

# Extracting Bulk QGP Properties from Experiment

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# MaDAI Collaboration

## Models and Data Analysis Initiative

### Michigan State University

RHIC Physics: Scott Pratt

Supernova: Wolfgang Bauer

Astrophysics: Brian O'Shea and Mark Voit

Atmospheric Modeling: Sharon Zhong

Statistics: Dan Dougherty



### Duke University

RHIC Physics: Steffen Bass and Berndt Müller

Statistics: Robert Wolpert

### University of North Carolina & Renaissance Computing Institute

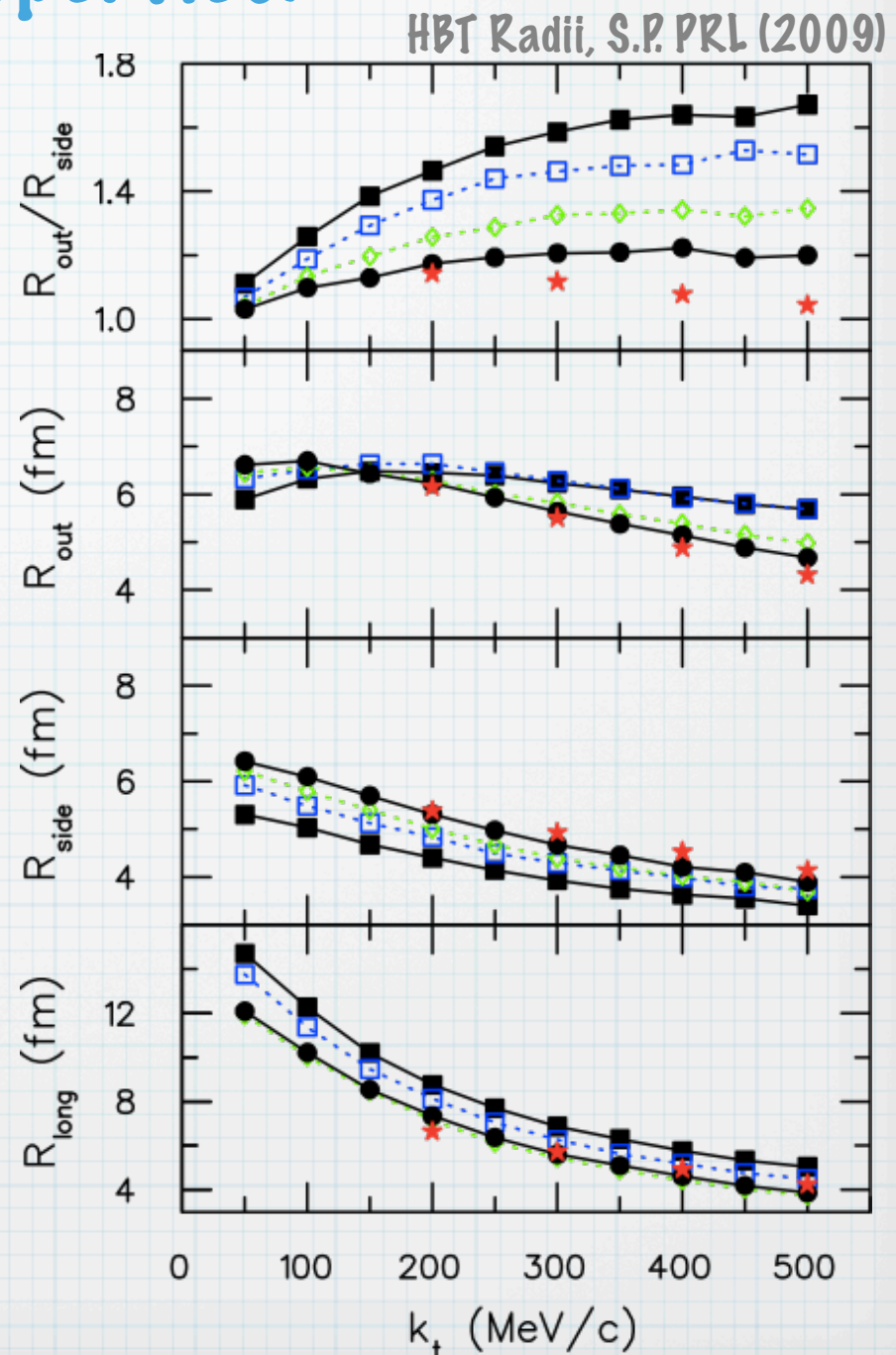
Visualization: Xunlei Wu

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# Lessons from RHIC (bulk properties)

- \* Matter is rather stiff:  
(no large latent heat,  
but softer than  $\pi$  gas)
- \* Early flow seems important:  
(otherwise difficult to fit HBT)
- \* Viscosity is low:  
(mostly from elliptic flow)

$$\frac{\eta}{s} \approx \left( \frac{1}{2} \text{ to } 3 \right) \times \frac{\hbar}{4\pi}$$



... but nothing is quantitative or rigorous

- \* EOS (min  $c_s^2$ , width of soft region, max  $c_s^2$ )
- \*  $??? < \eta / s < ???$  (energy dependence?)
- \*  $\epsilon$  for  $\tau < 0.5$  fm/c uncertain by factor of 2

1. Can parameters be stated rigorously? (with uncertainties)
2. Can important links be identified? (e.g. be viscosity and  $v_2$ )
3. Can prospects for future measurements be quantified?



# RHIC Analysis Challenge

## Observables

- $dE_{\dagger}/dy$
- $dn_{ch}/dy$
- $dn_B/dy$
- jets & high pt spectra
- hard photons
- $J/\psi \rightarrow$  dileptons
- int. mass direct photons
- dilepton continuum
- $\rho/\omega/\phi \rightarrow$  dilepton peaks
- $\pi/K/p$  spectra
- $K/\pi$  ratio
- hyperon yields
- antibaryon yields
- HBT Correlations
- low  $p_{\dagger} \pi$  spectra
- deuteron yields
- balance functions
- isospin fluctuations
- multiplicity fluctuations
- directed & elliptic flow
- radial flow

## Model Assumptions and Parameters

- QCD saturation scale
- QCD structure functions
- existence of quarks & gluons
- evolution of chiral condensate
- color correlations at high  $e$
- stopping mechanism
- initial density profiles
- kinetic thermalization
- initial chemistry
- chemical rates
- eq. of state (P vs.  $e$ )
- specific heat (T vs.  $e$ )
- in-medium mass shifts
- cross sections (in-medium)
- formation time
- viscosity
- phase instabilities
- hadronization mechanism:  
(entropy, particle creation)

**Individual elements cannot be isolated!!**  
**complex, non-linear network**

# Uncertainties and Parameters

Initial State	6	Energy density, profile shape, rapidity width, pressure, anisotropy of $T_{ij}$ , quark/gluon content
Hadronic Boltzmann	2-4	Mass changes
Eq. of State / Viscosity	3-8	Might be constrained by lattice, hadron gas
Chemical	3-6	Quark density, relaxation rates, hadronic scattering reduction
Jet Quenching	2-4	Dissipation rates
Systematic Experimental	?	Efficiencies, calibrations...

**$\approx 30$  parameters**  
**Some are unimportant**  
**Some combinations are unimportant**



Turn ALL the  
knobs!!!!





# Bayesian Analysis

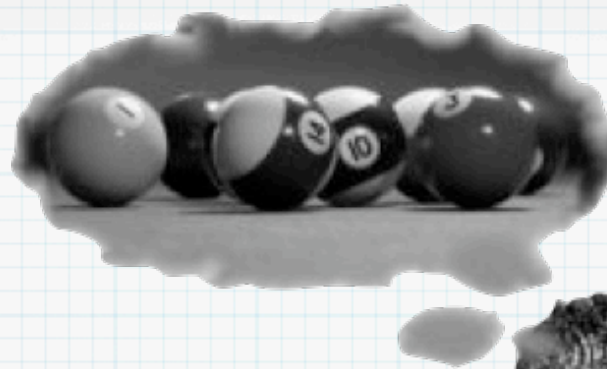
WIKIPEDIA: Bayesian inference is statistical inference in which evidence or observations are used to update or to newly infer the probability that a hypothesis may be true. The name "Bayesian" comes from the frequent use of Bayes' theorem in the inference process.

$$P(H | E) = \frac{P(E | H)P(H)}{P(E)}$$

- $P(H)$  is probability (in absence of  $E$ ) for parameter set  $H$   
a.k.a. the "prior distribution"
- $P(E|H)$  is probability of  $E$  given  $H$ , i.e.,  
$$P \sim \exp\left(-\sum \delta_i^2 / 2\sigma_i^2\right)$$
- $P(E)$  is net probability of  $E$ , i.e., a normalization factor
- $P(H|E)$  is probability of parameter set  $H$  given  $E$



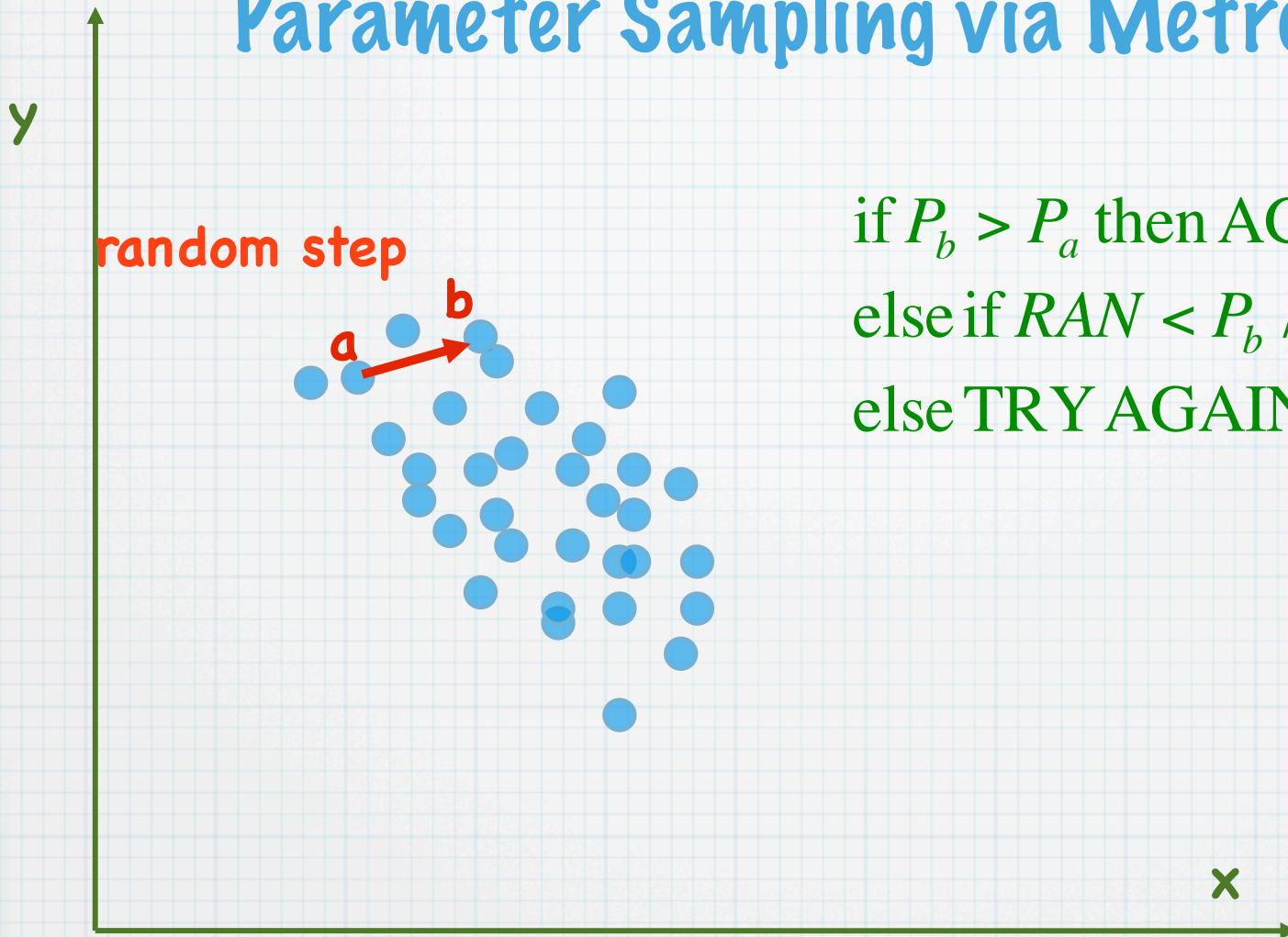
# Bayes' Theorem



$$P(E \& H) = P(E | H) \cdot P(H) = P(H | E) \cdot P(E)$$

$$P(H | E) = \frac{P(E | H) \cdot P(H)}{P(E)}$$

# Parameter Sampling via Metropolis



if  $P_b > P_a$  then ACCEPT

else if  $RAN < P_b / P_a$  ACCEPT

else TRY AGAIN

Can find disjoint regions  
No problem with undeterminable parameters



# Surrogate Models (a.k.a. Emulators, Meta-Models)

## Brute Force:

- Sampling requires millions of runs
- Each run requires 1 work-station day

## Alternative:

- Run  $10^2 - 10^3$  times at various points
- "Interpolate" to find values at all other points
- Competing "interpolation" schemes:
  - Gaussian fields
  - Multi-dimensional splines

*An Emerging Science*

# Identifying Key Links

Consider parameters  $\theta_\alpha$  and observables  $y_i$

$$\sigma_{ij} \equiv \langle \delta y_i \delta y_j \rangle, \quad \sigma_{\alpha\beta} \equiv \langle \delta \theta_\alpha \delta \theta_\beta \rangle, \quad \sigma_{i\alpha} \equiv \langle \delta y_i \delta \theta_\alpha \rangle$$

Averages taken from MCMC sampling

Resolving Power is

$$\sigma_{\alpha\beta}^{-1} = \sum_{ij, \gamma\delta} \sigma_{\alpha\gamma}^{-1} \sigma_{\gamma i} \sigma_{ij}^{-1} \sigma_{j\delta} \sigma_{\delta\beta}^{-1}$$

and can be divided into contributions

$$\sigma_{\alpha\beta}^{-1(s)} \equiv \sum_{ij \in s, \gamma\delta} \sigma_{\alpha\gamma}^{-1} \sigma_{\gamma i} \sigma_{ij}^{-1} \sigma_{j\delta} \sigma_{\delta\beta}^{-1}$$

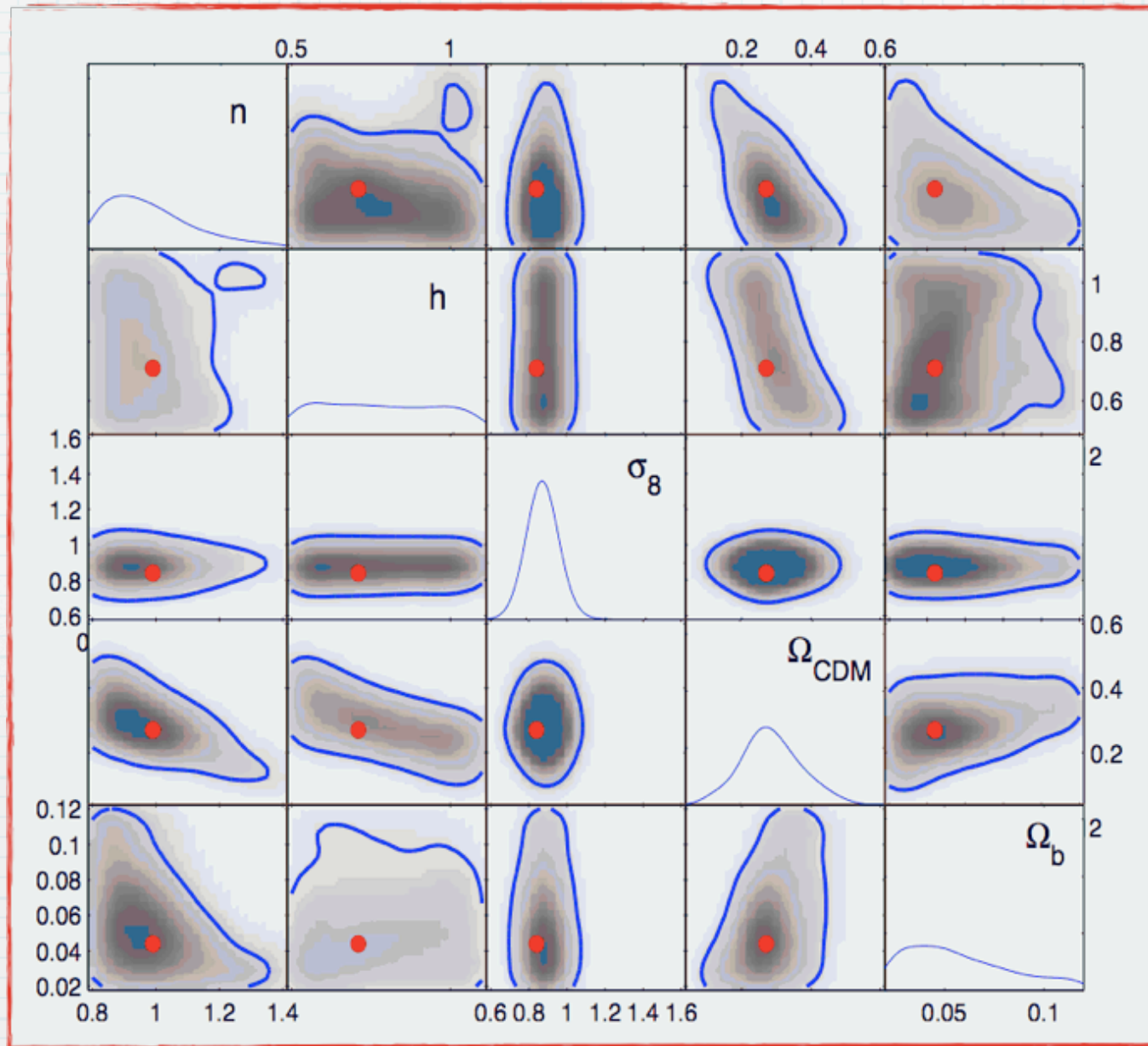
"s" refers to subset of data, e.g. elliptic flow



# Assessing Prospects for New Data

- **Make Simulated Data**
- **Analyze Like Real Data**
- **Calculate Resolving Power of Subset**

Other fields have done this ....



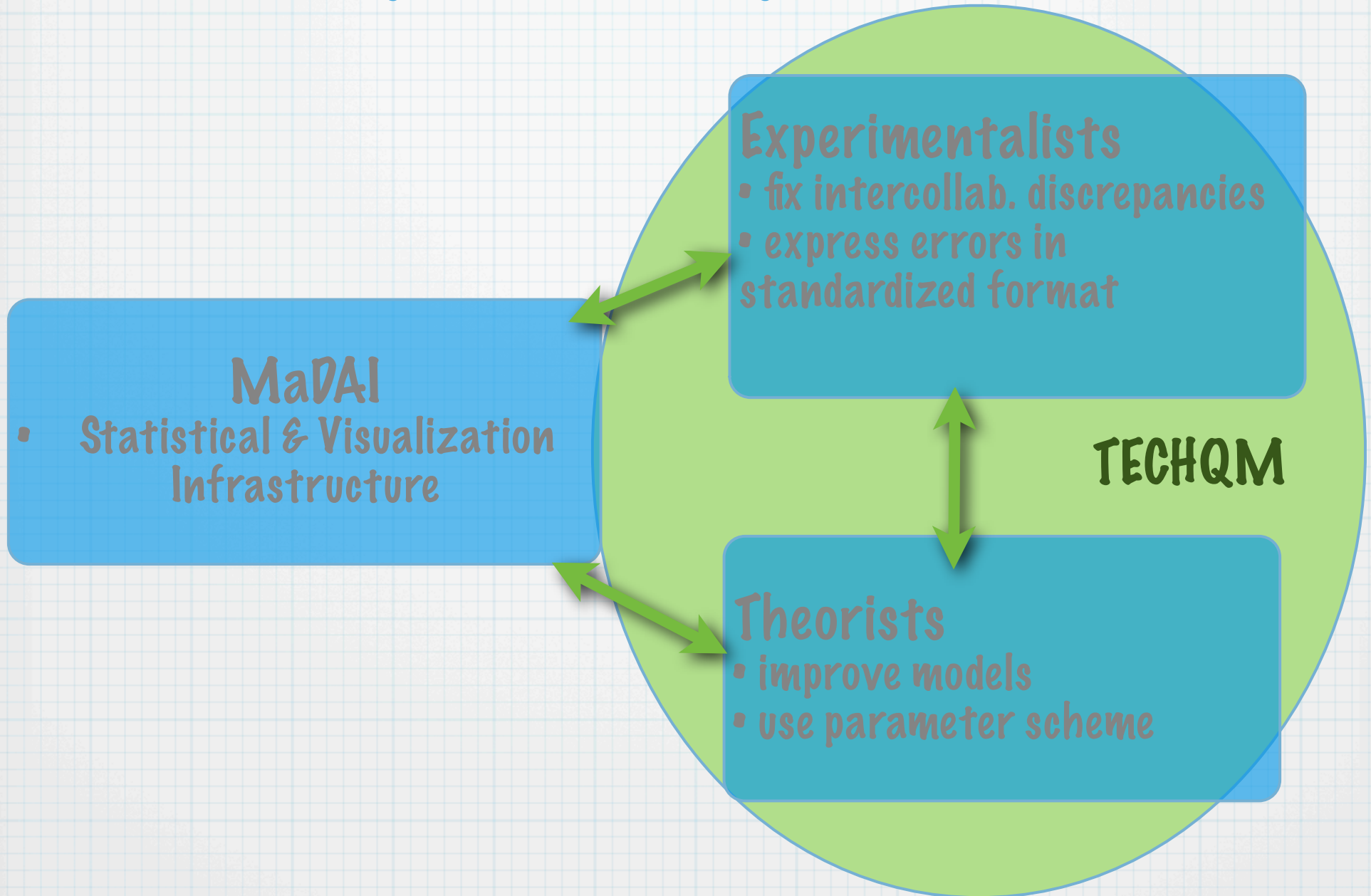
Cosmological parameters (Habib et al, astro-ph/0702348 1)



# Can this work for RHIC?

- \* **Must be amenable to parameterization**
  - \* **Model must contain basic truth**
  - \* **Not too many competing theories**
  - \* **First apply to soft observables (spectra, flow, HBT...)**
  - \* **Provide validated base for other calculations**
- \* **Must have well stated errors**
  - \* **Statistical & systematic for both theory and experiment**
  - \* **Cross correlated errors**
  - \* **May require re-expression of experimental results**
  - \* **Intimate theory/experimental discussions**

# Who Does What





# Outcomes

- **Rigorous Statement of Bulk Properties**
- **Validated Base for Calculating:  
Fluctuations, Jet Quenching, Heavy Flavors...**