

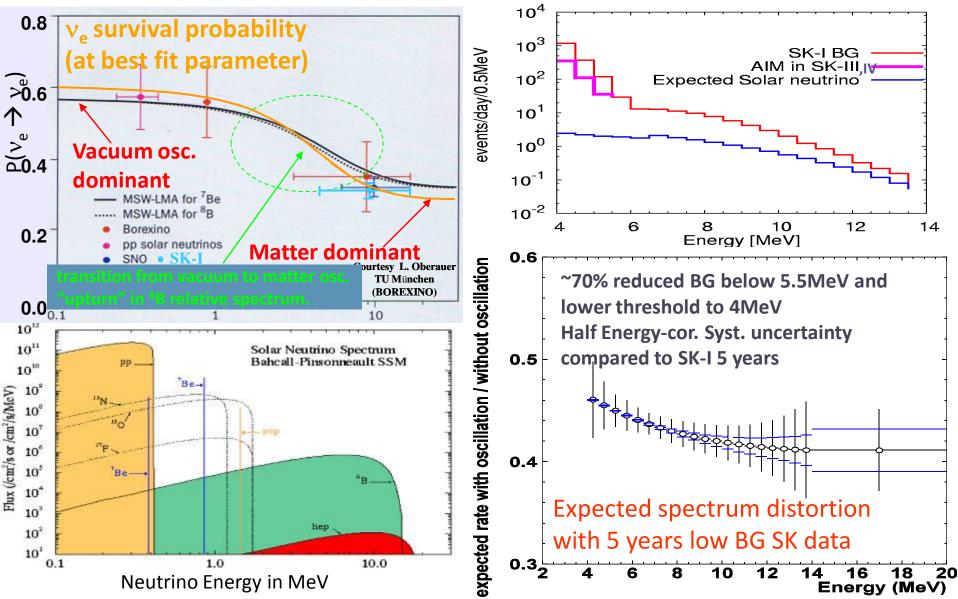
# **Solar Neutrino Measurement at SK-III**

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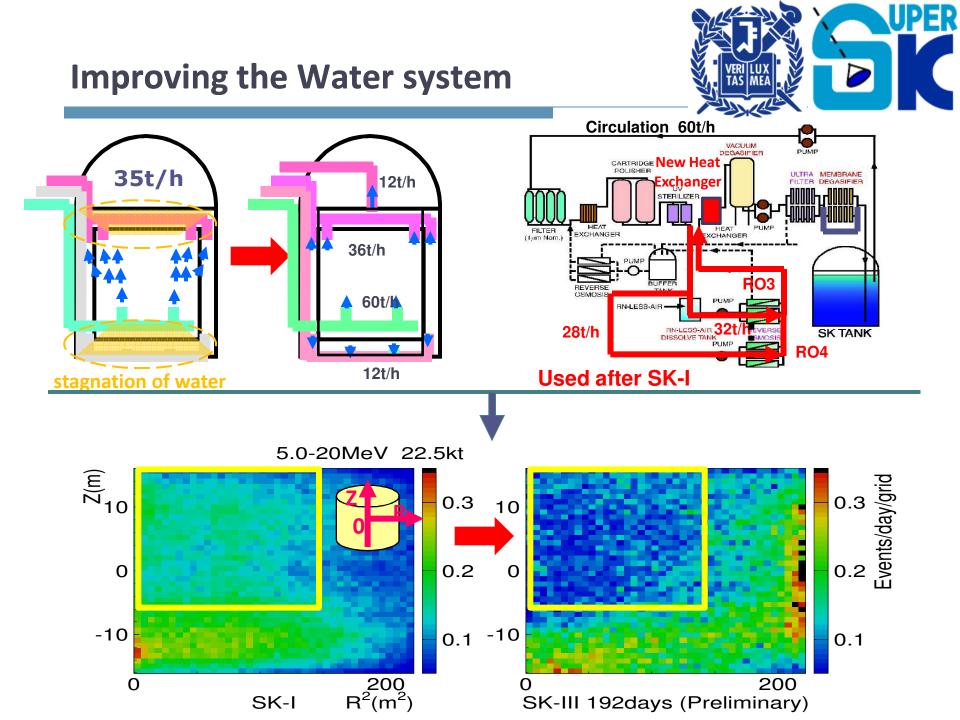
### **Solar Neutrino Future Prospects in SK**





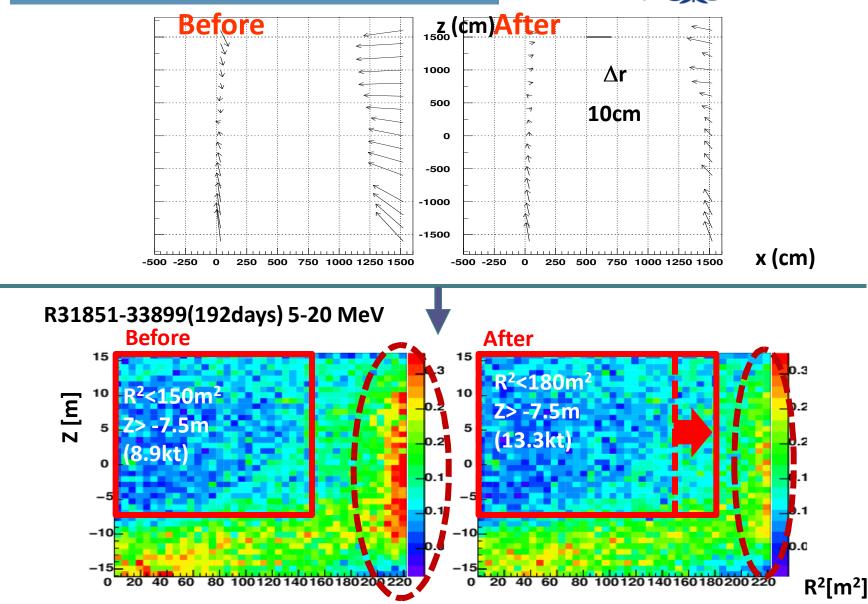


- To reduce BG and lower the energy threshold, improved the water system
- To reduce systematic errors
  - Found the reason of a large vertex shift(one source of the largest systematic error since SK-I) and shortened the shift
  - Installed the time and position dependency of the water quality into MC
  - Improved MC and a direction fitter and achieved ~10% better angular resolution
  - Most of the reconstruction tools & reduction criteria are retuned recently.
- Currently, estimation of the systematic errors are under way.
- **The improved results will be summarized in this year.**



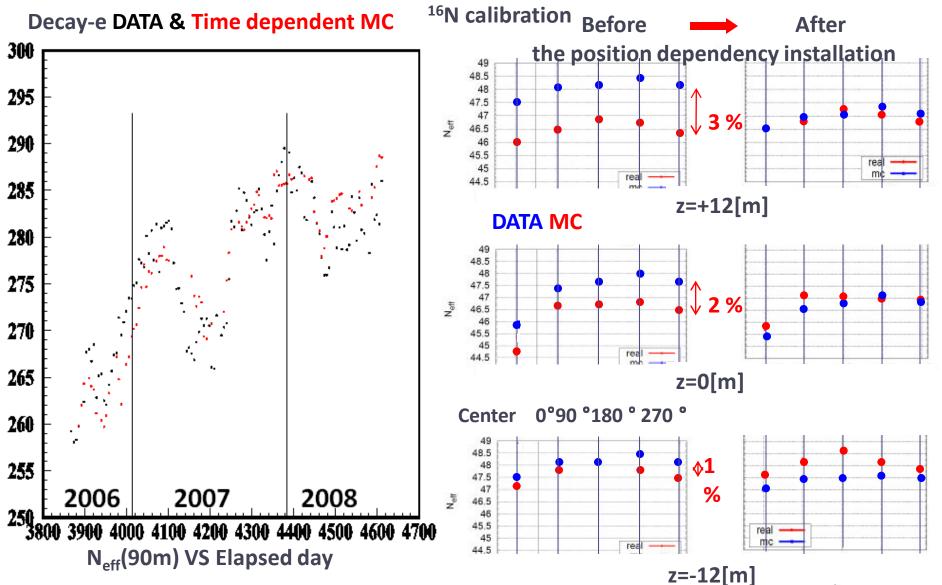


# **Shortening the Vertex shift**



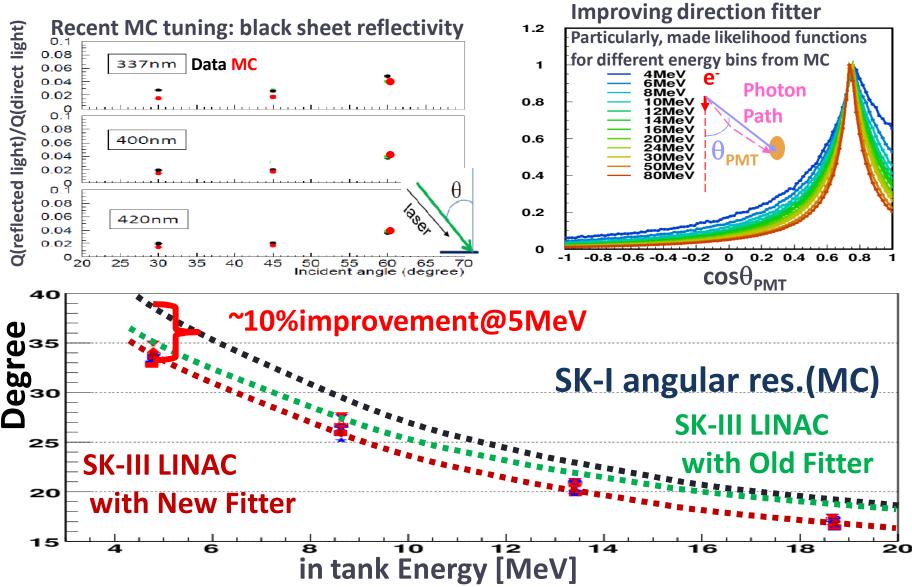
### **Time & Position dependent MC**





# **Reducing the angular resolution**





# The estimation of the systematic errors



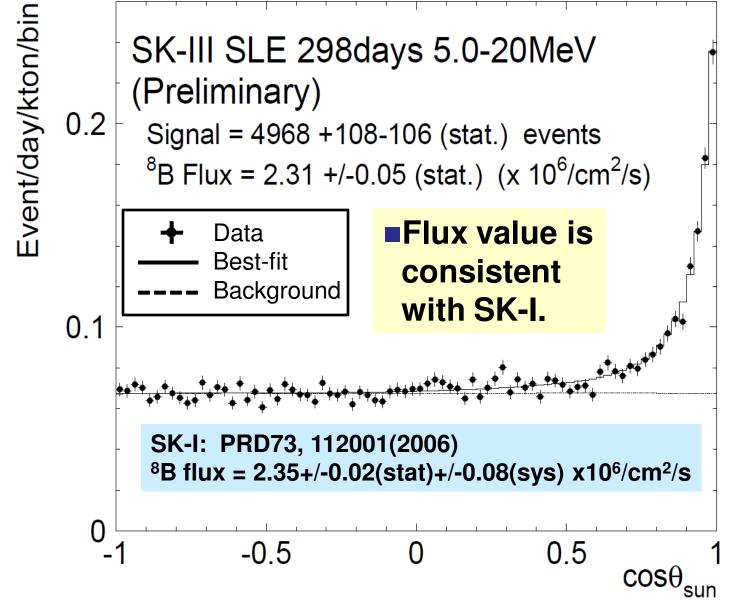
	SK-l Flux(%)	SK-III Flux(%)
Energy Scale	+/- 1.6	
Tirgger eff	+0.4 -0.3	
Spallation cut	+/- 0.2	
Reduction	+2.0 -1.6	Still Under way
Gamma Cut	+/- 0.5	
Vertex shift	+/- 1.3	Trying to reduce
Angular res.	+/- 1.2	
BG shape	+/- 0.1	
Livetime calc	+/- 0.1	
Total	+3.5 -3.2	



- **Observed** <sup>8</sup>B v flux in SK-III
- □ Angular distributions
- □ <sup>8</sup>B ∨ Flux Time Variation
- Recoil Electron Energy Spectrum
- Day/Night Asymmetry

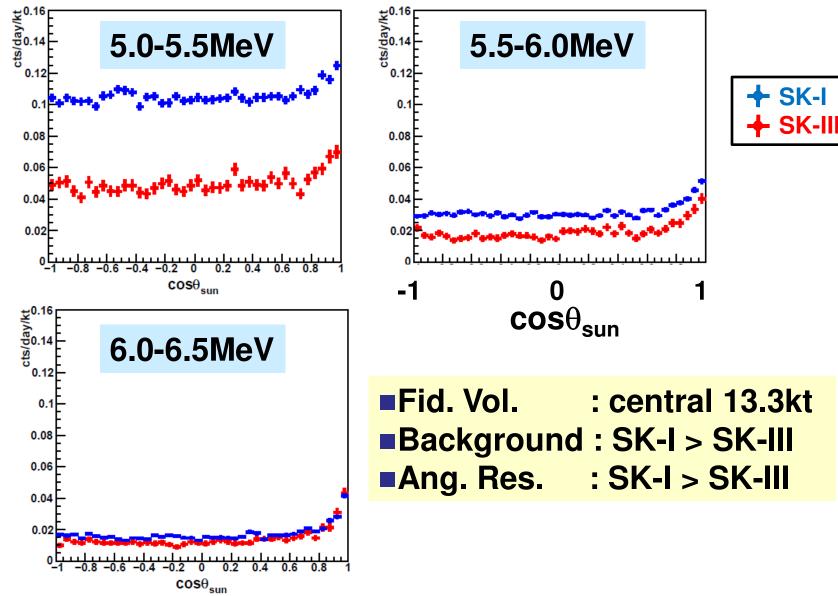
# **Observed <sup>8</sup>B flux in SK-III**





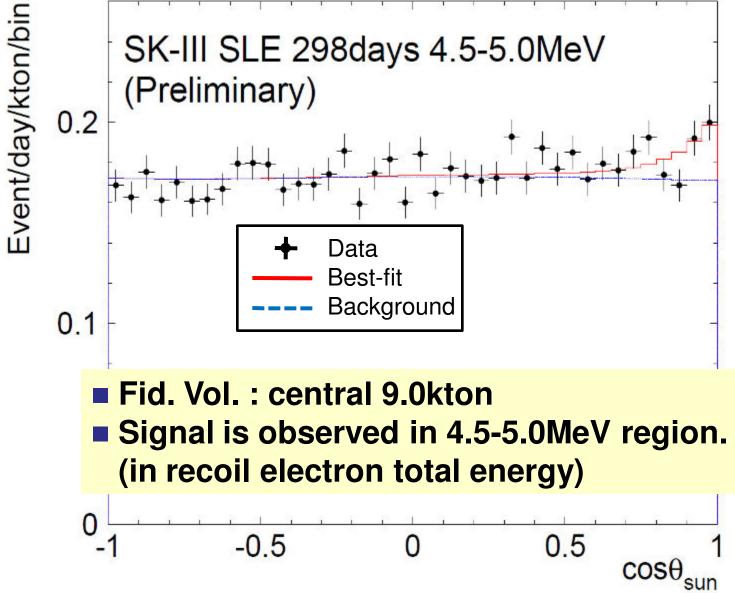
# **Angular distributions:1**





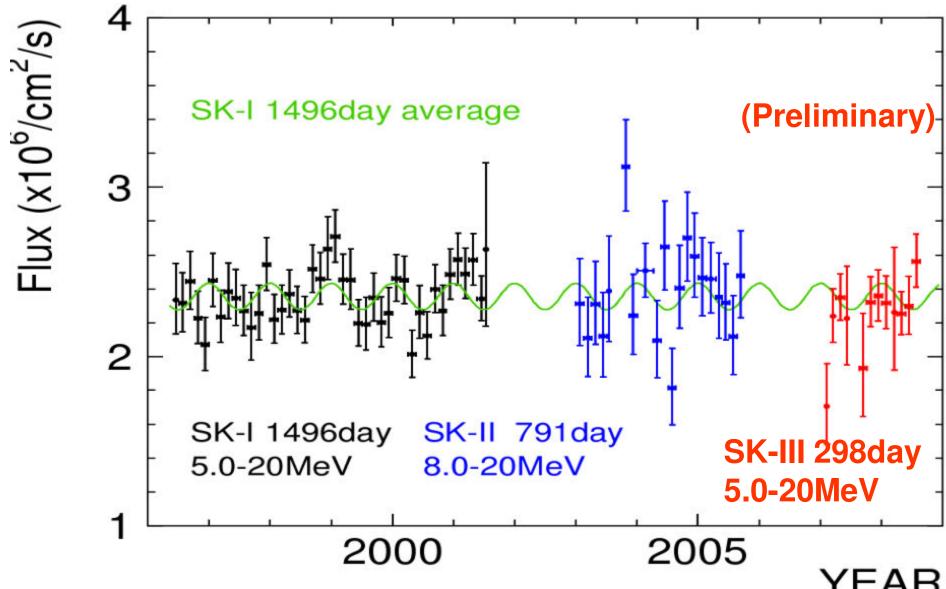
# **Angular distributions:2**





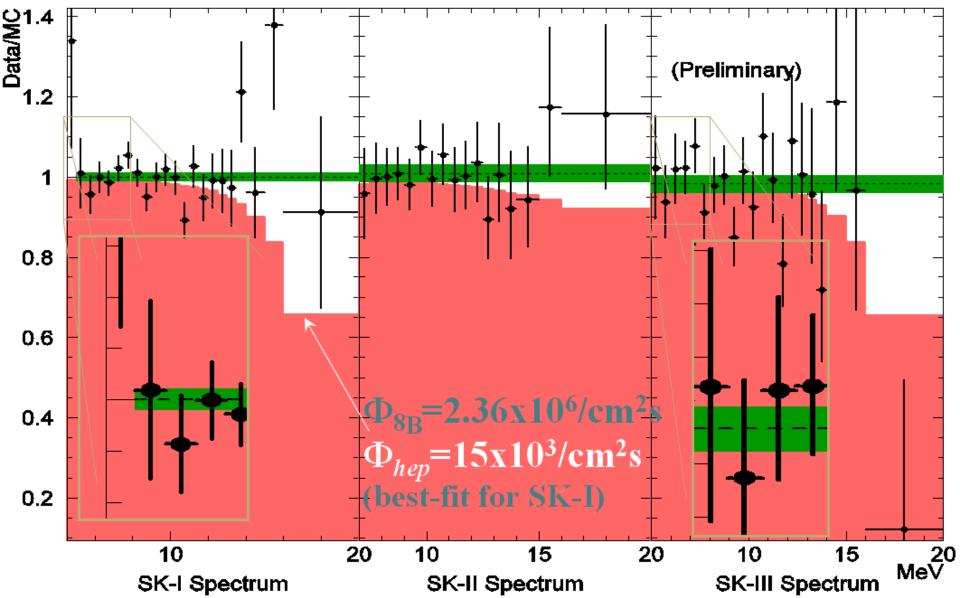
## $^{8}\text{B} \ v$ Flux Time variation





# **Recoil Electron Energy Spectrum**







- only direct test of matter effects on solar neutrino oscillations
- SK-I measured  $A_{DN}=2(D-N)/(D+N)=-2.1\pm2.0\%$ (stat)
- SK-I also fit LMA day/night variations; expressed as A<sub>DN</sub> the result is A<sub>DN</sub>=-1.8±1.6%(stat)
- SK-II measured A<sub>DN</sub>=-6.3±4.2%(stat)
- SK-III can measure A<sub>DN</sub> to ±4.3%(stat) with the shown 298 days of data; maybe to ±3.7%(stat) using the entire SK-III data set (including periods w/o SLE or high very low energy background runs)
- **SK-I-III can determine A<sub>DN</sub> to ±1.6%(stat)**
- **SK-I-III can fit LMA D/N variations to ±1.3%(stat)**

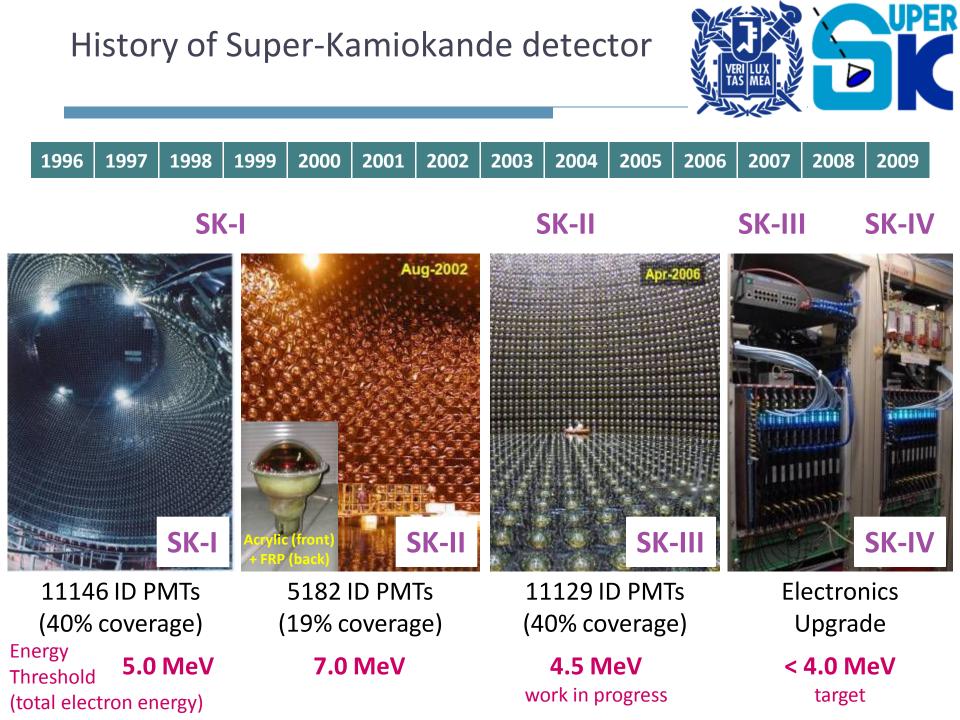
#### Summary



- Achieved lower backgrounds <5.5 MeV in the center of SK</p>
- Almost finished analyzing SK-III data
- Tried to reduce systematic errors compared to SK-I
- **Estimation of the systematic errors are under way**
- Consistent with SK-I and SK-II results within statistical uncertainties
- At SK-III, SK solar analysis energy threshold was lowered to 4.5 MeV
- With existing data, SK is statistically sensitive down to LMA day/night asymmetries of 1.3%
- **The improved results will be summarized in this year.**

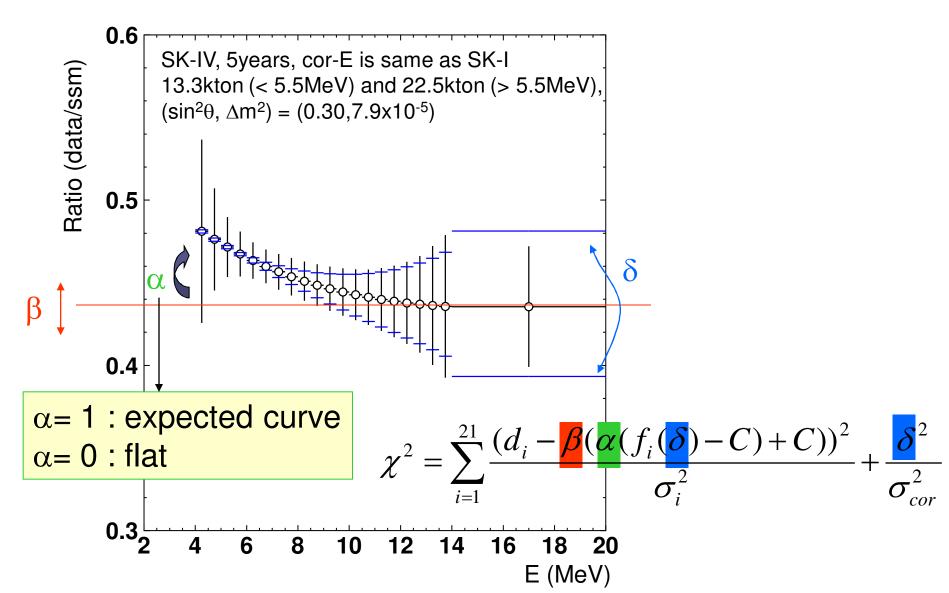
#### Backup





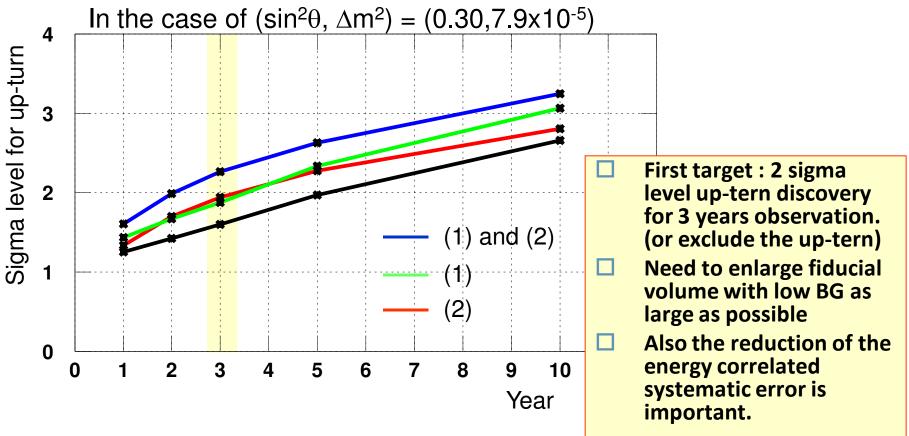






#### Sensitivity of the upturn measurement





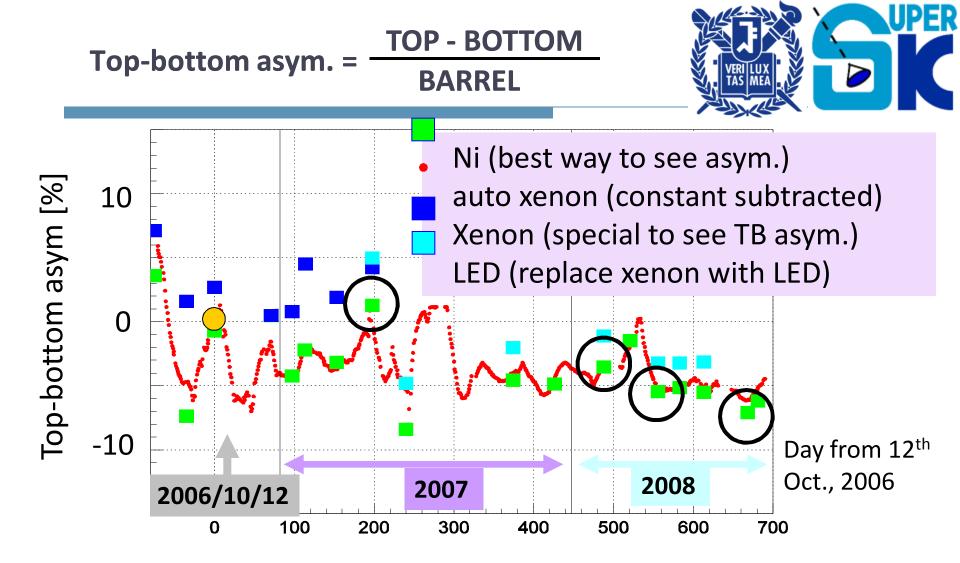
- (1) Enlarge fiducial volume to 22.5kton with low B.G.
- (2) Half energy correlated systematic error as SK-1.

The black line shows the 13.3kton (<5.5MeV), 22.5kton (>5.5MeV) fiducial volume with the same energy correlated error as SK-1





- Running 100% trigger efficiency at 4.5MeV now.
- □ The trigger threshold will be lowered in future.



check the energy scale in 4 Ni data sets with 3 positions; 2007/04/27, 2008/02/12, 2008/04/19, 2008/08/09 z=-12m, 0m, 12m



Right figure is example of WT function.

 $\lambda_0$ =100m,  $\alpha$ = -0.25.

If  $\lambda_0$  and  $\alpha$  are calculated for all SK-III period, the time and position dependence of water quality can be installed into MC.

Time table of  $\lambda_0$  is already calculated by Beongsu.

Try to make a relation between alpha and tba value.

