

Probing hidden sectors via the $U(1)$ portal

Adam Ritz

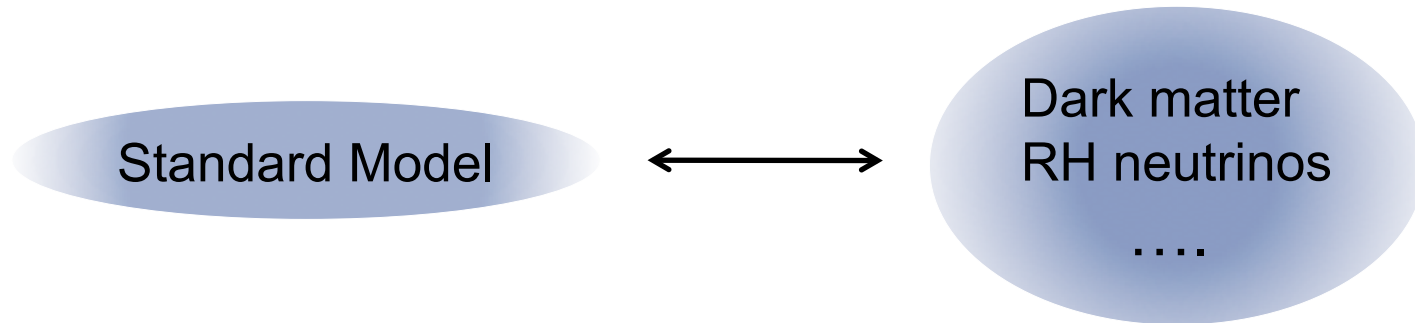
University of Victoria



With B. Batell and M. Pospelov,
0903.0363, 0906.5614

New physics in a hidden sector

Empirical evidence for new physics does not always point to the EW scale and above, but rather to a hidden sector that may also contain “light” states



Any “light” states should be neutral under SM gauge group

In particular, there are motivations for exploring dark matter as part of a larger (interacting, multi-component) hidden sector:

- a more generic setting to explore experimental sensitivity
- allows for possible enhancement/suppression, that could explain existing (astrophys) anomalies: DAMA, PAMELA, etc.

Probing a hidden (dark) sector

Standard Model



Hidden Sector

$$\mathcal{L}_{med} = \sum_{n,k,l}^{n=k+l-4} \frac{O_k^{(SM)} O_l^{(med)}}{\Lambda^n}$$

Generic interactions are irrelevant (dimension > 4), but there are three renormalizable “portals”

- Vector portal: $\mathcal{L} = -\frac{\kappa}{2} V^{\mu\nu} B_{\mu\nu}$
- Higgs portal: $\mathcal{L} = (-\lambda S^2 + \xi S) H^\dagger H$
- Neutrino portal: $\mathcal{L} = -y_{ij} \bar{L}_i H N_j$

Probing a hidden (dark) sector

Standard Model



Hidden Sector

$$\mathcal{L}_{med} = \sum_{n,k,l}^{n=k+l-4} \frac{O_k^{(SM)} O_l^{(med)}}{\Lambda^n}$$

Generic interactions are irrelevant (dimension > 4), but there are three renormalizable “portals”

- Vector portal: $\mathcal{L} = -\frac{\kappa}{2} V^{\mu\nu} B_{\mu\nu}$
- Higgs portal: $\mathcal{L} = (-\lambda S^2 + \xi S) H^\dagger H$
- Neutrino portal: $\mathcal{L} = -y_{ij} \bar{L}_i H N_j$

NB: The vector mediator V can naturally be light ($M \ll M_Z$)

A secluded U(1)

$$\mathcal{L} = -\frac{1}{4}V_{\mu\nu}^2 - \frac{\kappa}{2}V_{\mu\nu}F^{\mu\nu} + |D_\mu\phi|^2 - V(\phi) \quad [\text{Holdom '86}]$$

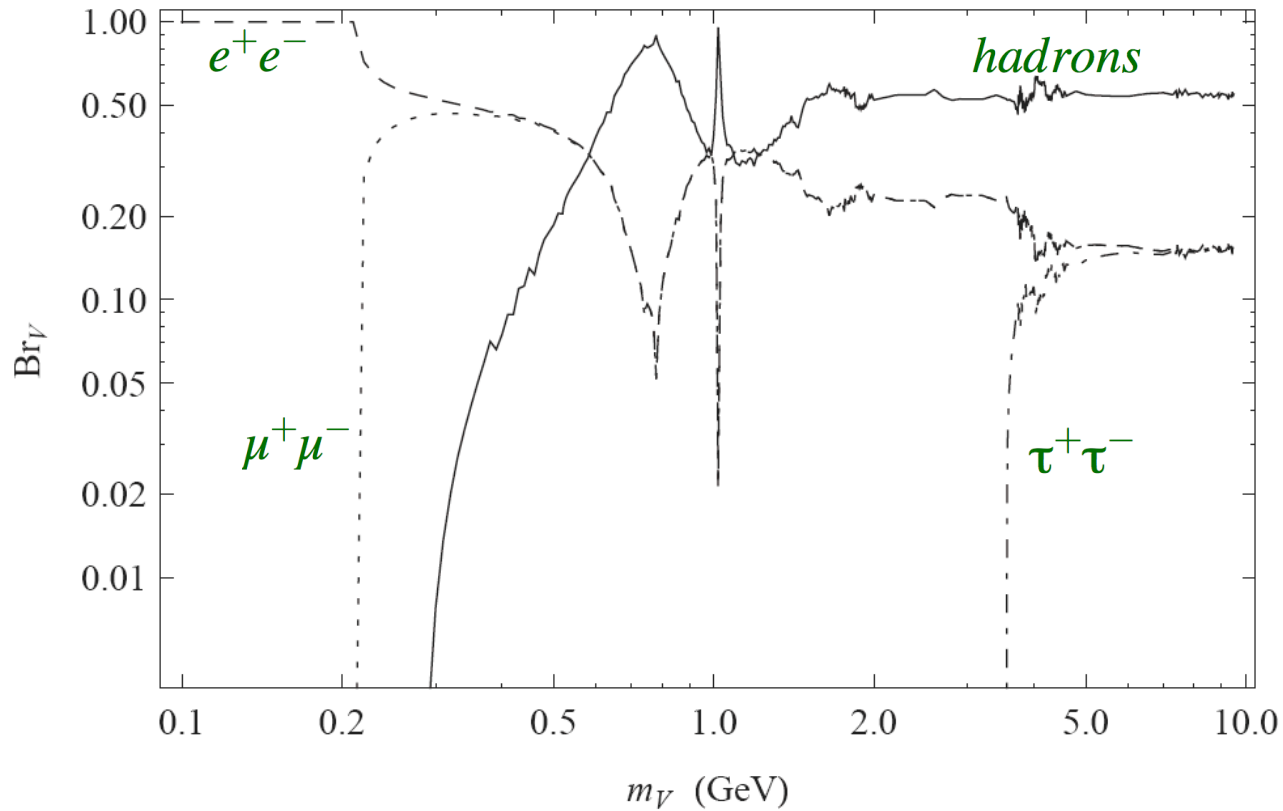


$$\mathcal{L}_{int} = -\kappa e V_\mu J_{em}^\mu + \frac{m_V^2}{v'} h' V_\mu V^\mu$$

V - production through mixing with EM current: $O(\kappa^2)$ h' - production through higgs'strahlung: $O(\kappa^2)$

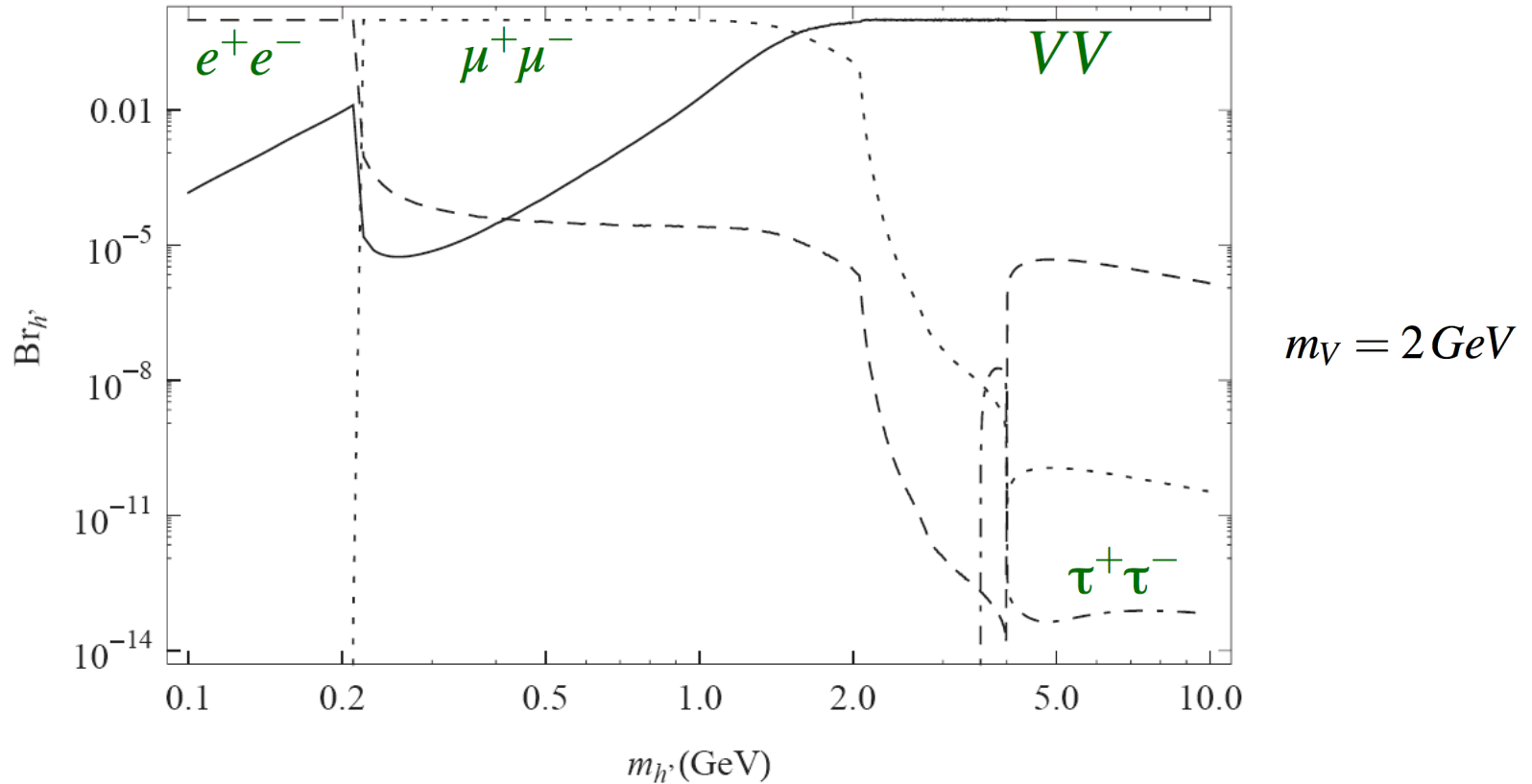
- Weak-scale states charged under $U(1)_S$ are WIMP dark matter candidates [Pospelov, AR, Voloshin '07], which can have enhanced galactic annihilation rates if V is light [Arkani-Hamed et al '08, Pospelov & AR '08]
- If kinetic mixing arises from integrating out heavy charged states at 1-loop $\Rightarrow \kappa \sim 10^{-3}$, SUSY D-terms then imply $m_V \sim O(\text{GeV})$ [Arkani-Hamed & Weiner '08, Baumgart et al '09; Cheung et al '09; Katz & Sundrum '09]
- We will take the secluded U(1) coupling $\alpha' = \alpha$, so the parameter space = $\{m_V, m_{h'}, \kappa\}$

Vector decays



Parametrically: $\Gamma_V(V \rightarrow 2l) \sim O(\kappa^2)$

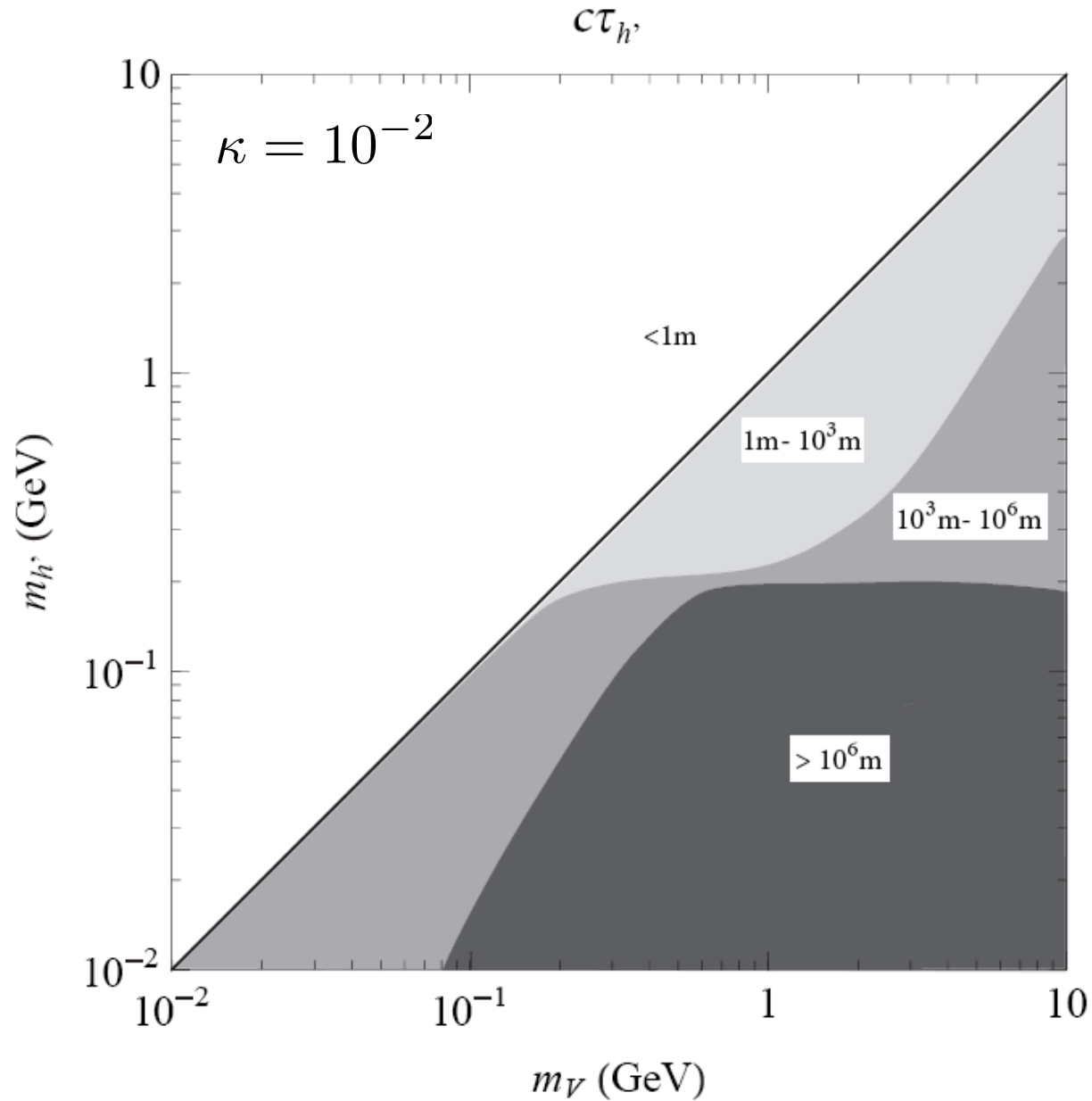
Higgs' decays



When the higgs' is light ($m_{h'} < m_{\nu}$), it is parametrically long-lived:

$$\Gamma_{h'}(h' \rightarrow 2l) \sim O(\kappa^4)$$

Higgs' decays



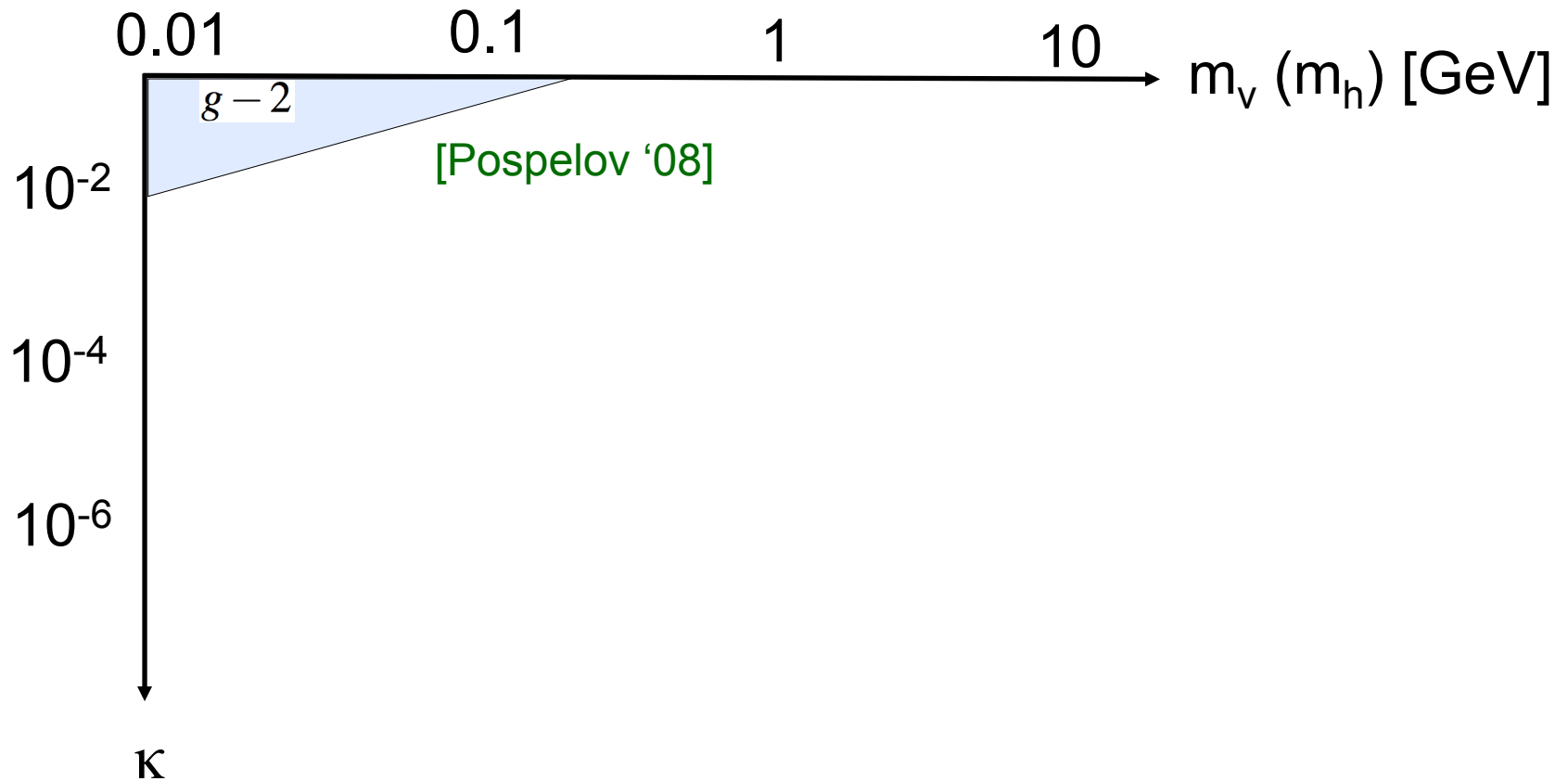
Experimental Sensitivity

Luminosity matters!

- Fixed targets (proton & electron beams dumps)
 - up to 10^{23} POT for modern neutrino sources
 - sensitive to long-lived states
- Medium energy colliders (BaBar, Belle)
 - large datasets $\sim O(1000) \text{ fb}^{-1}$
 - $\sqrt{s} \sim 10 \text{ GeV}$
- Rare decays
 - high statistics e.g. for kaon decays

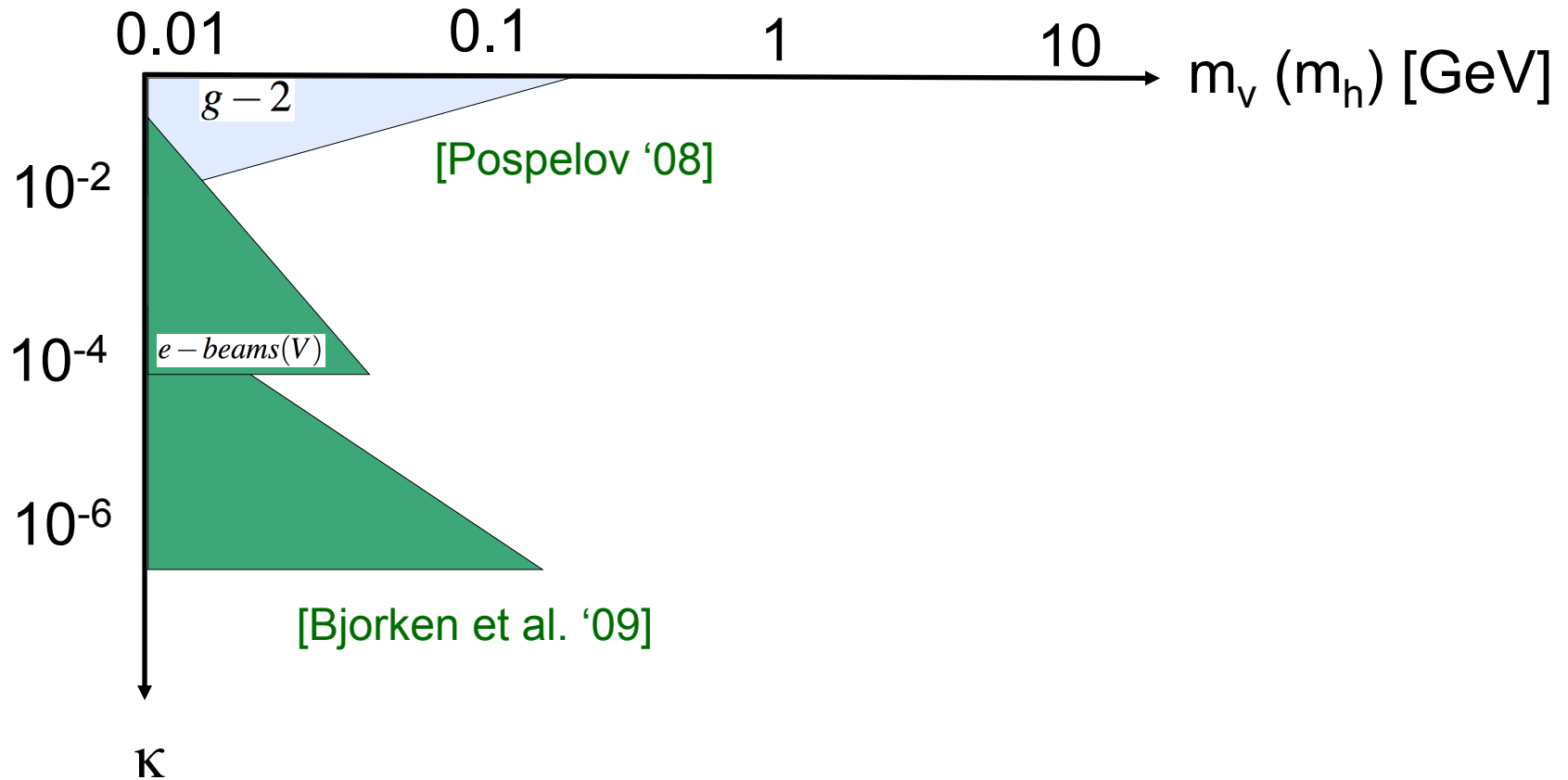
Experimental Sensitivity

In pictorial form



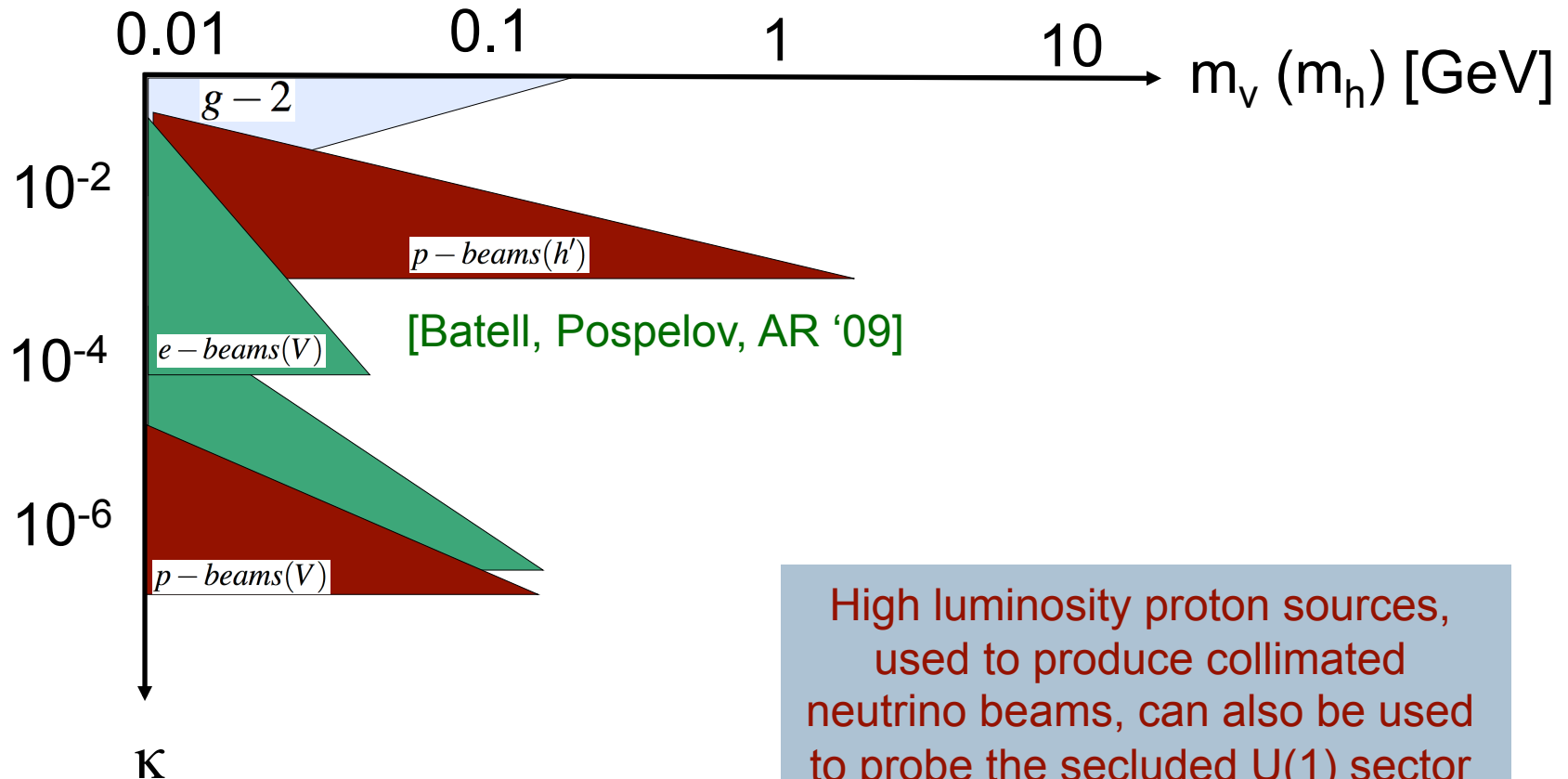
Experimental Sensitivity

In pictorial form



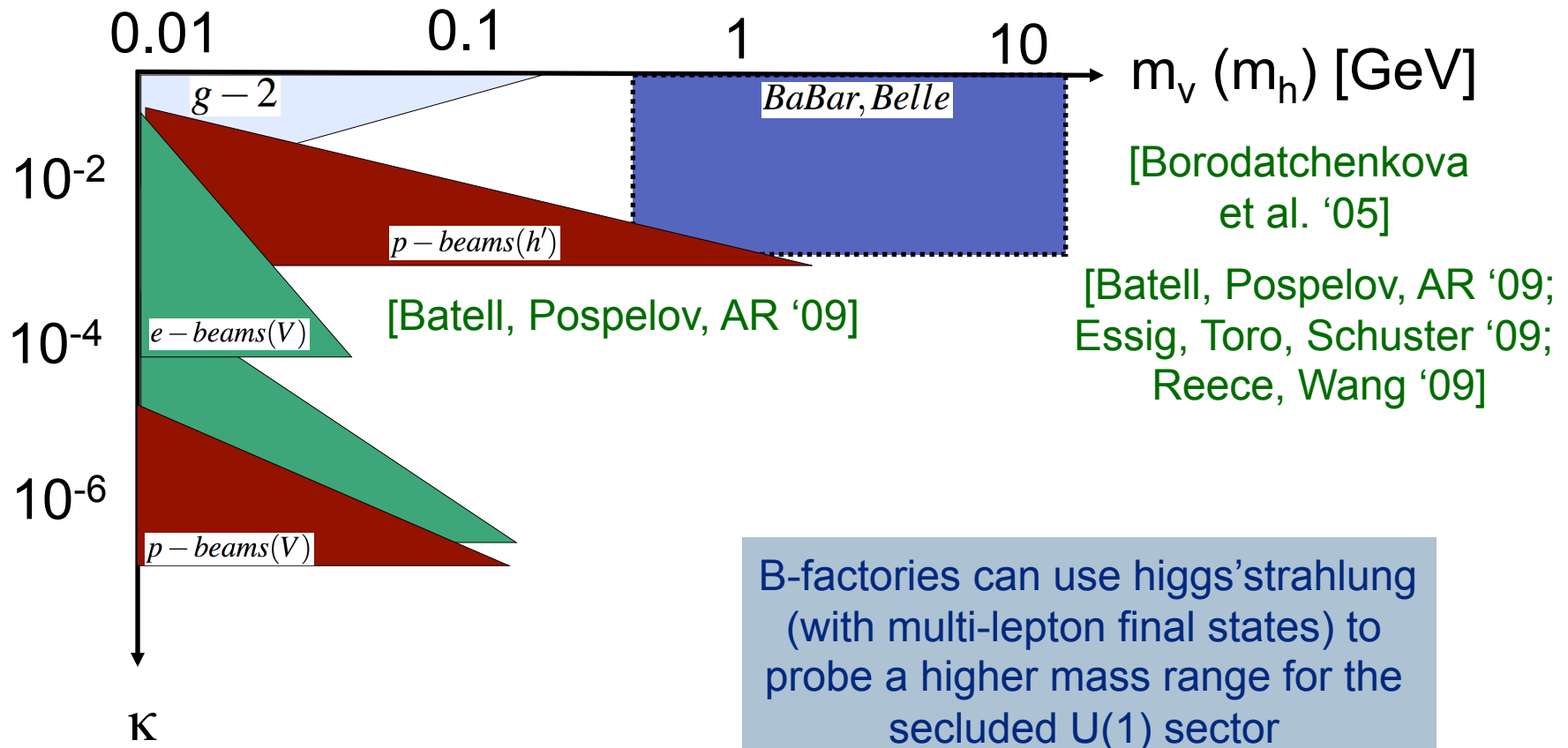
Experimental Sensitivity

In pictorial form

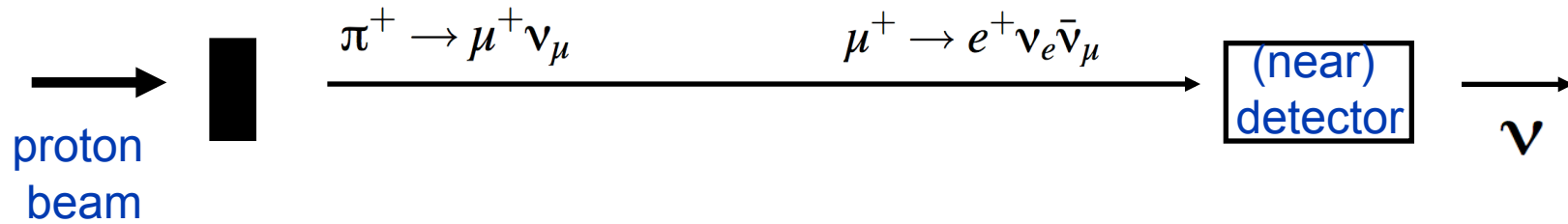


Experimental Sensitivity

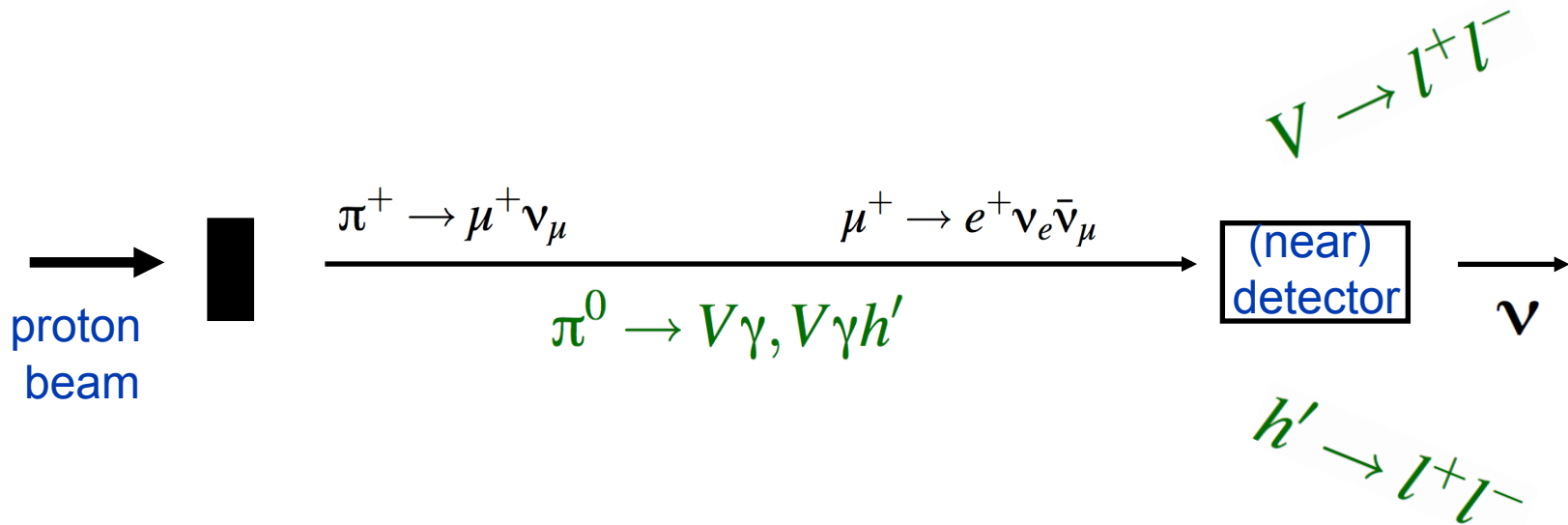
In pictorial form



Fixed target probes - Neutrino Beams



Neutrino Beams

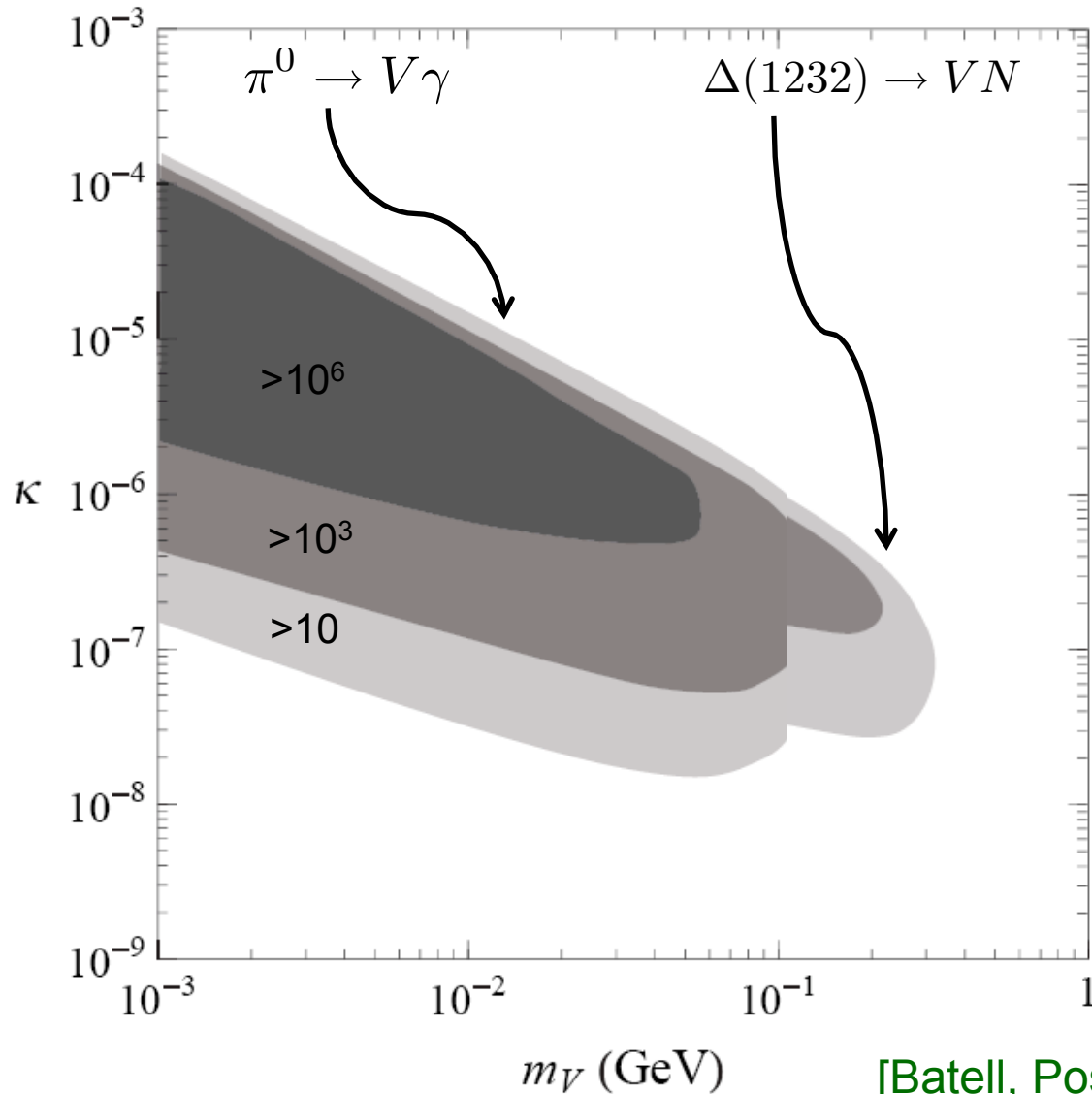


V sensitivity: $\Gamma_V \sim O(100m) \implies \kappa \sim 10^{-6} - 10^{-7}$

h' sensitivity: $\Gamma_{h'} \sim O(100m) \implies \kappa \sim 10^{-2} - 10^{-3}$

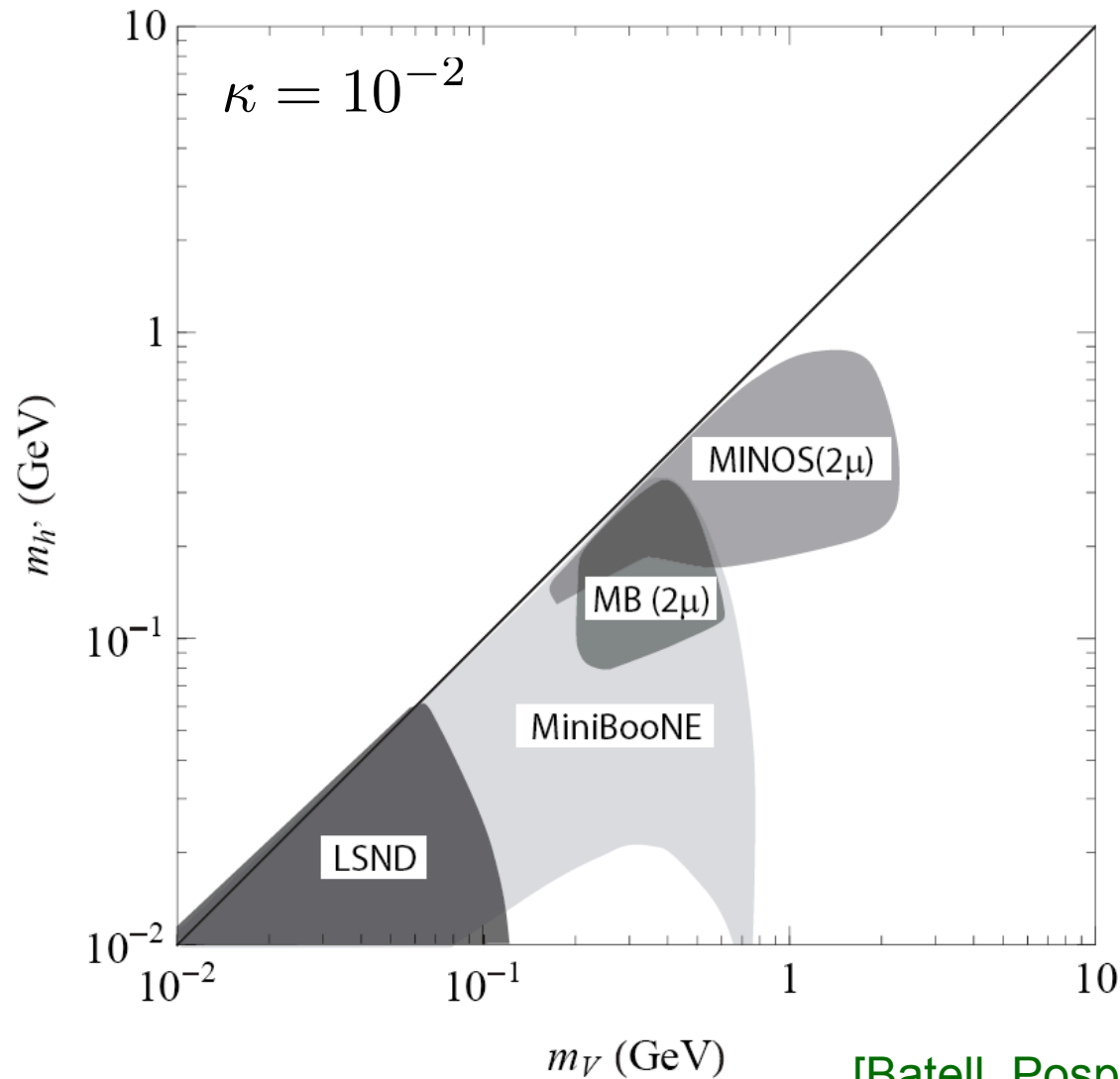
Sensitivity to Vectors

LSND: Distance to detector = 30m, 10^{23} protons on target



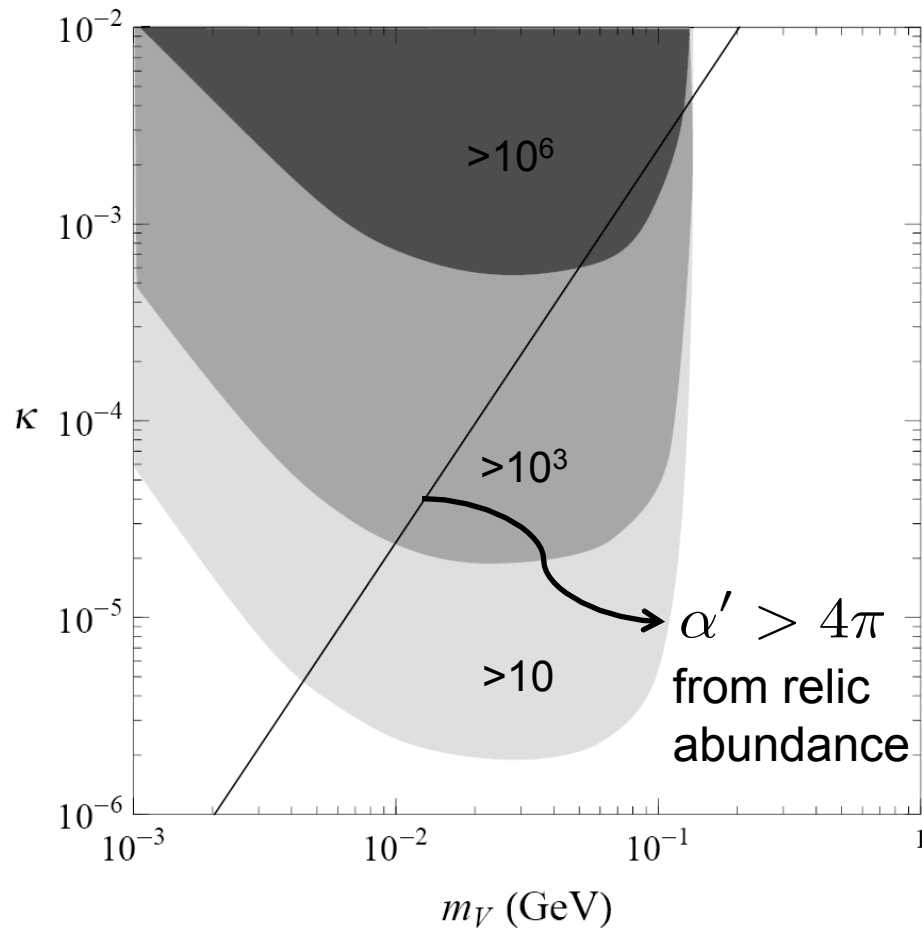
Sensitivity to Higgs'

From LSND, MiniBooNE, NuMI/MINOS, ...



Sensitivity to a “Dark Matter Beam”

If dark matter χ is light (MeV-scale [Boehm et al '03, Fayet '04,'07]), V may decay in the hidden sector ($V \rightarrow \chi\chi$) with NC-like signatures ($\chi+e \rightarrow \chi+e$) in the detector

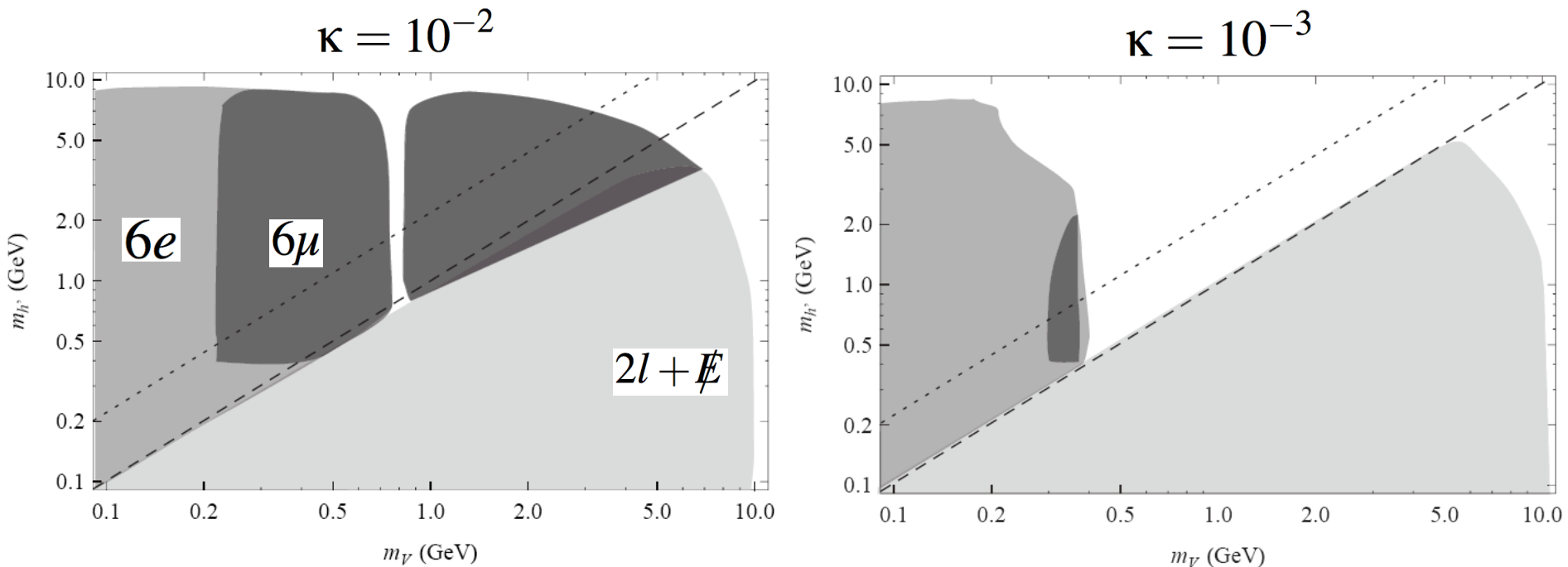


LSND provides the strongest constraints on certain models of MeV-scale DM!

Collider probes - B-factories

Higgs'strahlung: $e^+e^- \longrightarrow Vh' \longrightarrow 6l$ (or $2l + \cancel{E}$)

NB: primarily continuum V-production, Y 's give a small correction



[Batell, Pospelov, AR '09]

[see also: Essig et al. '09; Reece & Wang '09]

Probes via Rare Meson Decays

Can also search for resonant structures in rare meson decays with
(low mass) sensitivity $\kappa \sim 10^3 - 10^4$

- Kaon decays: $\text{Br}_{K_L \rightarrow V\gamma} \simeq 10^{-3} \times \kappa^2 \times \left(1 - \frac{m_V^2}{m_K^2}\right)^3$

$$\text{Br}_{K_L \rightarrow \mu^+ \mu^- \gamma} \simeq 3.6 \times 10^{-7} \quad [10^4 \text{ events @ KTeV '01}]$$

- $\Upsilon(nS)$ decays: B-factories now providing significant constraints of new light states (but no $V\gamma$ decay)

[Babar '09, see talks by Kolomensky & Guido]

Concluding Remarks

- A neutral hidden sector is an intriguing possibility, motivated by dark matter, RH neutrinos, SUSY breaking, ...
- Light degrees of freedom may interact with the SM at the renormalizable level via the vector, Higgs, and neutrino portals
- Sensitivity to these portals lies at the luminosity frontier, e.g. medium energy e^+e^- machines (B-factories), rare decays, fixed targets, ...
- Significant sensitivity to a secluded $U(1)$ from neutrino sources: LSND, MiniBooNE, NuMI/MINOS, T2K, NOvA, ...