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Nuclear astrophysics with gamma-ray line observations

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Gamma-ray spectrometers with high spectral resolution are operating in space since 2002: RHESSI and SPI on INTEGRAL. Understanding the instrumental response and backgrounds is major effort but a prerequisite of detailed science interpretations. While ^{44}Ti from core collapse supernovae could not be detected, this still adds constraints to Cas A ^{44}Ti ejection. Diffuse nucleosynthesis is studied through ^{26}Al , ^{60}Fe , and positron annihilation gamma-ray measurements. With SPI on INTEGRAL, the gamma-ray line from decay of radioactive ^{26}Al could be measured at unprecedented spectroscopic precision. This made possible a new determination of the total mass of ^{26}Al produced by stellar sources throughout the Galaxy, and an analysis of the properties of the interstellar medium around ^{26}Al sources. ^{60}Fe is clearly detected with SPI, its intensity ratio to ^{26}Al is confirmed to be on the lower side of theoretical predictions. Nucleosynthesis sources are probably minor contributors to Galactic positrons, as deduced from the bulge-centered spatial distribution of the annihilation gamma-ray emission.

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