

R-symmetric Gauge Mediation and the MRSSM

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Based on work with
S. De Lope Amigo, P. Fox and E. Poppitz
arXiv:0809.1112, ...

A Puzzle

- With all the new operators, the MSSM has 124 parameters – and that's the MINIMAL model!
- Many of those parameters are flavor mixing angles and phases, implying large FCNC's and CP violation. This is called the **SUSY Flavor/CP Puzzle**.
- We need some principle to eliminate this flavor violation to avoid conflicts with observation and reduce the parameter space to something manageable for experiments.

A New Solution

- An R-symmetry rotates fields within a supermultiplet differently.
- Kribs, Poppitz, Weiner (arXiv:0712.2039) found that by imposing an additional R-symmetry to the MSSM you can have sizable flavor-violating operators while not generating large FCNC's or CP violation, as long as gluinos are heavy.
- We impose a $U(1)_R$ symmetry, although a discrete symmetry would work as well.

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 - No A-terms for the scalars; hence no left-right squark/slepton mixing.
 - No μ -term, but there is a B-term (complicated Higgs sector).

K-Kbar Mixing

Strongest constraint in SUSY flavor physics.

Parametrize mixing: $\delta_L \equiv \frac{m_{\tilde{Q}12}^2}{M_{\tilde{q}}^2}$ $\delta_R \equiv \frac{m_{\tilde{d}12}^2}{M_{\tilde{q}}^2}$

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The Low-Energy Effective Lagrangian is:

$$\mathcal{L}_{\text{eff}} = \frac{\alpha_s^2(M_{\tilde{g}})}{216} \left(\frac{M_{\tilde{q}}^2}{M_{\tilde{g}}^4} \right) \sum_n C_n(\mu) \mathcal{O}_n(\mu)$$

$$\mathcal{O}_1 = (\bar{d}_L^i \gamma^\mu s_L^i) (\bar{d}_L^j \gamma_\mu s_L^j)$$

$$\mathcal{O}_4 = (\bar{d}_R^i s_L^i) (\bar{d}_L^j s_R^j)$$


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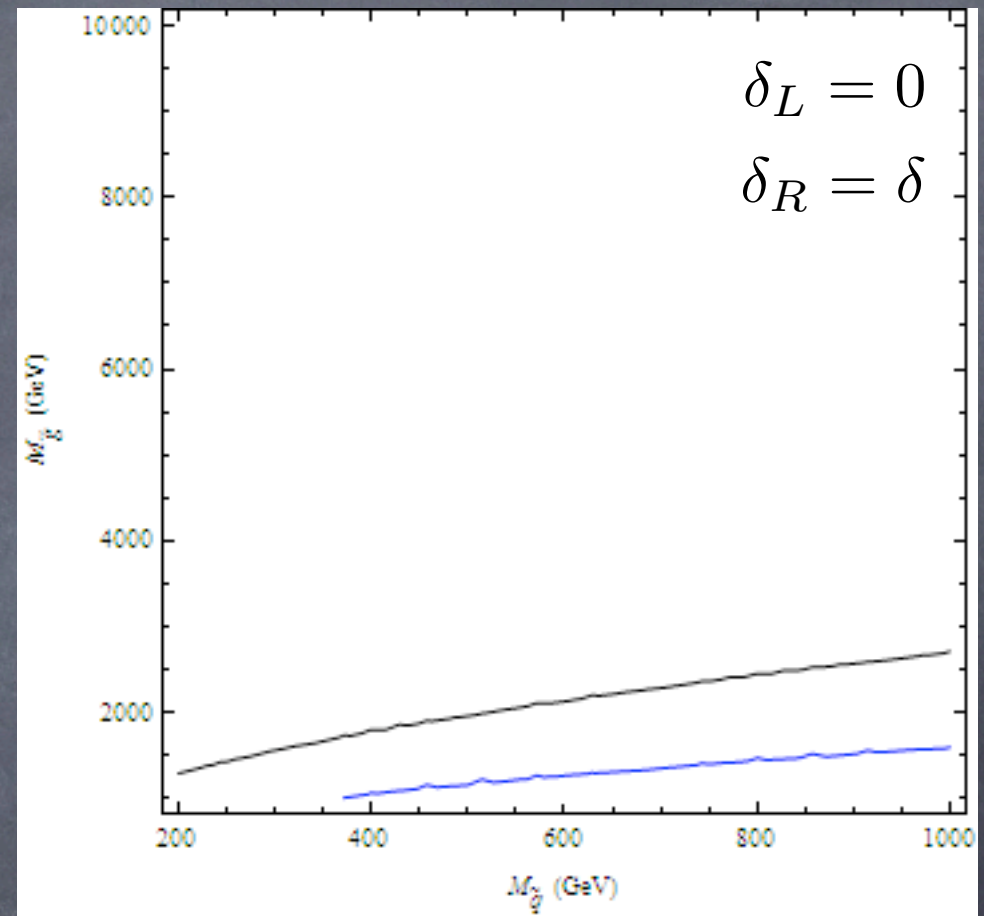
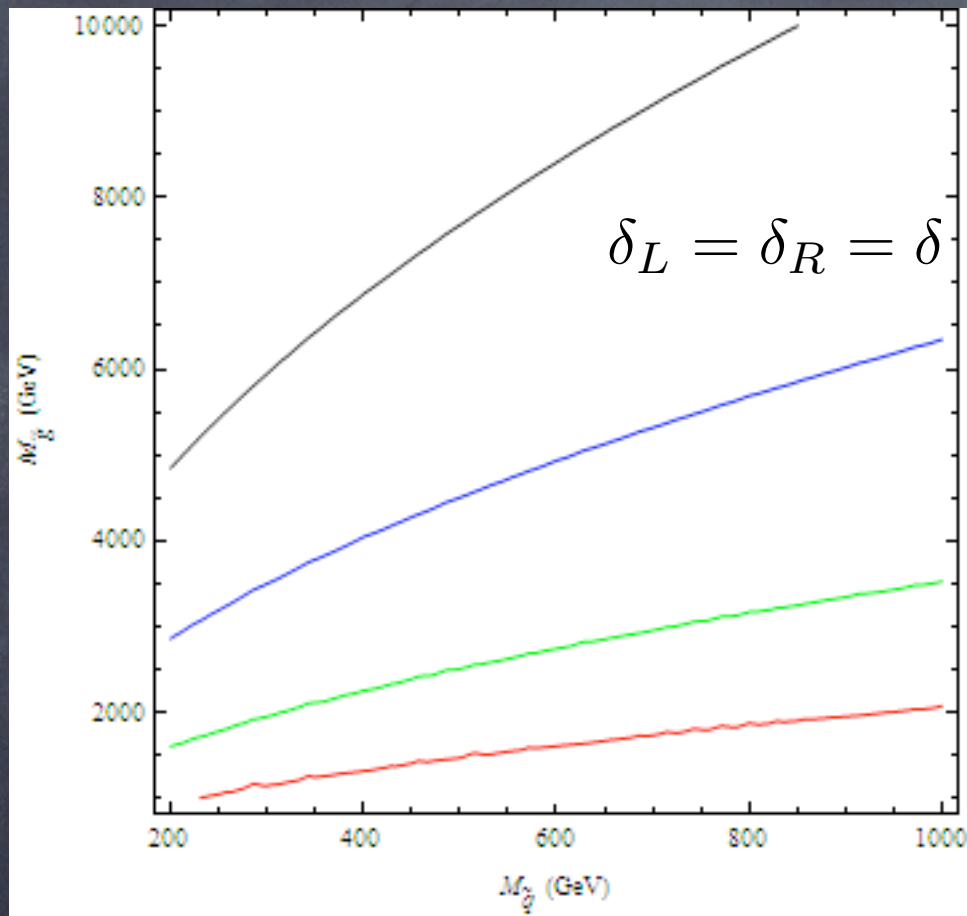
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QCD Corrections computed in
A.B., S.-P. Ng, arXiv:0803.3811

K-Kbar Mixing



- Red line: $\delta = 0.01$
- Green line: $\delta = 0.03$
- Blue line: $\delta = 0.1$
- Black line: $\delta = 0.3$

A Problem...

- Anytime you spontaneously break SUSY, you have a goldstino and a cosmological constant.
- When turning on supergravity, this goldstino is "eaten" by the gravitino (gauge field of SUSY) and it gains a mass which is R-violating (at least in N=1 SUSY).
- This mass term feeds into the MSSM through anomaly mediation, so it looks like this model can never be realized...

... A Solution!

- Gauge mediation has a very light gravitino, typically around 1 keV or so, rendering these troublesome effects irrelevant.
- Ordinary gauge mediation breaks the R-symmetry, but if we can find a framework where we can maintain the symmetry through a gauge-mediation-like mechanism, then we might realize the MRSSM naturally...

A Minimal Model

- Motivated originally by the “ISS” Models, we need a SUSY-breaking **AND** an R-preserving messenger scalar mass, as well as an R-preserving chirality flip.
- We can seek to capture these effects in a minimal model with a pair of vector-like chiral messenger superfields, and a single chiral superfield in the adjoint representation of the gauge group:

$$W_{mess} = \sum_{i=1}^{N_{mess}} (\Xi \bar{\varphi}^i \varphi^i + M_{mess} \bar{\varphi}^i N^i + M_{mess} \bar{N}^i \varphi^i + y \bar{\varphi}^i \Phi N^i - y \bar{N}^i \Phi \varphi^i)$$

The SUSY-breaking spurion is $\Xi = \theta^2 z M_{mess}^2$

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Note: with all this matter, QCD will develop a Landau Pole Λ_3

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- IR ("Gauge Mediation") contributions.

UV Contributions: The Size of Λ

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- $\Lambda \sim M_P$: UV operators are irrelevant; does not realize the large flavor-violating operators of the MRSSM, so we will not consider it here.
- $\Lambda \sim \frac{\Lambda_3}{4\pi}$: UV operators are important – this is the maximal size of the operators (using NDA). We will consider this case, but it could overestimate the size of these contributions.

UV Contributions

UV Dirac Gaugino Mass:

$$\int d^2\theta \frac{1}{\Lambda^3} (W^\alpha \Phi) \bar{D}^2 D_\alpha (\Xi^\dagger \Xi) \quad m_{1/2} \sim M_{UV} \left(\frac{M_{UV}}{\Lambda} \right)$$

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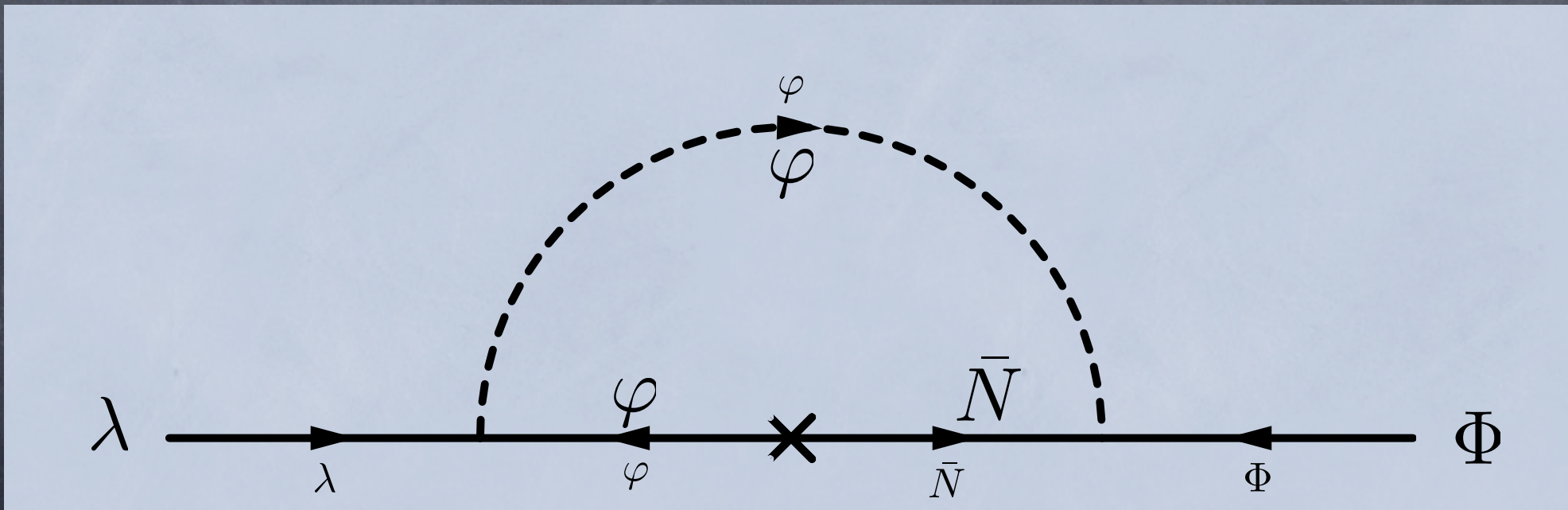
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UV Scalar Mass:

$$\int d^4\theta \frac{c_{ij}}{\Lambda^2} (\Xi^\dagger \Xi) Q_i^\dagger Q_j \quad m_{0\ ij} \sim M_{UV}$$

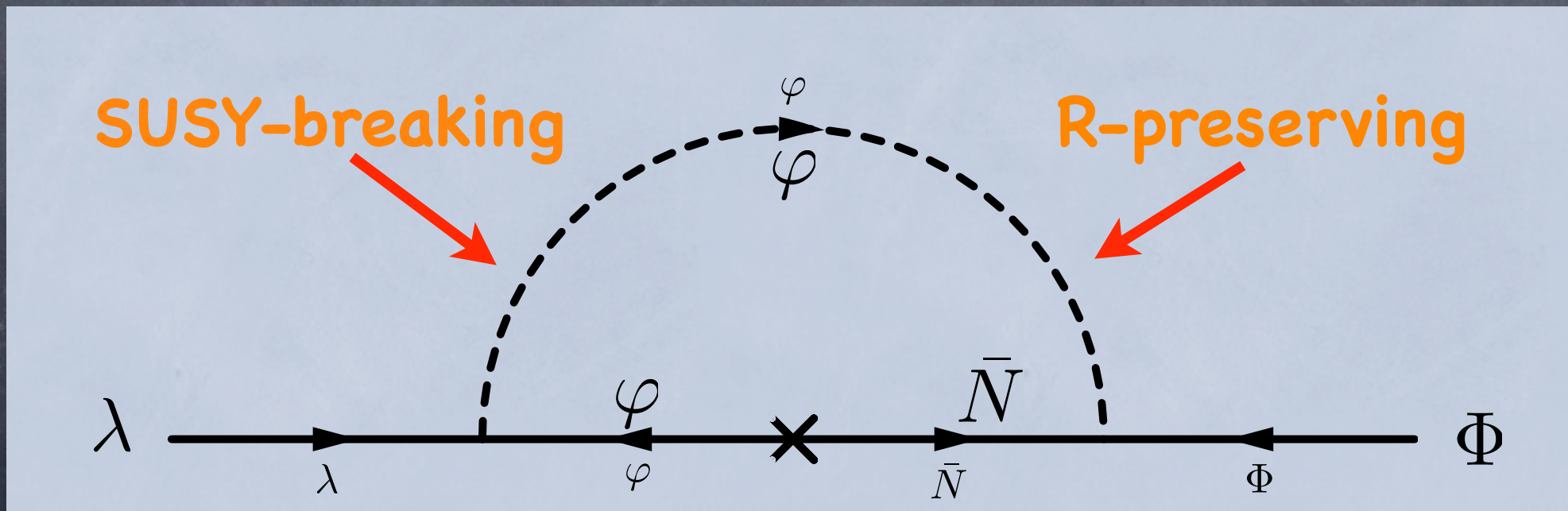
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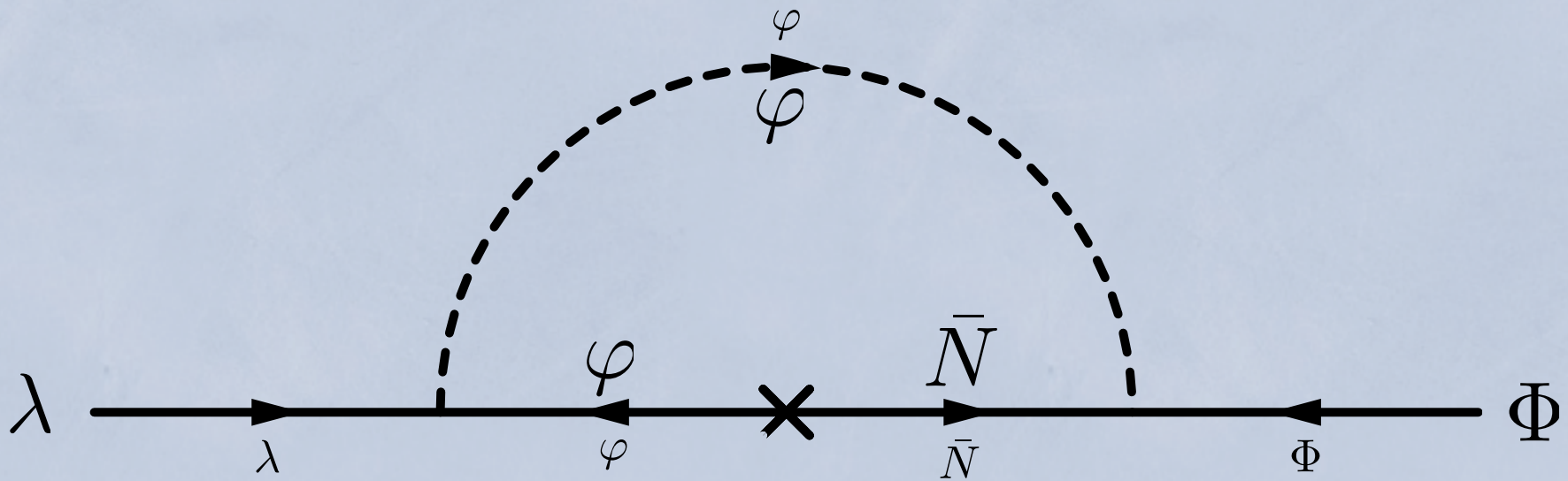
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This diagram gives:

Yukawa dependent!

$$m_{1/2} = \frac{gy}{16\pi^2} M_{\text{mess}} R(z) \cos\left(\frac{\langle \xi \rangle}{v}\right)$$

where we have the new function:

$$R(z) = \frac{1}{z} [(1+z) \log(1+z) - (1-z) \log(1-z) - 2z]$$

IR Contributions: Scalar Masses

Identical to Gauge Mediation with **one** messenger:

$$m_0^{(IR)2} = 2C_F^{(a)} \left(\frac{\alpha_a}{4\pi} \right)^2 M_{\text{mess}}^2 F(z)$$

where $F(z)$ is the usual GM function:

$$F(z) = (1+z) \left[\log(1+z) - 2\text{Li}_2 \left(\frac{z}{1+z} \right) + \frac{1}{2}\text{Li}_2 \left(\frac{2z}{1+z} \right) \right] + (z \rightarrow -z)$$

IR Contributions:

Adjoint Scalar Masses

- Messenger loops will also generate masses for the adjoint scalars (at **ONE** loop!):

$$m_{\phi}^2 = \frac{y^2}{16\pi^2} M_{\text{mess}}^2 R_s(z)$$

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- This leads to the prediction:

$$\frac{m_\phi}{m_{1/2}} \sim \frac{\sqrt{|B_\phi|}}{m_{1/2}} \sim \sqrt{\frac{4\pi}{\alpha}}$$

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- Recall that the MRSSM needed gauginos heavier than squarks by a factor of 5.

However:

$$\frac{m_{1/2}}{m_0^{IR}} = \frac{1}{\sqrt{2C_F}} \left(\frac{y}{g} \right) \left(\frac{R(z)}{\sqrt{F(z)}} \right)$$

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- This ratio function of z is strictly less than unity, so to make gauginos heavy requires a large Yukawa.

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- Both the gaugino mass and the squark mass squared are proportional to N_{mess} . Thus the gaugino:squark mass ratio $\propto \sqrt{N_{\text{mess}}}$.

Sample Spectrum 1

$$M_{\text{mess}} = 80 \text{ TeV} \quad N_{\text{mess}} = 6 \quad y = 3 \quad z = 0.4$$
$$\tan \beta = 5 \quad \tan \nu = 1 \quad \lambda = 0$$

All masses in GeV:

$$M_1 = 257 \quad M_2 = 370 \quad M_3 = 612$$

$$m_{\tilde{e}} = 130 \quad m_{\tilde{t}} = 986 \quad |\mu| = 447 \quad \sqrt{B_\mu} = 200$$

$$m_h = 80 \longrightarrow 134 \quad \text{tuning} = 4\%$$

$$m_{\text{adj}} = 4 \text{ TeV} \quad m_R = 312$$

Sample Spectrum 2

$$M_{\text{mess}} = 80 \text{ TeV} \quad N_{\text{mess}} = 6 \quad y = 3 \quad z = 0.95$$
$$\tan \beta = 5 \quad \tan \nu = 1 \quad \lambda = 0$$

All masses in GeV:

$$M_1 = 2100 \quad M_2 = 3000 \quad M_3 = 4900$$
$$m_{\tilde{e}} = 296 \quad m_{\tilde{t}} = 2200 \quad |\mu| = 800 \quad \sqrt{B_\mu} = 422$$
$$m_h = 54 \longrightarrow 152 \quad \text{tuning} = 0.5\%$$
$$m_{\text{adj}} = 9 \text{ TeV} \quad m_R = 693$$

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- To satisfy flavor constraints, there is a tuning of the off-diagonal mass terms in the UV contribution.

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Thus: $c_D \sim 1$ for General Models

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- From this formula, it is clear that it is difficult to avoid tuning.
- Using the results from $K - \bar{K}$ mixing:

$$m_0 = 1\text{TeV} \quad \delta = 0.07$$

$$t = 2.7\%$$

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- RGM: A modified SUSY-breaking mediator that provides a new and unique spectrum, both through what it allows, and the nature of the mass spectrum.
- For RGM to realize MRSSM, it is possible but requires a better understanding of the UV theory to avoid tuning.

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The End!