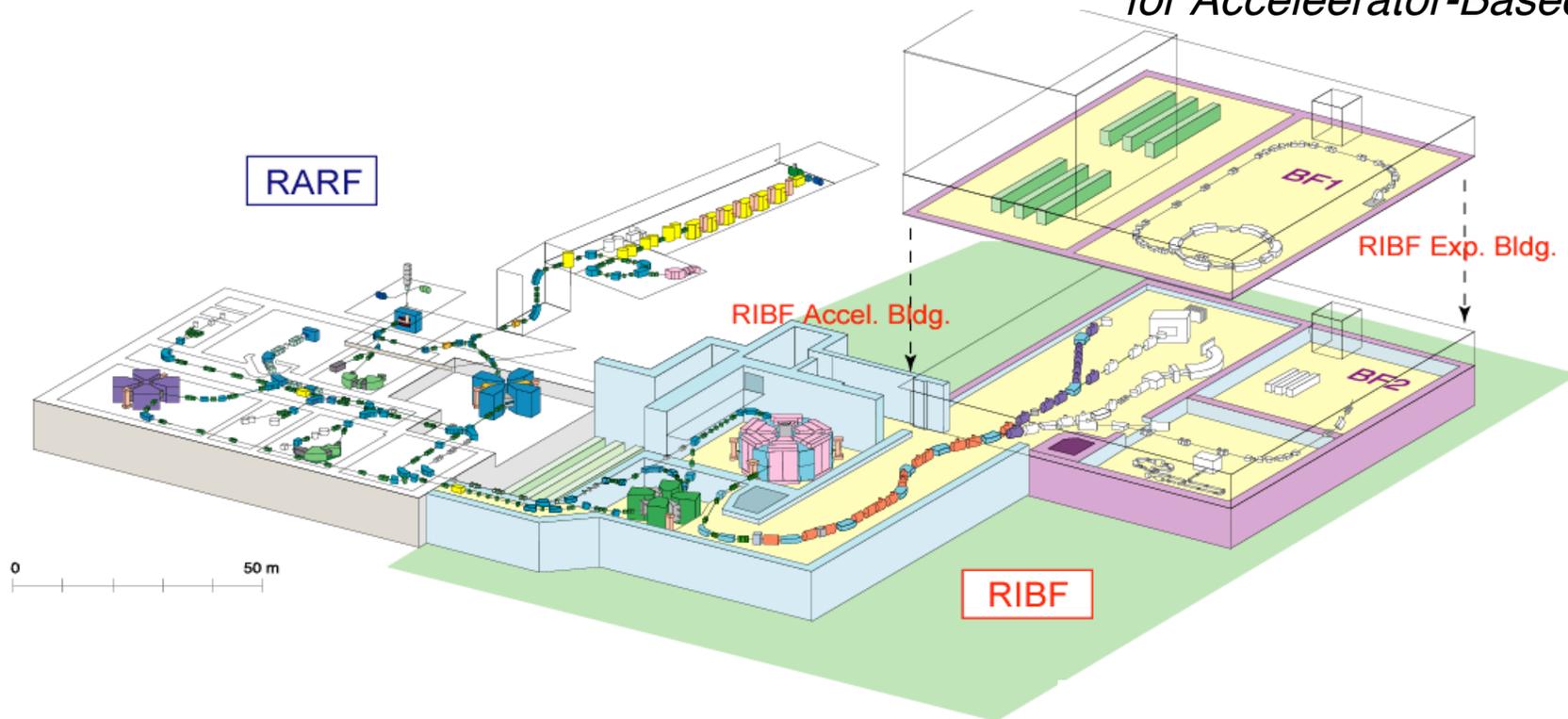


# Experimental determination of reaction rates *via* Coulomb dissociation

Tohru Motobayashi  
*RIKEN Nishina Center  
for Accelerator-Based Science*



# Coulomb dissociation (excitation) - radiative capture

# ANC method - radiative capture

# Trojan Horse method

<sup>6</sup>Li breakup  
Hammache

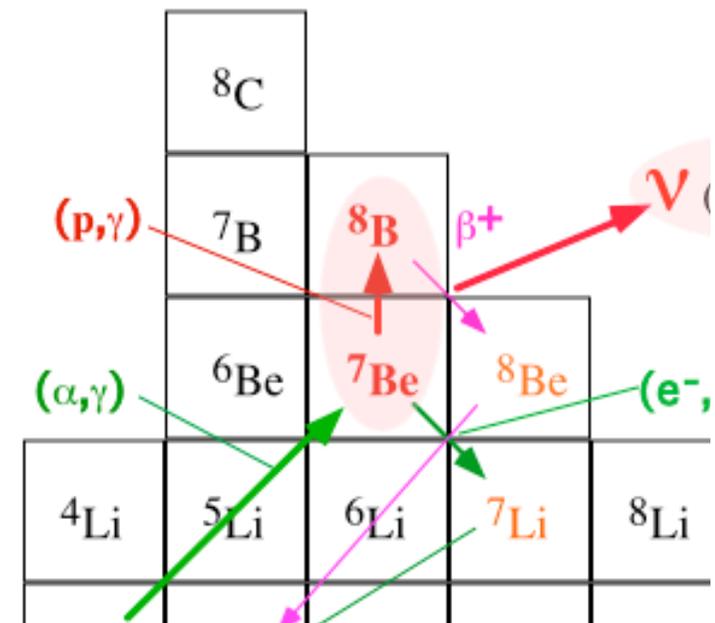
<sup>7</sup>Be(p,γ)<sup>8</sup>B reaction - a “bench-mark”, continuum, E1+E2

pp chain / solar neutrino

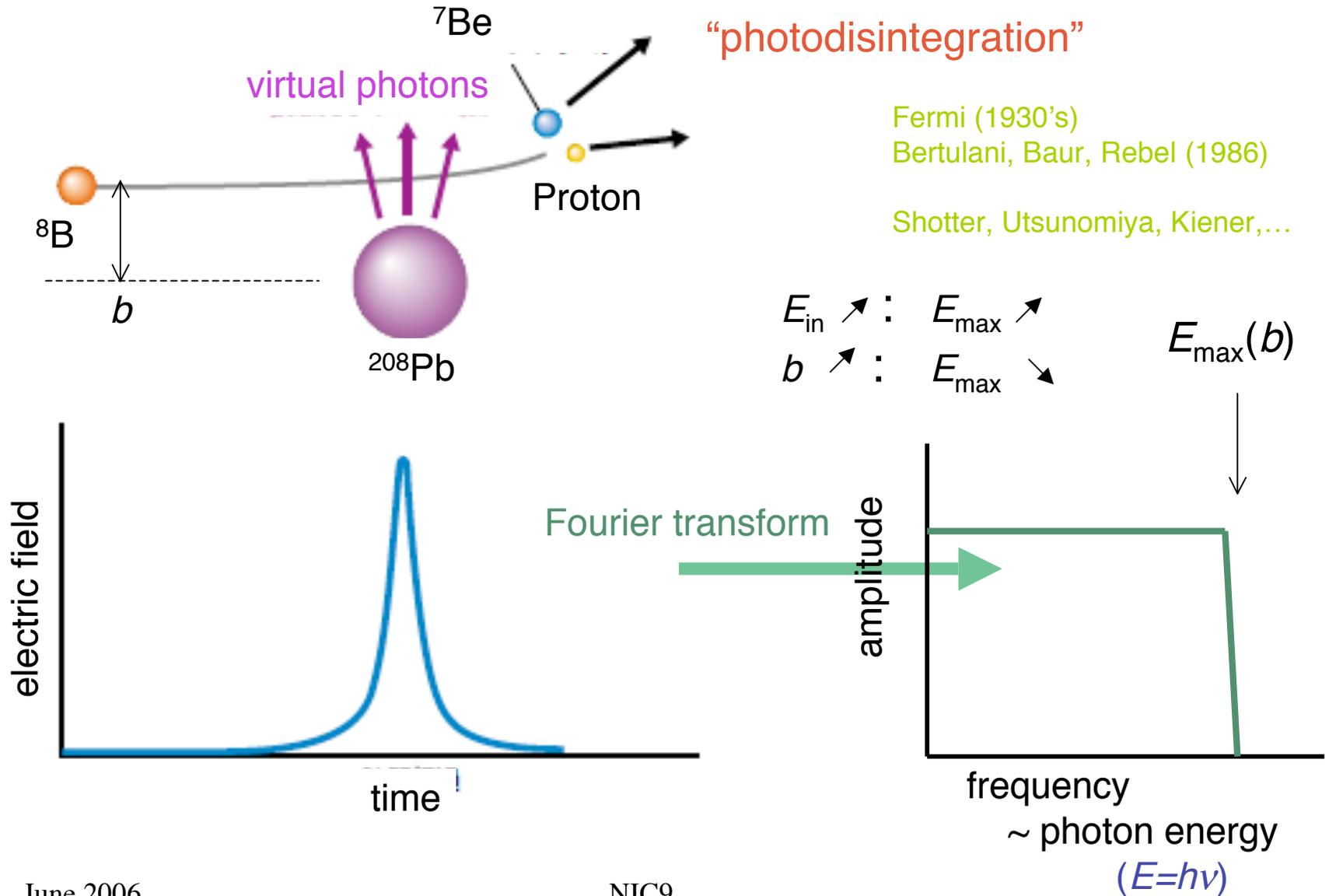
<sup>23</sup>Al breakup resonance, E2

rp process

RI Beam Factory



# electric field “felt” by the projectile





↓ virtual photon theory or DWBA



↓ detailed balance



Large yield

detailed balance

$$\sigma_{(\gamma, \text{p})} = \frac{(2j_7 + 1)(2j_1 + 1)}{2(2j_8 + 1)} \frac{k_{17}^2}{k_\gamma^2} \sigma_{(\text{p}, \gamma)} \quad 100 \sim 1000$$

virtual photon number (intermediate energy)

$$\left( \frac{d\sigma}{dE_\gamma} \right)_{\text{C.D.}} = \frac{n}{E_\gamma} \sigma_{(\gamma, \text{p})} \quad 100 \sim 1000$$

thick target

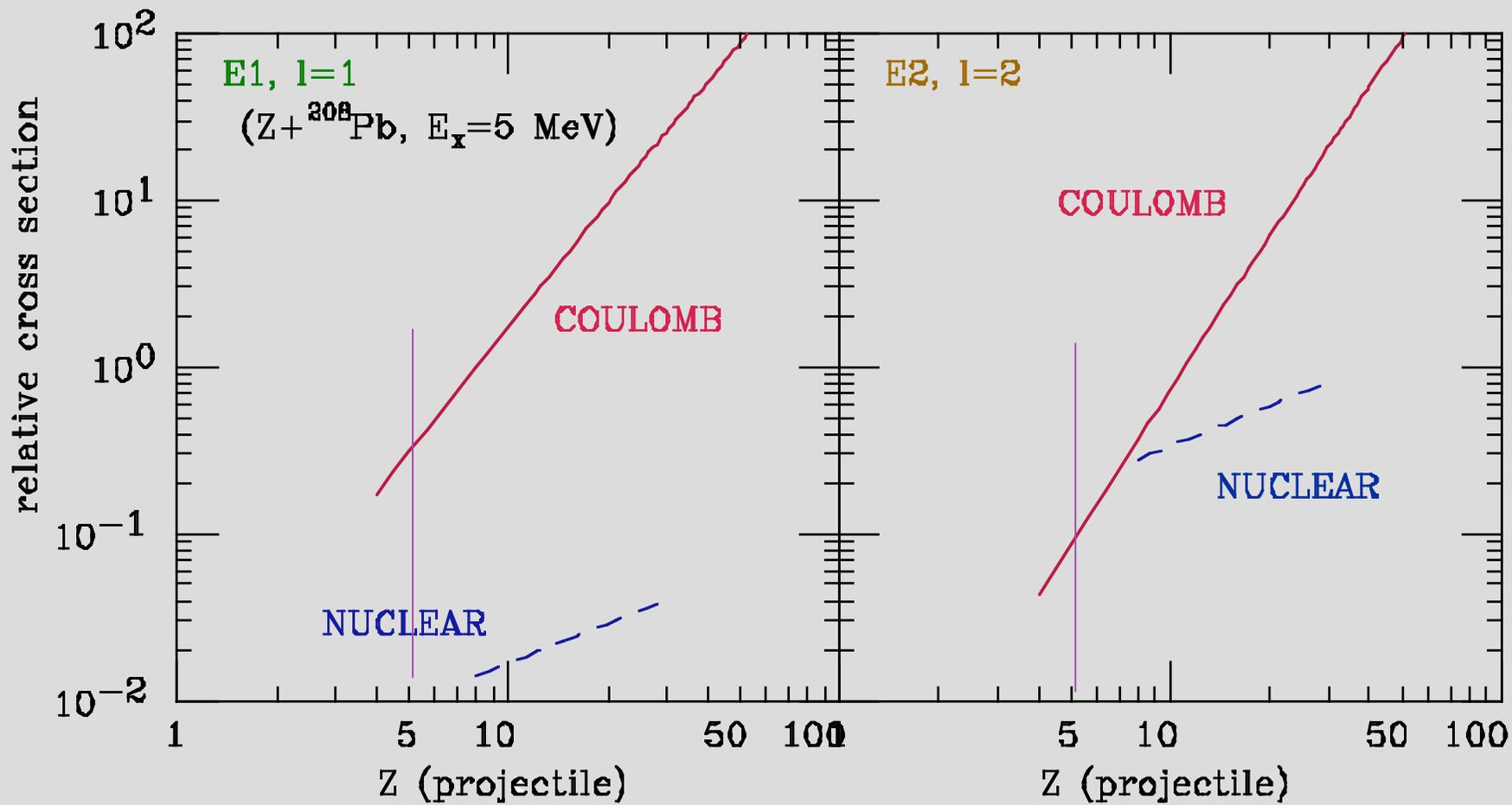
charged particle detection

experiments with R.I. beams

but

indirect *i.e.* nucl. force / higher order / E2

$l = 1$   $E_{in} = 50 \text{ AMeV}$   $l = 2$

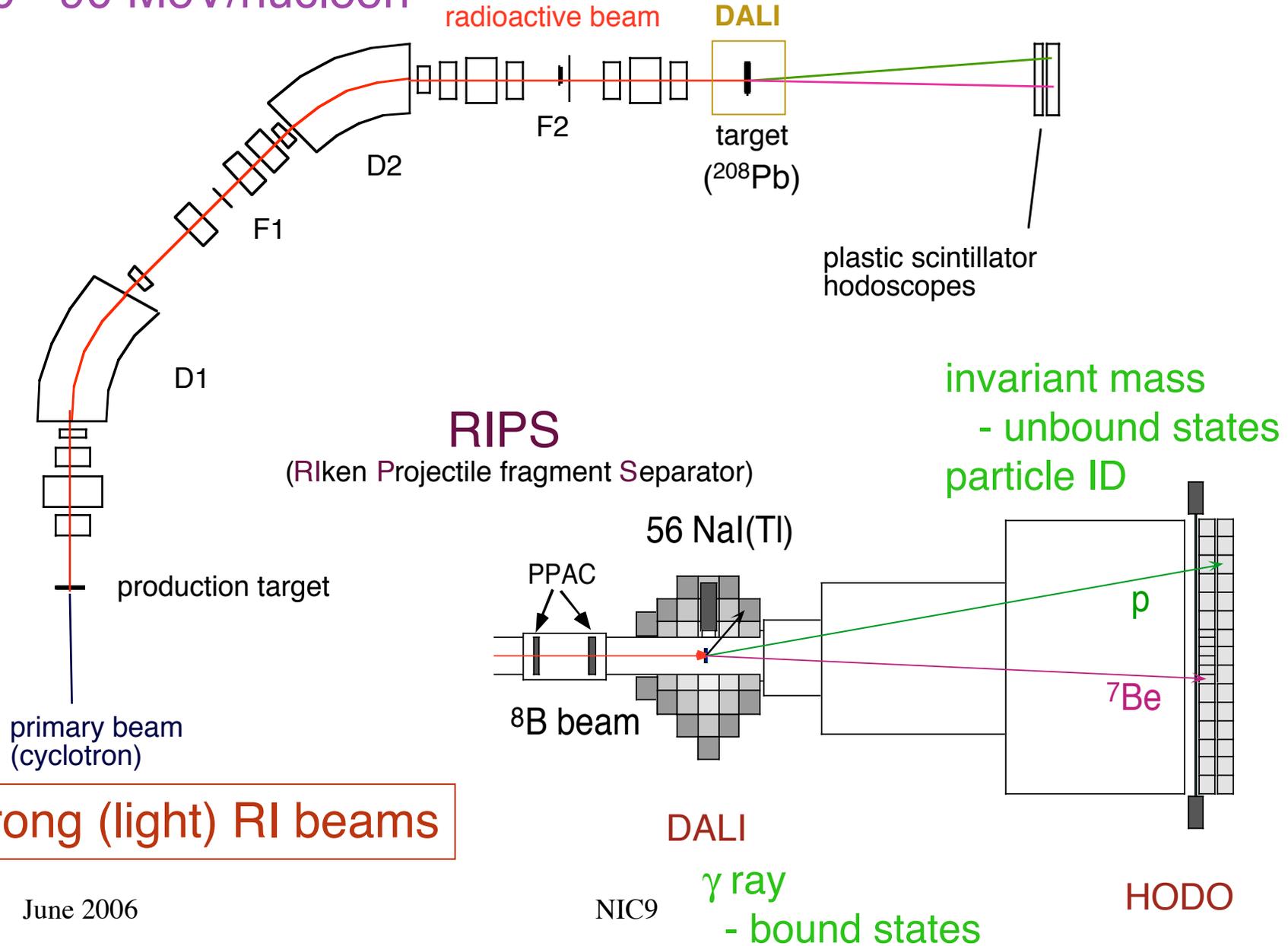


$C > N$

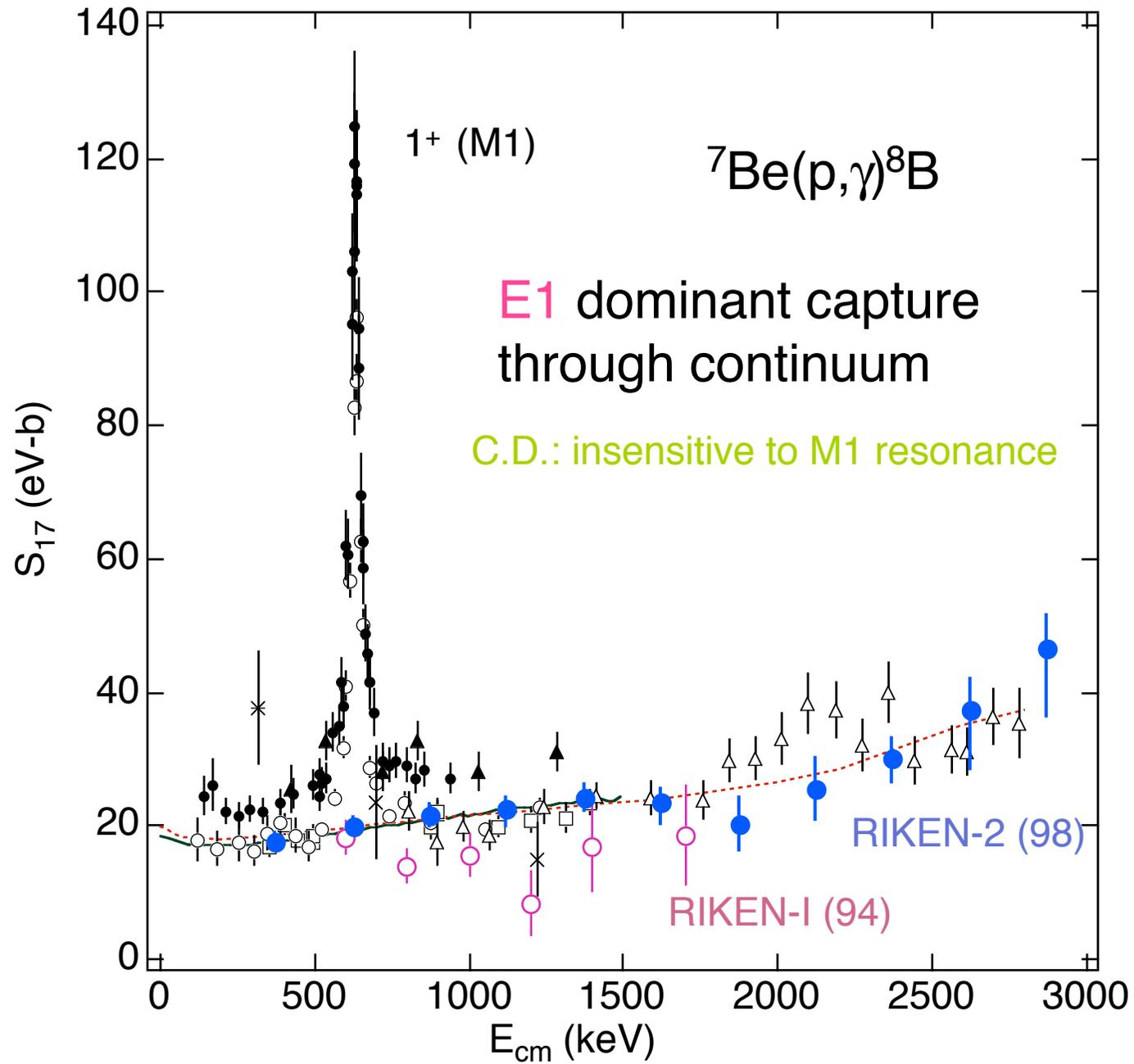
$C \sim N$  (low  $Z$ )

# spectroscopy of unstable nuclei

50 - 90 MeV/nucleon



June 2006

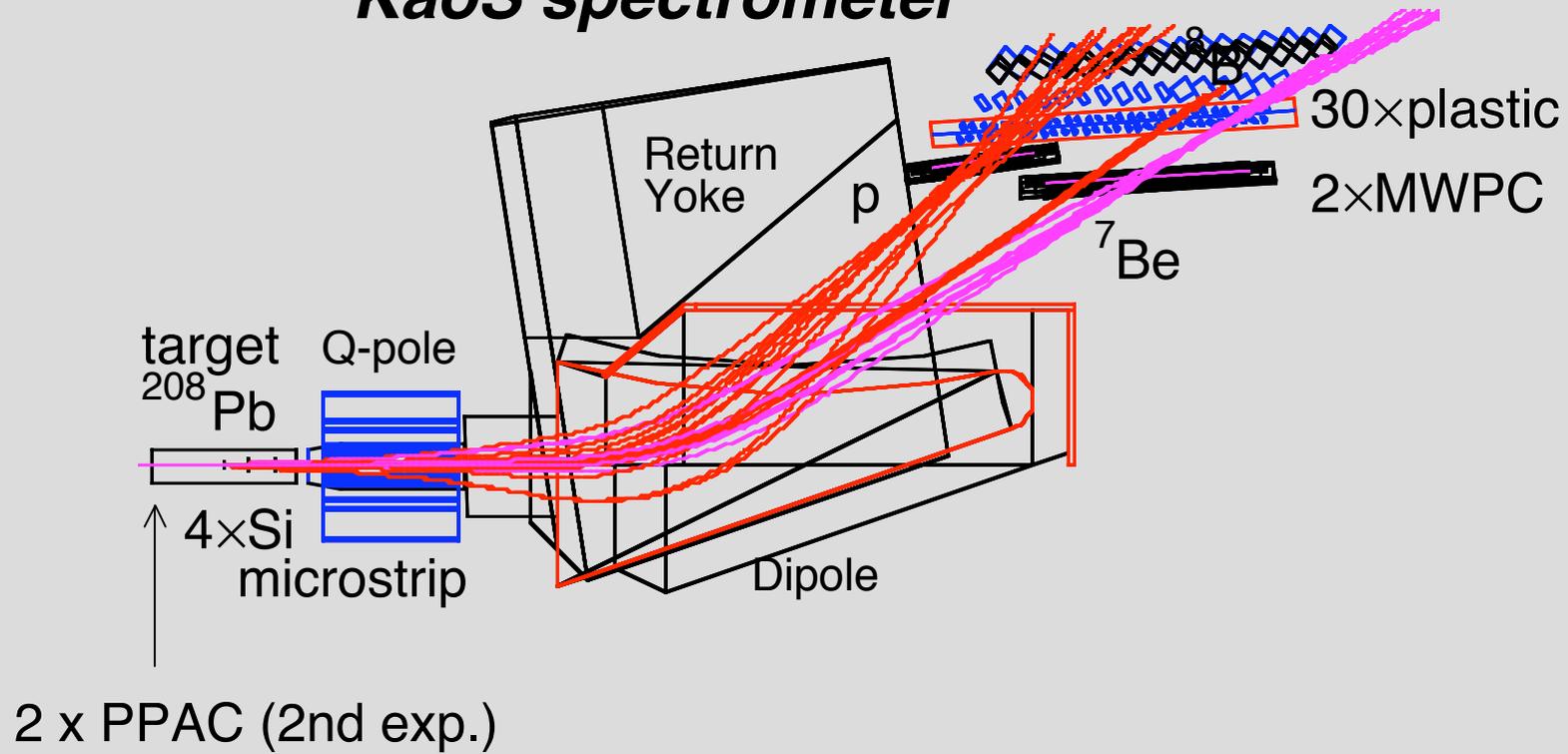


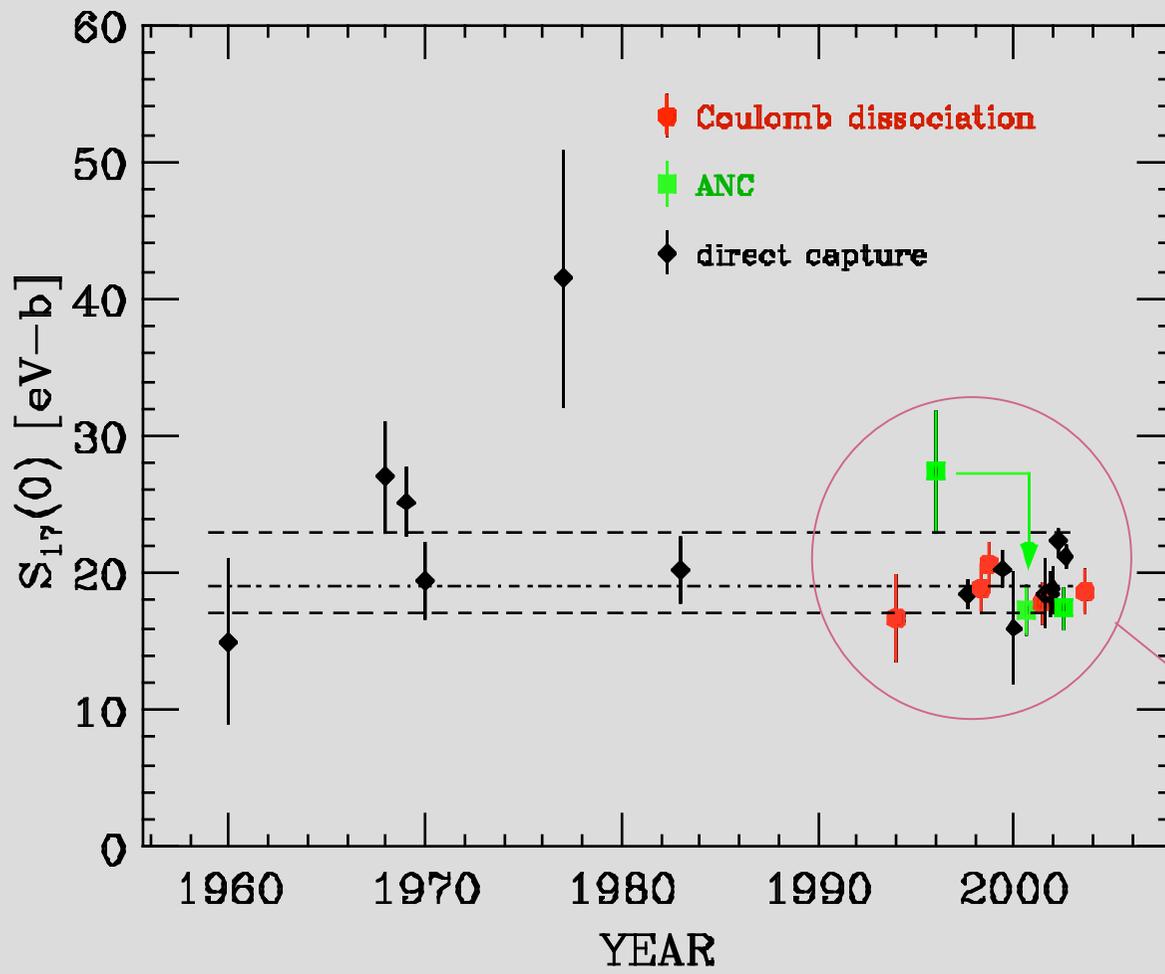
June 2006

NIC9

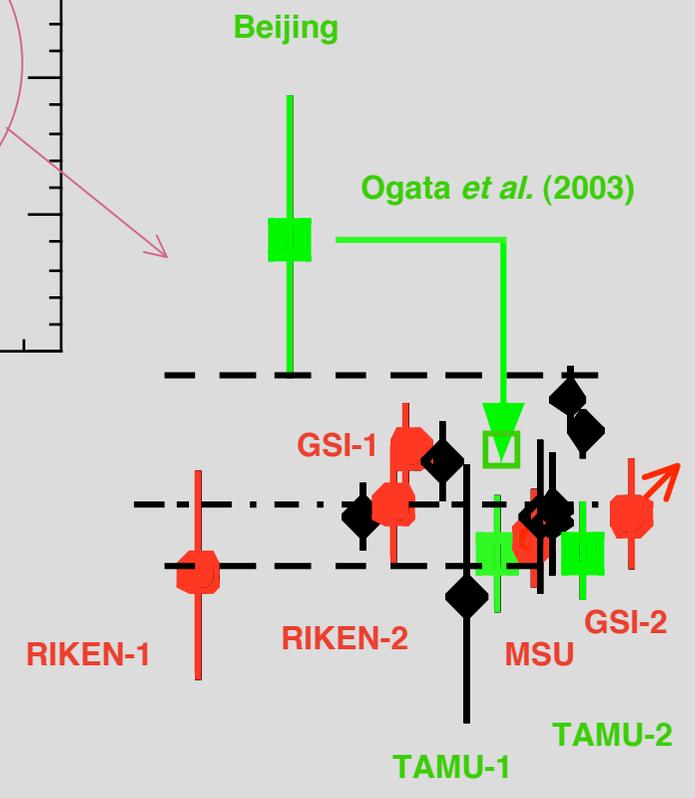
@ GSI 254 MeV/nucleon

## *KaoS spectrometer*

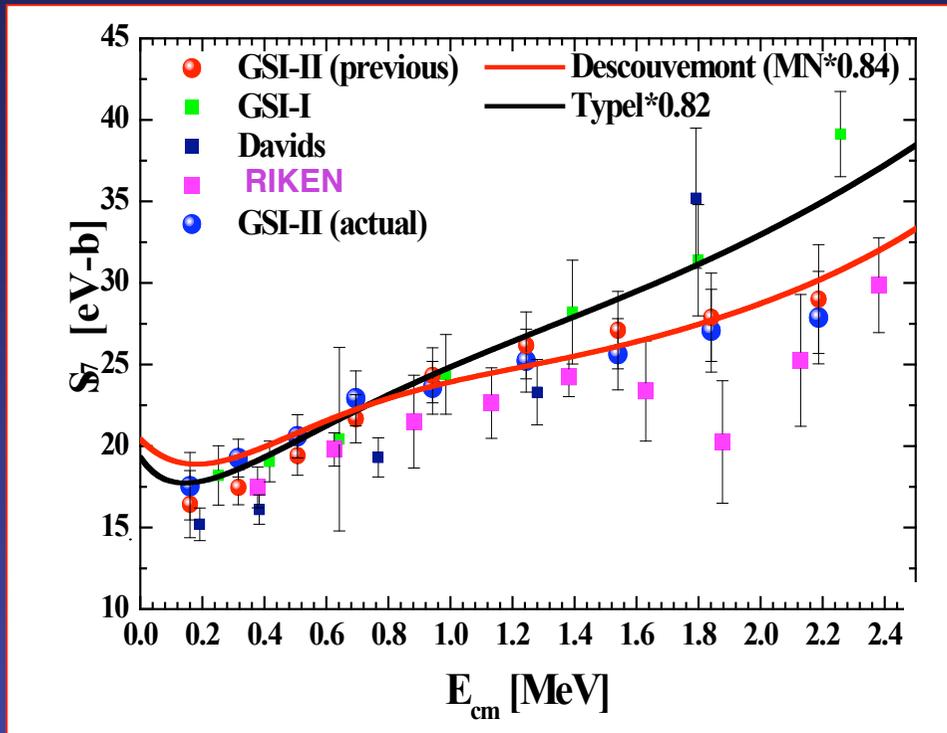




# $S_{17}$ at $E=0$



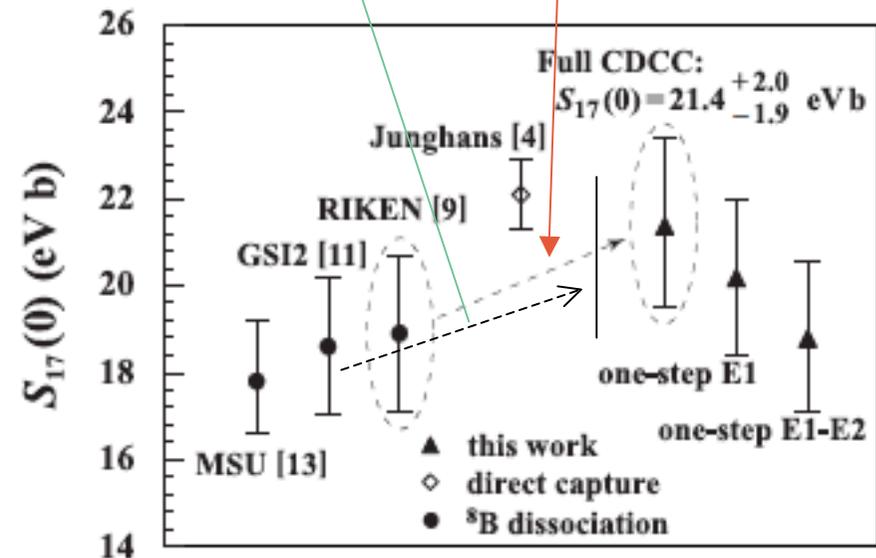
# $S_{17}$ - Factor



Possible ambiguities  
In optical potentials  
=> Bishop (poster 107)

E1/E2/nucl. Interference (CDCC)  
Ogata *et al.* PRC, 2005

Schumann *et al.*  
P. R. C 2005



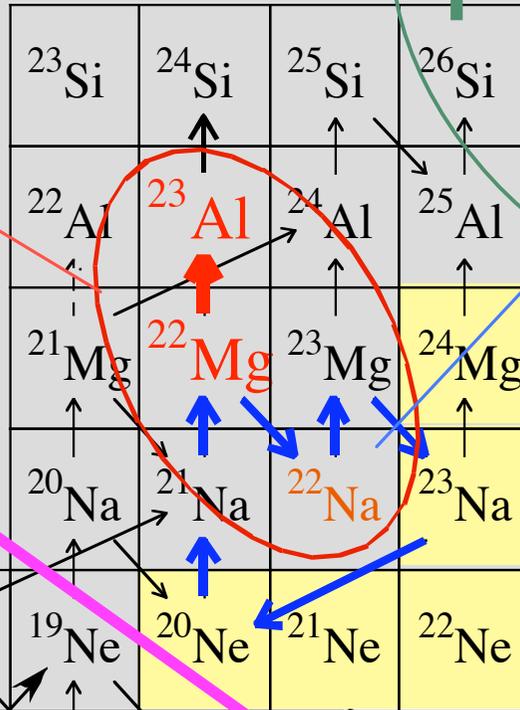
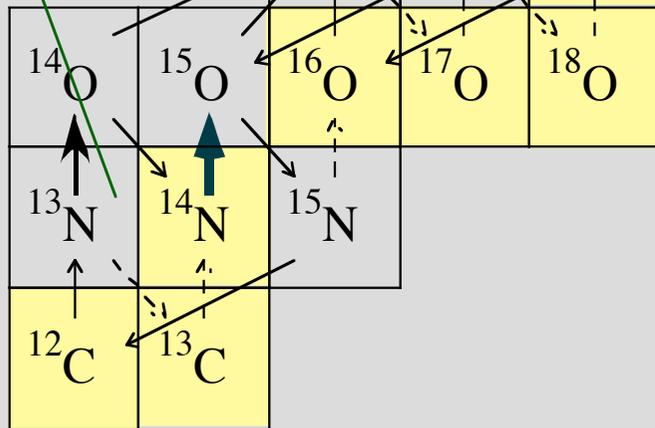
BreakOut of  
Hot-NeNa Cycle ?

$^{22}\text{Na}$  production

*Gomi et al.: poster*

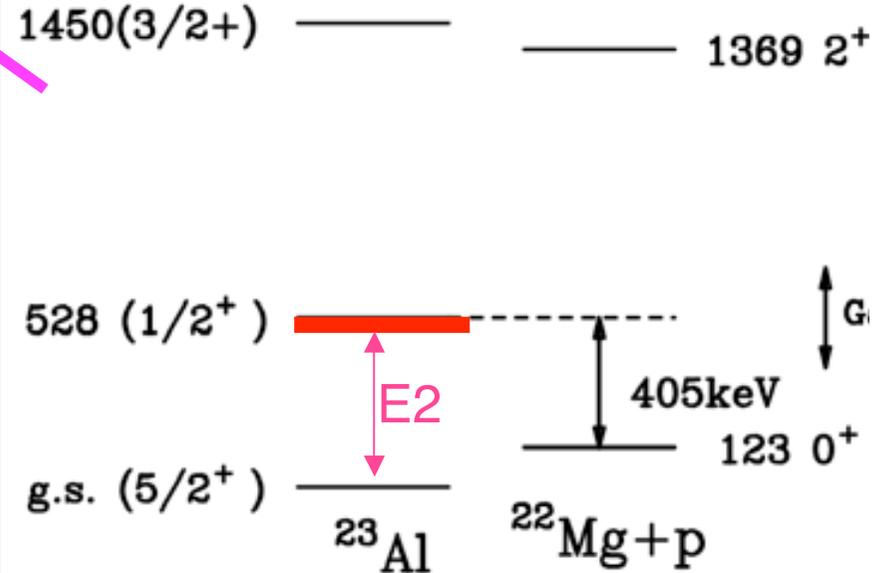
rp-process

(hot) CNO cycle

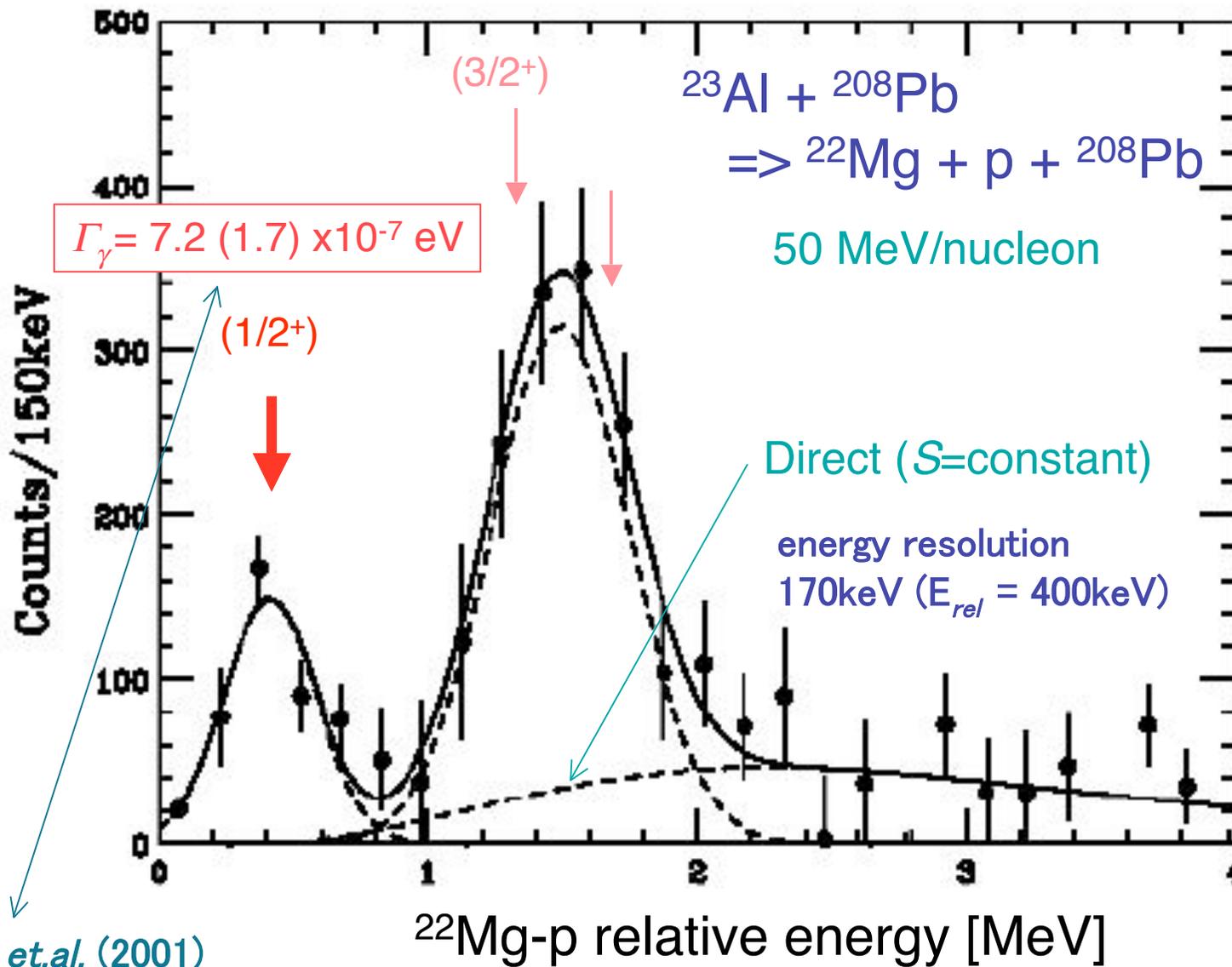


hot-NeNa cycle

Stable nuclei



$E_{rel}$  spectrum -  $10^4$  pps  $^{23}\text{Al}$   $\Leftrightarrow 10^{12}$  pps  $^{22}\text{Mg} ! + ^1\text{H}$



Caggiano *et.al.* (2001)  
 $5.49 \times 10^{-7}$  eV

Gomi et al., Nucl. Phys. A734 (04) E77

# Stellar reactions studied by Coulomb dissociation using RI beams

## Steady burning

pp chain  
(solar neutrino)  
CNO cycle



GSI (254A MeV)

RIKEN, MSU (50~80A MeV)

Direct, A NC

Notre Dame (3A MeV)



RIKEN (100A MeV)

D

(Coulomb excitation, sub-threshold state)

## Explosive burning

hot CNO cycle



RIKEN (78A MeV)

D



RIKEN (88A MeV), GANIL (70A MeV)

D

hot pp mode



RIKEN (70A MeV)

A



GANIL, RIKEN (70A MeV)

A



RIKEN (84A MeV)

A

rp-process



RIKEN (50A MeV)

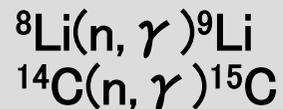


RIKEN (50A MeV)

Poster 163

Neutron  
capture

inhomogeneous  
BBN  
r-process



MSU (40A MeV)

GSI (605A MeV), RIKEN (70A MeV),

D

MSU (35A MeV)



RIKEN (67A MeV)

# RIKEN **R**adioactive **I**sotope **B**eam **F**actory **RIBF**

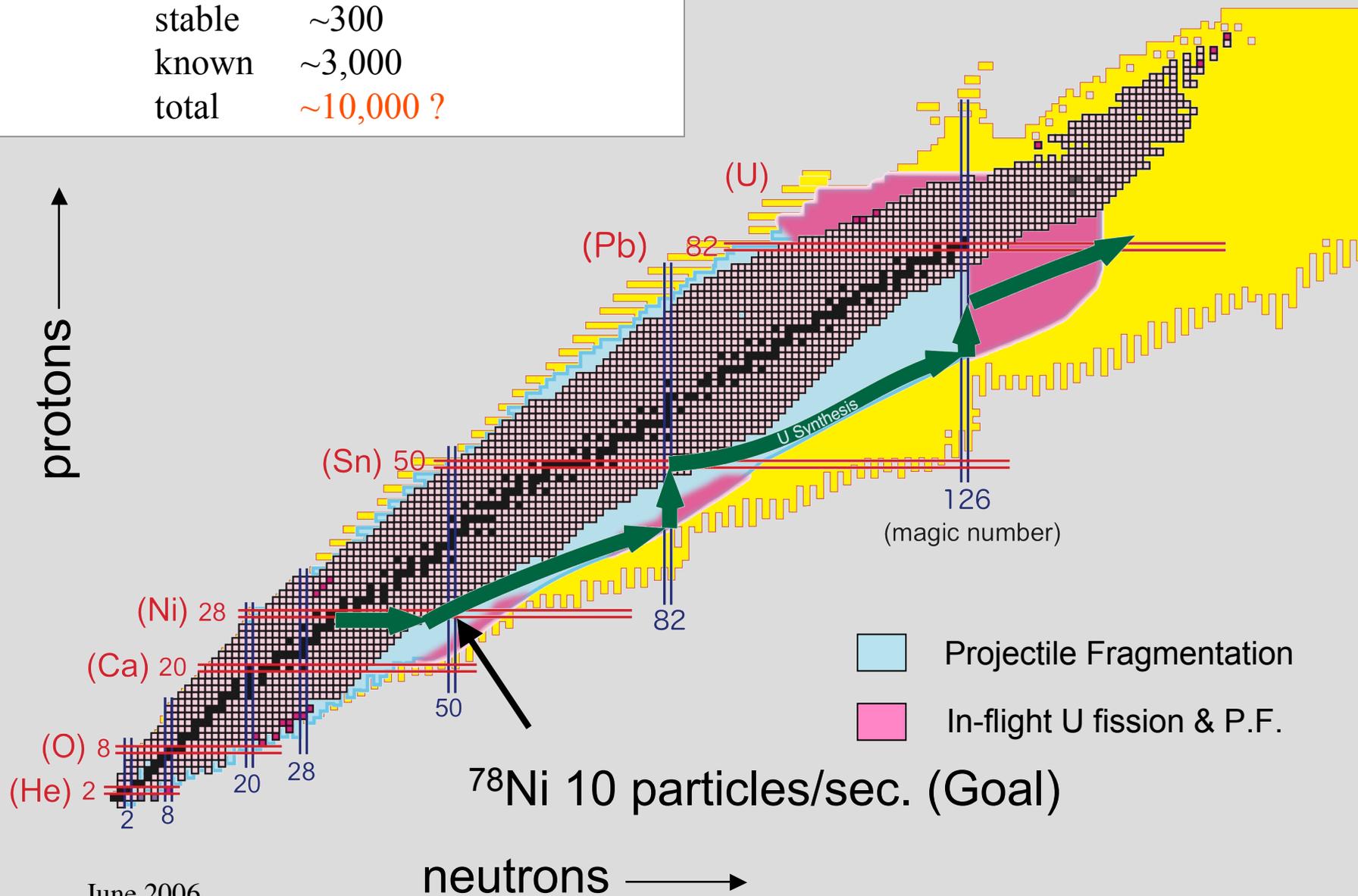
1st realization of next-generation RIB facilities  
\* RIA, GSI - FAIR

Intense primary beams (1 pμA up to U, 350 AMeV)  
=> farther from the stability valley  
by fragmentation and fission

1st beam: in 2006, experiments: 2007-

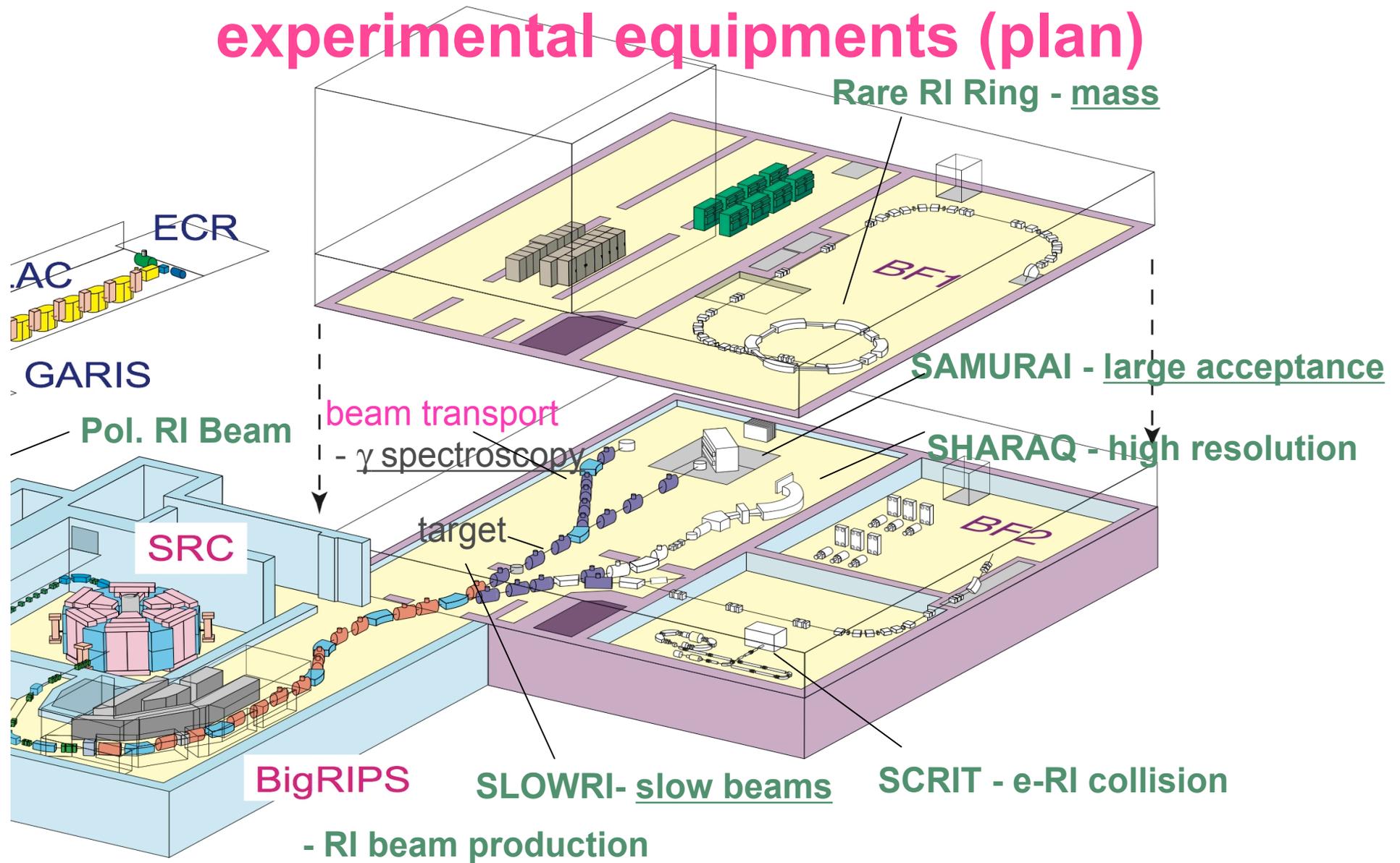
# Atomic nuclei (isotopes)

stable	~300
known	~3,000
total	~10,000 ?



June 2006

# RIKEN RIBF experimental equipments (plan)



## Summary

C.D. for astrophysical  $\sigma(p,\gamma)/\sigma(n,\gamma)$  on unstable nuclei

continuum

resonance: E1, E2, M1

RI-beam

Questions - control of reaction mechanism

higher order / E2-related corrections /  
nuclear excitations

theory  
experiment

Future: more  $(p,\gamma)$ ,  $(n,\gamma)$  /  $(2p,\gamma)$ ,  $(\alpha,\gamma)$  ....

Intense beam  
efficient setup

RIBF - 1st beam: 2006

covering the "r-process path"

*c.f.* Motizuki (poster 136)

**RIBF International Users Meeting:** Aug. 3, 4 (2006)

**Visit** the users group page:

[http://ribfwww.riken.go.jp/exp/RIBF\\_uec\\_eng/](http://ribfwww.riken.go.jp/exp/RIBF_uec_eng/)  
registration: [ribfusr\\_meeting06@rarf.riken.jp](mailto:ribfusr_meeting06@rarf.riken.jp)

**PAC:** end of 2006 or early 2007

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23rd INPC (INPC07): Tokyo, June 3-8, 2007

DREB (Direct Reaction with Exotic Beams) as a pre sympo.