Baryogenesis from the Decays of Exotic Vector-like Squarks

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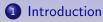
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#### Introduction

The Model Baryogenesis From Exotic Squark Decays The Cosmology Phenomenology Conclusions and Future Work

### Outline



- 2 The Model
- 3 Baryogenesis From Exotic Squark Decays
- 4 The Cosmology
- 5 Phenomenology
- 6 Conclusions and Future Work



- Determining the dynamical mechanism which generated the baryon asymmetry of the universe (BAU) is still an open question.
- The most accurate determination is given by WMAP5:

$$\eta \equiv \frac{n_B - n_{\overline{B}}}{s} = 6.225 \pm 0.170 \times 10^{-10}$$

- Any model for  $\eta$  must satisfy the Sakharov conditions:
  - Out-of-equilibrium
  - Baryon number violation
  - C and CP violation



- We will generate the BAU from the out-of-equilibrium decays of exotic vector-like squarks.
- Vector-like squarks are present in GUTs, string based models and gauge mediated SUSY breaking.
- Baryon-number and CP violation come from superpotential terms (i.e. we do not rely on SUSY breaking).



- We will generate the BAU from the out-of-equilibrium decays of exotic vector-like squarks.
- Vector-like squarks are present in GUTs, string based models and gauge mediated SUSY breaking.
- Baryon-number and CP violation come from superpotential terms (i.e. we do not rely on SUSY breaking).
- Previous related work includes the 1987 Dimopolous and Hall paper "Baryogenesis at the MeV Era."
- They use out-of-equilibrium decays of MSSM squarks to generate the BAU.
- Their universe could only reheat to  $\mathcal{O}(GeV)$  and they were constrained by electric dipole moment measurements.



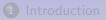
This model can exhibit the following characteristics:

- Natural splittings for the TeV scale masses of the two lightest exotic squarks.
- Maximizes the reheat temperature of the universe, thereby requiring degenerate TeV scale squarks.
- Allow the TeV exotics to exhibit displaced vertices at the LHC.
- Allow the exotics to be the messengers of gauge mediated SUSY breaking.

Please see the paper for specific benchmark parameters.



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- The relevant matter content is
  - the three generations of colored MSSM chiral superfields  $(u_i^c, d_i^c, q_i), i = 1 \dots 3,$
  - 2 families of exotic vector-like quark superfields  $(D_i, \overline{D}_i), i = 1 \dots 2.$
- There is an approximate Z<sub>2</sub> "exotic-parity."
- If this parity were exact, the lightest exotic would be stable.



• The superpotential is

$$\mathcal{W} = \mathcal{W}_{\mathrm{MSSM}} + \mathcal{W}_{\mathrm{Exotic}}$$

with

$$\mathcal{W}_{\text{Exotic}} = g'_{ijk} \, u_i^c \, D_j \, D_k + (\mu'_R)_{ij} \, d_i^c \, \overline{D}_j + \left(\frac{(\mu'_L)_{ij}}{v_d}\right) H_d \, q_i \, D_j + M_{ij} \, D_i \, \overline{D}_j$$

- In much of what follows  $\mu'_L = \mu'_R = \mu'.$
- Note that  $\mu'$  breaks exotic-parity so any interaction with an odd number of exotics is proportional to  $\mu'$ .



- Rotate to a basis without  $\mu^\prime$  mass terms.
- This mixes  $d_L$  with  $\overline{D}$  and  $d_R$  with D.
- There are new gauge interactions in this basis (including all SUSY counterparts):

$$\left(\frac{1}{2}\frac{g_w}{c_w}\frac{\mu'_L}{M}\right)(d_L)^{\dagger}\overline{\sigma}^{\mu}(\overline{D}) Z^0_{\mu} + \text{h.c}$$



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• There are superpotential interactions:

$$g' u^{c} D D$$
$$g' \left(\frac{\mu'_{R}}{M}\right) u^{c} d^{c} D$$
$$g' \left(\frac{\mu'_{R}}{M}\right)^{2} u^{c} d^{c} d^{c}$$

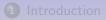


Typical values which lead to the correct value for the BAU and are consistent with the cosmological and phenomenological constraints are

- $g' \lesssim 10^{-2}$
- $\frac{\mu'}{M} \lesssim 10^{-2}$
- $M \sim 100 \,\mathrm{GeV} 10^6 \,\mathrm{GeV}$
- $\Delta M^2 \sim \mu'^2$
- $T_{\rm RH} \sim 10 \, {\rm GeV}$  for TeV scale exotics



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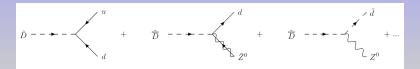


- We concentrate on representative contributions to the BAU, from g' and weak interactions.
- Need to interfere 1-loop with tree diagram.
- BAU will be generated without CP violation in soft terms.
- Define

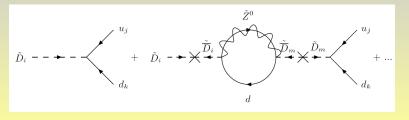
$$\epsilon \equiv \sum_{i=1}^{n} \frac{\Gamma((\tilde{D}_{\ell})_i \to u+d) - \Gamma((\tilde{D}_{\ell}^*)_i \to u^{\dagger} + d^{\dagger})}{\Gamma_{\text{total}}((\tilde{D}_{\ell})_i)}$$



• The dominant contributions to the total width are given by



• Then  $\epsilon$  is given by





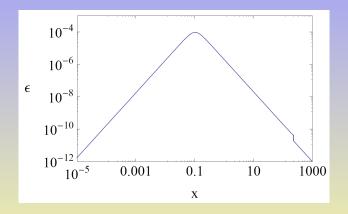
- Only the lightest exotic squarks make non-trivial contributions to  $\epsilon.$
- We can estimate

$$\epsilon \approx \frac{1}{16 \pi} \frac{g^{\prime 2} \left(\frac{g_w}{c_w}\right)^2}{\left(\frac{g_w}{c_w}\right)^2} \left(\frac{\mu^{\prime 2}}{\Delta M^2}\right)$$

where we have assumed a phase of  $\pi/2$  in one of the  $(g'\,\mu_R')$  couplings and  $g' < g_w.$ 

- Define mass splitting parameter, x, by  $(x \mu')^2 \equiv \Delta M^2$ .
- Then the  $\mu'$  suppression of  $\epsilon$  drops out.
- If  $\Delta M^2 \sim \mu'^2 \Leftrightarrow x \sim 1$ , then degenerate squarks compensate for small  $\mu'$ .





- This is a plot of  $\epsilon$  (please see paper for parameter choices).
- We have properly regulated the resonance via the width of D.
- The full expression for  $\epsilon$  is given in the paper.



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- Introduce a "inflaton" field  $\phi$ .
- Our cosmology begins in a  $\phi$  dominated phase.
- Then  $\phi$  decays to exotic squarks providing an out-of-equilibrium population given by

$$Y_D \equiv \frac{n_D}{s} \approx \mathrm{BR}\left(\frac{T_{\mathrm{RH}}}{m_{\phi}}\right),$$



- Introduce a "inflaton" field  $\phi$ .
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$$Y_D \equiv \frac{n_D}{s} \approx \mathrm{BR}\left(\frac{T_{\mathrm{RH}}}{m_{\phi}}\right),$$

• It is important that the squarks decay before they annihilate back to equilibrium.

$$\Gamma_{\text{decay}} > \Gamma_{\text{ann}}(T_{\text{RH}}).$$

where the dominant annihilation rate is  $\tilde{D} \ \tilde{D} \leftrightarrow g \ g$ 



• Then the BAU is given by

$$\eta \equiv \frac{n_B - n_{\overline{B}}}{s} = \epsilon \left(\frac{Y_D}{s}\right) = \epsilon \operatorname{BR} \left(\frac{T_{\mathrm{RH}}}{m_{\phi}}\right)$$

Note that

$$\Gamma_{\rm decay} > H(T_{\rm RH})$$

for all models considered.

- This means that the squarks decay instantaneously, i.e. at a temperature  $T_{\rm RH}.$ 



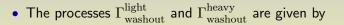
- We must be sure that the BAU is not erased.
- There are three dominant washout processes.
- We must check that the following conditions are satisfied

$H(T_{\rm RH})$	>	$\Gamma_{ m ID}(T_{ m RH})$
$H(T_{\rm RH})$	>	$\Gamma_{\rm washout}^{\rm heavy}(T_{\rm RH})$
$H(T_{\rm RH})$	>	$\Gamma_{\rm washout}^{\rm light}(T_{\rm RH})$

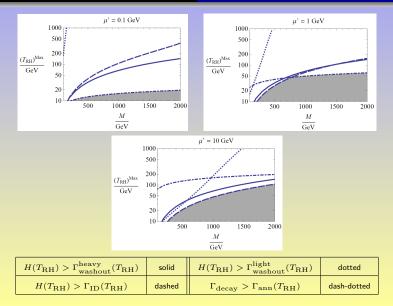


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For the scenario where the exotic squarks are the messengers of gauge mediation:

- There are new flavor violating contributions to the MSSM up-squark masses proportional to g'.
- There is the possibility of proton decay via  $p^+ \rightarrow K^+ + \tilde{G}$ .



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In general

- There are constraints from unitarity of the CKM matrix.
- There are constraints from  $D^0 \overline{D}^0$  mixing.
- Electric dipole moments vanish at 1-loop and the 2-loop effect is smaller then from the standard model.
- There is the spectacular LHC signal of long-lived tracks with baryon number violating decays.

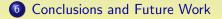


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#### Conclusions

- We have realized the BAU via the out-of-equilibrium decays of exotic vector-like squarks.
- We explored a variety of tests and predictions for this class of models.
- The exotic squarks can be the messengers of gauge mediation.



#### Conclusions

- We have realized the BAU via the out-of-equilibrium decays of exotic vector-like squarks.
- We explored a variety of tests and predictions for this class of models.
- The exotic squarks can be the messengers of gauge mediation. Future Work
  - Do a detailed calculation with specific g' and  $\mu'$  matrices.
  - Work out details when quarks are the lightest exotic states.
  - Find family symmetry when degenerate squarks are required.
  - Build a DM sector.
  - Do a full collider simulation to determine the viability of discovering Baryon-number violation at the LHC.



# THANK YOU



# Are there any questions?



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# **Backup Slides**



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Large $\tilde{D}$ mass splittings								
scenario	g'	$\frac{M}{(\text{GeV})}$	$\frac{\mu'}{(\text{GeV})}$	$\frac{\Delta M^2}{(\text{GeV})^2}$	$\epsilon$			
large splittings	0.4	500	4	$(57\mu')^2$	$2 \times 10^{-6}$			
Degenerate $\tilde{D}$ masses								
scenario	g'	$\frac{M}{(\text{GeV})}$	$\frac{\mu'}{(\text{GeV})}$	$\frac{\Delta M^2}{(\text{GeV})^2}$	$\epsilon$			
high $T_{\rm RH}$	0.005	1000	2		$8 \times 10^{-6}$			
displaced vertices	0.06	500	$10^{-5}$		$6  imes 10^{-3}$			
gauge mediation	0.01	$10^{6}$	1	$(0.4\mu')^2$	$6 \times 10^{-5}$			

I and D man and this an								
Large D mass splittings								
scenario	$\frac{T_{\rm RH}}{{\rm GeV}}$	$\frac{H}{\text{GeV}}$	$\frac{\Gamma_{\text{decay}}}{\text{GeV}}$	$\frac{\Gamma_{ann}}{\text{GeV}}$	$\frac{\Gamma_{\text{ID}}}{\text{GeV}}$	$\frac{\Gamma_{\text{washout}}^{\text{heavy}}}{\text{GeV}}$	$\frac{\Gamma_{\text{washout}}^{\text{light}}}{\text{GeV}}$	$\frac{m_{\phi}}{\text{GeV}}$
large splittings	18	$4 \times 10^{-16}$	$8 \times 10^{-4}$	$3 \times 10^{-6}$	$8 \times 10^{-17}$	$2 \times 10^{-24}$	$1\times 10^{-17}$	5000
Degenerate $\tilde{D}$ masses								
scenario	$\frac{T_{\rm RH}}{{\rm GeV}}$	$\frac{H}{\text{GeV}}$	$\frac{\Gamma_{\text{decay}}}{\text{GeV}}$	$\frac{\Gamma_{ann}}{\text{GeV}}$	$\frac{\Gamma_{\text{ID}}}{\text{GeV}}$	Γ <sup>heavy</sup> GeV	$\frac{\Gamma_{\text{washout}}^{\text{light}}}{\text{GeV}}$	$\frac{m_{\phi}}{\text{GeV}}$
high $T_{\rm RH}$	75	$8 \times 10^{-15}$	$4 \times 10^{-5}$	$1 \times 10^{-5}$	$1 \times 10^{-15}$	$4 \times 10^{-16}$	$6\times 10^{-20}$	$10^{5}$
displaced vertices				$2 \times 10^{-16}$	$\sim 0$	$\sim 0$	$\sim 0$	$10^{5}$
gauge mediation	1000	$1 \times 10^{-12}$	$1 \times 10^{-8}$	$1 \times 10^{-9}$	$\sim 0$	$\sim 0$	$\sim 0$	$10^{7}$



• We must assume a texture in the g' and  $\mu'$  matrices for the "large splittings" benchmark.

- One could imagine that the exotic squarks are the messengers of gauge mediated SUSY breaking.
- There are new contributions to the MSSM up-squark masses at 1-loop and 2-loops which can lead to FCNCs:

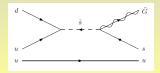
$$(\delta \tilde{m}_{u_R}^{1-\text{loop}})_{ij}^2 = -\frac{1}{8\pi^2} g'_{ikm} g'^*_{jkm} \frac{F_X^4}{M_X^6} (\delta \tilde{m}_{u_R}^{2-\text{loop}})_{ij}^2 \approx -\frac{1}{(16\pi^2)^2} g'_{ikm} g'^*_{jkm} g_s^2 \frac{F_X^2}{M_X^2}$$



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• There is also the possibility of proton decay via





There are a variety of phenomenological constraints one should consider:

• Unitarity of the CKM matrix:

$$|V_{ud}|^2 + |V_{us}|^2 + |V_{ub}|^2 + N\left(\frac{\mu'}{M}\right)^2 = 0.9999 \pm 0.001 \Rightarrow \frac{\mu'}{M} \lesssim 0.03$$

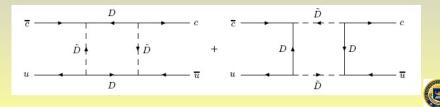


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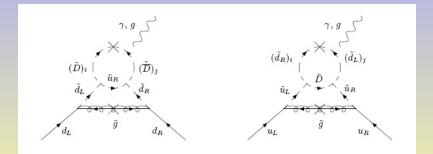
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• 
$$D^0 - \overline{D}^0$$
 mixing constrains  $g'$ :

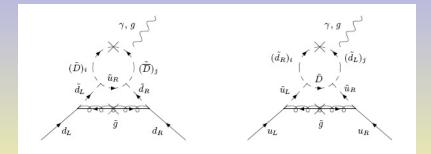


• EDMs are only generated at 2-loops and are sim4 orders smaller then the SM contribution:





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There are potential LHC signals to determine Baryon-number violating decays.

