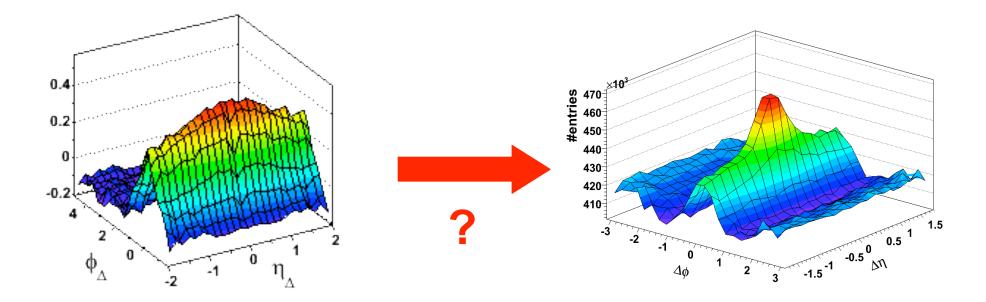
Soft Contribution to the Hard Ridge

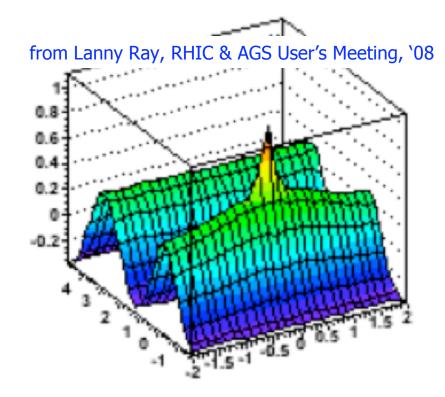


George Moschelli Sean Gavin DPF July 31, 2009 Wayne State University

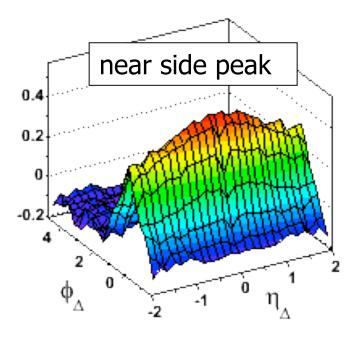
Outline:

- The soft ridge and bulk correlations
- p_t dependence on bulk correlations
- Jet correlations with bulk
- Relative contribution of bulk- bulk and jet-bulk correlations
- Summary

STAR: Soft Ridge



- Correlation function with no p_t trigger \Rightarrow near side peak
- Rapidity width and amplitude increase with centrality
- Azimuthal width roughly constant



Hard vs. Soft Ridge

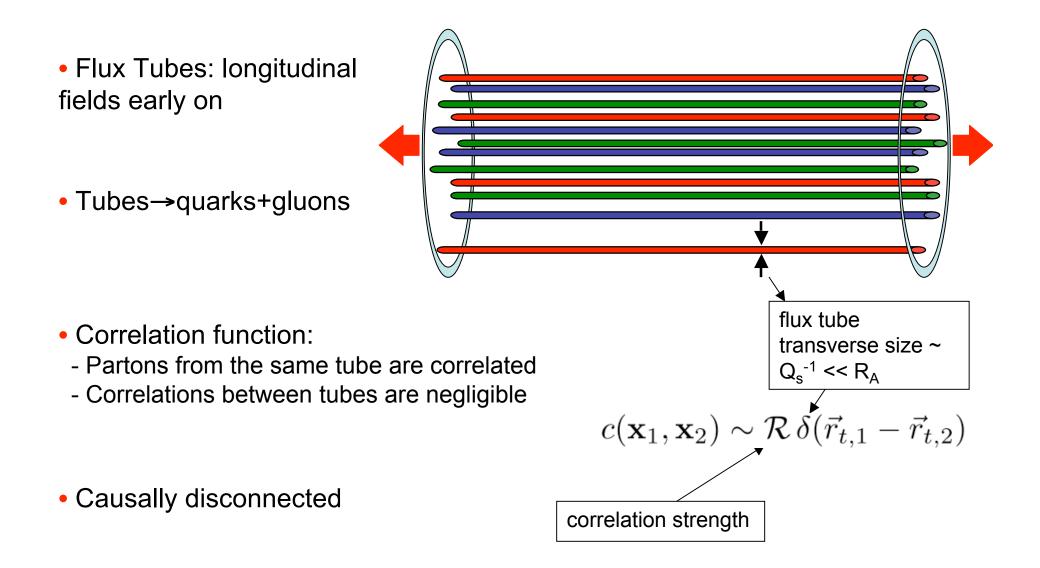
hard ridge explanations -- jet interactions with matter

- N. Armesto, C.A. Salgado, U.A. Wiedemann, Phys. Rev. Lett. 93, 242301 (2004)
- P. Romatschke, Phys. Rev. C 75, 014901 (2007)
- A. Majumder, B. Muller, S. A. Bass, Phys. Rev. Lett. 99, 042301 (2007)
- C. B. Chiu, R. C. Hwa, Phys. Rev. C 72, 034903 (2005)
- C. Y. Wong, arXiv:0712.3282 [hep-ph]
- R. C. Hwa, C. B. Yang, arXiv:0801.2183 [nucl-th]
- T. A. Trainor, arXiv:0708.0792 [hep-ph]
- A. Dumitru, Y. Nara, B. Schenke, M. Strickland, arXiv:0710.1223 [hep-ph]
- E. V. Shuryak, Phys. Rev. C 76, 047901 (2007)
- C. Pruneau, S. Gavin, S. Voloshin, Nucl.Phys.A802:107-121,2008

soft ridge -- similar but no jet -- collective behavior

- S. Gavin and M. Abdel-Aziz, Phys. Rev. Lett. 97, 162302 (2006)
- S. A. Voloshin, Phys. Lett. B 632, 490 (2006)
- S. Gavin and G. Moschelli, arXiv:0806.4366 [nucl-th]
- A. Dumitru, F. Gelis, L. McLerran and R. Venugopalan, arXiv:0804.3858 [hep-ph]
- S. Gavin, L. McLerran, G. Moschelli, arXiv:0806.4718 [nucl-th]
- F. Gelis, T. Lappi, R. Venugopalan, arXiv:0807.1306 [hep-ph]

Flux Tubes and Glasma



Flux Tubes and Correlations

•Gluon rapidity density

Kharzeev & Nardi:

$$\frac{dN}{dy} \propto \alpha_s^{-1}(Q_s)Q_s^2 R_A^2$$

Correlation Strength

$$\mathcal{R} \propto \langle \# tubes \rangle^{-1} = (Q_s R_A)^{-2}$$

• Long range **Glasma** fluctuations scale the phase space density

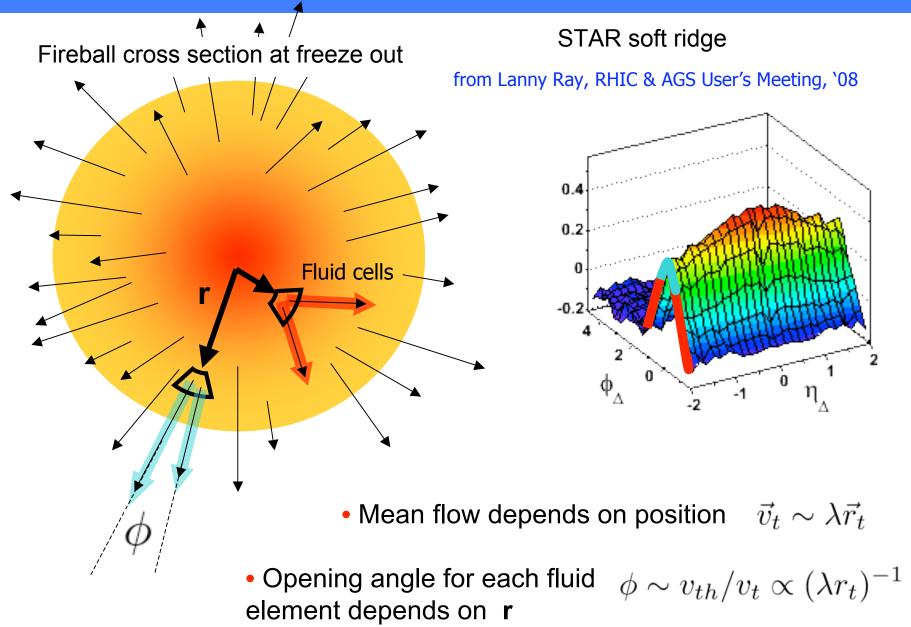
$$\mathcal{R}\frac{dN}{dy} \propto \alpha_s^{-1}(Q_s^2)$$

Dumitru, Gelis, McLerran & Venugopalan; Gavin, McLerran & Moschelli

 Energy and centrality dependence of correlation strength

$$Q_s^2 \propto (\sqrt{s})^{1/3} (N_{part})^{1/3}$$

Flow and Azimuthal Correlations



Blast Wave and the Correlation Function

Schnedermann, Sollfrank & Heinz

Single Particle Spectrum

Correlation Function

$$\rho_{1}(\vec{p}) \equiv \frac{dN}{dyd^{2}p_{t}} = \int_{freezout} f(t) dt$$

$$\vec{x}, \vec{p}) \\ \swarrow \\ \gamma_t \vec{v}_t = \lambda \vec{r}$$

A Hubble like expansion in used in a Boltzmann Distribution

 $\Delta \rho(\vec{p}_1, \vec{p}_2) \equiv pairs - (singles)^2$

$$\Delta \rho(\vec{p}_1, \vec{p}_2) = \iint_{freezout} c(\vec{x}_1, \vec{x}_2) f(\vec{x}_1, \vec{p}_1) f(\vec{x}_2, \vec{p}_2)$$

Normalization

$$\iint_{momenta} \Delta \rho = \iint_{positions} c = \langle N \rangle^2 \mathcal{R}$$

Glasma + Blast Wave

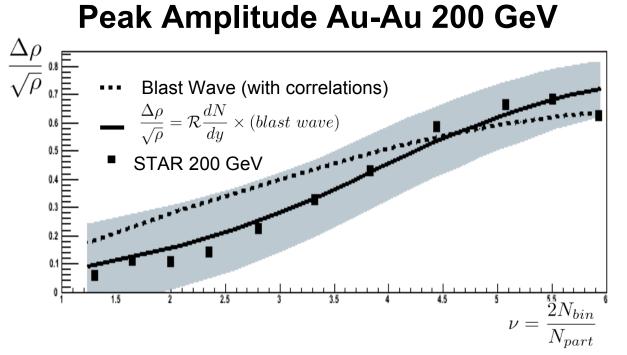
STAR measures:

We calculate:

$$\frac{\Delta\rho}{\sqrt{\rho_{ref}}} = \frac{\rho_2 - \rho_{2,mixed}}{\sqrt{\rho_{2,mixed}}}$$

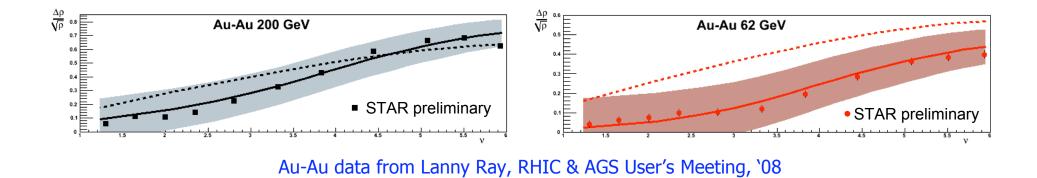
$$\Delta\rho(\eta,\phi) = \iint_{momenta} \Delta\rho(\vec{p_1},\vec{p_2})$$

$$\rho_{ref} = \iint_{momenta} \rho_1(\vec{p_1})\rho_1(\vec{p_2})$$

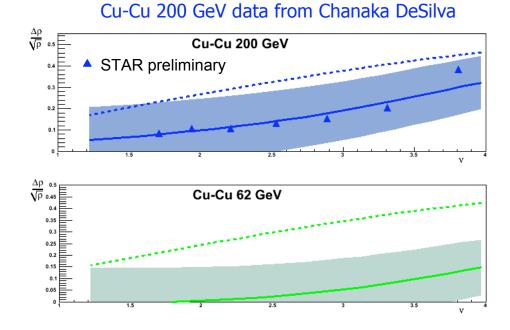


- BW scaled to fit 200 GeV
- v and T from Blast Wave Akio Kiyomichi, PHENIX
- ~10% uncertainty from BW
- uncertainty from Q_s greatest in peripheral collisions

Energy and System Dependence



- Dashed lines from Blast Wave Centrality and some energy dependence on blast wave parameters (v and T)
- CGC adds energy and centrality dependence and now system dependence
- Theoretical Cu-Cu centrality and BW parameterizations follow Au-Au



Angular Correlations

 $\frac{\Delta \rho}{\sqrt{
ho}}$ 0.6 Fit using Gaussian + offset 0.4 0.2 0 • Range: $-\pi/2 < \phi < \pi/2$ 0.5 1.5 0 1 2 ϕ Error band: 20% shift $\sigma_r^{1.3}$ STAR preliminary in fit range 1.2 Au-Au 200 GeV 1.1 • Au-Au 62 GeV Computed angular 1 0.9 width is approximately 0.8 independent of energy 0.7 0.6 0.5 1.5 2 2.5 3.5 4.5 5.5 ν

Cu-Cu Angular Correlations

0.6

0.4

0.2

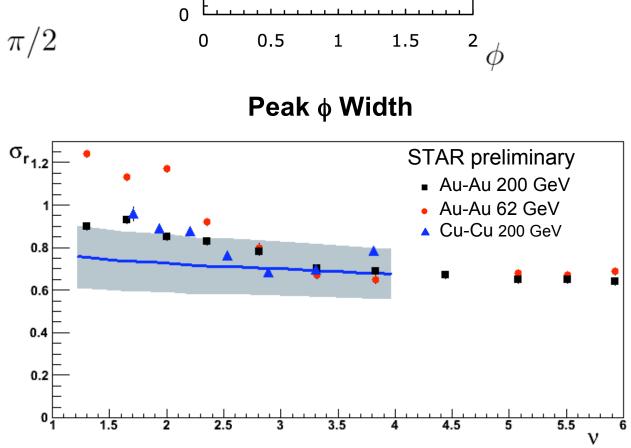
 $\frac{\Delta\rho}{\sqrt{\rho}}$

• Fit using Gaussian + offset

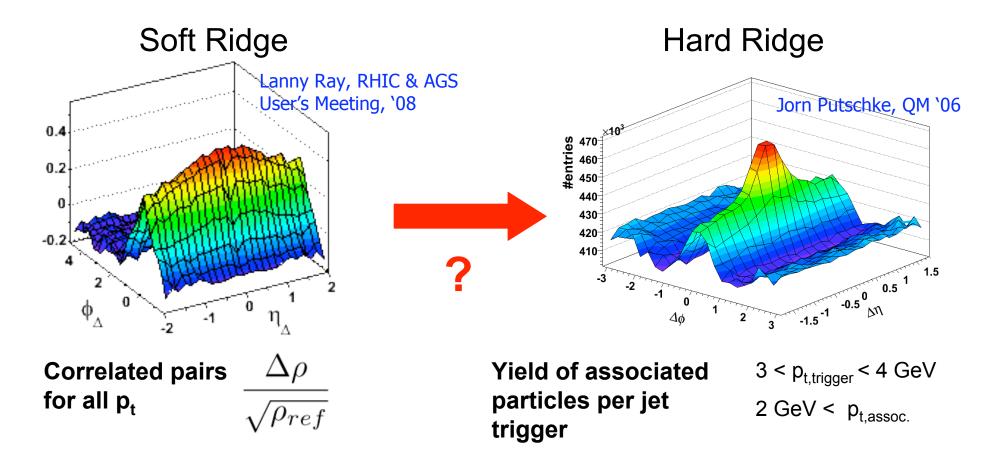
• Range:
$$-\pi/2 < \phi < \pi/2$$

• Error band: 20% shift in fit range

 Computed angular width is approximately independent of energy



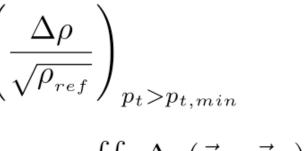
Are the soft and hard ridges related?



- Soft Ridge: no trigger, pairs not separated by p_t
- Hard Ridge: only high p_t particles, trigger and associated in different ranges
- Look for a change in the soft ridge as the minimum p_t range is increased

Soft Ridge vs. p_t

Peak Amplitude Au-Au 200 GeV Δρ Min p. limit all STAR preliminary 150 0.6 0.5 0.4 1300 1500 0.3 0.2 0.1 0 3.5 2.5 3 5.5 4.5 ν Most Central Amplitude vs. pt $\frac{\Delta \rho}{\sqrt{\rho}}$ 0.8 0.7 0.6 0.5 0.4 0.3 0.2 0.1 °0[□] 1.5 0.5 1 2 2.5 3.5



$$= \frac{\iint_{p_{t,min}} \Delta \rho(\vec{p}_{t1}, \vec{p}_{t2})}{\{\iint_{p_{t,min}} \rho_{1}(\vec{p}_{t1}) \rho_{1}(\vec{p}_{t2})\}^{1/2}}$$

 Increase the lower p_t limit of the soft ridge calculation toward the hard ridge range.

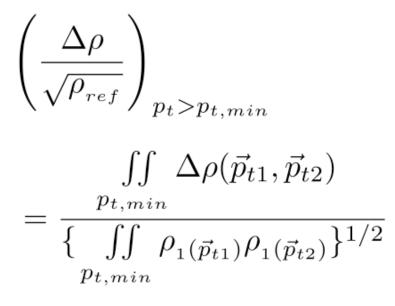
• As the lower p_t limit is increased less particles are available for correlations.

 Correlation amplitude for the most central collision plotted vs. the lower p_t limit.

p,

Soft Ridge vs. p_t

Peak Amplitude Au-Au 200 GeV Δρ Min p. limit all STAR preliminary 150 300 0.6 0.5 0.4 1300 1500 0.3 0.2 0.1 0 2.5 3 3.5 4.5 5 5.5 Most Central Amplitude vs. pt $\frac{\Delta \rho}{\sqrt{\rho}}$ 0.6 0.5 0.4 0.3 0.2 **0.1**⊟ 0 0.5 3.5 p,

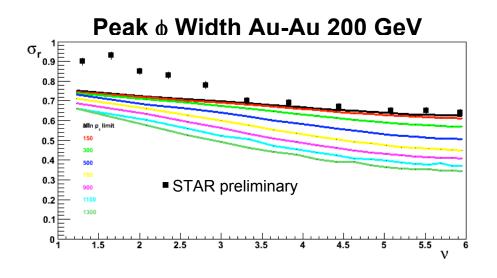


 Increase the lower p_t limit of the soft ridge calculation toward the hard ridge range.

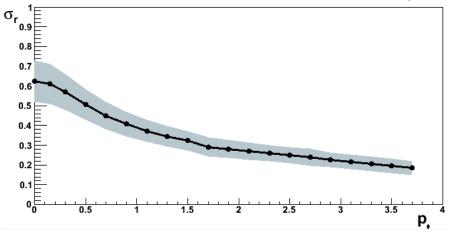
• As the lower p_t limit is increased less particles are available for correlations.

 Correlation amplitude for the most central collision plotted vs. the lower p_t limit.

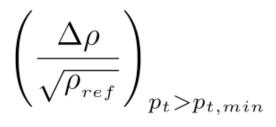
Soft Ridge vs. p_t



Most Central ϕ Width vs. p_t



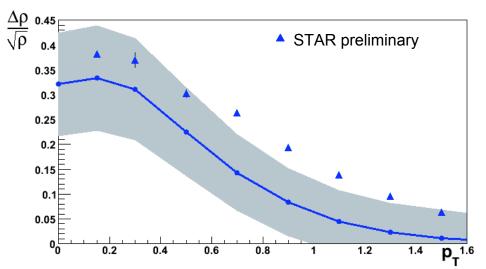
Angular width from



• Higher p_t particles received a larger radial push \Rightarrow narrower relative angle.

Cu-Cu Soft Ridge vs. p_t

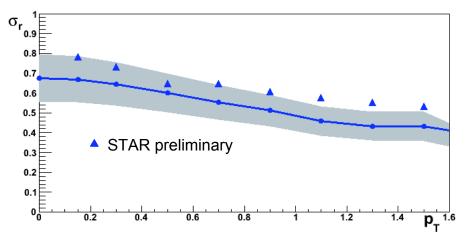
Cu-Cu 200 GeV Most Central Amplitude vs. pt



 Bulk correlations alone cannot explain the data at higher p_t

• The contribution from Jet-Bulk and Jet-Jet correlations should have an increasing effect with p_t

Cu-Cu 200 GeV Most Central ϕ Width vs. p_t



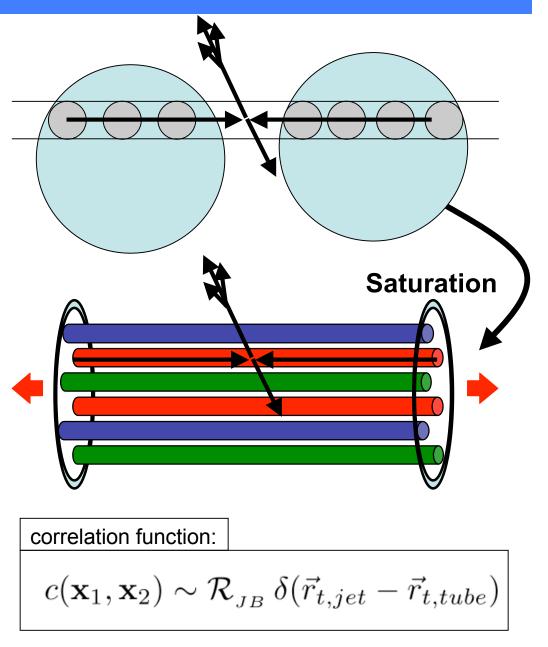
 \bullet The azimuthal width narrows with increasing $\ensuremath{\textbf{p}}_{t}$

 Jet contributions should force the correlation width to approach the jet correlation width

Jet Correlations with Flux Tubes

 Likelihood of having a jet depends on nuclear density (which depends on transverse position) and the momentum of colliding partons

- Low momentum fraction partons form CGC and tubes after the collision
- Correlations between a hard interaction and a tube at the same radial position



Quenching + Flow

- Surviving jets tend to be more radial, due to quenching.
- Jet path

$$L(r,\phi_{jet}) = \sqrt{R_A^2 - r^2 \sin^2(\phi_{jet}) - r \cos(\phi_{jet})}$$

E. Shuryak, Phys. Rev. C 76, 047901 (2007)

quenched away jet

 φ_{jet}

R

 ϕ_{bulk}

• Survival probability
$$S(\vec{x_1}, \vec{x_2}) = e^{\frac{-L(r, \phi_{jet})}{\rho\sigma}}$$

• Production probability
$$P_{prod}(r) \propto \left(1 - rac{r^2}{R^2}\right)$$

• Jet Distribution $f(\vec{x}, \vec{p}) = \frac{A}{p^n} P_{prod}(r) S(r, \phi_{jet})$

Jets and the Correlation Function

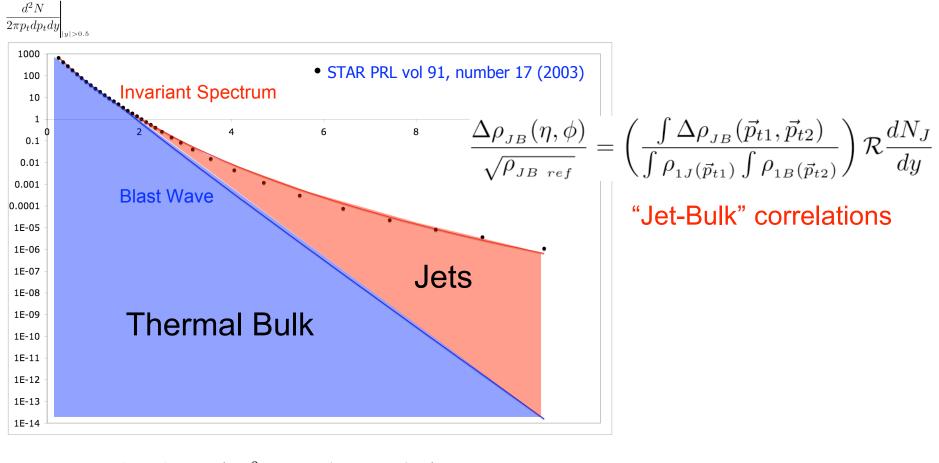
• Jet-Bulk Correlations

 $\Delta \rho_{JB}(\vec{p}_1, \vec{p}_2) \equiv (\text{pairs}) - (\text{Jet singles})(\text{Bulk singles})$

$$\Delta \rho_{JB}(\vec{p}_1, \vec{p}_2) = \iint_{freezout} c_{JB}(\vec{x}_1, \vec{x}_2) f_J(\vec{x}_1, \vec{p}_1) f_B(\vec{x}_2, \vec{p}_2)$$

• Normalization
$$\int_{momenta} \Delta \rho_{_{JB}} = \int_{positions} c_{_{JB}} = \langle N_{_J} \rangle \langle N_{_B} \rangle \mathcal{R}_{_{JB}}$$

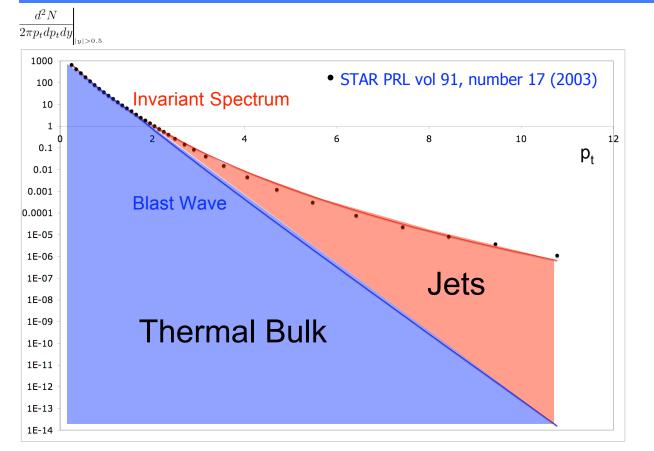
Two Contributions



$$\frac{\Delta \rho_{BB}(\eta, \phi)}{\sqrt{\rho_{BB\ ref}}} = \left(\frac{\int \Delta \rho_{BB}(\vec{p}_{t1}, \vec{p}_{t2})}{\int \rho_{1B}(\vec{p}_{t1}) \int \rho_{1B}(\vec{p}_{t2})}\right) \mathcal{R} \frac{dN_B}{dy}$$

"Bulk-Bulk" correlations

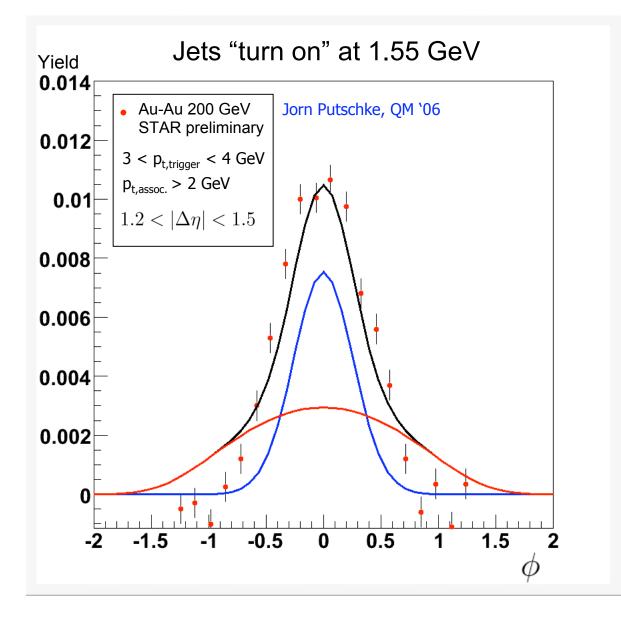
Two Contributions



• Both the Bulk-Bulk and Jet-Bulk contributions are weighted by the fraction of bulk or jet particles to the total.

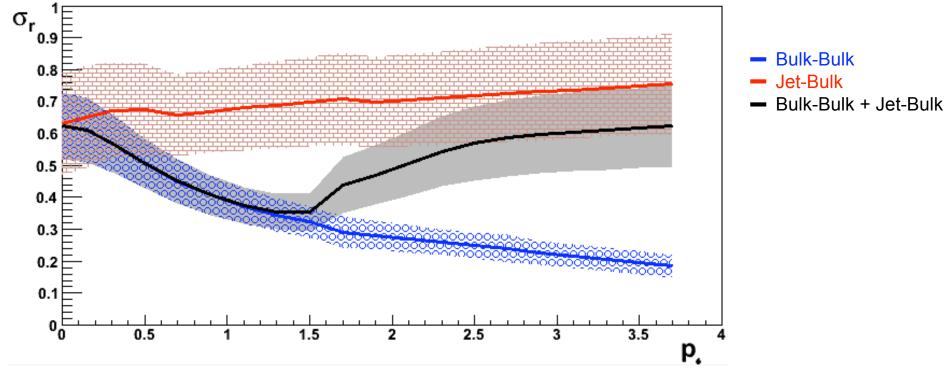
$$\frac{\Delta\rho(\eta,\phi)}{\sqrt{\rho_{ref}}} = \frac{\Delta\rho_{BB}(\eta,\phi)}{\sqrt{\rho_{BB\ ref}}} \left(\frac{Bulk}{fraction}\right) + \frac{\Delta\rho_{JB}(\eta,\phi)}{\sqrt{\rho_{JB\ ref}}} \left(\frac{Jet}{fraction}\right)$$

Hard Ridge



- Bulk-Bulk Yield
- Jet-Bulk Yield
- Bulk-Bulk + Jet-Bulk Yield
- Bulk-Bulk correlations still significant, but too narrow.
- Jet-Bulk correlations too wide.
- Bulk-Bulk + Jet-Bulk fits amplitude and width
- The choice of the starting energy for jets affects both the bulk and jet fractions
- Yield is a different observable

Azimuthal Width



- Broad angular width from quenching alone (Shuryak)
- Narrow width from flow alone (Gavin, McLerran, Moschelli)
- \bullet The jet contribution becomes more significant with increasing $\ensuremath{\textbf{p}}_t$
- Width of the combination seems to work

Angular width from

$$\left(\frac{\Delta\rho}{\sqrt{\rho_{_{ref}}}}\right)_{p_t > p_{t,min}}$$

Summary

Bulk correlations are a nontrivial contribution to the hard ridge

- Glasma provides energy, centrality, and system dependence
- Angular width from flow and jet quenching
- Experimental study is needed
 - p_t dependence of soft ridge: Chanaka DeSilva
 - A common observable between the hard and soft ridge could help

Jet Correlation Strength

$$\mathcal{R} = \frac{\langle N(N-1) \rangle - \langle N \rangle^2}{\langle N \rangle^2}$$

$$\mathcal{R}_{_{JB}} = \frac{\langle N_{_J}N_{_B}\rangle - \langle N_{_J}\rangle \langle N_{_B}\rangle}{\langle N_{_J}\rangle \langle N_{_B}\rangle}$$

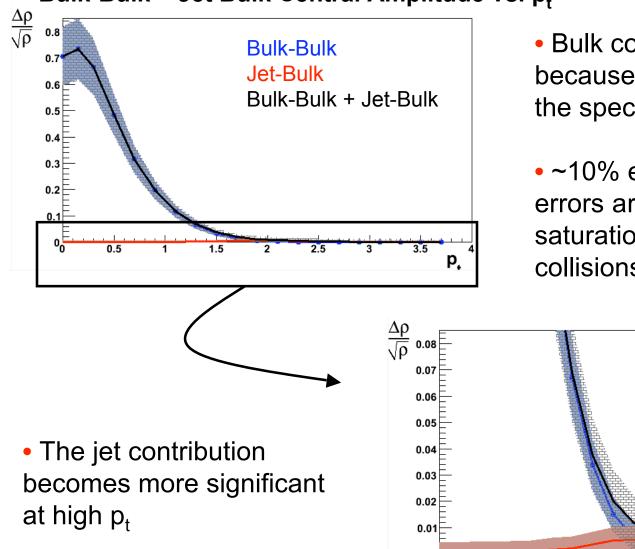
$$=\frac{\alpha\beta\langle N(N-1)\rangle-\alpha\beta\langle N\rangle\langle N\rangle}{\alpha\beta\langle N\rangle\langle N\rangle}$$

$$\begin{split} \langle N_{\scriptscriptstyle B} \rangle &= \beta \langle N \rangle \\ \langle N_{\scriptscriptstyle J} \rangle &= \alpha \langle N \rangle \\ \langle N_{\scriptscriptstyle J} N_{\scriptscriptstyle B} \rangle &= \alpha \beta \langle N(N-1) \rangle \end{split}$$

$$\implies \qquad \mathcal{R}_{_{JB}} = \mathcal{R}$$

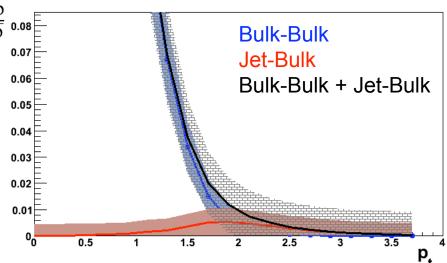
Amplitude vs. p_t



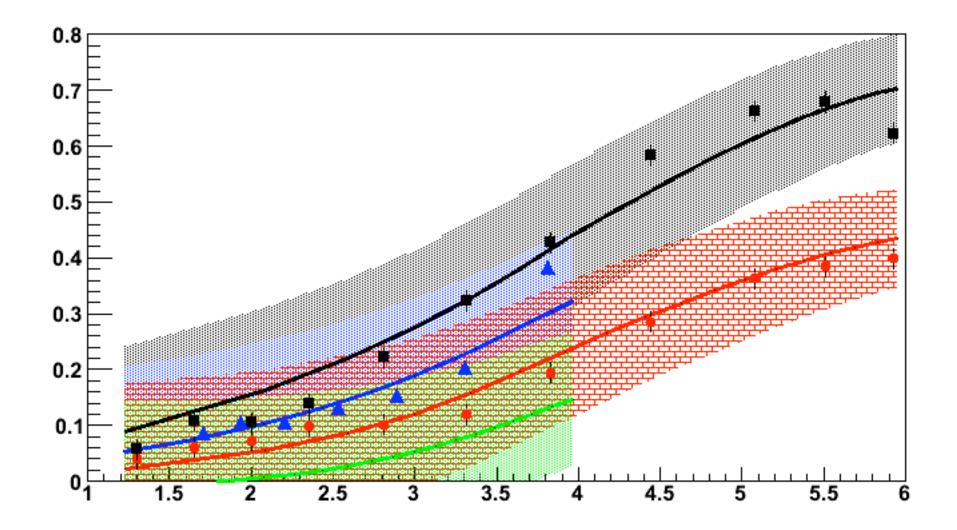


• Bulk correlations dominate because soft particles dominate the spectrum at low p_t

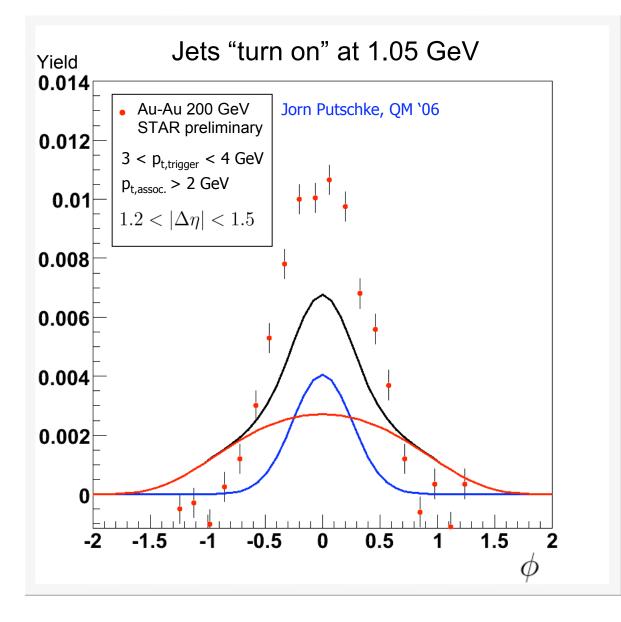
 ~10% error from BW; Glasma errors are at a minimum since saturation is strongest in central collisions



Amplitude Comparison



Hard Ridge



- Bulk-Bulk Yield
- Jet-Bulk Yield
- Bulk-Bulk + Jet-Bulk Yield