

# COMMISSIONING OF THE CMS ENDCAP MUON SYSTEM

**Phillip Killewald**

The Ohio State University

Meeting of the Division of Particles and Fields  
of the American Physical Society  
July 2009



## ① OVERVIEW

The CMS Detector

The CMS Endcap Muon System

Muons in TeV Physics

## ② COMMISSIONING FOR PHYSICS

Commissioning Time Line

Results from Beam Halo Muons

Results from Cosmic Rays

## ③ CONCLUSIONS

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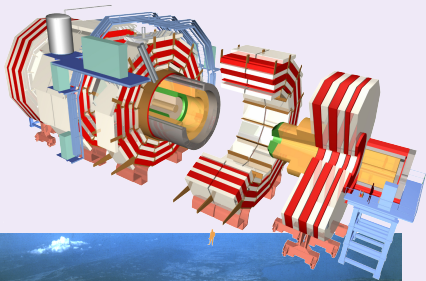
Results from Cosmic Rays

## 3 CONCLUSIONS



## CMS: The Compact Muon Solenoid

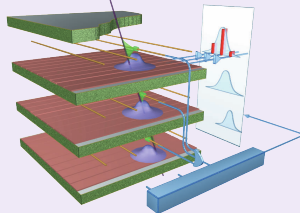
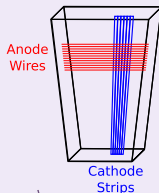
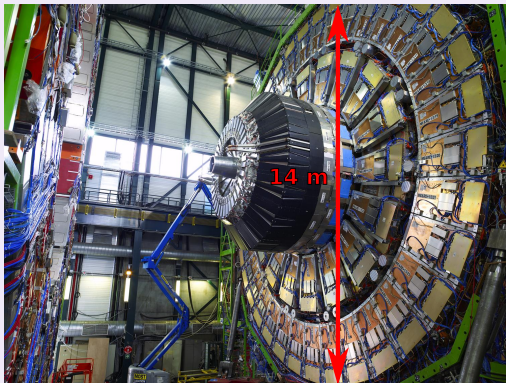
- General-purpose particle detector
- Centered around one of the interaction points of the Large Hadron Collider





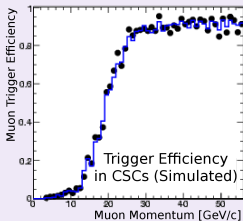
# THE CMS ENDCAP MUON SYSTEM

- CMS uses 468 Cathode Strip Chambers as a muon spectrometer in the endcaps
- Each CSC has 6 layers of cathode strips and anode wires
- Each layer makes an independent 2D measurement of a through-going muon's position



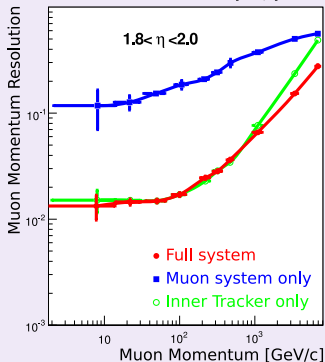


# THE CMS ENDCAP MUON SYSTEM



The CSCs are designed to provide:

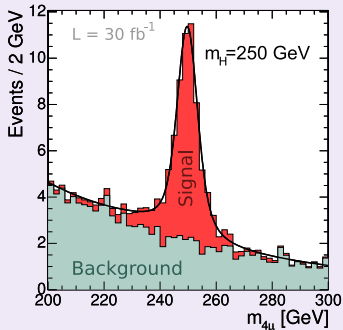
- Reliable muon identification
- Efficient muon triggering
- Improved muon momentum resolution for  $p > 500$  GeV



The CMS Collaboration 2006, CERN-LHCC-2006-001



# THE IMPORTANCE OF MUONS IN TEV PHYSICS



Muons are important in TeV-scale physics because they are:

- Easily identifiable with a clean experimental signature
- Present in many new physics signals
  - $H \rightarrow ZZ^{(*)} \rightarrow 4\mu$
  - $Z' \rightarrow 2\mu$
  - $W' \rightarrow \nu_\mu + \mu$
  - $\chi_2^0 \rightarrow \chi_1^0 + 2\mu$

The CMS Collaboration 2007, *J. Phys. G: Nucl. Part. Phys.* **34** 995-1579, doi:10.1088/0954-3899/34/6/S01

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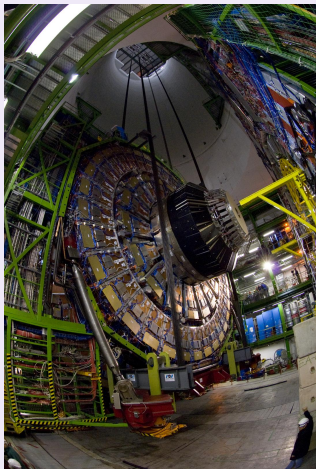
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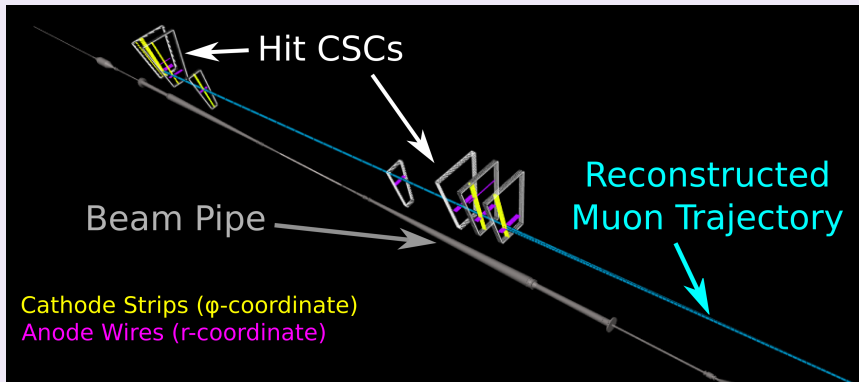
- Final CSC affixed to iron disk: February 2007
- Final disk lowered to underground cavern: January 2008
- First LHC beam: September 2008
  - Only 18 months to ready  $\sim 4 \times 10^5$  data channels from  $\sim 1.4 \times 10^4$  electronics boards!

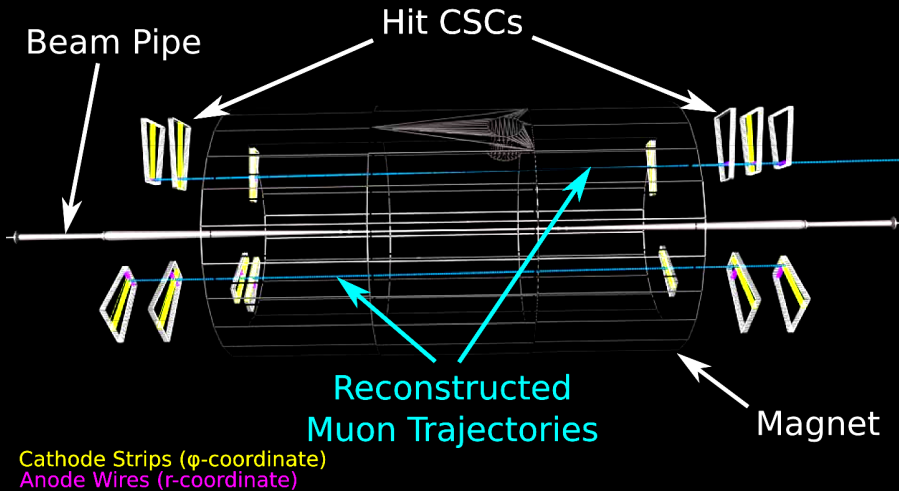


# FIRST LHC BEAM

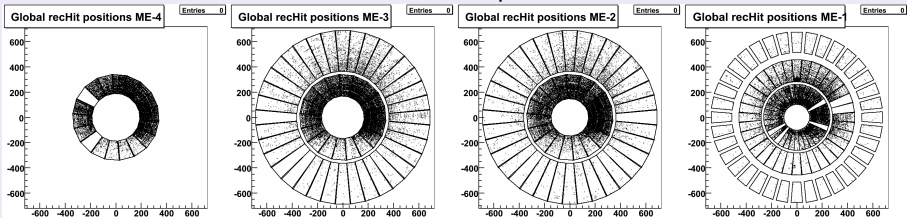
SEPTEMBER 11, 2008

- LHC successfully captured and circulated a beam of protons in the anti-clockwise direction for a total of 12 minutes
- Collisions of protons with beam-pipe gas generates a “halo” of muons around the beam-pipe traveling nearly parallel to it



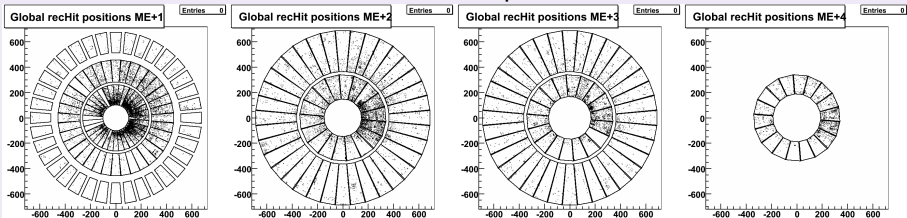


## -Z Endcap



Beam Direction 

## +Z Endcap



Beam Direction 



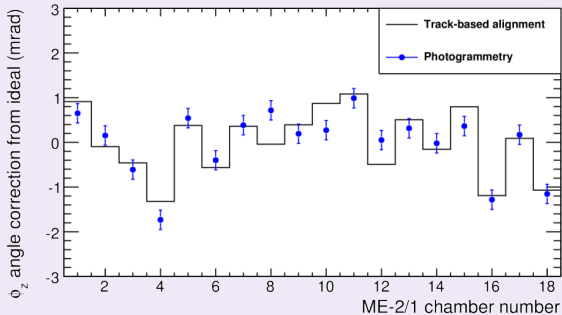
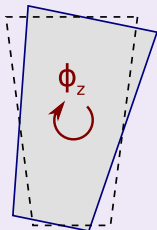
# TRACK-BASED ALIGNMENT

USING BEAM HALO MUONS

CSCs employ two independent methods of spacial alignment:

- Hardware-based
  - Uses lasers and photogrammetry to align points on the CSCs to other points on the detector
  - Rough estimate of positions
- Track-based
  - Uses through-going muons to align overlapping detector regions to each other
  - Ultimate alignment method (using IP-muons)





- Track-based method agrees with photogrammetry to  $\sim 0.5$  mrad

Jim Pivarski (Texas A&M)

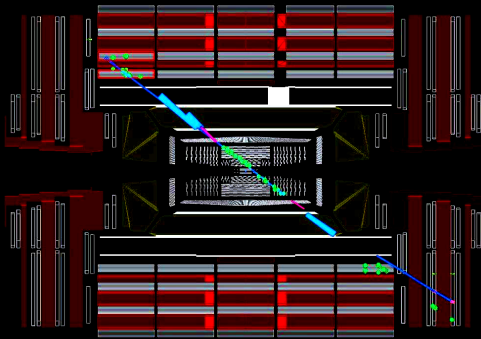
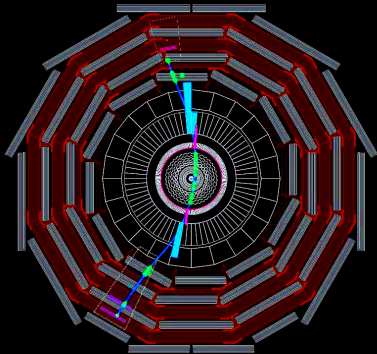


# THE ART OF MAKING LEMONADE

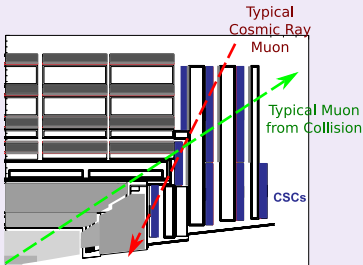
COSMIC RUN AT FOUR TESLA

- Since the LHC incident on September 18, 2008, CMS had only cosmic ray muons to use for commissioning. . .
- From October 13 to November 11, all of CMS participated in a global cosmic ray data-taking run dubbed “CRAFT”:
  - $\sim 3.7 \times 10^8$  cosmic ray events recorded
  - $\sim 2.9 \times 10^8$  of those with an operating magnetic field at 3.8T
  - $\sim 1.4 \times 10^7$  of those with muons reconstructed in the CSCs
  - $\sim 1.2 \times 10^5$  of those with muon trajectories that take them close to the interaction point (IP-pointing)

Run 66748, Event 8900172, LS 160, Orbit 167345832, BX 2011





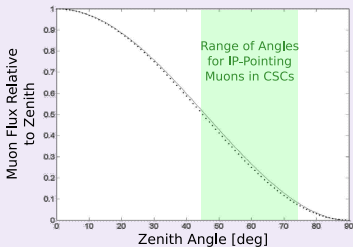


The CMS CSCs were designed for muons. . .

- coming from the interaction point
- at a rate of  $\sim 10^4$  Hz

Cosmic rays give us muons. . .

- coming mostly from the vertical direction
- at a rate of  $\sim 60$  Hz

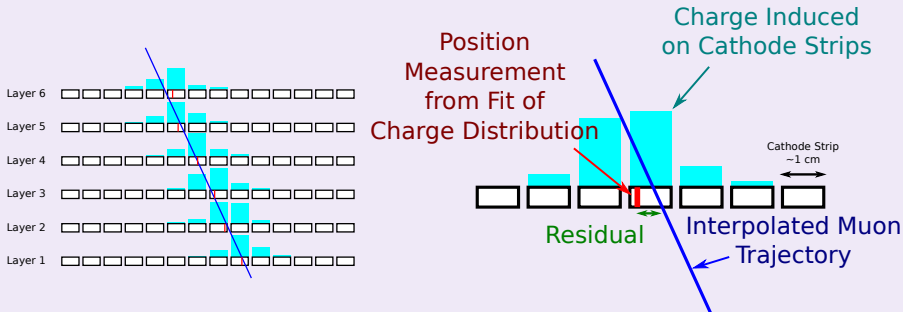


L.N. Bogdanova, et al., *Phys. Atom. Nucl.* **69** (2006) 1293; arXiv:nucl-ex/0601019v1

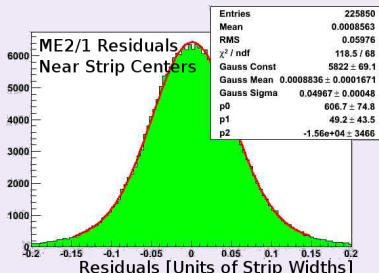
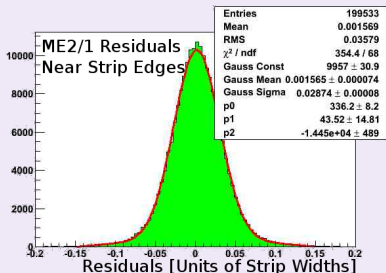


# USING COSMIC RAYS

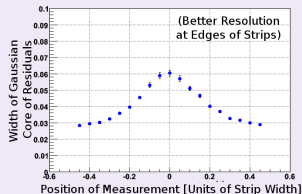
## LOCAL POSITION RESOLUTION



- Each CSC has six planes of anode wires and cathode strips
- The intersection of a strip and a wire group produces an independent 2D measurement (a “hit”) of a point on the muon’s trajectory in each layer
- Use hits from layers 1, 2, 4, 5, and 6 to interpolate where the hit in layer 3 should be and calculate residual

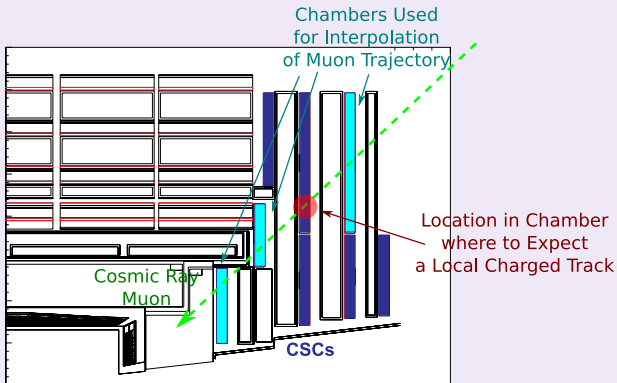


- $\sim 161 \mu\text{m}$  chamber resolution (after convolution of gaussians)
- Hits near strip edges have better resolutions because of the shape of the fit



Vladimir Palchik (JINR)

## LOCAL CHARGED TRACK (LCT) EFFICIENCIES

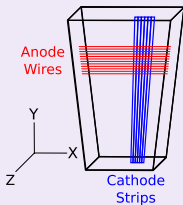
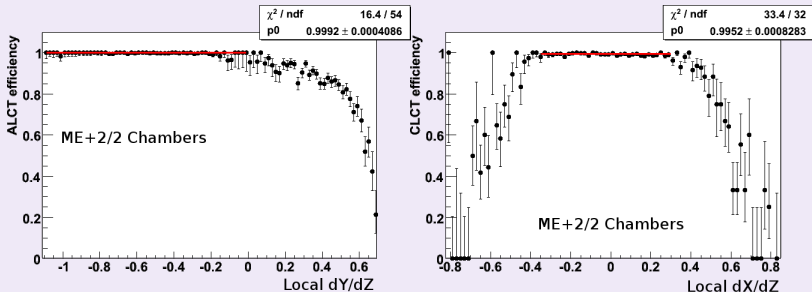


- Use 3/4 stations in a given endcap to interpolate where the muon trajectory should pass through the 4th station
- Look for local charged tracks (trajectory segments) in that neighborhood



# USING COSMIC RAYS

## RESULTS



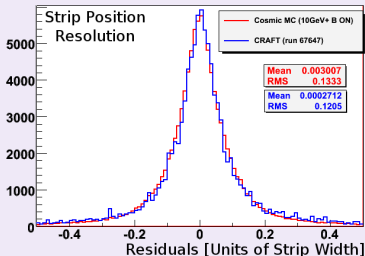
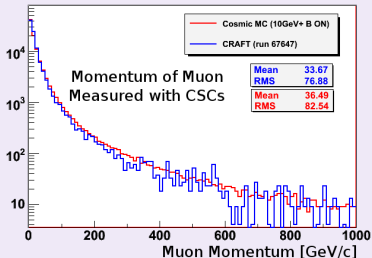
- $> 99.5\%$  trigger efficiency in trigger window (IP-pointing, high- $p_T$ )

Stoyan Stoynev (Northwestern)



# USING COSMIC RAYS

## MONTE CARLO COMPARISONS



- Cosmic ray muon MC run through a full CMS detector simulation
- Similarity in comparison demonstrates a good understanding of the detector

Andy Kubik (Northwestern)

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- We have built and deployed the world's largest cathode strip chamber muon spectrometer, and it works!
  - Efficient triggering
  - Improved momentum resolution
- We have been performing commissioning and performance studies on the CSCs using 12 minutes of beam halo data and 1 month of cosmic ray data
- These results show we have a strong understanding of the response of the system to real muons
- We are learning more every day how to maximize the CSCs' usefulness for the upcoming LHC beam



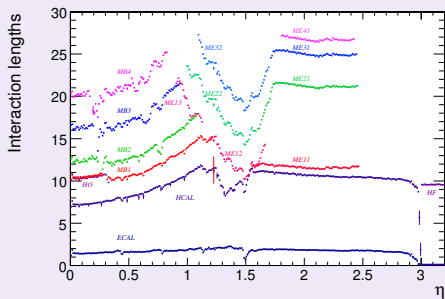
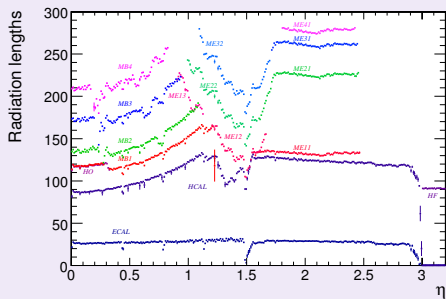
The success of this commissioning effort is due to the incredible work of the dedicated physicists, engineers, and graduate students of the CMS CSC collaboration.

Thank you.

Backup Slides



# RADIATION & INTERACTION LENGTHS



The CMS Collaboration 2006, CERN-LHCC-2006-001

