

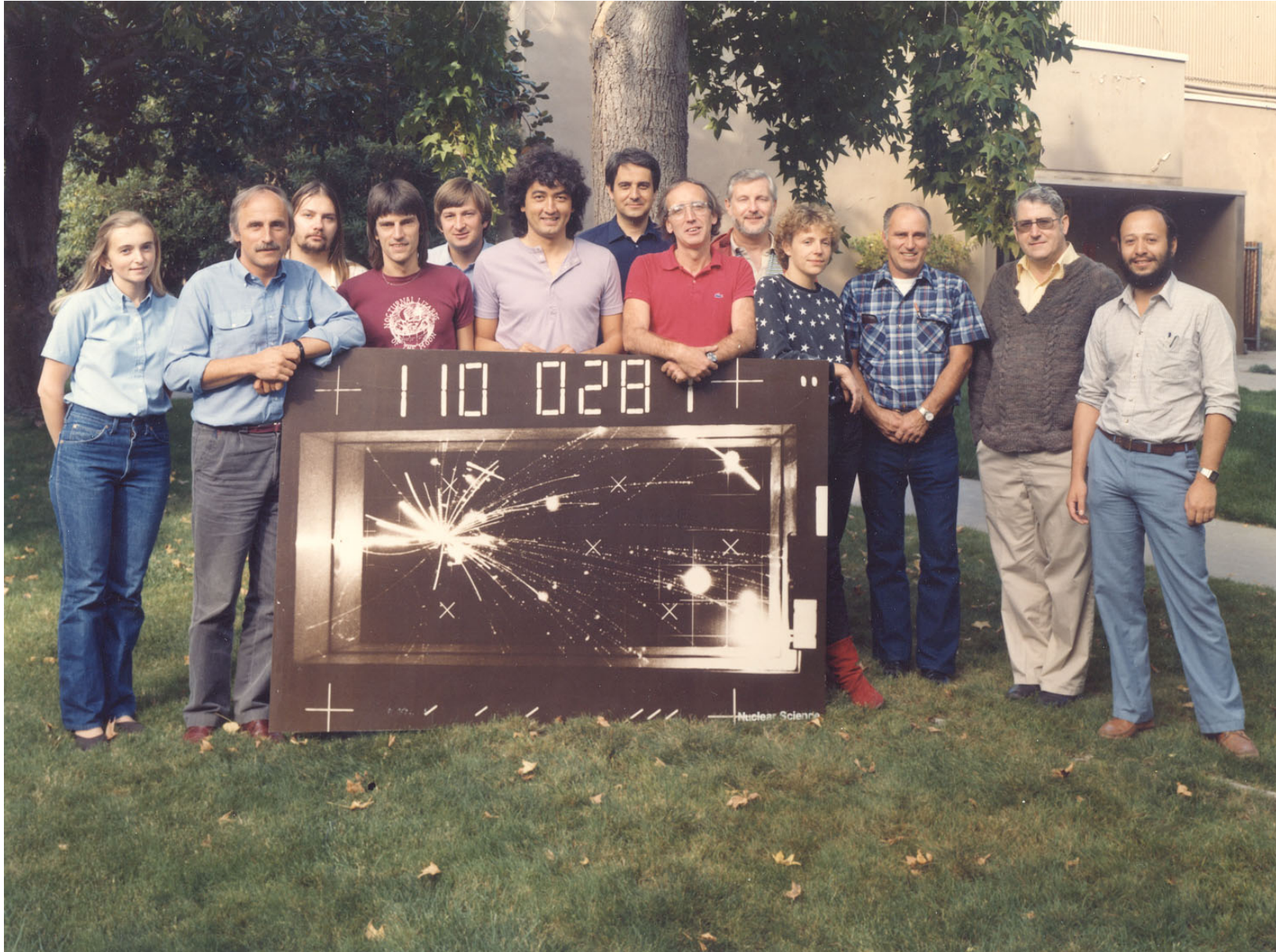
SQM 2016 Berkeley

**Pioneering Years:
with Grazyna Odyniec
at the Bevalac and SPS**

Reinhard Stock



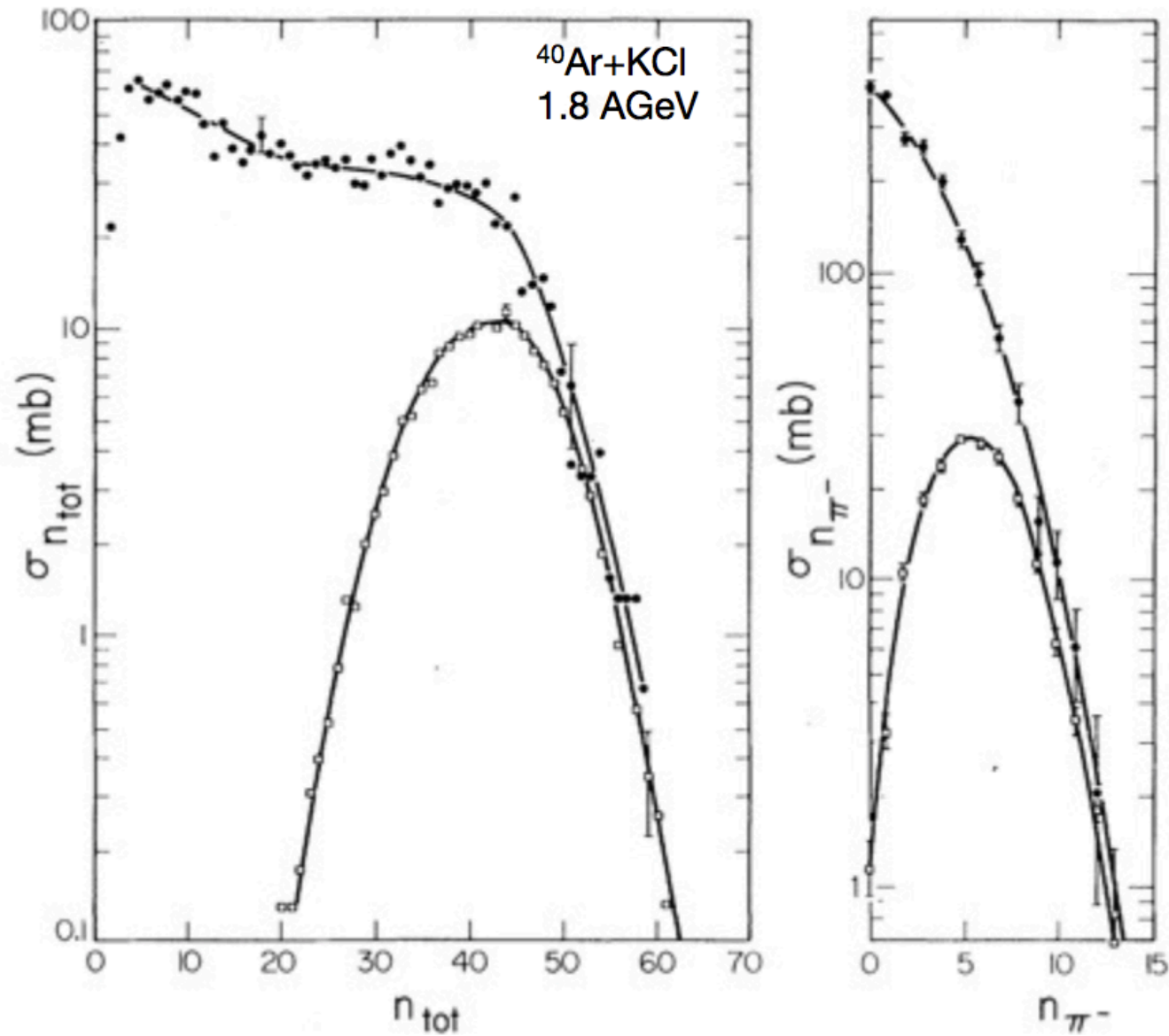
The Streamer Chamber Collaboration at LBL 1983



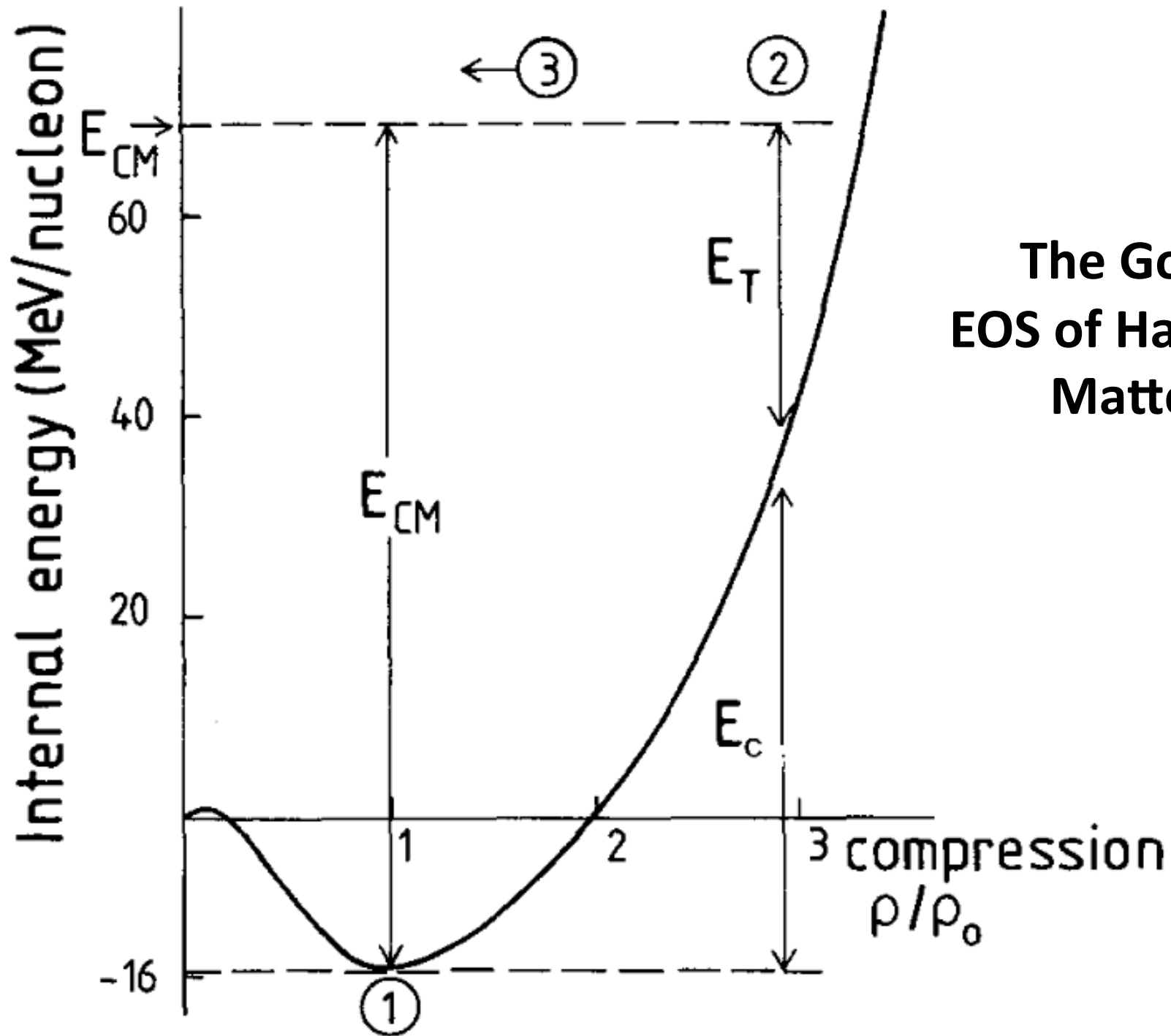
NA35 Collaboration at CERN SPS 1986



Historic Streamer Chamber Data

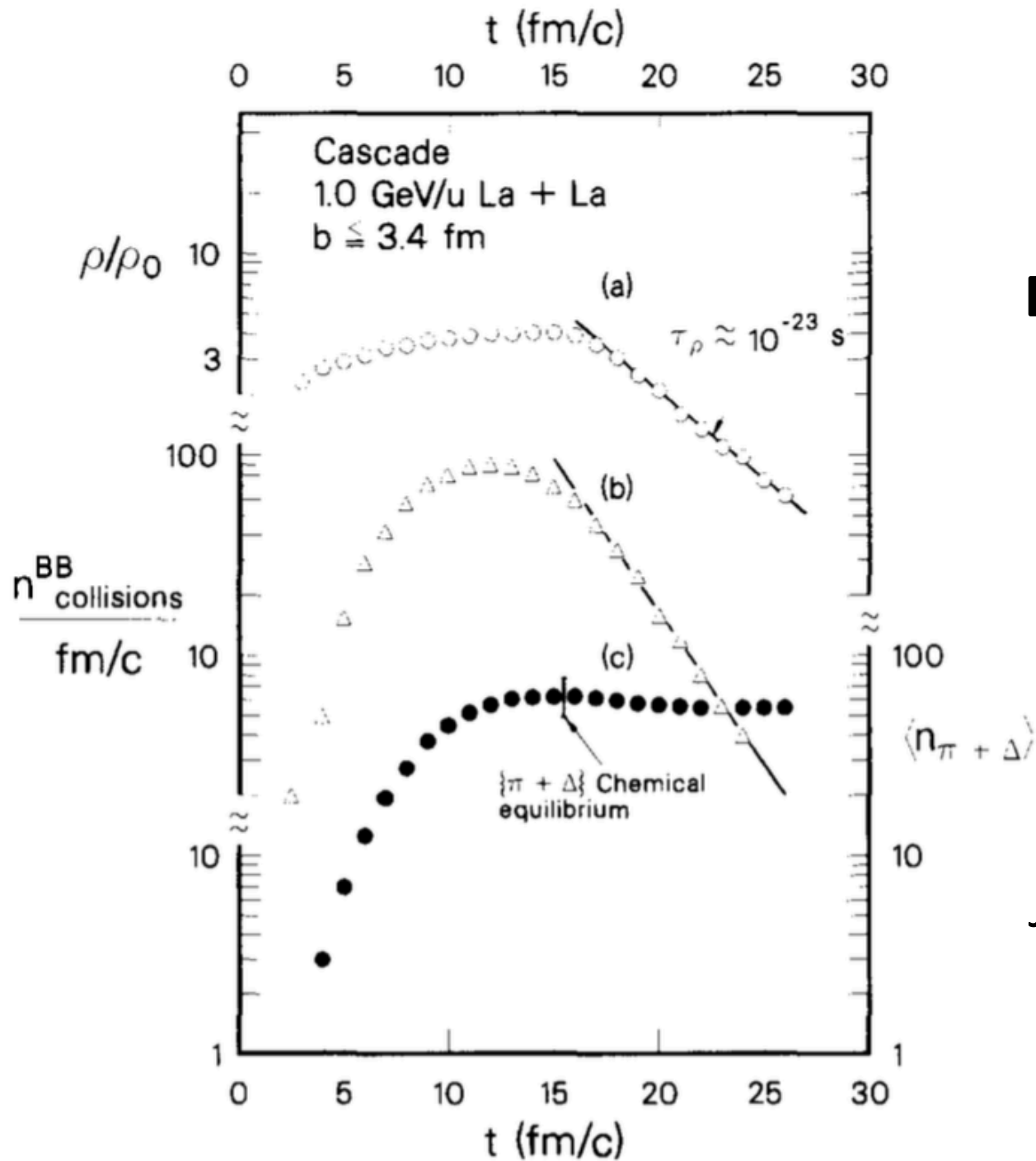


Phys. Rev. Lett. 45 (1980) 874



**The Goal:
EOS of Hadronic
Matter**

EOS Observables: sensitivity to high density stage



J. W. Harris & R. Stock
Notas de Fisica (1984)

Grazyna at Rio de Janeiro 1993



Grazyna at Rio de Janeiro 1997



Howel Pugh





VENUS

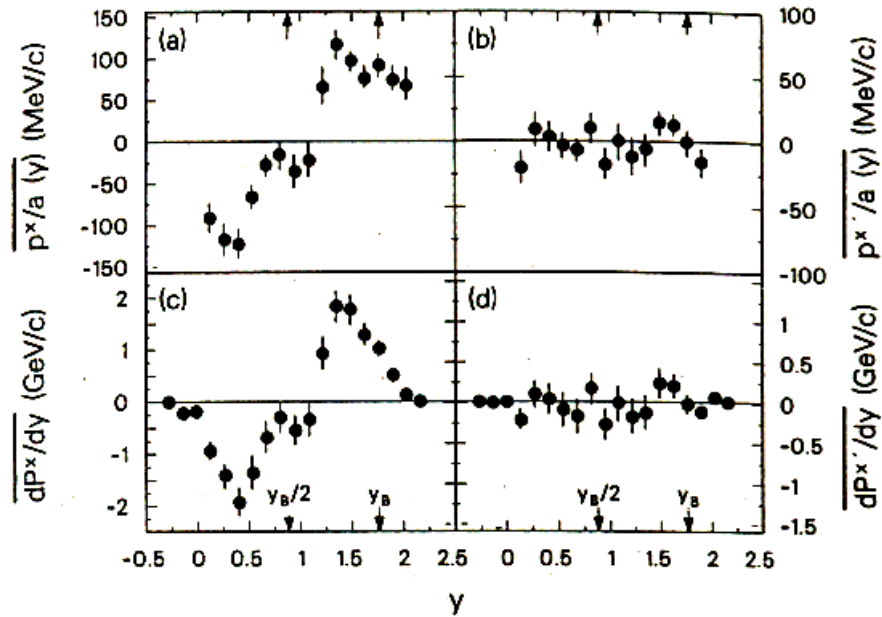
A PROPOSAL FOR A
RELATIVISTIC ION SYNCHROTRON & STORAGE RING

5 25 79

MBD WCH

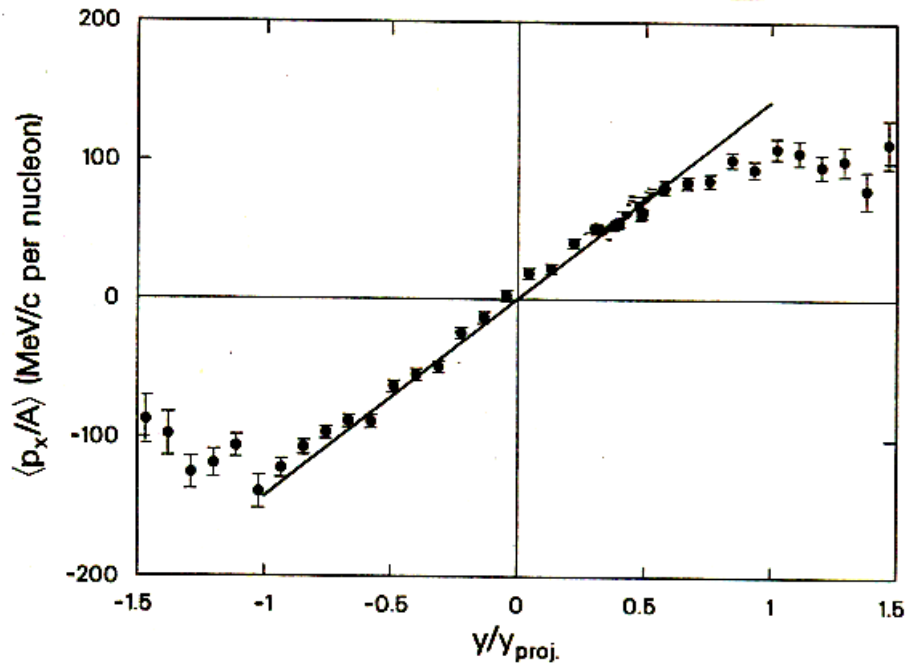
Transverse Momentum Balancing

P. Danielewicz and G. Odyniec, Phys. Lett 157B (1983) 146



Ar + KCl
Str.Ch.


EOS
Observables:
directed Flow
(v_1)



La + La
Pl. Ball

K.G.R. Doss et al
PRL 57 (1986) 302

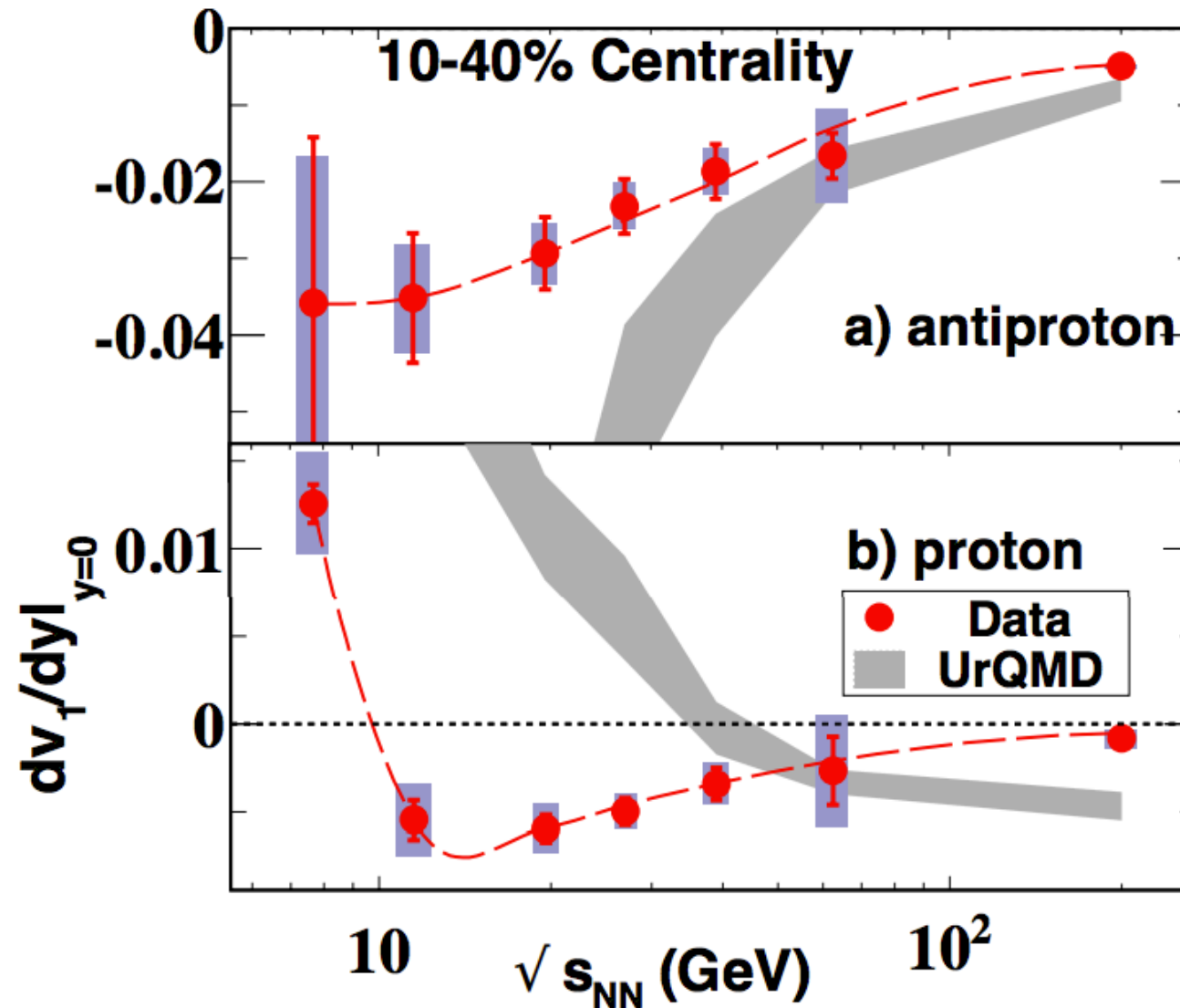
EOS from Sideflow Data

$$F(y) = P_{eff} \times S \times t_{pass}$$


EOS

A long story since 1989!

Softest Point of the EOS



H. Stöcker 2005

STAR
PRL 112 (2014) 162301

EOS in Neutron Stars

Oppenheimer, Volkoff, Tolman 1938

Hydrostatic equilibrium:

The inward pressure by gravity of each sphere-shell at radius R is counterbalanced by the EOS-compressional counter-force of the enclosed interior

Constraining the nuclear matter equation of state around twice saturation density

A. Le Fèvre^a, Y. Leifels^a, W. Reisdorf^a, J. Aichelin^b, Ch. Hartnack^b

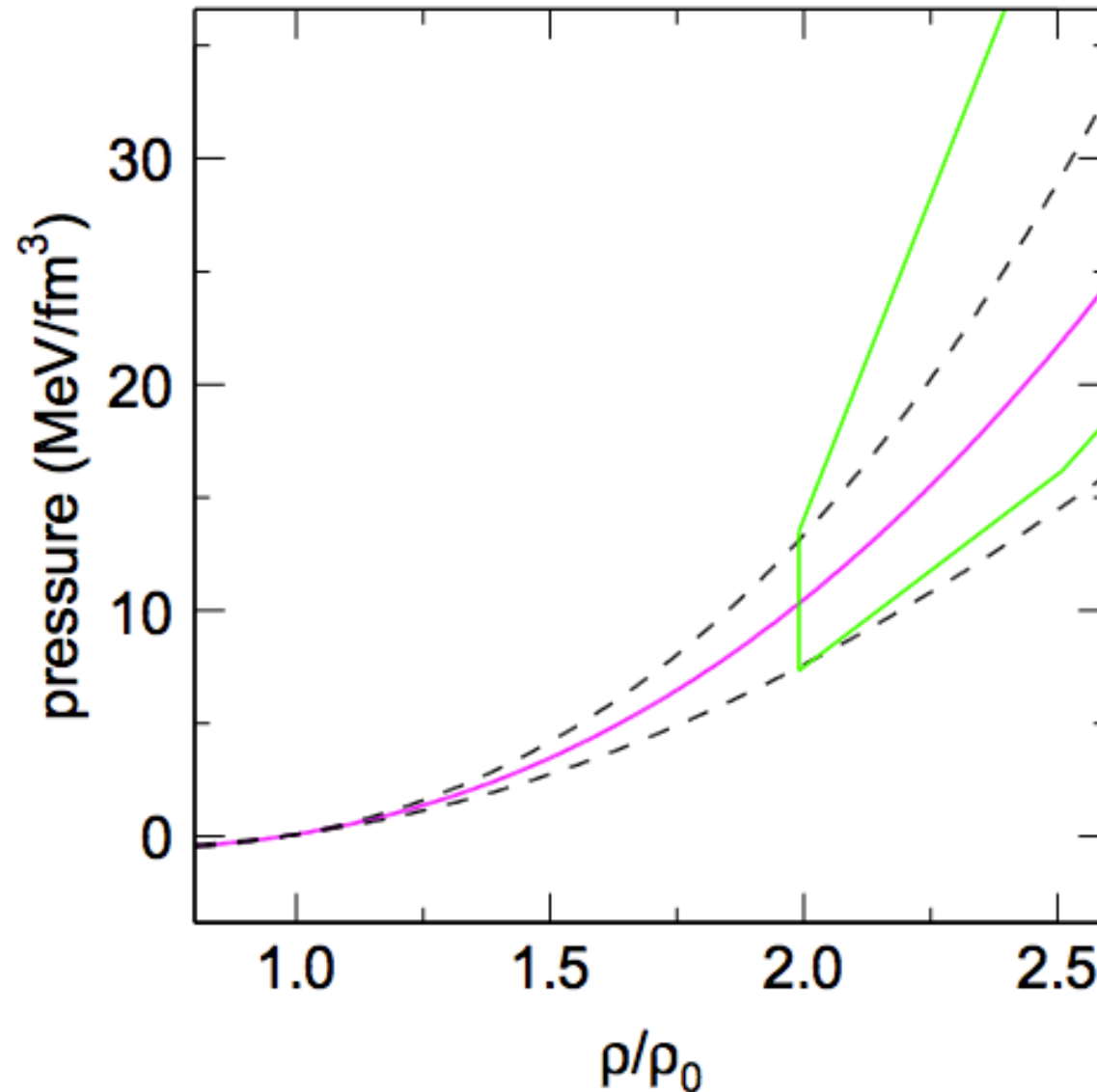
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Abstract

Using FOPI data on elliptic flow in Au+Au collisions between 0.4 and 1.5A GeV we extract constraints for the equation of state (EOS) of compressed symmetric nuclear matter using the transport code IQMD by introducing an observable describing the evolution of the size of the elliptic flow as a function of rapidity. This observable is sensitive to the nuclear EOS and a robust tool to constrain the compressibility of nuclear matter up to $2 \rho_0$.

EOS in Neutron Stars



A. Le Fèvre, Y. Leifels,
W. Reisdorf, J. Aichelin,
Ch. Hartnack

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