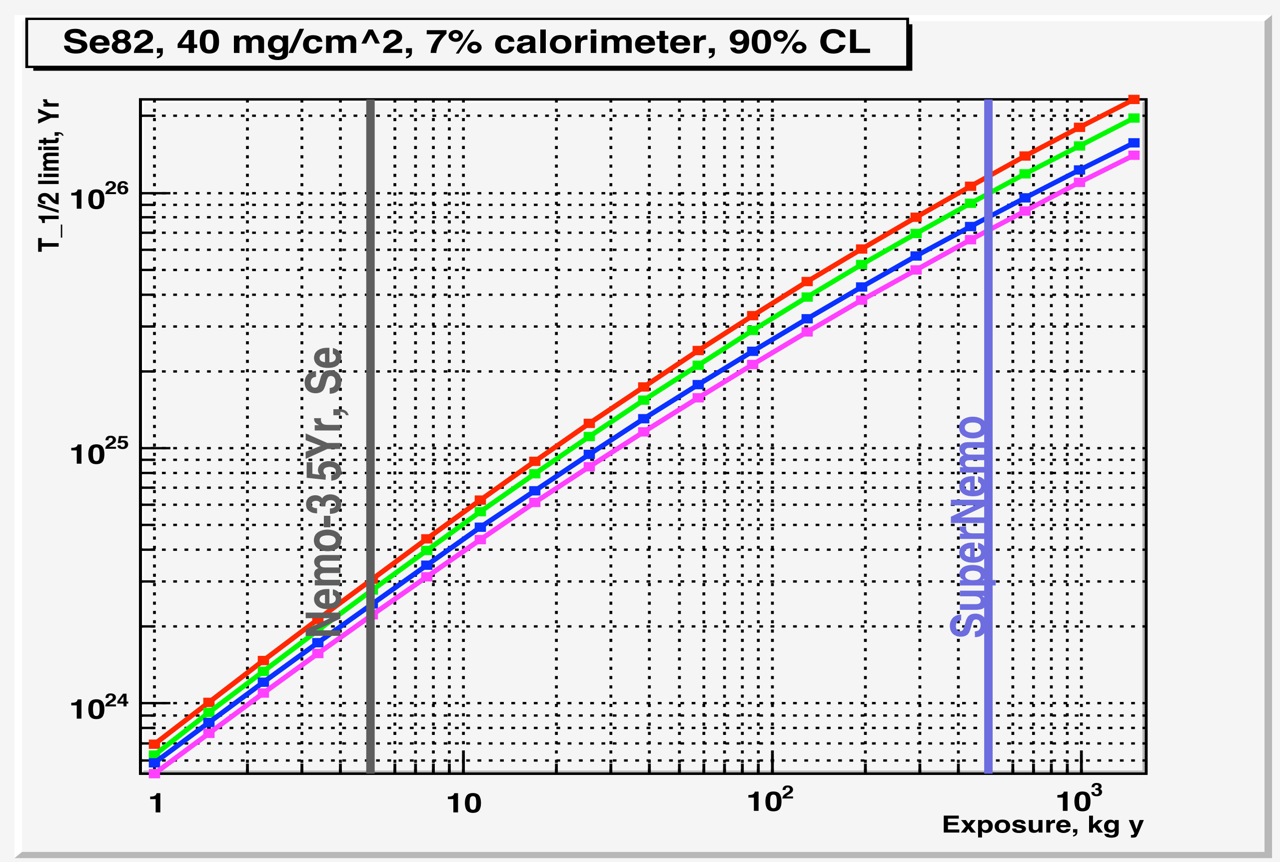
or ~5 expected "golden events" if Klapdor is right

Proposal submitted to PPRP

SuperNEMO sensitivity:

500 kg x yr, target bkg levels (2 and 10 µBq/kg of ^{208}Tl and ^{214}Bi)





^{82}Se:

$T_{1/2}(0\nu)$ = 10^{26} yr

$\langle m_\nu \rangle \leq 0.05 - 0.11$ eV

(range due to Nuclear Matrix Element uncertainties)

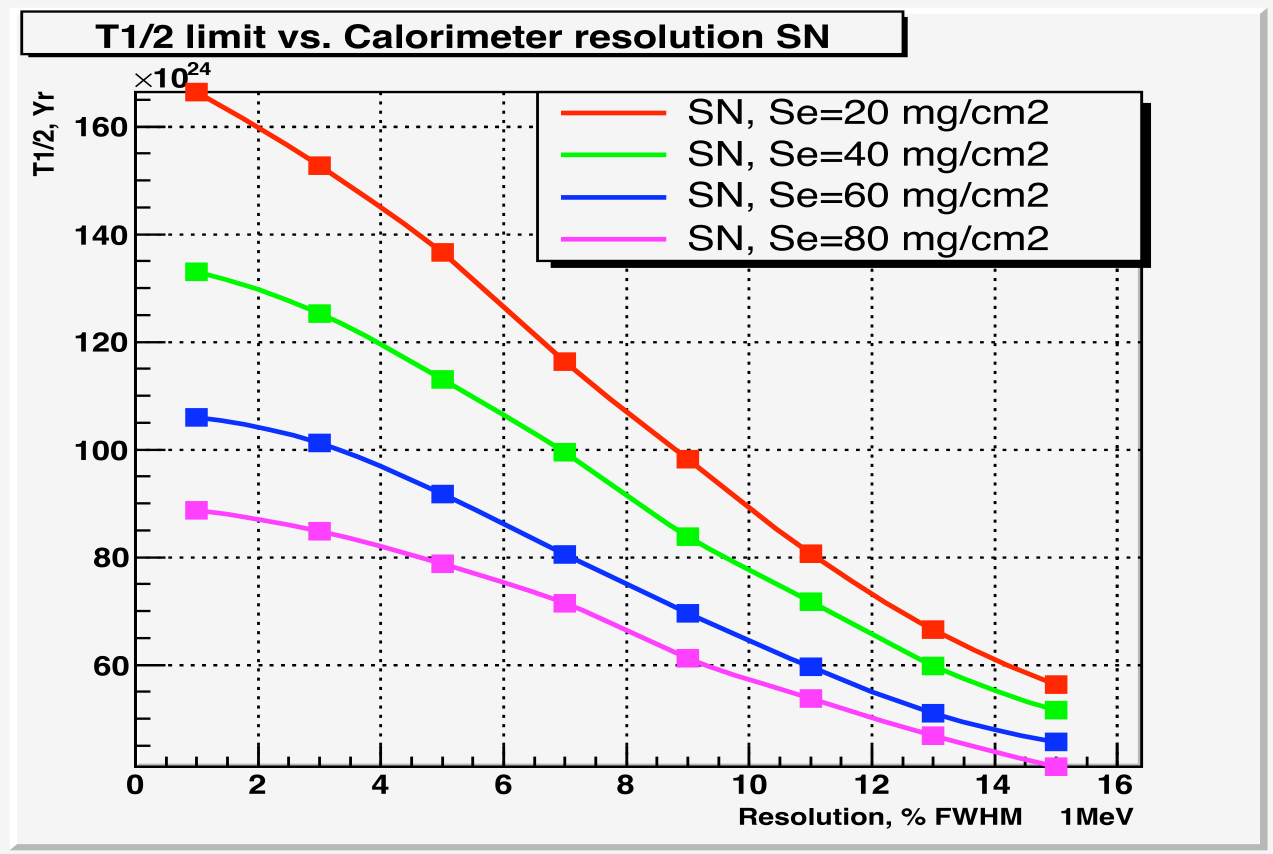
^{150}Nd:

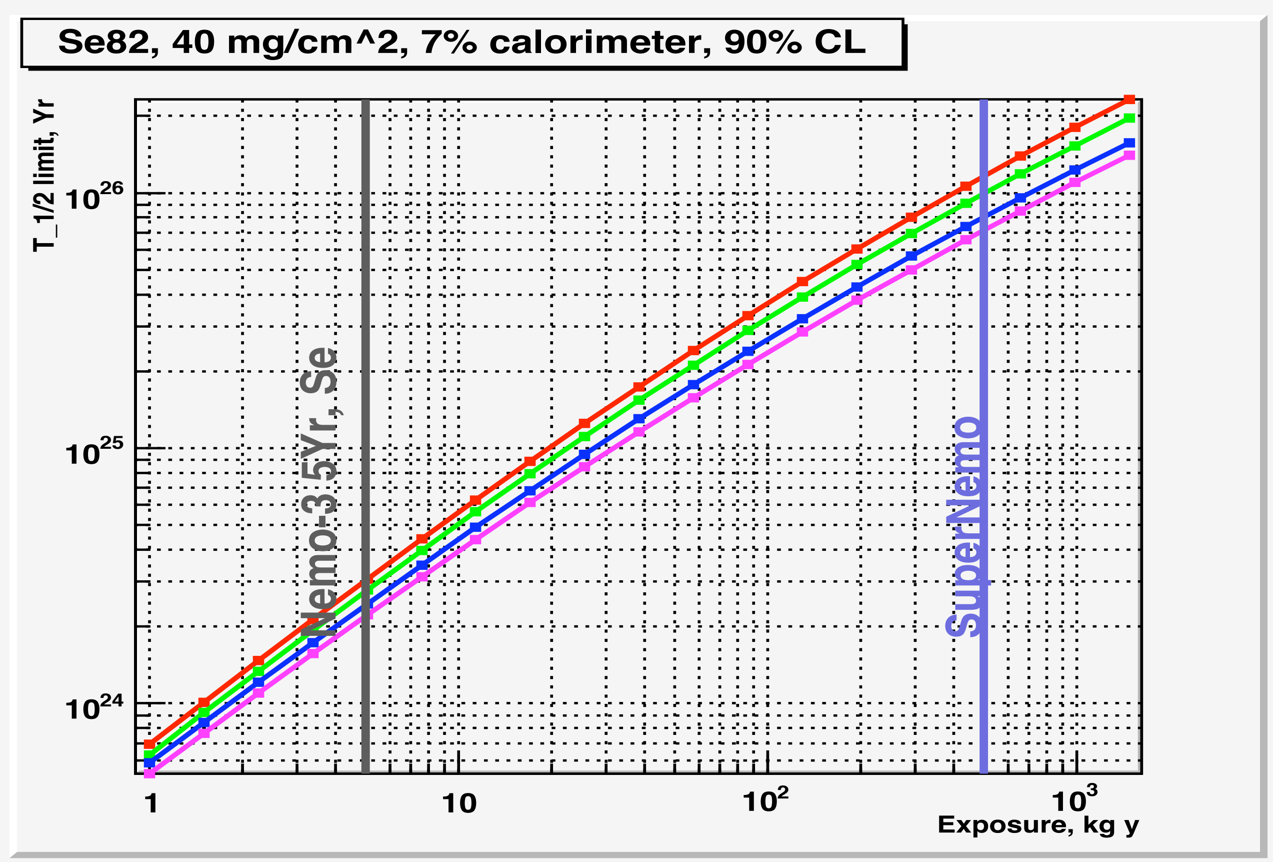
$T_{1/2}(0\nu)$ =5 10^{25} yr $\langle m_\nu \rangle \leq 0.045$ eV

(but deformation **not** taken into account)

SuperNEMO sensitivity:

500 kg x yr, target bkg levels (2 and 10 μBq/kg of ^{208}Tl and ^{214}Bi)





^{82}**Se**:

$T_{1/2}(0\nu) = 10^{26}$ yr

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(range due to Nuclear Matrix Element uncertainties)

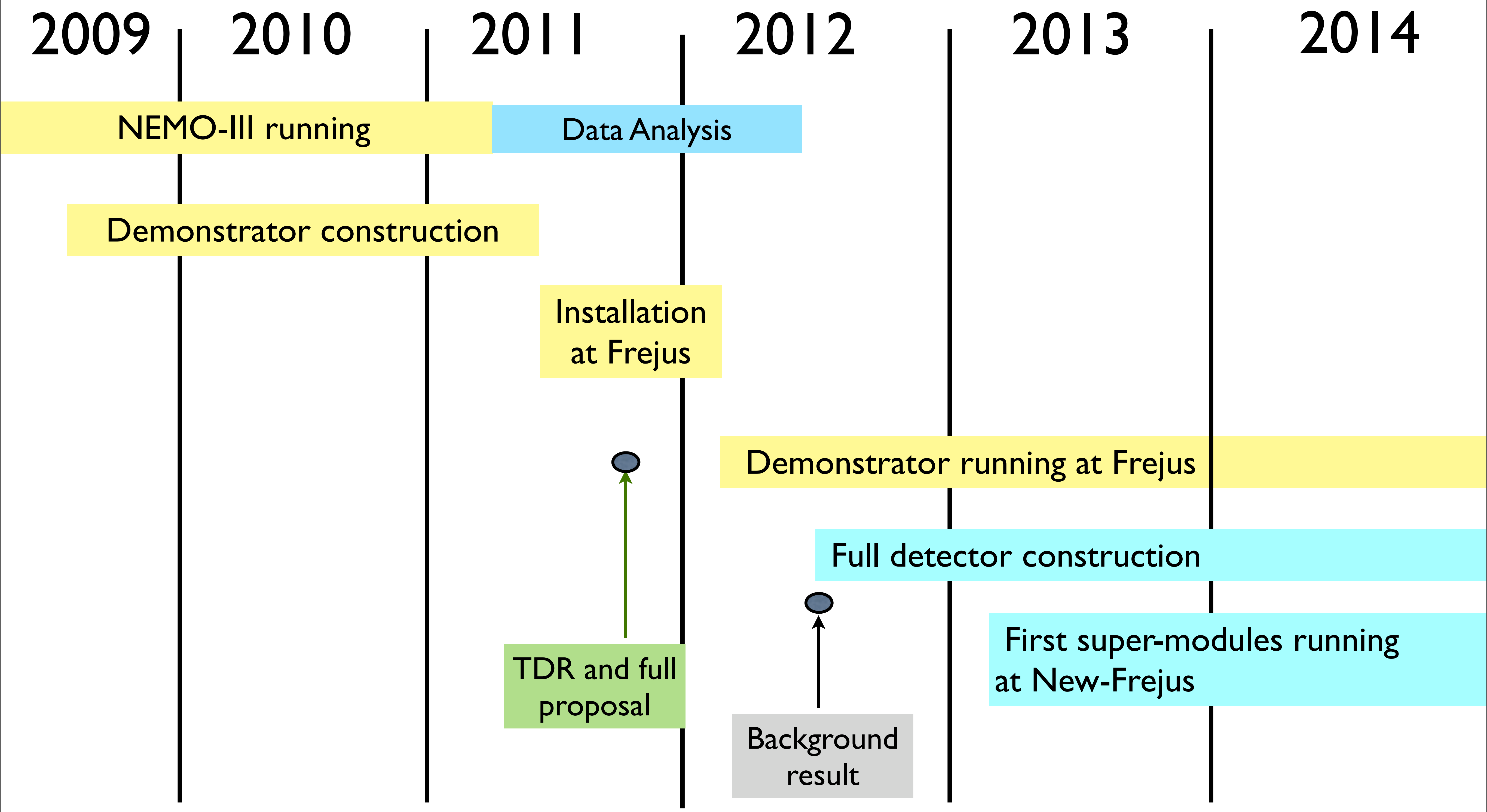
^{150}**Nd**:

$T_{1/2}(0\nu) = 5\ 10^{25}$ yr $\langle m_\nu \rangle \leq 0.045$ eV

(but deformation **not** taken into account)

Talk by C. Jackson later today

SuperNEMO schedule overview



Target sensitivity of 50-100 meV by 2018

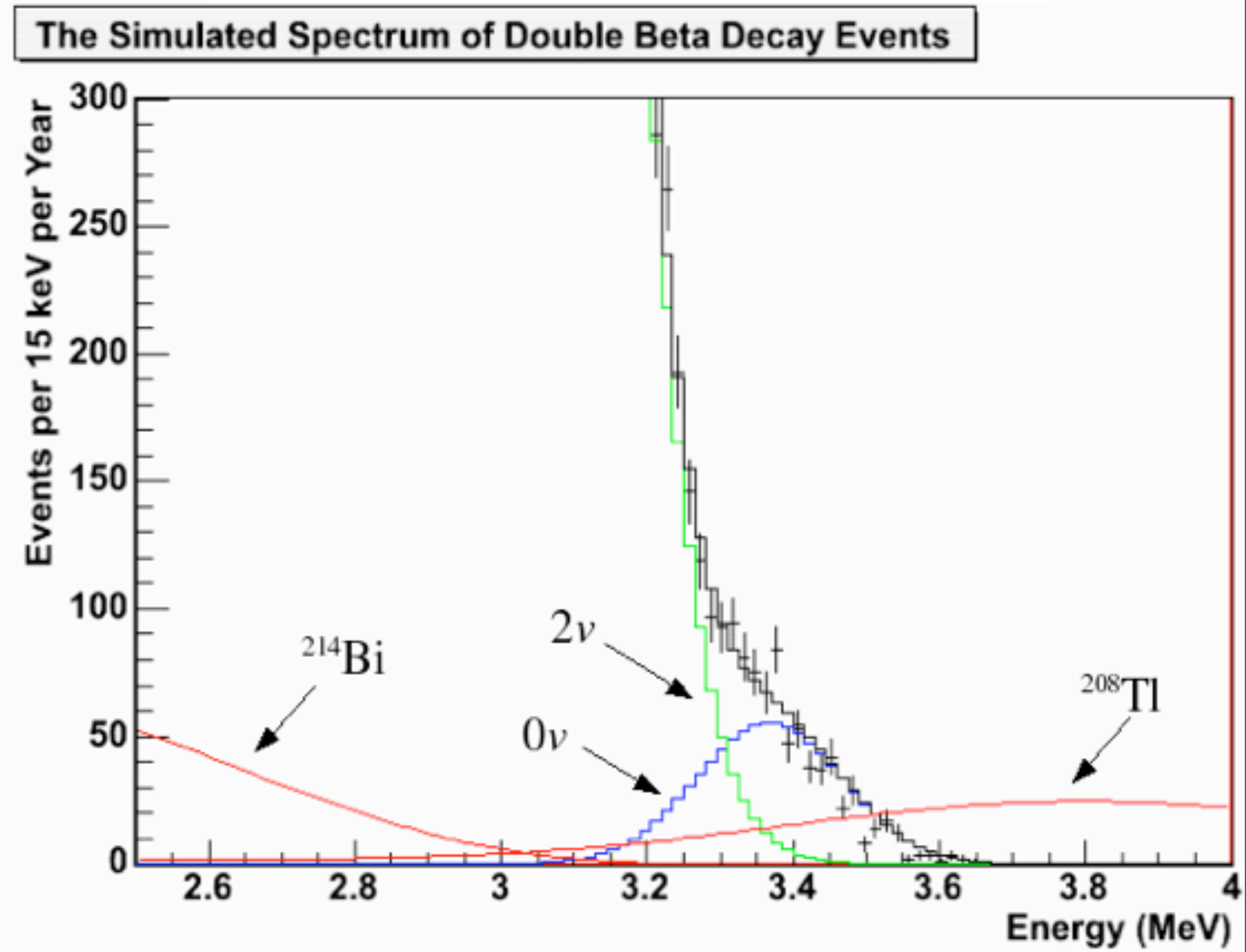
Source = detector approaches

- Scintillator calorimeters
- Semiconductors (HPGe, CdZnTe)
- Bolometers
- Liquid/Gas Xe TPC

SNO+

- ^{150}Nd double beta decays with an endpoint of 3.37 MeV (above most backgrounds).
- Poor energy resolution compensated by
  - little material near fiducial volume
  - meters of self-shielding
  - source in–source out capability

simulation of one year of data
testing$<m_\nu>$ = 150 meV (500kg of ^{150}Nd)

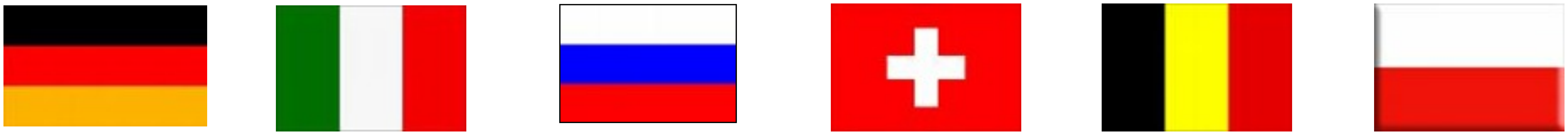


- Nd-loaded liquid scintillator
- 0.1% ^{nat}Nd in 1000 t of liquid scintillator
  - 56 kg of ^{150}Nd
- 0.1 eV* sensitivity with ^{nat}Nd
  - Quick "turnaround"
  - ^{nat}Nd radio-purity is one of key questions (10^{-14} g/g in U and Th required)
- Possibility to enrich Nd with AVLIS (joint R&D with SuperNEMO)
- If enriched, 50 meV* sensitivity possible

UK participation proposal submitted to PPRP

*^{150}Nd deformation not taken into account in NME calculation

GERDA. ^{76}Ge (HP Ge detector)



^{76}Ge – best way to check KKDC claim (free from NME uncertainties).

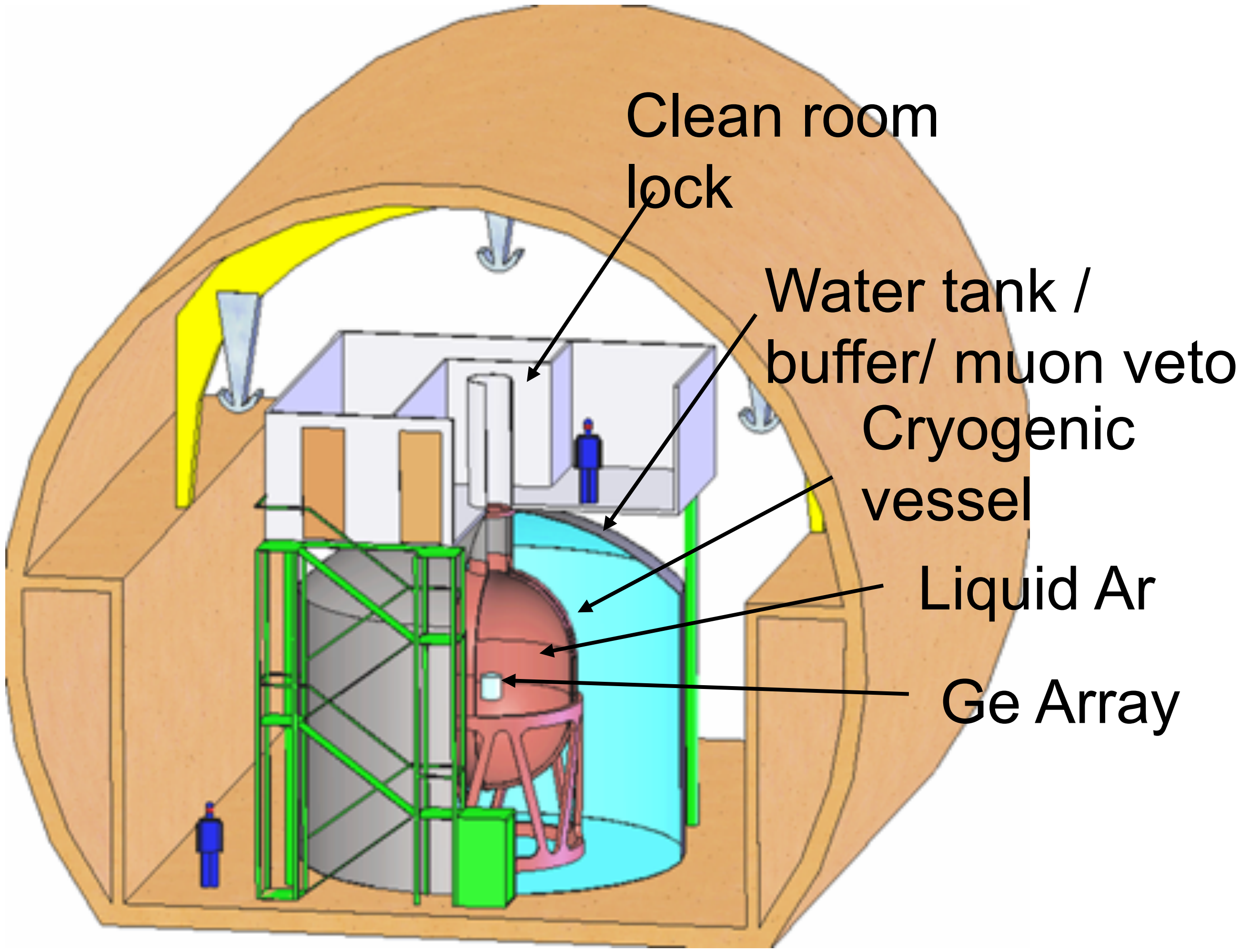
Naked enriched (86%) Ge-detectors
in LAr
Phased Approach.

Phase I:
Existing detectors (HM+IGEX)
17.9 kg enriched diodes
Bkg free proble of KKDC:
10^{-2} cts/kg keV yr

Phase II
Add new diodes (total: 40kg)
Bkg < 10^{-3} cts/kg keV yr

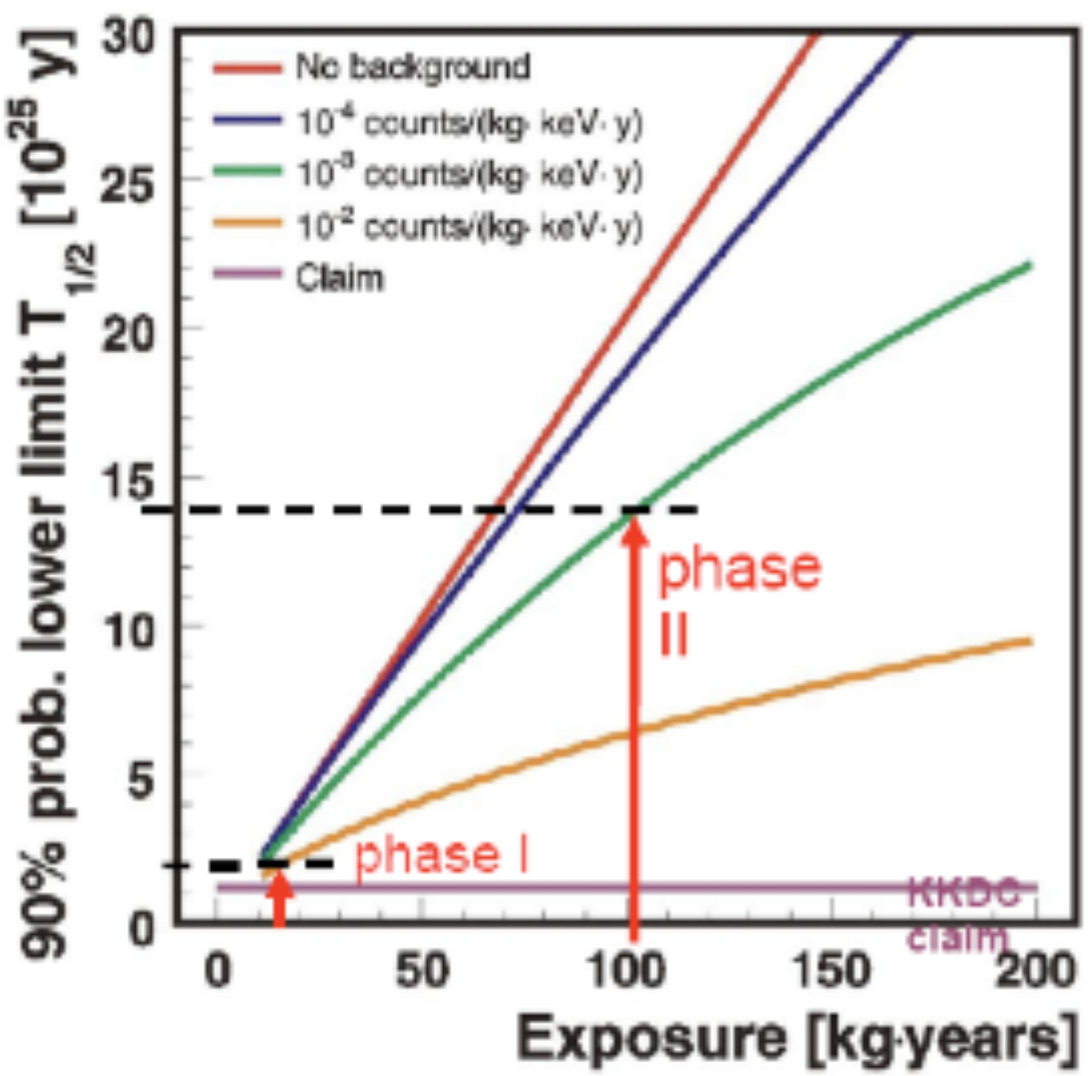
Both phases funded. Under construction

Location: Gran Sasso

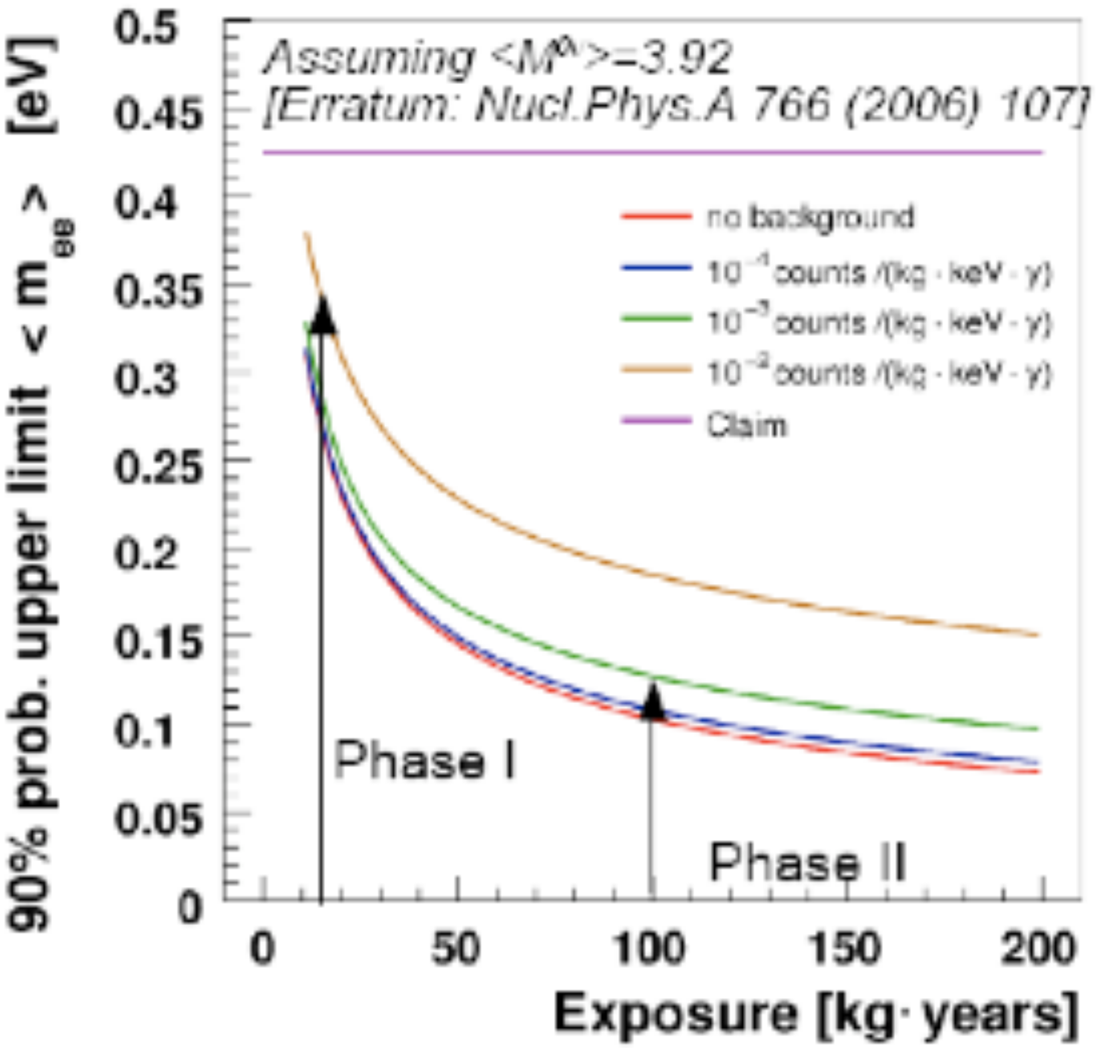


Next step: GERDA + Majorana

GERDA. ^{76}Ge



assumed energy resolution: ΔE = 4 keV

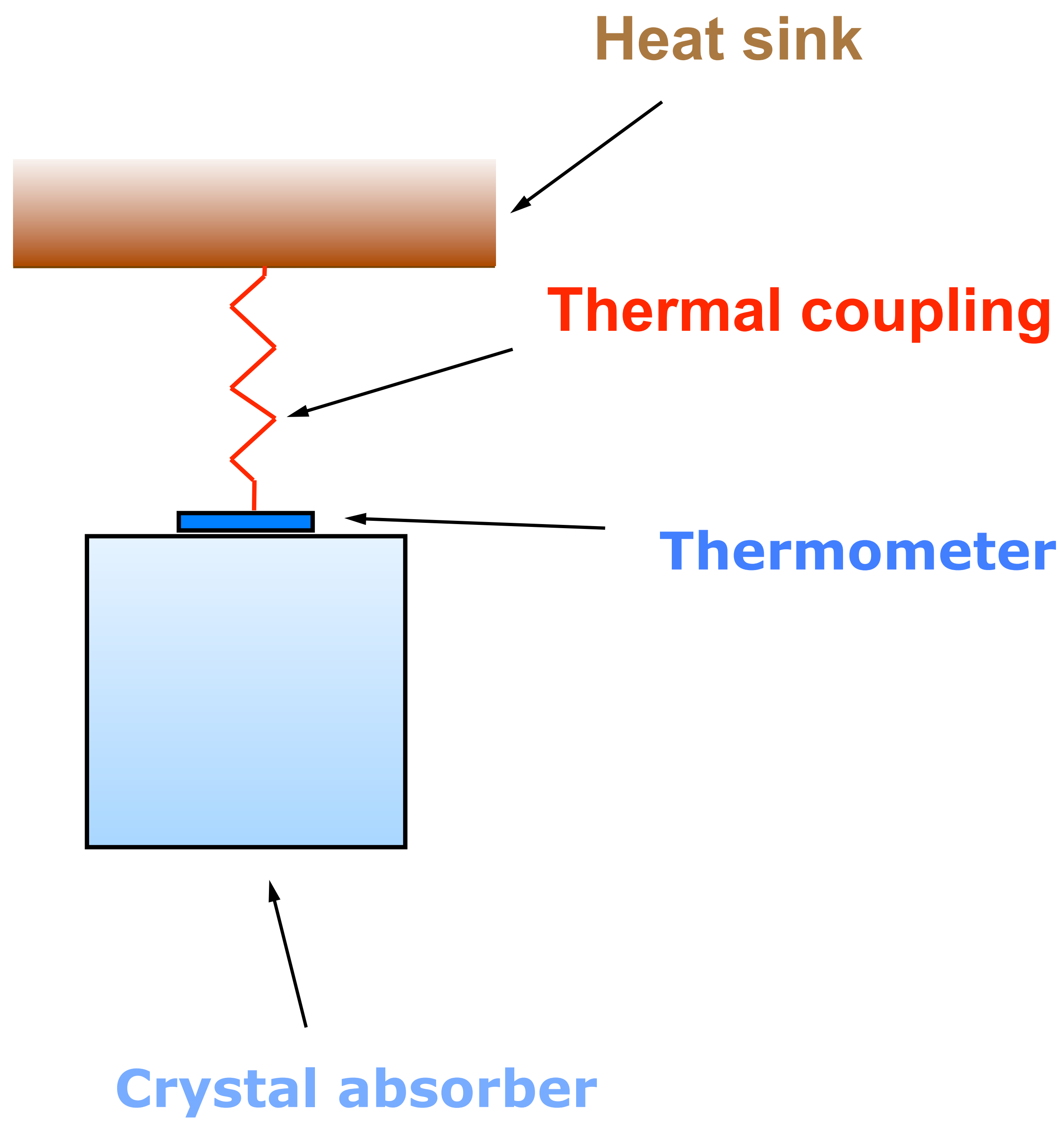


PhaseI (~15kg): Start data taking 2009

Phase III: GERDA + Majorana toward 1 ton detector
Depends heavily on background achieved in first two phases

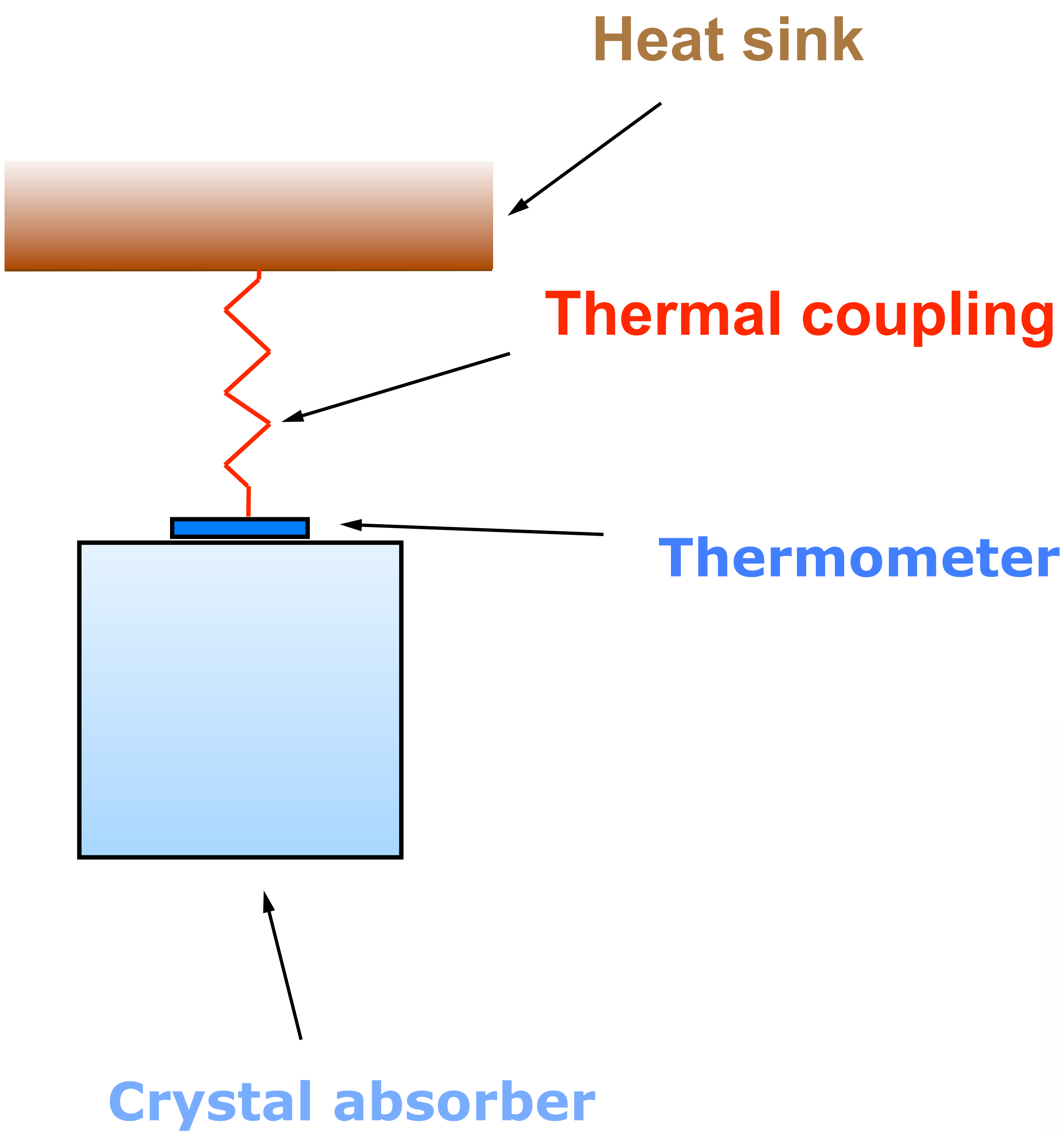
CUORICINO. ^{130}Te (Bolometer)

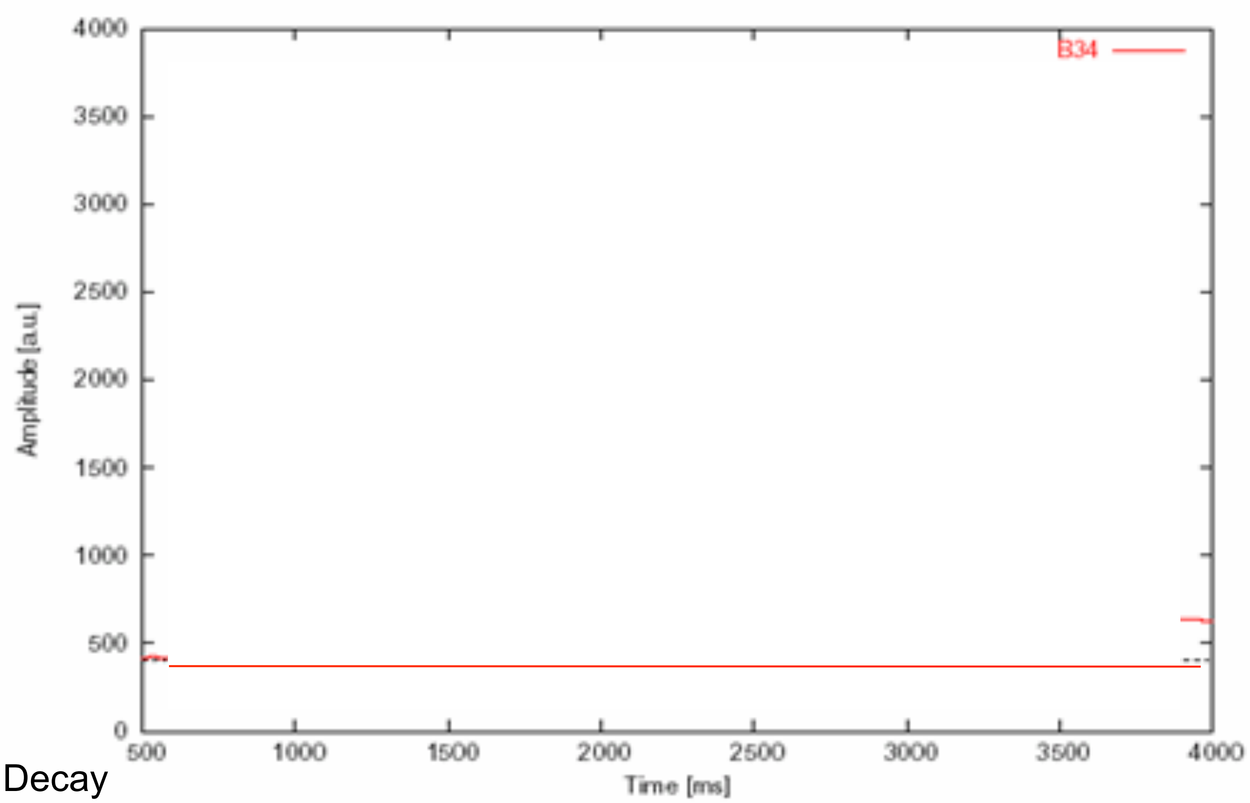
CUORICINO and NEMO-III are the only running DBD experiments at the moment



CUORICINO. ^{130}Te (Bolometer)

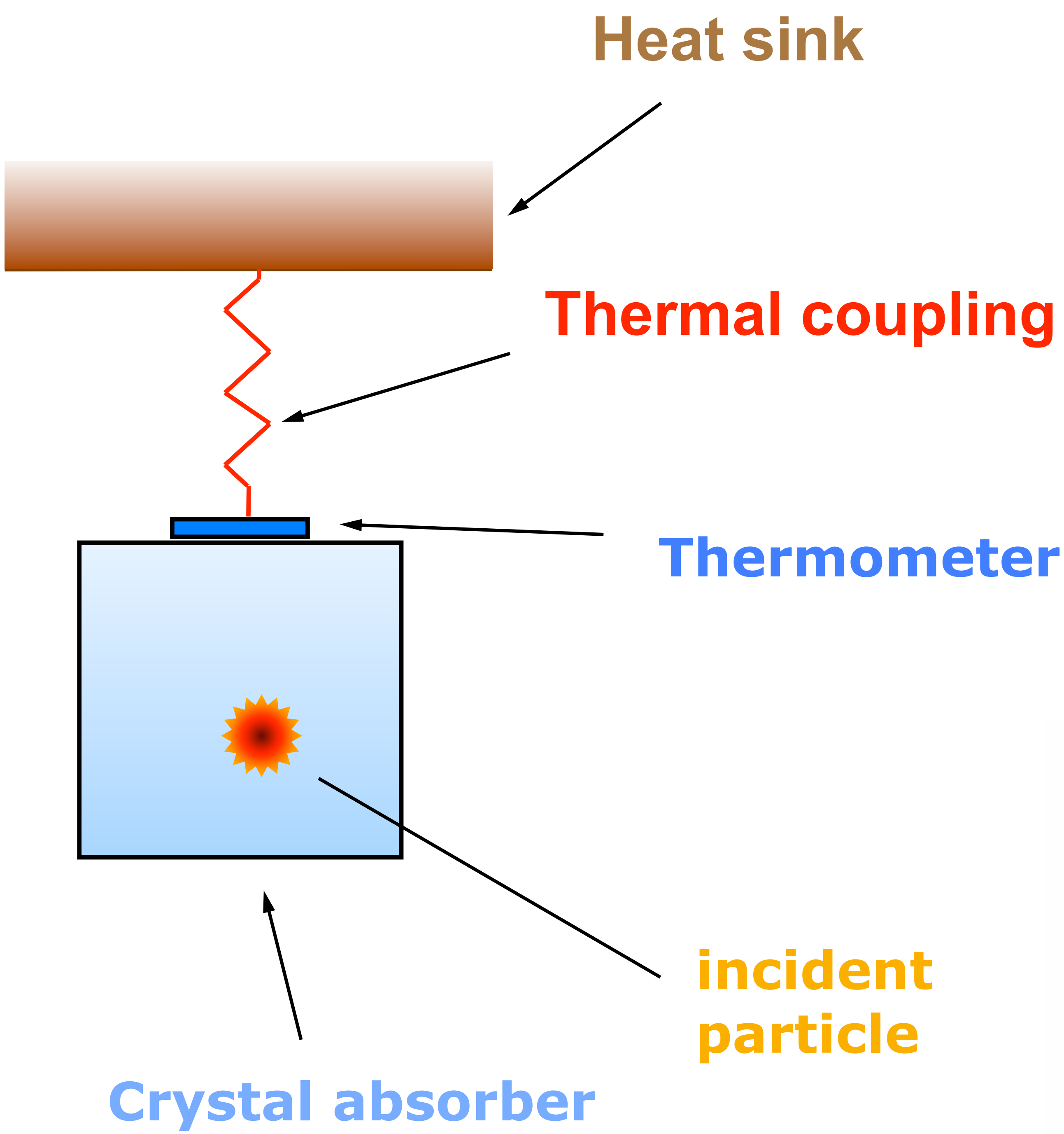
CUORICINO and NEMO-III are the only running DBD experiments at the moment

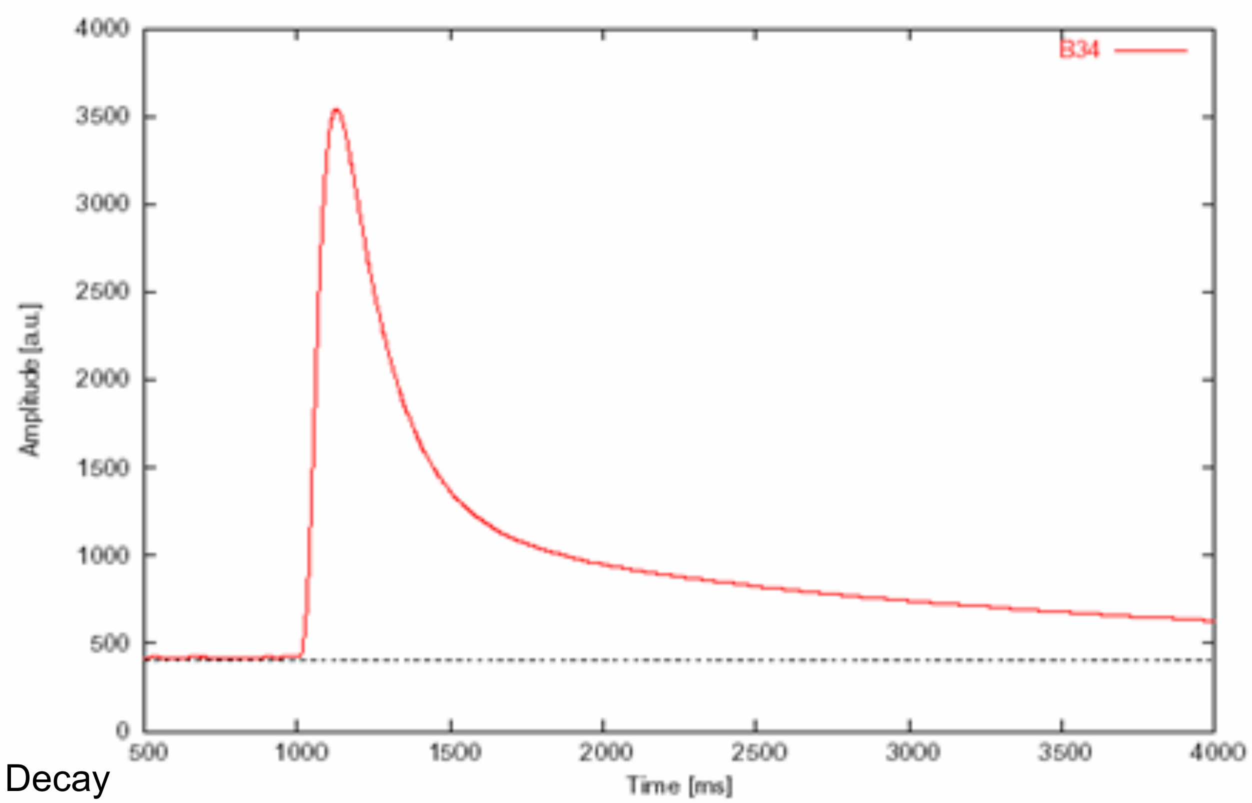




CUORICINO. ^{130}Te (Bolometer)

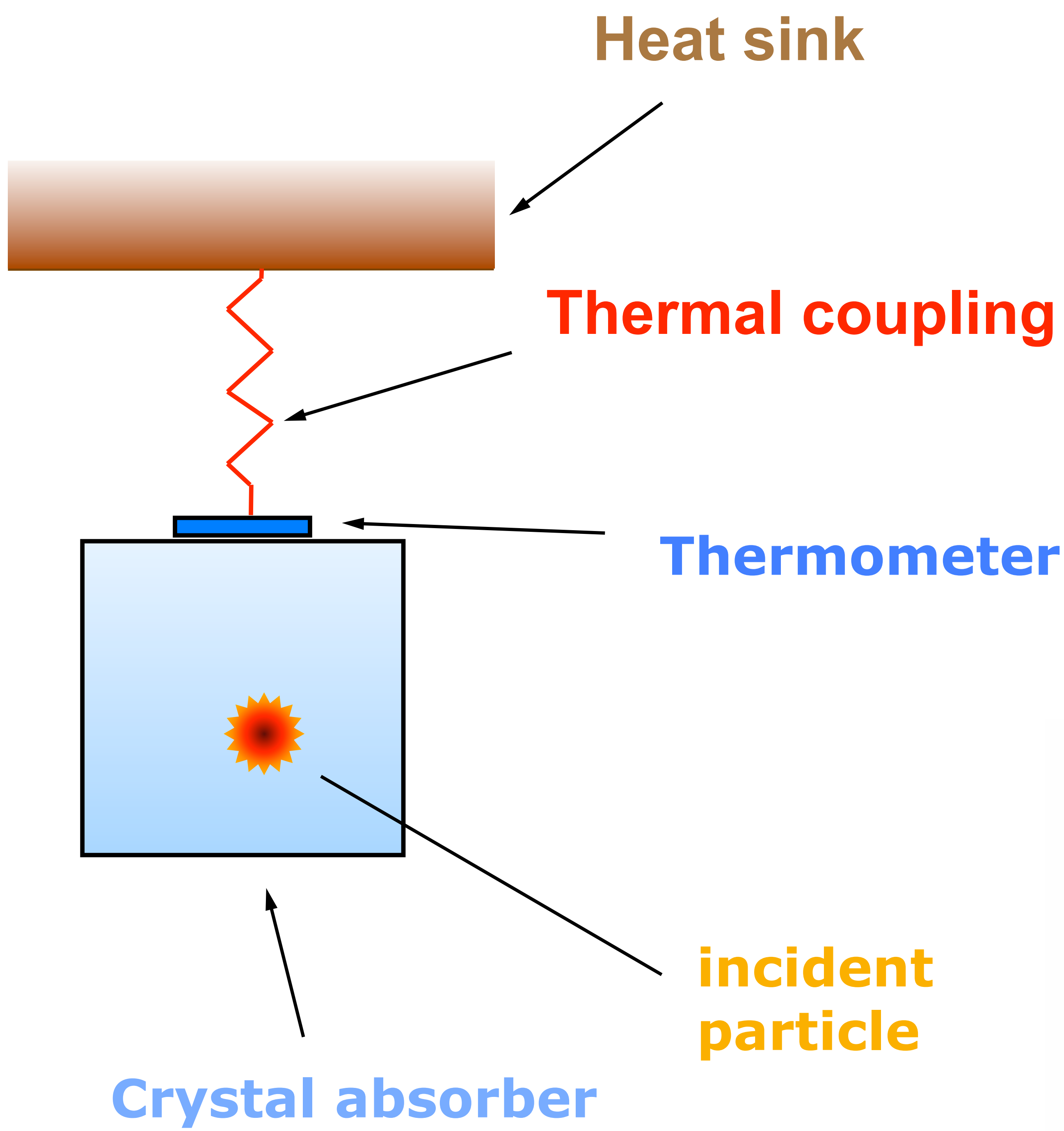
CUORICINO and NEMO-III are the only running DBD experiments at the moment





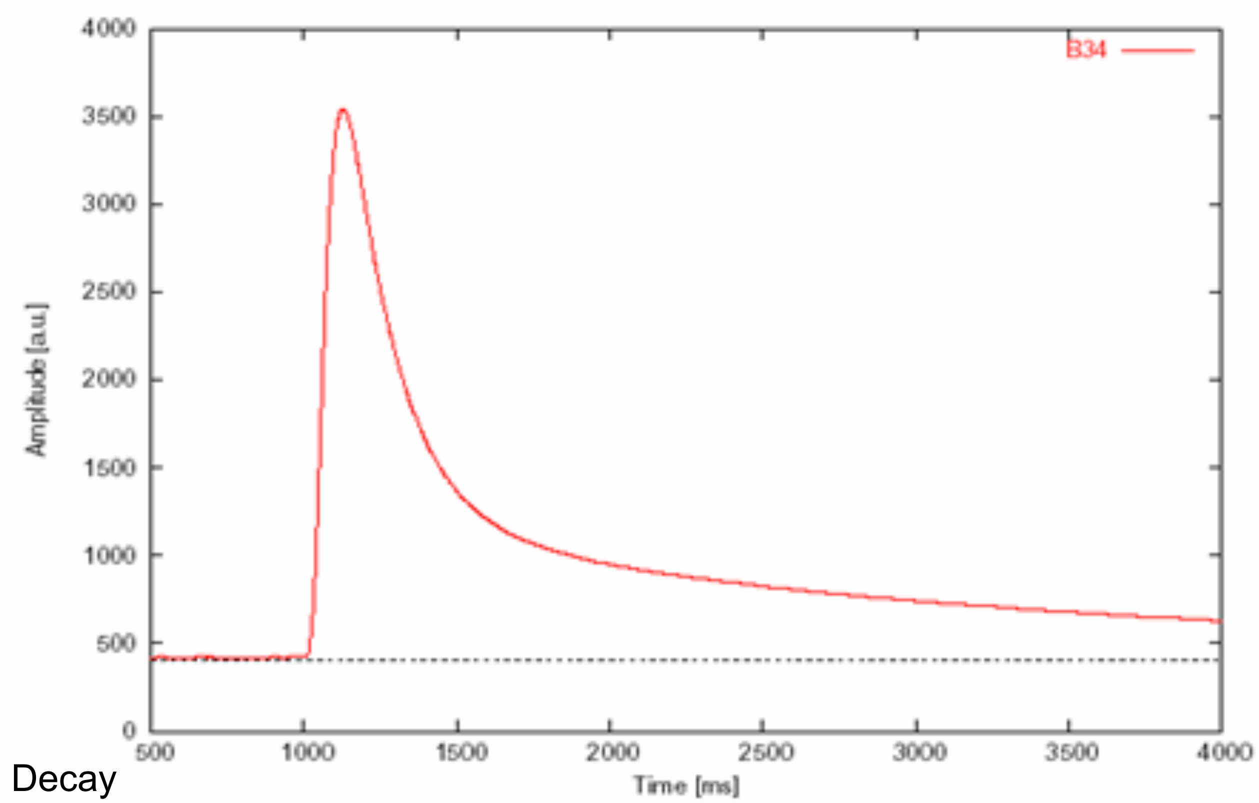
CUORICINO. ^{130}Te (Bolometer)

CUORICINO and NEMO-III are the only running DBD experiments at the moment



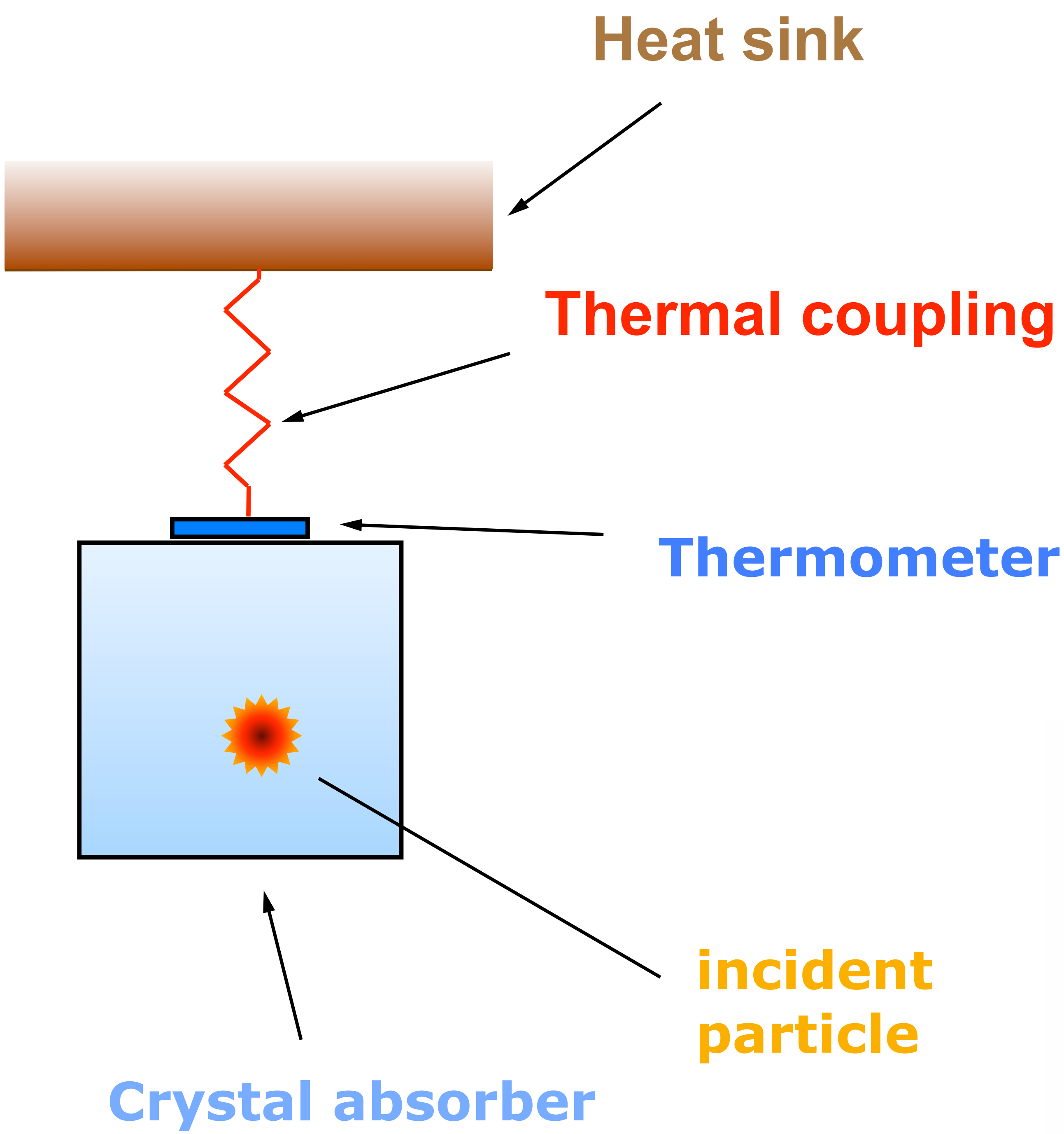
In a monolithic thermal model: thermal signal

$$\Delta T = E/C$$



CUORICINO. ^{130}Te (Bolometer)

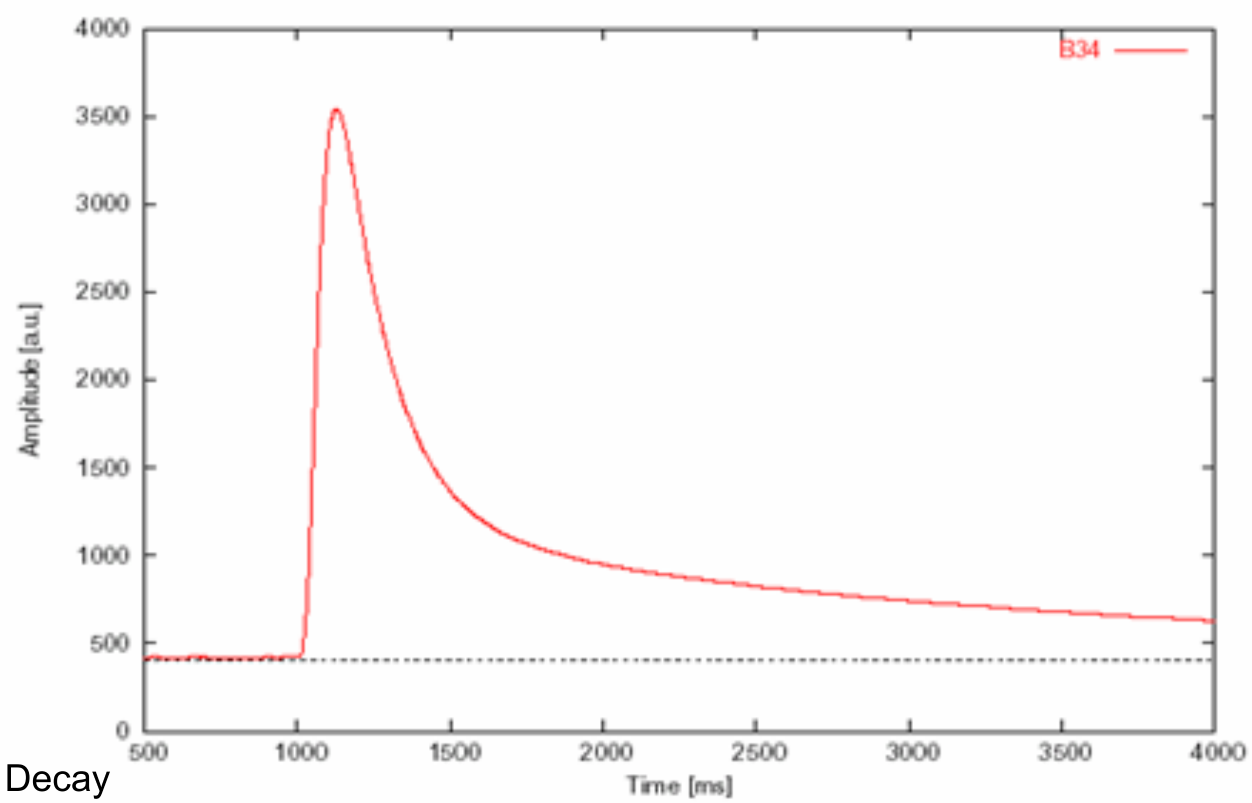
CUORICINO and NEMO-III are the only running DBD experiments at the moment



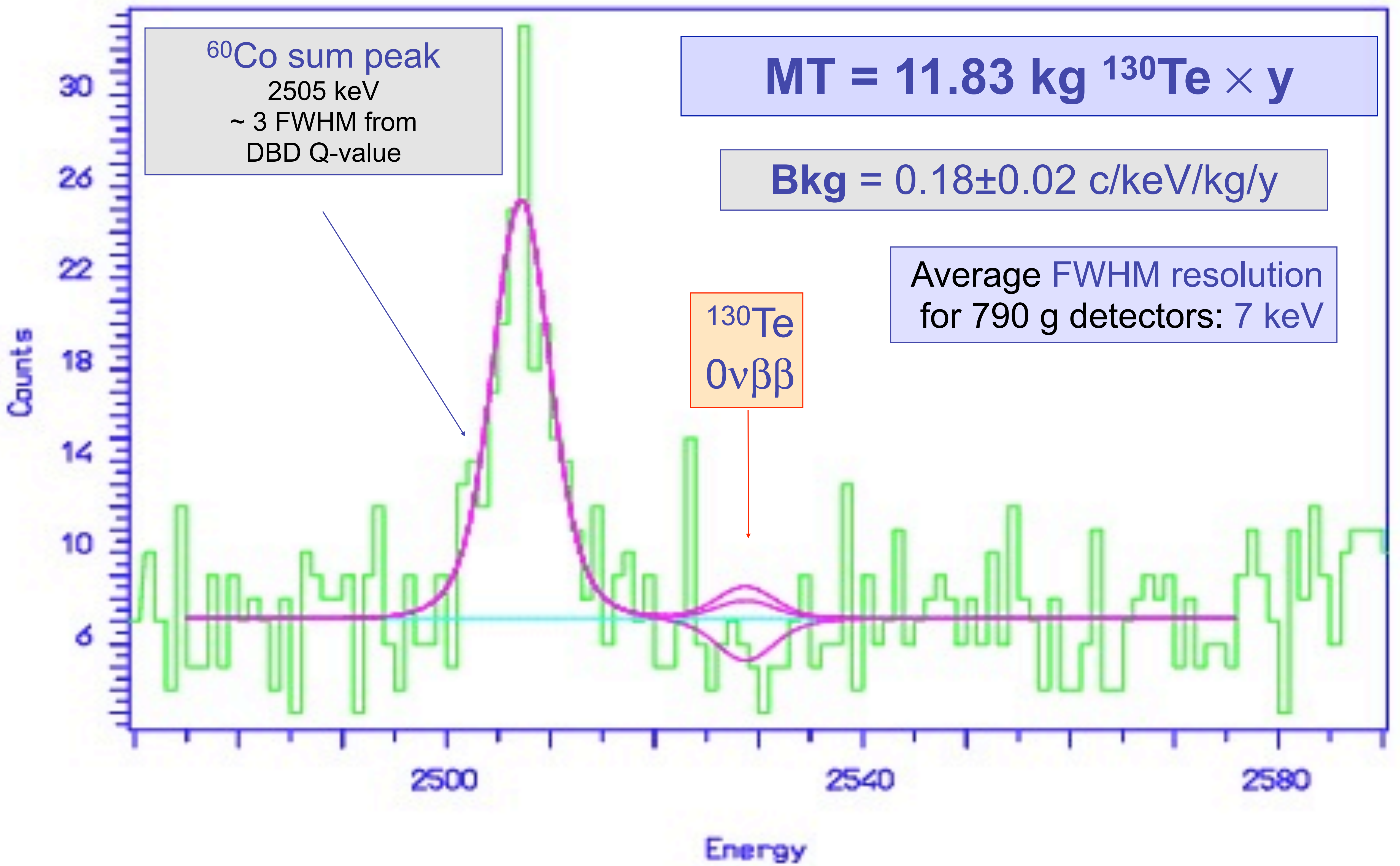
In a monolithic thermal model: thermal signal

$$\Delta T = E/C$$

The detector has to work at low temperature (10 mk) in order to develop high pulses



CUORICINO



$\tau_{1/2}^{0\nu}$ (y) > 3.0×10^{24} y (90% c.l.)

$\langle M_{\beta\beta} \rangle < 0.20 - 0.98$ eV

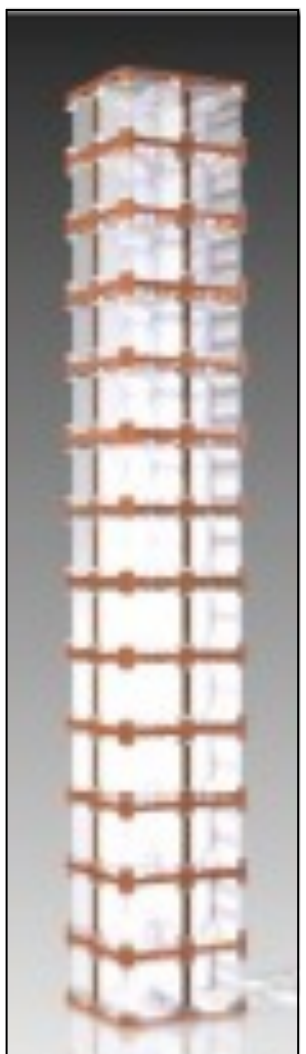
CUORE. ^{130}Te

Array of 988 detectors:
19 Cuoricino-like towers.

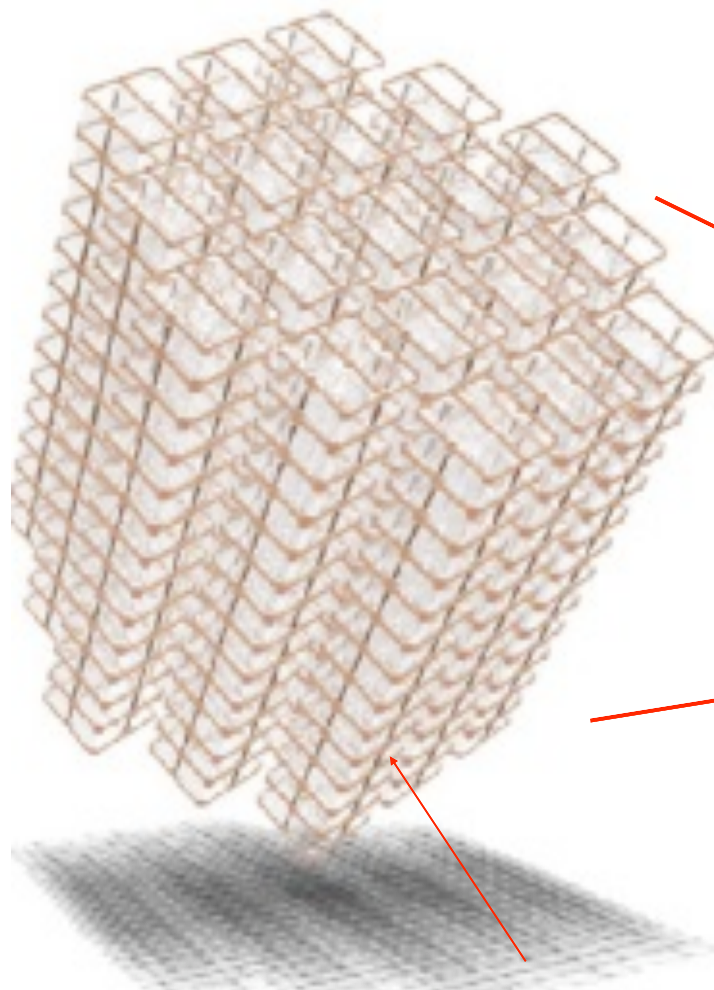
- $\Rightarrow$ M = 0.741 ton of TeO_2
- $\Rightarrow$ M = 600 kg of Te
- $\Rightarrow$ M = 203 kg of ^{130}Te

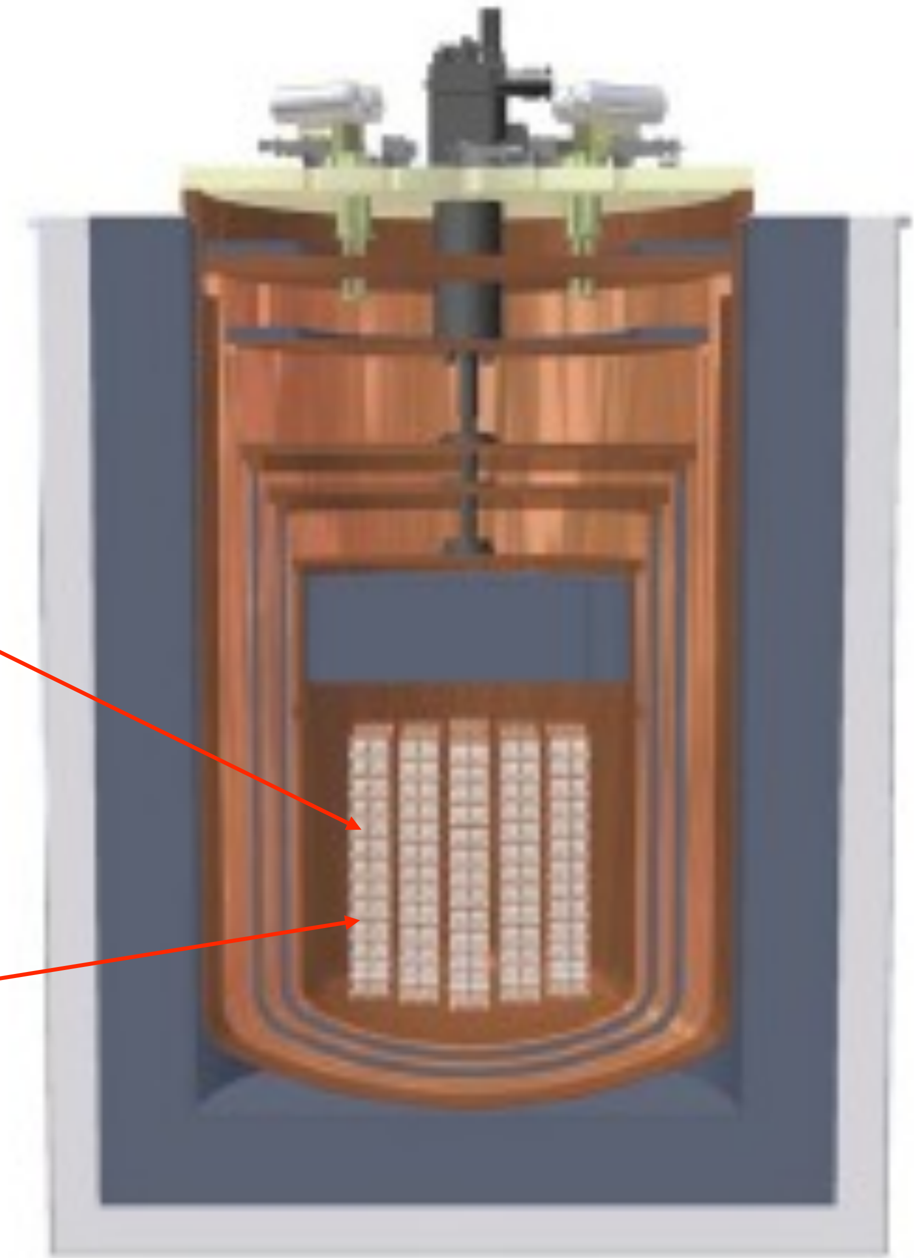
Location: Gran Sasso

Natural Te



90 cm





19 CUORICINO-like towers

CUORE

Main background: Surface contamination close to fiducial volume
(Recall CUORICINO background level 0.18±0.02 c/keV/kg/y)

Two-pronged approach to tackle this background:
Passive ⇒ surface cleaning
Active ⇒ event ID:

- Surface sensitive bolometers
- Scintillating bolometers

Enrichment is also possible

Expected sensitivities (5 years of data)

"Baseline" scenario

N_{bckg}=0.01 cts.keV^{-1}.kg^{-1}.yr^{-1}

$T_{½} > 2.1\ 10^{26}$ yr

$<m_\nu> < 0.03 – 0.17$ eV

"Aggressive" scenario

N_{bckg}=0.001 cts.keV^{-1}.kg^{-1}.yr^{-1}

$T_{½} > 6.6\ 10^{26}$ yr

$<m_\nu> < 0.015 – 0.1$ eV

Planned start-up: 2011

EXO – Enriched Xenon Observatory – ^{136}Xe

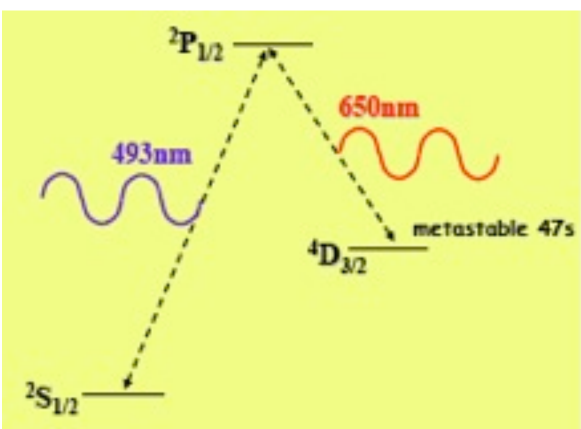



Liquid Xe TPC

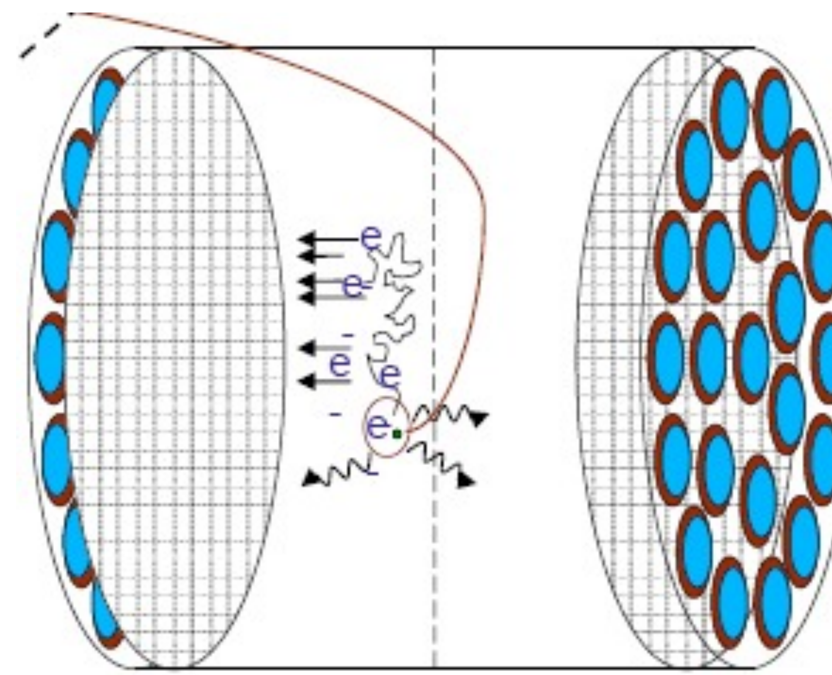
Energy measurement by ionization + scintillation

Tagging of Barium ion ($^{136}Xe \rightarrow {}^{136}Ba^{++} + 2e^-$)

Optical spectroscopy with Ba+



Ion Grabber/mover



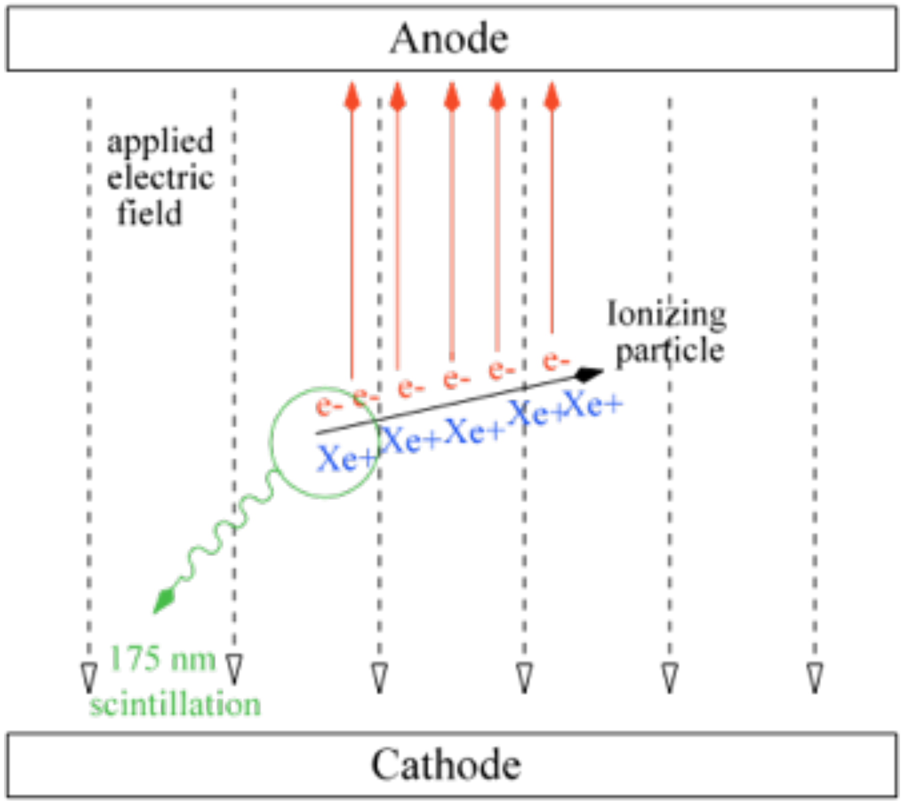
Case	Mass (ton)	Eff. (%)	Run Time (yr)	σ_E/E @ 2.5MeV (%)	$2\nu\beta\beta$ Background (events)	$T_{1/2}^{0\nu}$ (yr, 90%CL)	Majorana mass (meV) QRPA‡	NSM#
Conservative	1	70	5	1.6*	0.5 (use 1)	$2{*}10^{27}$	50	68
Aggressive	10	70	10	1†	0.7 (use 1)	$4.1{*}10^{28}$	11	15

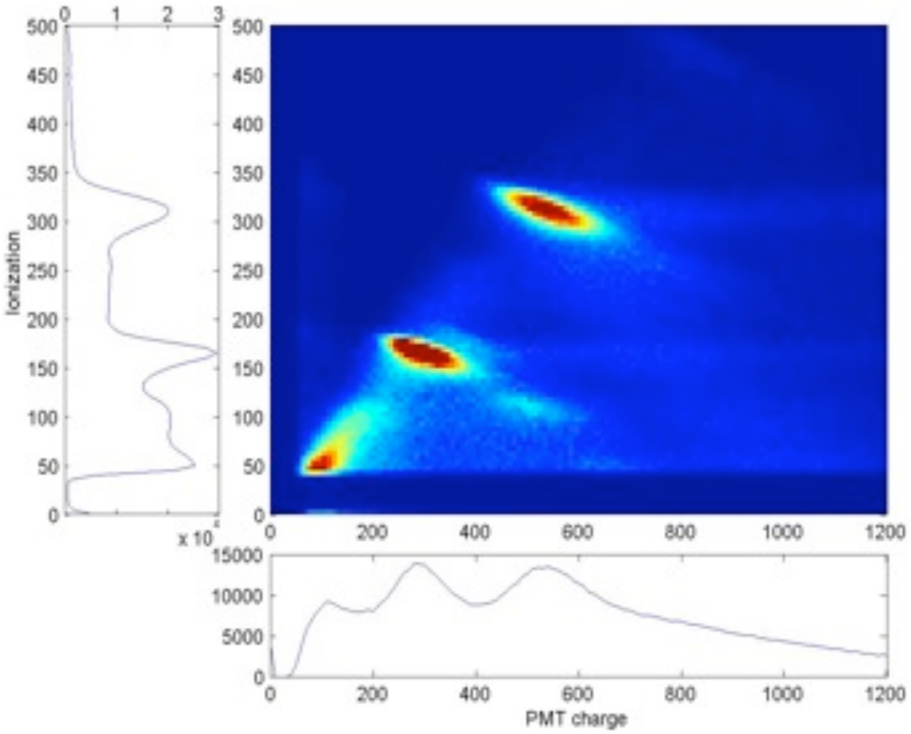
EXO

200 kg of LXe (80% enriched ^{136}Xe in hand)

No Ba+ tagging

Ionization + scintillation to improve $\Delta E/E$ and detect apha (BiPo bkg suppression)





Being commissioned in WIPP (New Mexico)

Physics run to start in 2009

Goals:

- Measure $2\nu\beta\beta$ of ^{136}Xe
- Search for $0\nu\beta\beta$ in ^{136}Xe with competetive
- sensitivity
  - $T_{1/2}^{0\nu} > 6.4\ 10^{25}$ y (after 2y)
- Understand the operation of a large LXe detector
  - Backgrounds
  - Resolution, Xe purification and handling

Apologies to

- Majorana
  - US-based ^{76}Ge experiment (segmented Ge detectors). Close collaboration with Ge. Merger between the two to go towards 1t.
- COBRA
  - R&D with CdZnTe semi-conductor detectors (room-ish temperature). Very interesting potential with pixelated detectors. Main isotopes ^{116}Cd and ^{130}Te
- CANDLES
  - Undoped CaF_2 scintillator crystal detectors. Isotope: ^{48}Ca. Start with natural, look into possibilities of enrichment in the future
- ....

Experiment	Isotope	kg	$T_{1/2}$ yr, 90% CL	m_ν*, meV	Start-up timescale	Status
HM	^{76}Ge	15	>1.9 10^{25}	230-560	1990	finished
KDHK claim	^{76}Ge	15	(0.7-4.2) 10^{25} (3σ)	150-920	1990	finished
NEMO 3	^{100}Mo	7	2 10^{24} (expect. 2010)	340-590	2003	running
CUORICINO	^{130}Te	11	>3 10^{24} (current)	260-610	2002	running
CUORE	^{130}Te	210	1.3 10^{26}	40-92	2011	approved
GERDA, Phase I	^{76}Ge	15	3 10^{25}	180-440	2009	approved
Phase II	^{76}Ge	~31	2 10^{26}	70-170	2011	approved
EXO 200	^{136}Xe	160	6.4 10^{25}	270-380	2009	approved
EXO 1t	^{136}Xe	800	2 10^{27}	50-68	2015	R&D
SuperNEMO	**^{82}Se/^{150}Nd**	**100**	**1 10^{26}**	**45-110**	**2012**	**Design Study**
COBRA	^{116}Cd	151	1.5 10^{26}	38-96	?	R&D
SNO+	^{150}Nd	500	(1-5) 10^{25}	50-100 (?)	2012?	R&D

CUORE, GERDA and SuperNEMO are on the European roadmap for astro-particle physics (ASPERA)

*** Matrix elements from MEDEX'07 or provided by experiments***

Summary

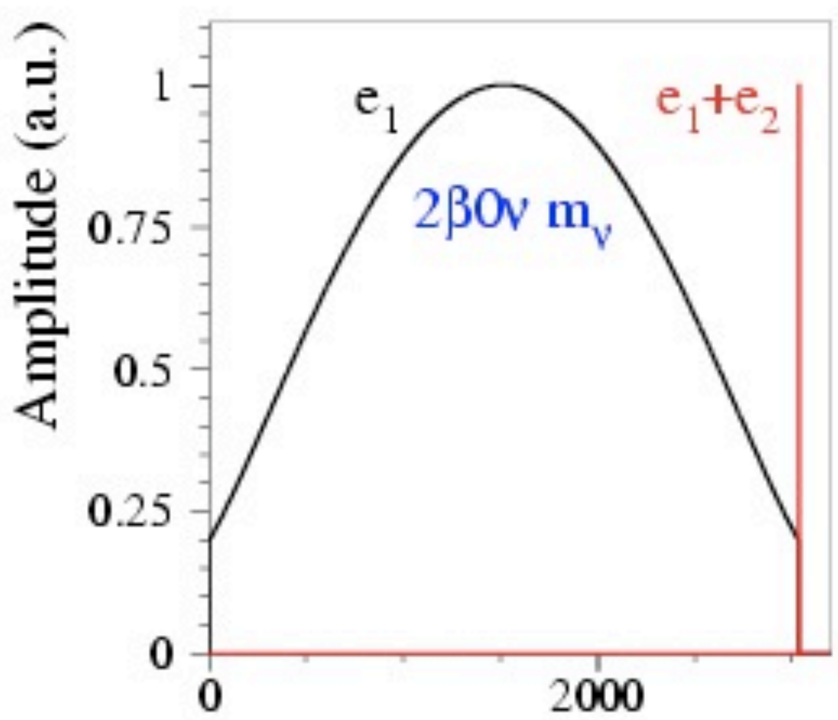
- $0\nu\beta\beta$ is the only practical way to establish neutrino mass nature (Majorana vs Dirac) and understand if $\Delta L \neq 0$.
- Vibrant and rapidly growing field
- Results will come in from several x100kg experiments in near-ish future
  - Big potential for a major discovery
  - Necessary step to converge on 1-2 ton-scale detector technology(ies)

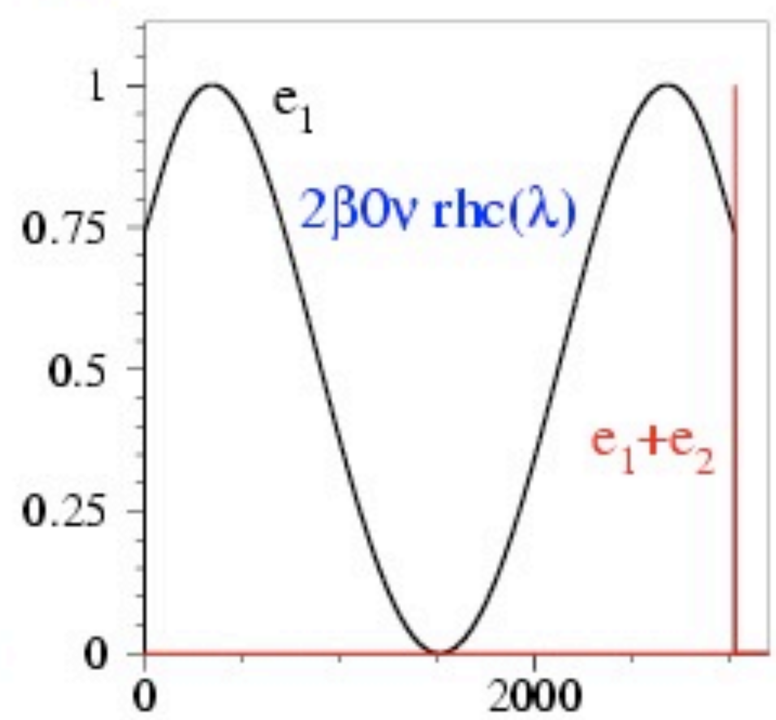
BACKUP

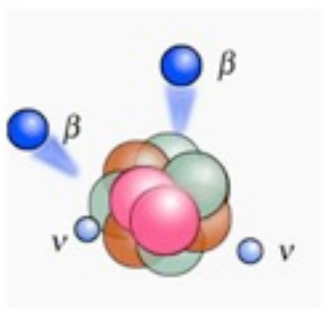
Topology detection may disentangle underlying physics mechanism

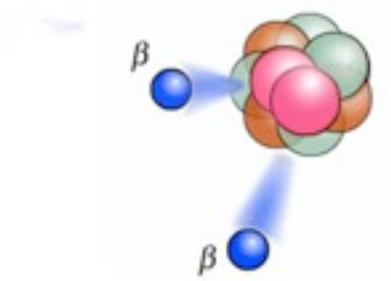
$\langle m_\nu \rangle$ $0\nu\beta\beta$ mechanism

V+A $0\nu\beta\beta$ mechanism

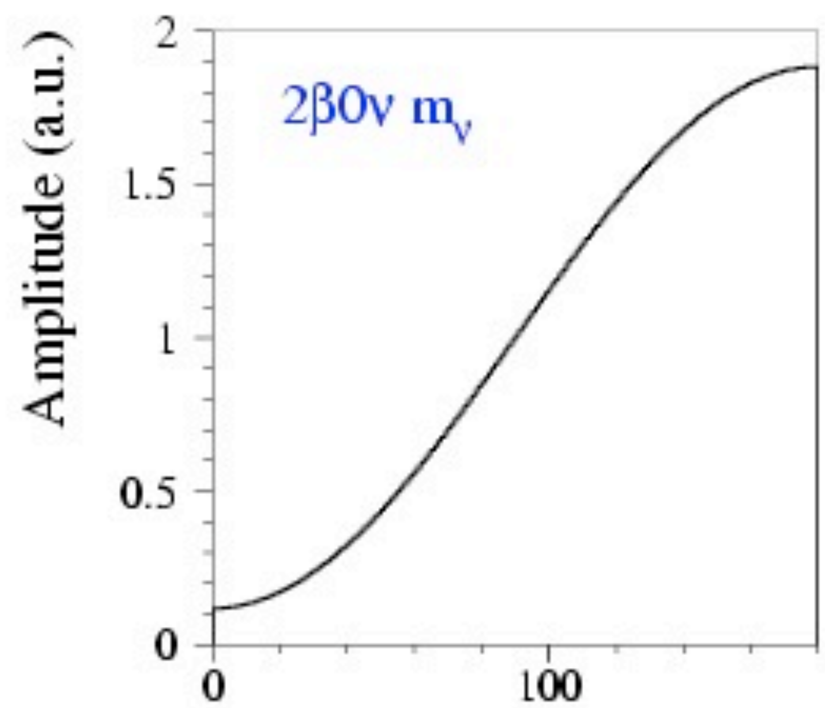


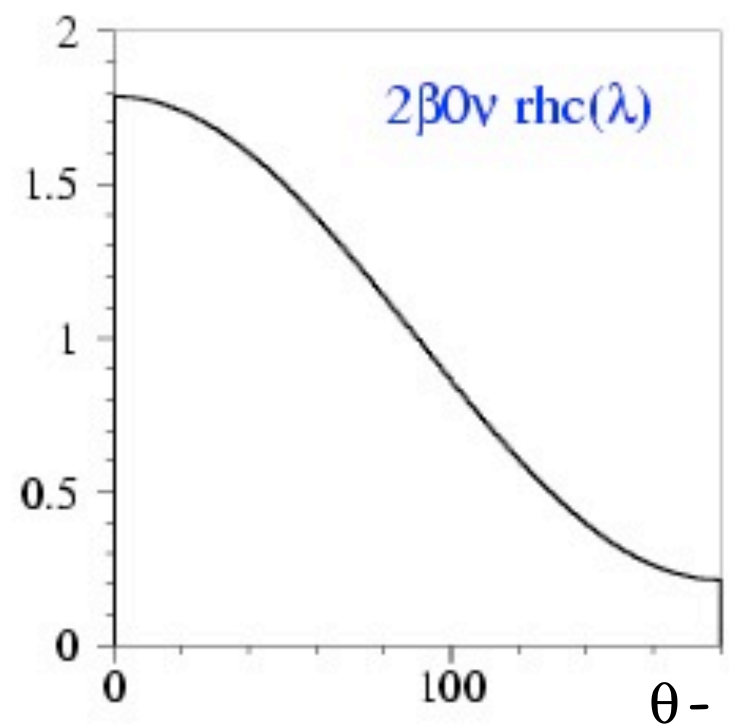






Energy (keV)





"Begging" for topological detection!

θ - angle between e_1 and e_2

$0\nu\beta\beta$ experiment is about BKG suppression!

detector efficiency — mass — live time

$$T_{1/2}(90\%CL) > N_A \frac{\varepsilon \ln 2}{n_{90} A} \sqrt{\frac{M\,T}{b\,\Delta E}}$$

bkg / (keV kg yr) in window ΔE

$N_{excluded}$ for 90% CL

Current best calorimeters (bolometers, Ge):

b = 0.2 counts/(keV kg yr) ΔE = 5 keV

bΔE = 1 c/kg yr

Next generation calorimeters (CUORE, GERDA):

b = 0.01 counts/(keV kg yr) ΔE = 5 keV

bΔE = 0.05 c/kg yr

$\Delta E/E$ = 0.2%

Current topological (NEMO3)

b = 0.002 counts/(keV kg yr) ΔE = 250 keV $\Delta E/E$ = 8%

bΔE = 0.5 c /kg yr

Next generation topological (SuperNEMO)

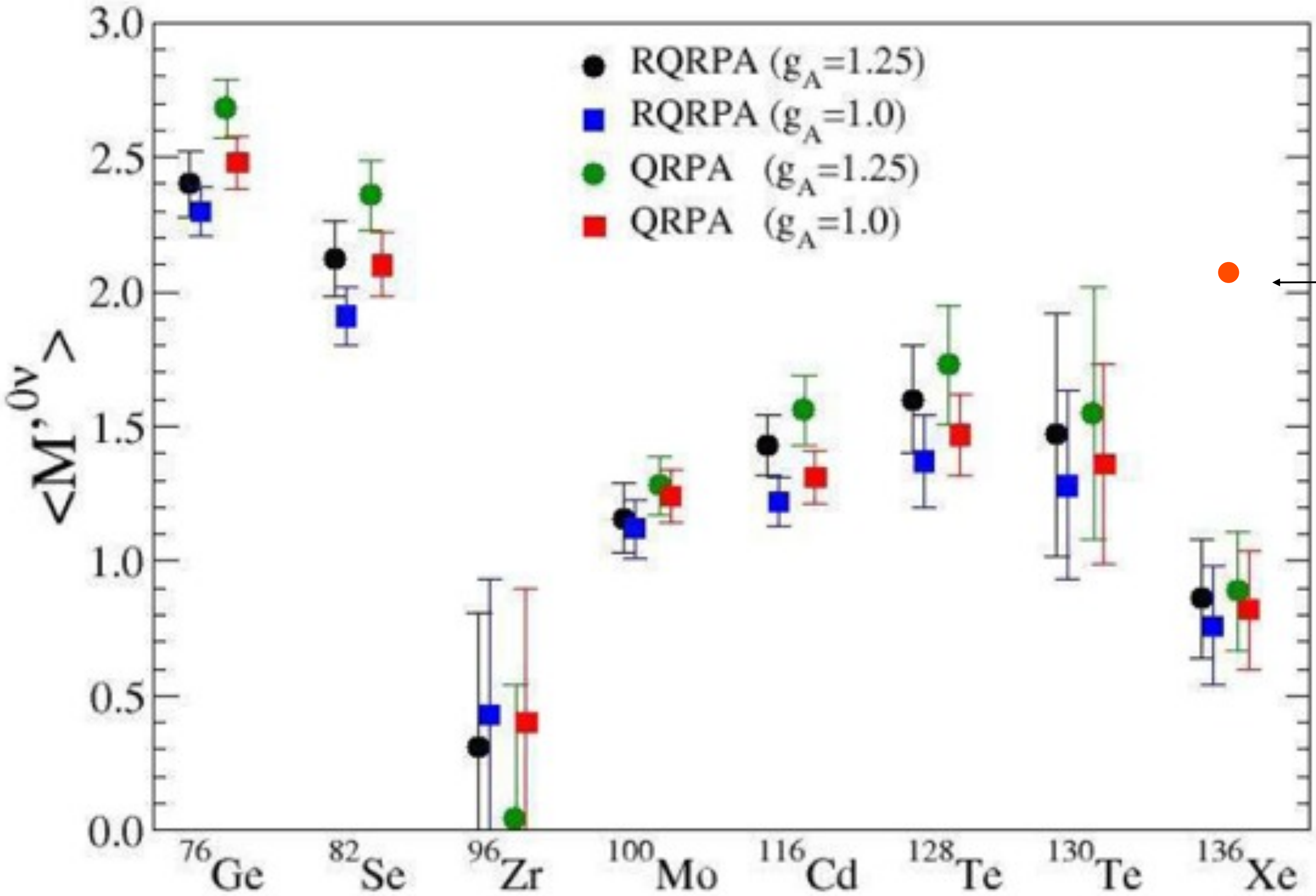
b = 5×10^{-5} counts/(keV kg yr) ΔE = 150 keV $\Delta E/E$ = 4%

bΔE= 0.0075 c/kg yr

"Dream" scenario for calorimeters:

b = 0.001 counts/(keV kg yr) ΔE = 5 keV $\Delta E/E$ = 0.2%

bΔE = 0.005 c/kg yr

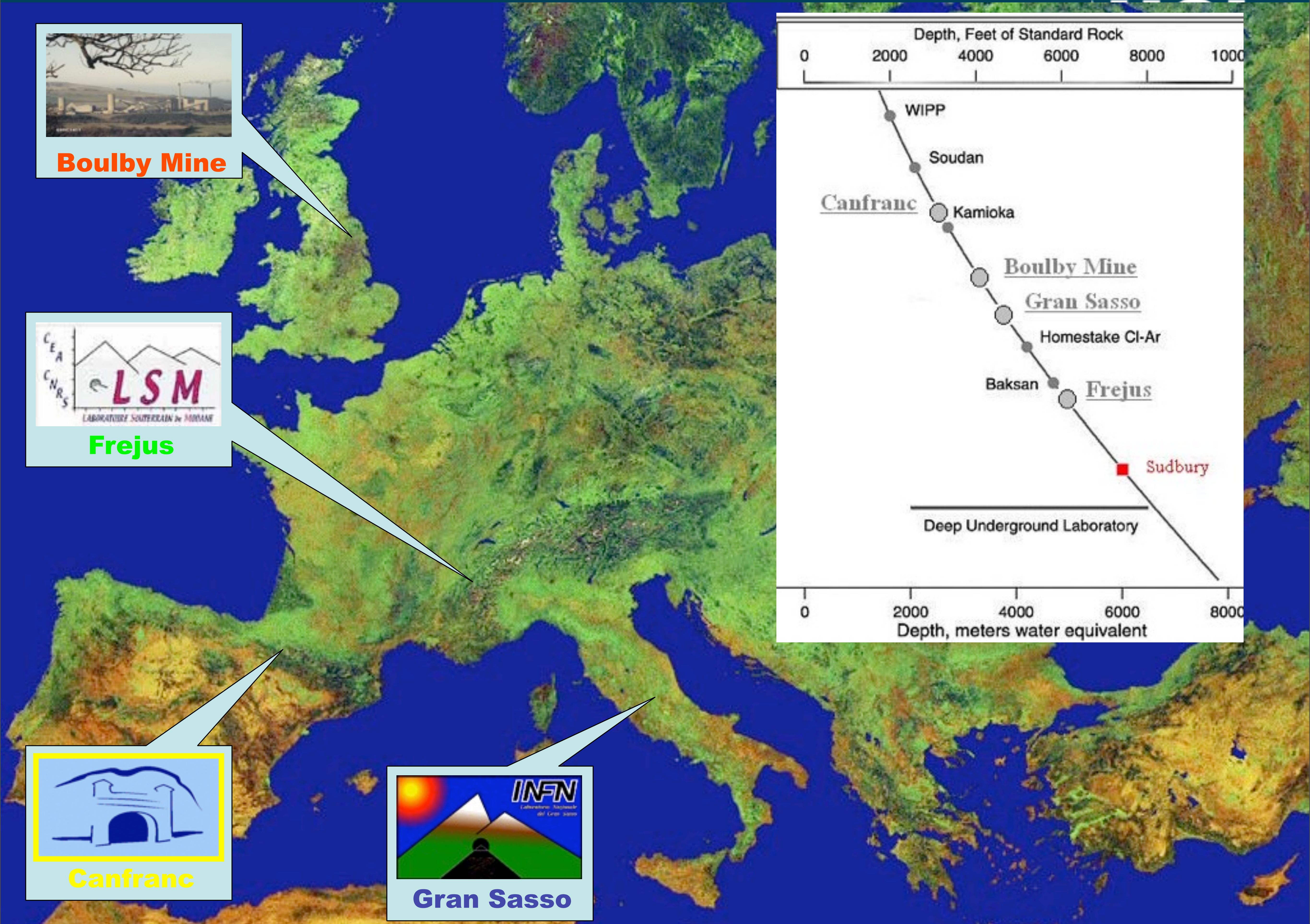


^{150}Nd (without deformation taken into account)

$$\Gamma^{0\nu} = G^{0\nu}(E_0, Z)\left|M^{0\nu}\right|^2 \left(\frac{< m_\nu >}{m_e}\right)^2$$

Phase space factor

Isotope	^{48}Ca	^{76}Ge	^{82}Se	^{96}Zr	^{100}Mo	^{116}Cd	^{130}Te	^{136}Xe	^{150}Nd
$G^{0\nu}$ $\times 10^{-15}$ yr^{-1}	75.8	7.6	33.5	69.7	54.5	58.9	52.8	56.3	249


Boulby Mine
Frejus
Canfranc
Gran Sasso
INFN
Depth, Feet of Standard Rock
0
2000
4000
6000
8000
WIPP
Soudan
Canfranc
Kamioka
Boulby Mine
Gran Sasso
Homestake Cl-Ar
Baksan
Frejus
Sudbury
Deep Underground Laboratory
0
2000
4000
6000
8000
Depth, meters water equivalent

LSM extension project. Plan to be ready in 2013.