

# Theoretical Considerations on Top Mass Observables

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LHC Top Working Group, November 2015

# Outline

- ▶ Some general consideration on top Mass measurement
- ▶ Top mass in production and decay
- ▶ Perturbative and non-perturbative theoretical errors
- ▶ Pole mass and  $\overline{\text{MS}}$  mass
- ▶ How to think of top mass measurements “Which mass are we measuring”: *is this the question?*
- ▶ New generators
- ▶ Hadronization errors

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# Measuring the Top Mass

- ▶  $m_t$  is a parameter of the Standard Model. We measure it by comparing **Standard Model predictions** with **experimental measurements** as a function of  $m_t$ .
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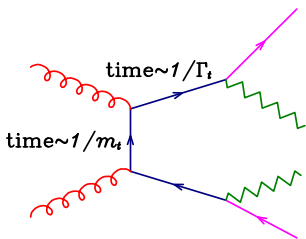
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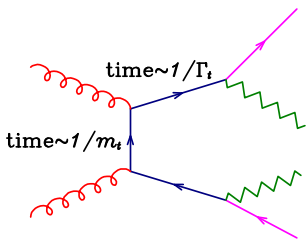
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## $m_t$ in production and decay



- ▶ The Top mass enters in both **production and decay**.
  - ▶ The two phenomena are relatively independent, they take place at two **widely different timescales**.
  - ▶ Top mass observables may be sensitive to **either or both** phenomena.
- ▶ From decay:
    - ▶ Reconstructing the mass of the  $W$  and  $b$  jet
    - ▶ End point spectrum of the lepton and the  $b$  jet/hadron
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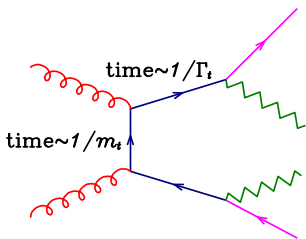
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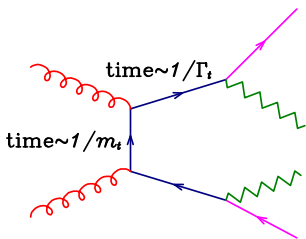
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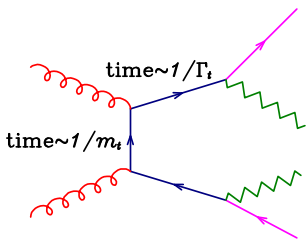
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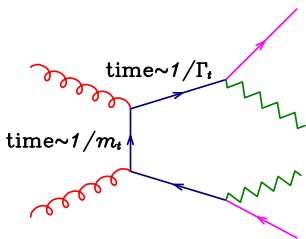
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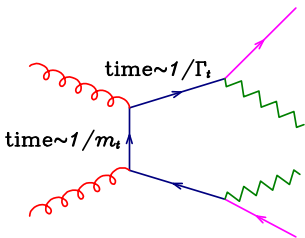
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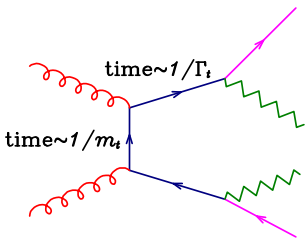
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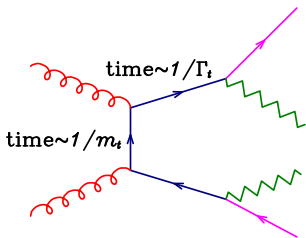
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# Theoretical errors: perturbative and non-perturbative

- ▶ “Perturbative” errors: vary scales, shower cutoff, etc.. In other words: **things that we know how to discuss**
- ▶ “Non Perturbative” errors: even the most elementary task, like assign the top decay products to the top quark, becomes a **difficult question to answer**.
  - ▶ Since the top is coloured, the associated  $b$  hadronic system must include particles not arising from top decay.
  - ▶ We don't have a theory, only models; we should explore variations in our models, constrain the variations using data, assess the error due to the **unknown** physics of hadron formation.

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## $\overline{\text{MS}}$ versus pole mass

Even if we demonstrate that hadronization effects yield a negligible error on the top mass, and that we should only worry about **perturbative** issues, a problem of **non-perturbative nature** arises if we measure the top pole mass (using observables that are sensitive to the mass of the top decay products), and then we relate it to short distance parameters of the theory like the top  $\overline{\text{MS}}$  mass.

It is often stated that when expressing the top  $\overline{\text{MS}}$  mass we are subject to an error ranging from 100 to 200 MeV related to infrared renormalons.

It is possible, also thank to the recent progress illustrated in the previous talk, to make some progress on this issue.

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(Marquard, A.V.Smirnov, V.A.Smirnov, Steinhauser, Feb.2015):

$$m_{t,\text{pole}} = m_{t,\overline{\text{MS}}}(1 + 0.4244\alpha_s + 0.8345\alpha_s^2 + 2.375\alpha_s^3 + (8.49 \pm 0.25)\alpha_s^4)$$

The last term corresponds to a correction of 200 MeV.

Pole mass affected by IR renormalons:

$$m_{t,\text{pole}} = m_{t,\overline{\text{MS}}} \left( 1 + \sum_{n=0}^{\infty} r_n \alpha_s^{n+1} \right),$$

For large  $n$  (Beneke, Braun, 1994; Beneke 1994):

$$(\partial \alpha_s / \partial \log \mu^2 = -b_0 \alpha_s^2 - b_1 \alpha_s^3 \dots)$$

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# Renormalons, Factorial growth and power corrections

Factorially divergent perturbative expansion: breaks down when  $(2b_0)^n \alpha_s^n n!$  for some  $n$  stops decreasing.

Take the ratio of two subsequent terms:

$$\frac{(2b_0)^{n+1} \alpha_s^{n+1} (n+1)!}{(2b_0)^n \alpha_s^n n!} \approx 1 \implies (2b_0) \alpha_s n \approx 1 \implies n \approx \frac{1}{2b_0 \alpha_s}$$

Size of the last good term (using  $n! \approx n^n e^{-n}$  and  $\alpha_s = \frac{1}{b_0 \log \frac{\mu^2}{\Lambda^2}}$ ):

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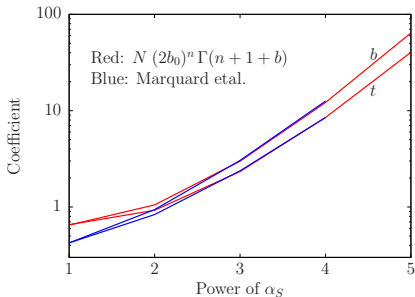
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Is Beneke's formula **consistent** with Marquard et al result? Fitting  $\alpha_s^4$  coefficient with Beneke's formula, we get  $N = 0.726$ , and fit well  $\alpha_s^3$  coeff. for  $t$  and  $b$ , and  $\alpha_s^4$  for  $b$ , **use it to predicts  $c_j, j \geq 5$  !!!**



Assuming  $\alpha_s = 0.1088$ , we get the  $\mathcal{O}(\alpha_s^5)$  contribution:

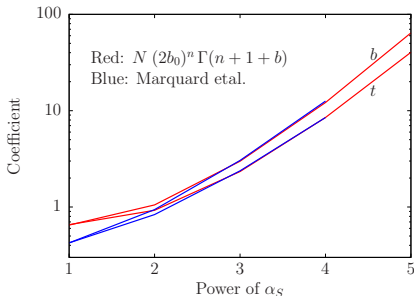
$$M_{\text{pole}} = 163.643 + 7.557 + 1.617 + 0.501 + 0.195 + \mathbf{0.10} \text{ GeV}$$

The terms in the perturbative expansion reach their minimum at order  $8 \sim 9$ , with last correction  $\approx 0.043$  GeV.

Alternatively:  $\Lambda_6 = 0.094 \implies \mathbf{N\Lambda_6 = 0.068} \text{ GeV}$  can be considered an estimate of the renormalon ambiguity in the determination of the pole mass.



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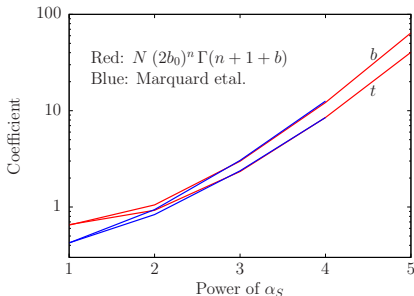
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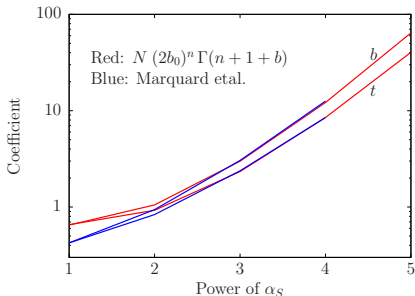
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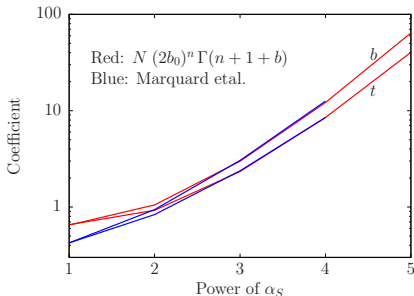
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- ▶ Compute these observables with the most precise tools available
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Simple (toy) example:

Use a generic, LO Shower MC to compute:

- ▶ The top production cross section:
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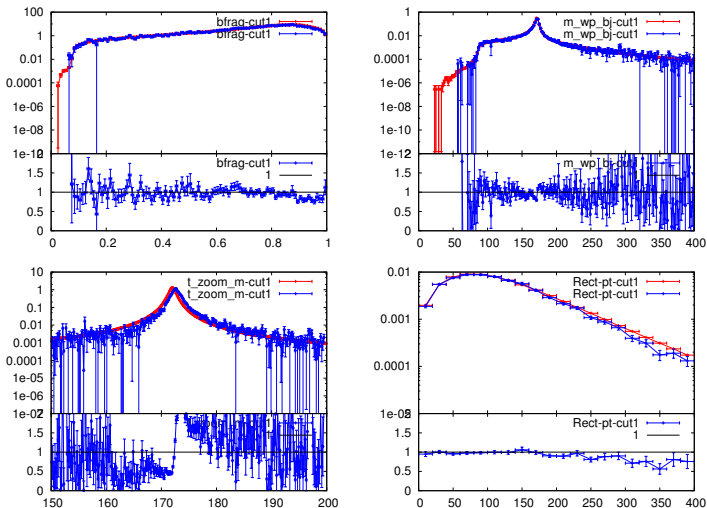
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First 64K of fully Pythia8 showered events from *bb4l* generator (blue) compared with Campbell, Ellis, Re, P.N, 2014.



(notice 0.5 GeV difference in mass ...)

# Non-perturbative error

- ▶ Rigorous approaches (Hoang): collect non-perturbative effects into **universal factors** that can be computed in other processes (**still far from practical**).
- ▶ Explore the freedom in changing parameters/features of hadronization model (Skands, Sjöstrand, colour reconnection sensitivity).
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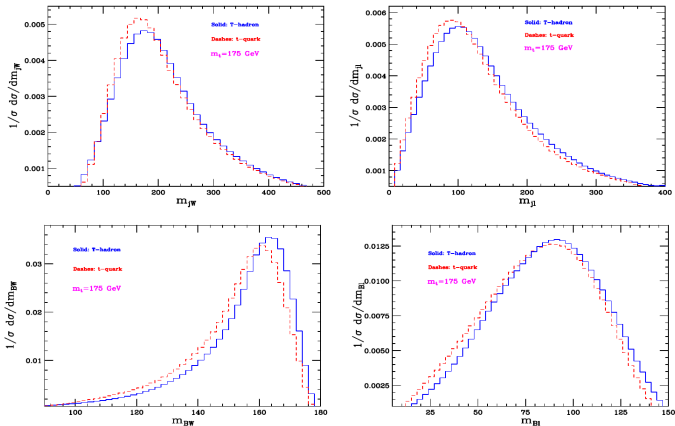


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# Another study: compare MC results for long-lived (hadronizing) top (in blue) to standard result (red) (Corcella, Mangano).

$pp$  collisions at  $\sqrt{s} = 8$  TeV, dilepton channel,  $k_T$  algorithm,  $R = 0.7$ ,  $p_{T,j} > 30$  GeV,  $p_{T,\ell} > 20$  GeV, MET > 20 GeV,  $|\eta_{j,\ell,\nu}| < 2.5$  (preliminary)



# Conclusions

- ▶ The renormalon ambiguity in the top pole mass may be much smaller than previously thought. We should not be afraid to focus upon variables that are pole mass dependent.
- ▶ “what kind of mass” in event generators: if we examine closely this question we see that it is more appropriate to ask instead “what kind of accuracy”.
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- ▶ “what kind of mass” in event generators: if we examine closely this question we see that it is more appropriate to ask instead “what kind of accuracy”.
- ▶ Generators/calculations with higher accuracy are available: we should use them.
- ▶ The estimate of non-perturbative effects on the top mass determination is the most difficult problem. More work on Monte Carlo hadronization models is desirable.

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